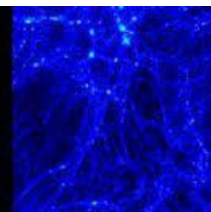


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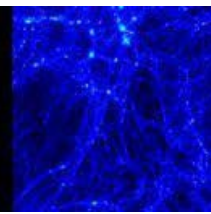


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- $U(1)_B$  and  $U(1)_L$  conservation has important consequences
  - **Baryon number conservation: Proton stability**
  - **Lepton number conservation: Dirac neutrinos**

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  - **Usual Choice:** Explicitly break  $U(1)_L \rightarrow Z_2$  in UV completions (seesaw or loop) of Weinberg Operator

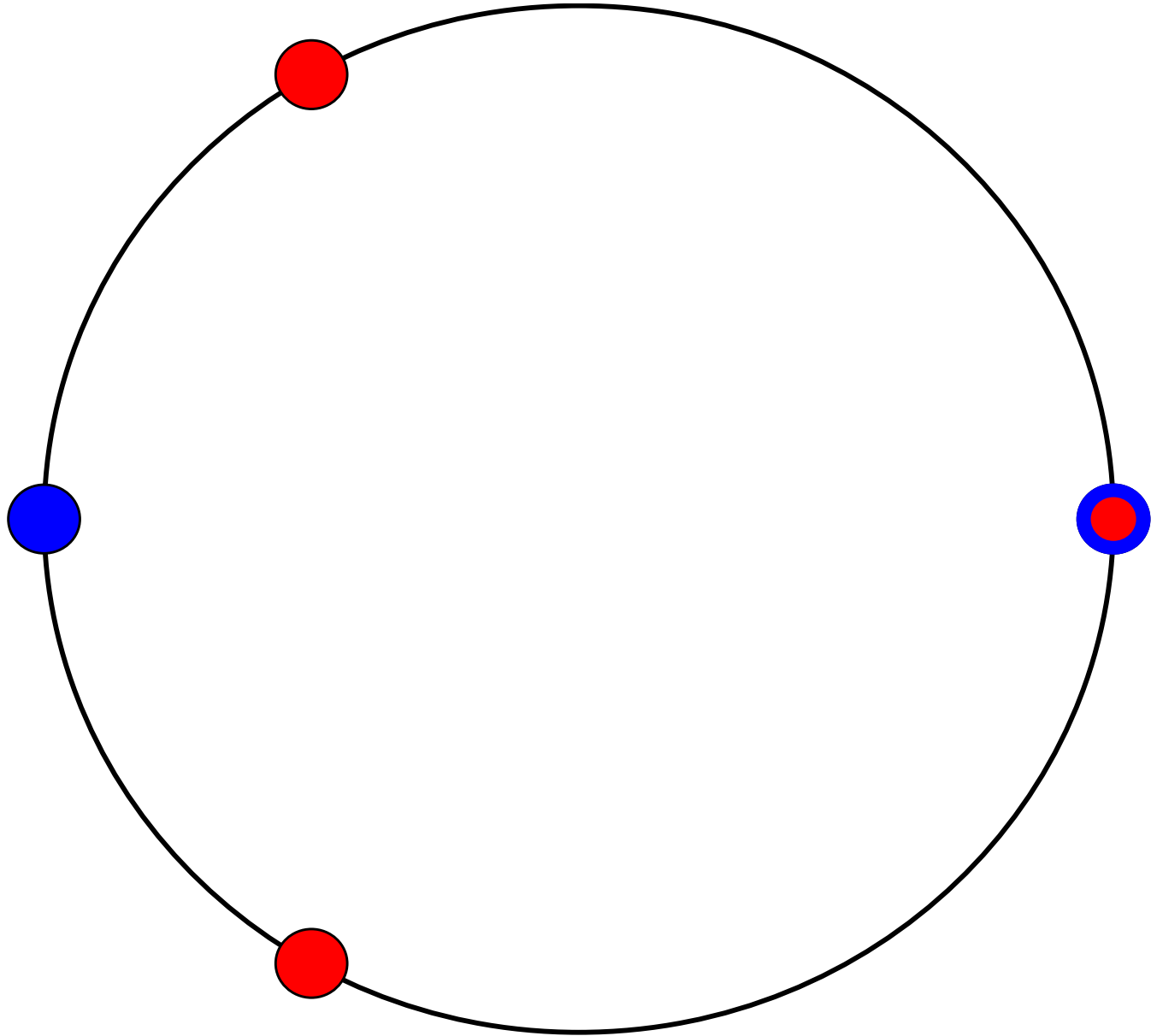
# Nature of Neutrinos

## Lepton Number Breaking Pattern [Hirsch, RS, Valle '17]

- If  $U(1)_L$  is conserved: **Neutrinos are Dirac**
  - Accidental Symmetry of SM: New physics beyond SM need not conserve it
- If  $U(1)_L$  is broken: **Symmetry breaking pattern will determine the nature of neutrinos**
  - $U(1)_L$  symmetry only admits  $Z_M$  subgroups i.e. cyclic groups of  $m$  elements
