

Search for the QCD phase transition

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Student Lecture
Quark Matter 2019

“A theory is something nobody believes,
except the person who made it.
An experiment is something everybody
believes, except the person who made it.”

A. Einstein

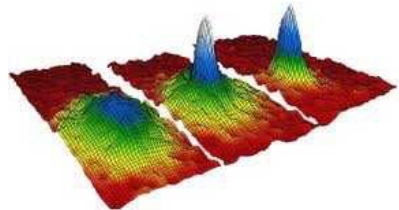
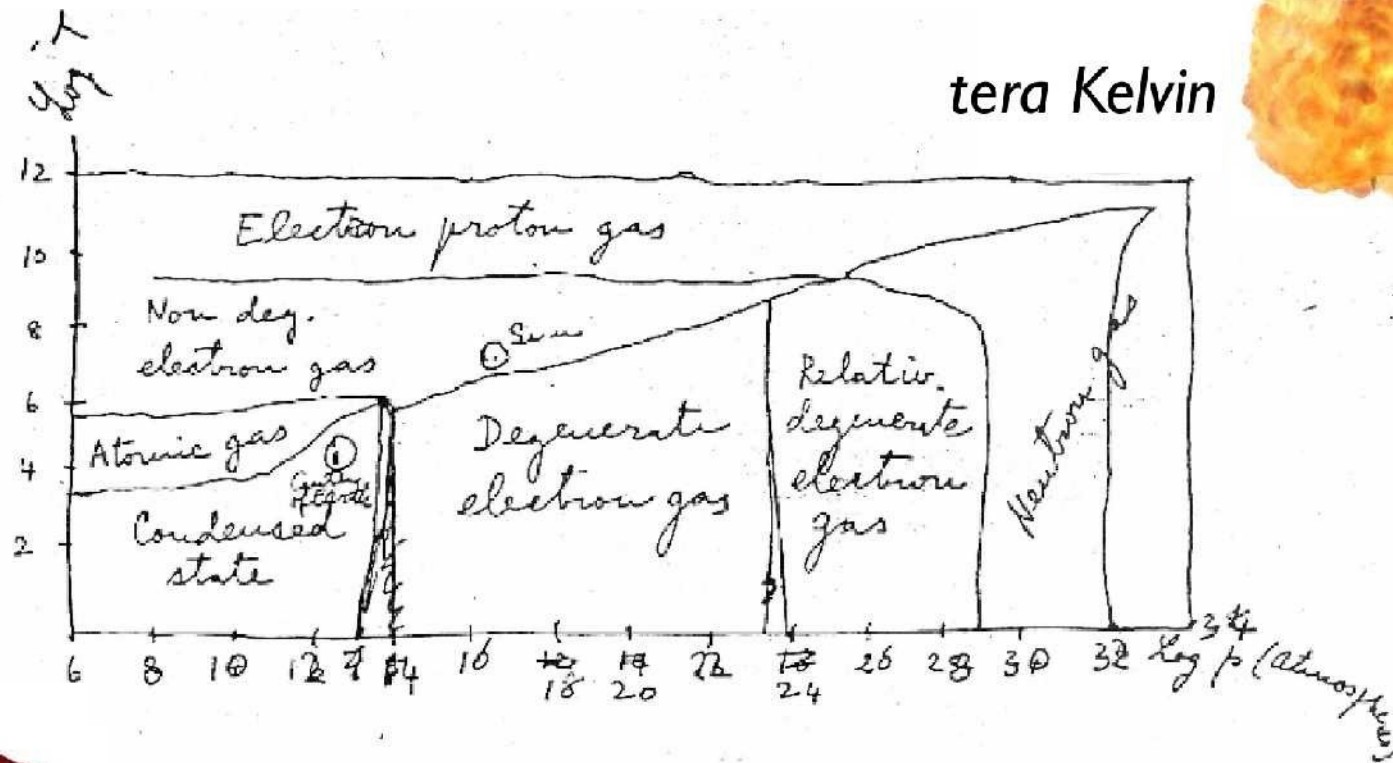
Outline

- Phase Transitions
- Cumulants: What are they and why are they useful
- Some preliminary experimental results and what they could mean
 - Some tricky experimental issues
 - Comparing data with Theory
 - Cumulants and correlations
- Spinodal instability
- Summary

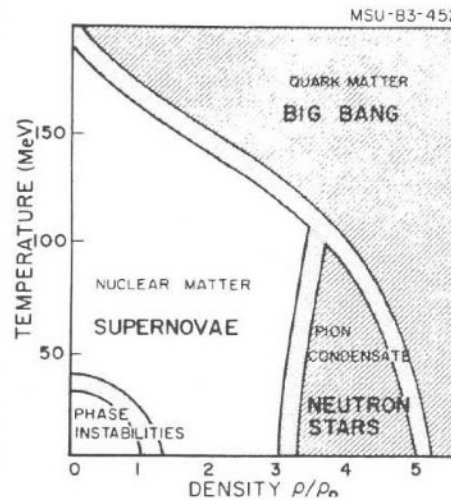
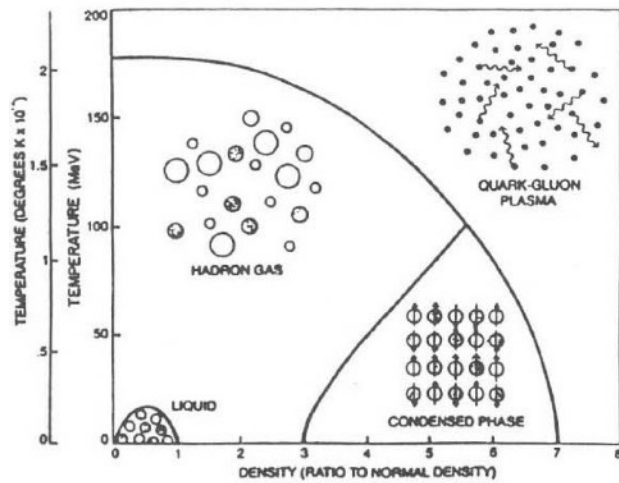
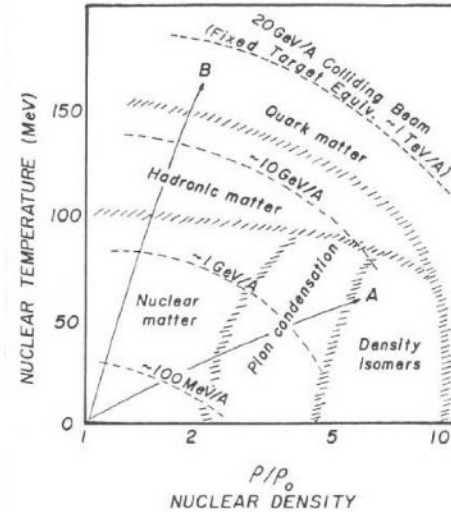
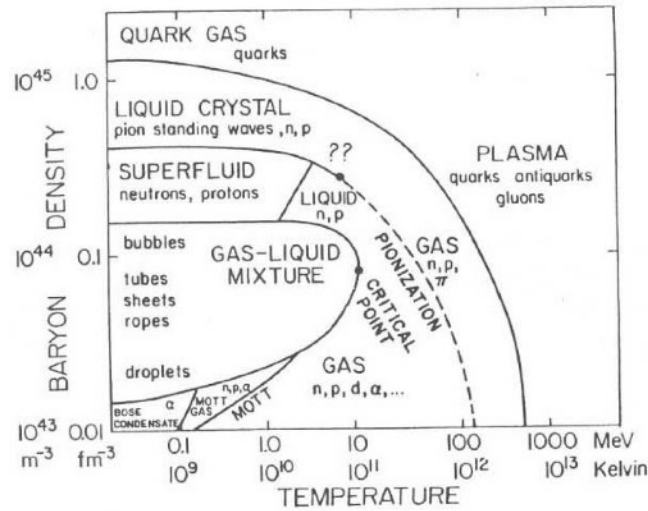
An old question



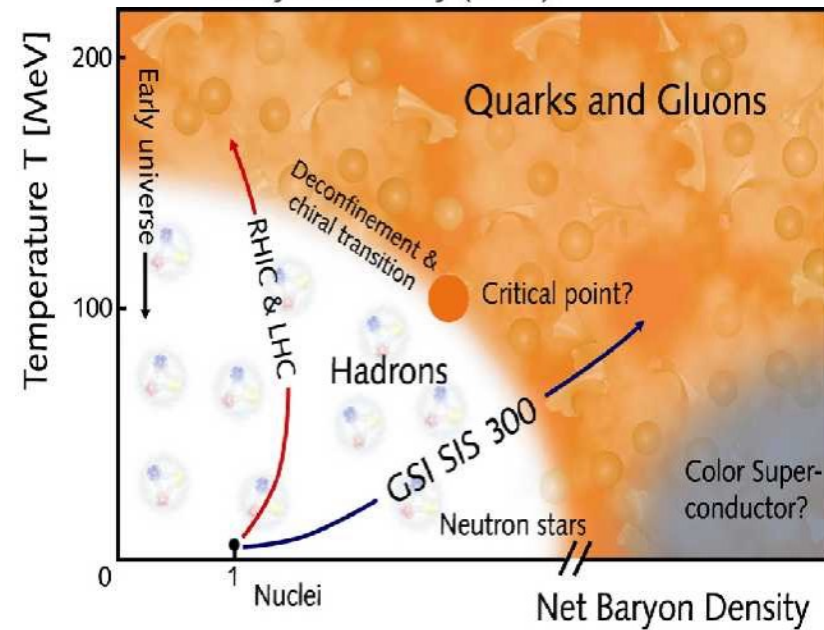
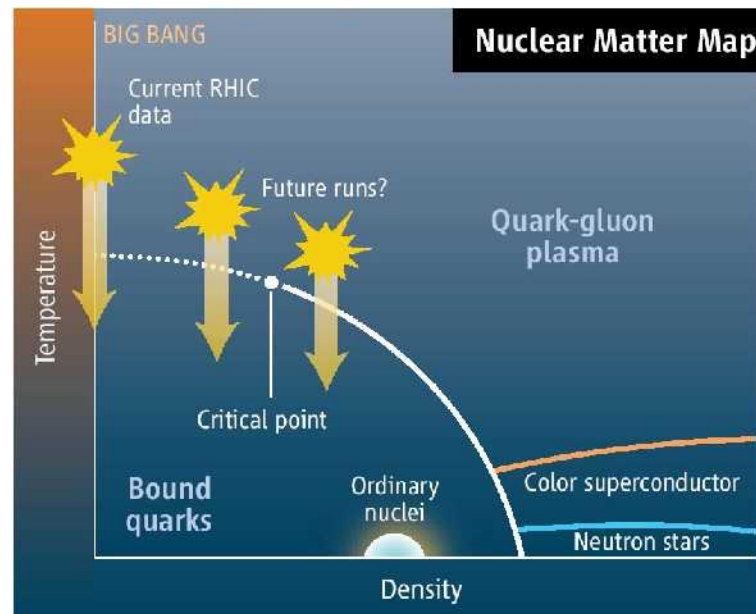
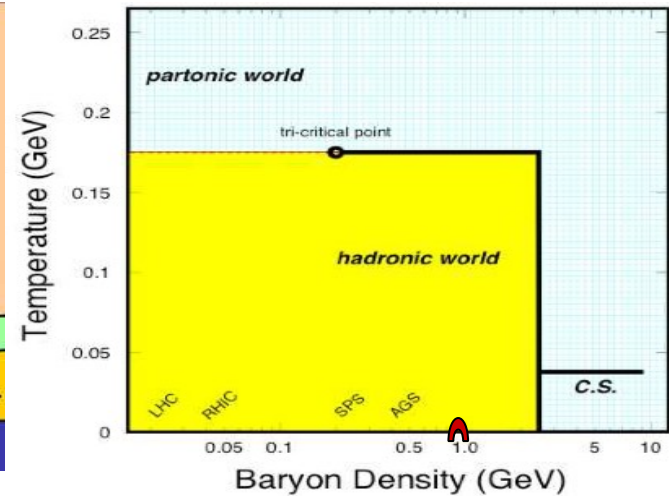
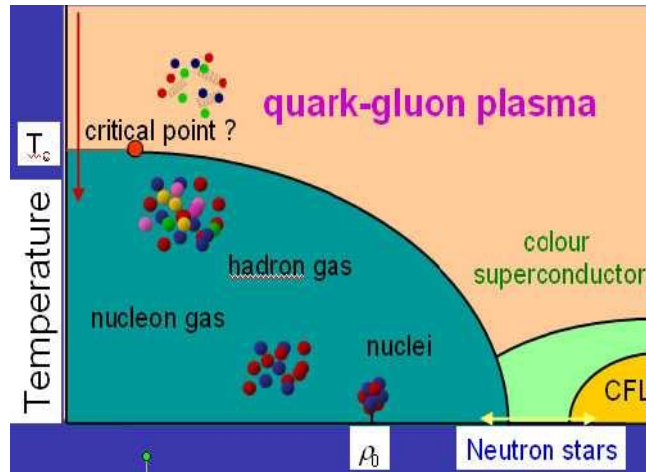
Fermi 1953



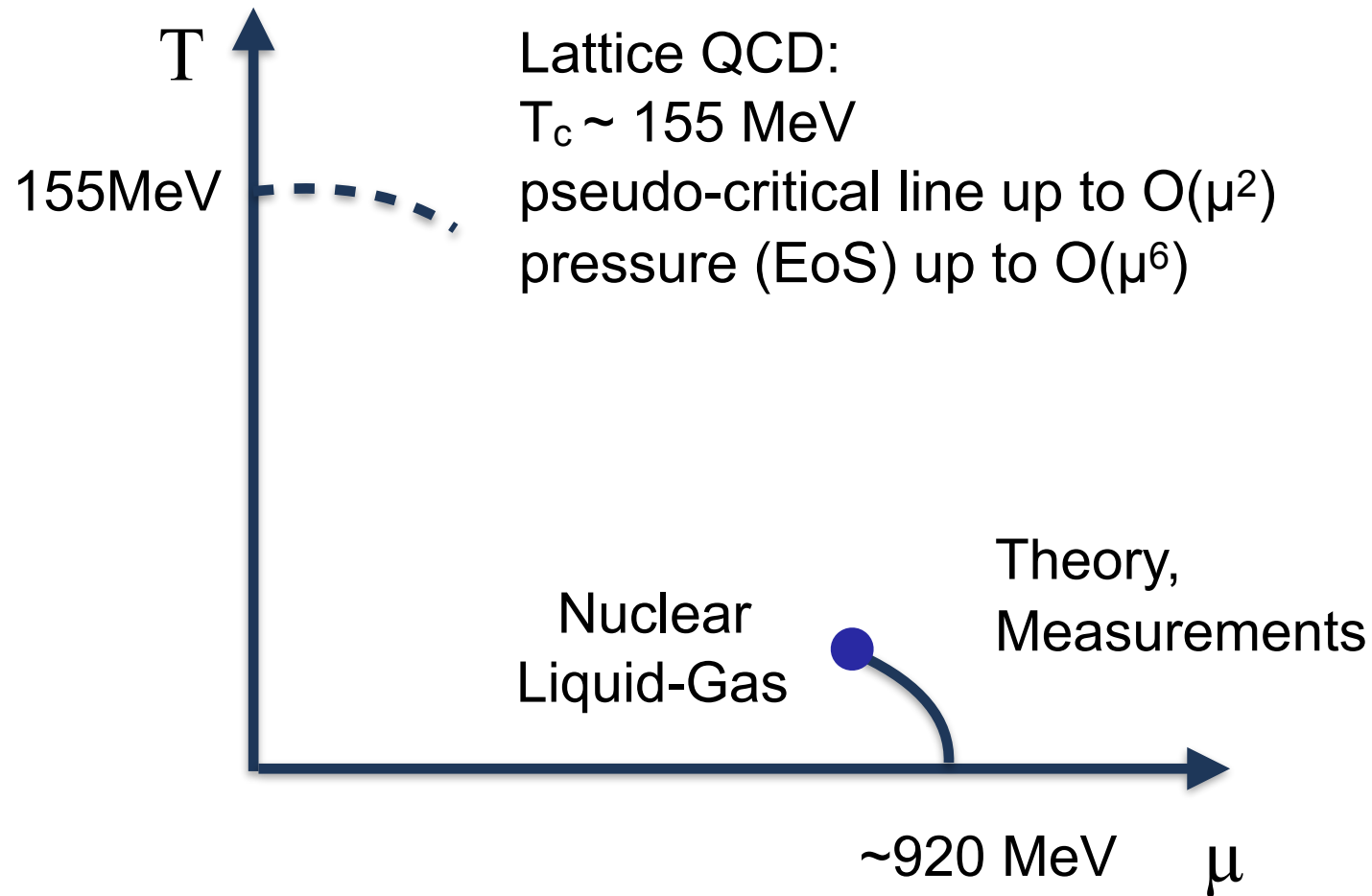
discussed for many years



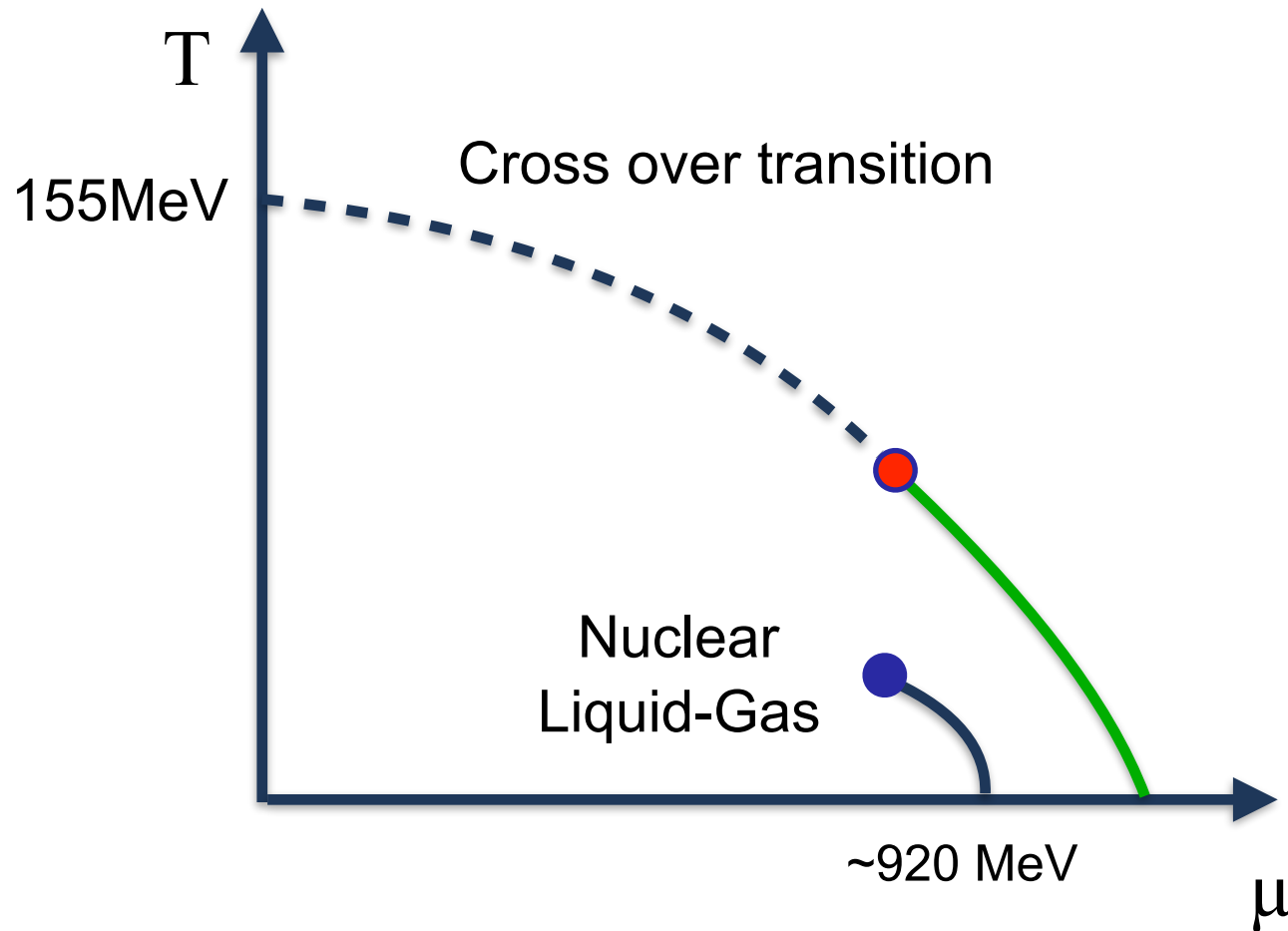
gets more colorful ...



What we know about the Phase Diagram



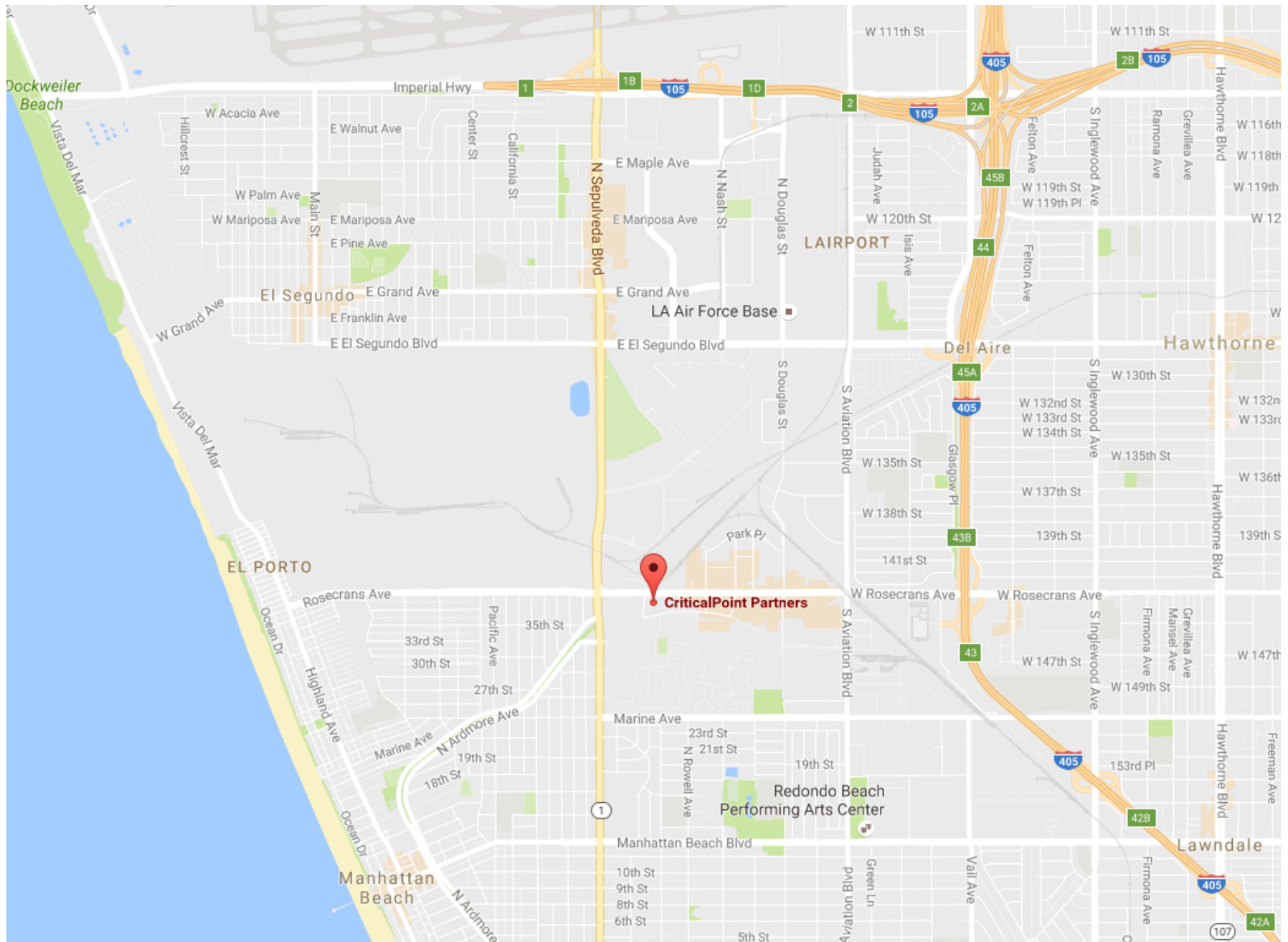
What we “hope” for



NB: critical point of water is at $T=647\text{ K}$ and $p=22.06\text{ MPa}$

Is there a critical point?

Google finds everything...



Phase Transitions

Examples:

Water - vapor (liquid - gas)

Water - ice

Ferromagnet

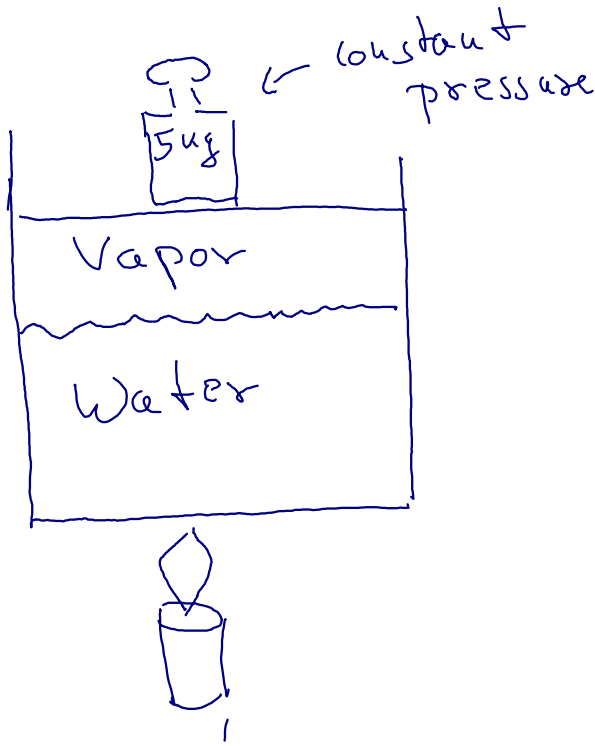
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Order parameter: Tells in which phase the system is
Examples ?

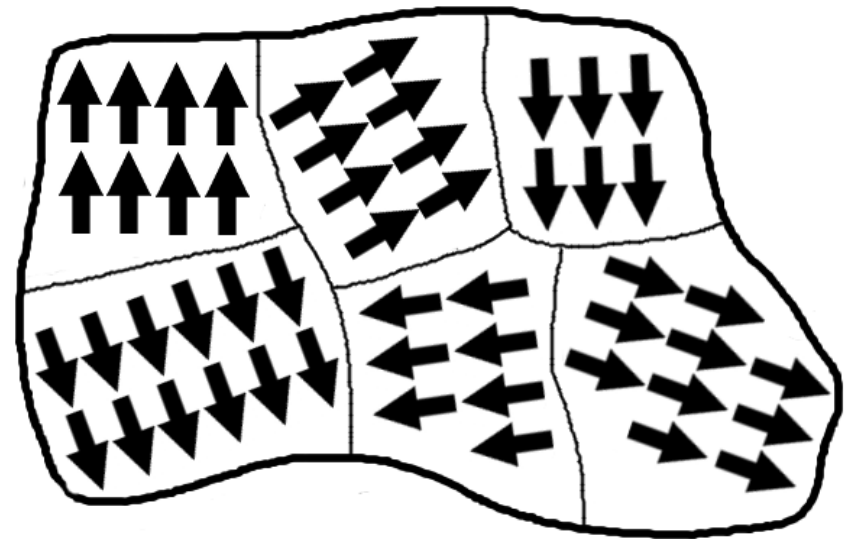
Control parameter: Moves system from one phase to another
Examples ?

Phase co-existence: Two or more phases can exist together
Examples ?

Phase Co-Existence

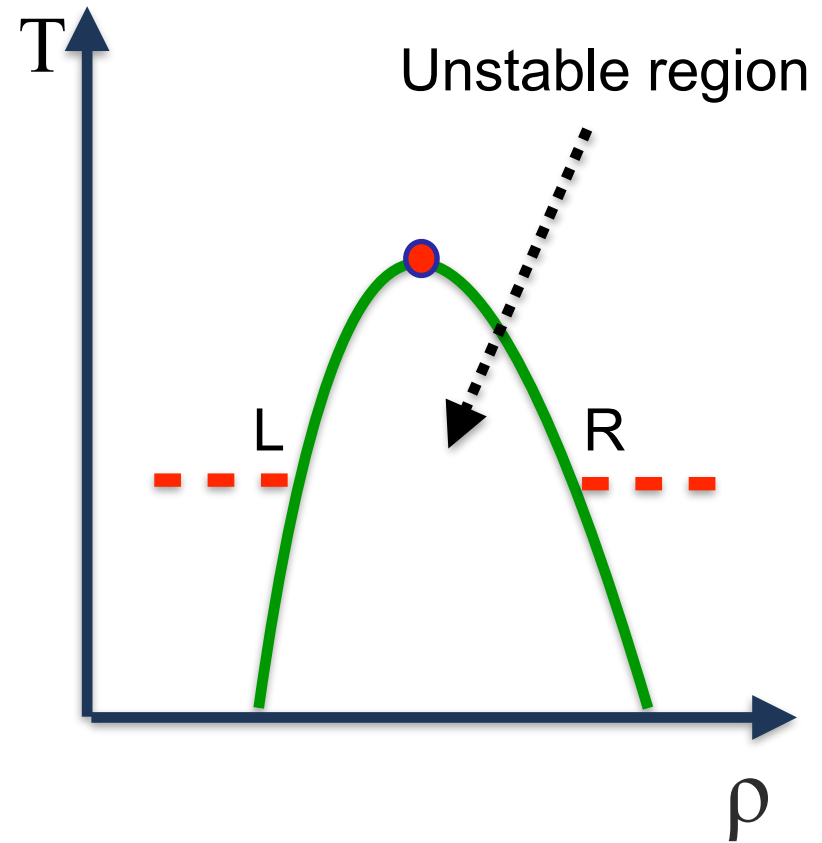
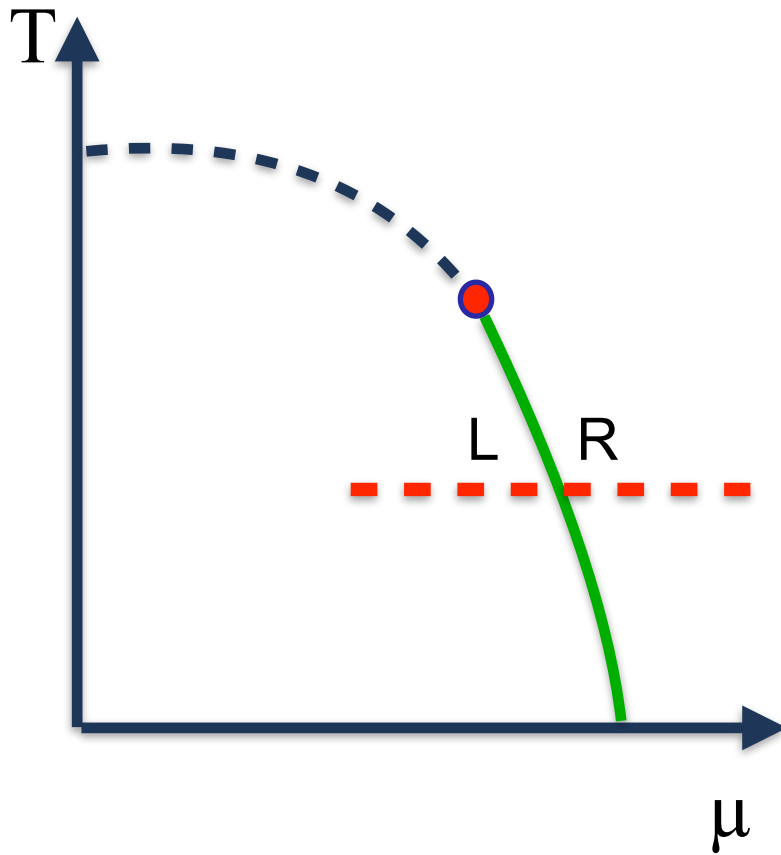


