

Relativistic Dynamics of Fluctuations Away and Near the QCD Critical Point

Xin An (安鑫)

With G. Başar, M. Stephanov[†] and H.-U. Yee

1902.09517 and paper to appear



Quark Matter

Wuhan, China, Nov 6, 2019

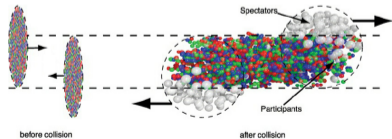


[†]Parallel talk in Session NTH I on Tuesday.

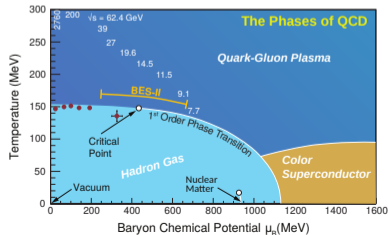
Introduction

- **Motivation:** Search for the QCD critical point.
- **Experimental approach:** Beam Energy Scan (BES) program.
- **Theory toolkit:** Generalized hydrodynamic formalism with

I) fluctuations in **arbitrary** relativistic flow.
Misha's talk (Tuesday)



II) **critical** slowing down (Hydro+).
Stephanov-Yin, 2017

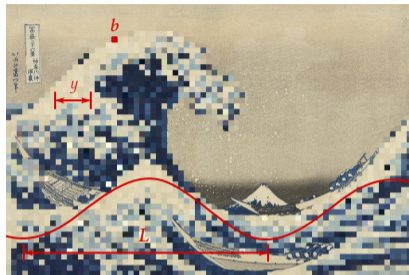


Historical review

authors	year	relativistic	arbitrary flow	non-conformal	charged	critical
Andreev	1970s	✗	✓	✓	✓	✗
Akamatsu-Mazeliauskas-Teaney	2017	✓	✗	✓	✗	✗
Martinez-Schäfer	2018	✓	✗	✗	✓	✗
XA-Başar-Stephanov-Yee	2019	✓	✓	✓	✓	✓

Separation of hydrodynamic scales

The Great Wave
K. Hokusai, 1830
Left: Original;
Right: Coarse-grained.



microscopic

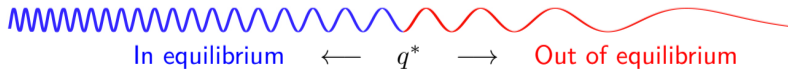
hydro cell

fluctuation

background

Length: $l_{\text{mic}} \ll b < y \ll L$

Wavenumber: $T \gg \Lambda > q \gg k$



Relaxation rate $\Gamma_{\text{relax}} = \gamma q^2 \xleftrightarrow{\text{competition}} \text{Evolution rate } \omega = c_s k \implies q^* \sim \sqrt{c_s k / \gamma}$.

From stochastic to deterministic approach

- Stochastic hydrodynamic equations:

$$\partial_t \check{\psi} = -\nabla \cdot (\text{flux}[\check{\psi}] + \text{noise}) \quad (\text{conservation})$$

where $\check{\psi}$ = energy, momentum and *baryon charge* densities. [click](#)

- Deterministic hydrodynamic equations:

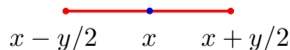
$$\begin{cases} \partial_t \psi = -\nabla \cdot \text{flux}[\psi; G], \\ \partial_t G = \text{relaxation}[G - G^{(\text{eq})}; \psi]. \end{cases}$$

where

$$\psi(x) = \langle \check{\psi}(x) \rangle, \quad \delta\psi(x) = \check{\psi} - \langle \check{\psi} \rangle; \quad (1\text{-pt})$$

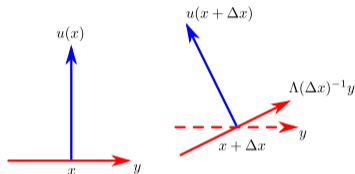
$$G(x, y) = \langle \delta\psi(x + y/2) \delta\psi(x - y/2) \rangle; \quad (2\text{-pt})$$

$$\langle \text{noise}(x + y/2) \text{noise}(x - y/2) \rangle \sim \delta^{(4)}(y). \quad (\text{FDT})$$

