

# Neutrino Masses and Modular Invariance

Hefei

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$$\tau \rightarrow \frac{a\tau + b}{c\tau + d}$$

## Flasy 2019

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Based on: - F.F. 1706.08749

- Juan Carlos Criado, F.F., 1807.01125
- Juan Carlos Criado, F.F., S.J.D King, in preparation

# Precision Era for Neutrino Physics

	IO	NO
$r \equiv  \Delta m_{sol}^2 / \Delta m_{atm}^2 $	0.0301(8)	0.0299(8)
$\sin^2 \theta_{12}$	0.303(13)	0.304(13)
$\sin^2 \theta_{13}$	0.0218(8)	0.0214(8)
$\sin^2 \theta_{23}$	0.56(3)	0.55(3)
$\delta/\pi$	1.52(14)	1.32(19)

independent global fits: de Salas, Gariazzo, Mena, Ternes , Tortola, 1806.11051, Gariazzo, Archidiacono, de Salas, Mena, Ternes, Tortola, 1801.04946 de Salas, Forero, Ternes, Tortola, J. W. F. Valle, 1708.01186 Esteban, Gonzalez-Garcia, Hernandez -Cabezudo, Maltoni and Schwetz 1811.05487
NO preferred over the IO

[F. Capozzi, E. Lisi, A. Marrone and A. Palazzo 1804.09678]

stimulating time for  
for models of neutrino masses  
and mixing angles.

$y_e(m_Z)$	$2.794745(16) \times 10^{-6}$
$y_\mu(m_Z)$	$5.899863(19) \times 10^{-4}$
$y_\tau(m_Z)$	$1.002950(91) \times 10^{-2}$

[Antusch and Maurer 1306.6879]

# Flavour Symmetry approach

One of the few tools we have, but with several obstacles

high number of free parameters

reviews:

Ishimori, Kobayashi, Ohki, Shimizu, Okada, Tanimoto , 1003.3552;

King, Luhn, 1301.1340;

King, Merle, Morisi, Shimizu, Tanimoto, 1402.4271;

King, 1701.04413

Hagedorn, 1705.00684;

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lowest order  
Lagrangian  
parameters

higher  
dimensional  
operators

$$m_{ij}(\tau) = m_{ij}^0 + m_{ij}^{1\alpha} \tau_\alpha + m_{ij}^{1\bar{\alpha}} \bar{\tau}_{\bar{\alpha}} + m_{ij}^{2\alpha\beta} \tau_\alpha \tau_\beta + \dots$$

vacuum alignment  
in SB sector

SUSY breaking effects  
RGE corrections  
( $\Lambda_{UV}, m_{SUSY}, \tan\beta$ )

# This proposal [F.F. 1706.08749]

a) neutrino masses and mixings  
depend on a small number of  
fields  
[ideally a single complex field  $\tau$ ]



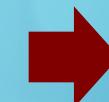
$$m_{ij}(\tau)$$

b) dependence of  $m_{ij}$  on  $\tau$  is holomorphic



supersymmetric  
model

c) flavour symmetry acts non-linearly  
[to determine all higher dimensional  
operators ]



$$\begin{cases} \tau \rightarrow F(\tau) \\ \varphi \rightarrow G(\tau, \varphi) \end{cases}$$

non-linear

a) + b) + c)



the functional form of  $m_{ij}(\tau)$  is  
completely determined

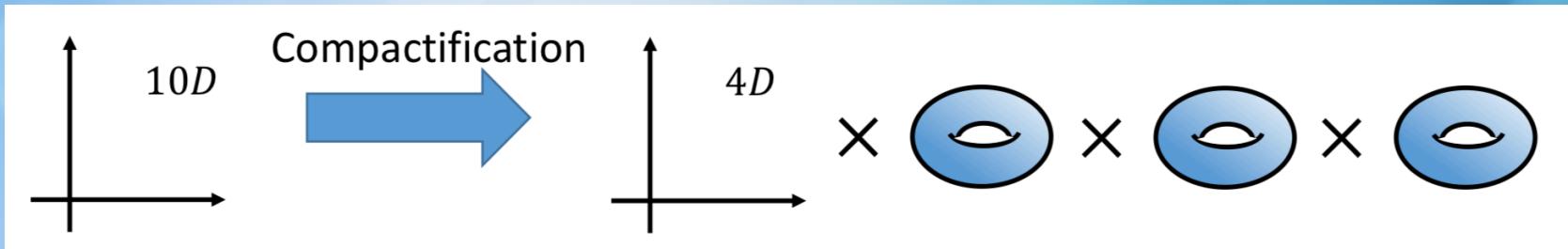
d)

the VEV  $\tau$  is selected by some unknown mechanism  
[anarchy in vacuum selection]