Water-vapor co-existence
a.k.a your water kettle

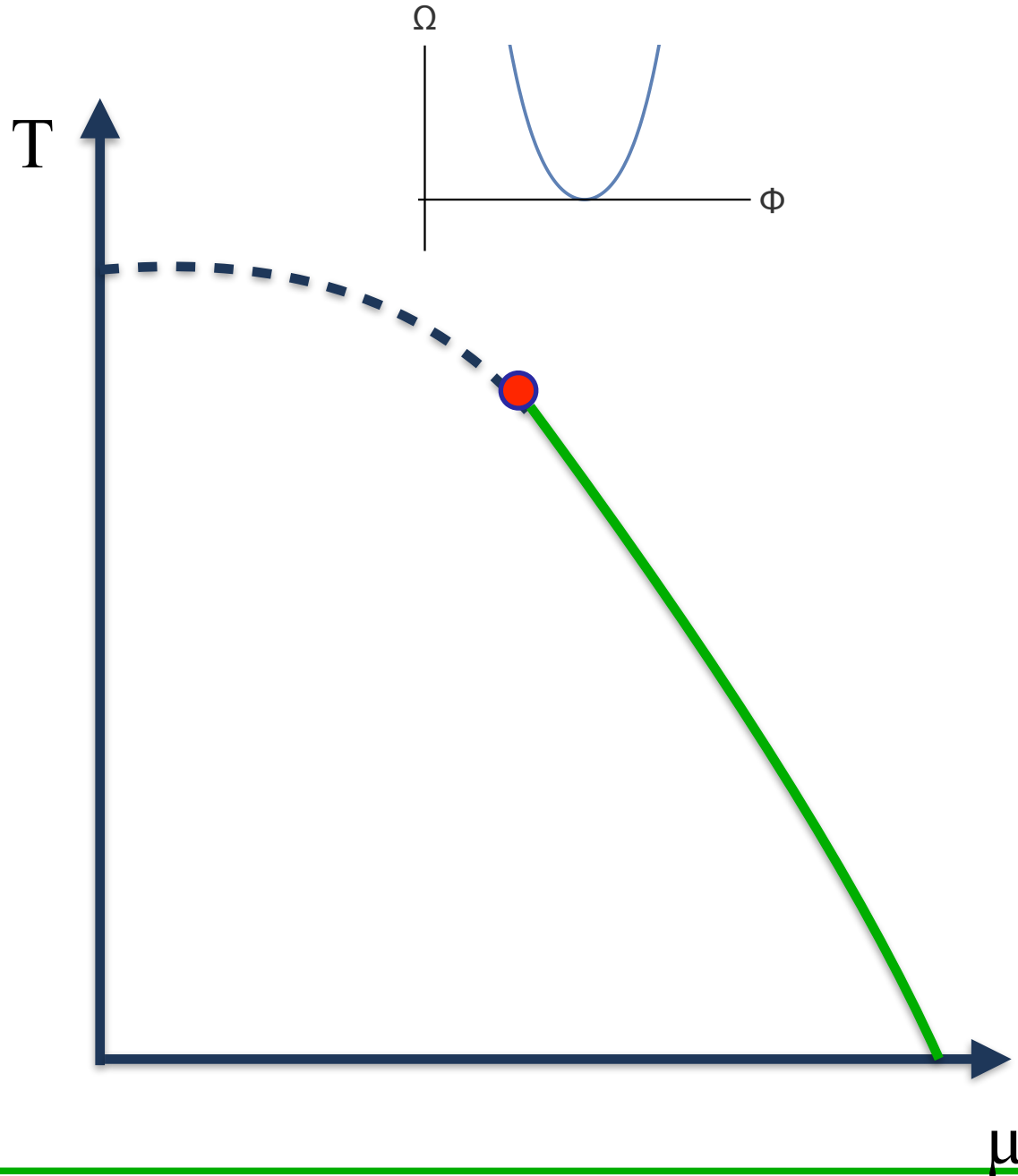


Ferro-magnet
Weiss domains

Phase diagrams



Free Energy



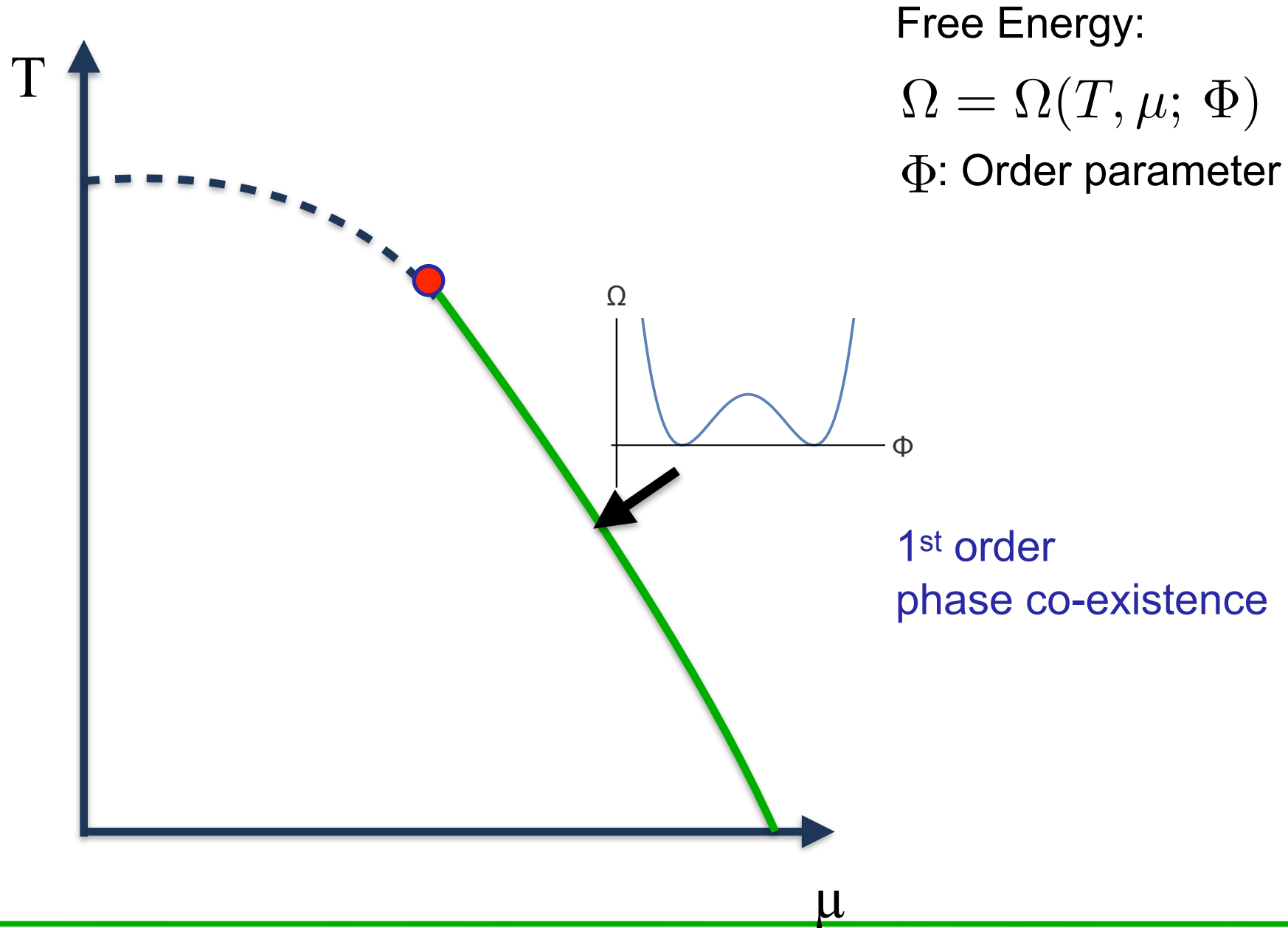
Free Energy:

$$\Omega = \Omega(T, \mu; \Phi)$$

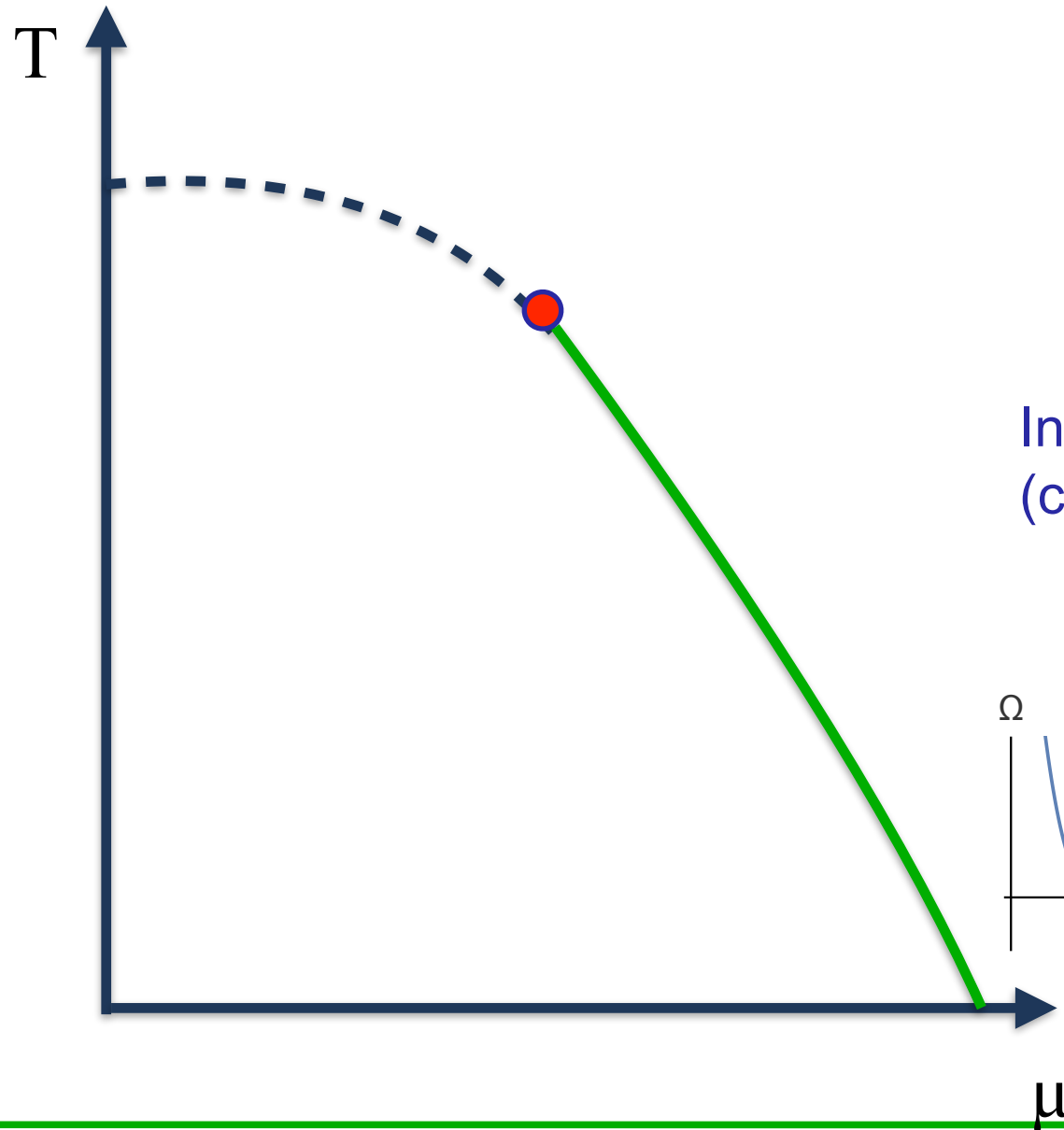
Φ : Order parameter

What we are used to:
One minimum

Free Energy



Free Energy

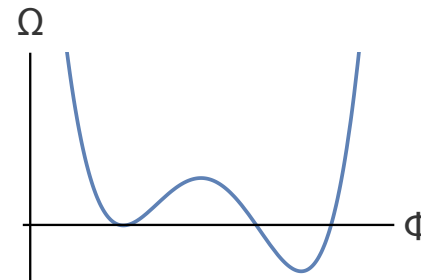


Free Energy:

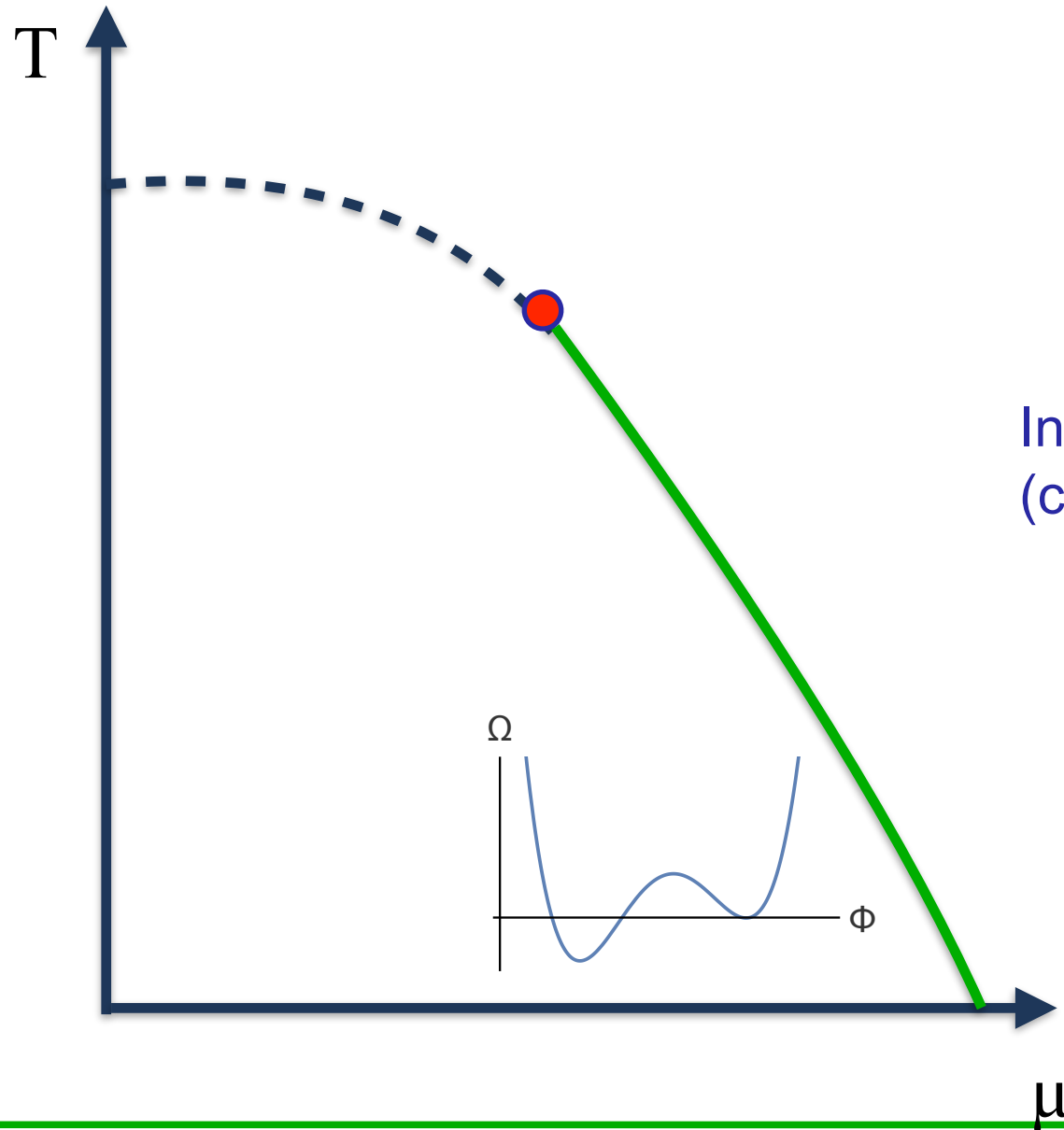
$$\Omega = \Omega(T, \mu; \Phi)$$

Φ : Order parameter

In “dense” phase
(close to transition)



Free Energy



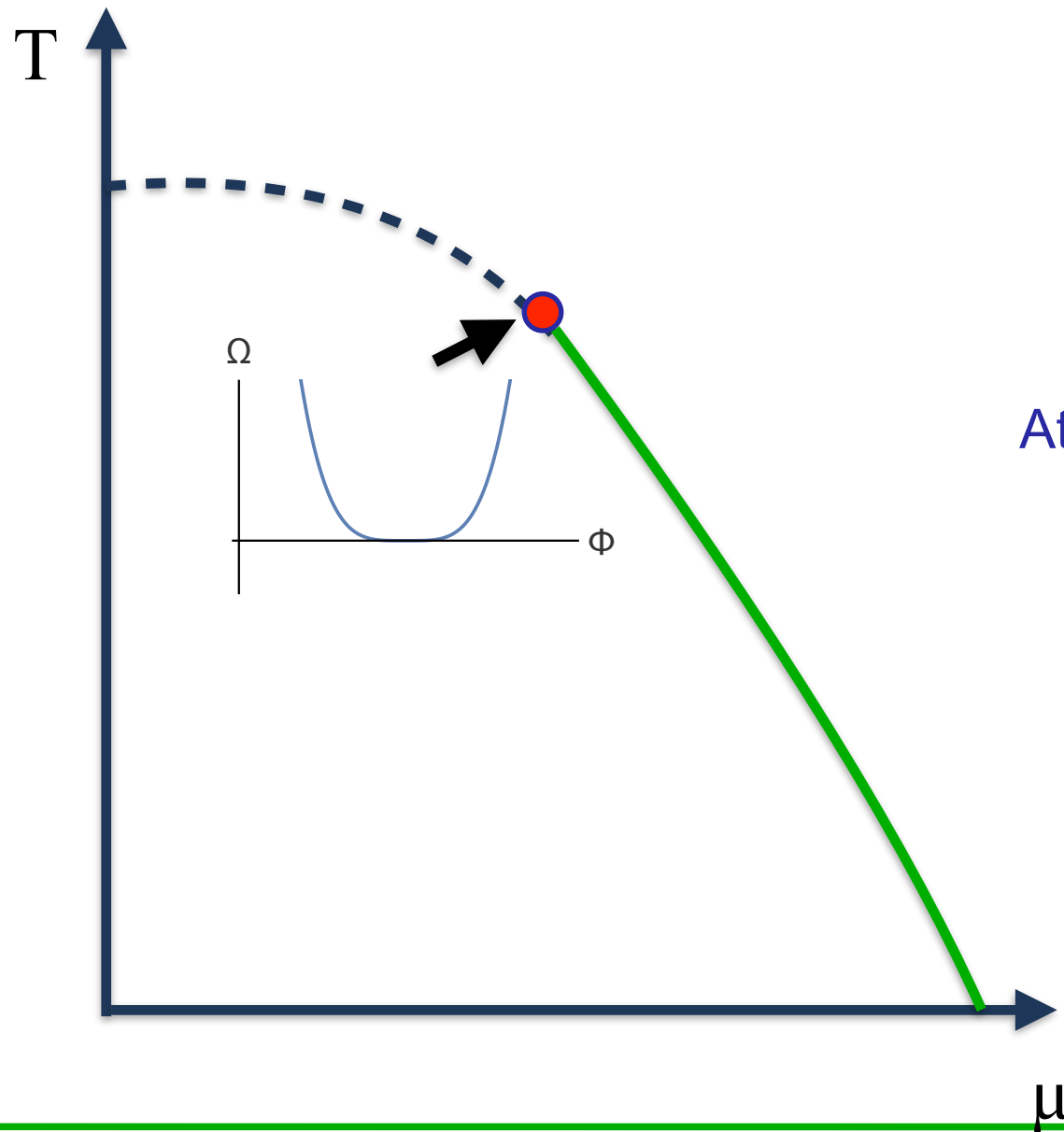
Free Energy:

$$\Omega = \Omega(T, \mu; \Phi)$$

Φ : Order parameter

In “dilute” phase
(close to transition)

Free Energy



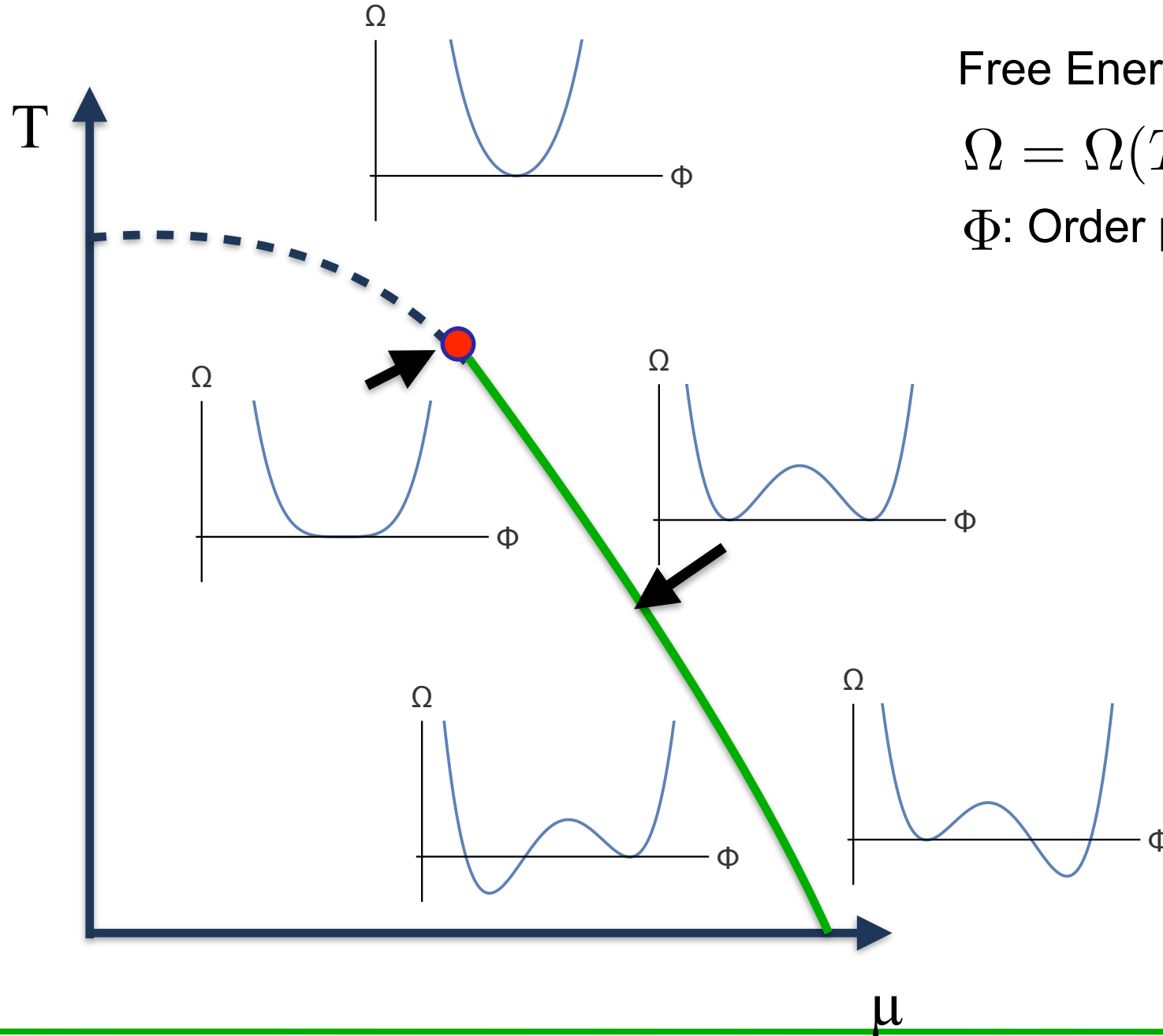
Free Energy:

$$\Omega = \Omega(T, \mu; \Phi)$$

Φ : Order parameter

At the critical point

Free Energy

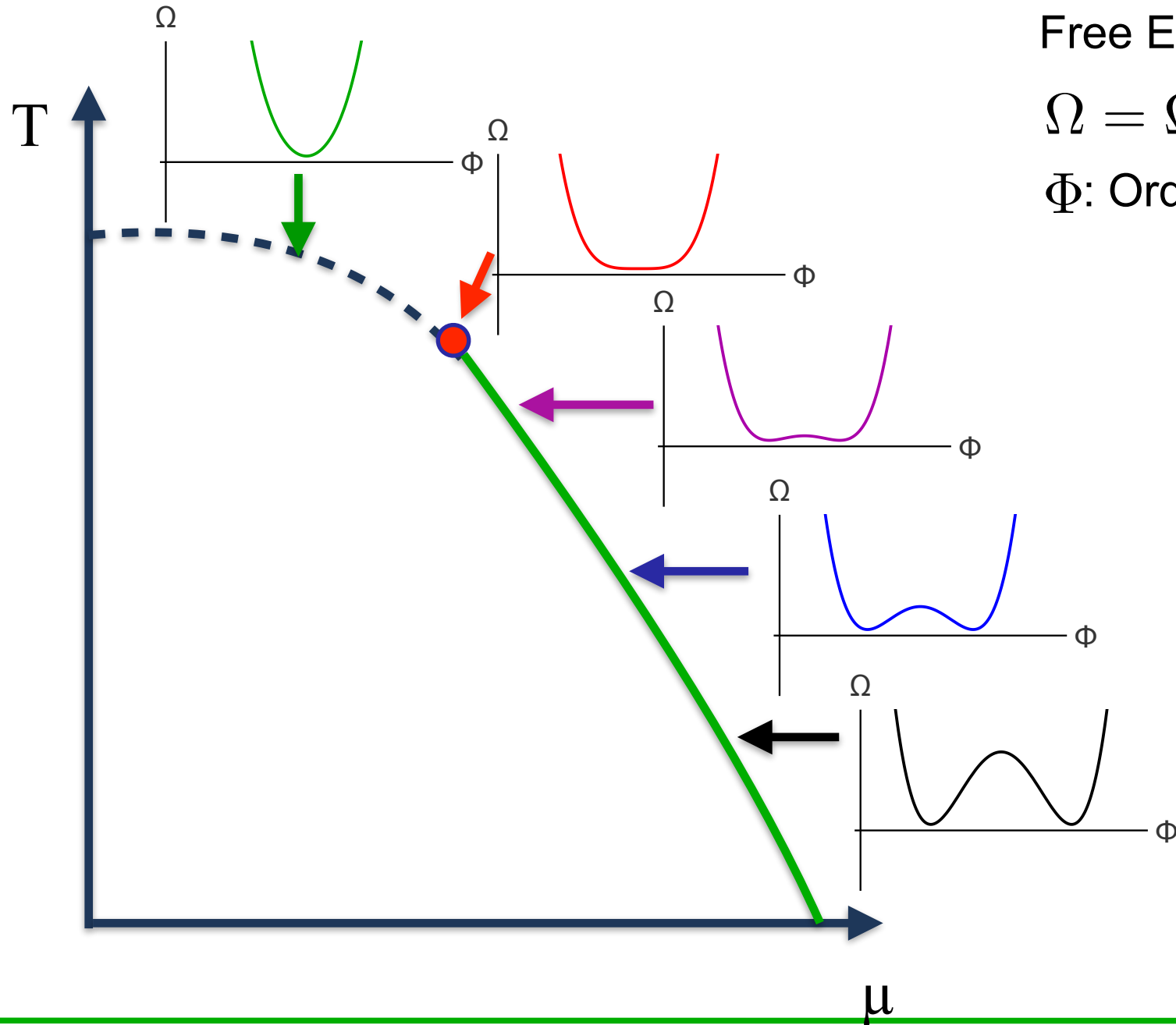


Free Energy:

$$\Omega = \Omega(T, \mu; \Phi)$$

Φ : Order parameter

Free Energy

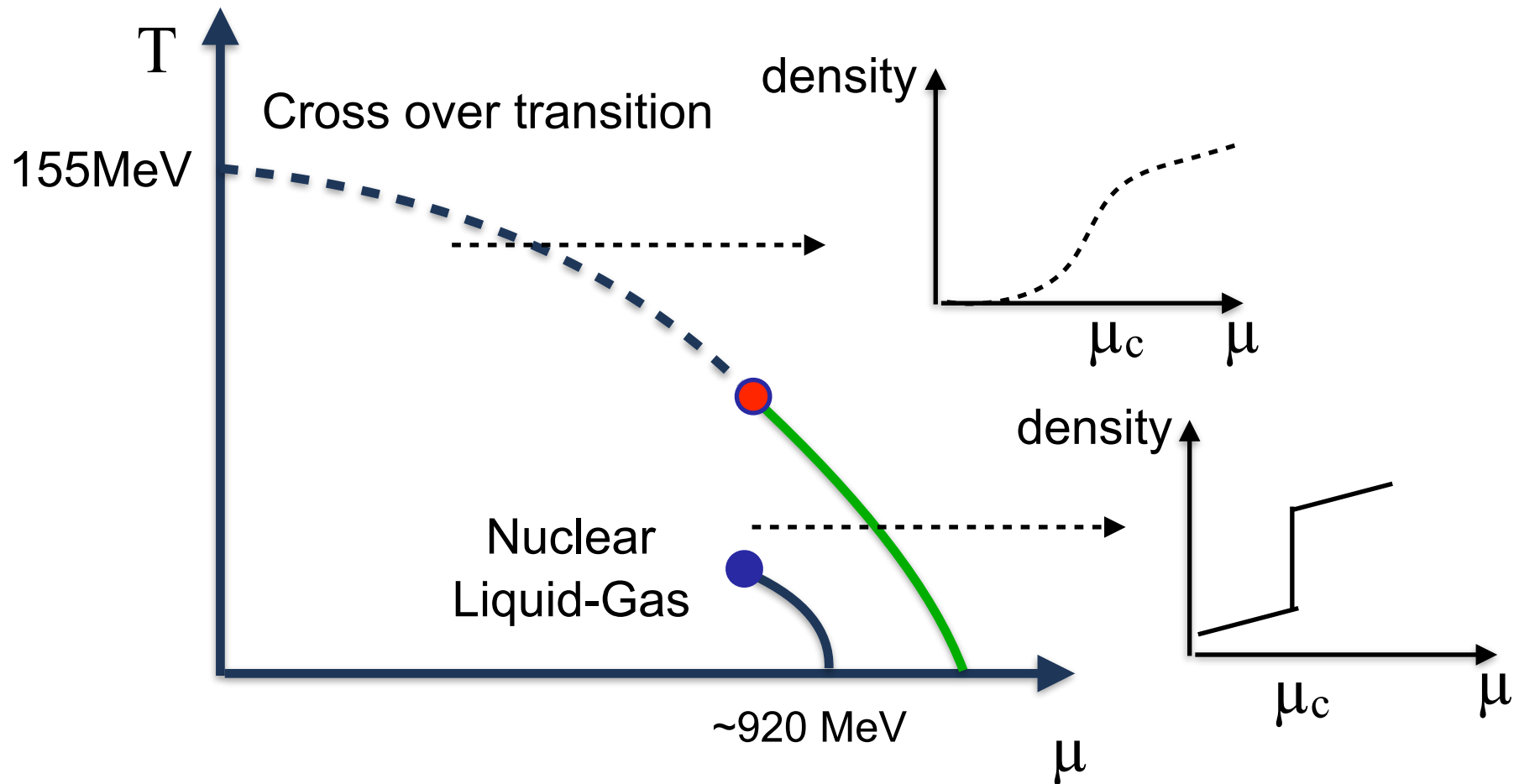


Free Energy:

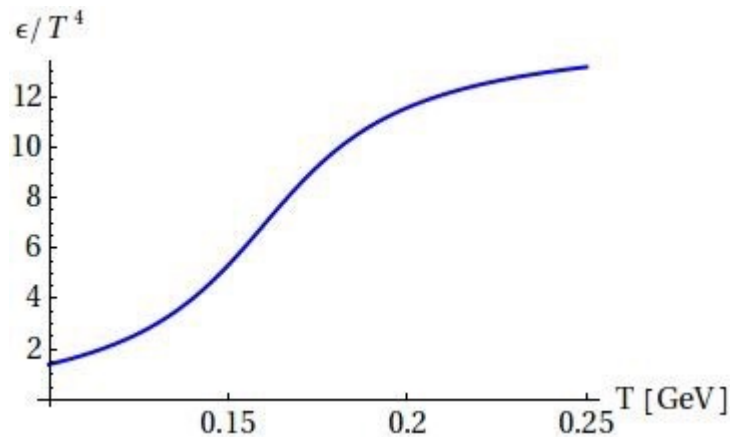
$$\Omega = \Omega(T, \mu; \Phi)$$

Φ : Order parameter

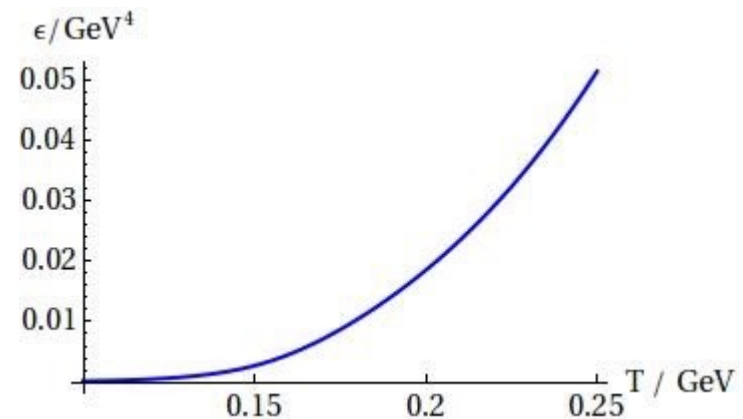
Looking for signs of a transition



Cumulants and phase structure



What we always see....

