

The QCD phase diagram and statistics friendly distributions



A. Bzdak, VK, N. Strodthoff: arXiv:1607.07375

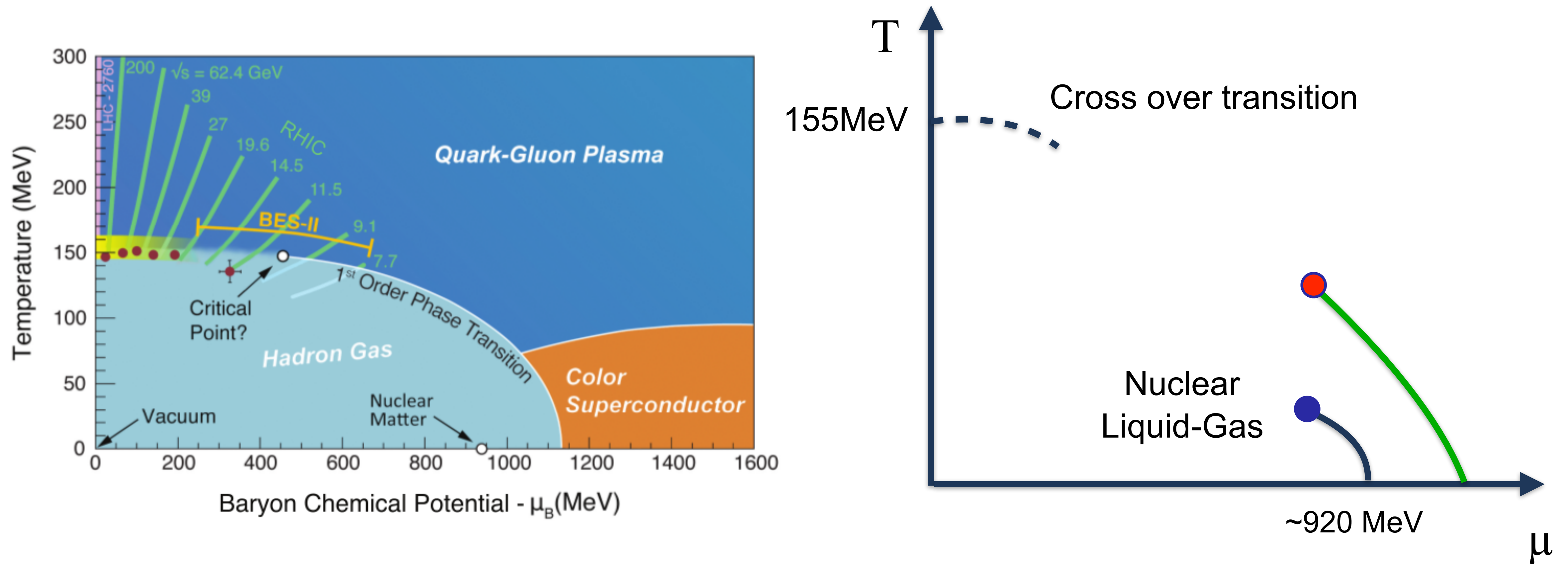
A. Bzdak, VK, V. Skokov: arXiv:1612.05128

A. Bzdak, D. Oliinychenko, J. Steinheimer, VK: arXiv:1804.04463

A. Bzdak and V.K: arXiv:1811.04456

BEST
COLLABORATION

The phase diagram



Cumulants of (Baryon) Number

$$K_n = \frac{\partial^n}{\partial(\mu/T)^n} \ln Z = \frac{\partial^{n-1}}{\partial(\mu/T)^{n-1}} \langle N \rangle$$

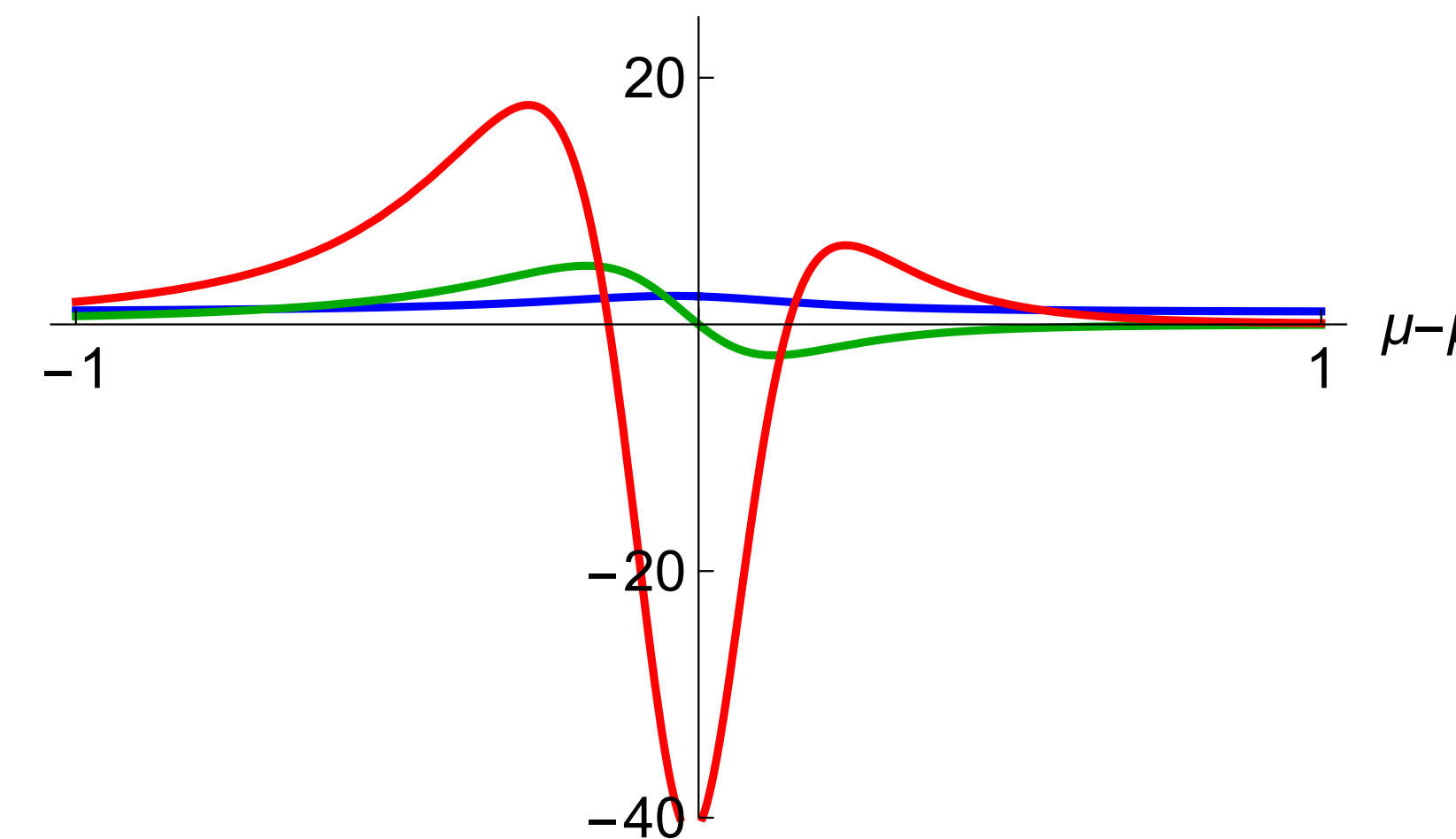
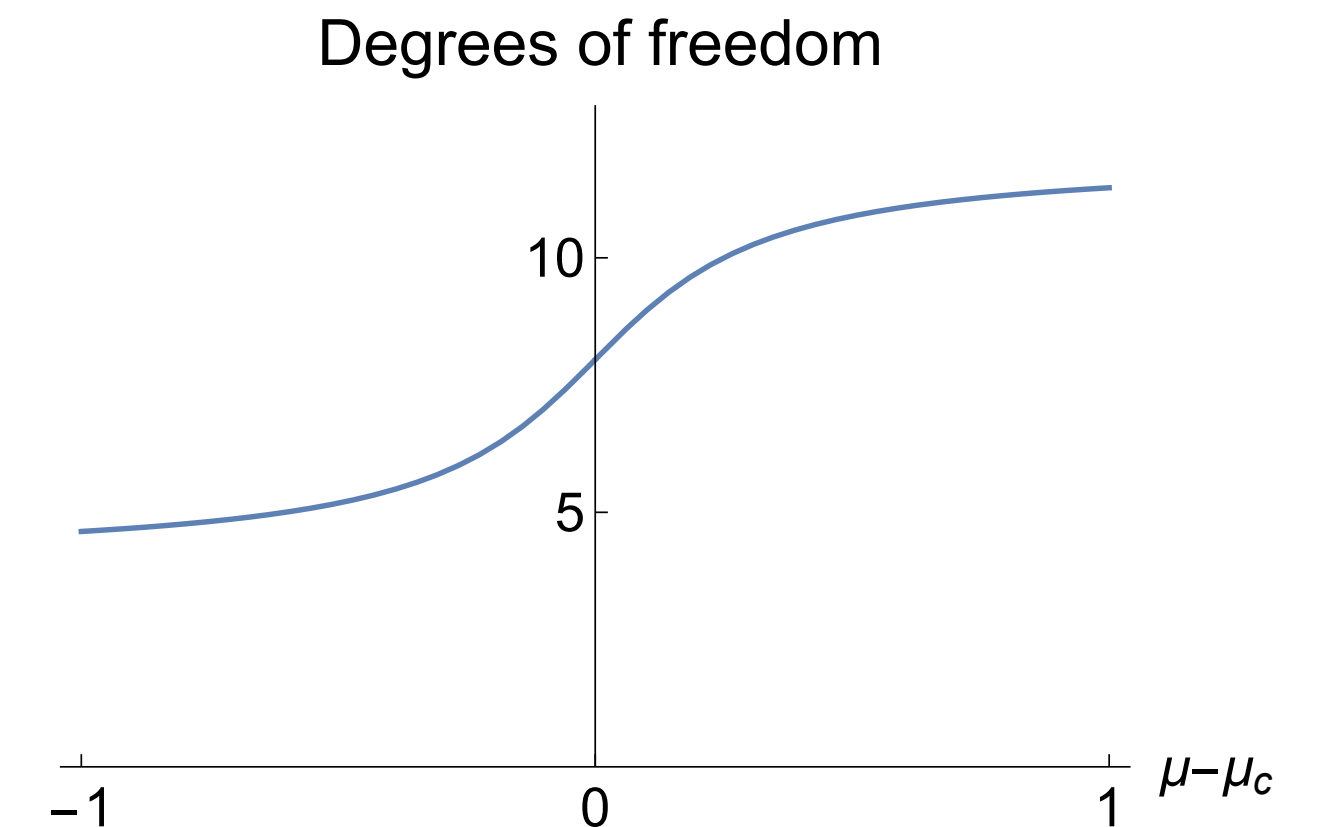
Cumulants measure derivatives of free energy w.r.t chemical potential

$$K_1 = \langle N \rangle, \quad K_2 = \langle N - \langle N \rangle \rangle^2, \quad K_3 = \langle N - \langle N \rangle \rangle^3$$

Cumulants scale with volume (extensive): $K_n \sim V$

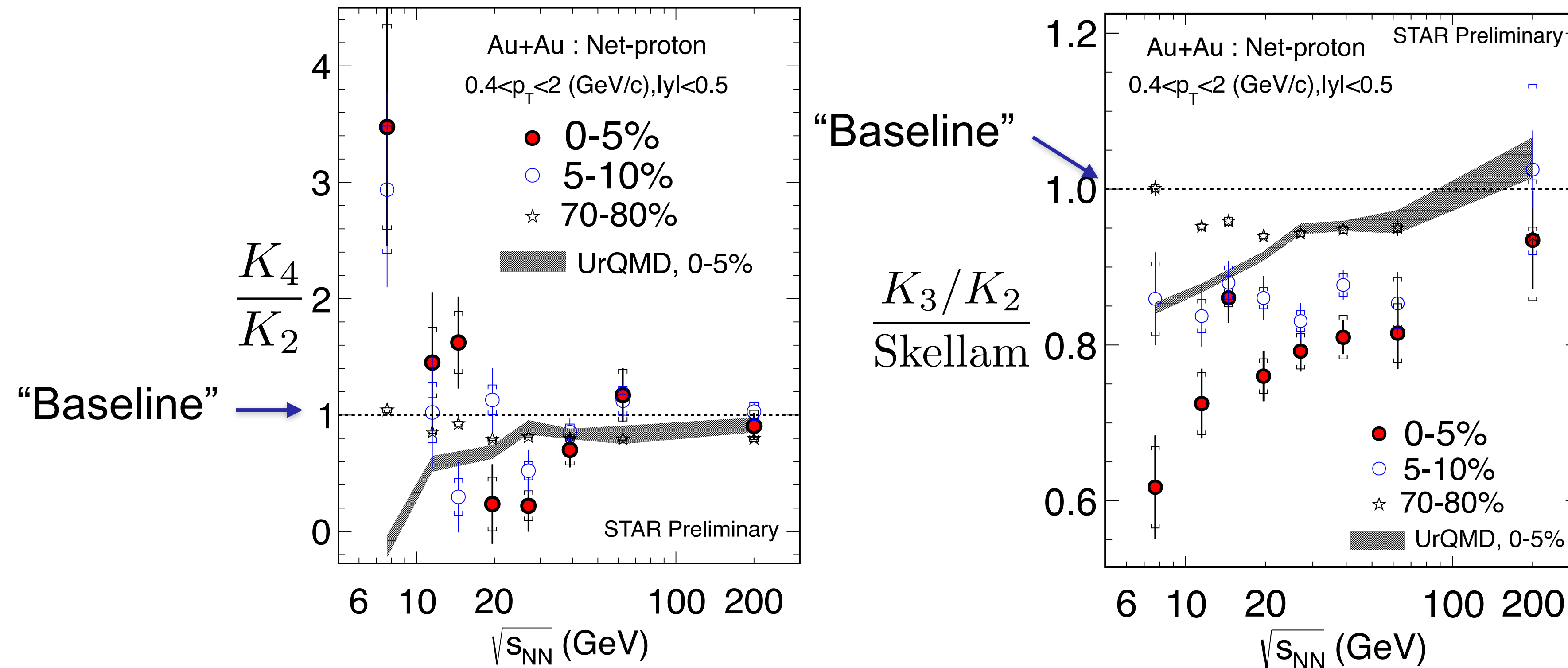
Volume not well controlled in heavy ion collisions

Cumulant Ratios: $\frac{K_2}{\langle N \rangle}, \quad \frac{K_3}{K_2}, \quad \frac{K_4}{K_2}$



Latest STAR result on net-proton cumulants

X. Luo, NPA 956 (2016) 75



K_4/K_2 follows expectation for CP, K_3/K_2 no so much.....
 URQMD totally fails to get trend for K_4/K_2 !

