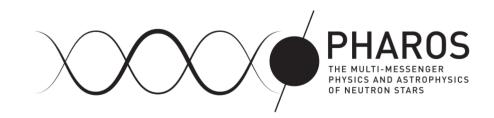
Signatures of Quark-Hadron Phase Transitions in General-Relativistic Neutron-Star Mergers

Veronica Dexheimer

in collaboration with Elias Most, Jens Papenfort, Matthias Hanauske, Luciano Rezzolla and Horst Stöcker

> Phys. Rev. Lett. (2019) J. Phys. G (2019) ArXiv 1910.13893 (2019)







CMF (Chiral Mean Field) model

- non-linear realization of the linear sigma model
- includes baryons (+ leptons) and quarks
- fitted to reproduce nuclear, astrophysical, lattice QCD, and heavy ion, constraints
- baryon and quark effective masses

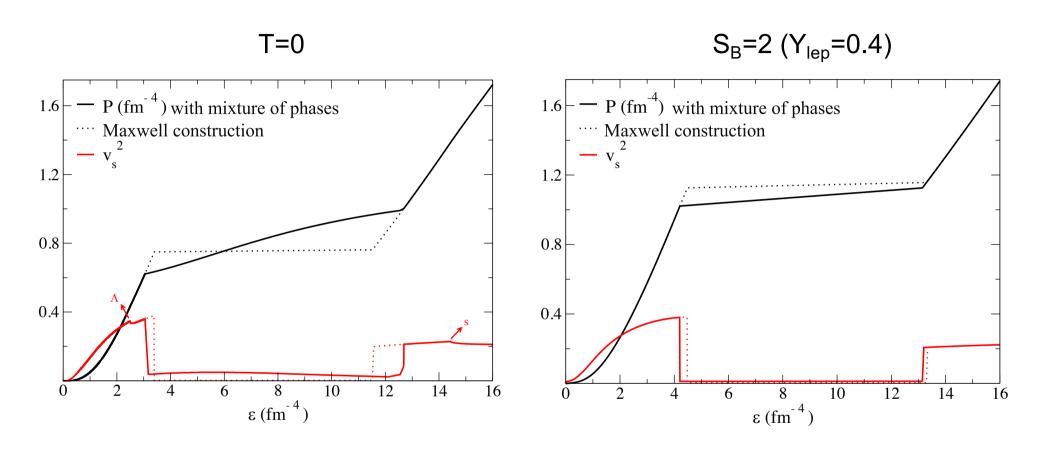
$$M_B^* = g_{B\sigma}\sigma + g_{B\delta}\tau_3\delta + g_{B\zeta}\zeta + M_{0_B} + g_{B\Phi}\Phi^2$$

$$M_q^* = g_{q\sigma}\sigma + g_{q\delta}\tau_3\delta + g_{q\zeta}\zeta + M_{0_q} + g_{q\Phi}(1 - \Phi)$$

- 1st order phase transitions or crossovers
- potential for Φ
 deconfinement
 order parameter

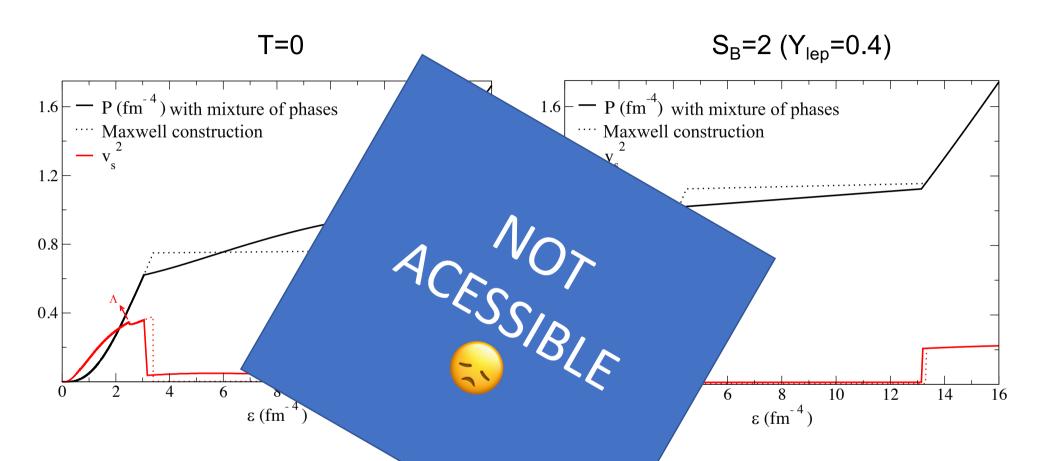
$$U = (a_o T^4 + a_1 \mu_B^4 + a_2 T^2 \mu_B^2) \Phi^2 + a_3 T_o^4 \ln(1 - 6\Phi^2 + 8\Phi^3 - 3\Phi^4)$$