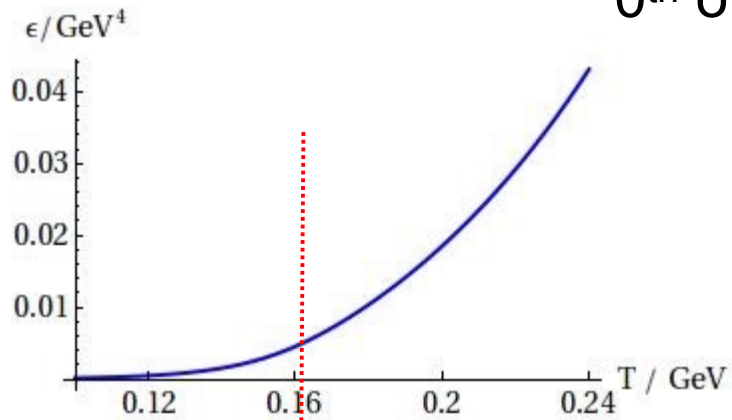


What it really means....

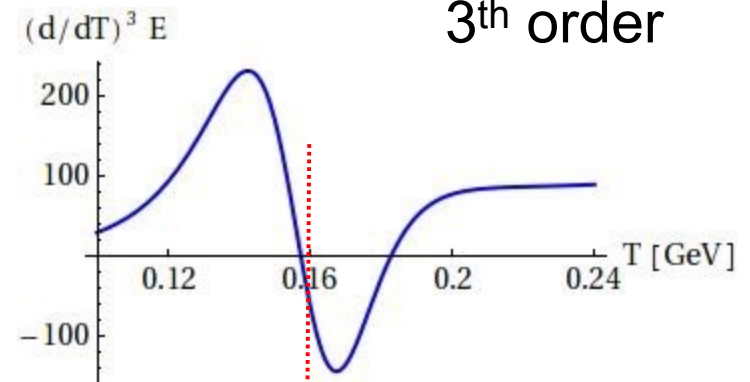
“ T_c ” \sim 160 MeV

Derivatives

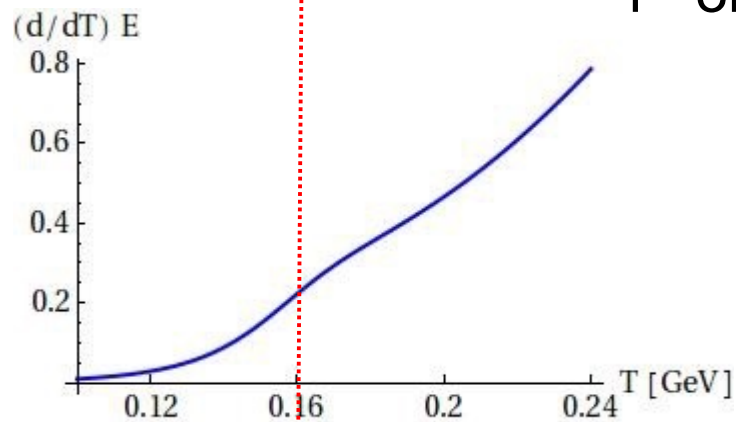
0th order



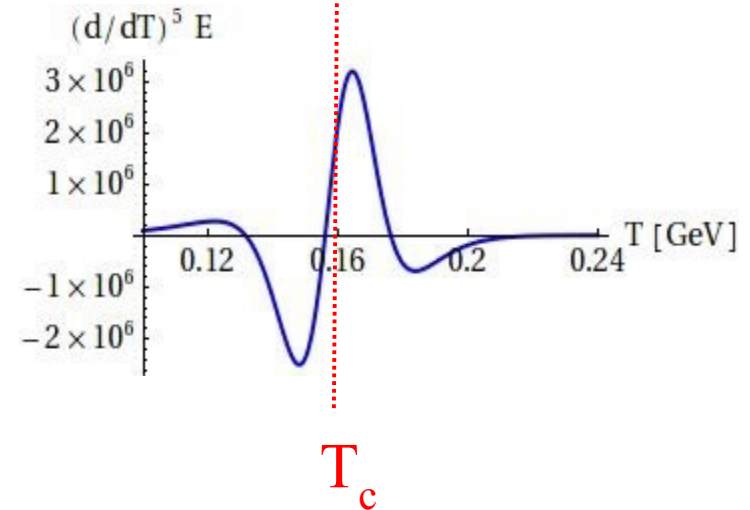
3th order



1st order



5th order



T_c

T_c

How to measure derivatives

At $\mu = 0$:

$$Z = \text{tr} e^{-\hat{E}/T + \mu/T \hat{N}_B}$$

$$\langle E \rangle = \frac{1}{Z} \text{tr} \hat{E} e^{-\hat{E}/T + \mu/T \hat{N}_B} = -\frac{\partial}{\partial 1/T} \ln(Z)$$

$$\langle (\delta E)^2 \rangle = \langle E^2 \rangle - \langle E \rangle^2 = \left(-\frac{\partial}{\partial 1/T} \right)^2 \ln(Z) = \left(-\frac{\partial}{\partial 1/T} \right) \langle E \rangle$$

$$\langle (\delta E)^n \rangle = \left(-\frac{\partial}{\partial 1/T} \right)^{n-1} \langle E \rangle$$

Cumulants of Energy measure the temperature derivatives of the EOS

Cumulants of Baryon number measure the chem. pot. derivatives of the EOS

Cumulants of (Baryon) Number

$$K_n = \frac{\partial^n}{\partial(\mu/T)^n} \ln Z = \frac{\partial^{n-1}}{\partial(\mu/T)^{n-1}} \langle N \rangle$$

$$K_1 = \langle N \rangle, \quad K_2 = \langle N - \langle N \rangle \rangle^2, \quad K_3 = \langle N - \langle N \rangle \rangle^3$$

Cumulants scale with volume (extensive): $K_n \sim V$

Volume not well controlled in heavy ion collisions

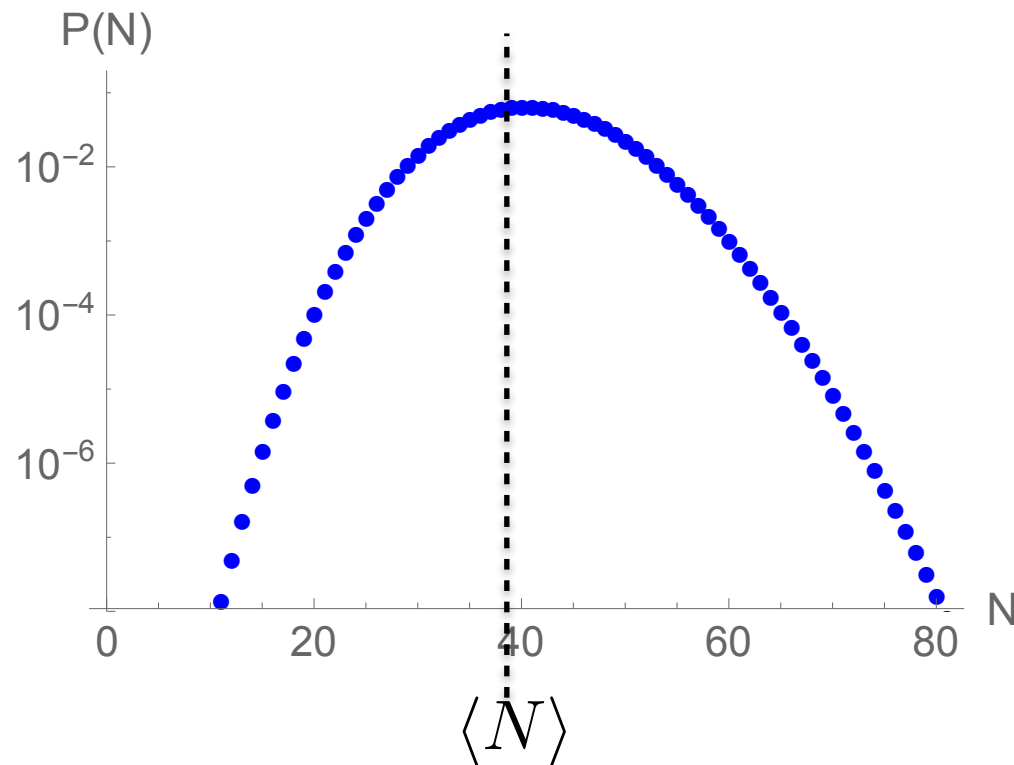
$$\text{Cumulant Ratios:} \quad \frac{K_2}{\langle N \rangle}, \quad \frac{K_3}{K_2}, \quad \frac{K_4}{K_2}$$

Measuring cumulants (derivatives)

$$K_2 = \langle N - \langle N \rangle \rangle^2 = \sum_N P(N) (N - \langle N \rangle)^2$$

$$K_3 = \langle N - \langle N \rangle \rangle^3 = \sum_N P(N) (N - \langle N \rangle)^3$$

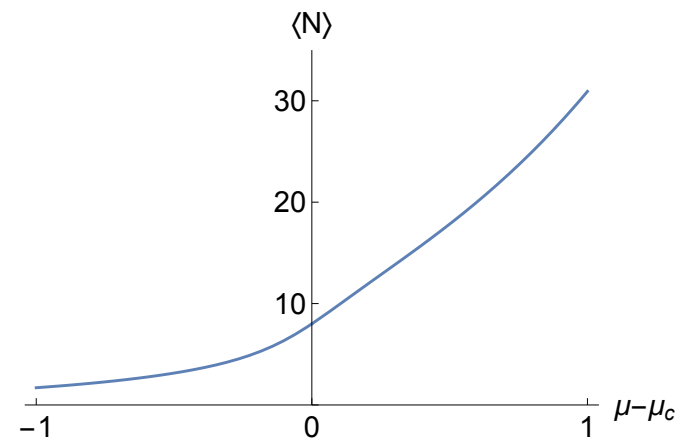
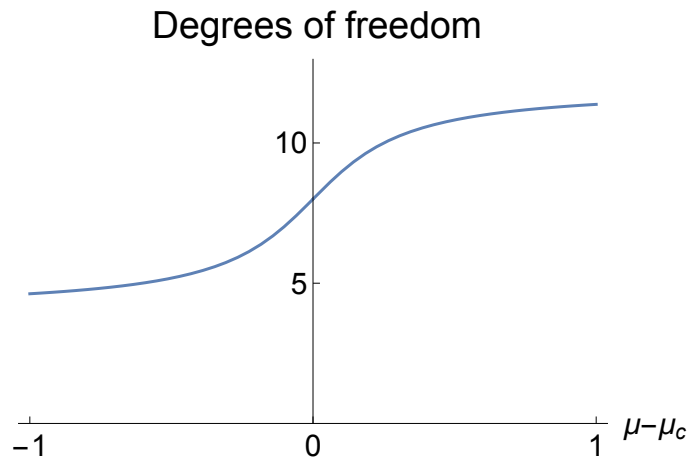
$$P(N) = \frac{N_{events}(N)}{N_{events}(total)}$$

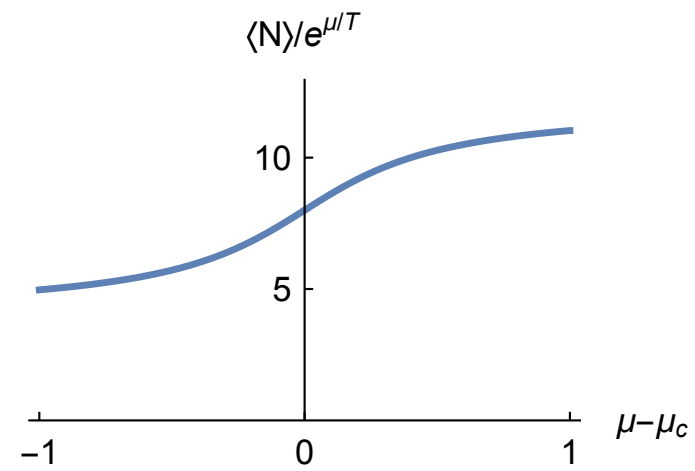
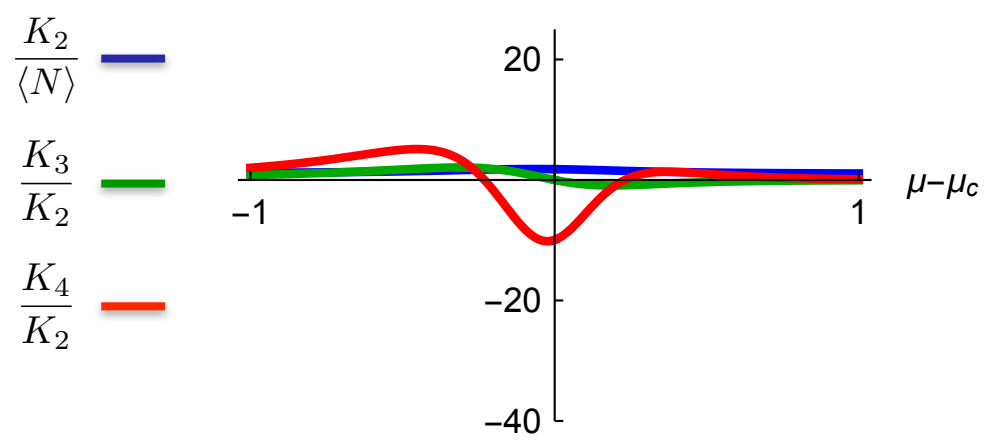
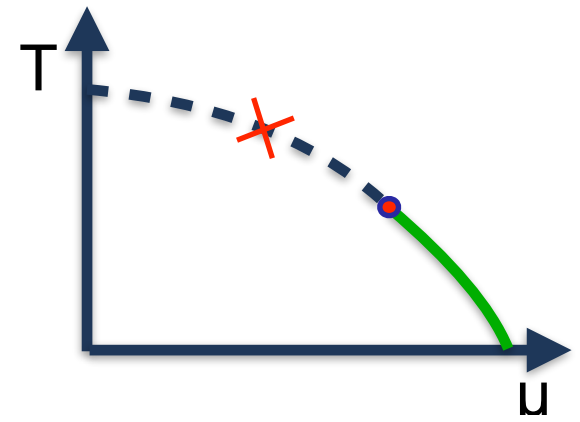


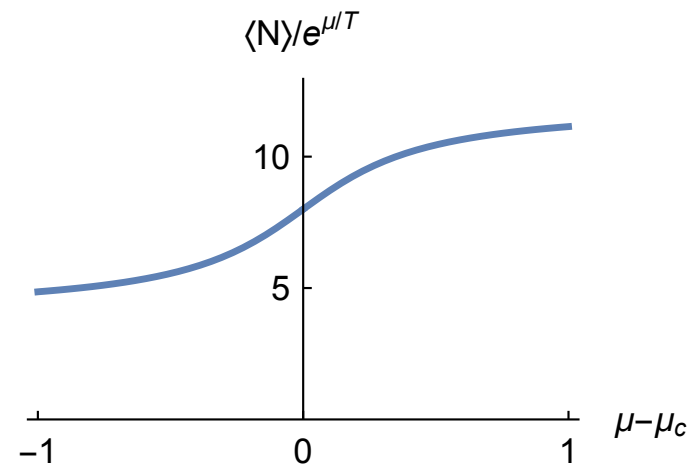
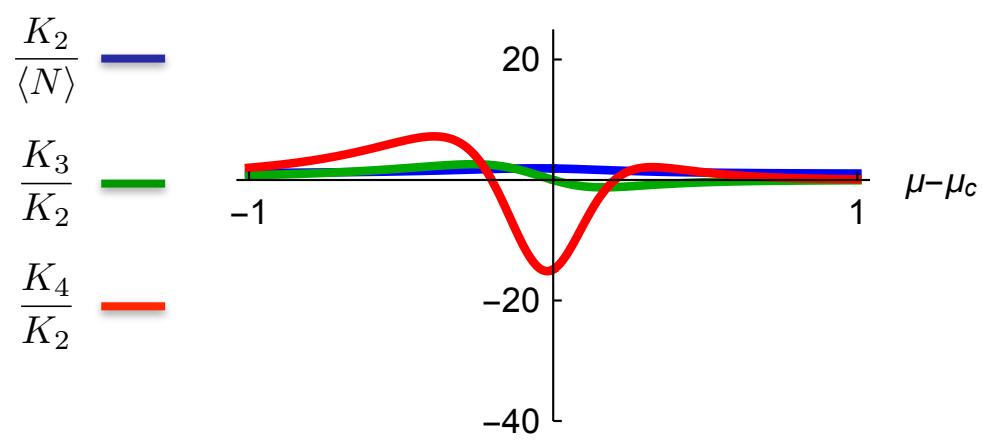
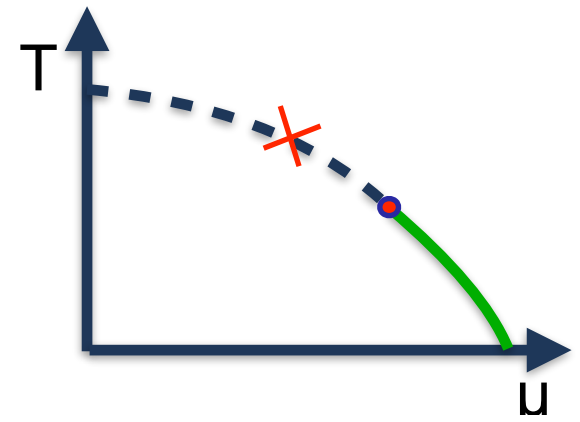
Simple model

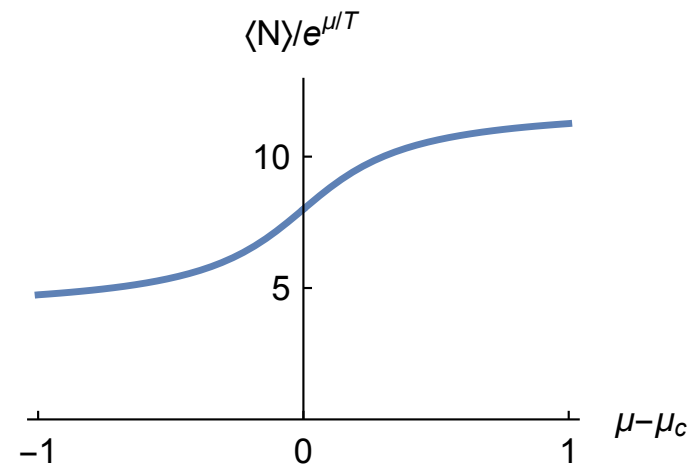
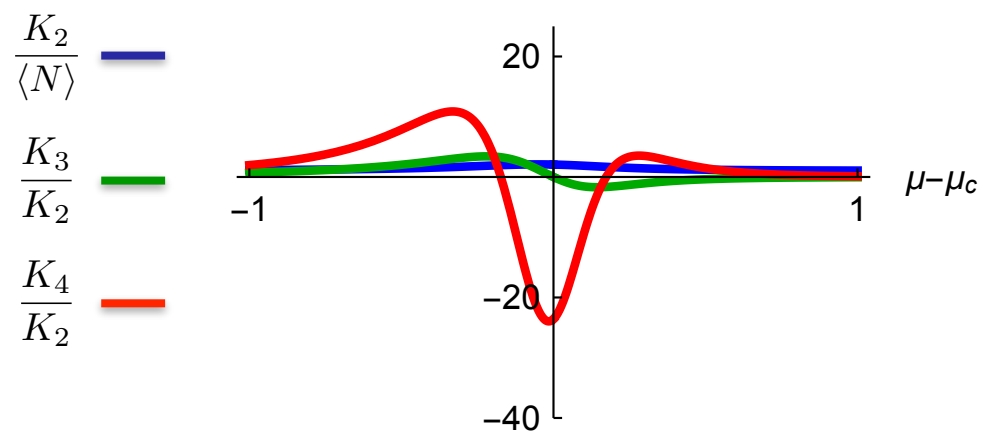
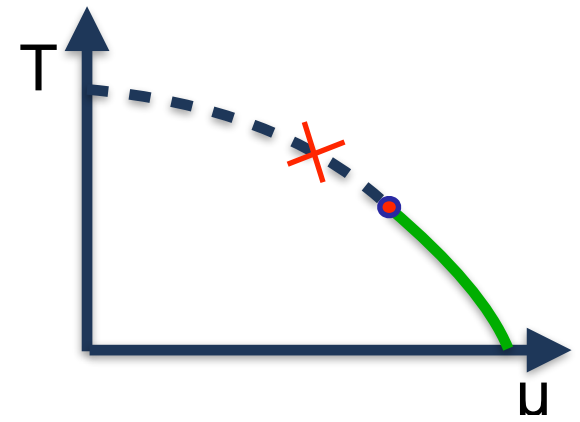
Change degrees of freedom
at phase transition

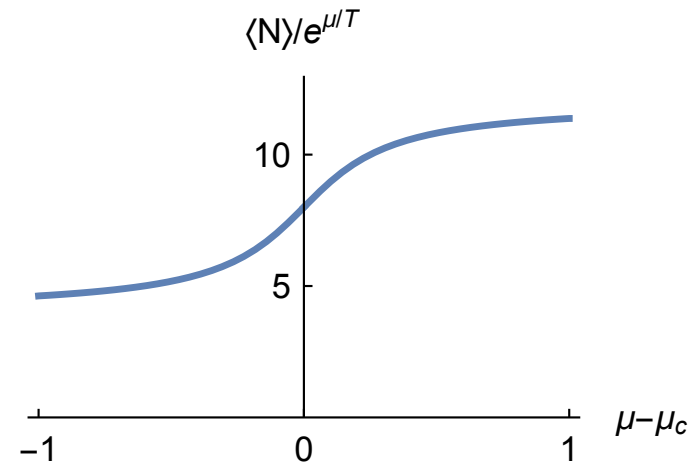
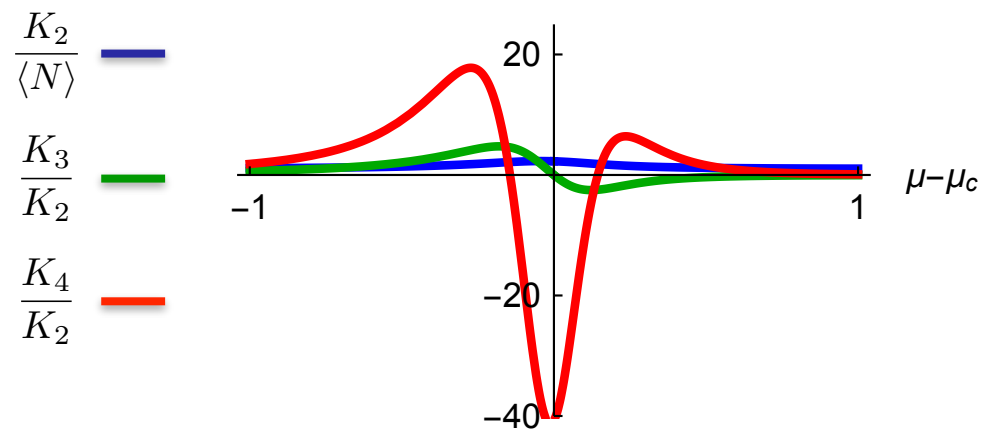
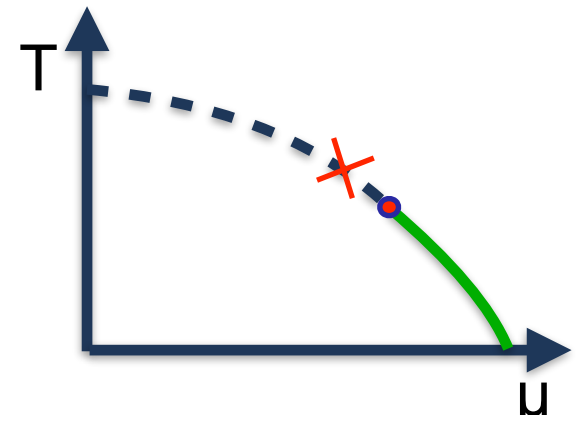
$$\langle N \rangle = \text{dof}(\mu) e^{\mu/T} \int d^3p e^{-E/T}$$







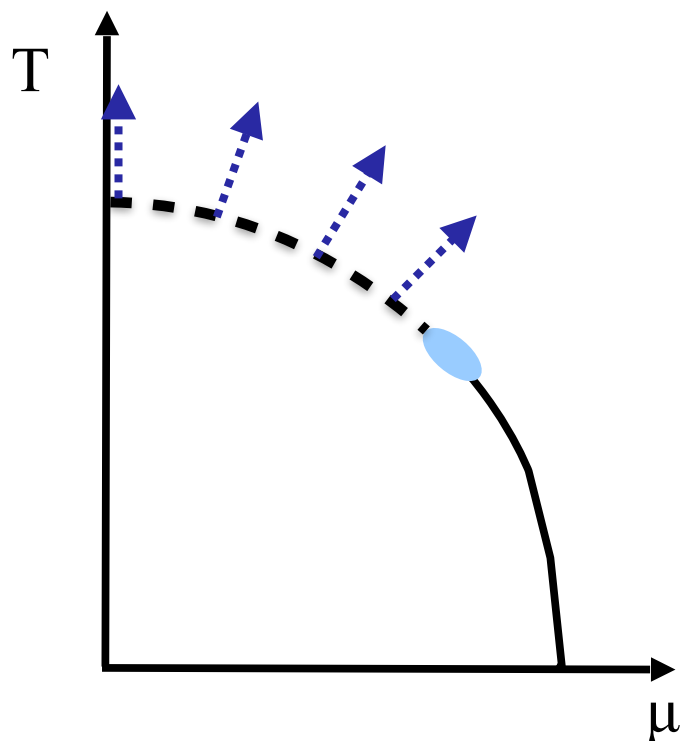




Close to $\mu=0$

$$F = F(r), \quad r = \sqrt{T^2 + a\mu^2}$$

$a \sim$ curvature of critical line



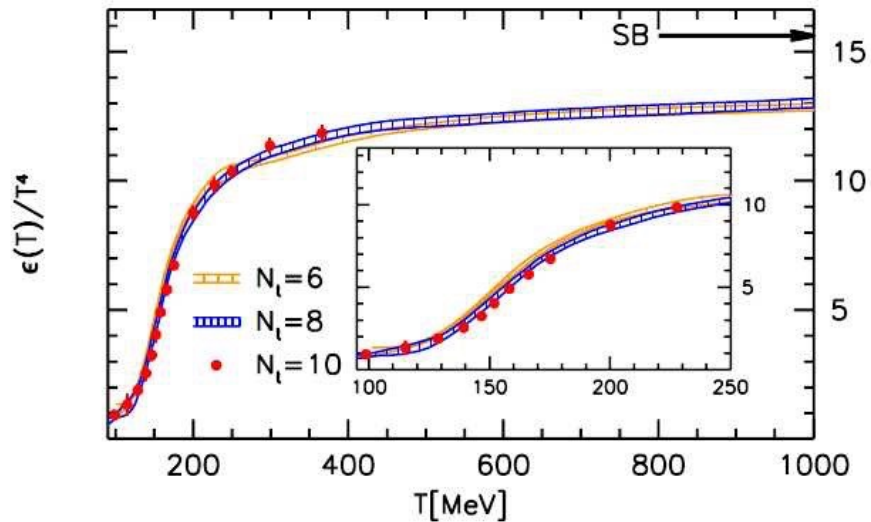
$$\frac{\partial^2}{\partial \mu^2} F(T, \mu)|_{\mu=0} = \frac{a}{T} \frac{\partial}{\partial T} F(T, \mu=0) \sim \langle E \rangle$$

Needs higher order cumulants (derivatives)
at $\mu \sim 0$

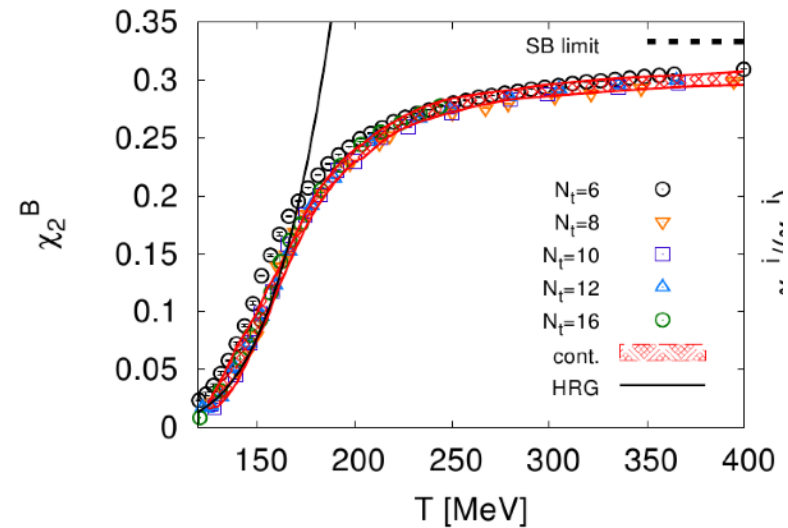
Lattice at $\mu=0$

Equation of state

S. Borsanyi et al, JHEP 1011 (2010) 077



Second order Cumulant



$$\frac{\partial^2}{\partial \mu^2} F(T, \mu)|_{\mu=0} = \frac{a}{T} \frac{\partial}{\partial T} F(T, \mu = 0) \sim \langle E \rangle$$

Cumulants: a closer look

$$Z = \text{tr} e^{-\hat{E}/T + \mu/T \hat{N}_B}$$

$$K_n = \frac{\partial^n}{\partial(\mu/T)^n} \ln Z = \frac{\partial^{n-1}}{\partial(\mu/T)^{n-1}} \langle N \rangle \quad \text{Cumulants are extensive: } K_n \sim V$$

$$K_2 = \langle N - \langle N \rangle \rangle^2 = \int d^3x d^3y \langle \delta\rho(x) \delta\rho(y) \rangle; \quad \delta\rho(x) = \rho(x) - \bar{\rho}$$

Susceptibility:

$$\chi_{(2) i,j} = \frac{1}{VT^3} \int d^3x d^3y \langle \delta\rho_i(x) \delta\rho_j(y) \rangle = \frac{1}{T^3} \bar{\rho}_i \delta_{i,j} + \frac{1}{T^3} \int d^3r C_{i,j}(r)$$

Correlation function (in configuration space!):

$$C_{i,j}(\vec{r}) = \langle \delta\rho_i(\vec{r}) \delta\rho_j(0) \rangle - \bar{\rho}_i \delta_{i,j} \delta(\vec{r}) \sim \frac{\exp[-r/\xi_{i,j}]}{r}$$

Correlation length (in configuration space!): $\xi_{i,j}$

Relation to cumulant: $K_2 = VT^3 \chi_{(2) i,i}$

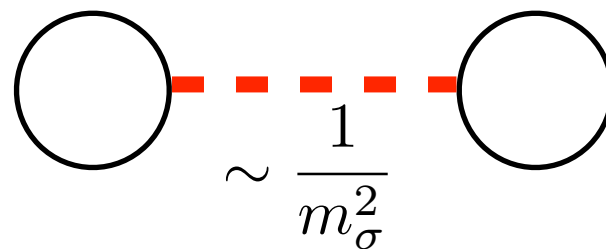
Correlation length

$$C(r) \sim \frac{\exp[-r/\xi]}{r}$$

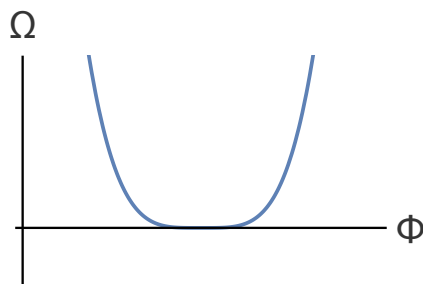
Static correlation function;
“Yukawa” potential with mass: $m \sim \frac{1}{\xi}$

$$\chi \sim \int C(r) d^3r \sim \xi^2 \sim \frac{1}{m^2}$$

simple “sigma” exchange

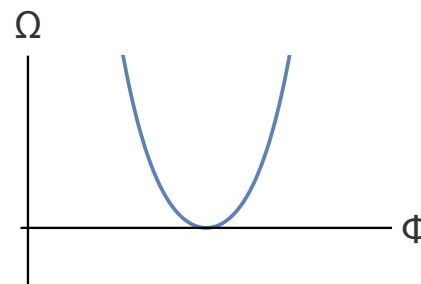


Critical point (second order)



$$m_\sigma \rightarrow 0, \quad \xi \rightarrow \infty$$

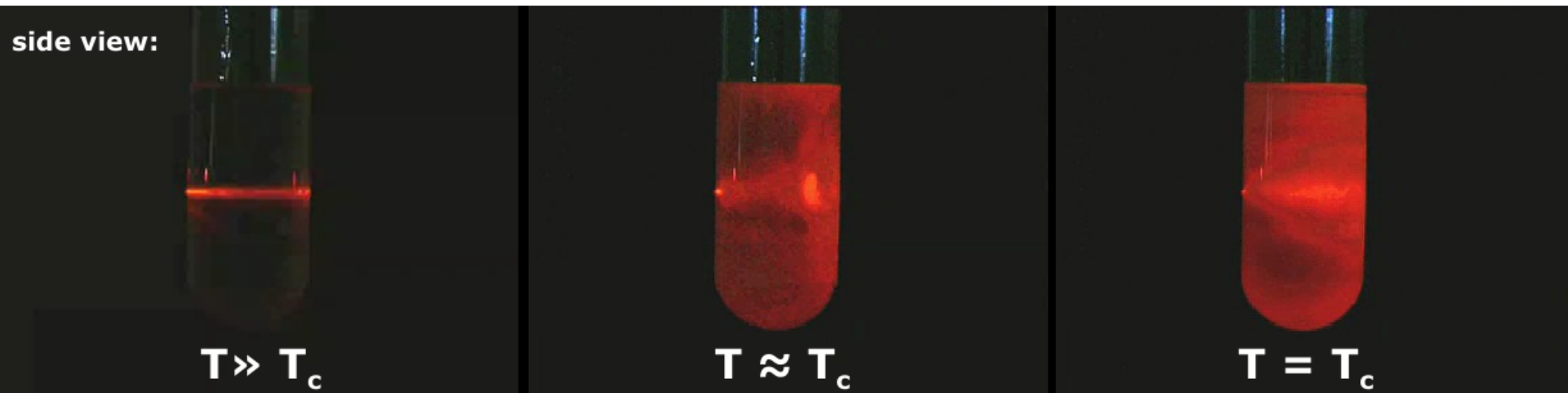
Cross over



$$m_\sigma, \xi \text{ finite}$$

Critical point

- Second order phase transition
- Fluctuations at all length scales
 - Critical opalescence