Further insights: Correlations

Cumulants $K_n = \frac{\partial^n}{\partial \hat{\mu}^n} P/T^4$

$$K_2 = \langle N - \langle N \rangle \rangle^2 = \langle (\delta N)^2 \rangle$$

$$\rho_2(p_1, p_2) = \rho_1(p_1)\rho_1(p_2) + C_2(p_1, p_2), \quad \text{C}_2: \text{Correlation Function}$$

$$K_3 = \langle (\delta N)^3 \rangle$$

$$\begin{aligned} \rho_3(p_1, p_2, p_3) = & \rho_1(p_1)\rho_1(p_2)\rho_1(p_3) + \rho_1(p_1)\underline{C_2(p_2, p_3)} + \rho_1(p_2)\underline{C_2(p_1, p_3)} \\ & + \rho_1(p_3)\underline{C_2(p_1, p_2)} + \underline{C_3(p_1, p_2, p_3)} \end{aligned}$$

From Cumulants to Correlations

Defining integrated correlations function a.k.a **factorial cumulants**

$$C_n = \int dp_1 \dots dp_n C_n(p_1, \dots, p_n)$$

Simple Algebra leads to relation between correlations C_n and K_n

$$C_2 = -K_1 + K_2,$$

$$C_3 = 2K_1 - 3K_2 + K_3,$$

$$C_4 = -6K_1 + 11K_2 - 6K_3 + K_4, .$$

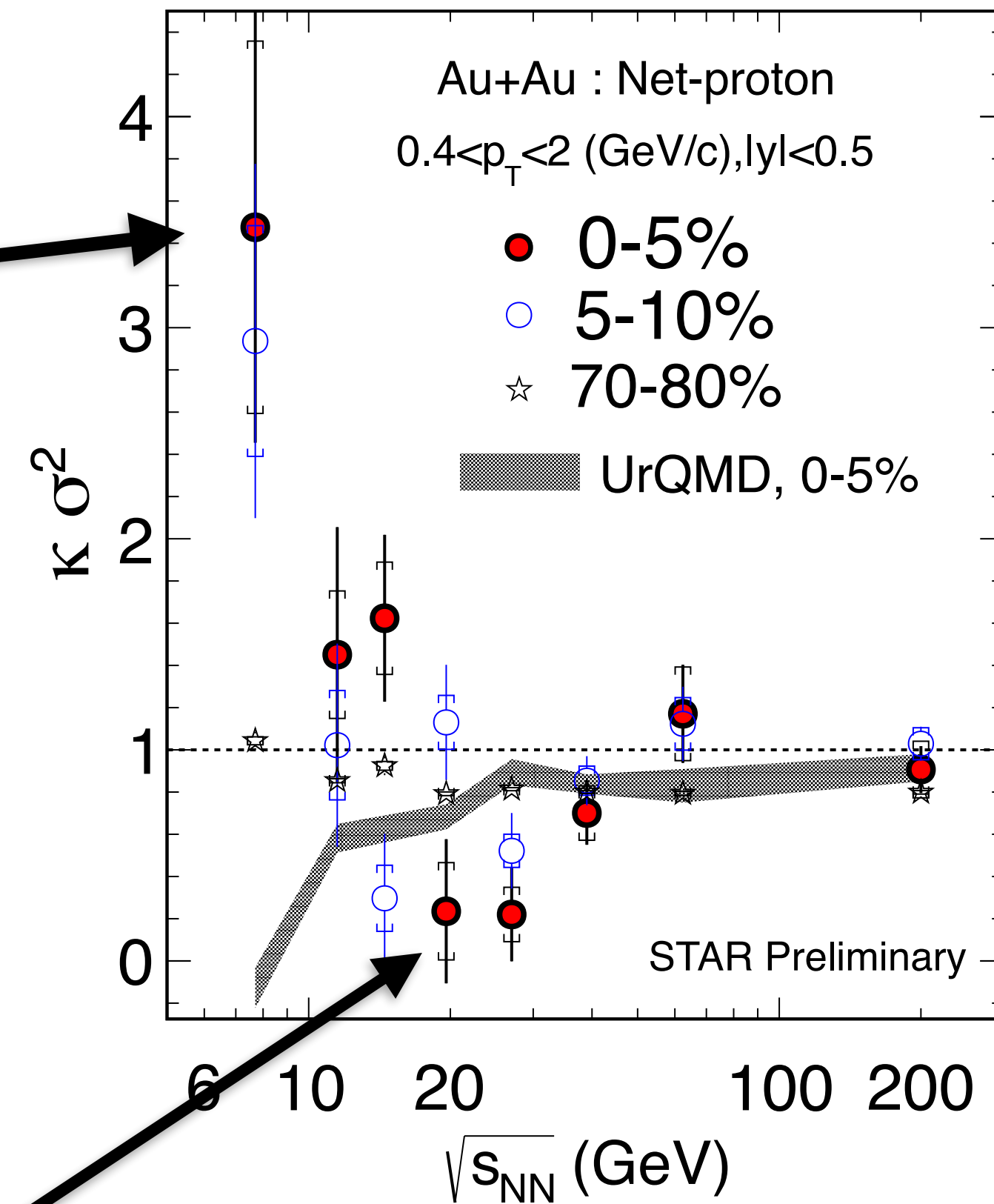
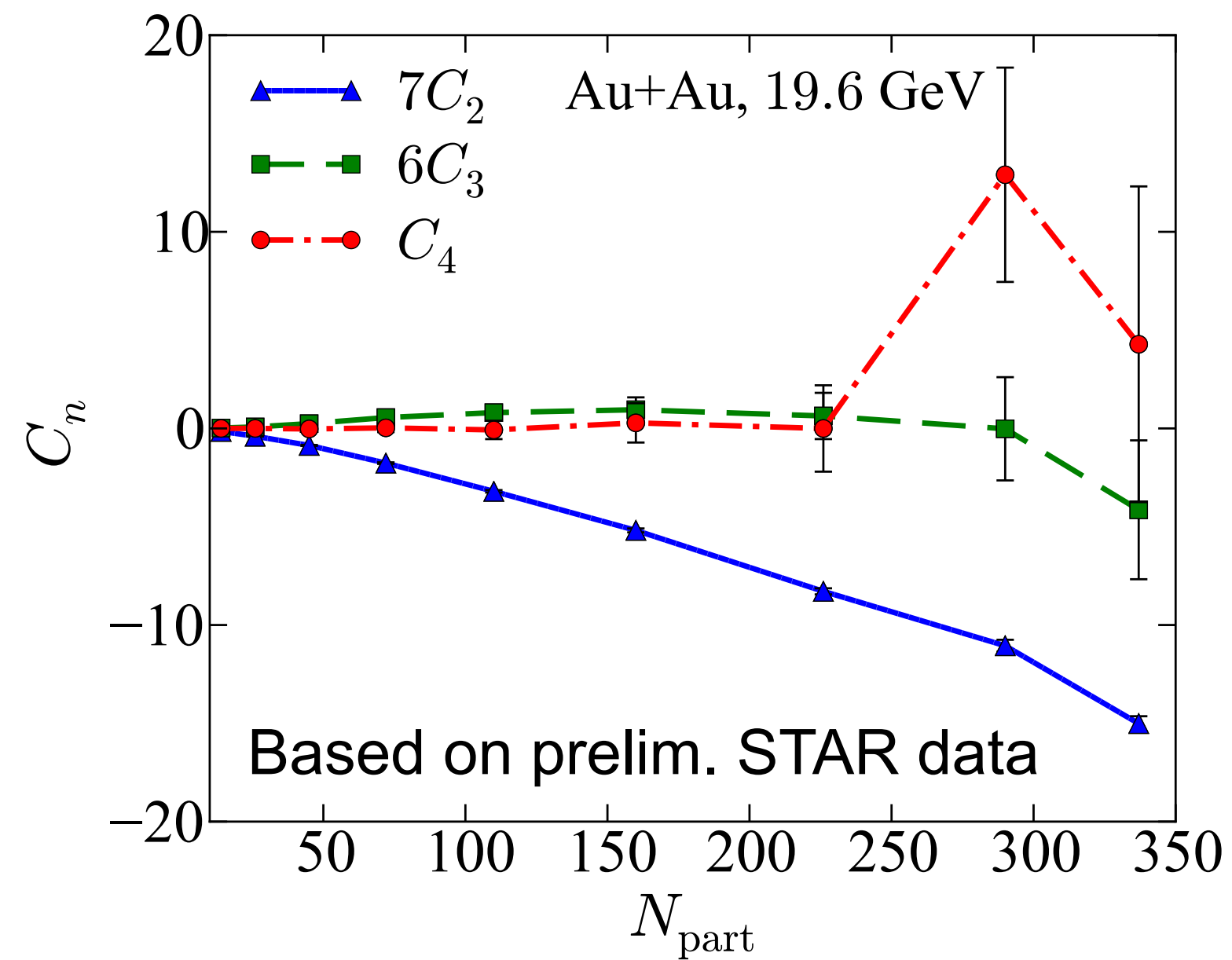
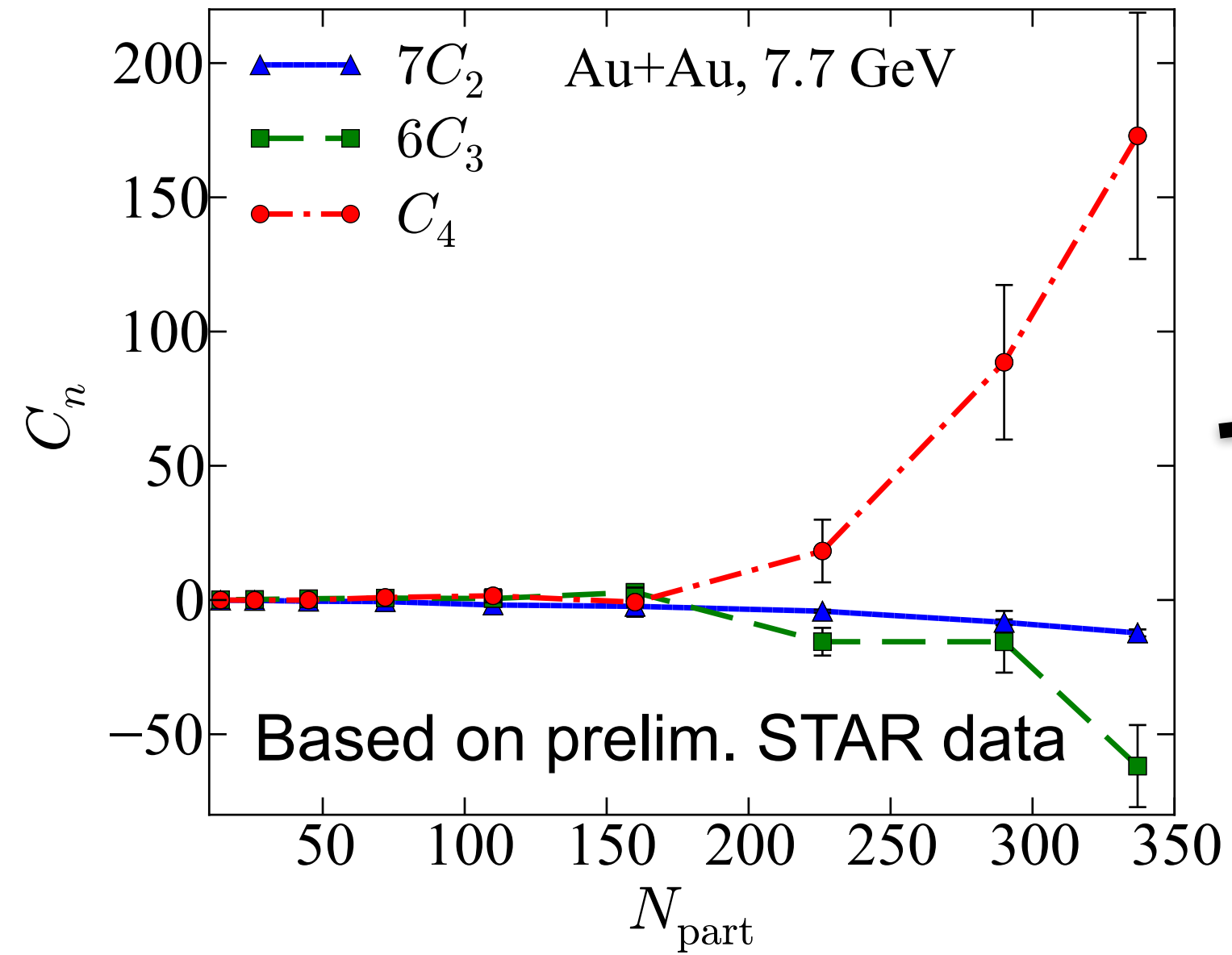
or vice versa

$$K_2 = \langle N \rangle + C_2$$

$$K_3 = \langle N \rangle + 3C_2 + C_3$$

$$K_4 = \langle N \rangle + 7C_2 + 6C_3 + C_4$$

Correlations



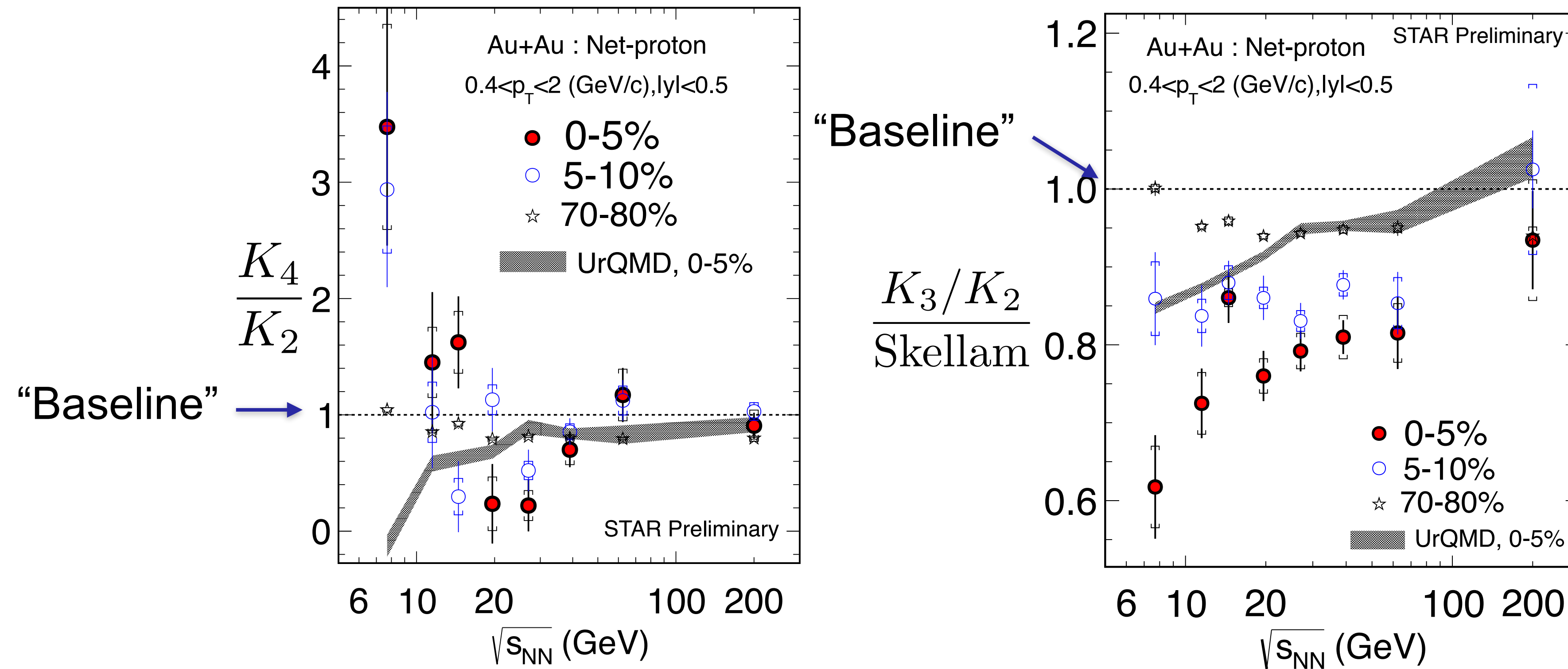
Dip at 19.6 GeV from
NEGATIVE C_2 !