  - If  $x$  is a non-identity group element of  $Z_M$ , then  $x^{M+1} \equiv x$
  - The  $Z_M$  groups only admit one-dimensional irreducible representations
  - Conveniently represented by using the  $n$ -th roots of unity,  $\omega = \text{Exp}[2\pi I/M]$  where  $\omega^M = 1$



# Residual Subgroups



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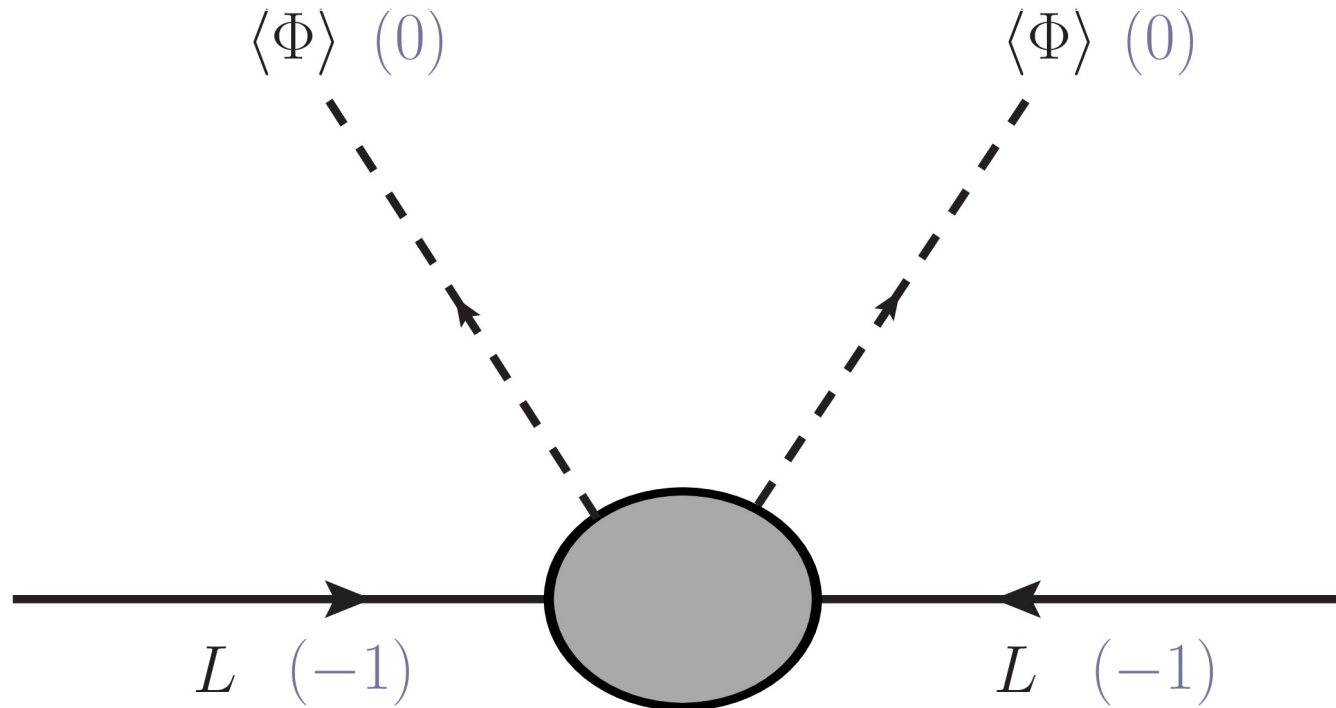
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- From symmetry point of view: **Dirac neutrinos are more natural !!!**

# Majorana Neutrinos: Weinberg Operator

- Weinberg Operator: Provides “effective” description of how Majorana neutrino mass can be generated **[S. Weinberg '79]**



- Breaks  $U(1)_L \rightarrow \mathcal{Z}_2$ 
  - Both reps of  $\mathcal{Z}_2$  satisfy the Majorana condition
  - **All UV completions of Weinberg operator will always lead to Majorana neutrinos**