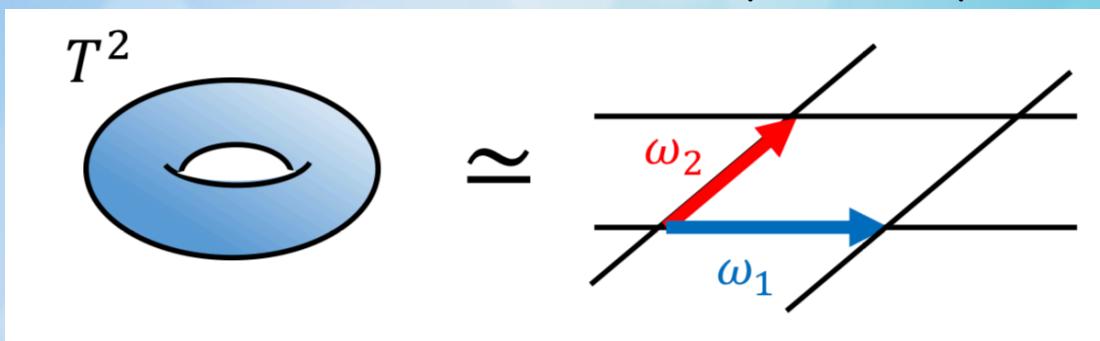
Here: a) + b) + c) from **modular invariance** as flavour symmetry

# Modular Invariance as Flavour Symmetry

string theory in d=10 need 6 compact dimensions



simplest compactification: 3 copies of a torus  $T^2$



completely  
characterized by

$$\tau = \frac{\omega_2}{\omega_1} \quad Im(\tau) > 0$$

$$\mathcal{L}_{eff} = \mathcal{L}_{eff}(g(\tau), \varphi)$$

lattice left invariant by modular transformations:

$$\tau \rightarrow \gamma\tau = \frac{a\tau + b}{c\tau + d}$$

a,b,c,d integers  
 $ad-bc=1$



$\mathcal{L}_{eff}$  modular invariant

they form the (discrete, infinite) modular group  $\bar{\Gamma}$  generated by

$$S : \tau \rightarrow -\frac{1}{\tau} ,$$

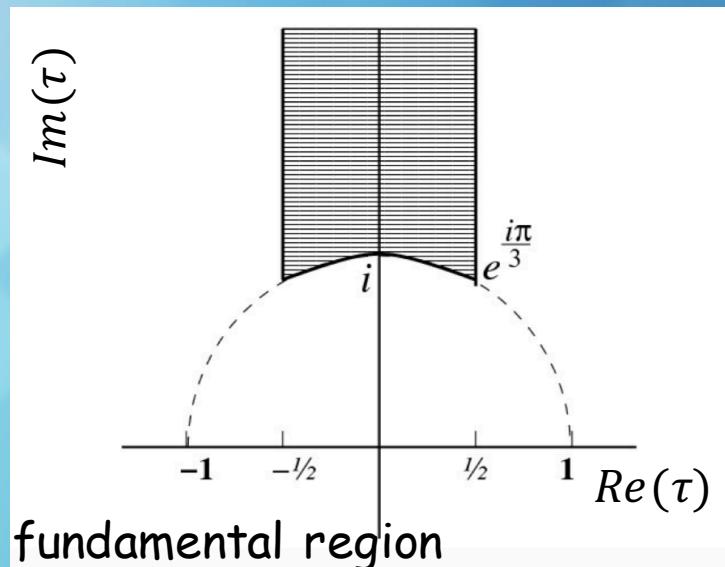
duality

$$T : \tau \rightarrow \tau + 1$$

discrete shift symmetry

$$S^2 = \mathbf{1} , \quad (ST)^3 = \mathbf{1}$$

- can be thought of as a gauge symmetry
- with a "gauge choice"  $\tau$  can be restricted to a fundamental region



most general transformation on a set of  $\mathcal{N}=1$  SUSY chiral multiplets  $\varphi^{(I)}$

$$\begin{cases} \tau \rightarrow \gamma\tau \equiv \frac{a\tau + b}{c\tau + d} \\ \varphi^{(I)} \rightarrow (c\tau + d)^{k_I} \rho^{(I)}(\gamma) \varphi^{(I)} \end{cases}$$

the weight,  
a real number

unitary representation  
of the finite modular group

e.g.