Higher moments (cumulants) and ξ

- Consider probability distribution for the order-parameter field:

$$P[\sigma] \sim \exp \{ -\Omega[\sigma]/T \} ,$$

$$\Omega = \int d^3x \left[\frac{1}{2} (\nabla \sigma)^2 + \frac{m_\sigma^2}{2} \sigma^2 + \frac{\lambda_3}{3} \sigma^3 + \frac{\lambda_4}{4} \sigma^4 + \dots \right] . \quad \Rightarrow \quad \xi = m_\sigma^{-1}$$

- Moments (connected) of $q = 0$ mode $\sigma_V \equiv \int d^3x \sigma(x)$:

First approximation:
count σ propagators

$$\kappa_2 = \langle \sigma_V^2 \rangle = VT \xi^2 ; \quad \kappa_3 = \langle \sigma_V^3 \rangle = 2VT^2 \lambda_3 \xi^6 ;$$

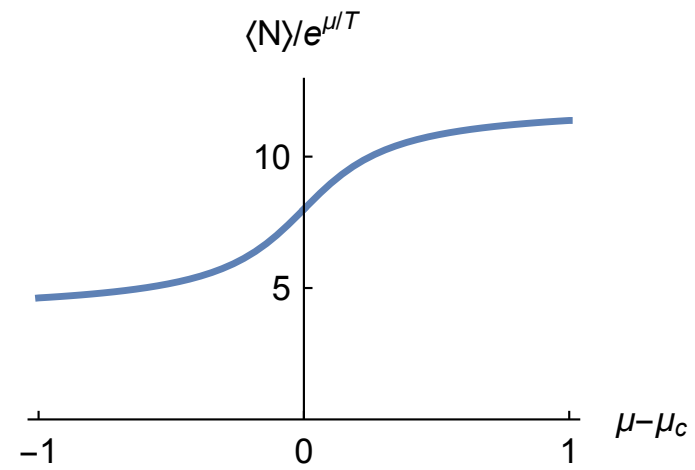
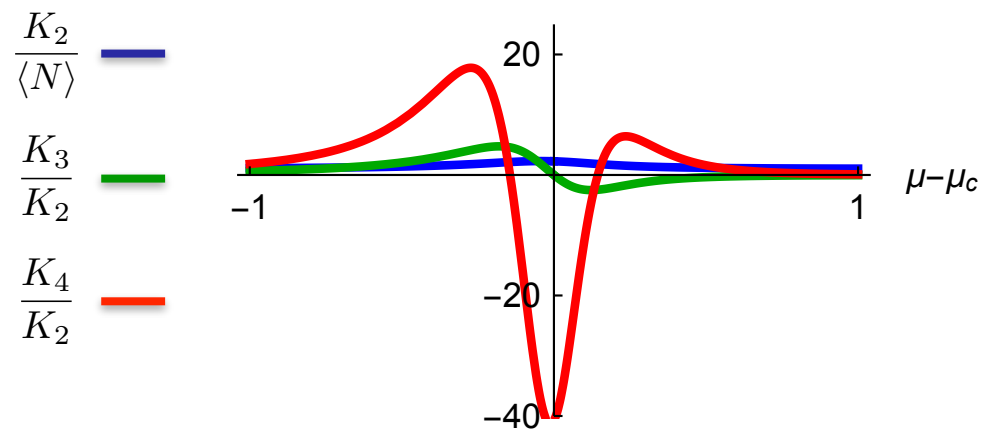
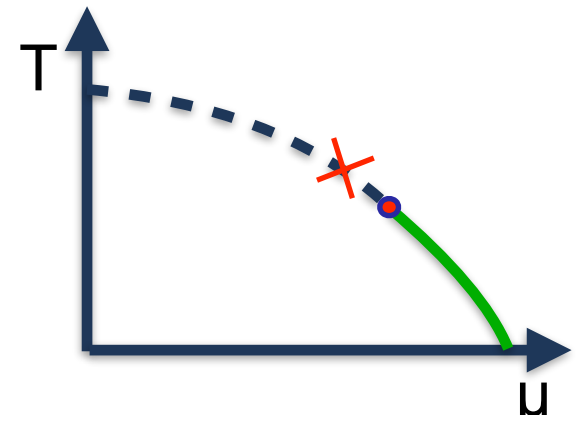
$$\kappa_4 = \langle \sigma_V^4 \rangle_c \equiv \langle \sigma_V^4 \rangle - 3 \langle \sigma_V^2 \rangle^2 = 6VT^3 [2(\lambda_3 \xi)^2 - \lambda_4] \xi^8 .$$

- Tree graphs. Each propagator gives ξ^2 .



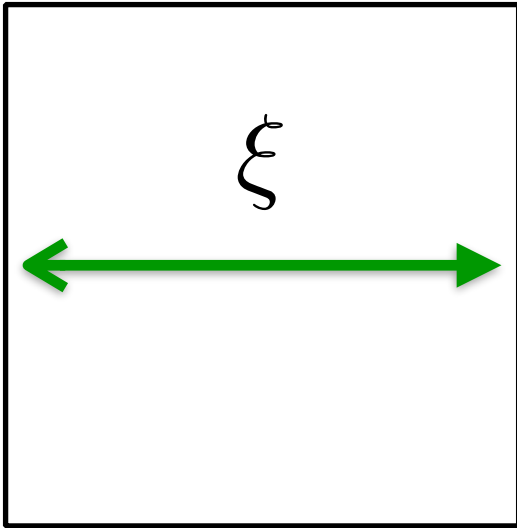
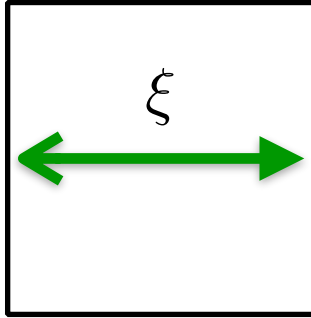
- Scaling requires “running”: $\lambda_3 = \tilde{\lambda}_3 T (T\xi)^{-3/2}$ and $\lambda_4 = \tilde{\lambda}_4 (T\xi)^{-1}$, i.e.,

$$\kappa_3 = \langle \sigma_V^3 \rangle = 2VT^{3/2} \tilde{\lambda}_3 \xi^{4.5} ; \quad \kappa_4 = 6VT^2 [2(\tilde{\lambda}_3)^2 - \tilde{\lambda}_4] \xi^7 .$$



Finite size scaling

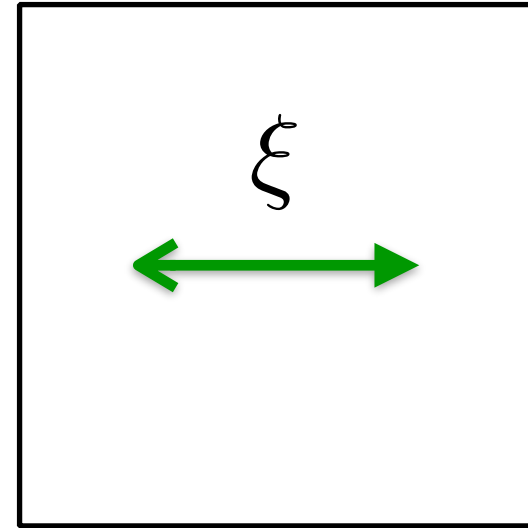
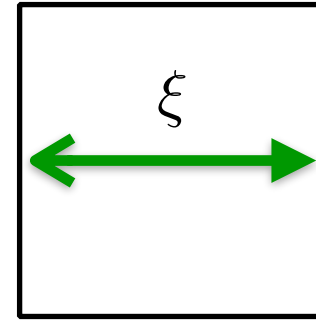
Second order (critical point)



$$\xi \sim V^{2/3}, \quad \chi \sim V^{4/3}$$

(mean field)

Cross over

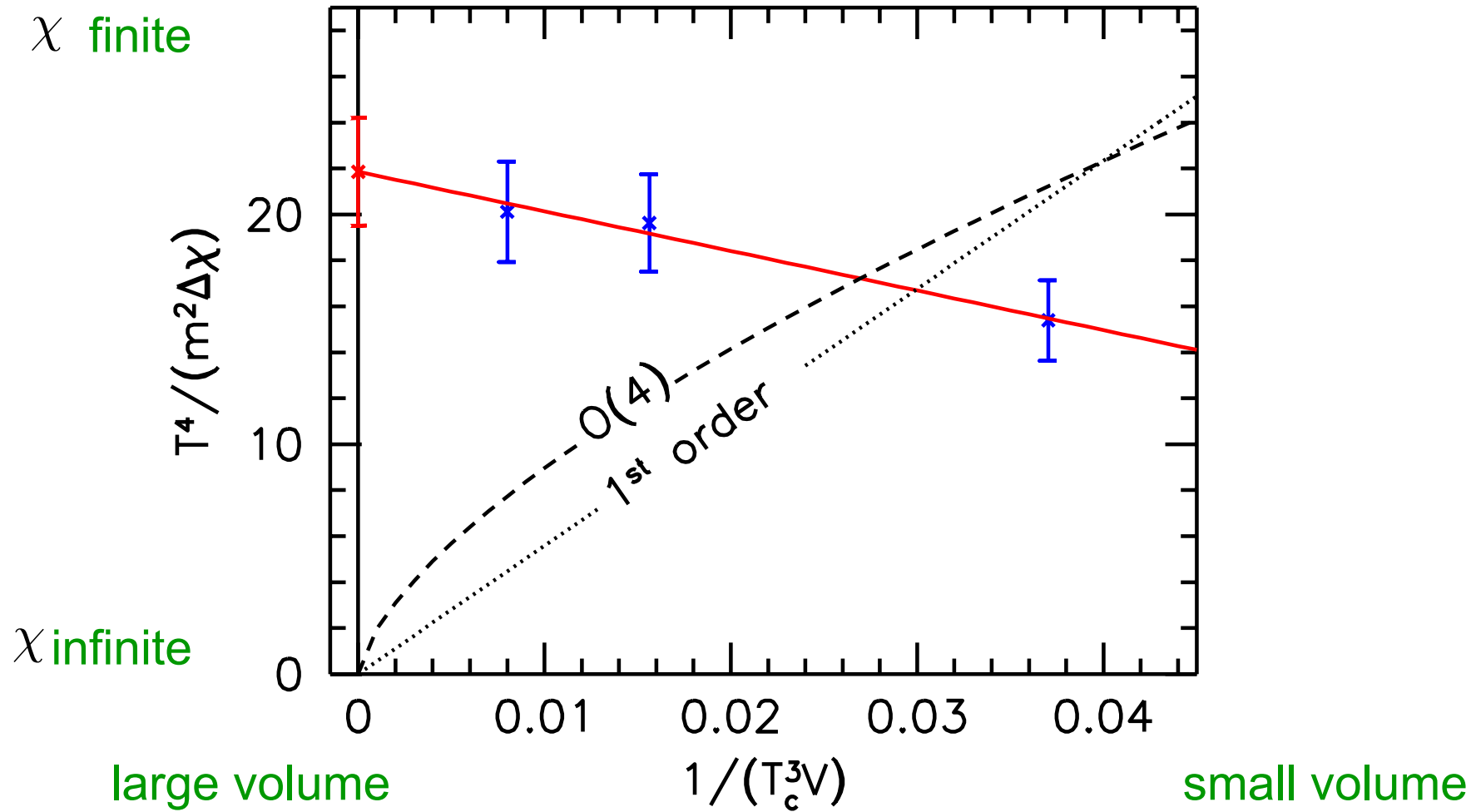


$$\xi = \text{const}, \quad \chi = \text{const}$$

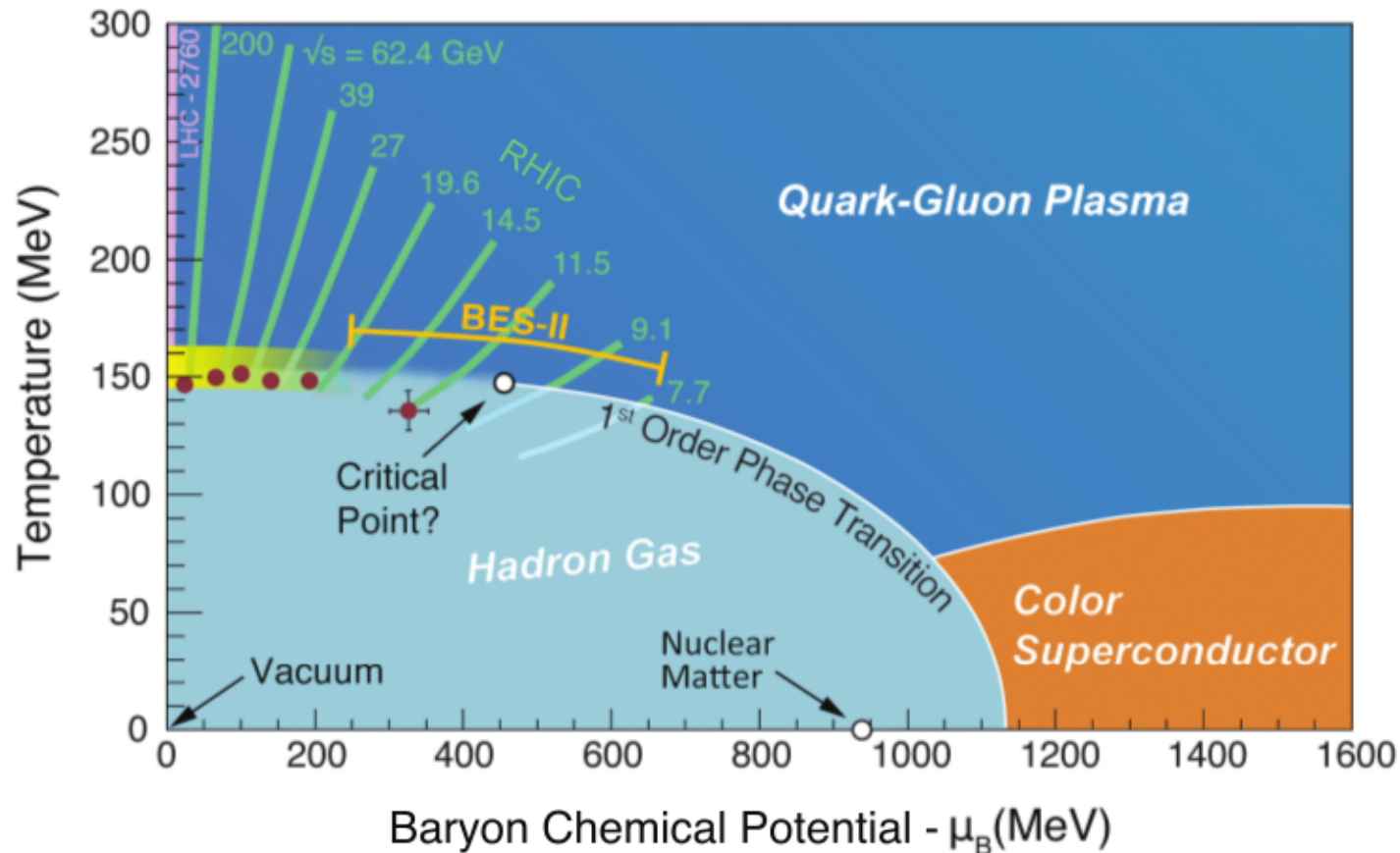
NB: 1st order: $\chi \sim V$

QCD at $\mu=0$ is cross-over

Aoki et al, Nature 43:675-678,2006



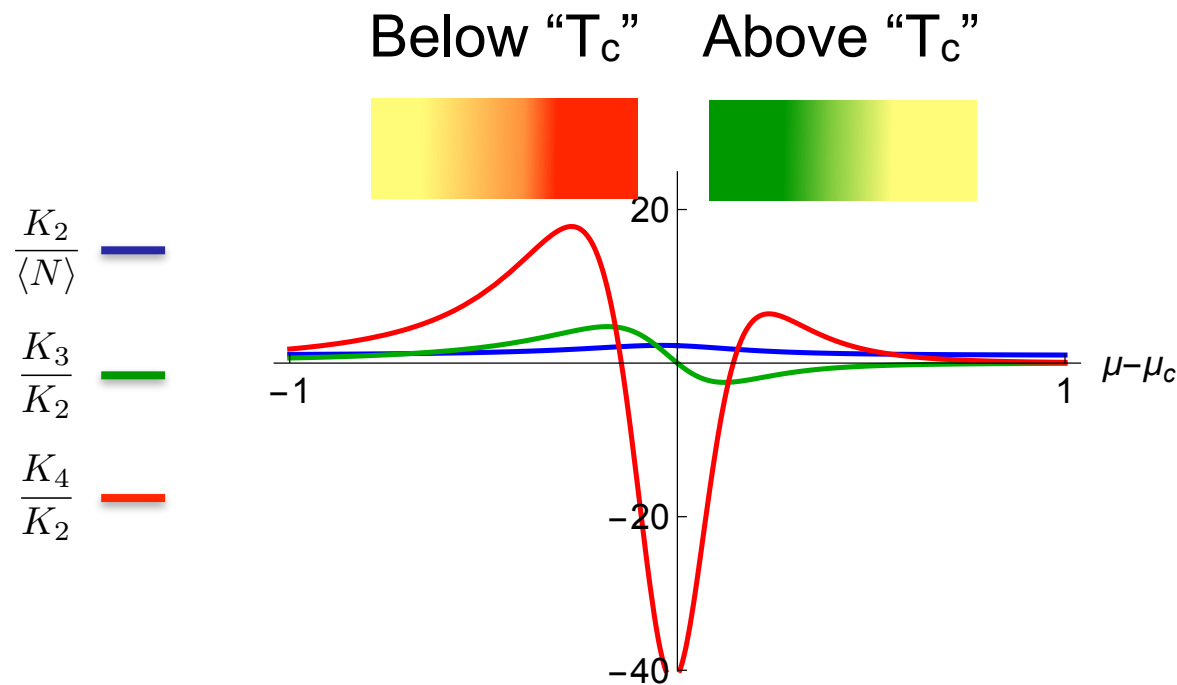
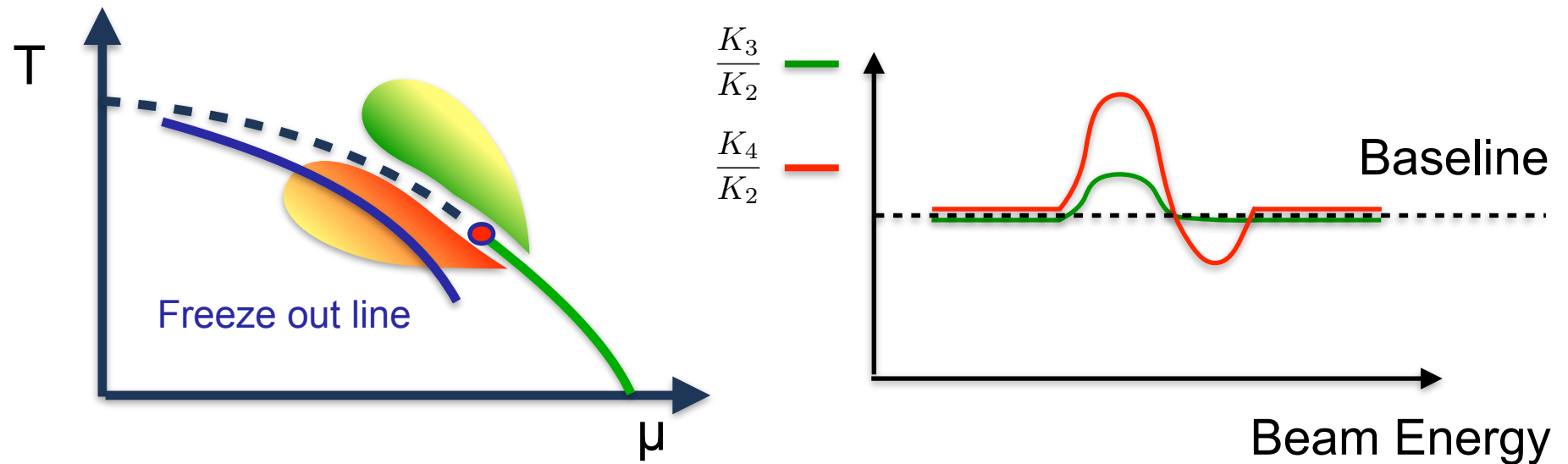
The phase diagram



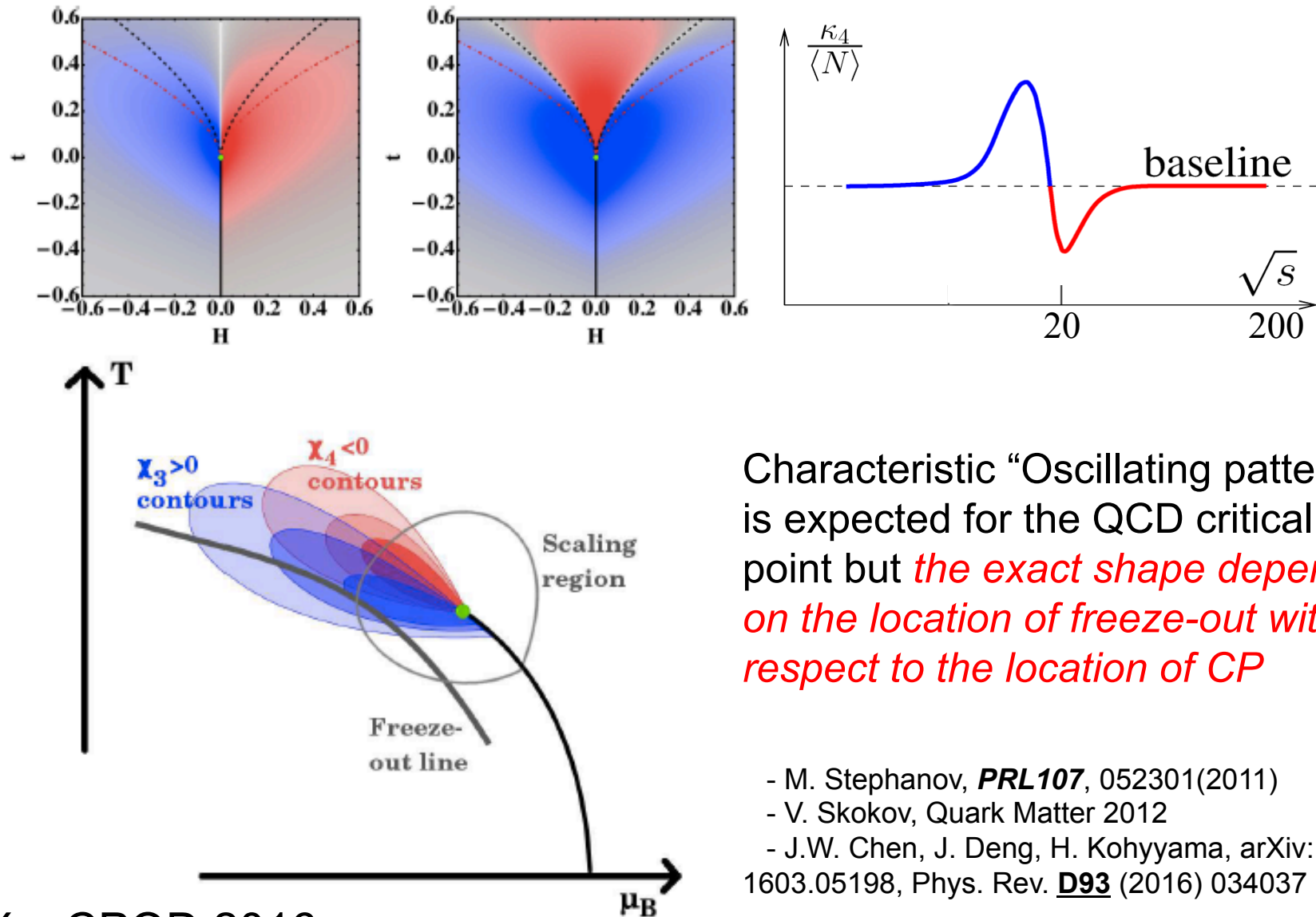
Increase chemical potential by lowering the beam energy

In reality, we add baryons (nucleons) from target and projectile to mid-rapidity

What to expect from experiment?



Expectation from Calculations

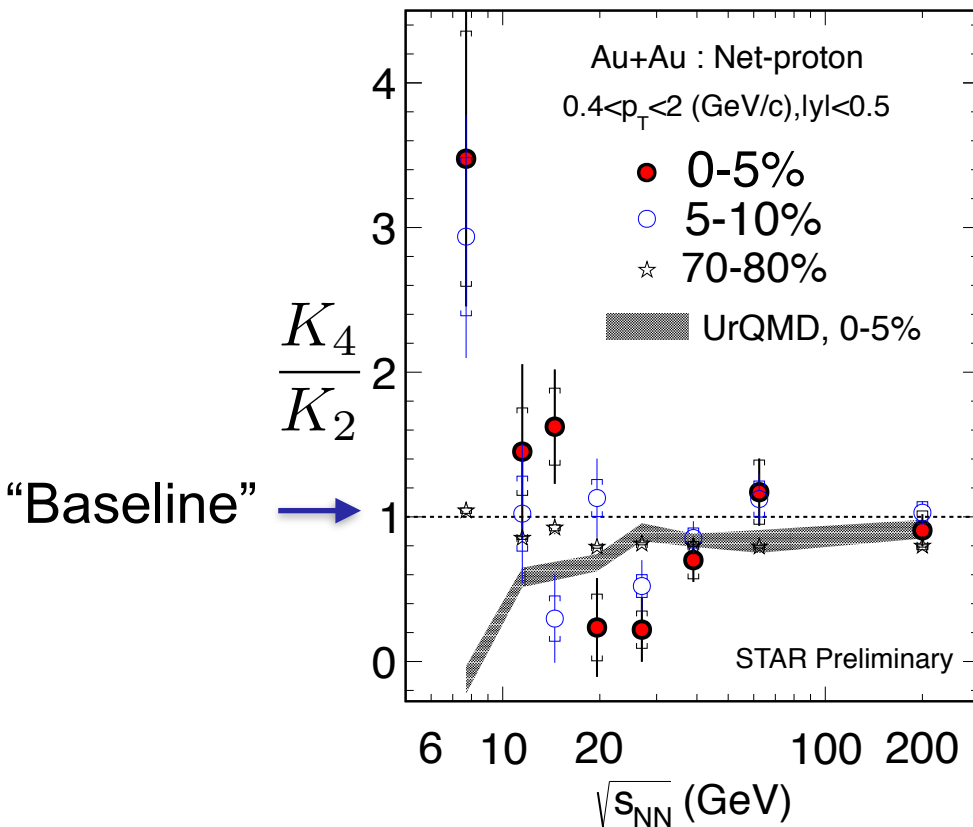
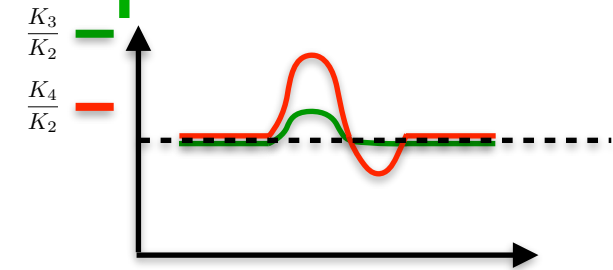


Characteristic “Oscillating pattern” is expected for the QCD critical point but *the exact shape depends on the location of freeze-out with respect to the location of CP*

- M. Stephanov, **PRL****107**, 052301(2011)
- V. Skokov, Quark Matter 2012
- J.W. Chen, J. Deng, H. Kohyama, arXiv: 1603.05198, Phys. Rev. **D93** (2016) 034037

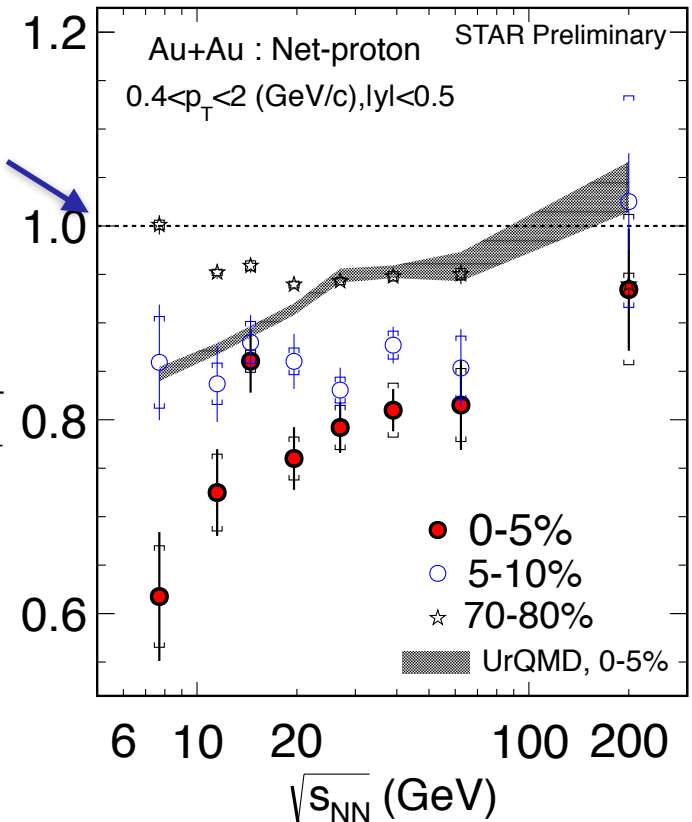
Latest STAR result on net-proton cumulants