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  - No extra explicit or accidental symmetries

# Lepton Number of Right Handed Neutrinos

- Baryon and Lepton number of all SM particles are fixed
- What is the Lepton number of Right Handed Neutrinos?
  - B and L symmetries are anomalous
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- Vector solution : Add three right handed neutrinos with B-L charges of  $(-1,-1,-1)$

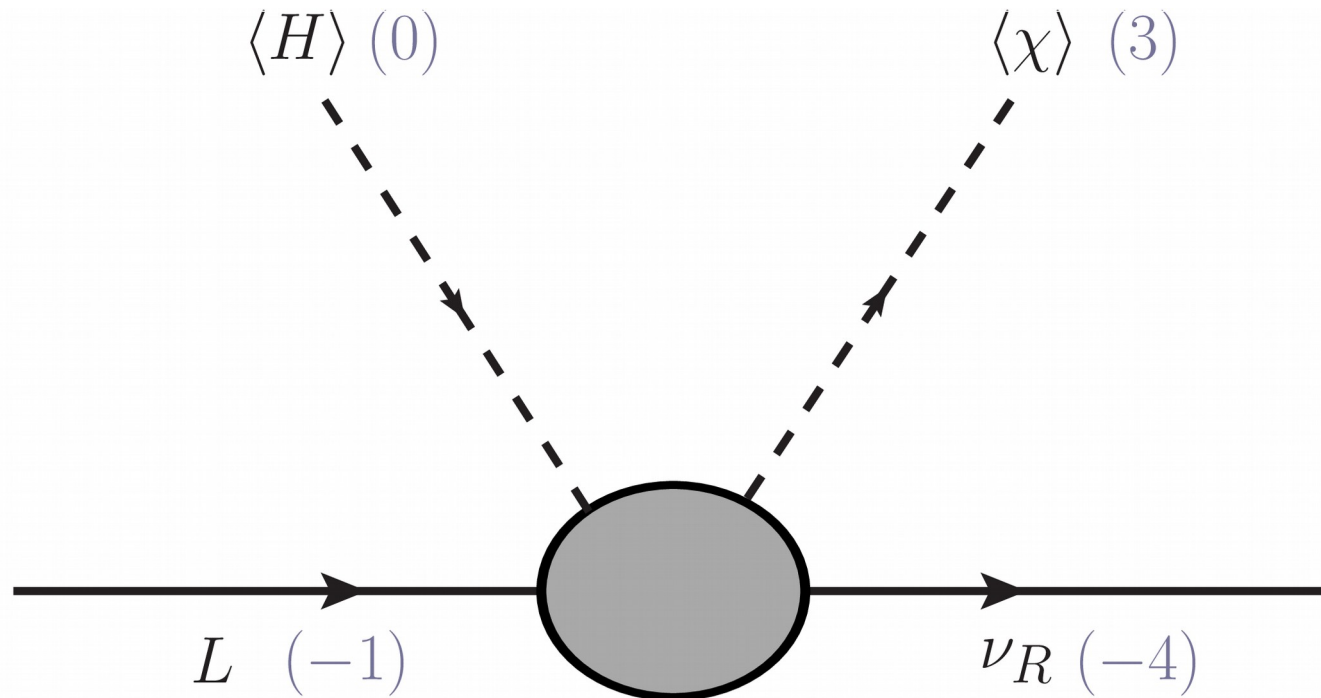
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- **New Chiral Solution: Right handed neutrinos with B - L charges of (-4,-4,5) [Ma, RS '14, Pollard, Ma, RS, Reza '15]**
  - Yukawa term  $L\tilde{\Phi}\nu_R$  automatically forbidden
  - Paves way for “naturally small” Dirac neutrino masses:  
**Dirac neutrino mass mechanisms**  
**[RS et.al '15,'16,'17'18'19, Several other]**

# Generalized Weinberg Operator

- Neutrino Mass can be generated at dim-5 level

[C.Bonilla,S.C.Chulia,R.Cepedello,E.Peinado,RS '18,'19]