$$\varphi^{(I)} = \begin{pmatrix} e \\ \mu \\ \tau \end{pmatrix}$$

[Ferrara, Lust, Shapere and Theisen, 1989]

$$\Gamma_N \equiv \bar{\Gamma}/\bar{\Gamma}(N)$$

$N = 1, 2, 3, \dots$

# $\mathcal{N}=1$ SUSY modular invariant theories

Yukawa interactions in  $\mathcal{N}=1$  global SUSY [extension to  $\mathcal{N}=1$  SUGRA straightforward]

$$S = \int d^4x d^2\theta w(\tau, \varphi) + h.c + \text{kinetic terms}$$

invariance  
satisfied by  
"minimal"  
Kahler potential

$$w(\tau, \varphi) = \sum_n Y_{I_1 \dots I_n}(\tau) \varphi^{(I_1)} \dots \varphi^{(I_n)}$$

field-dependent  
Yukawa couplings

invariance of  $w(\Phi)$  guaranteed by an holomorphic  $Y_{I_1 \dots I_n}(\tau)$  such that

$$Y_{I_1 \dots I_n}(\gamma\tau) = (c\tau + d)^{k_Y(n)} \rho(\gamma) Y_{I_1 \dots I_n}(\tau)$$

1.  $k_Y(n) + k_{I_1} + \dots + k_{I_n} = 0$

2.  $\rho \times \rho^{I_1} \times \dots \times \rho^{I_n} \supset 1$

modular forms  
of level N and weight  $k_Y$



form a linear space  $\mathcal{M}_k(\Gamma_N)$   
of finite dimension

# Example

$$\Gamma_3 \approx A_4$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} \rightarrow (c\tau + d)^{-1} \rho(\gamma) \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

$k_\nu = -1$

$\sim 3$  of  $\Gamma_3$

$$w(\tau, \nu) = m_0 \nu Y(\tau) \nu + h.c.$$

modular form of level 3  
 $k = +2$  and  $\rho \subset 3 + 1 + 1' + 1''$

$$\begin{aligned} d(\mathcal{M}_2(\Gamma_3)) &= 3 \\ \rho &= 3 \end{aligned}$$

$$m(\tau) = m_0 \begin{pmatrix} 2Y_1(\tau) & -Y_3(\tau) & -Y_2(\tau) \\ -Y_3(\tau) & 2Y_2(\tau) & -Y_1(\tau) \\ -Y_2(\tau) & -Y_1(\tau) & 2Y_3(\tau) \end{pmatrix}$$

[F.F. 1706.08749]

mass matrix completely determined in terms of  $\tau$   
 up to an overall constant

no corrections from higher order operators in the exact SUSY limit

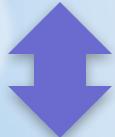
# Modular Invariance and CP

a unique CP law consistent  
with the modular group  
[ $\text{Im}(\tau) > 0$ ]

[Novichkov, Penedo, Petcov and Titov 1905.11970  
see also: Baur, Nilles, Trautner and Vaudrevange,  
1901.03251]

$$\tau \rightarrow -\tau^*$$

[up to modular transformations]



outer automorphism of  $\bar{\Gamma}$

$$S \rightarrow S \quad T \rightarrow T^{-1}$$

CP on matter multiplets

$$\varphi^{(I)} \rightarrow X_{CP} \varphi^{(I)*}$$

$X_{CP} = \mathbb{I}$  not restrictive if  
S and T symmetric matrices  
[canonical CP basis]