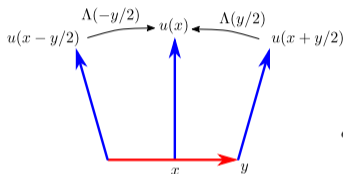


“Confluent” quantities

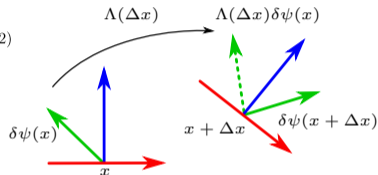
- “Confluent” quantities: correlators and derivatives adjusted by the fluid. *XA, et al, 2019*



equal time: $u \cdot y = 0$



correlator: $G \xrightarrow{\Lambda} \bar{G}$



derivative: $\partial \xrightarrow{\Lambda} \bar{\partial}$

- The “confluent” Wigner function

$$W(x, q) = \int_y e^{-iq \cdot y} \bar{G}(x, y)$$

satisfies a “confluent” evolution equation. [click](#)

“Confluent” evolution equations

- Upon averaging over fast modes on relevant time scales, **seven** slow modes survive. The sound mode decouples and matches the kinetic equation for phonons: [XA, et al, 2019](#)

$$\mathcal{L}[W_{++}] = \left[(u + v) \cdot \bar{\nabla} + \text{force} \cdot \frac{\partial}{\partial q} \right] W_{++} = -\Gamma_L q^2 (W_{++} - T/E),$$

where $v = c_s \hat{q}$, $E = c_s |q|$ and



“Finally, after about six months of work off and on, all the pieces suddenly fitted together, producing miraculous cancellation, and I was staring at the amazingly simple final result.” (C.N. Yang)

Renormalization and long-time tails

- flux = flux $[\psi; G]$ where

$$G(x) = \int^{\Lambda} d^3q W(x, q) \sim \underbrace{\Lambda^3 + \Lambda \partial(u, \mu/T)}_{\text{divergent}} + \underbrace{\tilde{G}}_{\text{remaining}} .$$

Divergent part: renormalize flux, i.e., $\psi \rightarrow \psi_R$ (EOS), $\gamma \rightarrow \gamma_R$ (transp.coeff.).

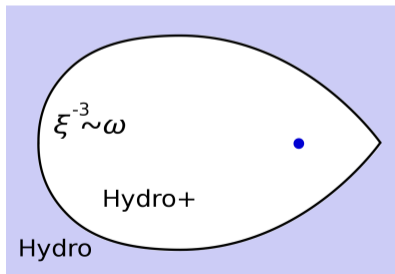
Remaining part: non-analytic terms (long-time tails)

$$\tilde{G} \sim q^{*3} \sim (c_s k / \gamma)^{3/2} \gg \mathcal{O}(k^2) .$$

- We obtain a closed set of **cutoff-independent** equations suitable for simulation : [XA, et al, 2019](#)

$$\begin{cases} \partial_t \psi = -\nabla \cdot \text{flux}[\psi; G], \\ \partial_t G = \text{relaxation}[G - G^{(\text{eq})}; \psi]. \end{cases} \longrightarrow \begin{cases} \partial_t \psi_R = -\nabla \cdot \text{flux}[\psi_R; \tilde{G}], \\ \bar{\partial}_t \tilde{G} = \text{relaxation}[\tilde{G}; \psi_R]. \end{cases}$$

- **Critical slowing down:** $\Gamma_{\text{relax}} \sim \xi^{-z} \rightarrow 0$ as the correlation length $\xi \rightarrow \infty$ near the critical point.
- Hydrodynamics breaks down when relaxation rate \lesssim evolution rate (i.e., $\Gamma_{\text{relax}} \lesssim \omega$).
- **Hydro+:** Extending Hydro by adding **one** critical slow mode G relaxing with $\Gamma_\lambda \sim \xi^{-3}$. Stephanov-Yin, 2017
- Our formalism merges with and extends Hydro+. XA, et al, 2019