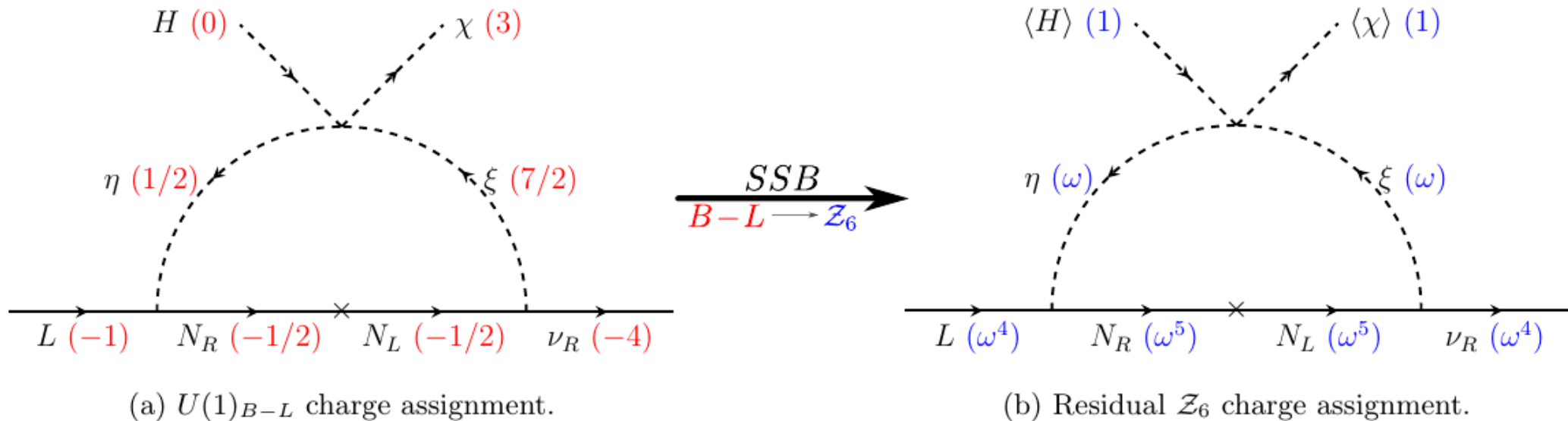
- Since  $\chi \sim 3$ , its vev breaks  $U(1)_{B-L} \rightarrow \mathbb{Z}_{3m}; m \in \mathbb{Z}^+$ 
  - The exact residual subgroup depends on UV completion

# UV Completion

- One loop completion: Dark Sector particles in the loop *a la* Scotogenic models  
**[C.Bonilla,S.C.Chulia,R.Cepedello,E.Peinado,RS '18,'19]**
- The Residual  $\mathcal{Z}_{3m}$  subgroup should protect Diracness and Dark Matter stability
  - Exact subgroup fixed by the smallest B-L charge in model
  - If SM leptons have smallest charge then  $U(1)_{B-L} \rightarrow \mathcal{Z}_3$
  - Turns out  $\mathcal{Z}_3$  is too small **[C.Bonilla,E.Peinado,RS '19]**
    - Cannot insure DM stability on its own
- Break  $U(1)_{B-L} \rightarrow \mathcal{Z}_6$ 
  - Can be achieved if the particles running in loop carry half integral B-L charges



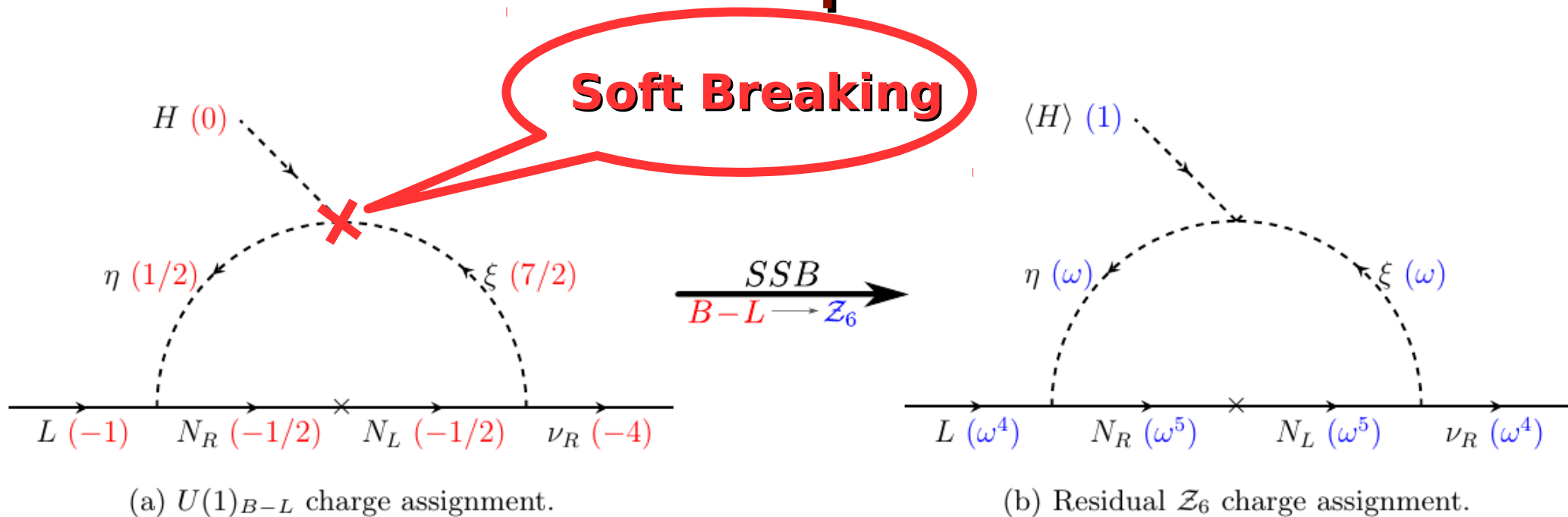
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- Here  $\omega = e^{2\pi I/6}$ ;  $\omega^6 = 1$  is the 6<sup>th</sup> root of unity.
- All particles carrying fractional B-L charges belong to Dark Sector
- **Lightest Dark Sector particle will be Stable Dark Matter Candidate**