Speed of sound for neutron-star matter



- $v_s^2 \sim 1/3$ at very large energies
- tested against PQCD results at zero/finite temperature

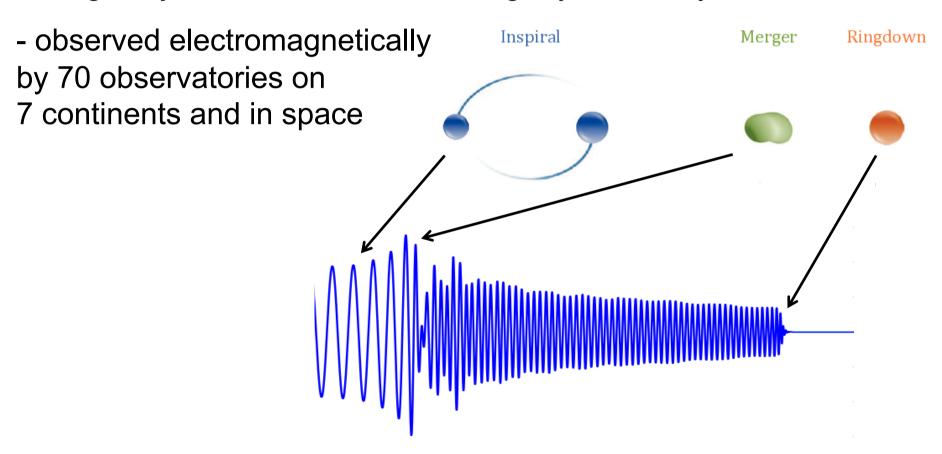
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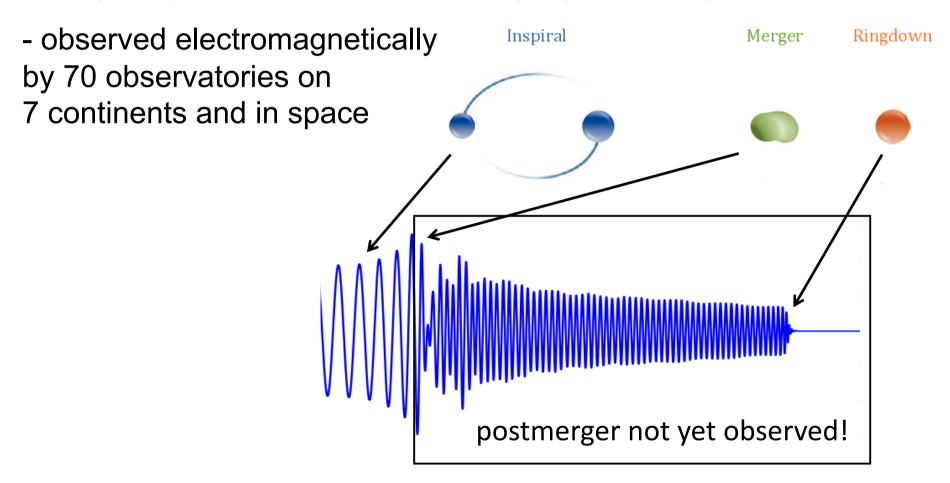
Neutron-star merger 170817

- gravitational waves observed by LIGO/VIRGO in 17 August 2017 from galaxy NGC 4993 140 million light-years away



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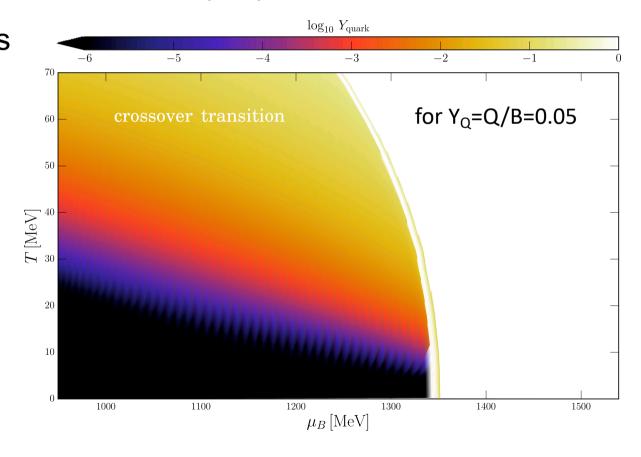


 dephasing due to deconfinement to quark matter (Luciano Rezzolla's plenary talk on Saturday)



