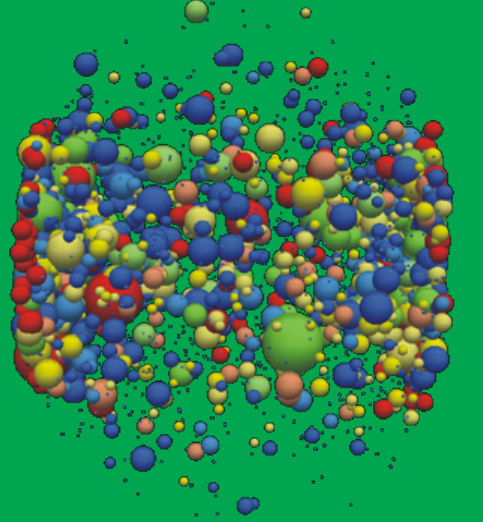


Fluctuations of anisotropic flow in transport

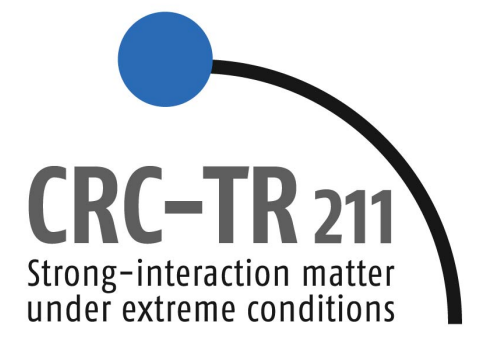
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Introduction

- We investigate anisotropic flow for a 2D system of massless particles, within the approach of C. Gombeaud and J.-Y. Ollitrault [1].
- For controlled initial geometries, we study the change in v_2 , v_3 , v_4 as the Knudsen number Kn is varied.
- Using a MC Glauber model as input for the initial condition, we show how the resulting fluctuations in v_2 and v_3 depend on the mean number of rescatterings in the system.

Dependence of anisotropic flow on Kn for controlled initial conditions

- The spatial part of our initial-state distribution function:

$$f(r, \theta) = \frac{1}{2\pi R^2} e^{-\frac{r^2}{2R^2}} \left[1 - 4\varepsilon_2 e^{-\frac{r^2}{2R^2}} \left(\frac{r}{R}\right)^2 \cos(2\theta) - \sqrt{2\pi}\varepsilon_3 e^{-\frac{r^2}{2R^2}} \left(\frac{r}{R}\right)^3 \cos(3\theta) - \frac{4}{3}\varepsilon_4 e^{-\frac{r^2}{2R^2}} \left(\frac{r}{R}\right)^4 \cos(4\theta) \right]$$

- We performed calculations at different Kn with only one $\varepsilon_n = 0.15$ and the other $\varepsilon_{p \neq n} = 0$. To reduce statistical fluctuations of v_n about 0 in the initial state, we averaged over 500 runs with $2.5 \cdot 10^5$ particles for each (initial) geometry.

- Every anisotropic flow harmonic behaves as $v_n = \frac{v_n^{\text{hydro}}}{1 + \frac{\text{Kn}}{\text{Kn}_0}}$, as anticipated in [2] and observed in [1] for v_2 .

- At fixed Kn , higher harmonics are more suppressed.

- As n grows, v_n sets on at increasingly larger number of rescatterings (smaller Kn).