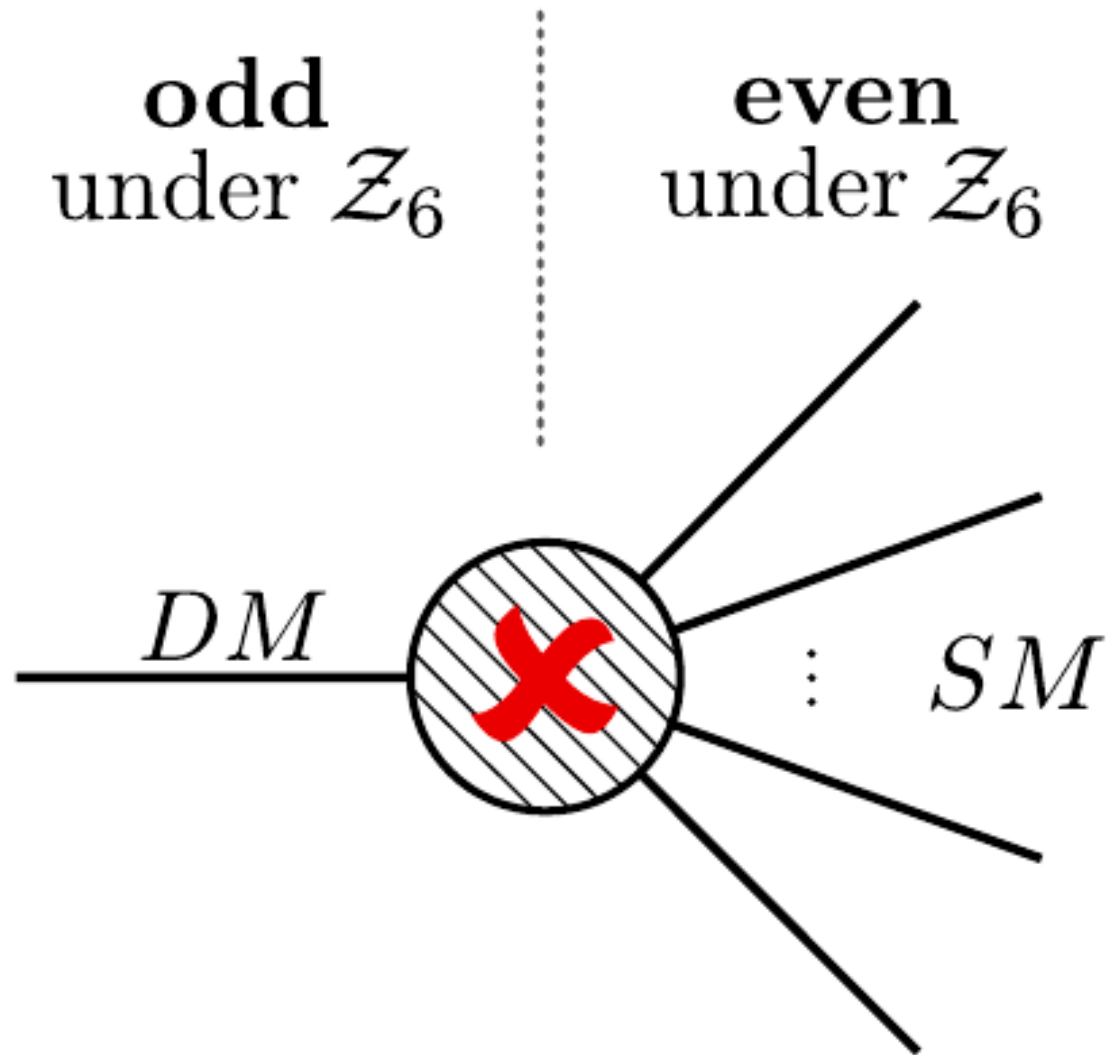
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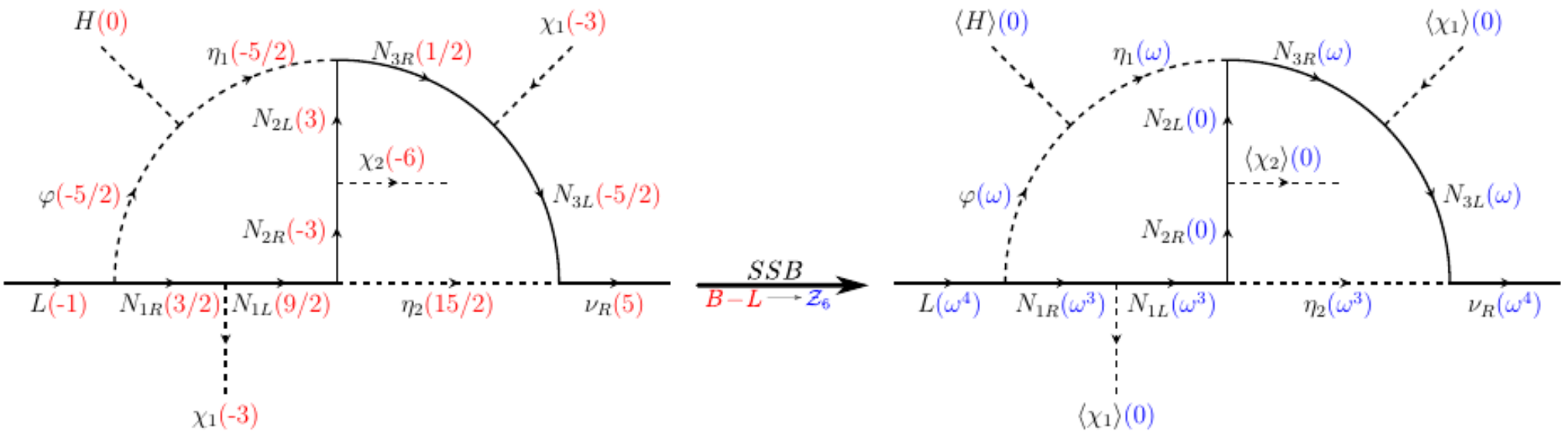
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# Two Loop Model