[in such a basis]  
CP invariance



$$g_i^* = g_i$$

on Lagrangian parameters  $g_i$

CP conserved  $\leftrightarrow \tau$  imaginary or  
at the border of the fundamental  
region

otherwise CP spontaneously broken  
by  $\langle \tau \rangle$

# Modular Forms

$k > 0$  even integer

	$d(\mathcal{M}_k(\Gamma_N))$	$k = 2$	$k = 4$	$k \geq 6$	
$\Gamma_2 \approx S_3$	$k/2 + 1$	2	1 + 2	...	[TTT]
$\Gamma_3 \approx A_4$	$k + 1$	3	1 + 1' + 3	...	[F]
$\Gamma_4 \approx S_4$	$2k + 1$	$2 + 3'$	$1 + 2 + 3 + 3'$	...	[PP]
$\Gamma_5 \approx A_5$	$5k + 1$	$3 + 3' + 5$	$1 + 3 + 3' + 4 + 5 + 5$	...	[NPPT DKL]

$$\left. \begin{array}{l} \Gamma_8 \supset \Delta(96) \\ \Gamma_{16} \supset \Delta(384) \end{array} \right\} k = 2 \quad \rho = 3 \quad [KT]$$

[TTT = T. Kobayashi, K. Tanaka and T. H. Tatsuishi,  
1803.10391

F = F. Feruglio 1706.08749

PP = J. T. Penedo and S. T. Petcov 1806.11040

NPPT = P. P. Novichkov, J. T. Penedo,  
S. T. Petcov and A. V. Titov 1812.02158

DKL = G. J. Ding, S. F. King and X. G. Liu 1903.12588

KT = T. Kobayashi and S. Tamba, 1811.11384]

built in terms of  
Dedekind eta function  
Klein forms  
Jacobi theta functions

$k > 0$  odd/even integer

fall in representations of homogeneous finite modular groups  $\Gamma'_N$   
e.g. N=3 and k=1 gives a doublet of  $\Gamma'_3 = T'$

[Ding, Liu 1907.01488]

# Selection of results

→ See talks by King, Novichkov, Penedo, Tanimoto, Titov, Zhou

freedom in a bottom-up approach:

$$\Gamma_N, \rho^{(I)}, k_I$$

Majorana neutrinos

[Dirac case explored in  
T. Kobayashi, N. Omoto, Y. Shimizu, K. Takagi,  
M. Tanimoto and T. H. Tatsuishi, 1808.03012]

weight  $k = 2$  in  $Y(\tau)$  of neutrino sector (too many parameters if  $k > 2$ )

number of free parameters  $p$ :

$$p = \text{Total} - 3 \text{ (charged fermion masses)}$$

number of observables in neutrino sector = 9  
(3 angles + 3 masses + 3 phases)



(9-p) predictions

$p \geq 3$  (always includes  $\text{Re}(\tau), \text{Im}(\tau)$  and one overall scale)  
Here  $p > 5$  not considered

# Example 1: $m_\nu(\tau)$ and $m_e(\tau)$

[P. P. Novichkov, J. T. Penedo, S. T. Petcov and A. V. Titov, 1905.11970]

$\Gamma_4$

no additional flavons beyond  $\tau$   
CP spontaneously broken by  $\langle \tau \rangle$

1 extra parameters  
from Dirac neutrino  
mass



$p = 4$

$\text{Re } \tau$	$\pm 0.09922$
$\text{Im } \tau$	1.016
$g'/g$	-0.02093
$v_u^2 g^2 / \Lambda$ [eV]	0.0135

$r$	0.02981
$\delta m^2$ [ $10^{-5}$ eV $^2$ ]	7.326
$ \Delta m^2 $ [ $10^{-3}$ eV $^2$ ]	2.457
$\sin^2 \theta_{12}$	0.305
$\sin^2 \theta_{13}$	0.02136
$\sin^2 \theta_{23}$	0.4862

$N\sigma$	1.012
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Ordering	NO
$m_1$ [eV]	0.01211
$m_2$ [eV]	0.01483
$m_3$ [eV]	0.05139
$\sum_i m_i$ [eV]	0.07833
$ \langle m \rangle $ [eV]	0.01201
$\delta/\pi$	$\pm 1.641$
$\alpha_{21}/\pi$	$\pm 0.3464$
$\alpha_{31}/\pi$	$\pm 1.254$

# Example 2:

$m_\nu(\tau) \text{ and } m_e(\varphi)$

[Juan Carlos Criado, F.F., 1807.01125]