Correlations

- Four particles correlations kick in for $E < 15$ GeV
- Data are consistent with long range correlations in rapidity and transverse momentum $\Rightarrow |C_n| \sim \langle N \rangle^n$
- fourth order correlations are large
 - effects from participant fluctuations $\sim 10^{-3}$
 - Cluster model:
 - magnitude requires that 28 out of 30 protons result from 4-proton clusters
- Cluster model is not able to describe both C_4 and C_3

Latest STAR result on net-proton cumulants

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K_4/K_2 above baseline K_3/K_2 below baseline

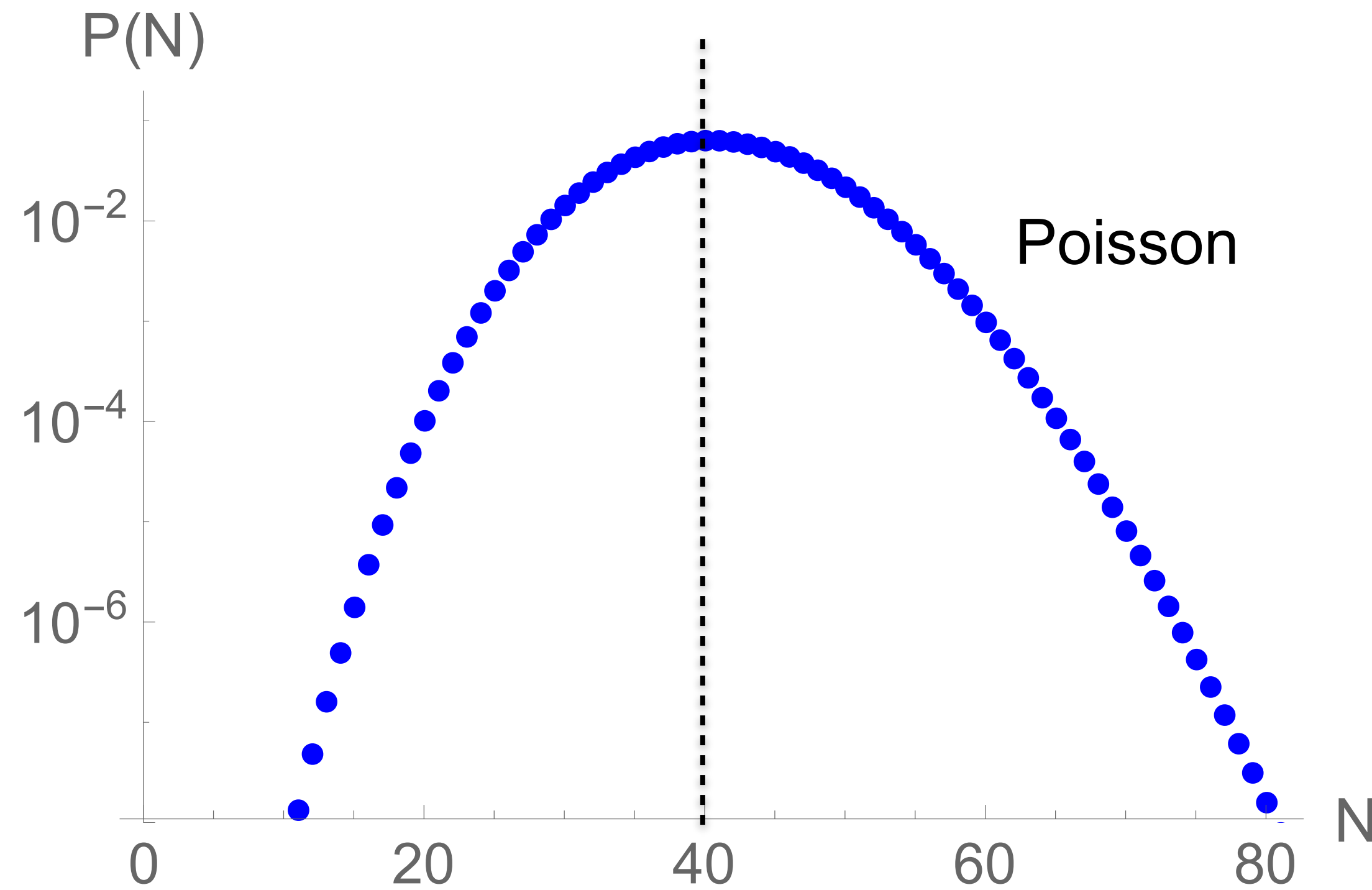
Shape of probability distribution

$$K_3 < \langle N \rangle$$

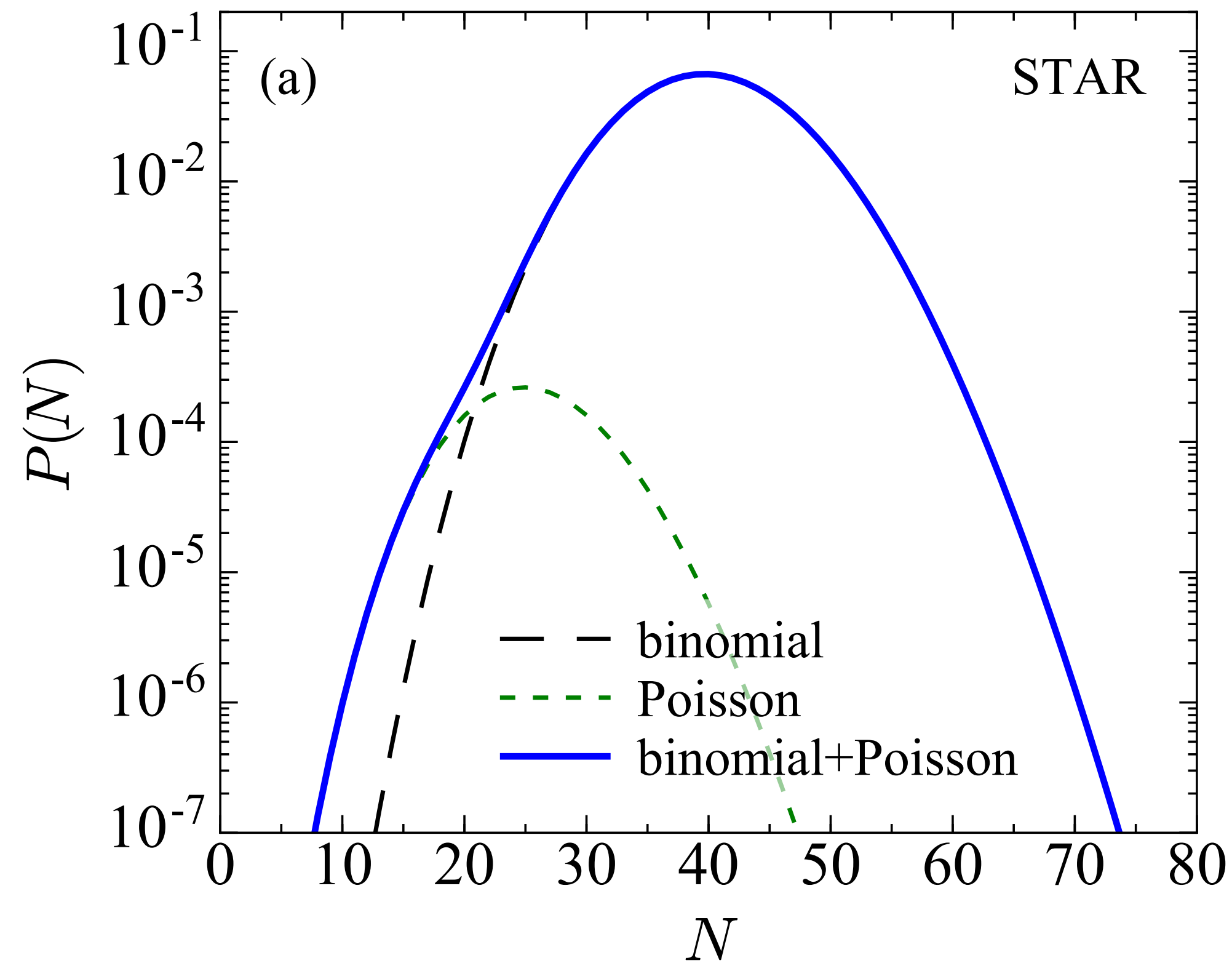
$$K_4 > \langle N \rangle$$

$$K_3 = \langle N - \langle N \rangle \rangle^3$$

$$K_4 = \langle N - \langle N \rangle \rangle^4 - 3 \langle N - \langle N \rangle \rangle^2$$



Simple two component model



Weight of small component: $\sim 0.3\%$

Left bump not visible raw (efficiency) uncorrected distribution

Two component model

$$P(N) = (1 - \alpha)P_{(a)}(N) + \alpha P_{(b)}(N)$$

$$\bar{N} = \langle N_{(a)} \rangle - \langle N_{(b)} \rangle$$

$$C_2 = C_2^{(a)} - \alpha \{ \bar{C}_2 - (1 - \alpha) \bar{N}^2 \}$$

$$C_3 = C_3^{(a)} - \alpha \{ \bar{C}_3 + (1 - \alpha) [(1 - 2\alpha) \bar{N}^3 - 3\bar{N} \bar{C}_2] \}$$

$$C_4 = C_4^{(a)} - \alpha \{ \bar{C}_4 - (1 - \alpha) [(1 - 6\alpha + 6\alpha^2) \bar{N}^4 - 6(1 - 2\alpha) \bar{N}^2 \bar{C}_2 + 4\bar{N} \bar{C}_3 + 3(\bar{C}_2)^2] \}$$

$$\bar{C}_n = C_n^{(a)} - C_n^{(b)},$$

For Poisson, $C_{(a)}, C_{(b)}=0$

Fit to STAR data: $\langle N_{(a)} \rangle \simeq 40, \quad \langle N_{(b)} \rangle \simeq 25, \quad \alpha \simeq 0.003$

Two component model

$$P(N) = (1 - \alpha)P_{(a)}(N) + \alpha P_{(b)}(N)$$

$$\bar{N} = \langle N_{(a)} \rangle - \langle N_{(b)} \rangle > 0$$

For $P_{(a)}$, $P_{(b)}$ Poisson, or (to good approximation) Binomial

$$C_n = (-1)^n K_n^B \bar{N}^n \quad n \geq 2$$

K_n^B : Cumulant of Bernoulli distribution

$$\alpha \ll 1, K_n^B = \alpha \Rightarrow C_n \simeq \alpha (-1)^n \bar{N}^n$$

$\Rightarrow |C_n| \sim \langle N \rangle^n$ as seen by STAR (i.e. “infinite” correlation length)

predict:

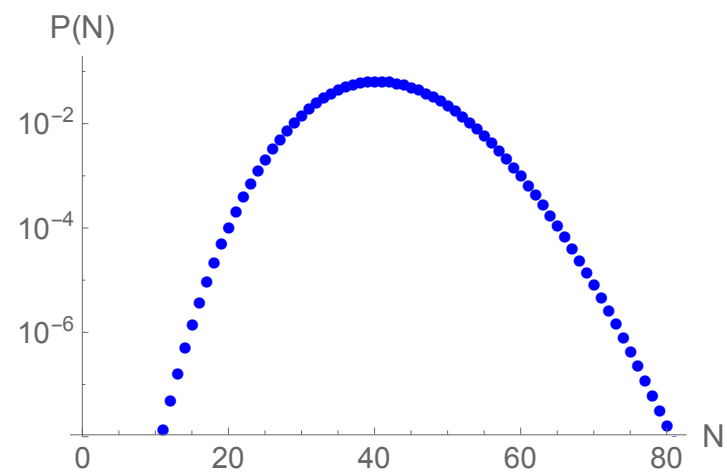
$$\frac{C_4}{C_3} = \frac{C_5}{C_4} = \frac{C_{n+1}}{C_n} = -\bar{N}$$

$$\bar{N} \simeq 15$$

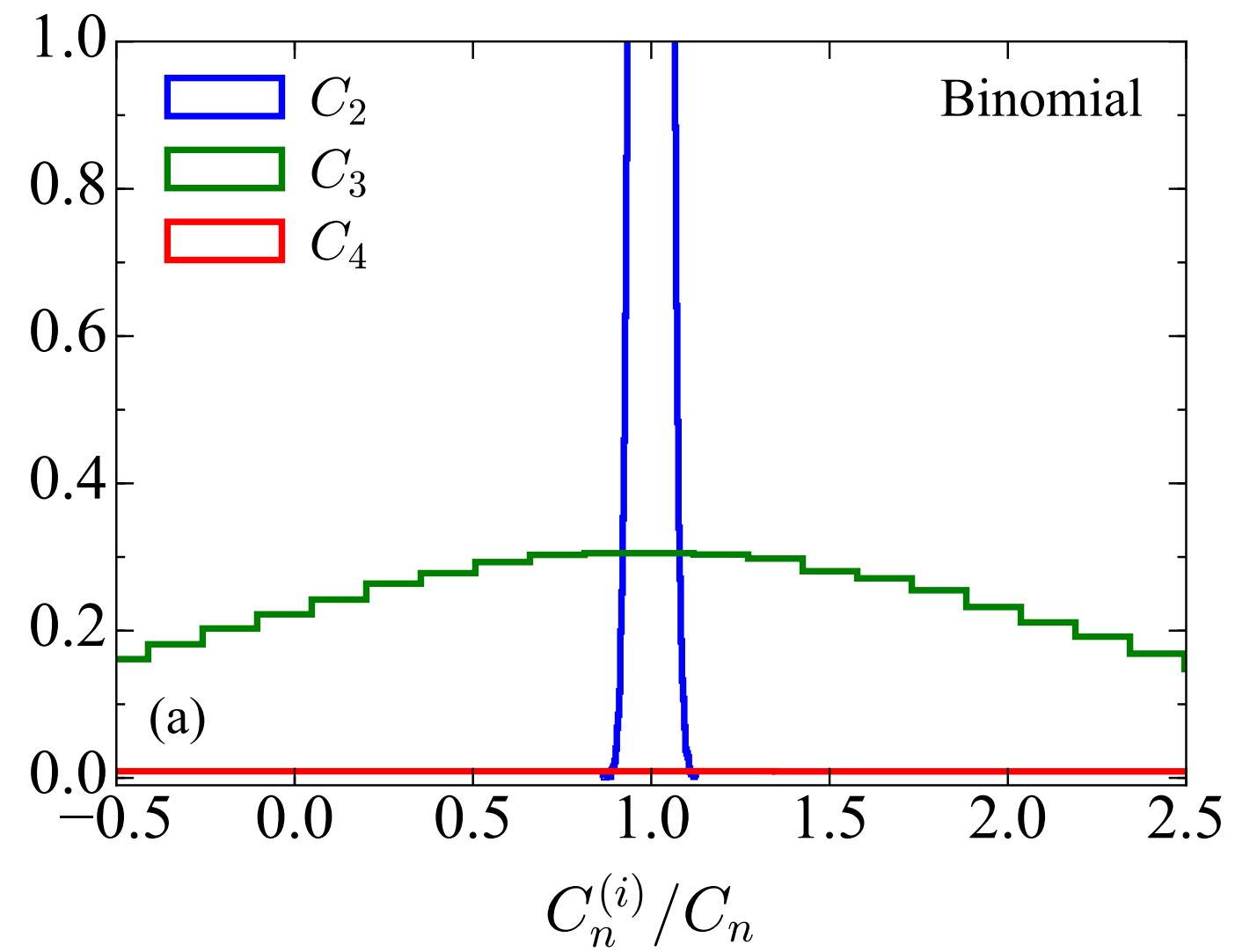
Clear and falsifiable prediction:

$$C_5 \approx -2650 \quad C_6 \approx 41000$$

Two component model is “statistics friendly”