(a)  $U(1)_{B-L}$  charge assignment.

(b) Remnant  $Z_6$  charge assignment.

**[C.Bonilla,S.C.Chulia,R.Cepedello,E.Peinado,RS; Coming Soon]**

# General Two Loop Model

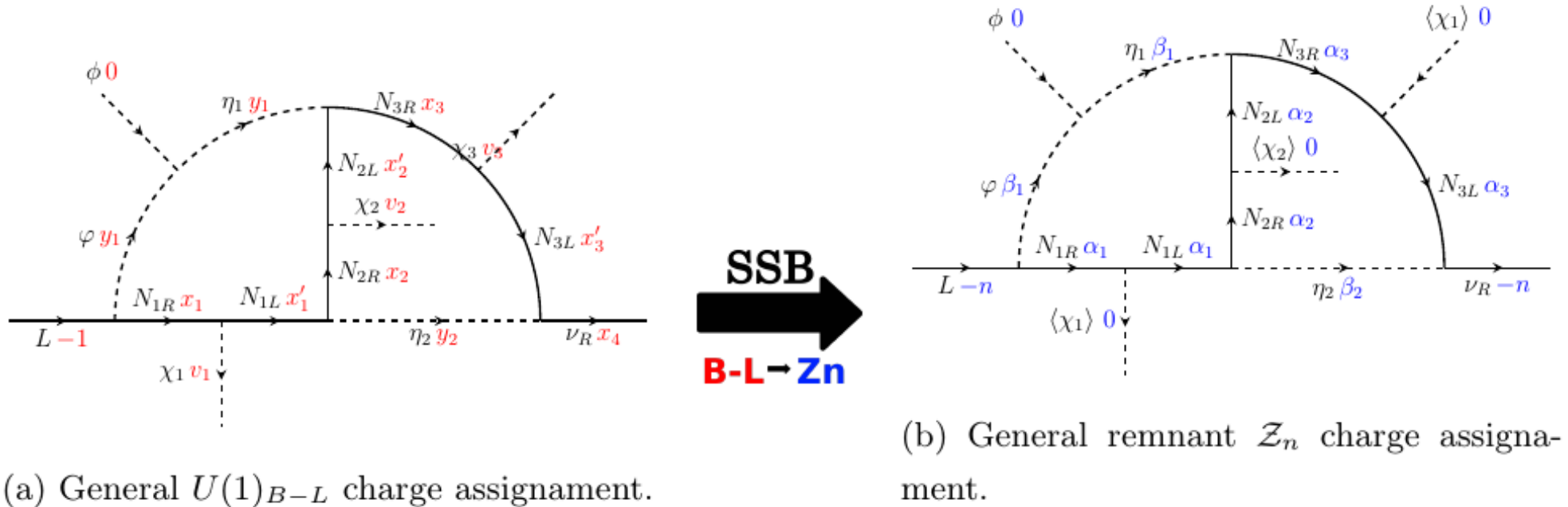