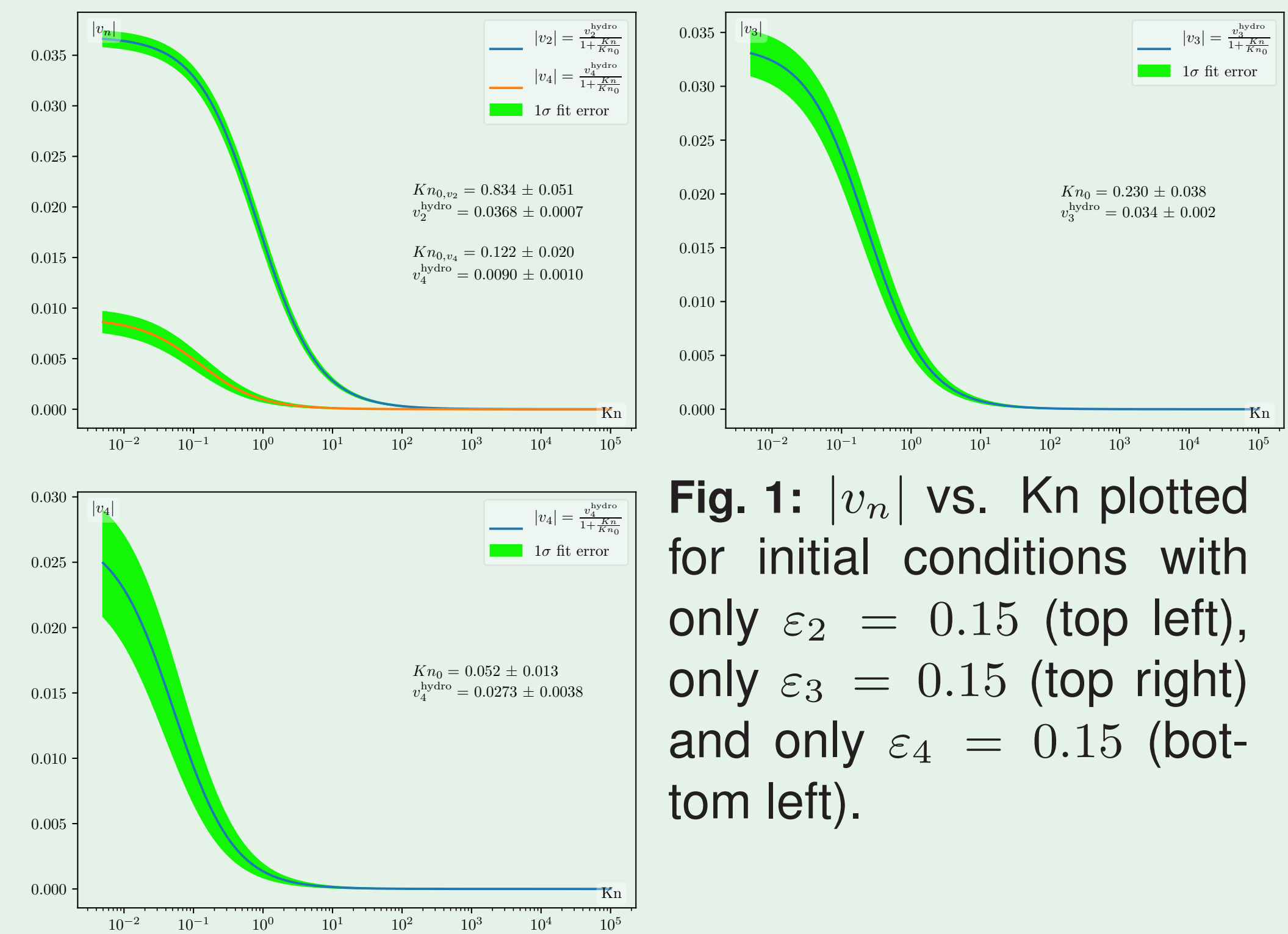


Fig. 1: $|v_n|$ vs. Kn plotted for initial conditions with only $\varepsilon_2 = 0.15$ (top left), only $\varepsilon_3 = 0.15$ (top right) and only $\varepsilon_4 = 0.15$ (bottom left).

Setup for MC initial condition

- Input: TGlauberMC [3] (Pb-Pb $\sqrt{s_{\text{NN}}} = 5.02$ TeV) $\rightarrow N_{\text{coll}}(x, y), N_{\text{part}}(x, y)$
- Energy density: $e(x, y)$
 - $N(x, y) = (1 - \xi)N_{\text{part}} + \xi N_{\text{coll}}$ with $\xi \approx 0.15$.
 - Smear the energy density as a Gaussian with width $R_N = \frac{1}{2} \sqrt{\frac{\sigma_{\text{inel}}^{\text{NN}}}{\pi}}$.
- For the particlization we convert $e(x, y)$ to $n(x, y)$ with the equations for an ideal gas in 2D. We checked the energy conservation in the process.
- Momentum isotropy, i.e. $v_n = 0$.
- We compute 10 runs over one initial condition with $5 \cdot 10^5$ particles each to reduce statistical errors.

Fluctuation characterization

- Fluctuations in the eccentricity probability distribution can be characterized by the elliptic-power law [4]:

$$p(\varepsilon_n) = \frac{2\alpha\varepsilon_n}{\pi} (1 - \varepsilon_n^2)^{\alpha-1} (1 - \varepsilon_0^2)^{\alpha+\frac{1}{2}} \int_0^\pi d\varphi (1 - \varepsilon_0\varepsilon_n \cos \varphi)^{-2\alpha-1}.$$

- For vanishing mean anisotropy ε_0 in the reaction plane:

$$p(\varepsilon_n) = 2\alpha\varepsilon_n (1 - \varepsilon_n^2)^{\alpha-1}.$$

- The relation between ε_n and v_n (for $n = 2, 3$) is: $v_n \approx \mathcal{K}_{n,n}\varepsilon_n$.