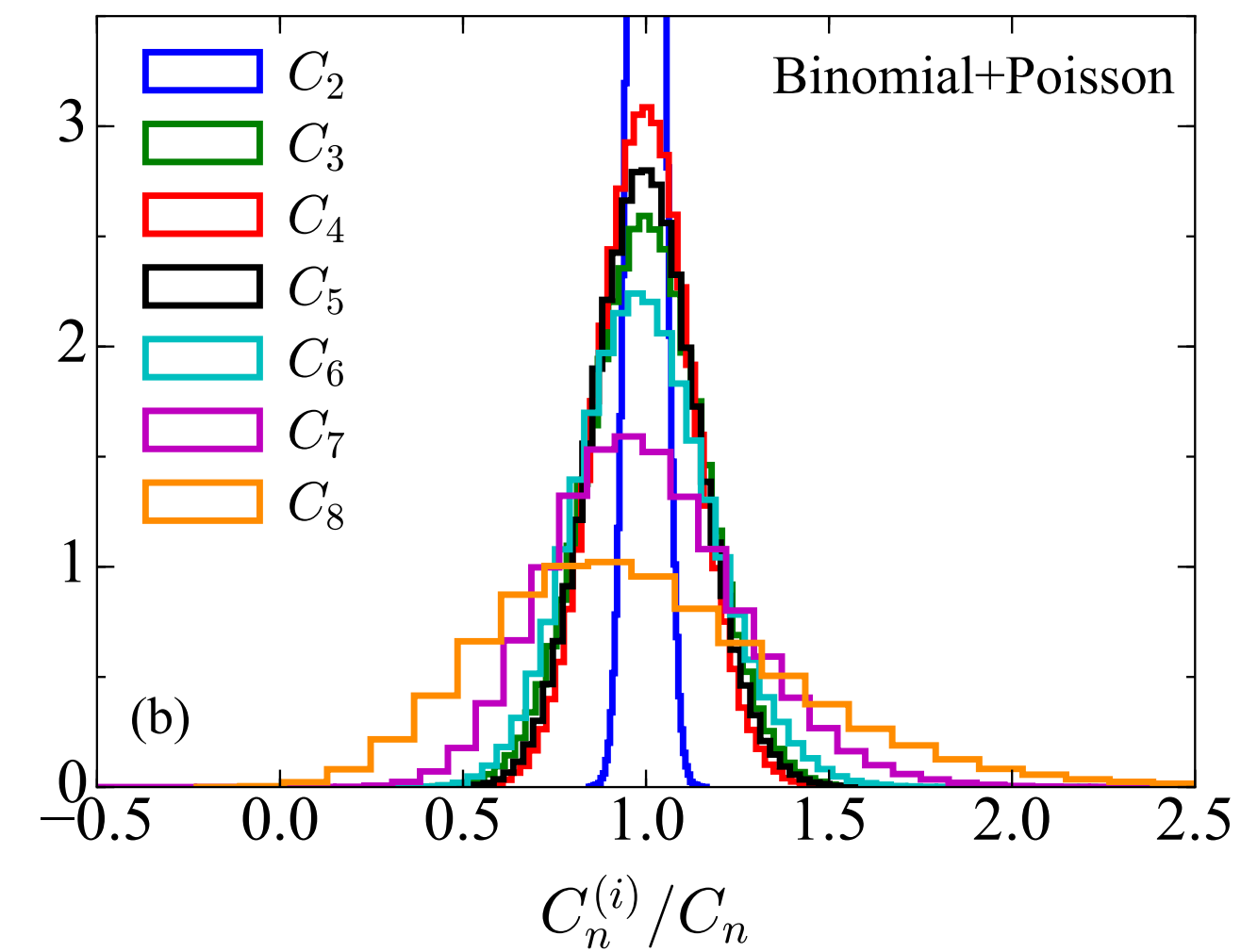


One component

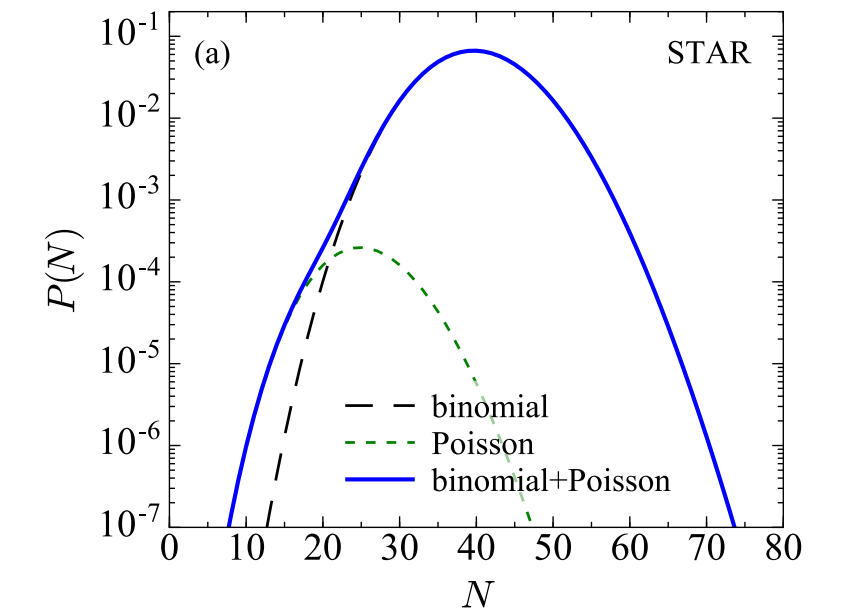


Single component (binomial)

Two component

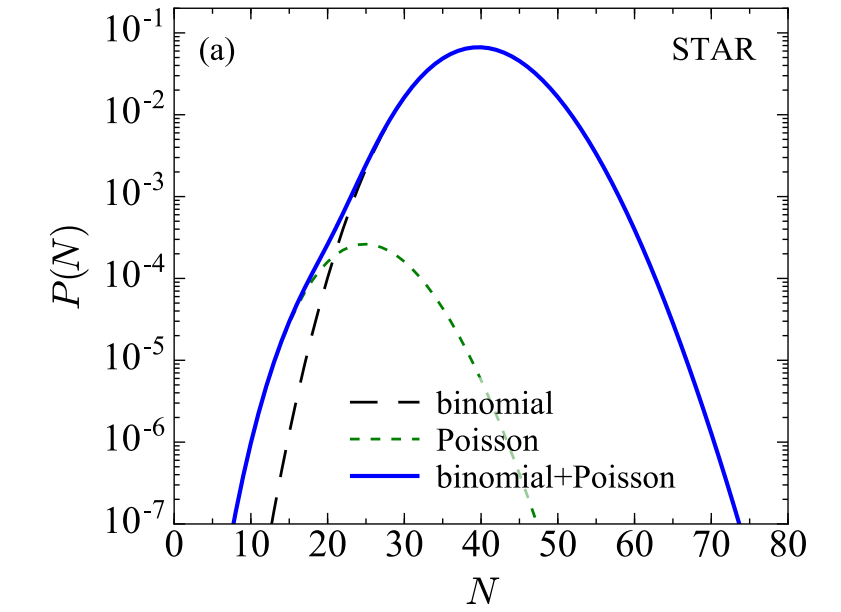
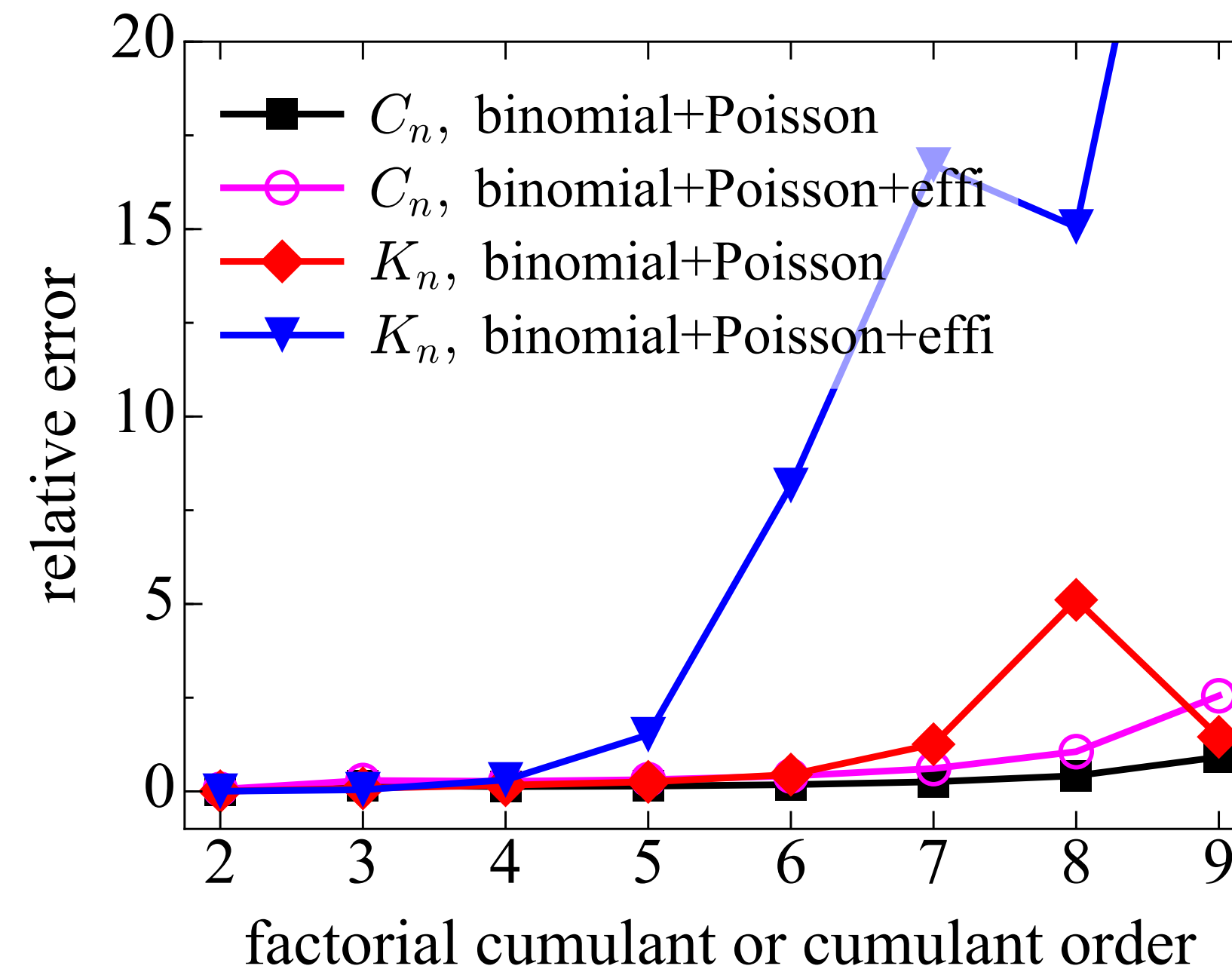
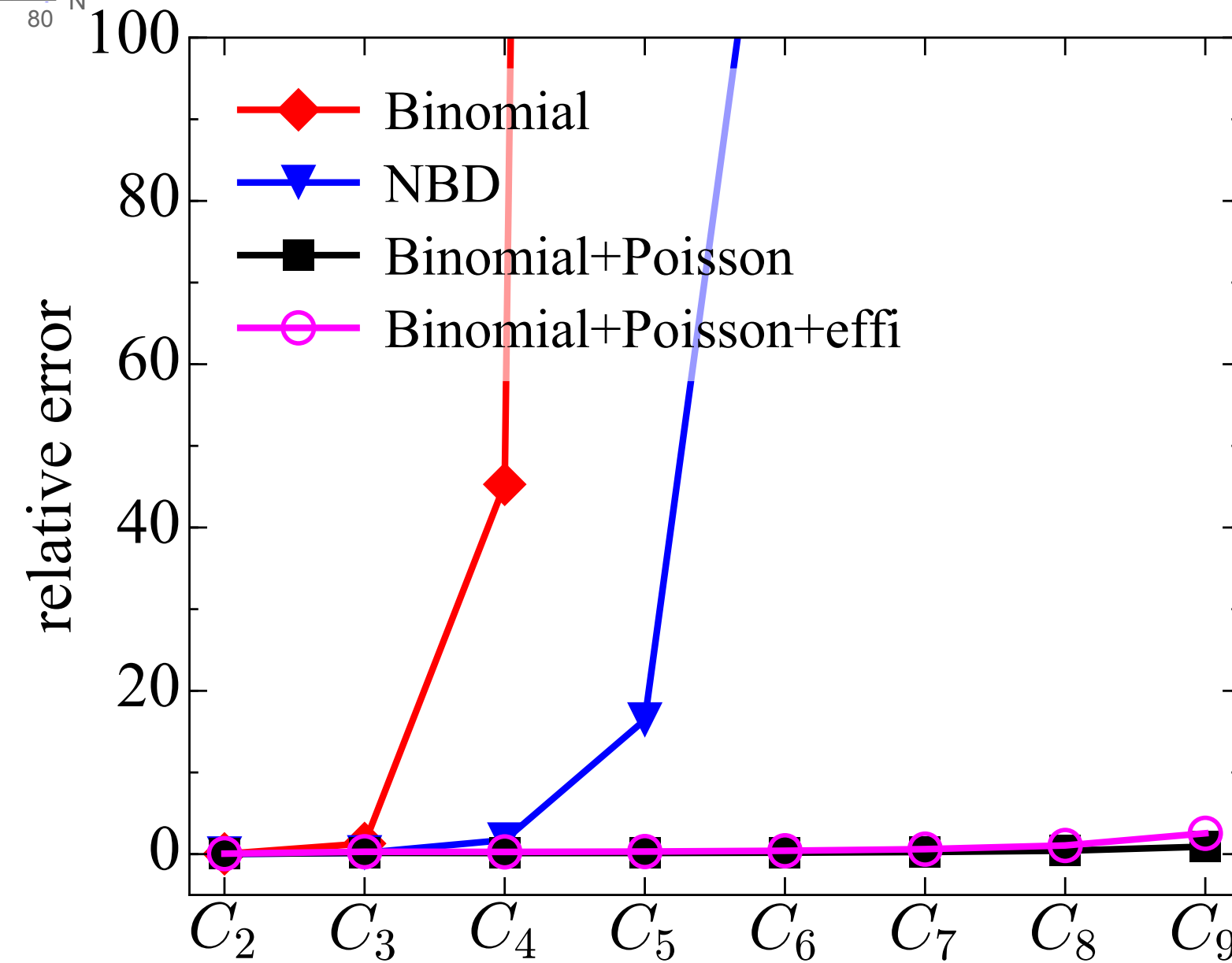
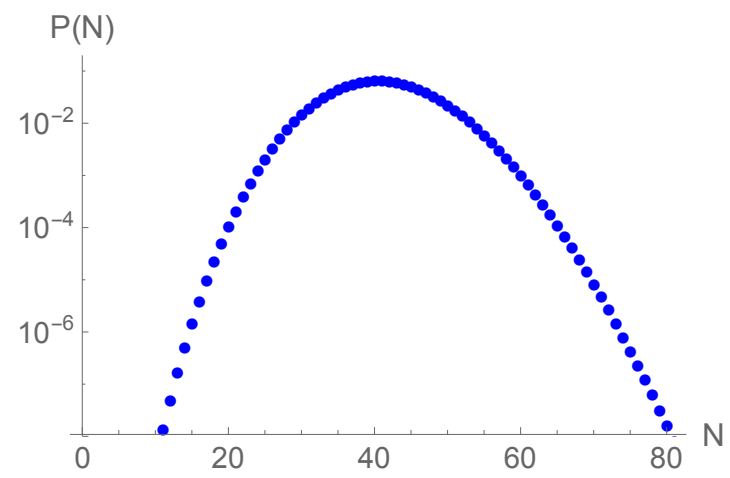


Two component (binomial + Poisson)



Based on 144393 events (same as STAR 0-5% at 7.7 GEV)

Two component model is “statistics friendly”



Cumulants are less “statistics friendly”

Based on 144393 events (same as STAR 0-5% at 7.7 GEV)

This model can be tested RIGHT NOW!