X. Luo, NPA 956 (2016) 75



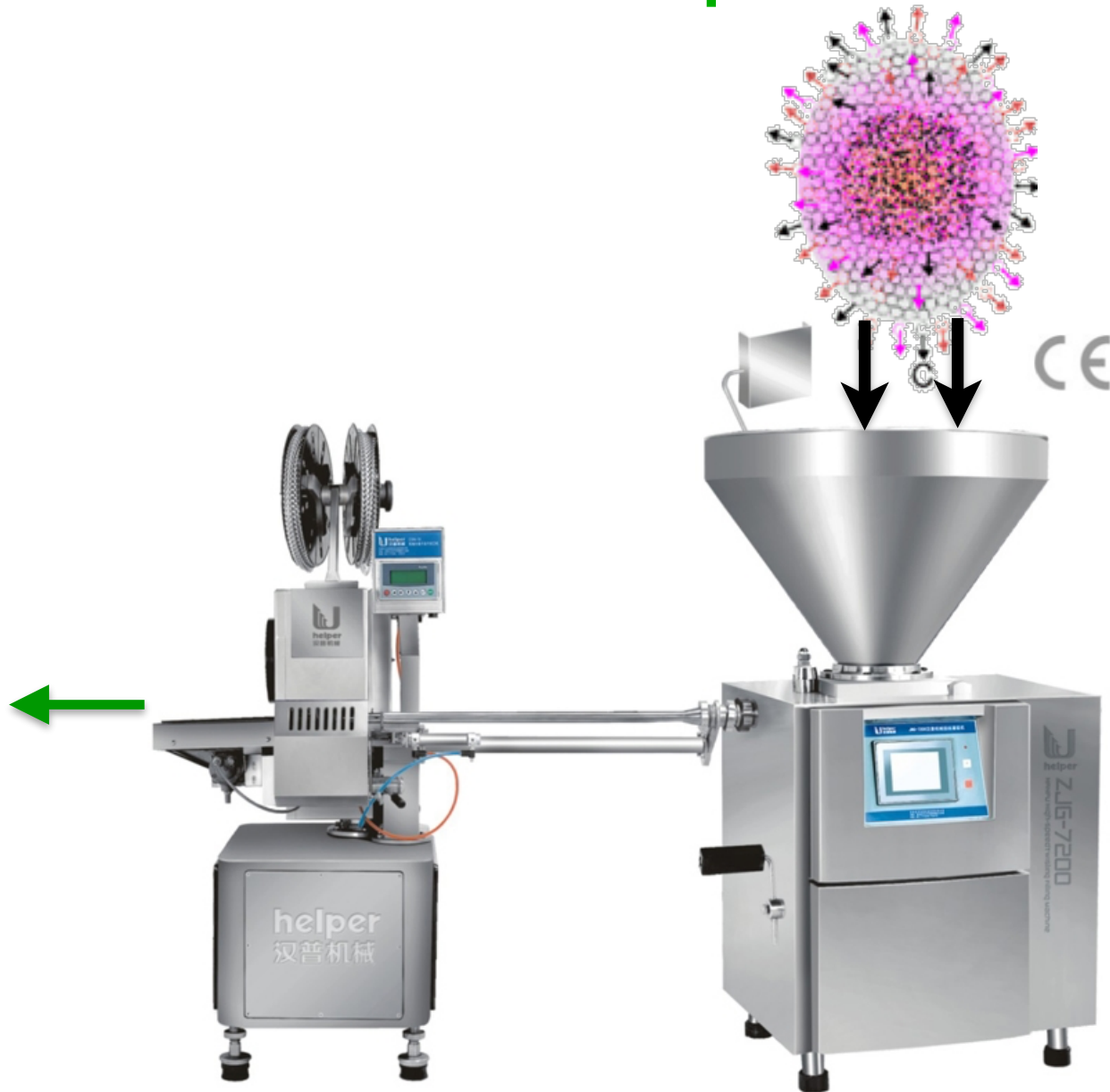
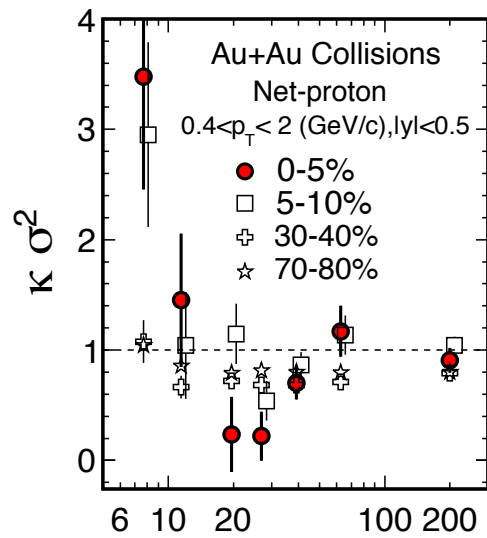
“Baseline” →

$$\frac{K_3/K_2}{\text{Skellam}}$$

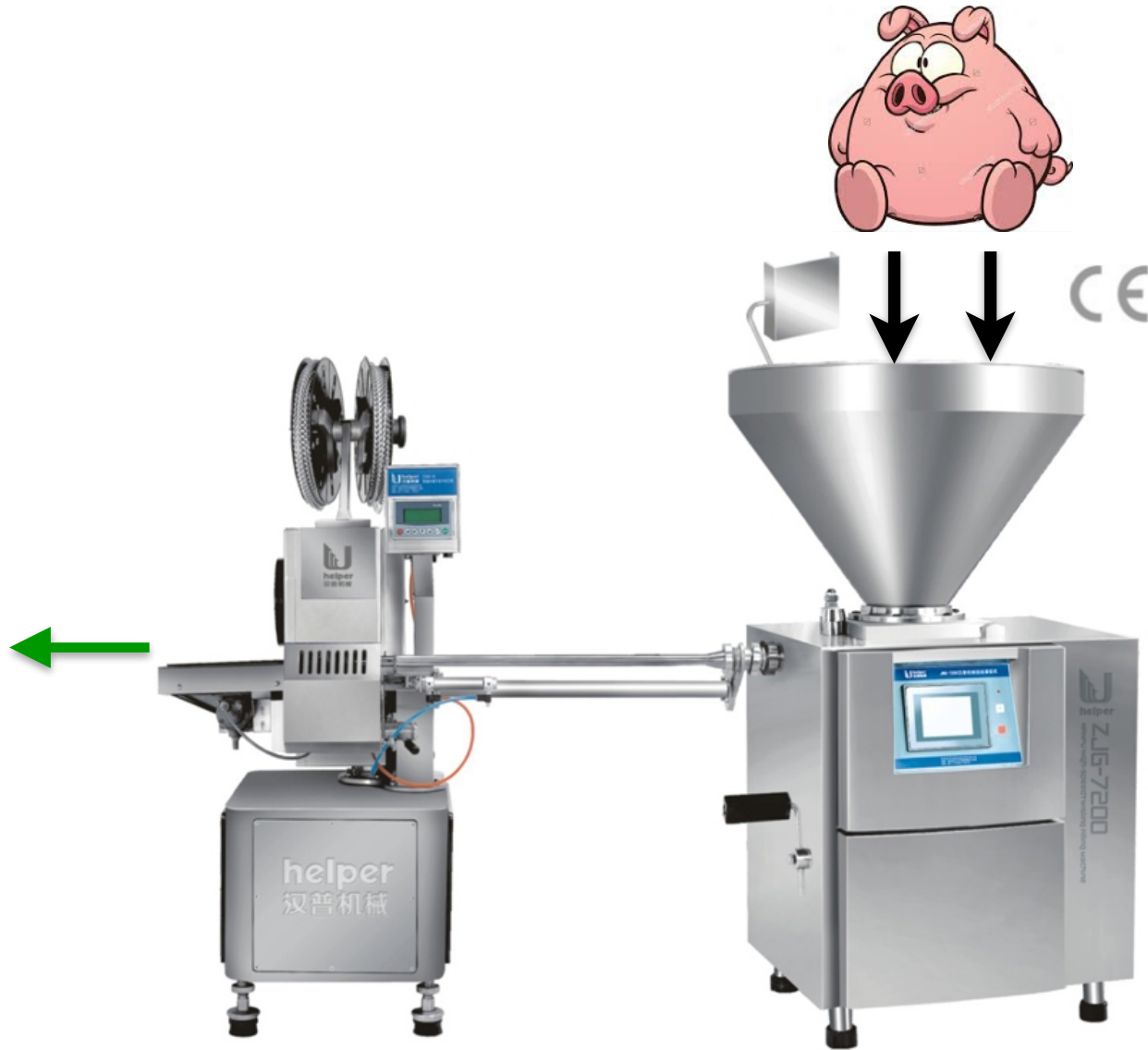


K_4/K_2 follows expectation, K_3/K_2 no so much.....
 URQMD totally fails to get trend for K_4/K_2 !

The measurement process



Or in the real world.....



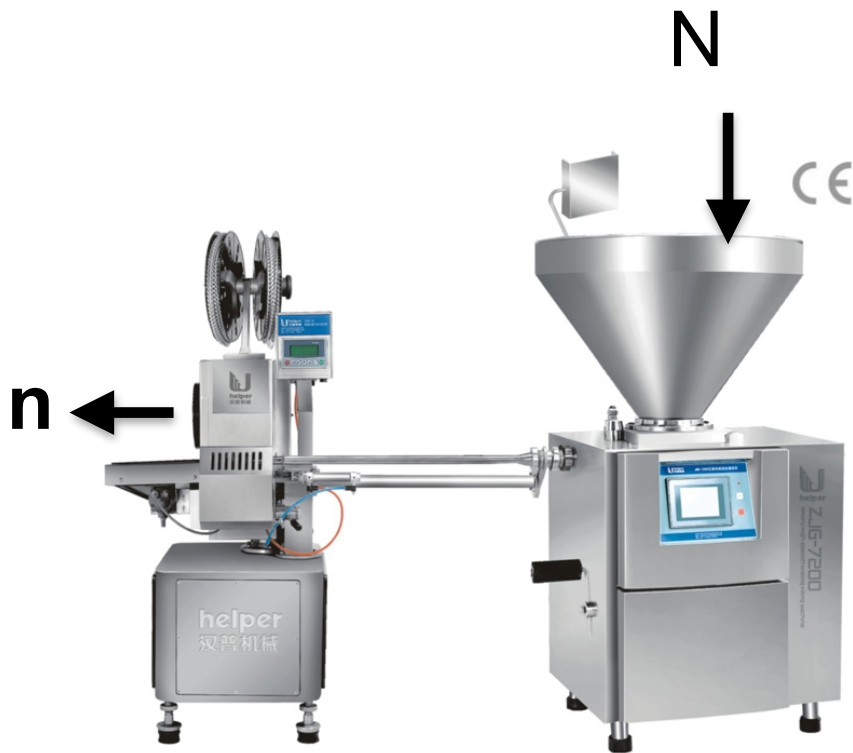
Modeling the detector (multiplicities only)

Detector maps TRUE number of particles
onto OBSERVED number of particles

$$p(n) = \sum_N B(n, N, \epsilon, \dots) P(N)$$

Observed

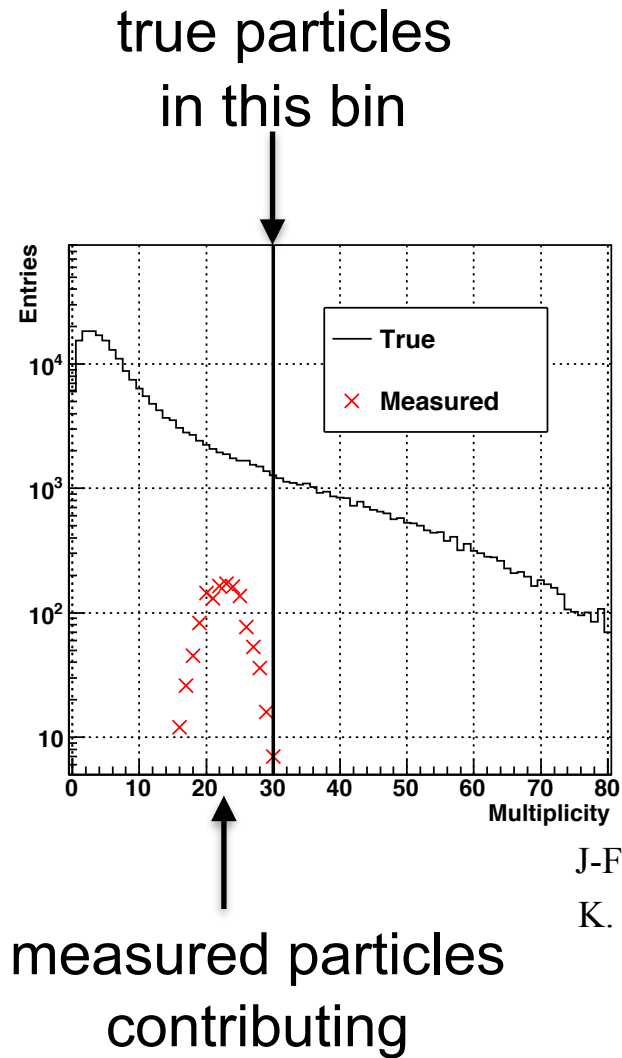
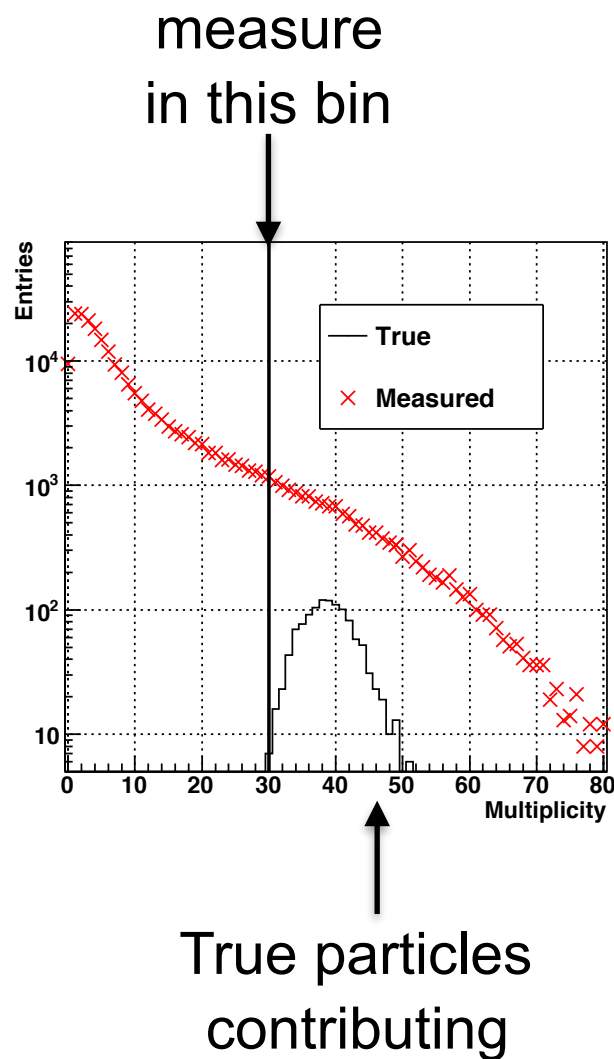
True



$$B(n, N, \epsilon, \dots)$$

B is matrix which
controls the mapping

$$p_n = B_{n,N} P_N$$



J-F. Grosse-Oetringhaus,
K. Reygers arXiv:0912.0023v2

Unfolding

$$\underbrace{p(n)}_{\text{Observed}} = \sum_N B(n, N, \epsilon, \dots) \underbrace{P(N)}_{\text{True}}$$

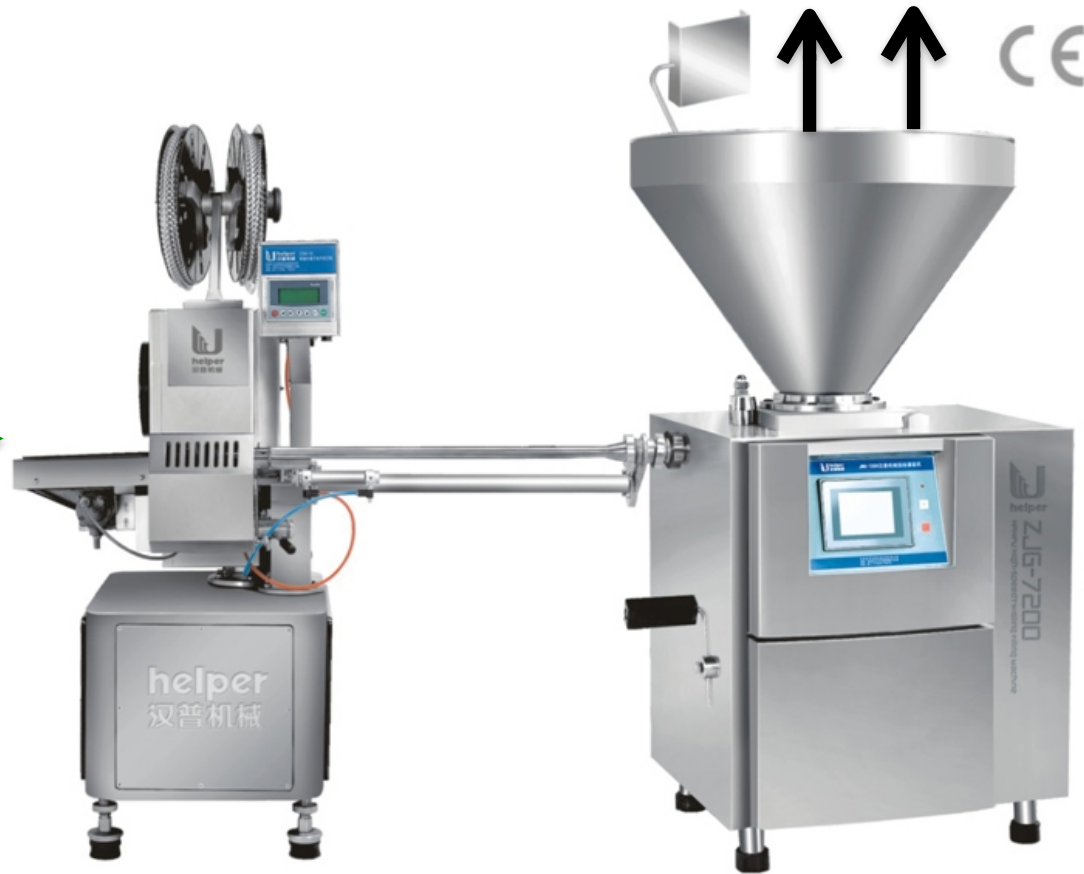
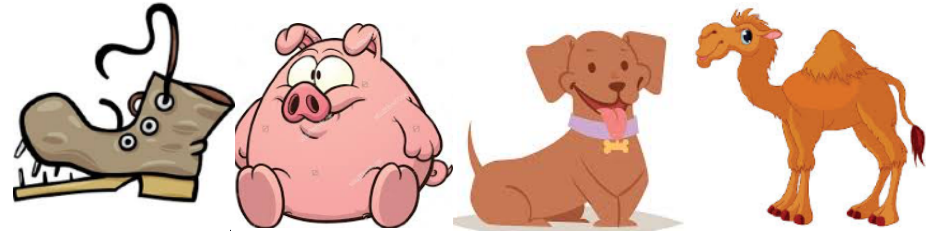
To get TRUE $P(N)$ we need to invert matrix B so that

$$\underbrace{P(N)}_{\text{True}} = \sum_n B^{-1}(N, n, \epsilon, \dots) \underbrace{p(n)}_{\text{Observed}}$$

This is called **UNFOLDING**

In practice simple inverting does not work!

Or in the real world...



Example: Binomial

$$\mathbf{p}_n = \mathbf{B}_{n,N} \mathbf{P}_N$$
$$\begin{pmatrix} p(0) \\ p(1) \\ p(2) \\ p(3) \\ p(4) \end{pmatrix} = \begin{pmatrix} 1 & 1-\epsilon & (1-\epsilon)^2 & (1-\epsilon)^3 & (1-\epsilon)^4 \\ 0 & \epsilon & 2\epsilon(1-\epsilon) & 3\epsilon(1-\epsilon)^2 & 4\epsilon(1-\epsilon)^3 \\ 0 & 0 & \epsilon^2 & 3\epsilon^2(1-\epsilon) & 6\epsilon^2(1-\epsilon)^2 \\ 0 & 0 & 0 & \epsilon^3 & 4\epsilon^3(1-\epsilon) \\ 0 & 0 & 0 & 0 & \epsilon^4 \end{pmatrix} \begin{pmatrix} P(0) \\ P(1) \\ P(2) \\ P(3) \\ P(4) \end{pmatrix}$$

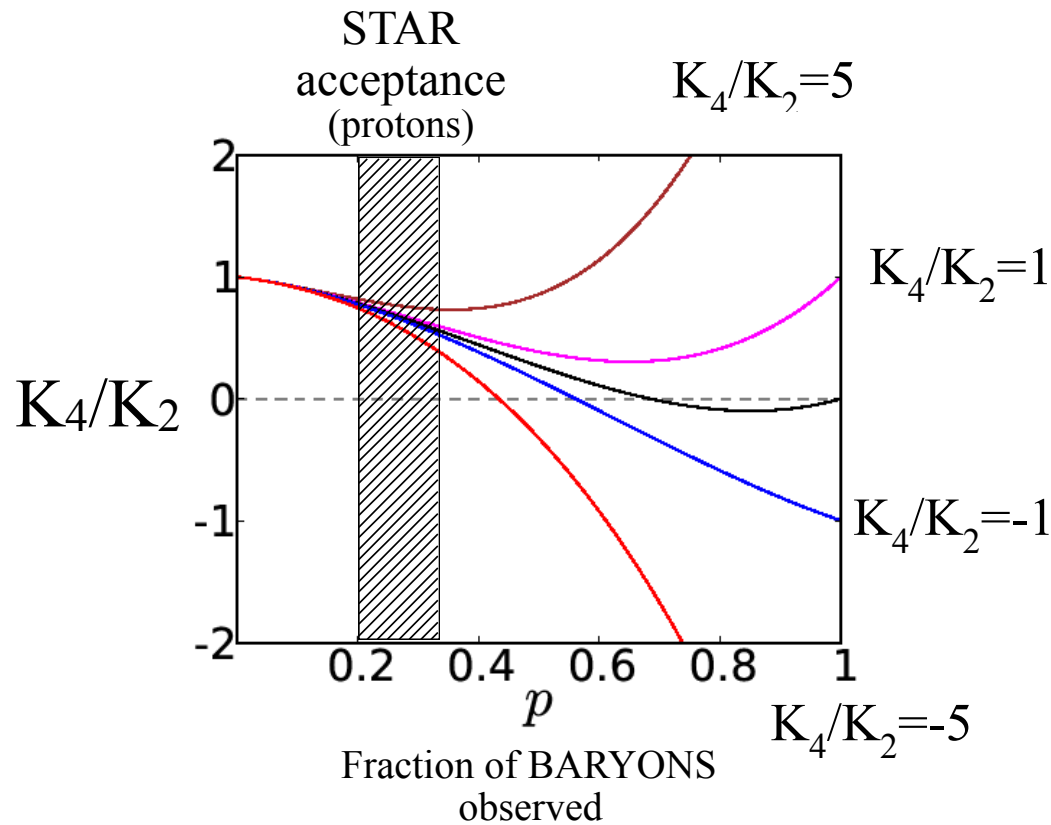
Binomial probability $\epsilon < 1$ is often called “efficiency”

Theoretically: $n_{Obs.} \leq N_{True} \Rightarrow$ B is triangular

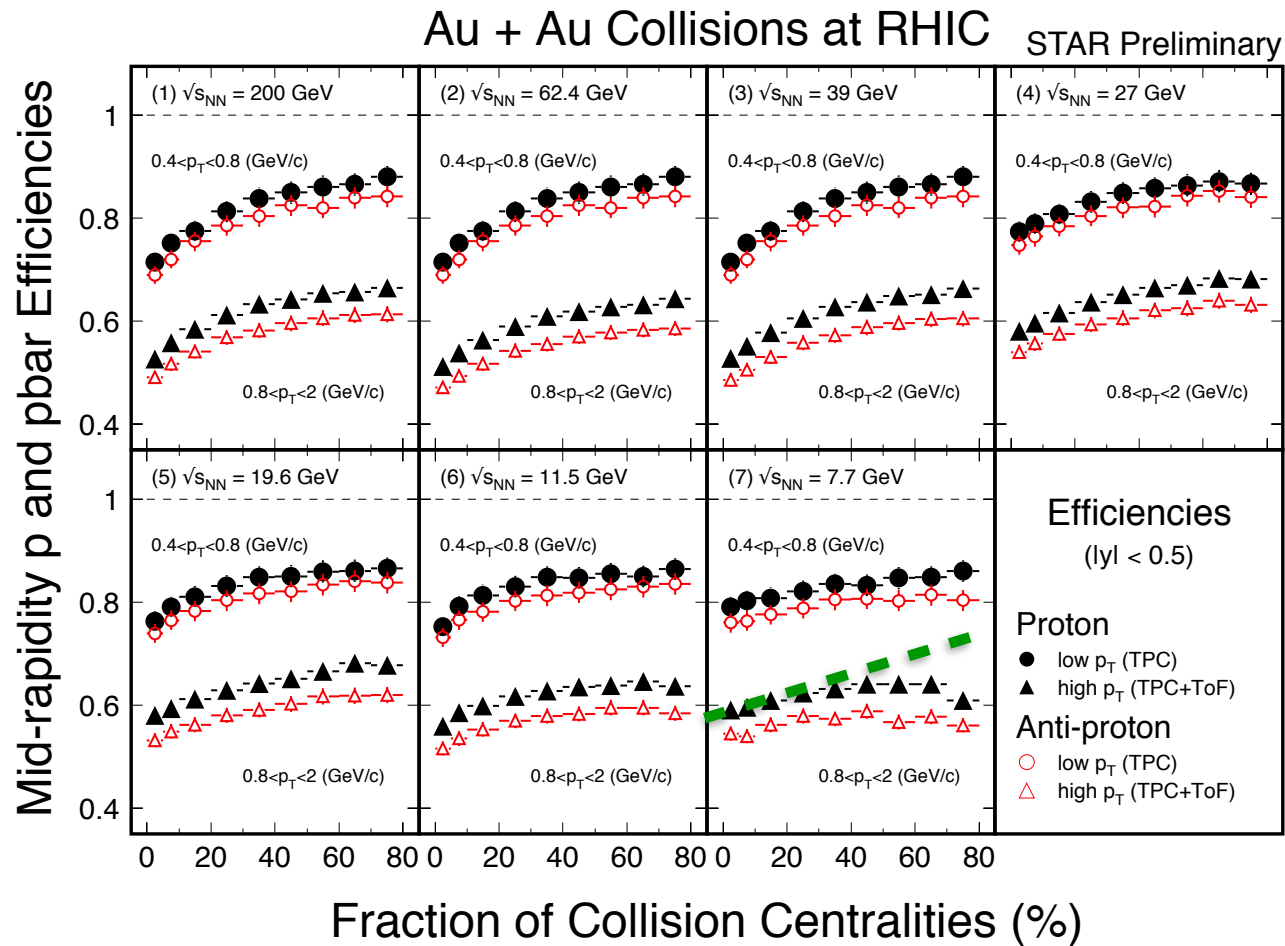
$\mathbf{B}_{n,N}$ almost singular ! STAR: $0.6 < \epsilon < 0.8$

In Practice: Who knows... is the detector even “binomial”

Binomial allows to invert (at least for cumulants)



Is $B(n,N)$ binomial ?



Efficiency depends on multiplicity!

Binomial distributions and real detectors

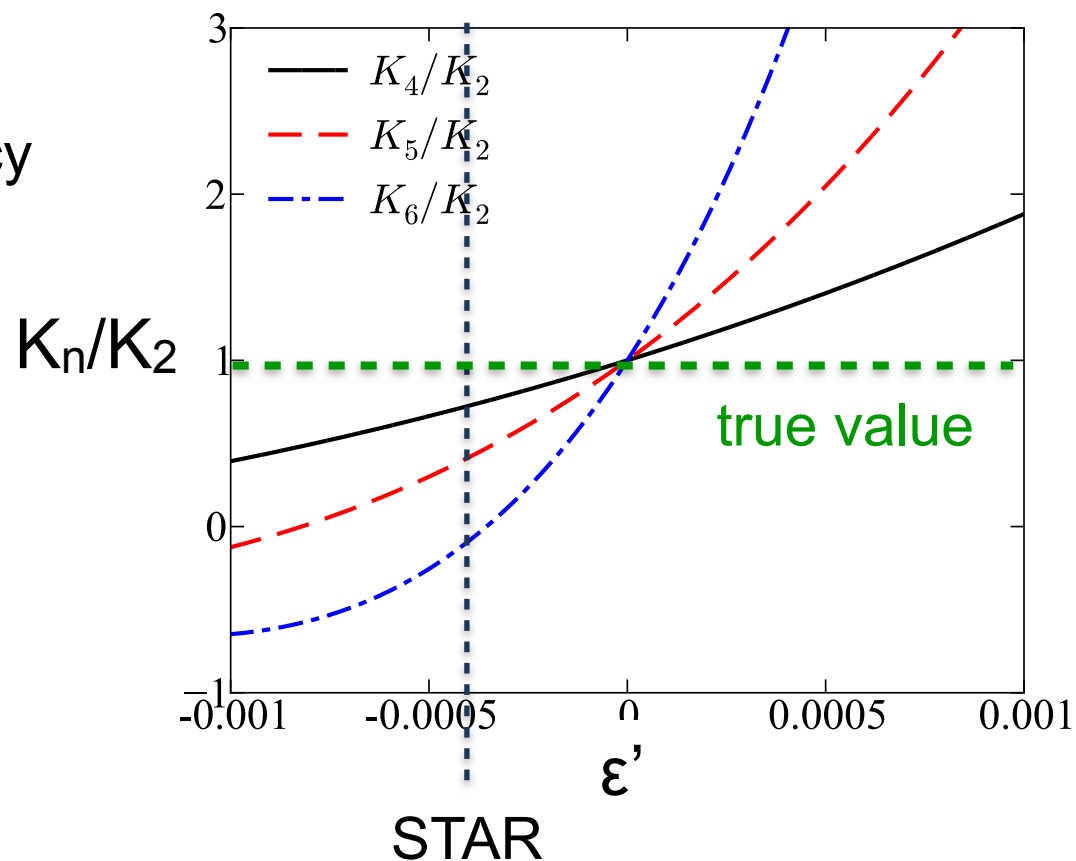
The most obvious correction:
Multiplicity dependence of efficiency

$$\epsilon(N) = \epsilon_0 + \epsilon'(N - \langle N \rangle)$$

More details:

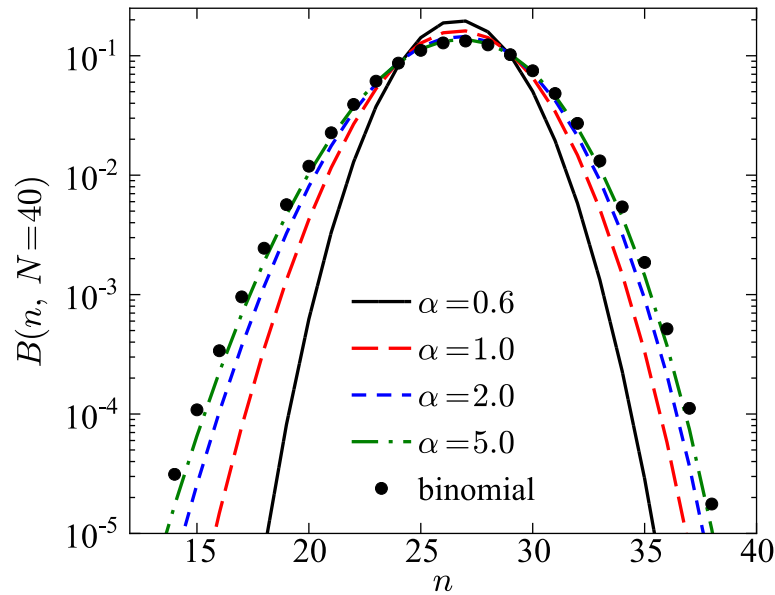
A. Bzdak, R. Holzmann et al.

arXiv:1603.09057



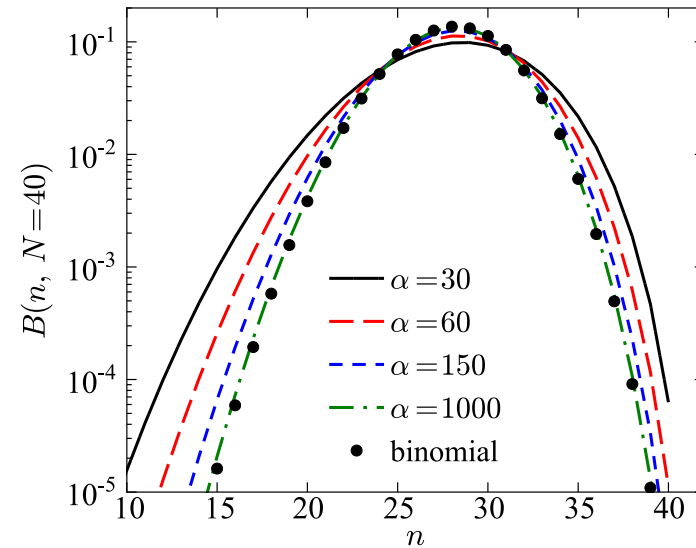
Other models for $B(n, N)$

Hypergeometric



Hypergeometric	$\alpha = 0.6$	$\alpha = 1.0$	$\alpha = 2.0$	$\alpha = 5.0$
K_3/K_2	1.16	1.12	1.07	1.03
K_4/K_2	0.66	0.88	0.98	1.00
K_5/K_2	2.19	1.68	1.23	1.05
K_6/K_2	-3.99	-1.38	0.31	0.89

Beta Binomial



Beta-binomial	$\alpha = 30$	$\alpha = 60$	$\alpha = 150$	$\alpha = 1000$
K_3/K_2	1.28	1.24	1.13	1.02
K_4/K_2	0.82	1.45	1.35	1.07
K_5/K_2	-1.11	1.15	1.63	1.16
K_6/K_2	5.71	-0.44	1.80	1.32