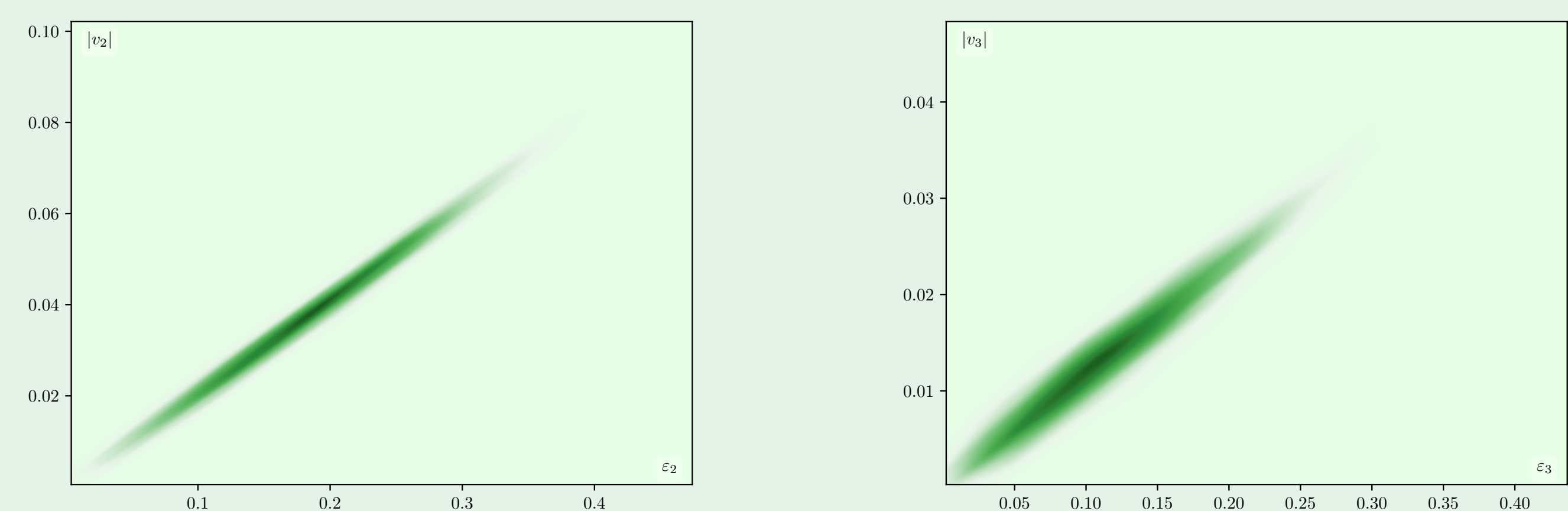


Fig. 2: Distribution of ε_2 and v_2 (left) and the distribution of ε_3 and v_3 (right) for $\langle \text{Kn} \rangle = 0.29$.

- The anisotropic flow distribution reads: $p(v_n) = \frac{1}{\mathcal{K}_{n,n}} p\left(\frac{v_n}{\mathcal{K}_{n,n}}\right)$.

Propagation of fluctuations

- The integral form of the power law is better suited for the distributions with non vanishing mean anisotropy.