Model prediction:

$$\begin{aligned} C_5 &= -2645 (1 \pm 0.14), & C_6 &= 40900 (1 \pm 0.18), \\ C_7 &= -615135 (1 \pm 0.26), & C_8 &= 8520220 (1 \pm 0.42) \end{aligned}$$

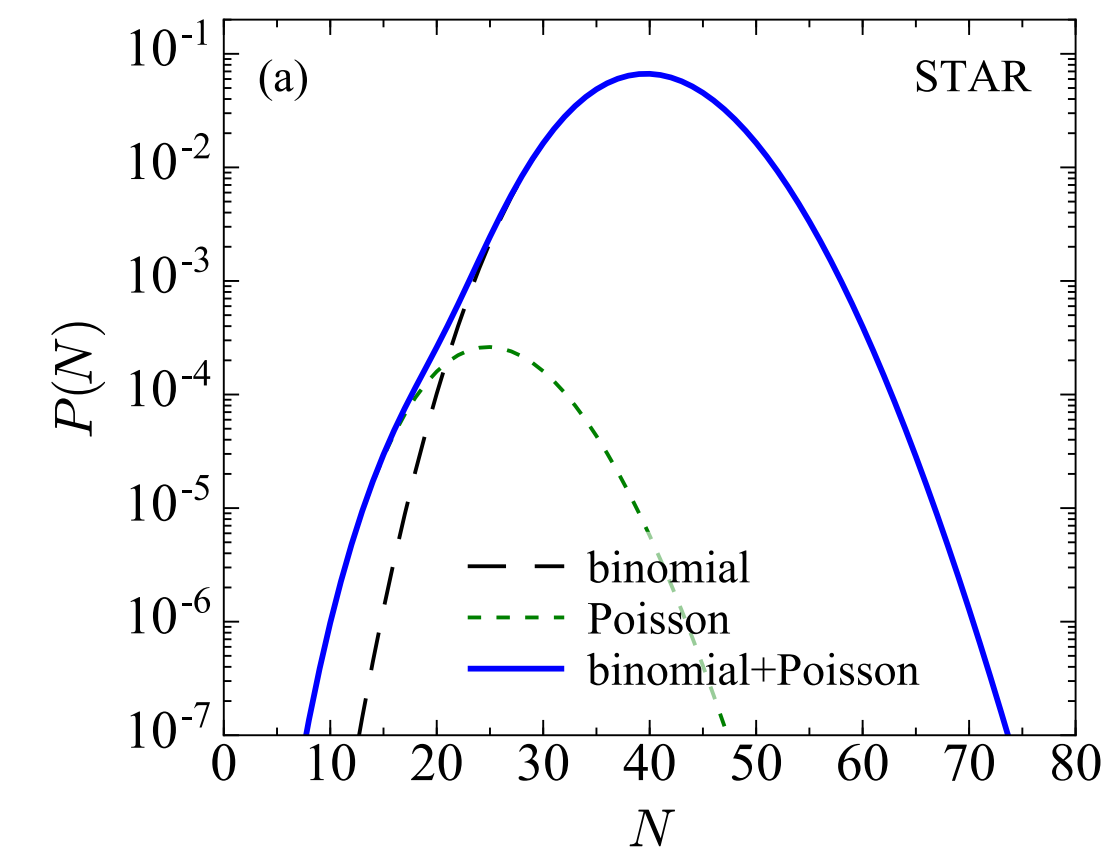
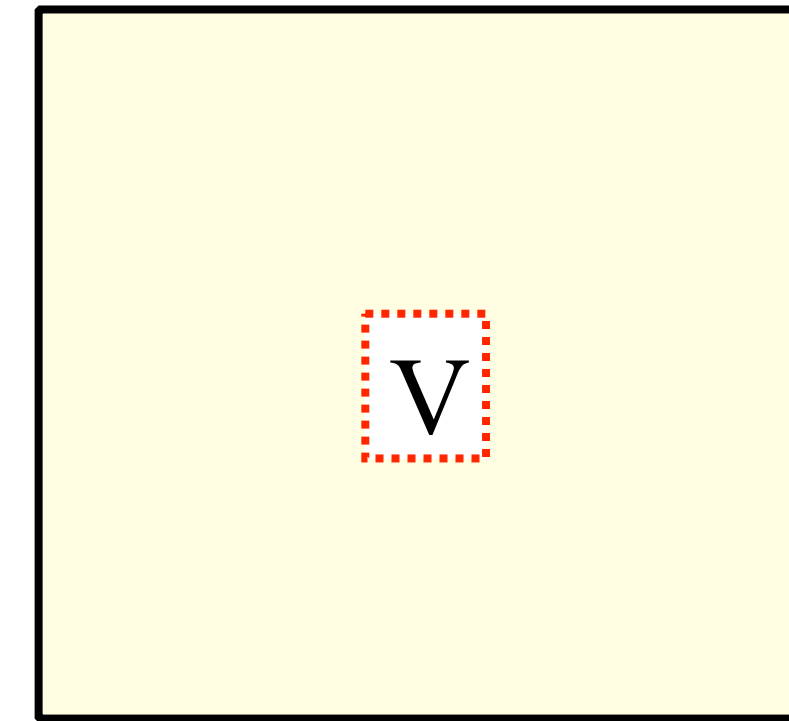
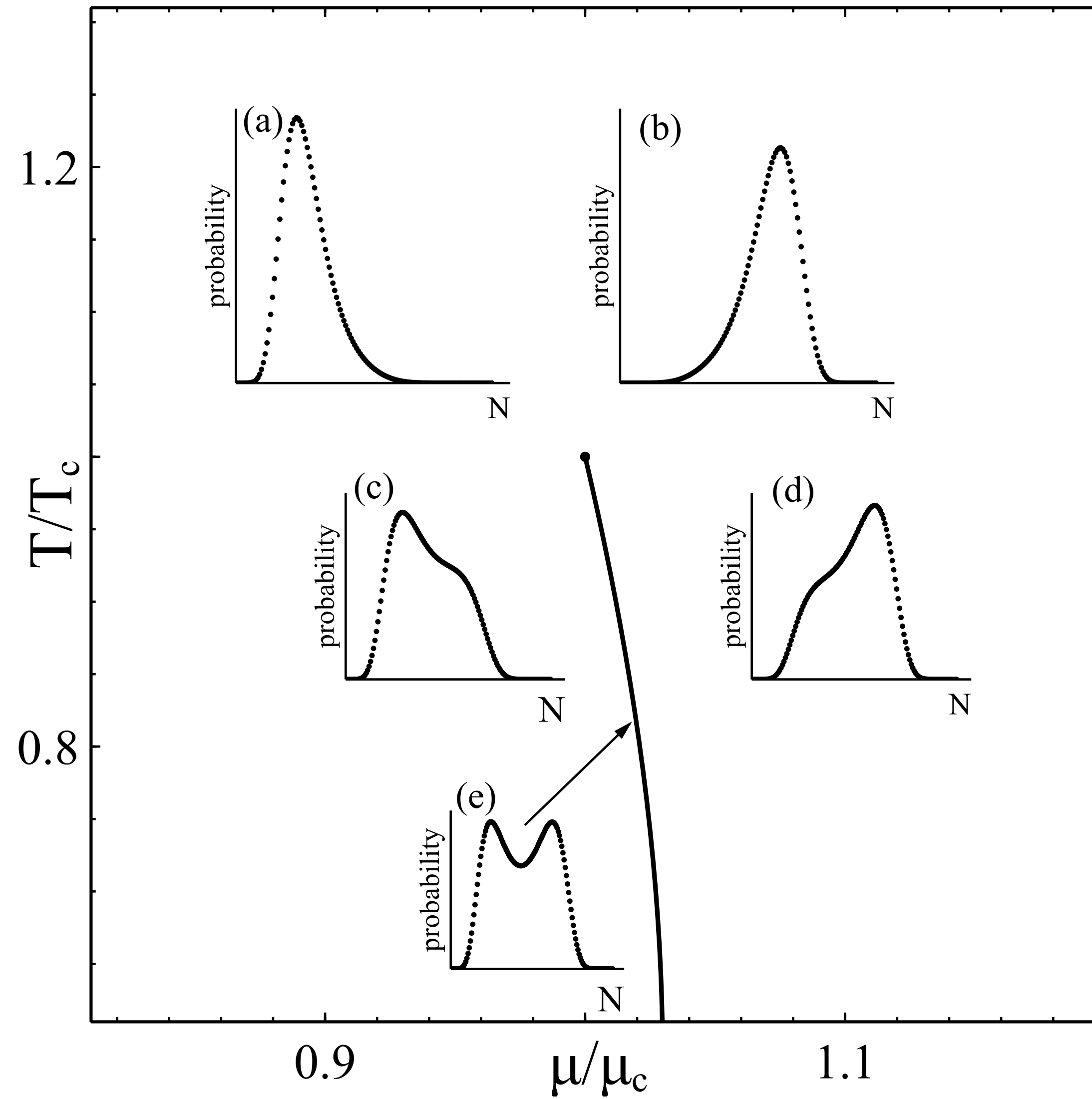
Efficiency
corrected

$$\begin{aligned} C_5 &= -307 (1 \pm 0.31), & C_6 &= 3085 (1 \pm 0.41), \\ C_7 &= -30155 (1 \pm 0.61), & C_8 &= 271492 (1 \pm 1.06), \end{aligned}$$

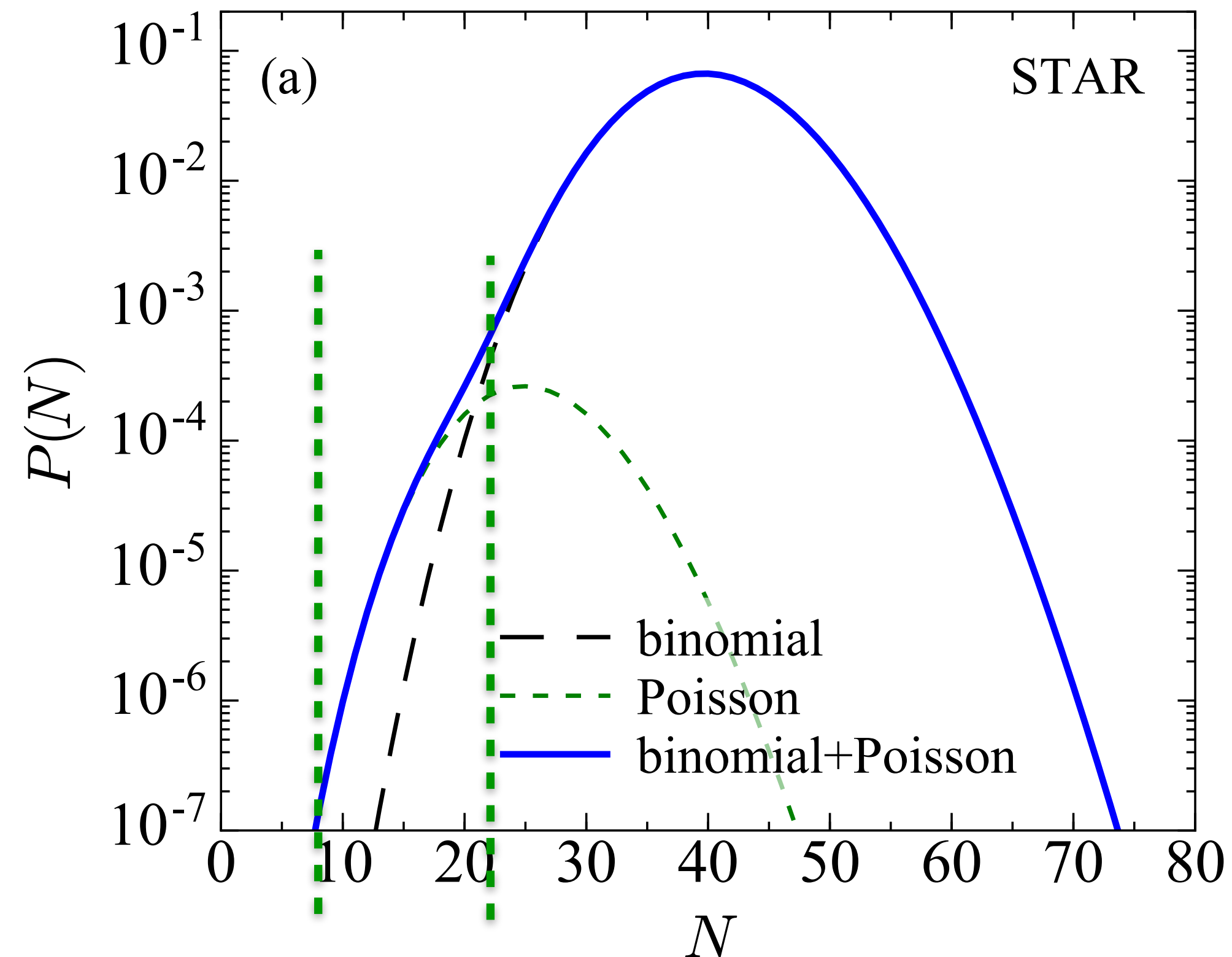
Efficiency
UN-corrected

Based on 144393 events (same as STAR 0-5% at 7.7 GEV)

Speculation



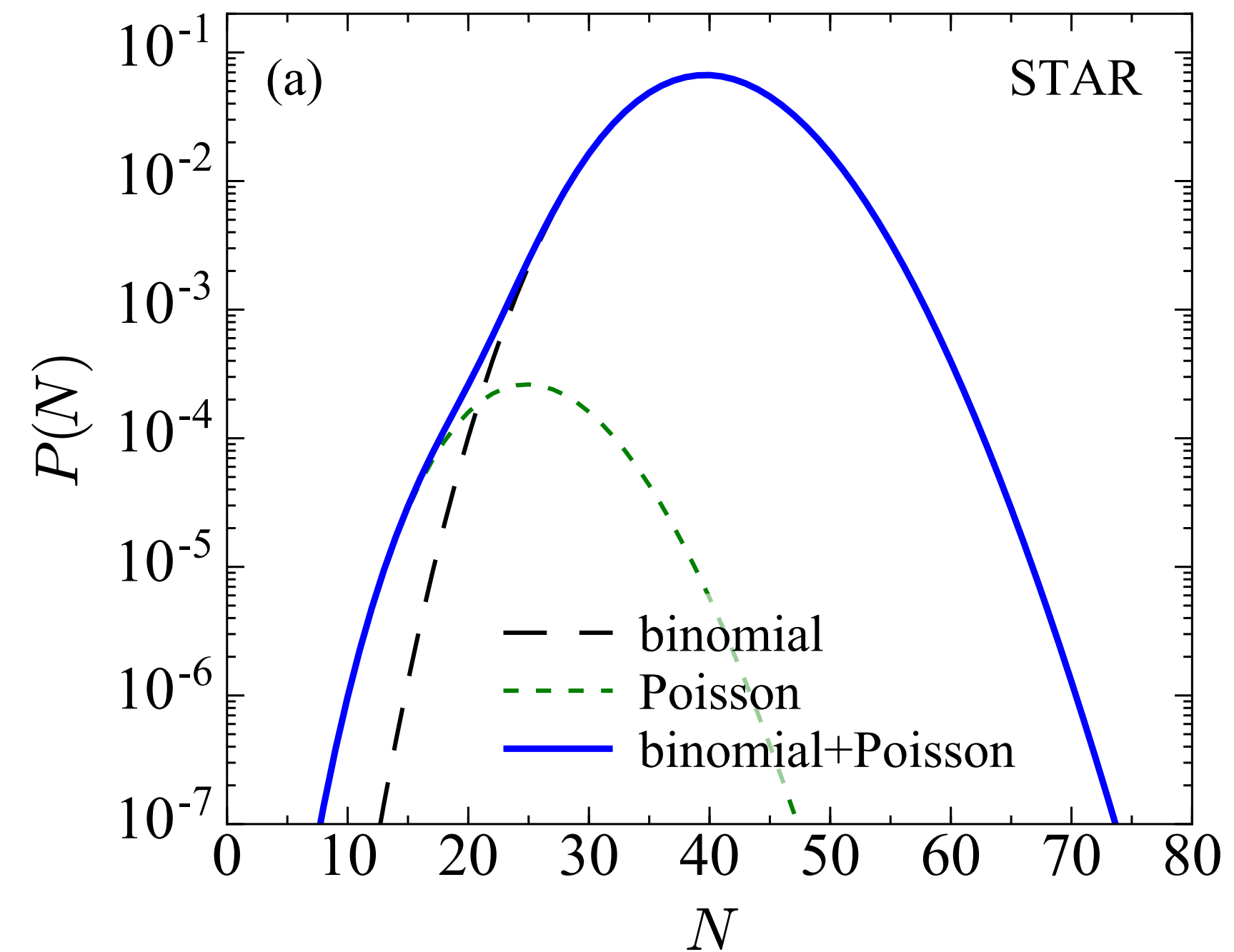
Simple two component model



Analyse data for $N_p < 20$

- Is flow etc different?
- “Inspect by eye (<1% of all events)

