Measurement of the $\gamma\gamma \rightarrow e^+e^-$ Process and its Angular Correlations in UPC and Peripheral Au+Au Collisions with the STAR Detector

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武法 2019

Ultra-Peripheral Collisions



Ultra-relativistic charged nuclei produce highly Lorentz contracted electromagnetic field

Weizäcker-Williams Equivalent Photon Approximation (EPA): \rightarrow In a specific phase space, EM fields can be quantized as a flux of **real photons**

Weizsäcker, C. F. v. Zeitschrift für Physik 88 (1934): 612

 $Z\alpha \approx 1 \rightarrow$ High photon density Magnetic field strength $\vec{B} \approx 10^{14} - 10^{16}$ T Skokov, V., et. al. Int. J. Mod. Phys. A 24 (2009): 5925–32

Test QED under extreme conditions

 $\gamma\gamma \rightarrow e^+e^-$ Process

1934 Breit & Wheeler : "Collision of two Light Quanta"

G. Breit and J. A. Wheeler. Physical Review 46 (1934): 1087



1. Identifying $\gamma \gamma \rightarrow e^+e^-$ process in ultra-peripheral heavy-ion collisions

2. Ultra-peripheral vs. peripheral

3. First Earth-based observation of vacuum birefringence

4. Applications

Signatures of the $\gamma\gamma \rightarrow e^+e^-$ Process



1. Exclusive production of e^+e^- pair

- 2. Smooth invariant mass spectra (No vector mesons)
- 3. Individual e^+/e^- preferentially aligned in beam direction
- 4. Production peaked at very low P_{\perp} (pair transverse momentum)

Total $\gamma\gamma \rightarrow e^+e^-$ cross-section in STAR Acceptance



STARLight: S. R. Klein, et. al. *Comput. Phys. Commun.* 212 (2017) 258 gEPA & QED : W. Zha, J.D.B., Z. Tang, Z. Xu arXiv:1812.02820 [nucl-th]

Pure QED 2 \rightarrow 2 scattering : $d\sigma/dM \propto E^{-4} \approx M^{-4}$

No vector meson production \rightarrow Forbidden for real photons with helicity ± 1 (i.e. 0 is forbidden)

$$\sigma(\gamma\gamma \rightarrow e^+e^-)$$
 in STAR Acceptance:

Data : 0.261 ± 0.004 (stat.) ± 0.013 (sys.) ± 0.034 (scale) mbSTARLightgEPA0.22 mb0.26 mb0.29 mb

Measurement of total cross section agrees with theory calculations at $\pm 1\sigma$ level

$$d\sigma(\gamma\gamma \to e^+e^-)/d\cos\theta'$$

 $\gamma\gamma \rightarrow e^+e^-$: Individual e^+/e^- preferentially aligned along beam axis [1]:

$$G(\theta) = 2 + 4\left(1 - \frac{4m^2}{W^2}\right) \frac{\left(1 - \frac{4m^2}{W^2}\right)\sin^2\theta\cos^2\theta + \frac{4m^2}{W^2}}{\left(1 - \left(1 - \frac{4m^2}{W^2}\right)\cos^2\theta\right)^2}$$

 Highly virtual photon interactions should have an <u>isotropic distribution</u>

 \circ Measure θ' , the angle between the e^+ and the beam axis in the pair rest frame.

Data are fully consistent with $G(\theta)$ distribution expected for $\gamma\gamma \rightarrow e^+e^-$



[1] S. Brodsky, T. Kinoshita and H. Terazawa, Phys. Rev. **D4**, 1532 (1971) STARLight: S. R. Klein, et. al. *Comput. Phys. Commun.* 212 (2017) 258

 $d\sigma(\gamma\gamma \to e^+e^-)/dP_+$



STARLight: S. R. Klein, et. al. *Comput. Phys. Commun.* 212 (2017) 258 QED : W. Zha, J.D.B., Z. Tang, Z. Xu arXiv:1812.02820 [nucl-th] \circ Cross-section peaks at low P_{\perp} , as expected for <u>real photon collisions</u>

 \odot Data are well described by leading order QED calculation ($\gamma\gamma
ightarrow e^+e^-$)

 \circ STARLight predicts significantly lower $\langle P_{\perp} \rangle$ than seen in data

 \circ STARLight calculations do not have centrality dependent P_{\perp} distribution

 Experimentally investigate impact parameter dependence :