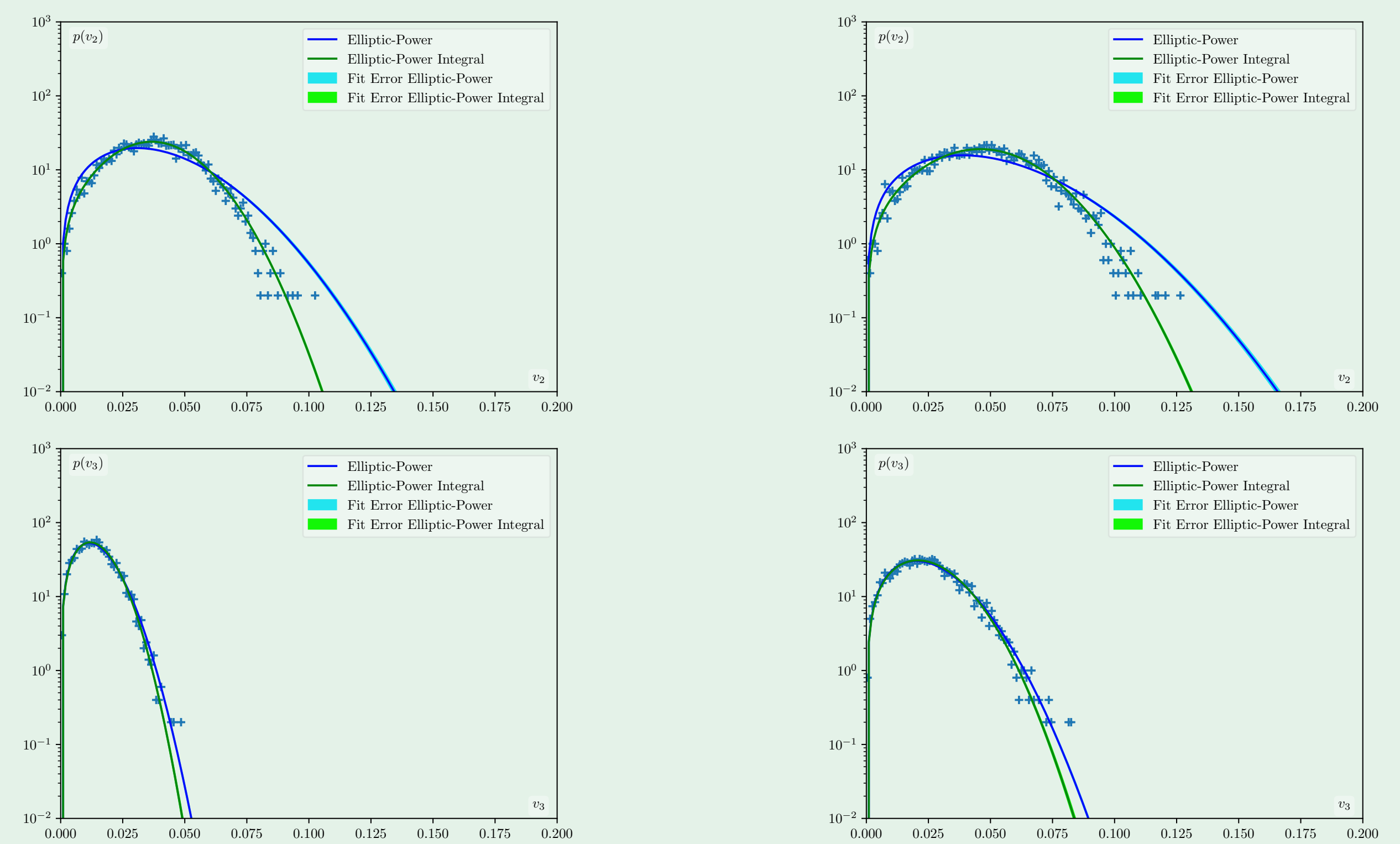


Fig. 3: Fitted distributions of v_2 (top row) and v_3 (bottom row) for $\langle \text{Kn} \rangle = 0.29$ (left column) and $\langle \text{Kn} \rangle = 0.07$ (right column).

- We find that for ε_2 and ε_3 the fluctuations in the distributions are washed out during the time evolution, resulting in larger values of α for the v_n distributions.
- The value of α decreases and that of v_0 increases with growing number of rescatterings. The computation with the largest Kn , approaching the free-streaming limit, yields a peaked $p(v_n)$, for which the value of α is limited by numerical fluctuations in the initial momentum distribution.

$b = 6$ fm			$b = 6$ fm		
ε_2	ε_0 or v_0	α	ε_3	ε_0 or v_0	α
$v_2, \langle \text{Kn} \rangle = 2.91$	-	62 ± 2	$v_3, \langle \text{Kn} \rangle = 2.91$	-	71 ± 4
$v_2, \langle \text{Kn} \rangle = 0.29$	0.0327 ± 0.0002	6400 ± 30	$v_3, \langle \text{Kn} \rangle = 0.29$	0.0088 ± 0.0002	802100 ± 700
$v_2, \langle \text{Kn} \rangle = 0.07$	0.0408 ± 0.0002	1560 ± 20	$v_3, \langle \text{Kn} \rangle = 0.07$	0.0153 ± 0.0003	5707 ± 100
		980 ± 20			1840 ± 40

Tab. 1: Fit values for the $\varepsilon_{2,3}$ and $v_{2,3}$ distributions for collisions at $b = 6$ fm. Cells with a "-" indicate fits with the distribution function where $v_0 = 0$.

Outlook Further transport calculations with smaller Kn are needed to see if the value of α decreases further and finally approaches the α value of the initial state eccentricity distribution.

The calculation will also be performed for different impact parameters.

References

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- [3] C. Loizides, J. Nagle, and P. Steinberg. Improved version of the PHOBOS Glauber Monte Carlo. *SoftwareX*, 1-2:13 – 18, 2015.
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