Insights from Theory



Compare Data with Lattice QCD and other field theoretical models

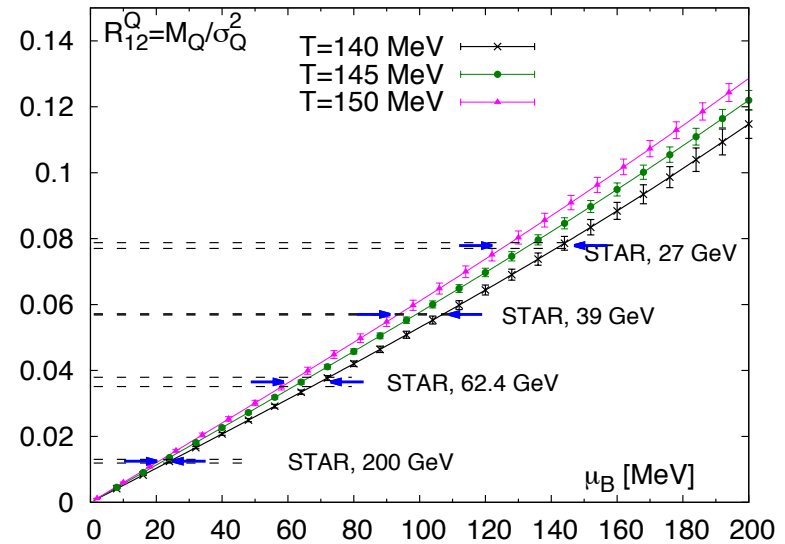
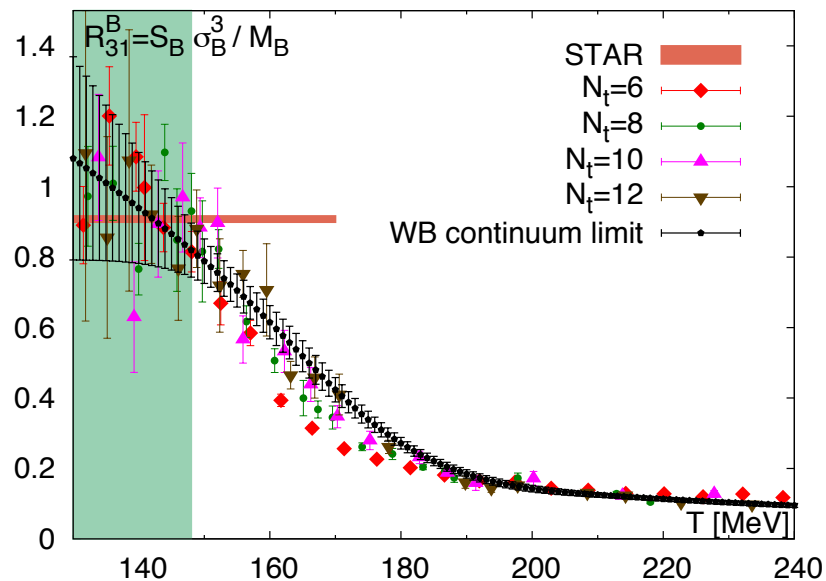


Compare Data with Lattice QCD and other field theoretical models

- Lattice cannot calculate hadron abundances
- Cumulants are well defined quantities
- Compare cumulants !?
 - Baryon number conservation
 - Experiment measures protons not all baryons
 - Volume is not fixed in experiment
 - Experiment has finite momentum space coverage (usually)

Compare Data with Lattice QCD

For example: Wuppertal-Budapest (arXiv:1305.5161)
(similar from Hot QCD)

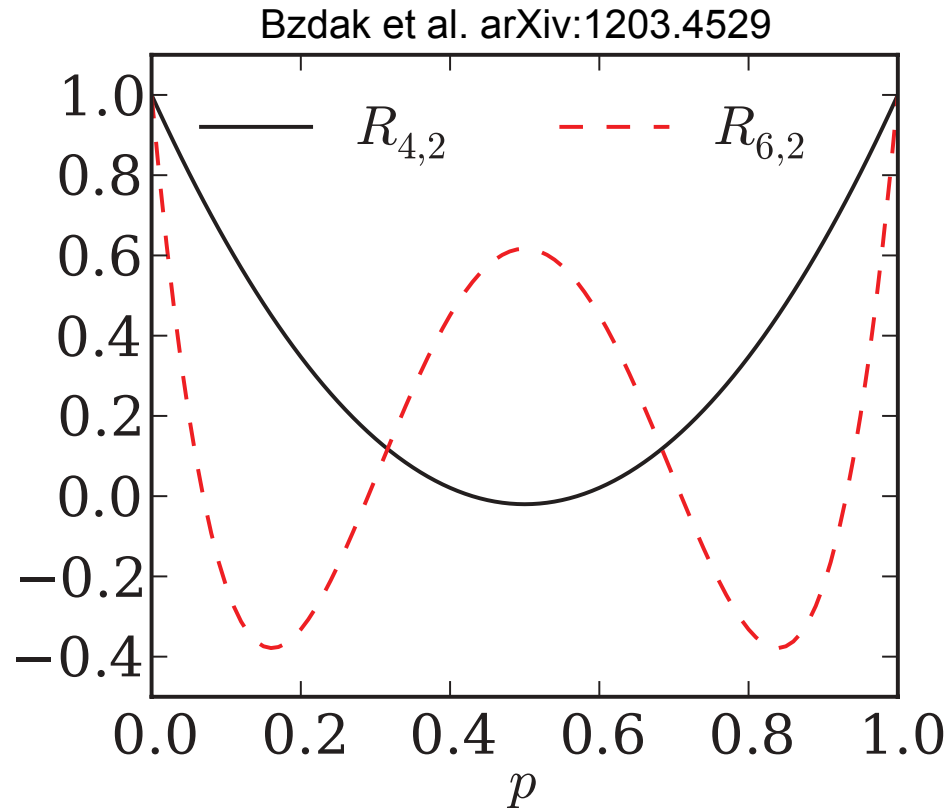


Baryon number conservation

Lattice works in grand-canonical ensemble:
Baryon number conserved only on average

Experiment: Baryon number is conserved event-by-event

No physics
other than
baryon number
conservation



Fraction of total baryons in detector

$$R_{4,2} = \frac{K_4}{K_2}$$
$$R_{6,2} = \frac{K_6}{K_2}$$

Protons vs Baryons

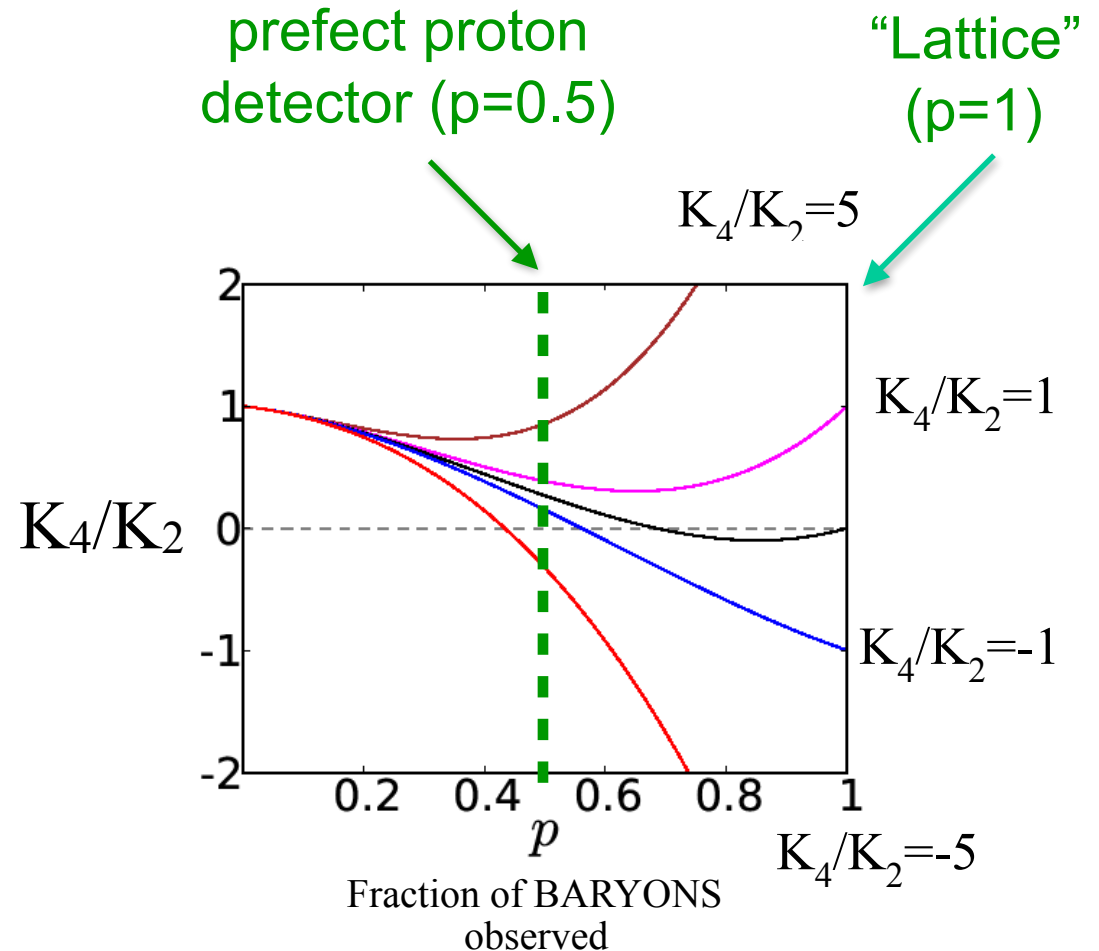
Fast isospin exchange
a.k.a lots of pions:

protons and neutrons follow
binomial distribution

$$P(N_p) = \frac{B!}{N_p!(B - N_p)!} p^{N_p} (1 - p)^{B - N_p}$$

with $p \sim 0.5$

(Kitazawa, Asakawa arXiv:1107.2755)



Finite acceptance

Example: “Charge” susceptibility

$$\chi_Q = \int d^3x \langle \rho(x) \rho(0) \rangle = \int d^3p \langle \tilde{\rho}(p) \tilde{\rho}(0) \rangle$$

Equivalence of *Integrated* coordinate space and momentum space correlation function

Experiment almost never integrates ALL of momentum space!

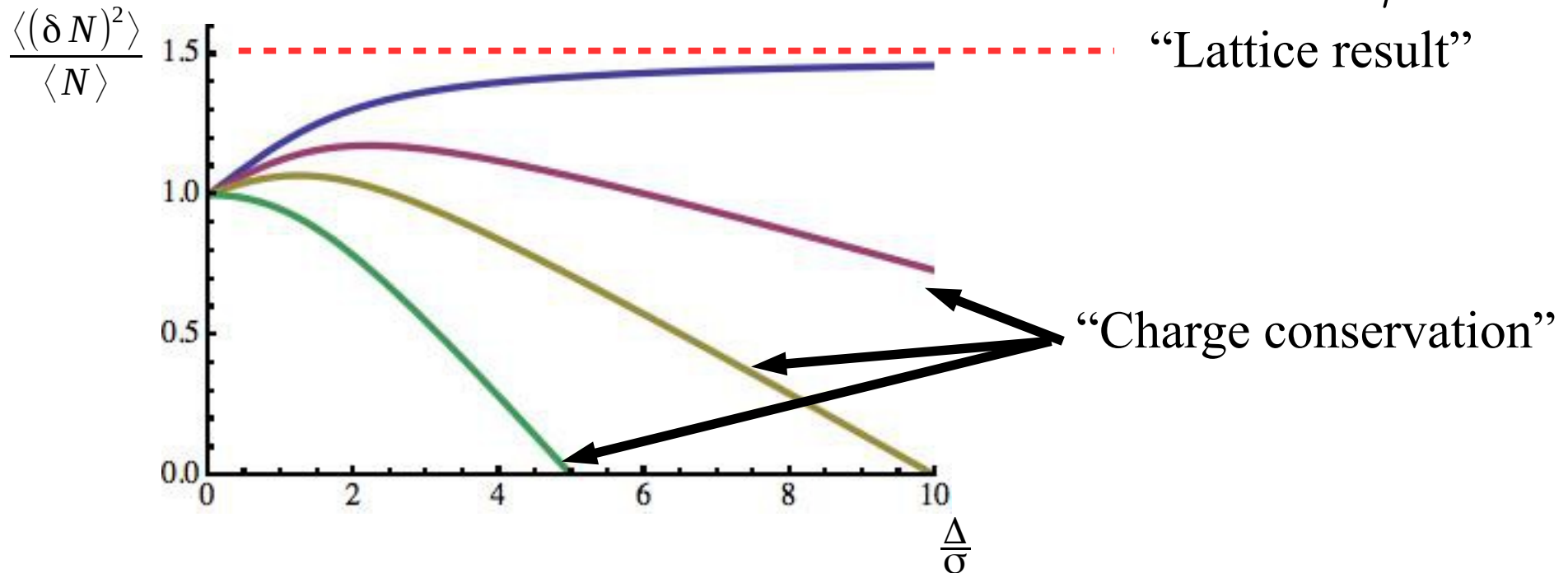
Lattice (hopefully) does integrate over all coordinate space

Correlations: Lattice vs Data

$$\langle n(y_1)(n(y_2) - \delta(y_1 - y_2)) \rangle = \langle n(y_1) \rangle \langle n(y_2) \rangle (1 + C(y_1, y_2))$$

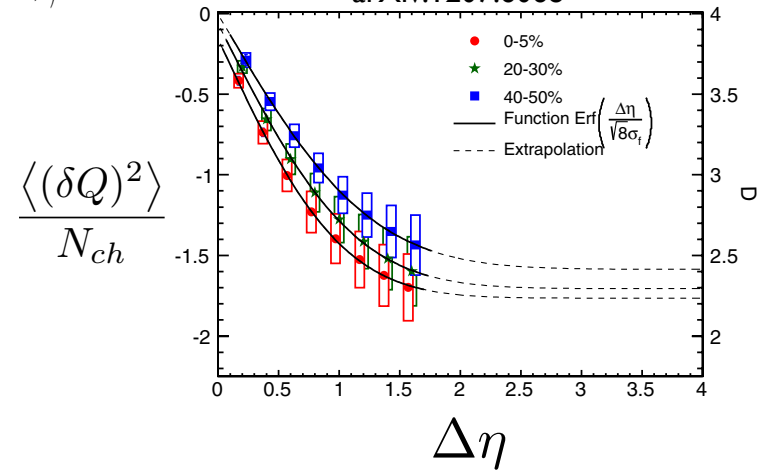
$$C(y_1, y_2) \sim \exp\left(-\frac{(y_1 - y_2)^2}{2\sigma^2}\right)$$

$$\frac{\langle (\delta N)^2 \rangle}{\langle N \rangle} = 1 + \langle N \rangle \int_{-\Delta/2}^{\Delta/2} C(y_1, y_2) dy_1 dy_2$$



Alice Charge Flucts

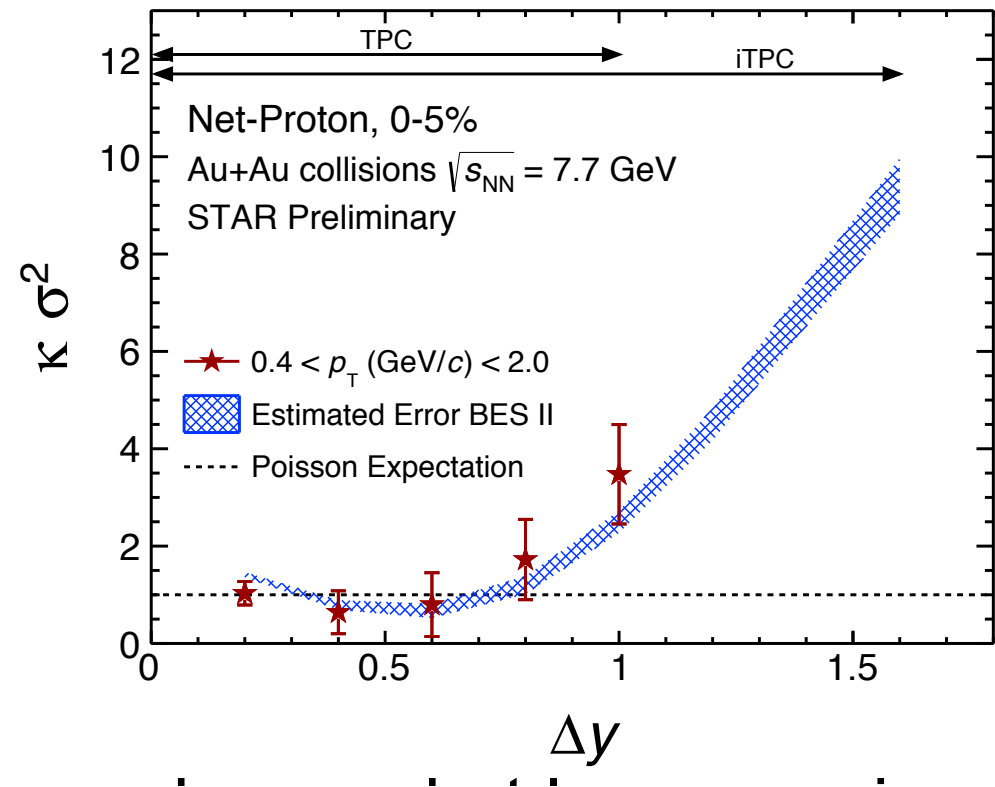
arXiv:1207.6068



Dependence on Rapidity window

X. Luo, EMMI Workshop, Nov. 2015

- Kurtosis depends strongly on Rapidity window
- Comparison with Lattice:
 - Lattice catches the full correlation length
 - need to expand rapidity window until signal saturates (after correcting for charge conservation)

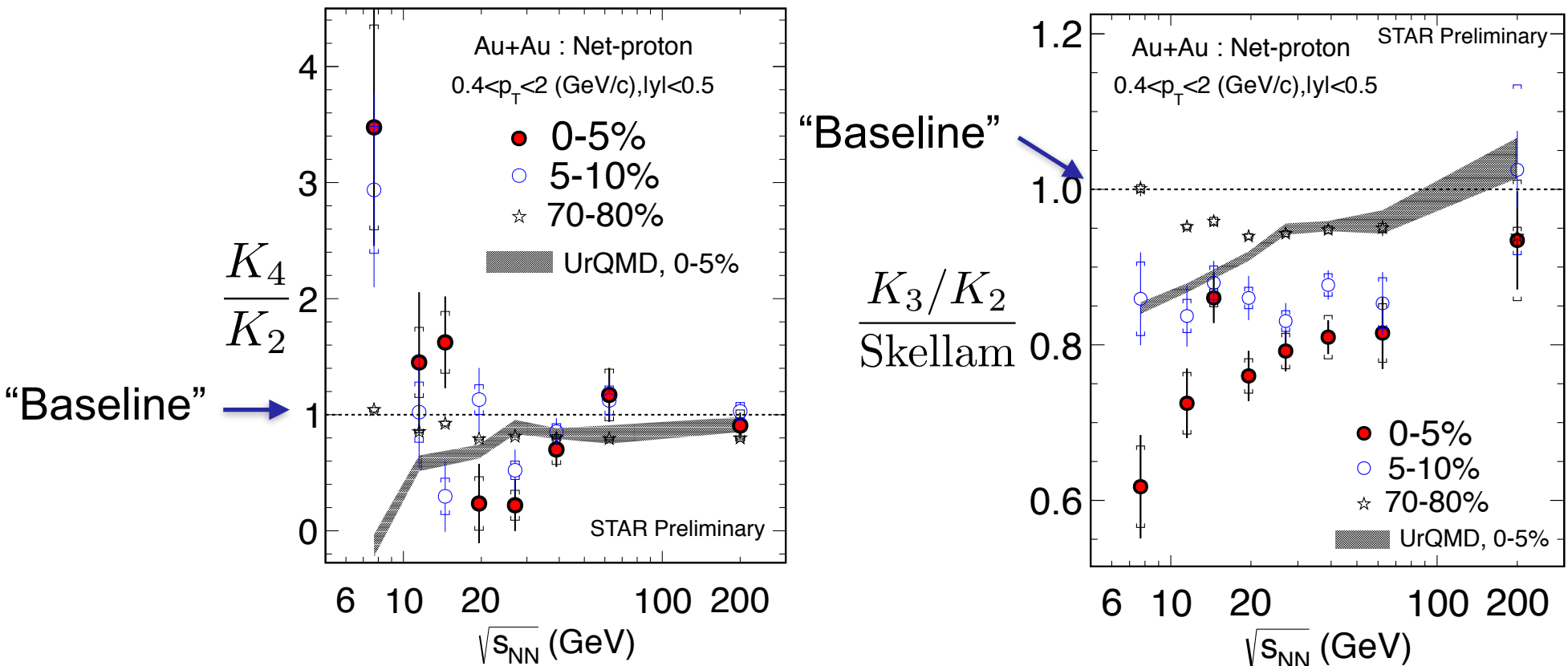


Any comparison of Lattice to Data needs to assure that cumulants reach asymptotic value in experiment.

So far this has NOT been established for proton cumulants

Back to data assuming that STAR has done their job

X. Luo, NPA 956 (2016) 75



K_4/K_2 follows expectation, K_3/K_2 no so much.....
URQMD totally fails to get trend for K_4/K_2 !

Further insights: Correlations

Cumulants $K_n = \frac{\partial^n}{\partial(\mu/T)^n} \ln Z$

$$K_2 = \langle N - \langle N \rangle \rangle^2 = \langle (\delta N)^2 \rangle$$

$$\rho_2(p_1, p_2) = \rho_1(p_1)\rho_1(p_2) + C_2(p_1, p_2), \quad \mathbf{C_2: Correlation Function}$$

$$K_3 = \langle (\delta N)^3 \rangle$$

$$\begin{aligned} \rho_3(p_1, p_2, p_3) = & \rho_1(p_1)\rho_1(p_2)\rho_1(p_3) + \rho_1(p_1)\underline{C_2(p_2, p_3)} + \rho_1(p_2)\underline{C_2(p_1, p_3)} \\ & + \rho_1(p_3)\underline{C_2(p_1, p_2)} + \underline{C_3(p_1, p_2, p_3)} \end{aligned}$$

More details: Bzdak et al, arXiv:1607.07375, Lin et al arXiv:1512.09125

From Cumulants to Correlations (no anti-protons)

Defining integrated correlations function

$$C_n = \int dp_1 \dots dp_n C_n(p_1, \dots, p_n)$$

Simple Algebra leads to relation between correlations C_n and K_n

$$C_2 = -K_1 + K_2,$$

$$C_3 = 2K_1 - 3K_2 + K_3,$$

$$C_4 = -6K_1 + 11K_2 - 6K_3 + K_4, .$$

or vice versa

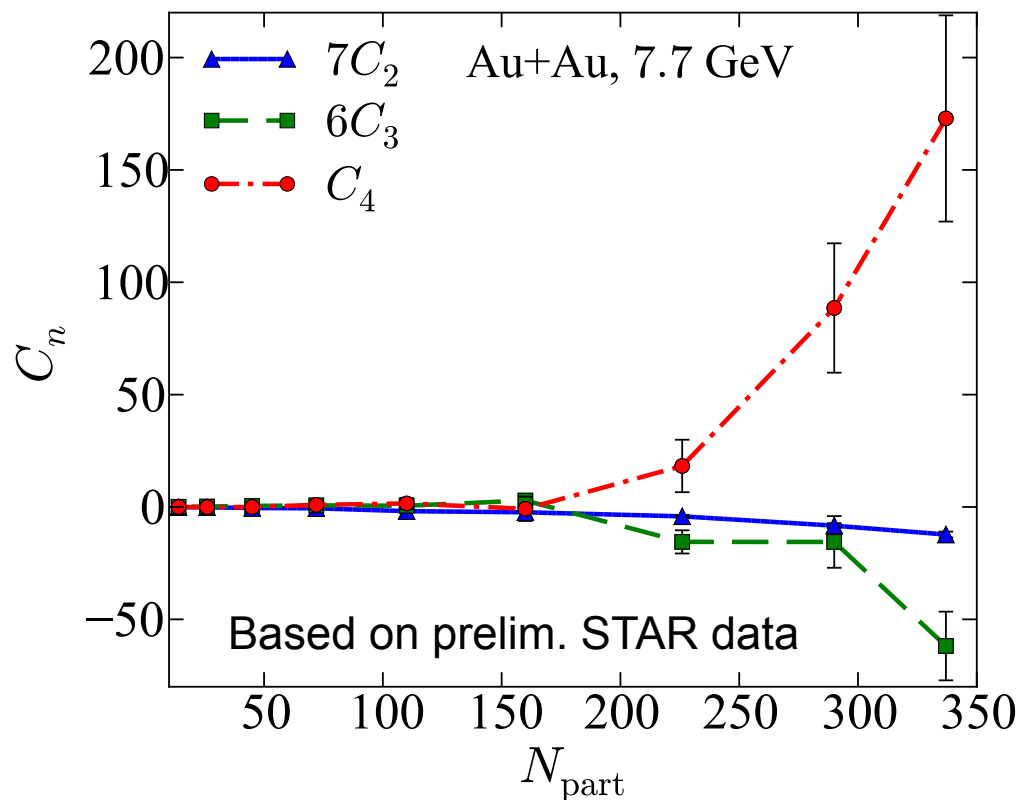
$$K_2 = \langle N \rangle + C_2$$

$$K_3 = \langle N \rangle + 3C_2 + C_3$$

$$K_4 = \langle N \rangle + 7C_2 + 6C_3 + C_4$$

Preliminary Star Data

(X. Luo, PoS Cpod 2014 (019))



Significant four particle correlations!

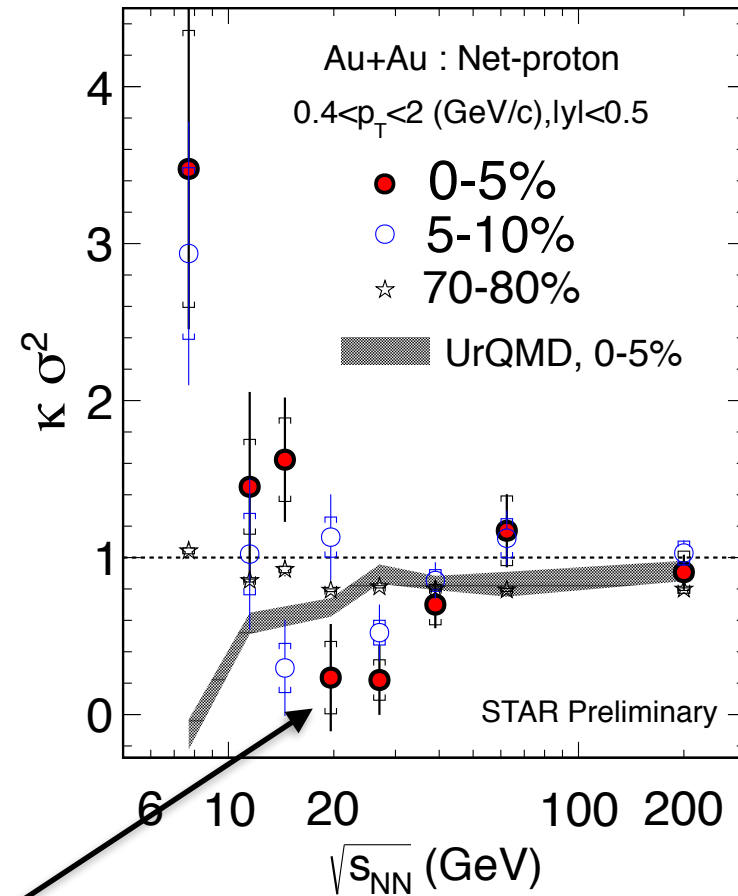
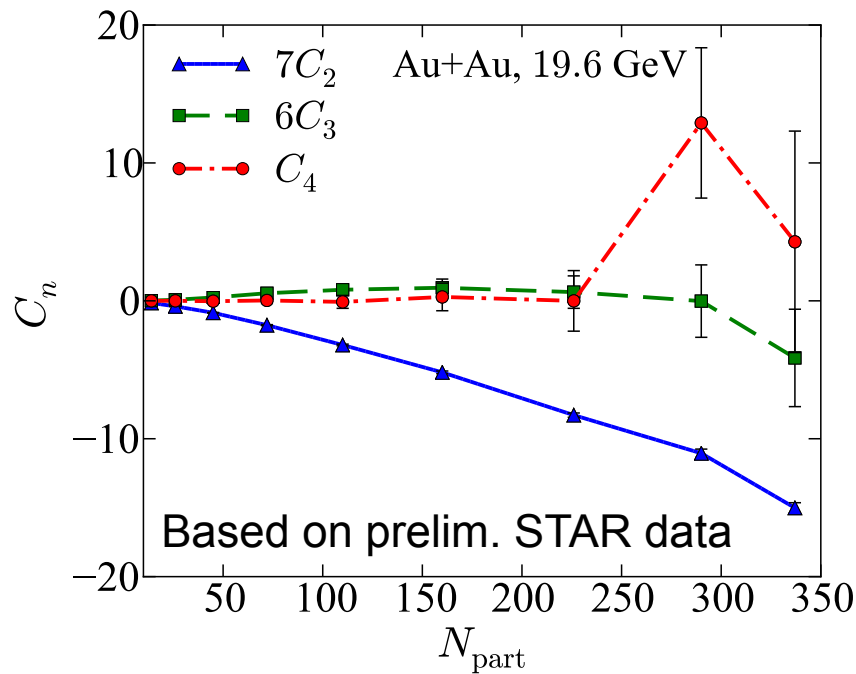
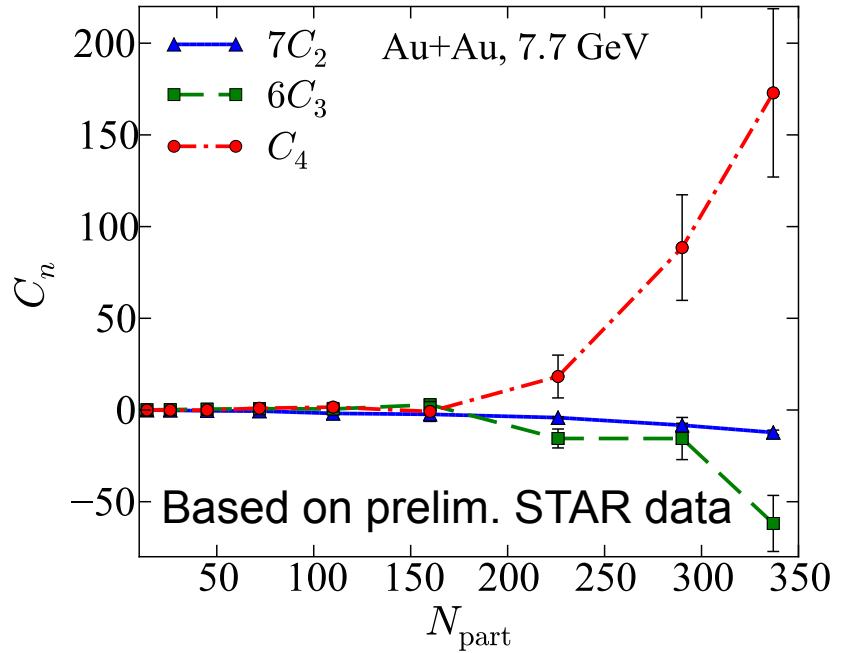
Four particle correlation dominate K_4 for central collisions at 7.7 GeV

$$K_2 = \langle N \rangle + C_2$$

$$K_3 = \langle N \rangle + 3C_2 + C_3$$

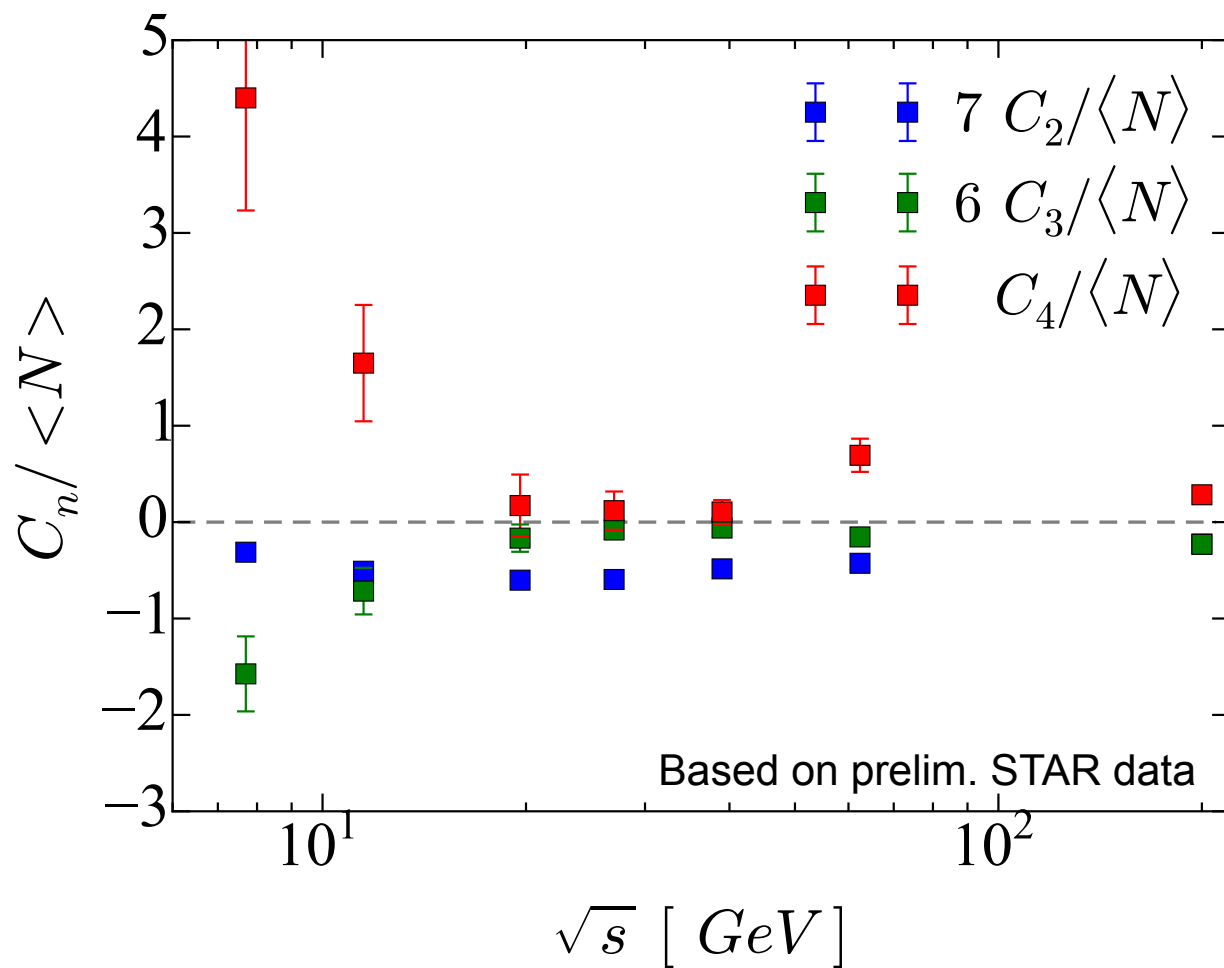
$$K_4 = \langle N \rangle + 7C_2 + 6C_3 + C_4$$

Correlations



Dip at 19.6 GeV from
NEGATIVE C_2 !