 \rightarrow Compare UPC vs. peripheral collisions

$\gamma\gamma \rightarrow e^+e^-$: UPC vs. Peripheral

[1] STAR, Phys. Rev. Lett. 121 (2018) 132301



Characterize difference in spectra via $\sqrt{\langle P_{\perp}^2 \rangle}$					
$\sqrt{\langle P_{\perp}^2 angle}$ (MeV/c)	UPC Au+Au	60-80% Au+Au			
Measured	38.1 ± 0.9	50.9 <u>+</u> 2.5			
QED	37.6	48.5			
b range (fm)	≈ 20	$\approx 11.5 - 13.5$			
 Leading order QED calculation of 					
$\gamma\gamma \rightarrow e^+e^-$ describes both spectra ($\pm 1\sigma$)					



[1] STAR, Phys. Rev. Lett. 121 (2018) 132301

[2] S. R. Klein, et. al, Phys. Rev. Lett. 122, (2019), 132301
[3] ATLAS Phys. Rev. Lett. 121 (2018), 212301

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• Leading order QED calculation of $\gamma\gamma \rightarrow e^+e^-$ describes both spectra (±1 σ)

• Best fit for spectra in 60-80% collisions found for QED shape plus 14 ± 4 (stat.) ± 4 (syst.) MeV/c broadening

 Proposed as a probe of trapped magnetic field or Coulomb scattering in QGP [1-3]

STAR observes 4.8 σ difference between UPC and 60-80% Au+Au collisions

Optical Birefringence

Birefringent material: Different index of refraction for light polarized parallel (n_{\parallel}) vs. perpendicular (n_{\perp}) to material's ordinary axis

ightarrow splitting of wave function when $\Delta n = n_{\parallel} - n_{\perp}
eq 0$



Birefringence of the QED Vacuum

Vacuum birefringence : Predicted in 1936 by Heisenberg & Euler. Index of refraction for γ interaction with \vec{B} field <u>depends on relative</u> <u>polarization angle</u> i.e. $\Delta \sigma = \sigma_{\parallel} - \sigma_{\perp} \neq 0$

Lorentz contraction of EM fields \rightarrow

Quasi-real photons should be linearly polarized $(\vec{E} \perp \vec{B} \perp \vec{k})$



Can we observe vacuum birefringence in ultra-peripheral collisions?

Feynman Diagram for Vacuum Birefringence



S. Bragin, et. al., *Phys. Rev. Lett.* 119 (2017), 250403 R. P. Mignani, *et al., Mon. Not. Roy. Astron. Soc.* 465 (2017), 492

Birefringence of the QED Vacuum [1] C. Li, J. Zhou, Y.-j. Zhou, Phys. Lett. B 795, 576 (2019) QED calculation: arxiv : 1911.00237

Recently realized, $\Delta \sigma = \sigma_{\parallel} - \sigma_{\perp} \neq 0$ leads to $\cos(n\Delta \phi)$ modulations in polarized $\gamma \gamma \rightarrow e^+ e^-$ [1]

$\Delta \phi = \Delta \phi [(e^+ + e^-), (e^+ - e^-)]$
$pprox \Delta \phi[(e^+ + e^-), e^+]$

Ultra-Peripheral

Quantity	Measured	QED	χ^2/ndf	
$-A_{4\Delta\phi}(\%)$	16.8 ± 2.5	22	18.8 / 16	
	Peripheral (60–80%)			
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\rightarrow First Earth-based observation (6.7 σ level) of vacuum birefringence

Application : Mapping the Magnetic Field

Total and differential cross-sections (e.g. $d\sigma/dP_{\perp}$) for $\gamma\gamma \rightarrow e^+e^-$ are related to field strength and configuration

photon density is related to energy flux of the electromagnetic fields [1]

$$n \propto \vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$$

 \rightarrow Report \vec{B} (single ion) that matches measured cross-section





Diffractive Photoproduction of the ho^0 Meson

Employ the same observable for $\rho^0 \rightarrow \pi^+\pi^-$ (and direct $\pi^+\pi^-$)

- $\circ \ \ \ {\rm Use \ the \ polarized \ } \gamma \ {\rm as \ a \ probe} \\ {\rm of \ the \ nucleus \ }$
- Calculate coefficients $\langle \cos(n\Delta\phi) \rangle$

$$\Delta \phi = \Delta \phi [(\pi^+ + \pi^-), (\pi^+ - \pi^-)]$$



11/05/19

Photoproduction of the ho^0 Meson



- Amplitude of the $\cos(2\Delta\phi)$ modulation appears to be related to diffraction peaks
- o <u>Theory input needed</u> for quantitative description of data