Dynamics of fluctuations near the critical point

- Sound mode decouples, while remaining slow modes mix :

$$\begin{cases} \mathcal{L}[W_{\sigma\sigma}] = -2\Gamma_{\lambda} \left(W_{\sigma\sigma} - W_{\sigma\sigma}^{(\text{eq})} \right) + L[W_{\sigma i}], \\ \mathcal{L}[W_{\sigma i}] = -(\Gamma_{\eta} + \Gamma_{\lambda}) W_{\sigma i} + L[W_{\sigma\sigma}, W_{\sigma i}, W_{ij}], \\ \mathcal{L}[W_{ij}] = -2\Gamma_{\eta} \left(W_{ij} - W_{ij}^{(\text{eq})} \right) + L[W_{ij}, W_{\sigma i}]. \end{cases}$$

- The correlation length ξ increases as the system approaches the critical point, thus

$$\Gamma_{\lambda} \sim \xi^{-3} \ll \Gamma_{\eta} \sim \xi^{-2} \ll \xi^{-1}.$$

- Don't forget that Γ_{relax} and ω compete!

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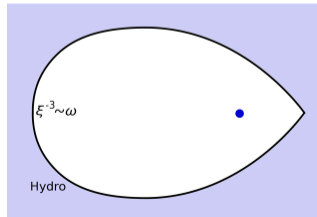
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- $\omega \ll \Gamma_\lambda \ll \Gamma_\eta$:



toward the smaller system (**increasing ω**)



approach the critical point (**increasing ξ**)

Dynamics of fluctuations near the critical point

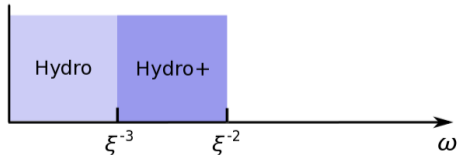
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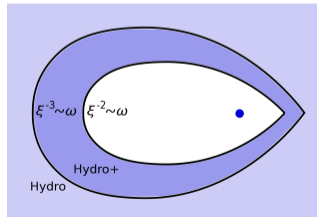
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Dynamics of fluctuations near the critical point

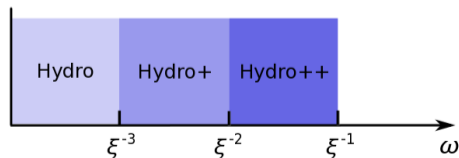
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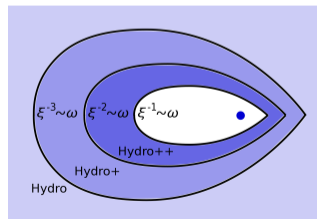
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- $\Gamma_\lambda \ll \Gamma_\eta \ll \omega \ll \xi^{-1}$:



toward the smaller system (increasing ω)



approach the critical point (increasing ξ)

Dynamics of fluctuations near the critical point

- Sound mode decouples, while remaining slow modes mix:

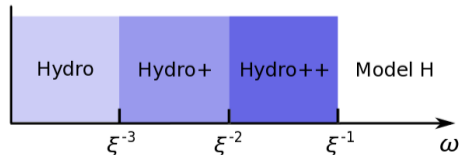


Equations need **nonlocal** modification! \longrightarrow Model H

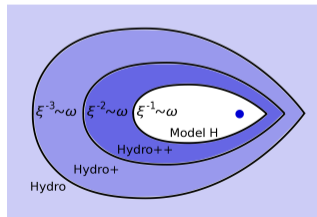
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- $\Gamma_\lambda \ll \Gamma_\eta \ll \xi^{-1} \lesssim \omega$:



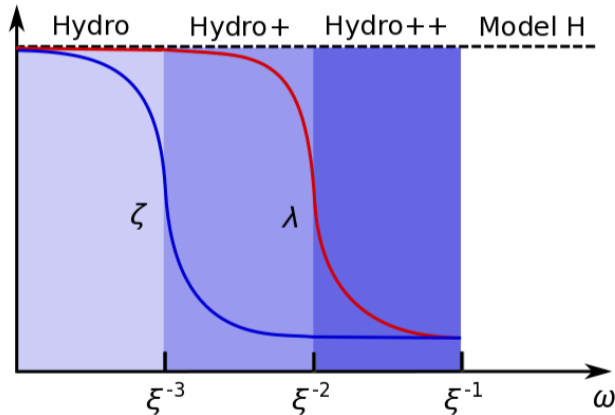
toward the smaller system (**increasing ω**)



approach the critical point (**increasing ξ**)