Energy dependence



Note: anti-protons are non-negligible above 19.6 GeV

Rapidity dependence

$$C_k(\Delta Y) = \int_{\Delta Y} dy_1 \dots dy_k \rho_1(y_1) \dots \rho_1(y_k) c_k(y_1, \dots, y_k)$$

Assume: $\rho_1(y) \simeq \text{const.}$

short range correlations:

$$c_k(y_1, \dots, y_k) \sim \delta(y_1 - y_2) \dots \delta(y_{k-1} - y_k)$$

$$C_k(\Delta Y) \sim \Delta Y \rightarrow K_k \sim \Delta Y$$

Long range correlations:

$$c_k(y_1, \dots, y_k) = \text{const.}$$

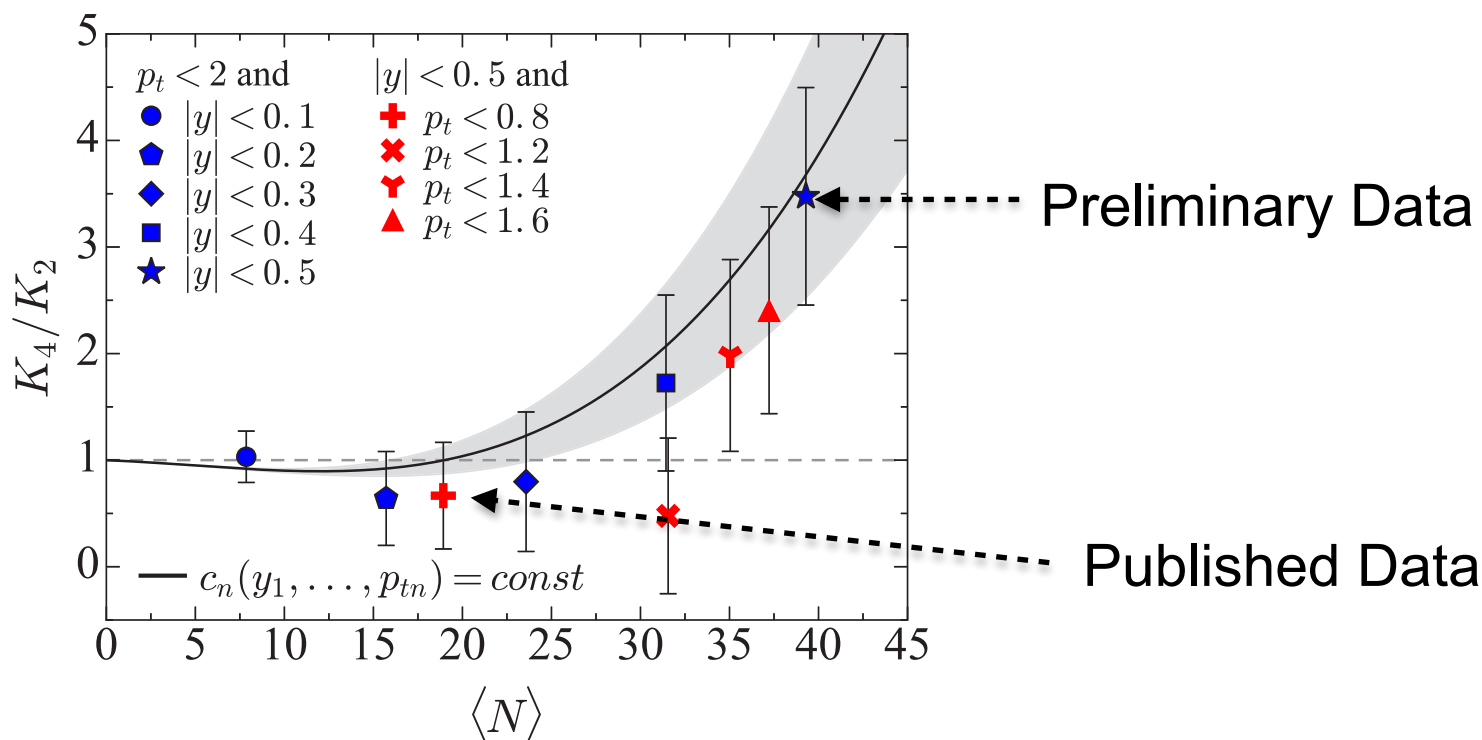
$$C_k(\Delta Y) \sim (\Delta Y)^k \sim \langle N \rangle^k$$

$$\Rightarrow K_n = K_n(\langle N \rangle)$$

Long range correlations

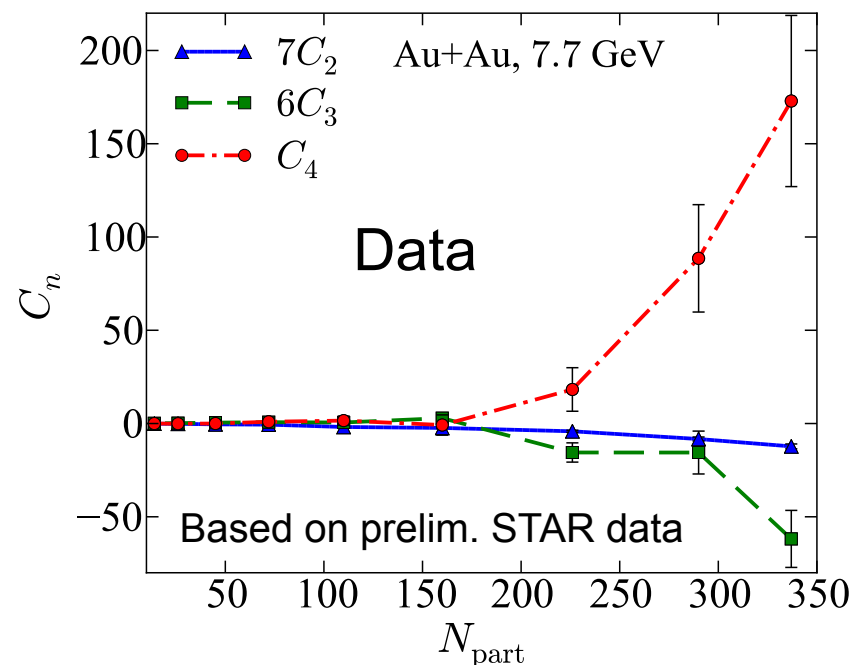
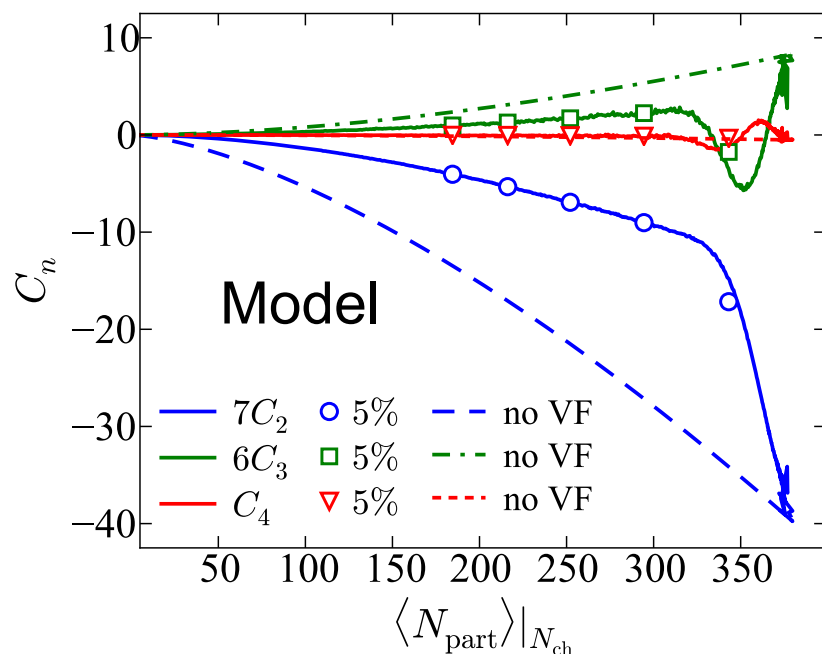
$$C_k = \langle N \rangle^k c_k$$

$$c_k = \text{const.} \Rightarrow K_n = K_n(\langle N \rangle)$$



Can we understand these correlations?

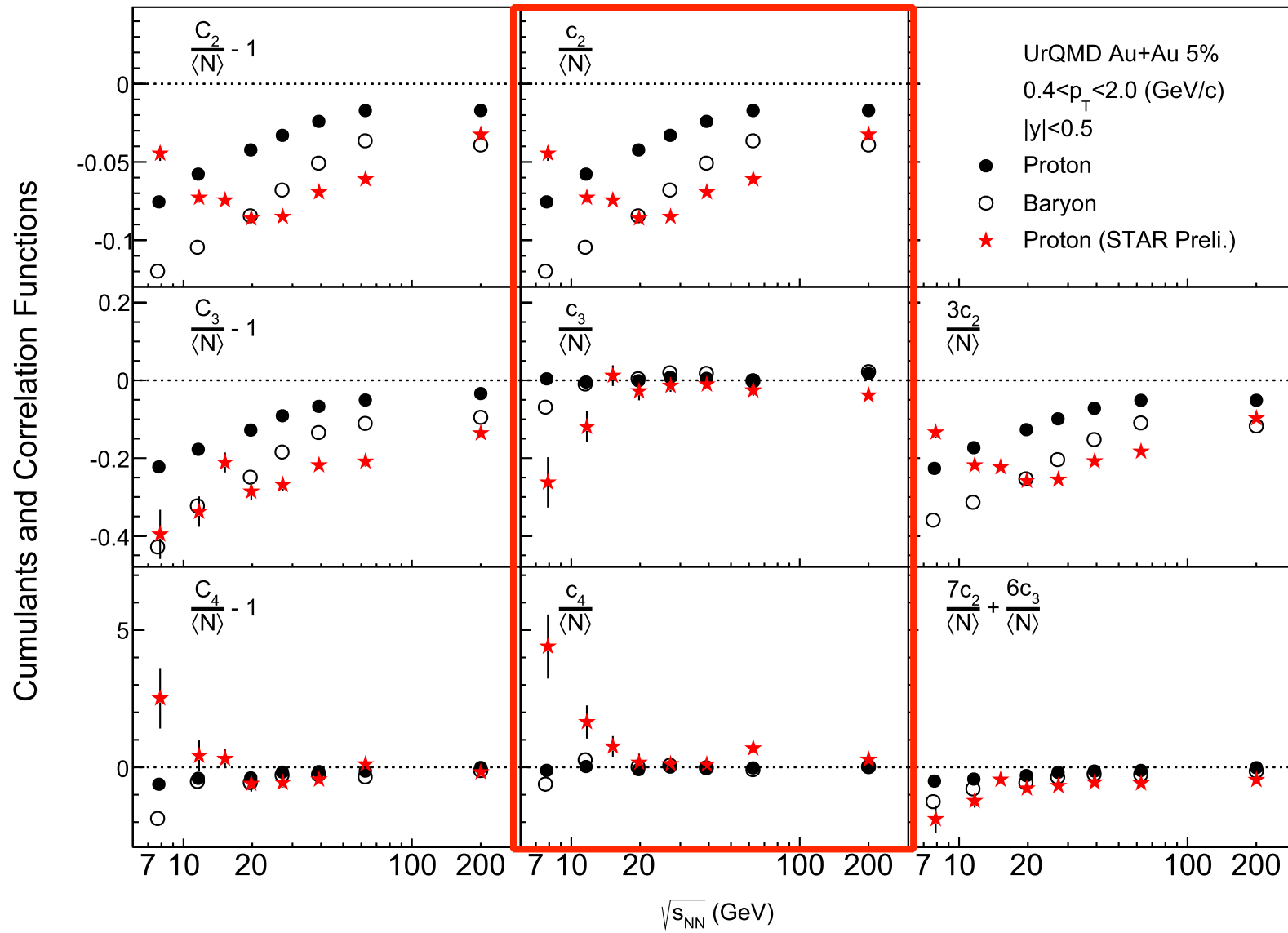
- Two particle correlations can be understood by simple Glauber model + Baryon number conservation



Four particle correlations are orders of magnitudes larger in the data
Also seen in URQMD calculations by He et al. PLB774 (2017) 623

Need to assume the ~40% of protons come from 8-nucleon cluster in order to get magnitude right!

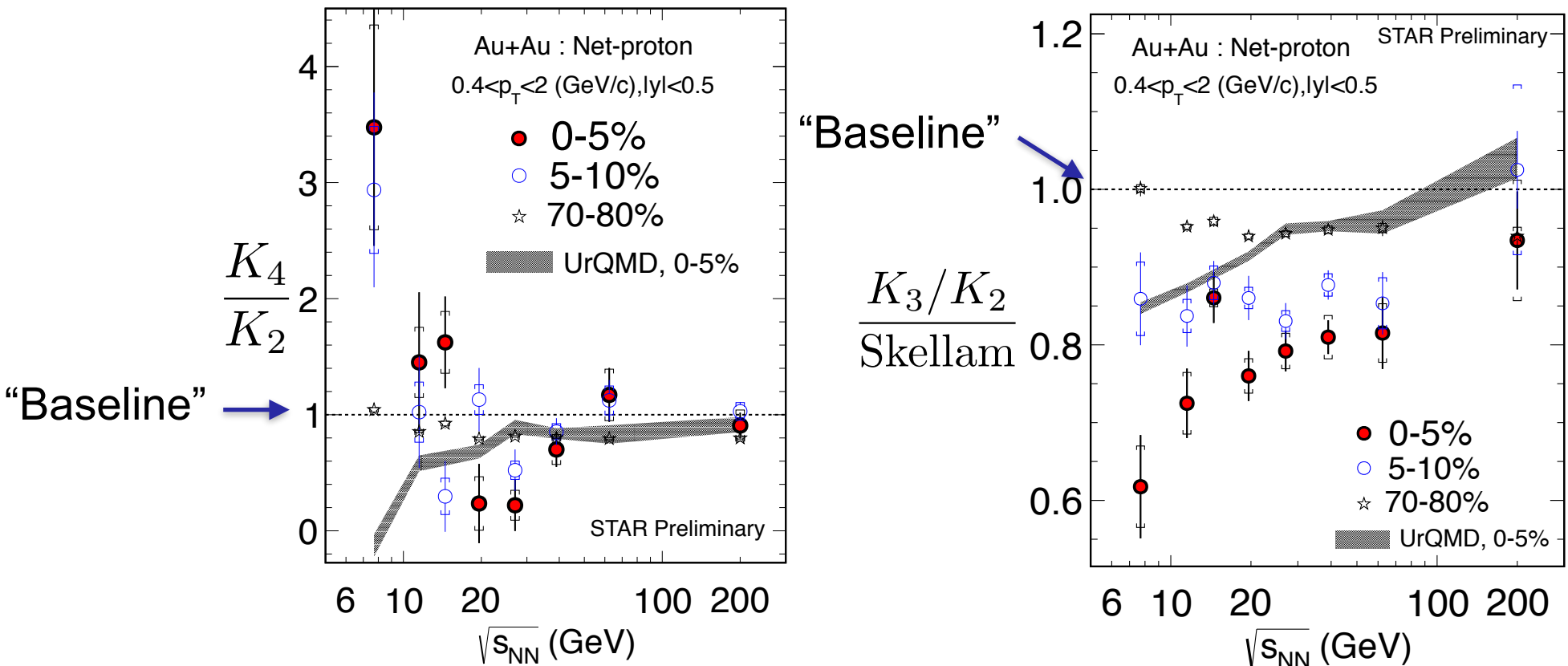
URQMD



He, Luo PLB774 (2017) 623

Latest STAR result on net-proton cumulants

X. Luo, NPA 956 (2016) 75

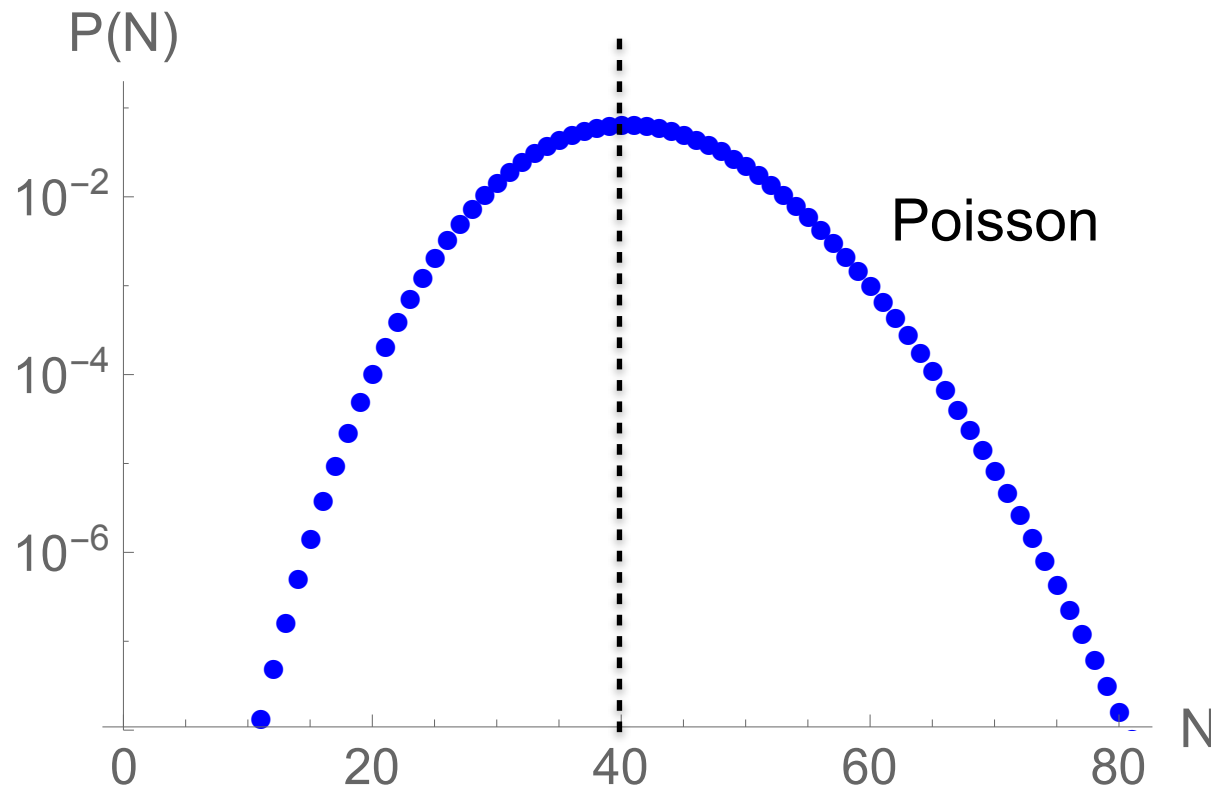


K_4/K_2 above baseline K_3/K_2 below baseline

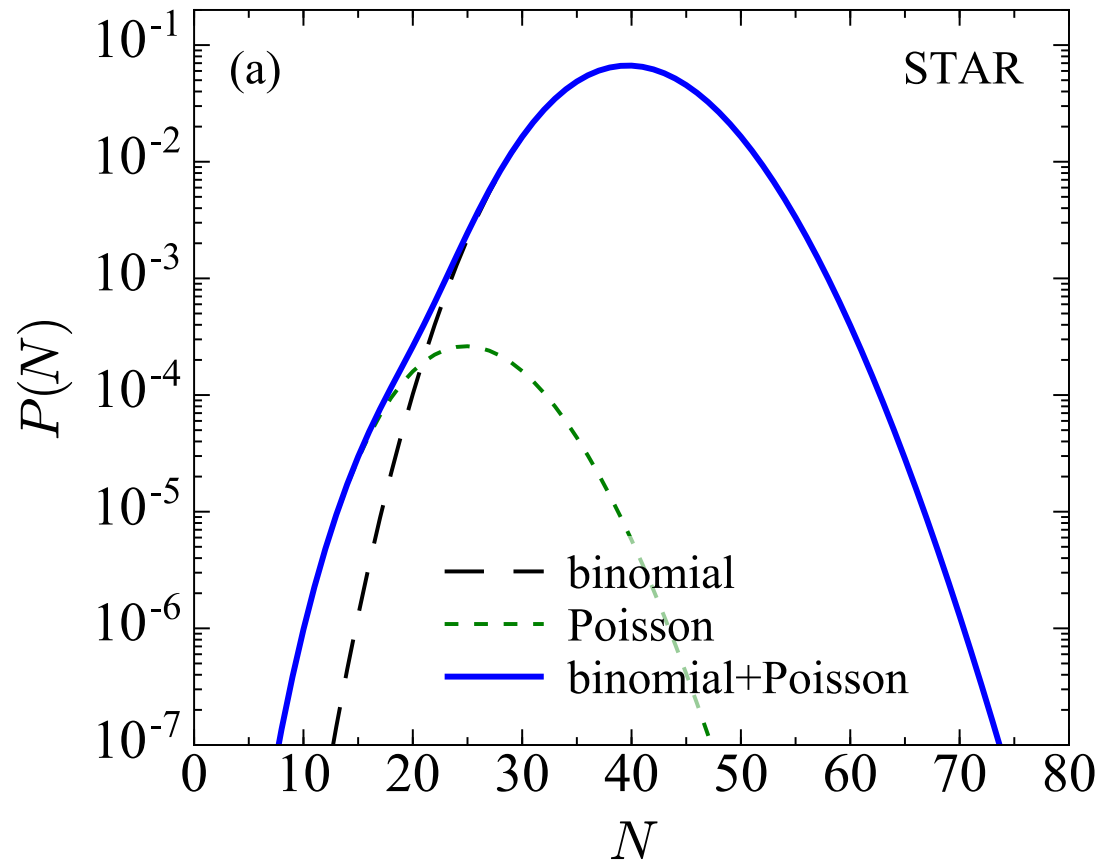
Shape of probability distribution

$$K_3 < \langle N \rangle \quad K_3 = \langle N - \langle N \rangle \rangle^3$$

$$K_4 > \langle N \rangle \quad K_4 = \langle N - \langle N \rangle \rangle^4 - 3 \langle N - \langle N \rangle \rangle^2$$



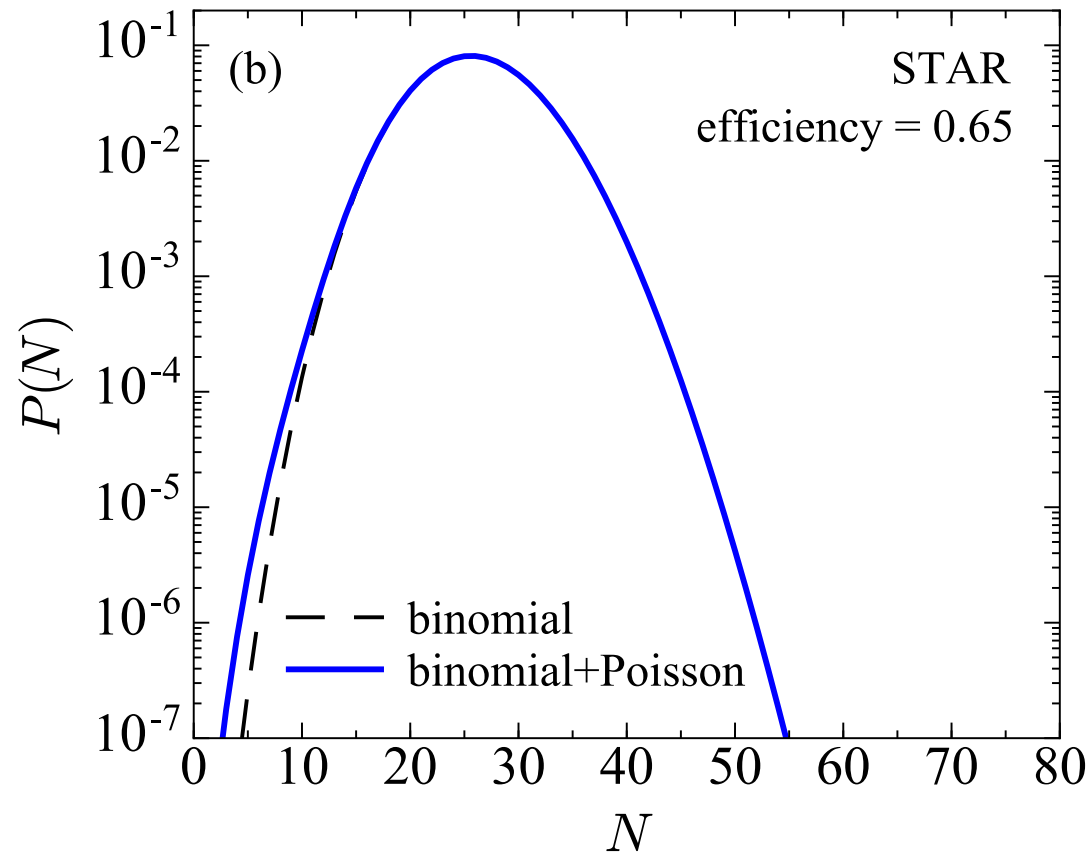
Simple two component model



Weight of small component: $\sim 0.3\%$

Simple two component model

Difficult to see in the real data with efficiency $\varepsilon=0.65$



Two component model

$$P(N) = (1 - \alpha)P_{(a)}(N) + \alpha P_{(b)}(N)$$

$$\bar{N} = \langle N_{(a)} \rangle - \langle N_{(b)} \rangle$$

$$C_2 = C_2^{(a)} - \alpha \{ \bar{C}_2 - (1 - \alpha) \bar{N}^2 \}$$

$$C_3 = C_3^{(a)} - \alpha \{ \bar{C}_3 + (1 - \alpha) [(1 - 2\alpha) \bar{N}^3 - 3\bar{N} \bar{C}_2] \}$$

$$C_4 = C_4^{(a)} - \alpha \{ \bar{C}_4 - (1 - \alpha) [(1 - 6\alpha + 6\alpha^2) \bar{N}^4 - 6(1 - 2\alpha) \bar{N}^2 \bar{C}_2 + 4\bar{N} \bar{C}_3 + 3(\bar{C}_2)^2] \}$$

$$\bar{C}_n = C_n^{(a)} - C_n^{(b)},$$

For Poisson, $C_{(a)}, C_{(b)}=0$

Fit to STAR data: $\langle N_{(a)} \rangle \simeq 40, \quad \langle N_{(b)} \rangle \simeq 25, \quad \alpha \simeq 0.003$

Two component model

$$P(N) = (1 - \alpha)P_{(a)}(N) + \alpha P_{(b)}(N)$$

$$\bar{N} = \langle N_{(a)} \rangle - \langle N_{(b)} \rangle > 0$$

For $P_{(a)}$, $P_{(b)}$ Poisson, or (to good approximation) Binomial

$$C_n = (-1)^n K_n^B \bar{N}^n \quad n \geq 2$$

K_n^B : Cumulant of Bernoulli distribution

$$\alpha \ll 1, K_n^B = \alpha \Rightarrow C_n \simeq \alpha (-1)^n \bar{N}^n$$

$\Rightarrow |C_n| \sim \langle N \rangle^n$ as seen by STAR (i.e. “infinite” correlation length)

predict:

$$\frac{C_4}{C_3} = \frac{C_5}{C_4} = \frac{C_{n+1}}{C_n} = -\bar{N}$$

$$\bar{N} \simeq 15$$

Clear and falsifiable prediction:

$$C_5 \approx -2650 \quad C_6 \approx 41000$$

This model can be tested RIGHT NOW!