Summary



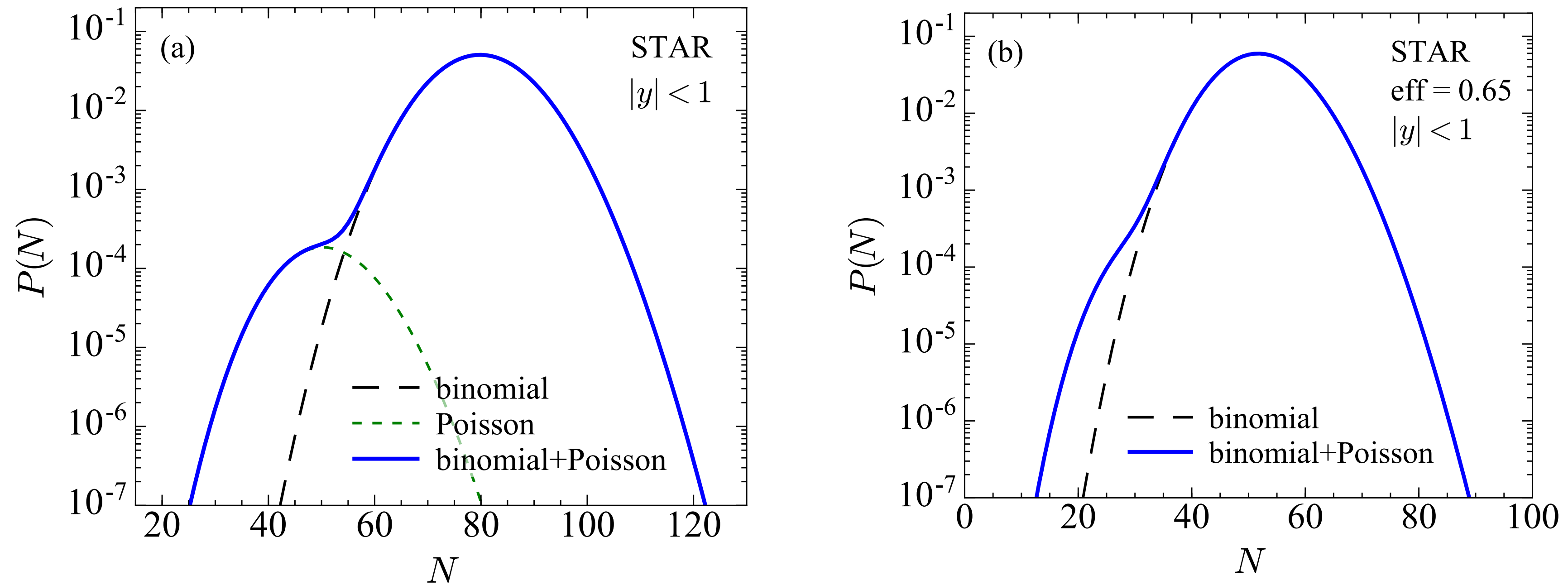
Prediction:

$$C_5 = -2645 (1 \pm 0.14), \quad C_6 = 40900 (1 \pm 0.18),$$

$$C_7 = -615135 (1 \pm 0.26), \quad C_8 = 8520220 (1 \pm 0.42)$$

Thank You

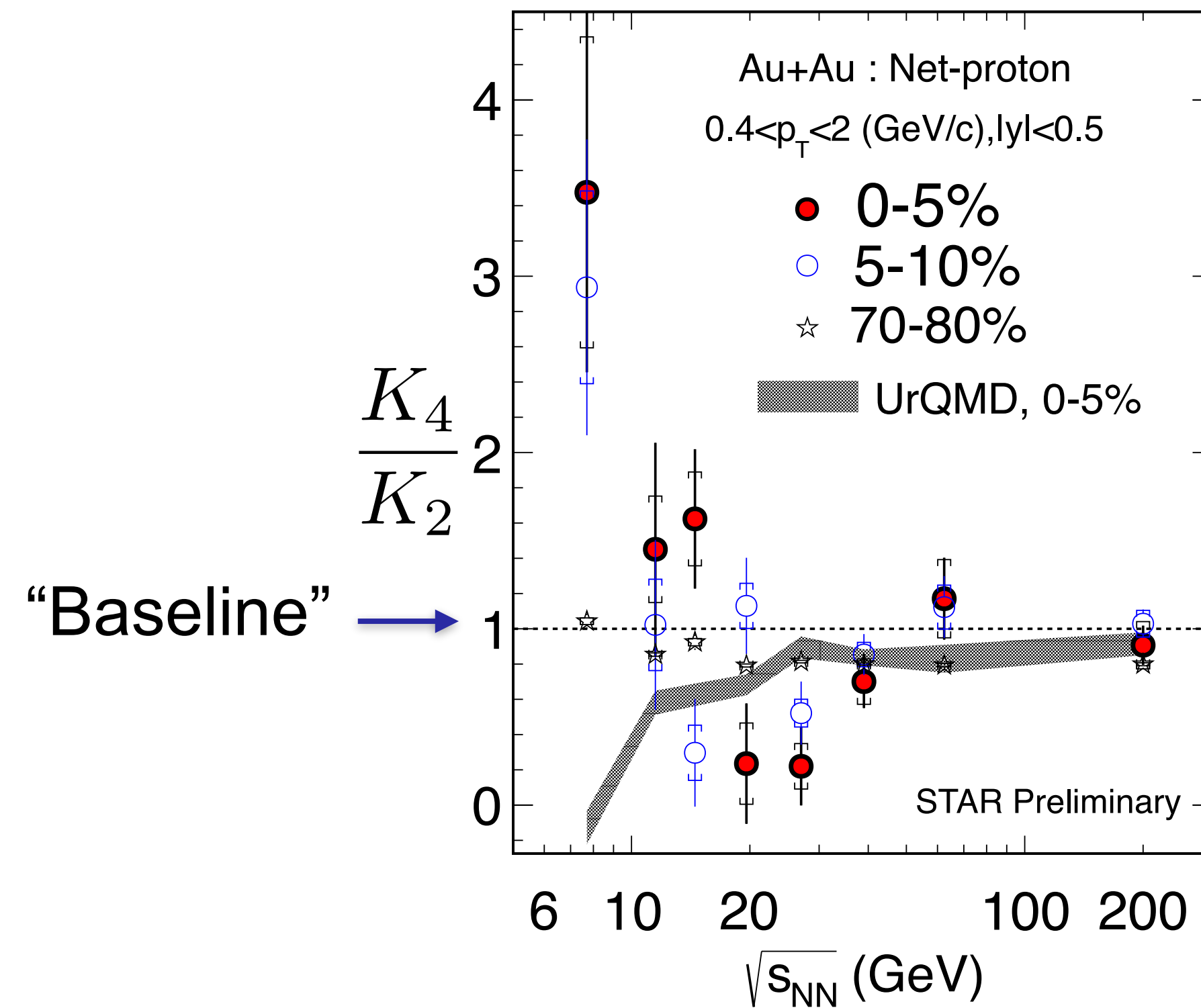
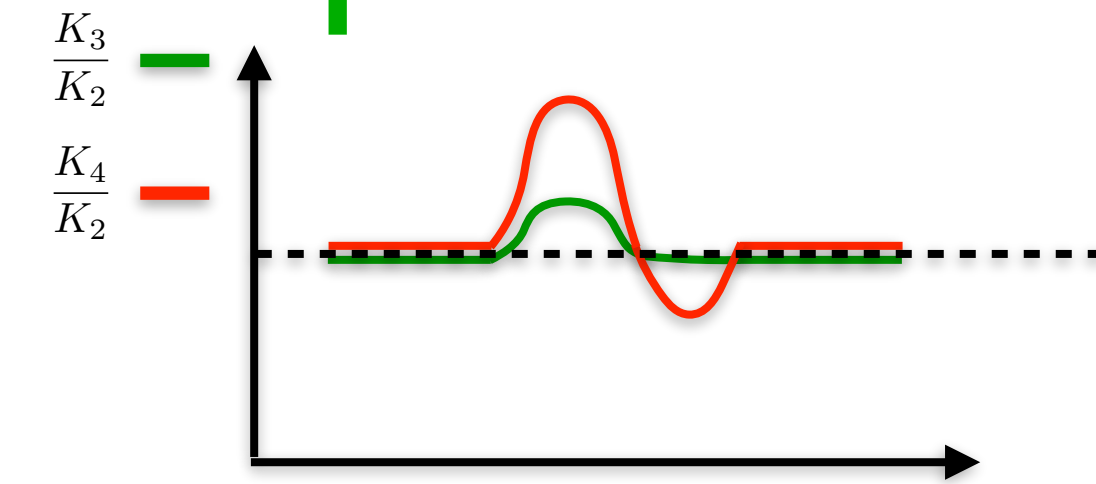
Double the acceptance



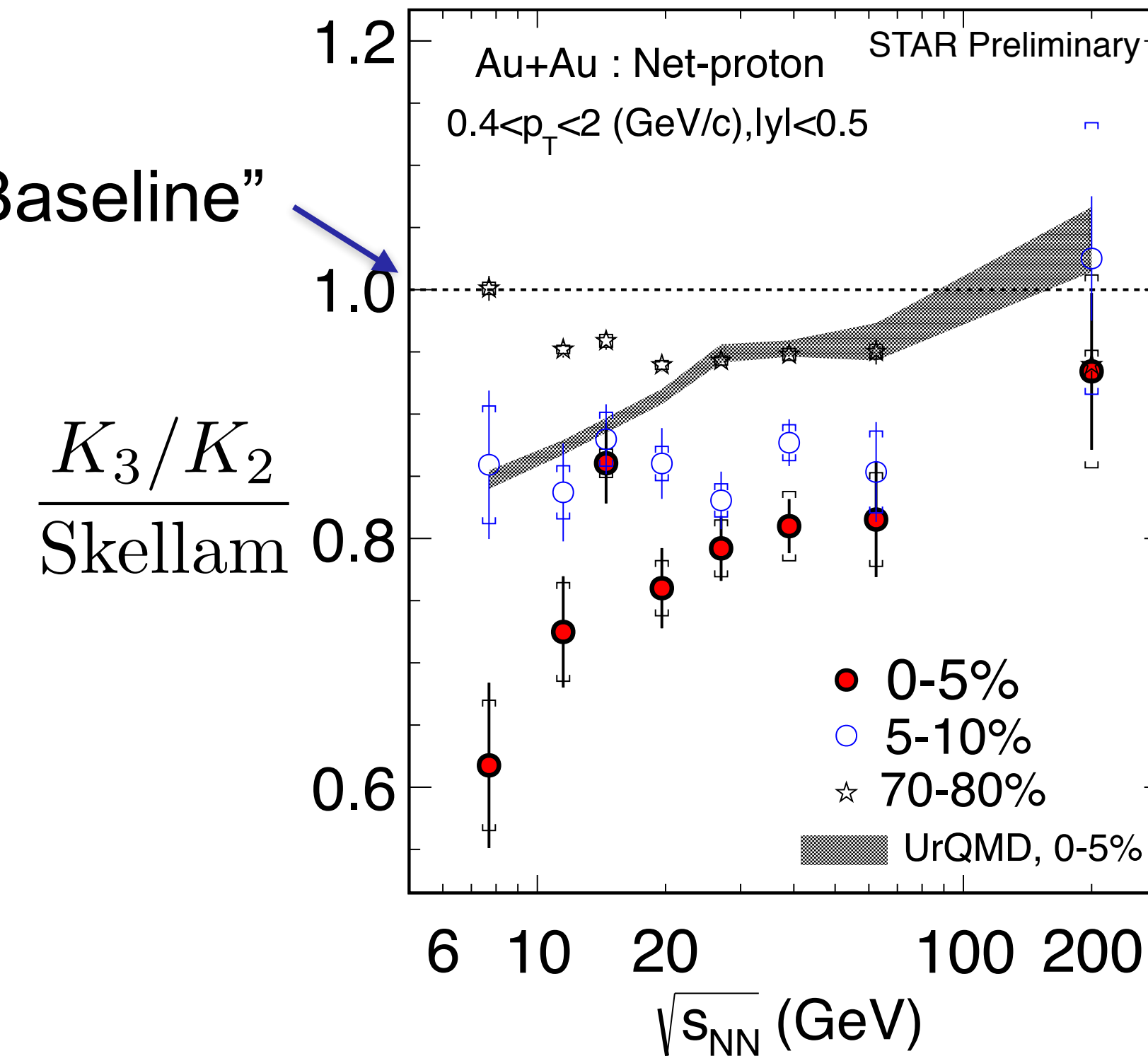
Should be visible in raw (unfolded) data

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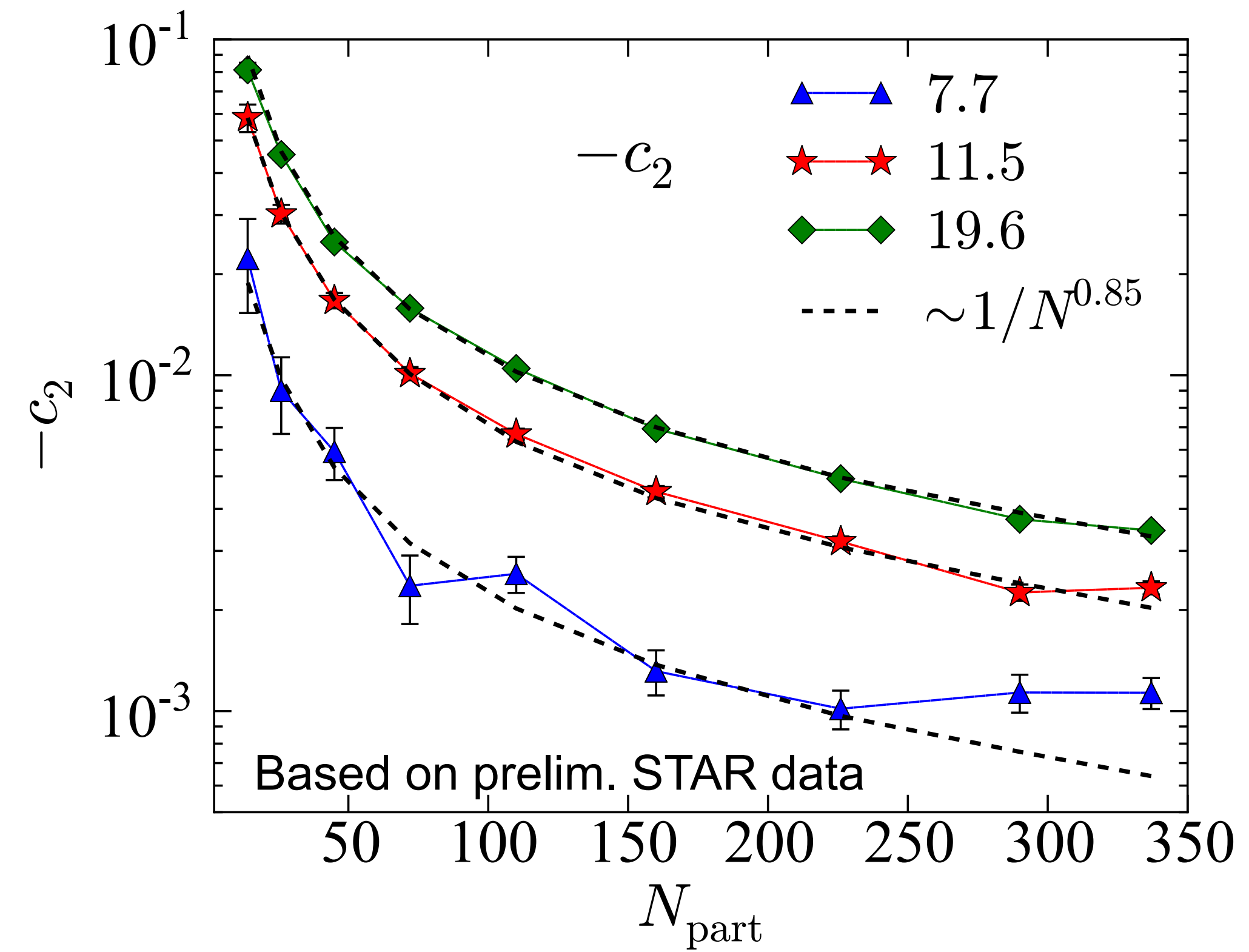