Figure 1: General charge assignment for a given topology and its spontaneous symmetry breaking

**[C.Bonilla,S.C.Chulia,R.Cepedello,E.Peinado,RS; Coming Soon]**

# Completely General N Loop Formalism

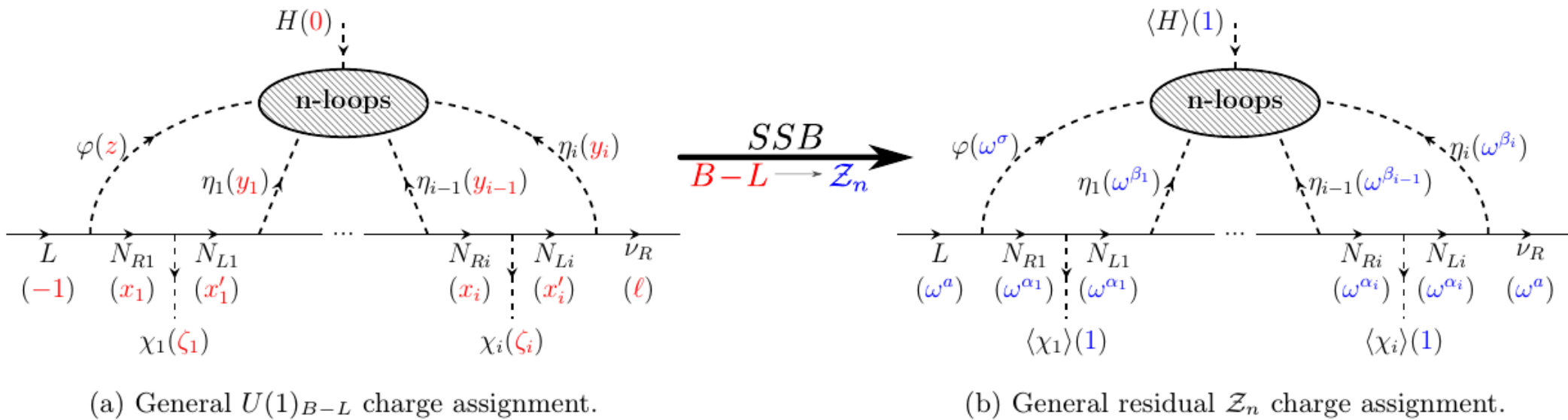
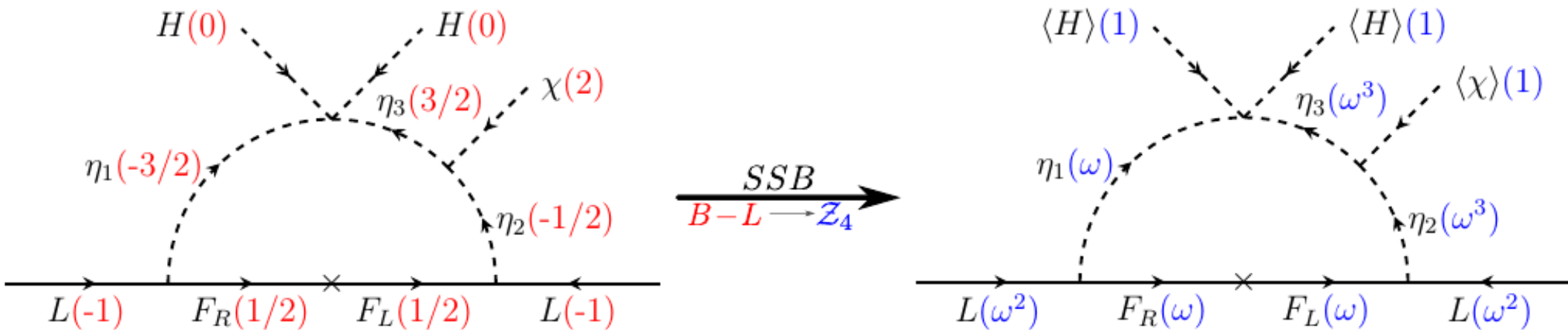