Model prediction:

$$\begin{aligned} C_5 &= -2645 (1 \pm 0.14), & C_6 &= 40900 (1 \pm 0.18), \\ C_7 &= -615135 (1 \pm 0.26), & C_8 &= 8520220 (1 \pm 0.42) \end{aligned}$$

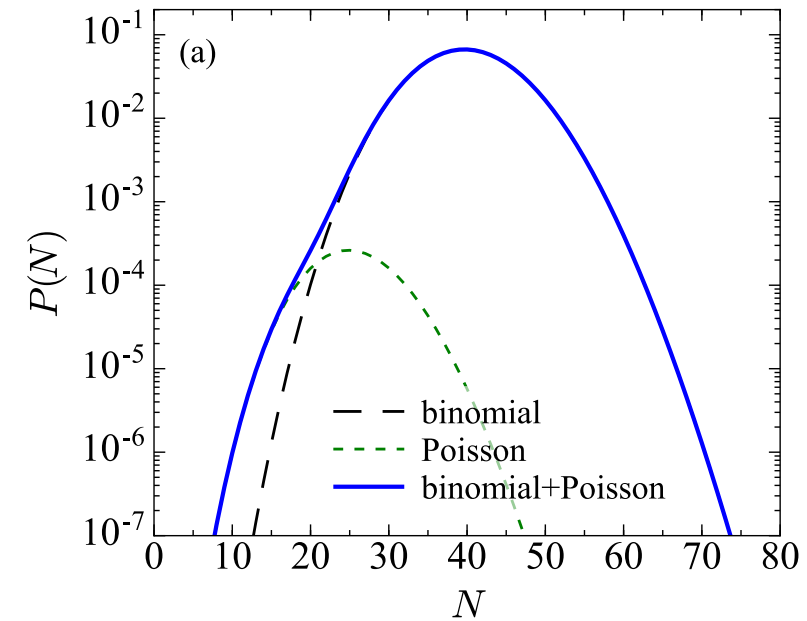
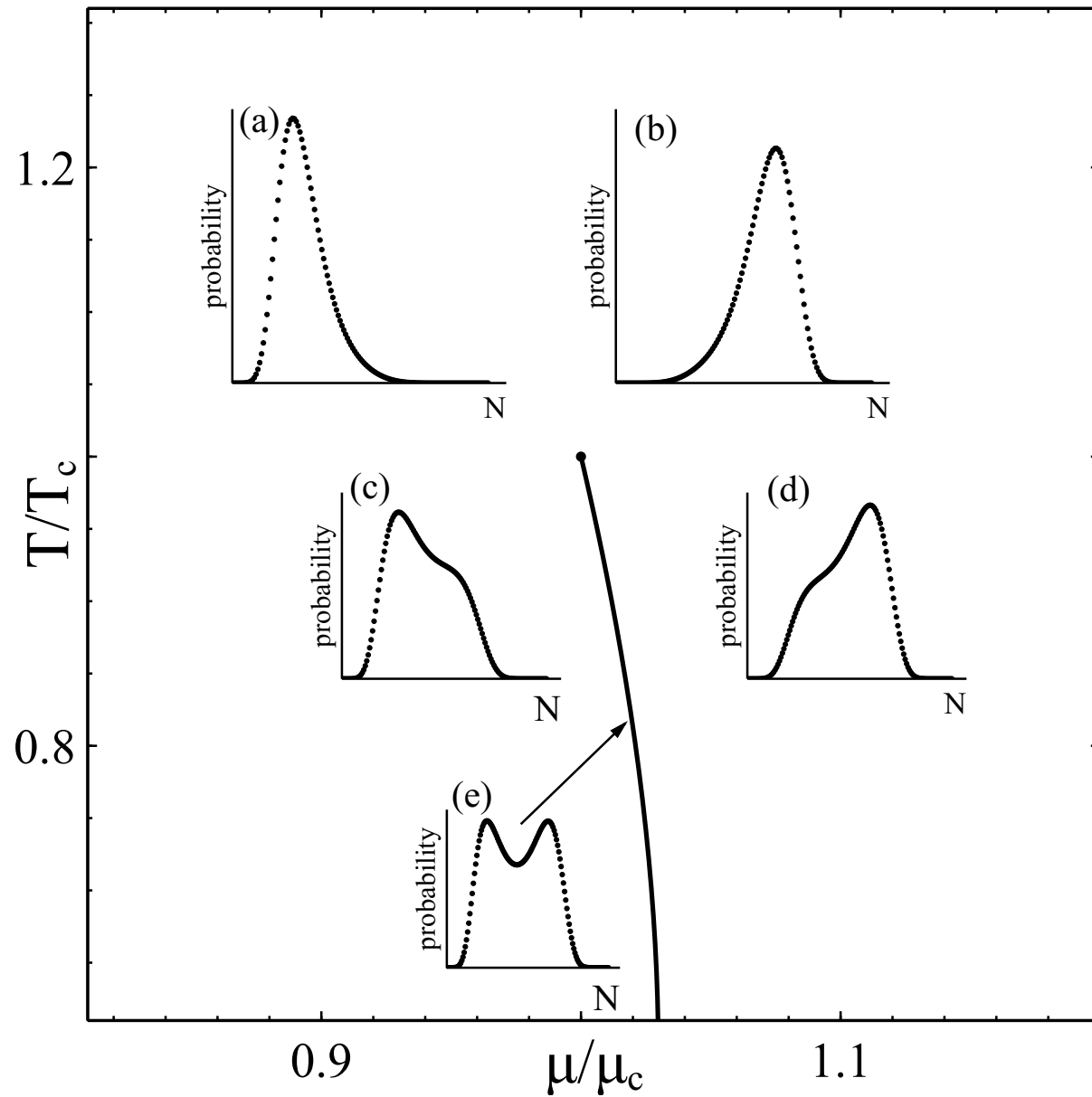
Efficiency
corrected

$$\begin{aligned} C_5 &= -307 (1 \pm 0.31), & C_6 &= 3085 (1 \pm 0.41), \\ C_7 &= -30155 (1 \pm 0.61), & C_8 &= 271492 (1 \pm 1.06), \end{aligned}$$

Efficiency
UN-corrected

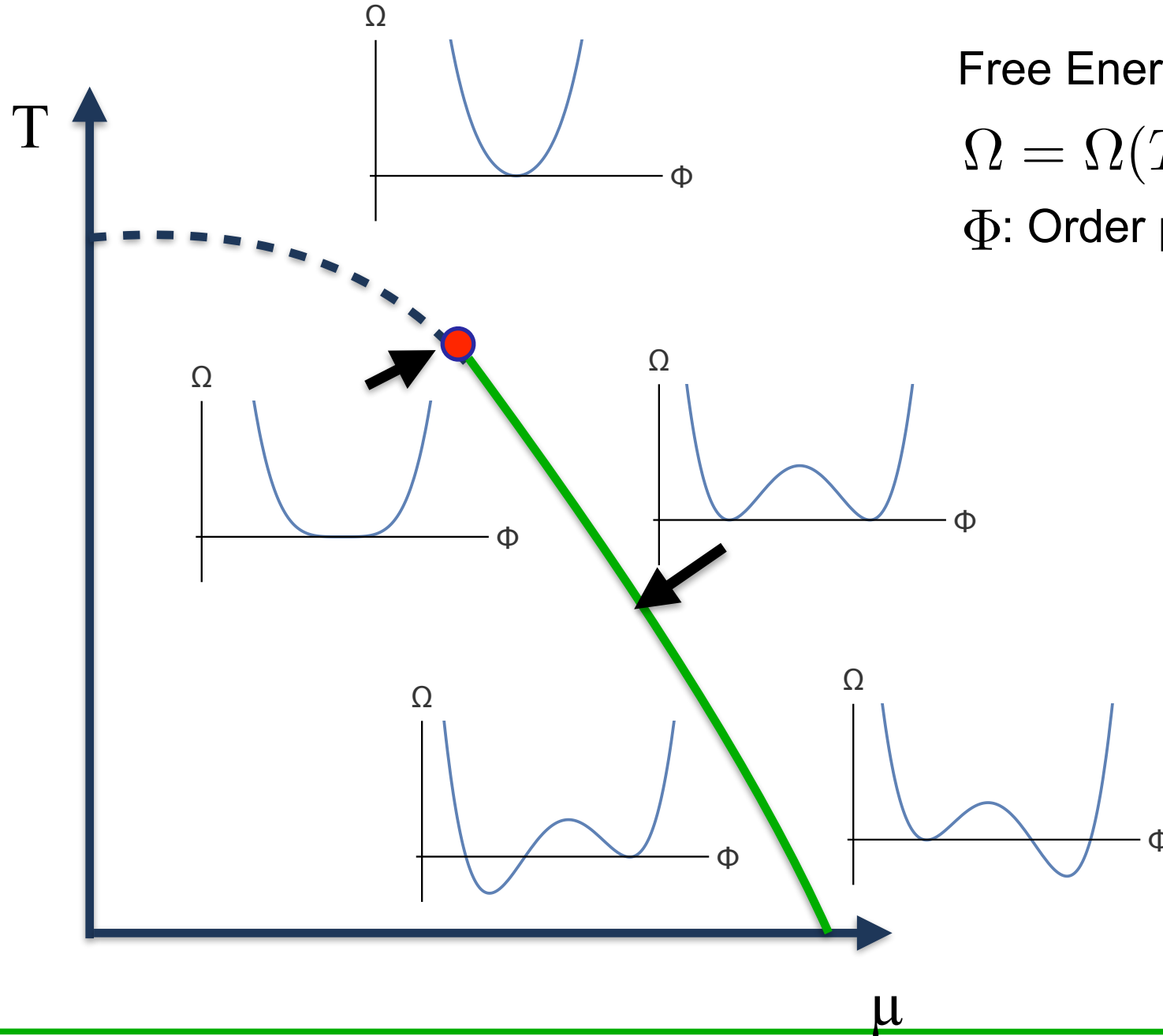
Based on 144393 events (same as STAR 0-5% at 7.7 GEV)

Speculation



Bzdak et al, arXiv:1804.04463

Free Energy

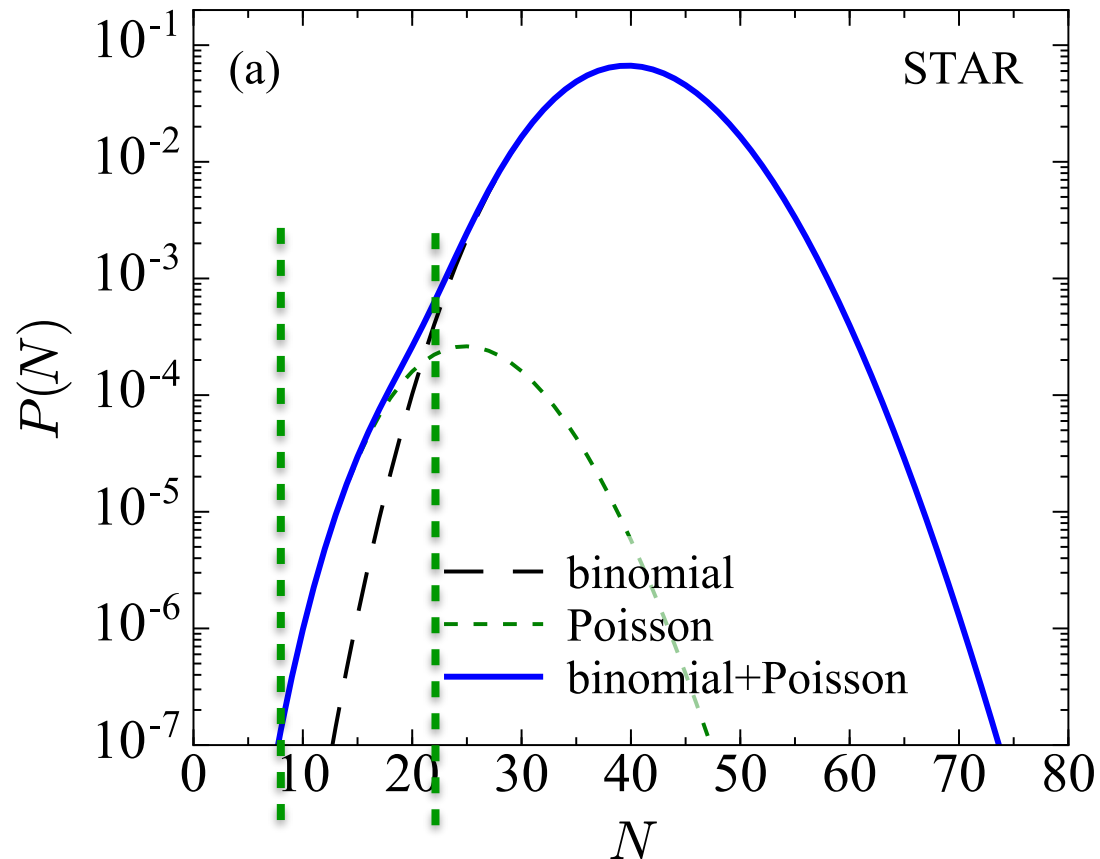


Free Energy:

$$\Omega = \Omega(T, \mu; \Phi)$$

Φ : Order parameter

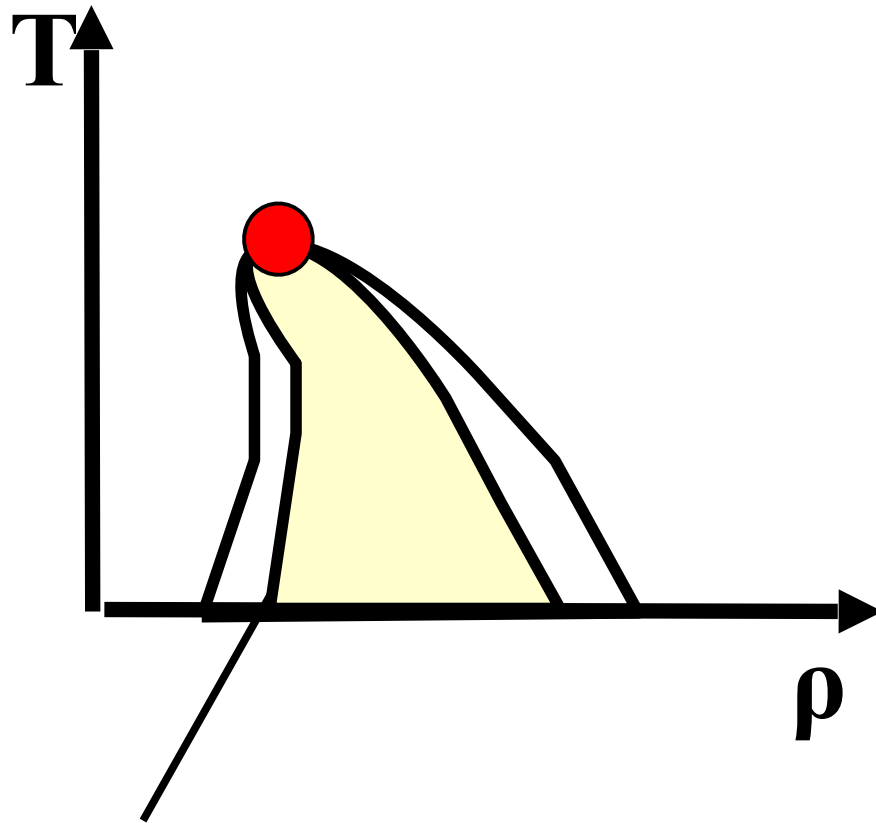
Simple two component model



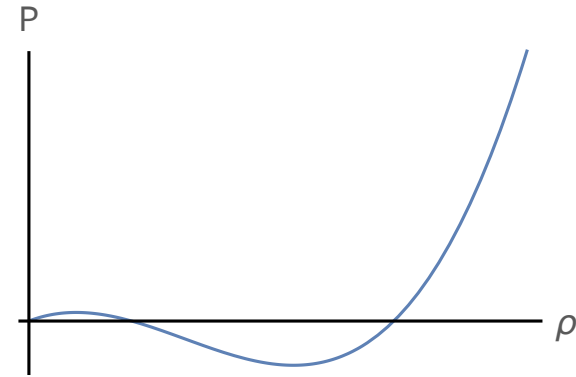
Analyse data for $N_p < 20$

- Is flow etc different?
- “Inspect by eye (<1% of all events)”

Co-existence region



System should spent long time
in spinodal region



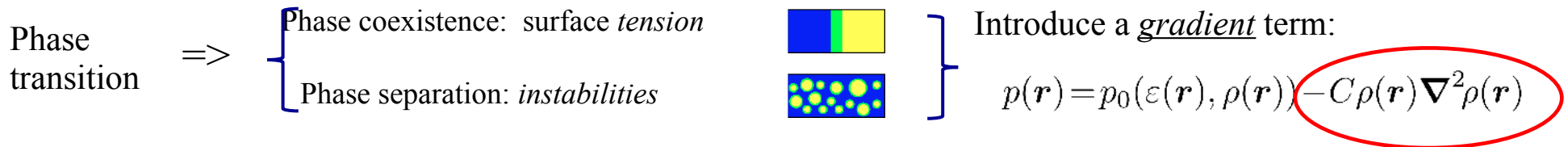
Spinodal instability:
Mechanical instability

$$\frac{\partial P}{\partial \rho} < 0$$

Exponential growth of clumping

Non-equilibrium phenomenon!

Phase-transition dynamics: Density clumping

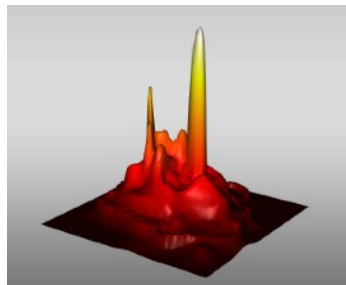


Insert the modified pressure into existing ideal finite-density fluid dynamics code

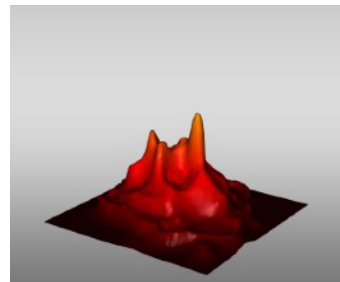
Use UrQMD for pre-equilibrium stage to obtain fluctuating initial conditions

Simulate central Pb+Pb collisions at ≈ 3 GeV/A beam kinetic energy on fixed target, using an Equation of State either with a phase transition or without (Maxwell partner):

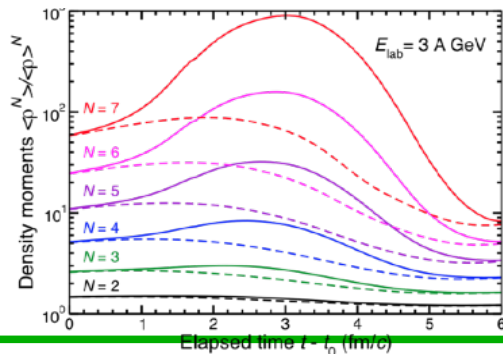
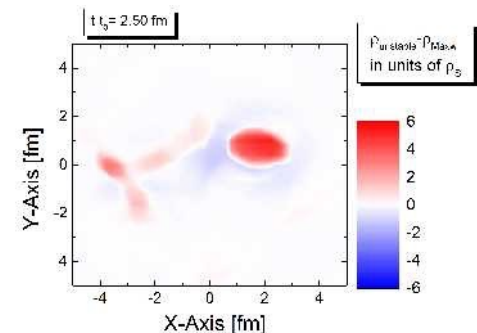
With phase transition:



Without phase transition:



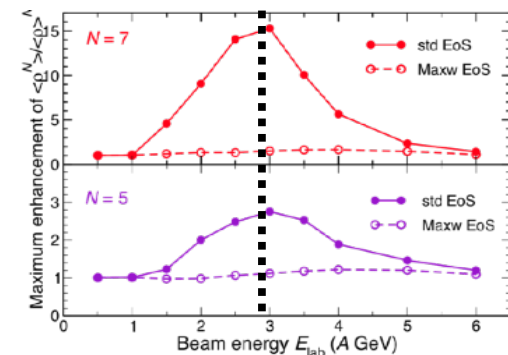
Density enhancement:



Evolution of density moments

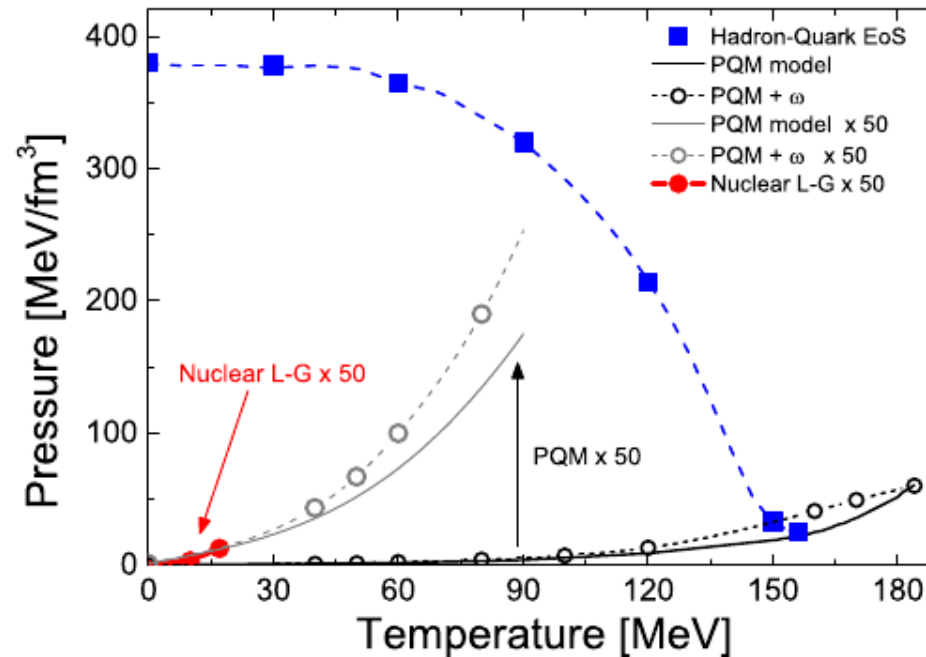
$$\langle \rho^N \rangle \equiv \frac{1}{A} \int \rho(\mathbf{r})^N \rho(\mathbf{r}) d^3\mathbf{r}$$

J. Steinheimer & J. Randrup,
PRL 109, 212301(2012)
PRC 87, 054903 (2013)

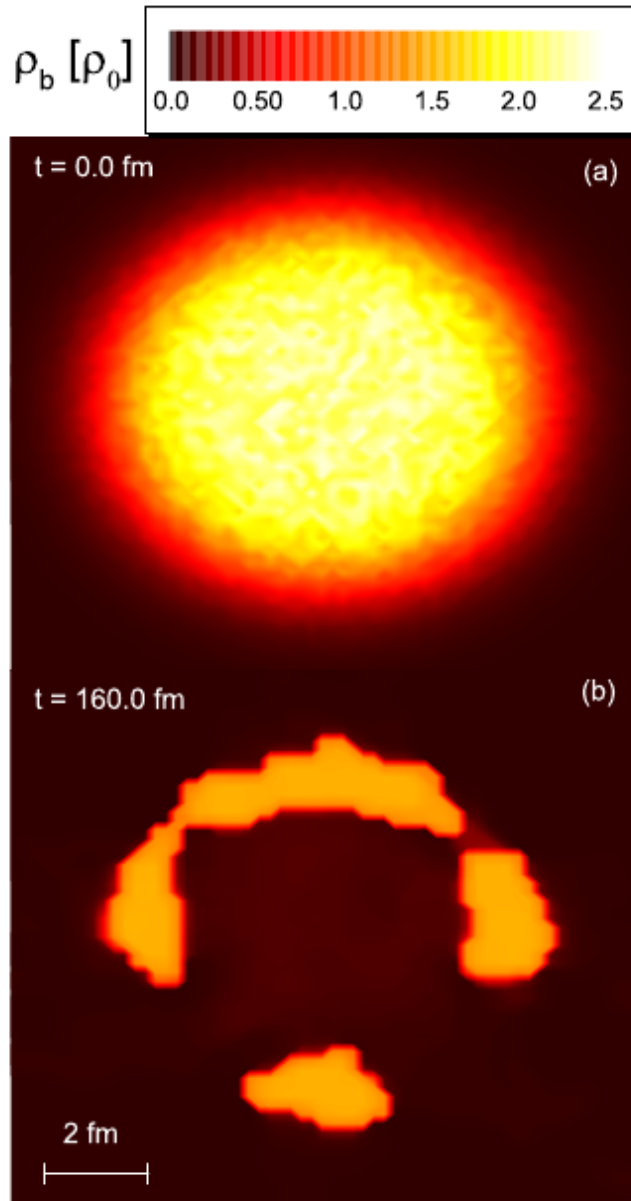


$E_{Lab} = 3$ GeV

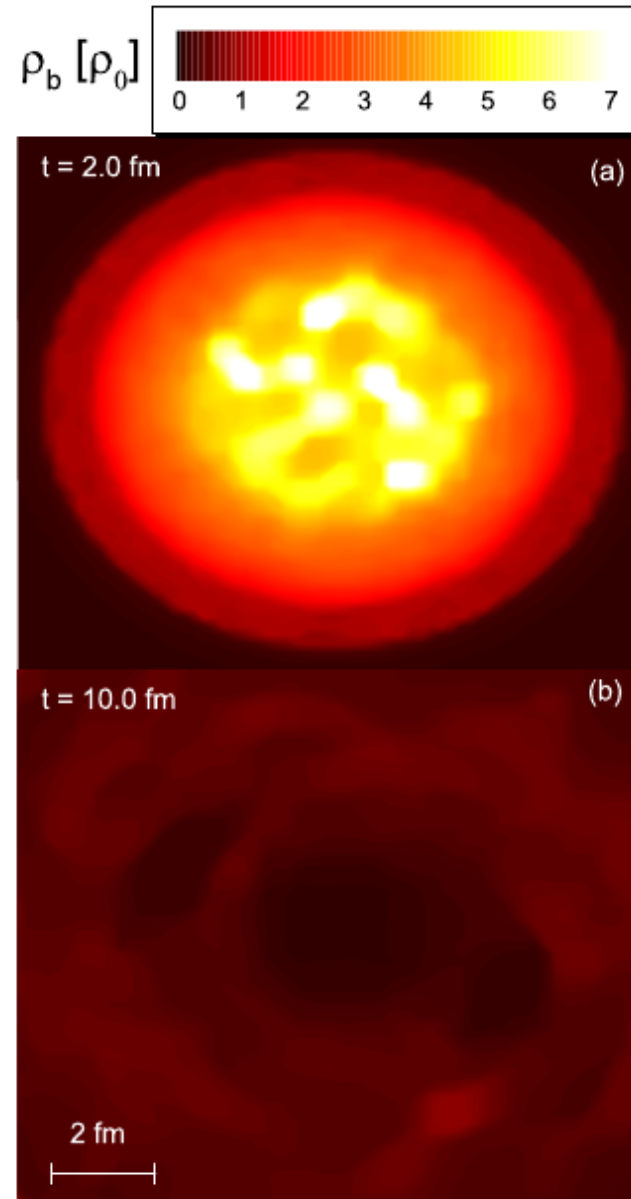
Consider two Equations of State



Steinheimer et al,
Phys.Rev. C89 (2014) 034901

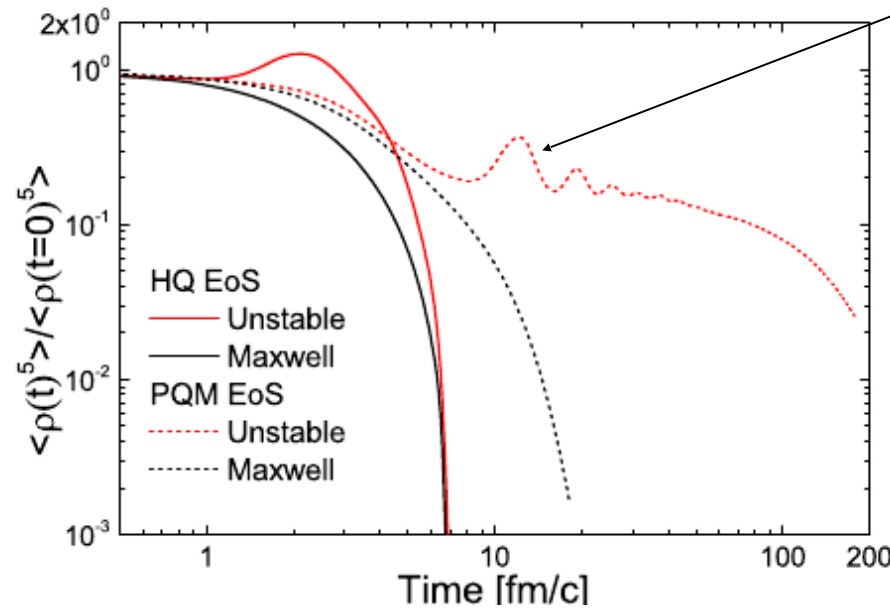


PQM (“liquid-gas”)



“QCD”

Time evolution



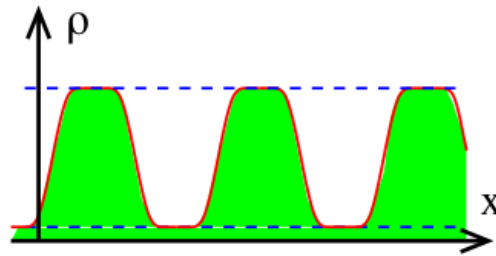
Oscillation of nearly stable droplets for “liquid-gas” EoS

Higher pressure leads to faster evolution of “QCD” EoS.

Steinheimer et al,
Phys.Rev. C89 (2014) 034901

Cluster a.k.a. nuclei

Even if total baryon number does not fluctuate the baryon **density** does

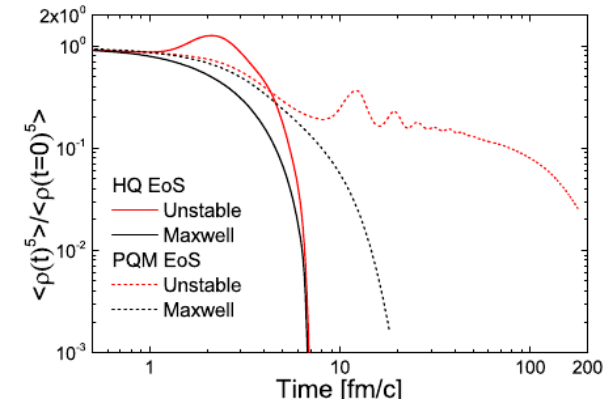


Therefore measure production of NUCLEI: d, ^3He , ^4He , ^7Li

$$\langle d \rangle \sim \langle \rho^2 \rangle \quad \langle {}^3\text{He} \rangle \sim \langle \rho^3 \rangle \quad \langle {}^7\text{Li} \rangle \sim \langle \rho^7 \rangle$$

Extracts higher moments of the baryon **density** at freeze out

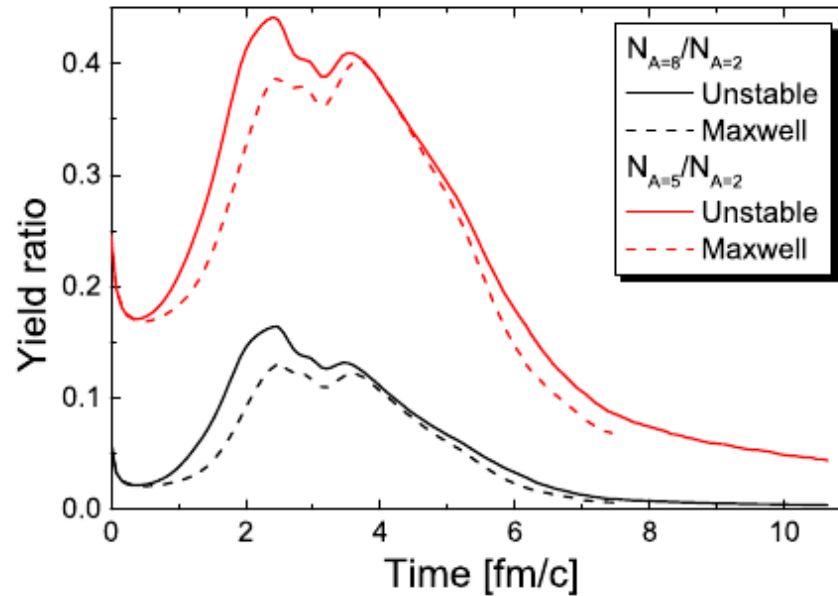
Nice Idea, but...



“Cluster” formation

“QCD” EoS

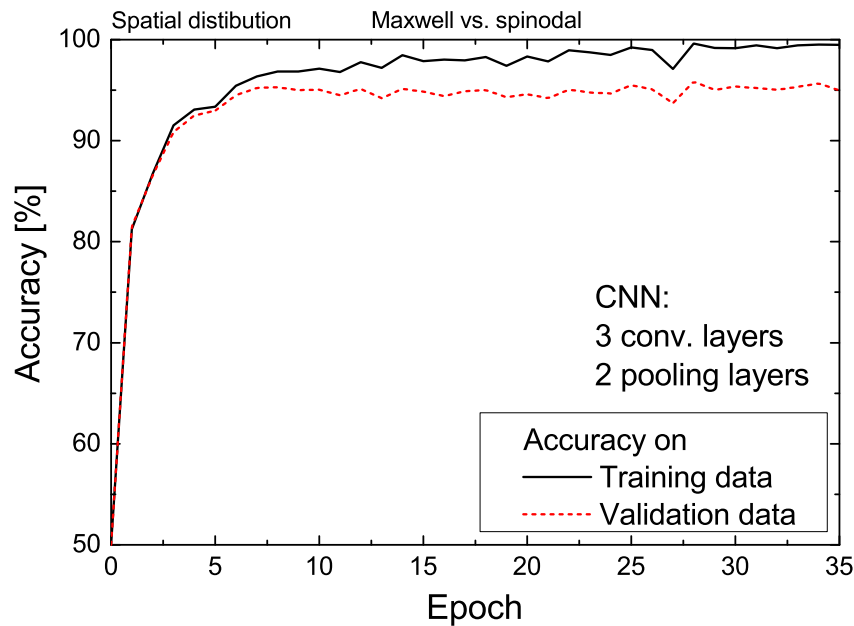
$$\left(\frac{S}{B}\right)_{hadron-gas} < \left(\frac{S}{B}\right)_{QGP-liquid}$$



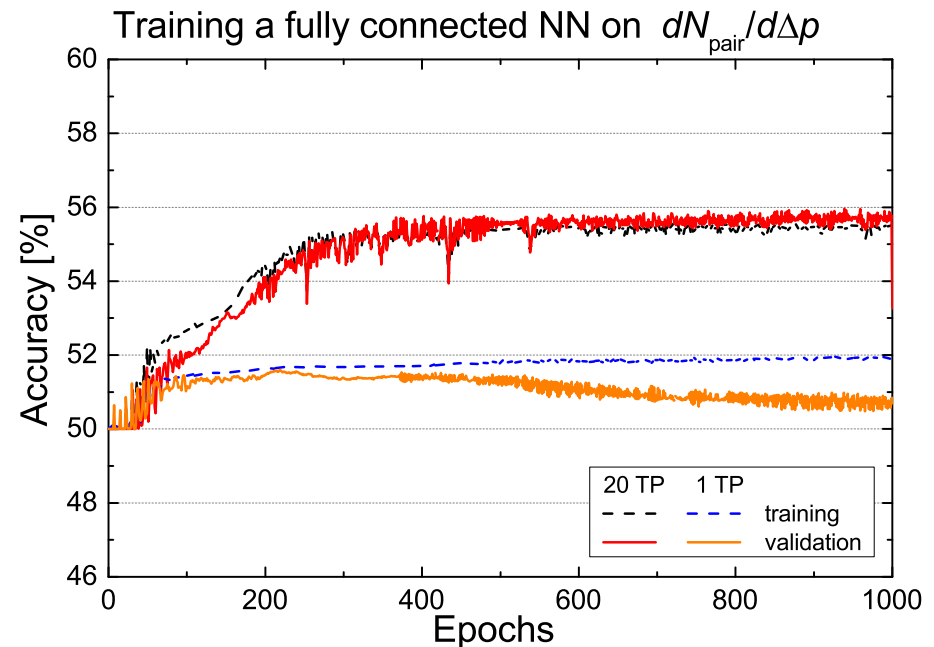
Clumping in coordinate space is compensated by dilution in momentum space → tiny effect

Steinheimer et al,
Phys.Rev. C89 (2014) 034901

“Deep” learning fails as well....



Coordinate space 😊



Momentum space 😞

Steinheimer et al., arXiv:1906.06562

Summary

- Cumulants measure derivatives of the free energy (equation of state)
 - Sensitive to “wiggles” a.k.a. “remnants” of phase transition
- Experiments are difficult: detector needs to be understood well
- Careful when comparing theory with measured cumulants
- Correlations a.k.a. factorial cumulants provide complementary insights
 - strong four particle correlation at low energies
- Don't forget the first order phase transition
 - Spinodal instability
- Very active field, both in experiment and theory

VERY INTERESTING TIMES ADHEAD