Photoproduction of the ho^0 Meson

Observation of significant $\cos(4\Delta\phi)$ modulation with respect to background

Predicted to be sensitive to the gluon Generalized Transverse Momentum Dependent (GTMD) Distribution [1]

"...offers direct access to the second derivative of the saturation scale with respect to b_{\perp}^{2} " [1]

Tensor Pomeron model may also lead to $\cos 4\Delta\phi$ modulations



• <u>Theory input needed</u> for quantitative description of data

Summary 1

1. Measurements of exclusive $\gamma \gamma \rightarrow e^+ e^-$ process

2. Experimental demonstration that the $\sqrt{\langle P_{\perp}^2 \rangle}$ spectra from $\gamma \gamma \rightarrow l^+ l^-$ depends on impact parameter (4.8 σ observation)

3. First Earth-based observation of Vacuum Birefringence :

Observed (6.7 σ) via angular modulations in linear polarized $\gamma\gamma \rightarrow e^+e^-$ process

Summary 2 (Applications)

- 1. Photons originate from colliding electromagnetic fields \circ Measurement of the $\gamma\gamma \rightarrow e^+e^-$ process : Lorentz invariant process for mapping strength of initial \vec{B}
- 2. New observable measured in photo-produced ρ⁰ Meson
 Significant cos 2Δφ and cos 4Δφ modulations observed
 May be sensitive to gluon Generalized Transverse Momentum Dependent (GTMD) Distribution or spin of Pomeron (in Pomeron model)
 Theory input needed for quantitative description
 - <u>Theory input needed</u> for quantitative description

Thank you for your attention

Backup

The $\gamma\gamma \rightarrow e^+e^-$ Process

1934 Breit & Wheeler : "Collision of two Light Quanta"

G. Breit and J. A. Wheeler. Physical Review 46 (1934): 1087



• Trigger on neutrons in ZDC \rightarrow Select events with mutual Coulomb excitation followed by dissociation

$d\sigma/dM$ for events with 1n1n events



1n1n: events with 1 neutron in each ZDC

Application : Mapping the Magnetic Field

The colliding photons in the $\gamma\gamma \rightarrow e^+e^$ process <u>originate from the Lorent-</u> <u>contracted Electromagnetic fields</u>

photon density is related to energy flux of $\overset{(\widehat{\underline{E}})}{>}$ the electromagnetic fields

$$n \propto \vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$$

For highly Lorentz contracted fields $|E| \approx |B|$ with $\vec{E} \perp \vec{B}$ and $\vec{S} \propto |E|^2 \approx |B|^2$



Equivalent Photon Approximation, photon density (single ion):

$$n(\omega;b) = \frac{1}{\pi\omega} |E_{\perp}(b,\omega)|^2 = \frac{1}{\pi\omega} |B_{\perp}(b,\omega)|^2 = \frac{4Z^2\alpha}{\omega} \left| \int \frac{d^2k_{\perp}}{(2\pi)^2} k_{\perp} \frac{F(k_{\perp}^2 + \omega^2/\gamma^2)}{k_{\perp}^2 + \omega^2/\gamma^2} e^{-i b \cdot k_{\perp}} \right|^2$$

M. Vidović, et al., Phys. Rev. C 47, 2308 (1993).
 C. F. v. Weizsa cker, Z. Phys. 88, 612 (1934).

Example : Light-by-Light Scattering



ATLAS Observed Light-by-Light Scattering in UPCs:

Purely quantum mechanical process (α⁴_{em})
 Light-by-Light scattering involves real photons by definition

ATLAS, Nature Physics 13 (2017), 852

 $d\sigma(\gamma\gamma \rightarrow e^+e^-)/dP_+$



QED Calculation: W. Zha, J.D.B., Z. Tang, Z. Xu arXiv:1812.02820 [nucl-th]

Photoproduction of the ho^0 Meson

Use similar observable for $\rho^0 \rightarrow \pi^+\pi^-$

- Calculate coefficients $\langle \cos(n\Delta\phi) \rangle$
- Sensitive to gluon distribution and gradients
- n = 1: Closure test, no modulation expected
- Background estimates:
- 1. STARLight (does not include polarization effects)
- 2. Data-driven (like-sign pairs)



J. Zhou Phys. Rev. D **94** (2016), 114017