"Baseline"



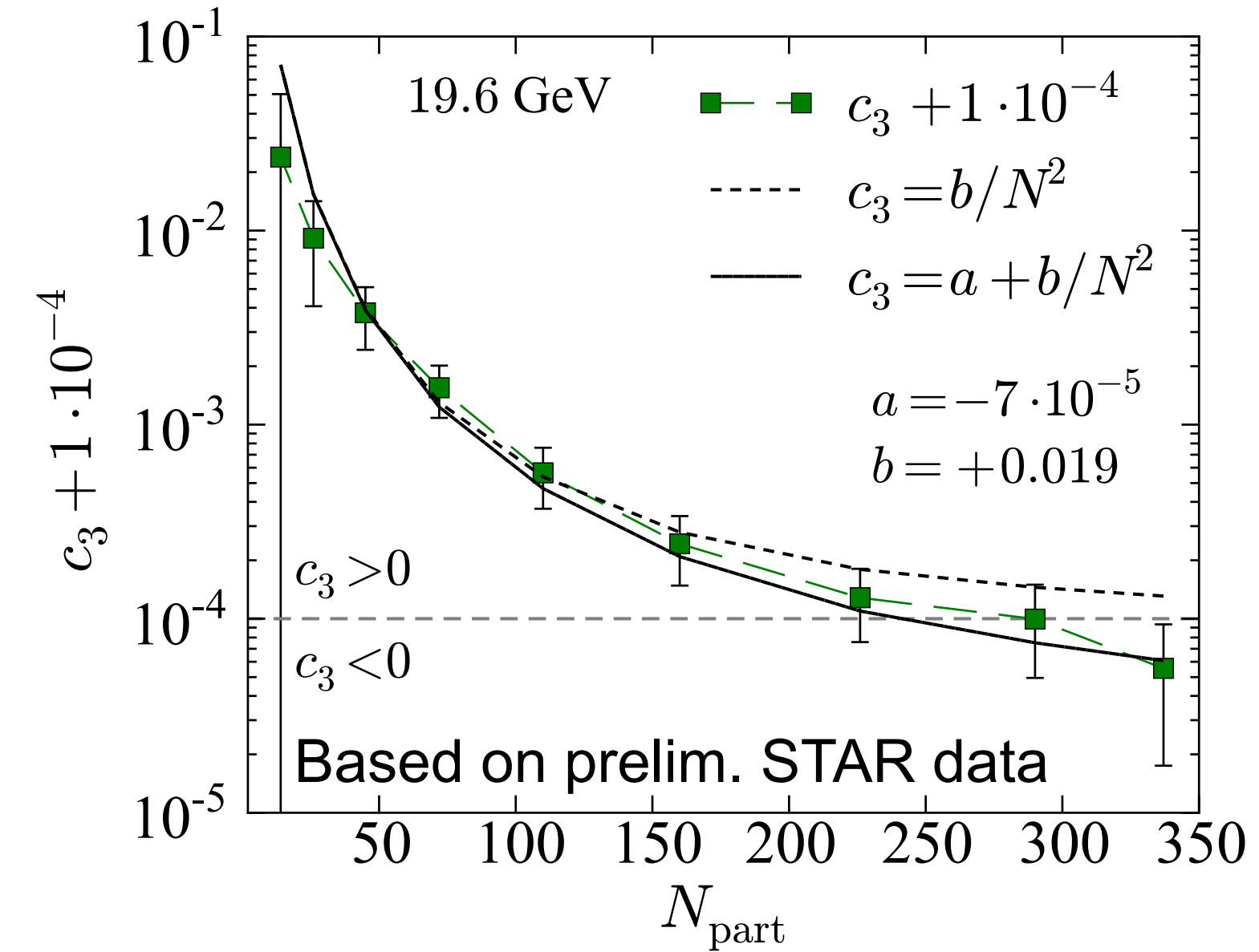
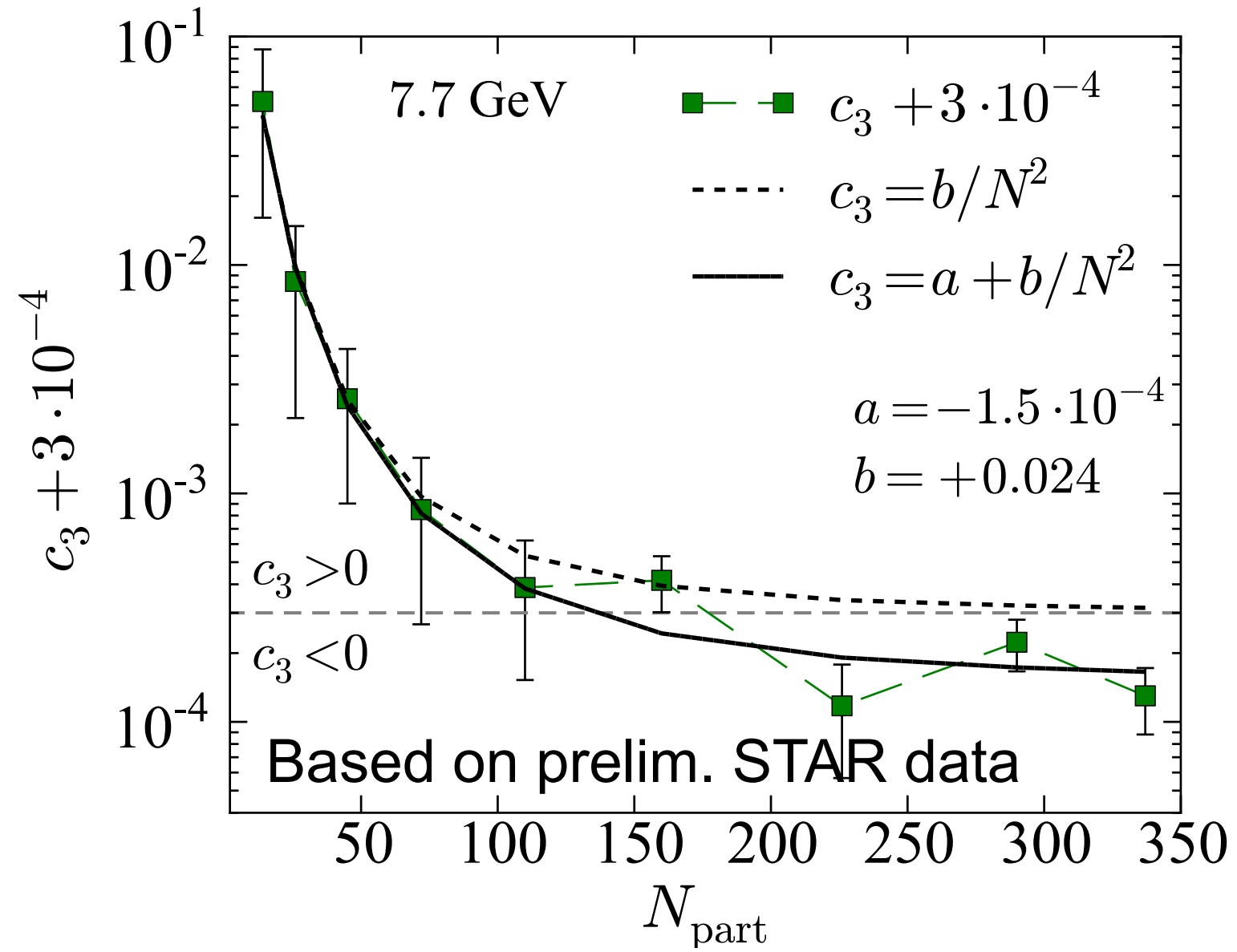
K_4/K_2 follows expectation, K_3/K_2 no so much.....
URQMD totally fails to get trend for K_4/K_2 !

Centrality dependence

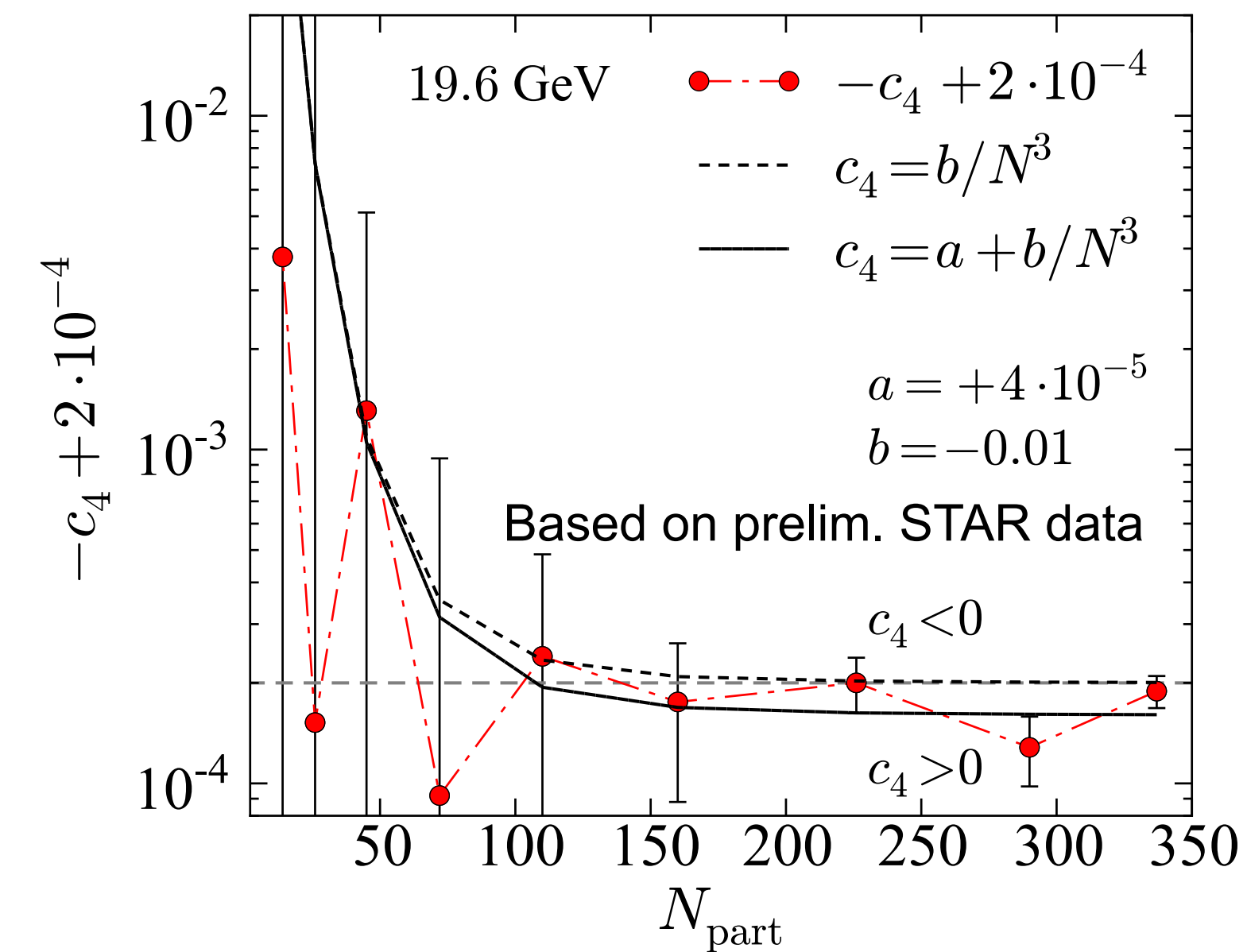
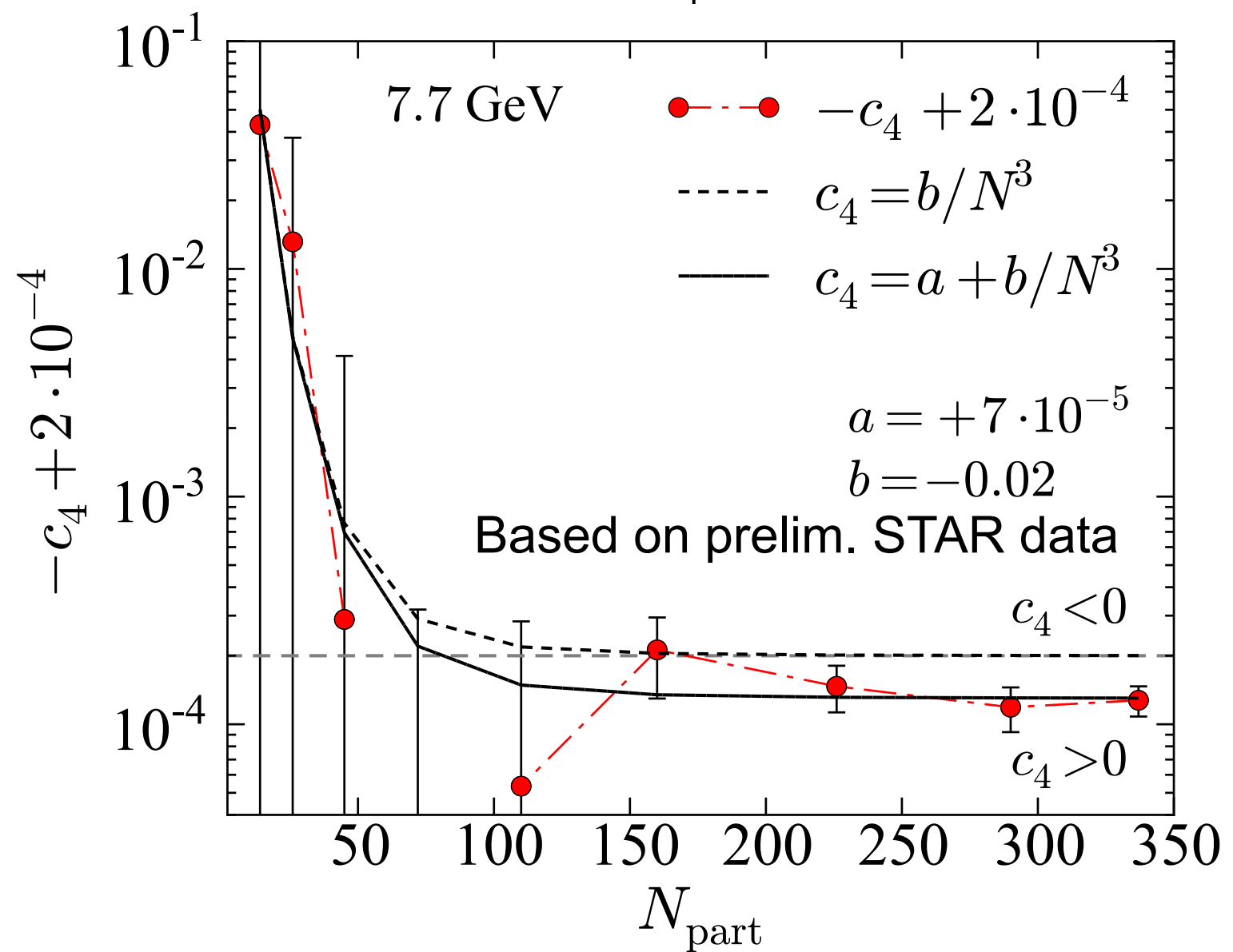