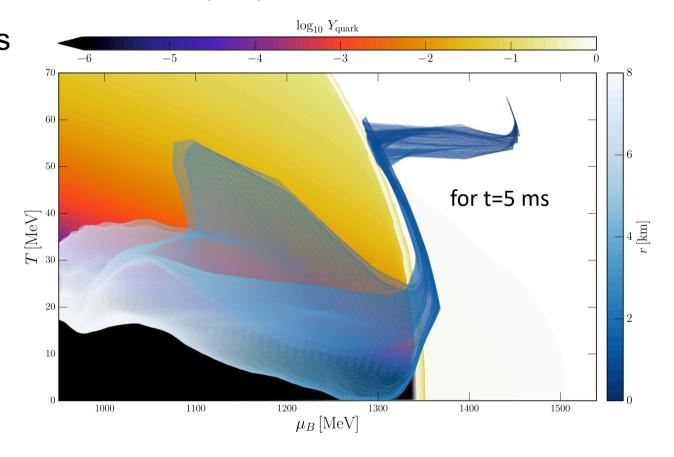
NS-merger simulation with deconfinement

- 3D (T,n_B,Y_Q) CMF EoS with 1st order phase transition
- solve coupled Einstein-hydrodynamics system using Frankfurt/IllinoisGRMHD code (FIL)
- interesting results
 for binaries with
 final masses
 of 2.8 and
 2.9 M_{Sun} after
 deconfinement
 but before
 collapse to
 black hole



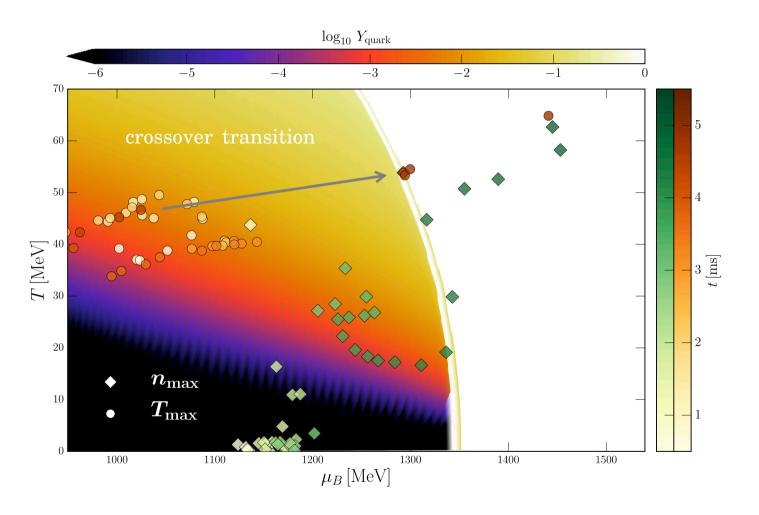
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NS-merger in the QCD phase diagram