Dynamics of fluctuations near the critical point

- Schematic frequency dependence of the transport coefficients due to fluctuations:



Summary and outlook

- Our formalism is designed to
 - 1) apply to arbitrary relativistic flow;
 - 2) extend hydrodynamics to smaller system;
 - 3) extend Hydro+ closer to the QCD critical point.
- To Do List:
 - 1) higher-points functions;
 - 2) first order phase transition;
 - 3) freeze-out of heavy ion collisions;
 - 4) hydrodynamics with other DOFs.

A lot of work could to be done and interesting physics is just ahead!

Backup

Hydrodynamic quantities and equations

- The conserved quantities

$$\psi = (T^{00}, T^{i0}, J^0), \quad \text{flux} = (T^{i0}, T^{ij}, J^i),$$

where the stress tensor $T^{\mu\nu}$ and charge current J^μ are given by [back](#)

$$T^{\mu\nu}(\epsilon, n, u) = w(\epsilon, n)u^\mu u^\nu + p(\epsilon, n)g^{\mu\nu} + \Pi^{\mu\nu}, \quad w \equiv \epsilon + p(\epsilon, n),$$
$$J^\mu(\epsilon, n, u) = nu^\mu + \nu^\mu,$$

and

$$\Pi^{\mu\nu} = -2\eta \left(\theta^{\mu\nu} - \frac{1}{3} \Delta^{\mu\nu} \theta \right) - \zeta \Delta^{\mu\nu} \theta, \quad \nu^\mu = -\lambda \partial_\perp^\mu (\mu/T),$$
$$\theta^{\mu\nu} \equiv \frac{1}{2} (\partial^\mu u^\nu + \partial^\nu u^\mu), \quad \theta \equiv \partial \cdot u, \quad \Delta^{\mu\nu} \equiv g^{\mu\nu} + u^\mu u^\nu, \quad \Delta^{\mu\nu} \partial_\nu \equiv \partial_\perp^\mu.$$

- Conservation equations

$$\partial_\mu \check{T}^{\mu\nu} = 0, \quad \partial_\mu \check{J}^\mu = 0.$$

“Confluent” fluctuation evolution equation

- Fluctuation evolution equation: [back](#)

$$\begin{aligned} u \cdot \bar{\nabla} W(x, q) = & - \left[i\mathbb{L}^{(q)} + \mathbb{K}^{(a)}, W \right] - \left\{ \frac{1}{2} \bar{\mathbb{L}} + \mathbb{D}^{(q)} + \mathbb{K}^{(s)}, W \right\} + \theta W + 2\mathbb{Q}^{(q)} \\ & + (\partial_{\perp\lambda} u_{\mu}) q^{\mu} \frac{\partial W}{\partial q_{\lambda}} + \frac{1}{2} \left(a_{\lambda} + \frac{\partial_{\perp\lambda} c_s}{c_s} \right) \left\{ \mathbb{L}^{(q)}, \frac{\partial W}{\partial q_{\lambda}} \right\} \\ & + \frac{\partial}{\partial q_{\lambda}} \left(\left\{ \Omega_{\lambda}^{(s)}, W \right\} + [\Omega_{\lambda}^{(a)}, W] - \frac{1}{4} [\mathbb{H}_{\lambda}, [\mathbb{L}^{(q)}, W]] \right), \end{aligned}$$

where schematically

$$\begin{aligned} \mathbb{L}^{(q)} &\sim c_s q, & \bar{\mathbb{L}} &\sim c_s \bar{\nabla}_{\perp}, & \mathbb{D}^{(q)} &\sim \mathbb{Q}^{(q)} \sim \gamma q^2, \\ \mathbb{K} &\sim \partial(u, \alpha), & \mathbb{H} &\sim c_s \partial u, & \Omega &\sim c_s^2 \omega q. \end{aligned}$$