Centrality dependence

C_3



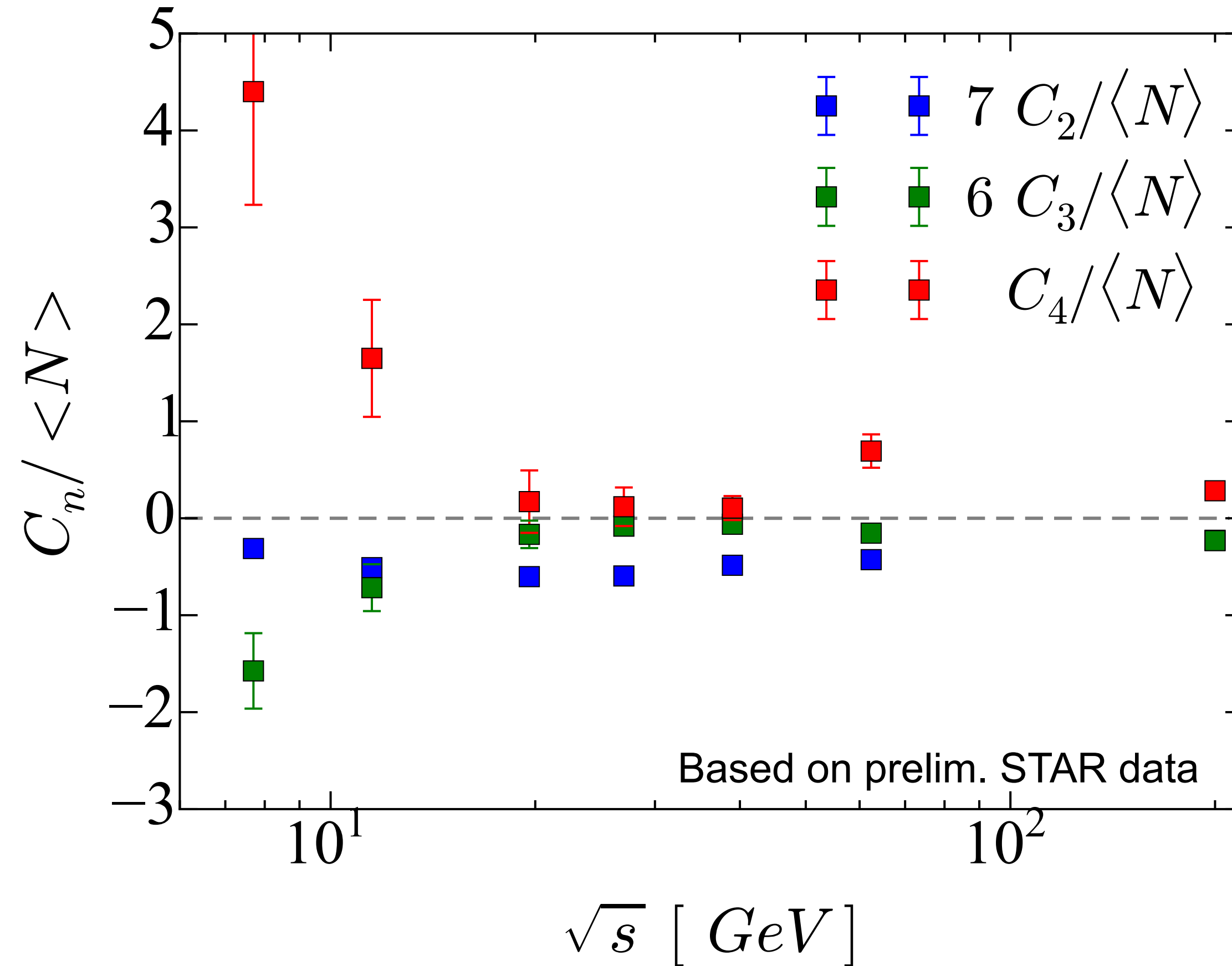
C_4



7.7 GeV

19.6 GeV

Energy dependence

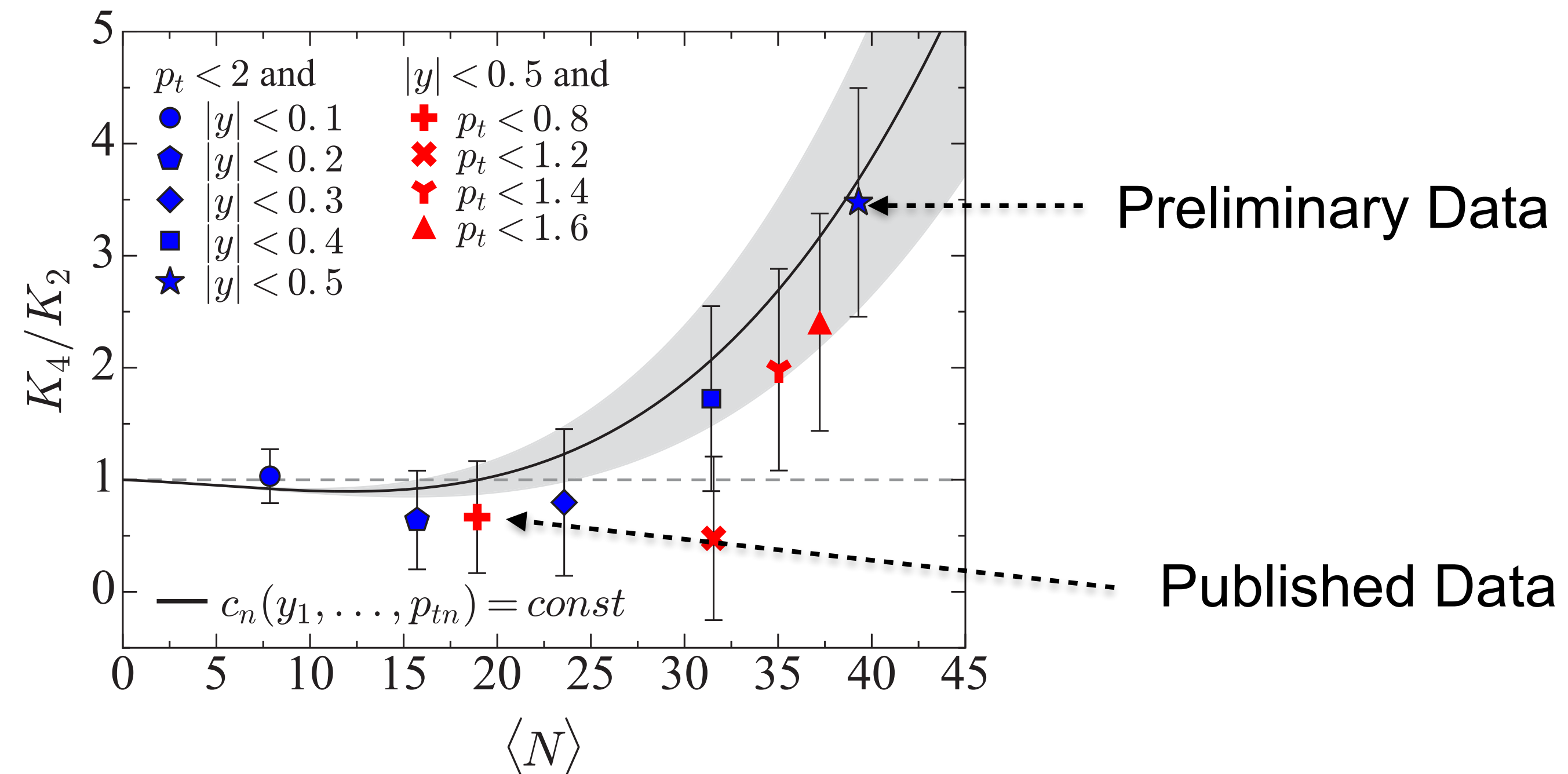


Note: anti-protons are non-negligible above 19.6 GeV
Data are protons only

Long range correlations

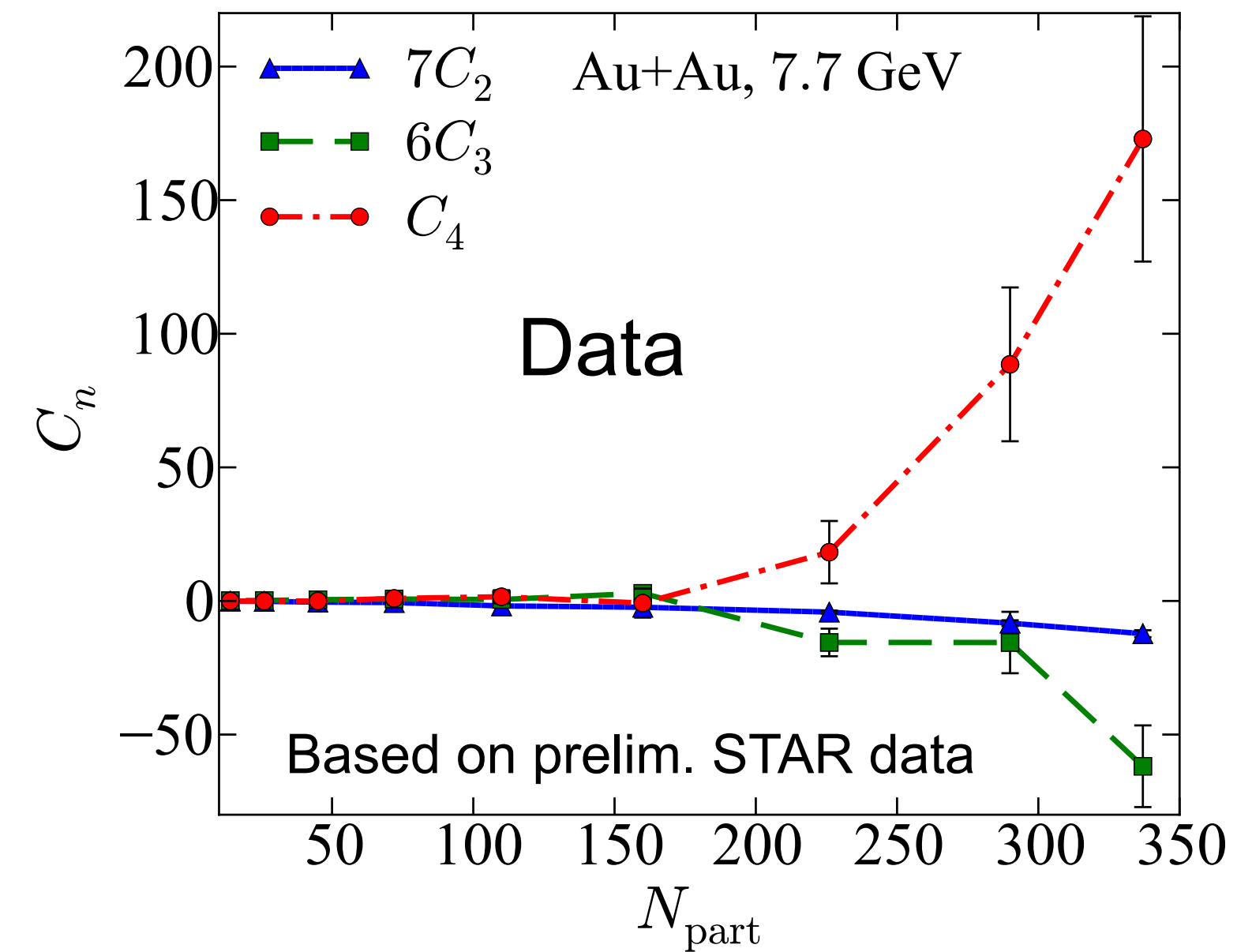
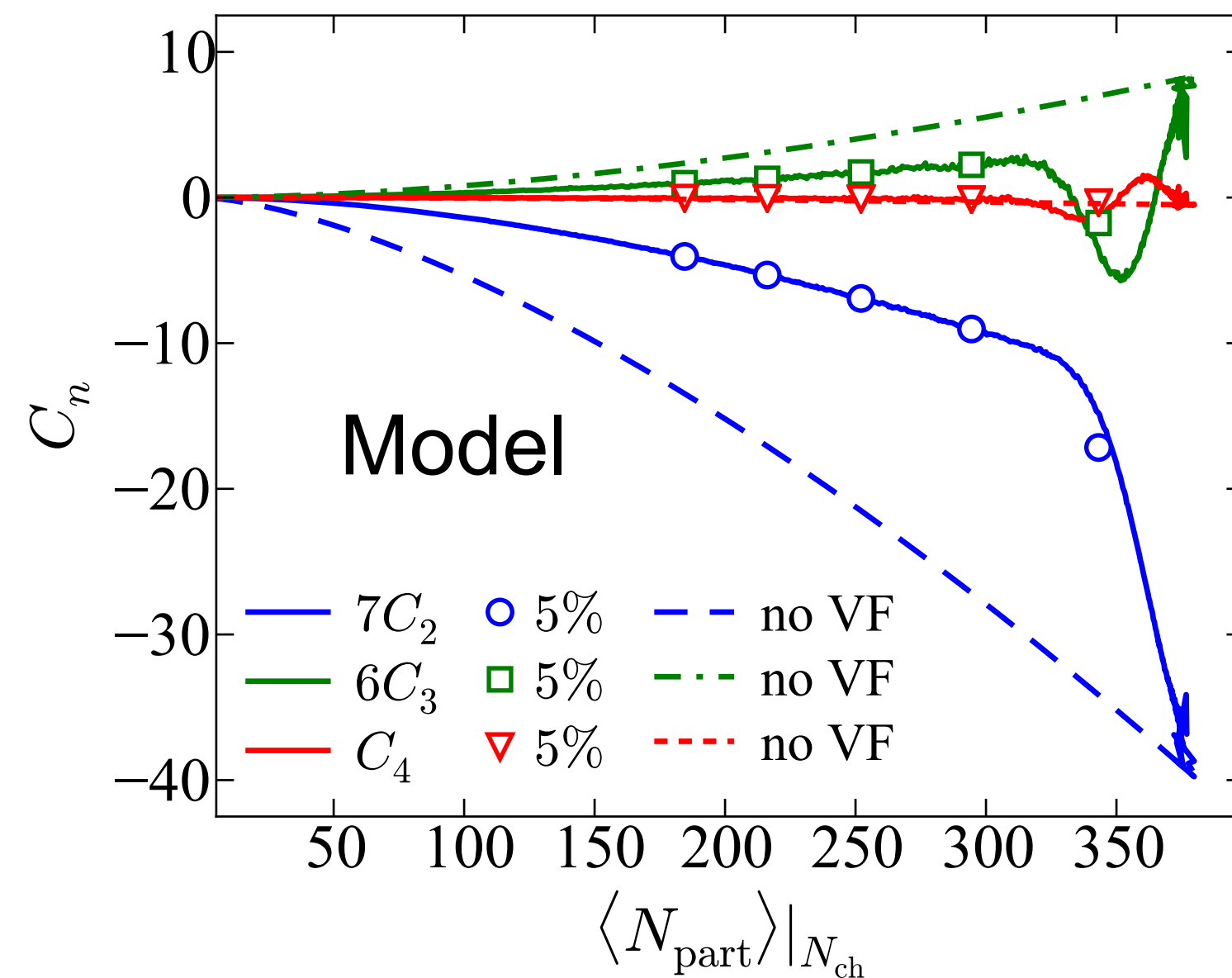
$$C_k = \langle N \rangle^k c_k$$

$$c_k = \text{const.} \Rightarrow K_n = K_n(\langle N \rangle)$$



Can we understand these correlations?

- Two particle correlations can be understood by simple Glauber model + Baryon number conservation



Four particle correlations are orders of magnitudes larger in the data
Also seen in URQMD calculations by He et al. PLB774 (2017) 623

Need to assume the ~40% of protons come from 8-nucleon cluster
in order to get magnitude right!

URQMD