- tracking maximum temperature ● and density ◆ in merger

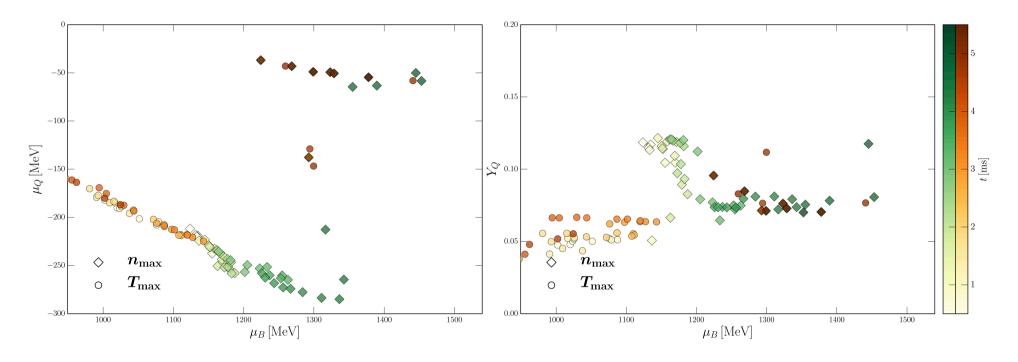


More phase diagrams

tracking maximum temperature

 and density

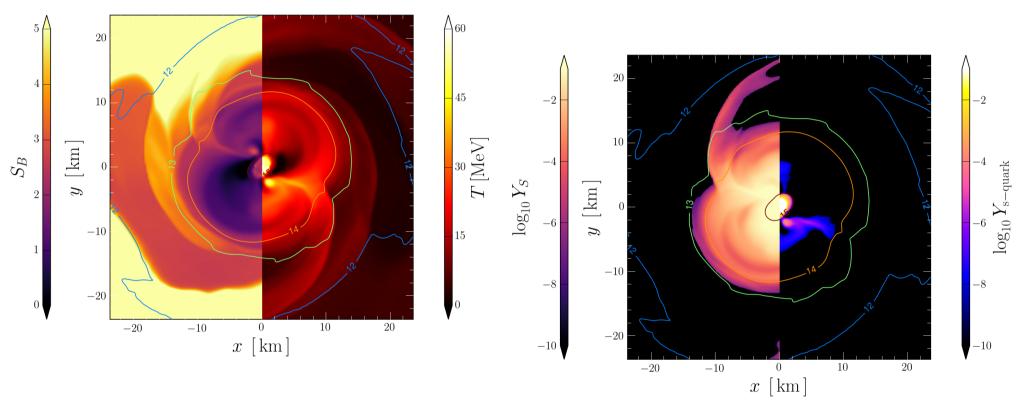
 in merger



- increase in abs. value of charged chemical potential until phase transition, when it drops
- decrease in charge fraction of core after when quarks appear

Inside 1 hypermassive neutron star

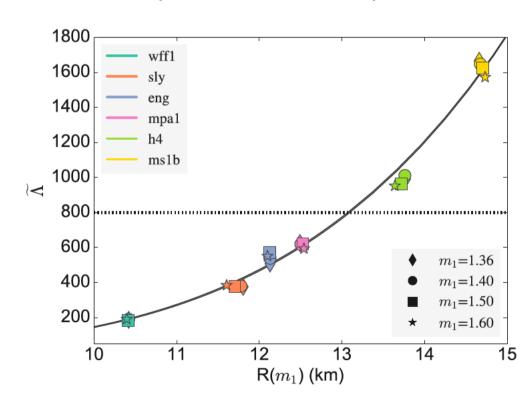
- total mass 2.9 M_{Sun} 5 ms after merger



- increase of temperature, entropy per baryon, and s-quark fraction at phase transition
- total strangeness (hyperons → s-quarks) remains the same

Tidal deformability

- normalized stellar quadrupole deformation by companion
- calculated from finite-size
 effects in end of inspiral:
 76 → 1045 with 90% confidence (De et. al 2018)
- related to NS radius of M=1.4 M_{sun} (Raithel et. al 2018)
- universal relation?



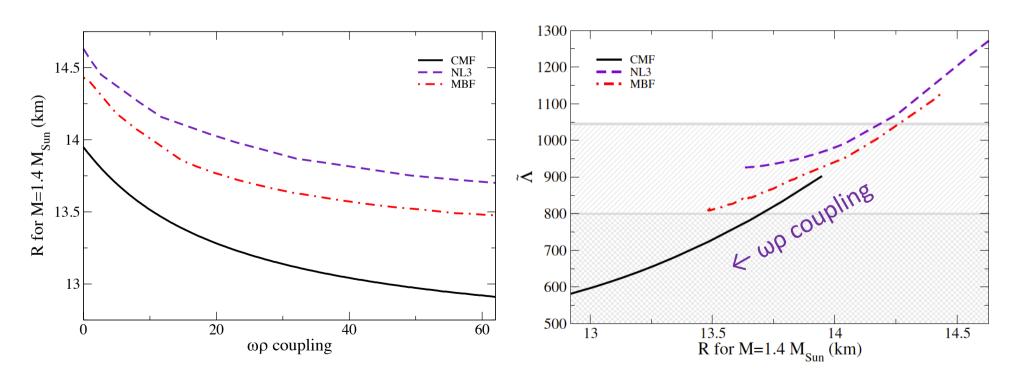
Inspiral

Ringdown

Merger

Exploring hadronic isovector coupling

- using 3 relativistic EoS's that fulfill standard nuclear and astrophysical constraints
- new repulsive vector-isovector channel $L_{\omega\rho} = g_{\omega\rho}g_{\omega}^2g_{\rho}^2\omega_{\mu}\omega^{\mu}\rho_{\mu}\rho^{\mu}$ suggested by Horowitz and Piekarewicz (2002)



- non-trivial relation between $\widetilde{\Lambda}$ and $\mathsf{R}_{\mathsf{1.4M_{Sun}}}$

Outlook

- GW170817 provided us a new way to study dense and hot matter in an environment not quite the same as heavy-ion or supernova ones
- more quantum relativistic models with temperature/exotic degrees of freedom needed to study
 - relation between tidal deformability and nuclear physics
 - realistic neutron-star merger simulations
- more LIGO/VIRGO merger data coming soon ...so, maybe, there will be a clear signature for deconfinement from astrophysics!