Figure 1: General charge assignment for any topology and its spontaneous symmetry breaking pattern.

**[C.Bonilla,S.C.Chulia,R.Cepedello,E.Peinado,RS '18]**

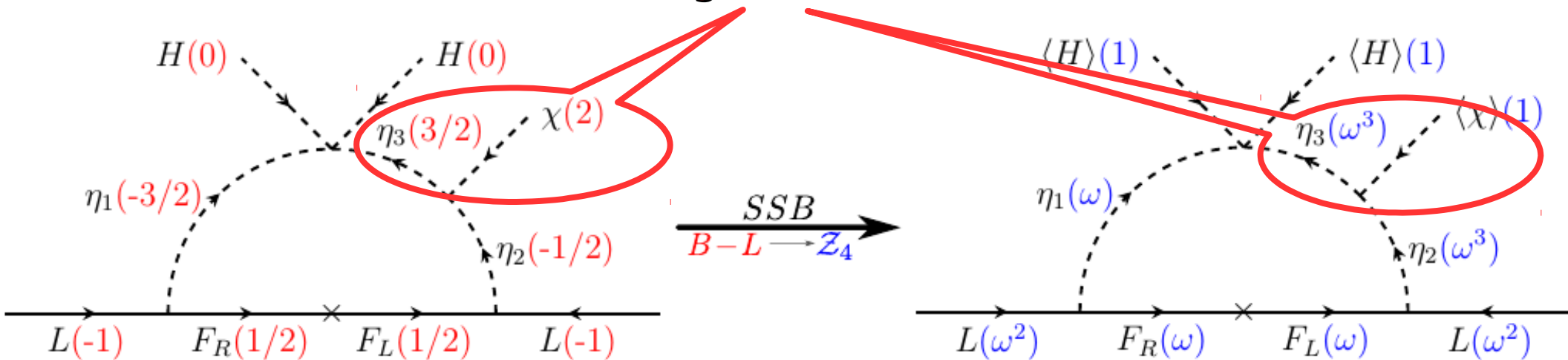
# B-L Scotogenic for Majorana Neutrinos



[S.C.Chulia,R.Cepedello,E.Peinado,RS '19]

# B-L Scotogenic for Majorana Neutrinos

Hard Breaking: Remove these fields



[S.C.Chulia,R.Cepedello,E.Peinado,RS '19]



# Conclusions

- **Nature of Neutrinos and Dark matter are two of the most important open questions**
- **We definitely need additional particles beyond those in SM to account for Dark Matter as well as mass of neutrinos**
- **However, I hope I convinced you that the symmetries present in SM are enough to**
  - Account for Dark Matter stability
  - Protect Diracness of neutrinos
  - Explain the smallness of neutrino mass
- **The Dirac nature of neutrinos and Dark Matter Stability are intimately related**
  - Guaranteed by the same Residual Subgroup of B-L

# Conclusions

- **The relation between Diracness and Dark Matter Stability is even deeper**
  - Also holds true for Dirac Seesaw Mechanisms  
[S.C.Chulia,E. Ma, RS, J.W.F.Valle '16] [SCC, RS, JWFV, '17,'18, '19]
  - The relation actually holds independent of the mass generation mechanism for Dirac neutrinos  
[S.C.Chulia,RS, J.W.F.Valle '18]
- **For certain special cases, the formalism discussed here can also be adopted for Majorana neutrinos**  
[S.C.Chulia,R.Cepedello,E.Peinado,RS '19]
  - Leads to a Scotogenic like mechanism

**Thank You**