$\Gamma_3$

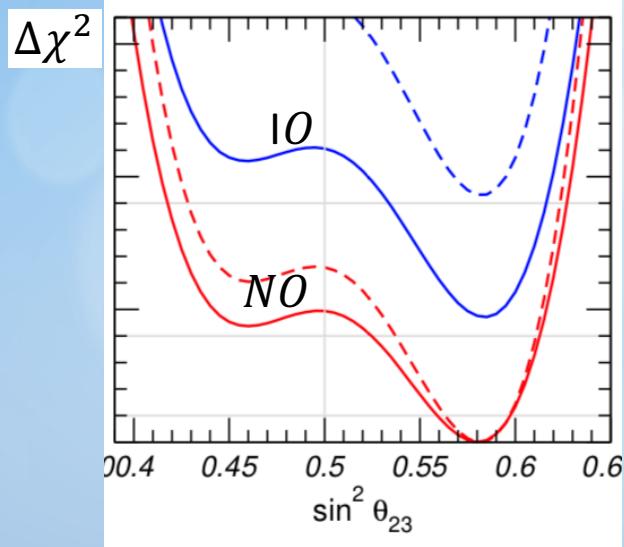
charged lepton sector depending on additional flavon

$\varphi = (1, 0, \varphi_3) \quad \varphi_3 \ll 1 \text{ real}$

CP spontaneously broken by  $\langle \tau \rangle$



$p = 4$



T. Schwetz @ Neutrino Telescopes, Venice, 19 March 2019

Seesaw $\rightarrow$ NO	best value
$r \equiv  \Delta m_{sol}^2 / \Delta m_{atm}^2 $	0.0299(12)
$m_3/m_2$	3.68(5)
$\sin^2 \theta_{12}$	0.306(11)
$\sin^2 \theta_{13}$	0.0211(12)
$\sin^2 \theta_{23}$	0.459(5)
$\delta/\pi$	1.438(8)
$\alpha_{21}/\pi$	1.704(5)
$\alpha_{31}/\pi$	1.201(16)

all dimensionless neutrino data determined in terms of 3 vacuum parameters

$Re(\tau), Im(\tau), \varphi_3$

$\tau$	$-0.2005 + i 1.0578$
$\varphi_3$	0.117



1<sup>st</sup> octant

$N\sigma = 2.5 \div 3$

$m_1 = 1.09(3) \times 10^{-2} \text{ eV} \quad ,$

$m_2 = 1.39(2) \times 10^{-2} \text{ eV} \quad ,$

$m_3 = 5.11(4) \times 10^{-2} \text{ eV}$

$|m_{ee}| = 1.04(2) \times 10^{-2} \text{ eV}$

# Other examples

$m_\nu(\tau)$  and  $m_e(\varphi)$

[Criado, F, S.J.D. King, in preparation]

$\Gamma_4$

$$\left[ \begin{pmatrix} 0 & Y_1 & -Y_2 \\ Y_1 & -Y_2 & 0 \\ -Y_2 & 0 & Y_1 \end{pmatrix} + \xi \begin{pmatrix} 2Y_3 & -Y_5 & -Y_4 \\ -Y_5 & 2Y_4 & -Y_3 \\ -Y_4 & -Y_3 & 2Y_5 \end{pmatrix} \right]$$

Type I seesaw  
CP-invariant

1 extra parameters  
in  $\nu^c$  mass



$p = 5$

	best value
$\Delta m_{21}^2 \cdot 10^5 \text{ eV}^{-2}$	7.39
$\Delta m_{3\ell}^2 \cdot 10^3 \text{ eV}^{-2}$	2.53
$\sin^2 \theta_{12}$	0.317
$\sin^2 \theta_{13}$	0.0224
$\sin^2 \theta_{23}$	0.580
$\delta/\pi$	1.25

$$\chi^2_{min} = 0.34$$

$$\tau = 2.506 + i 0.591$$

$$\xi = -2.595$$

$$\varphi_3 = i 0.108$$

	best value
$m_1 \cdot 10^2 \text{ eV}^{-1}$	4.26
$m_2 \cdot 10^2 \text{ eV}^{-1}$	4.35
$m_3 \cdot 10^2 \text{ eV}^{-1}$	6.59
$\alpha_{21}/\pi$	0.11
$\alpha_{31}/\pi$	0.30
$m_{ee} \cdot 10^2 \text{ eV}^{-1}$	4.25

$\Gamma_5$  $y_e$  diagonalno known solutions  
with  $p \leq 5$  and  $y_e(\tau)$ 

Weinberg

 $p = 3$  not viable

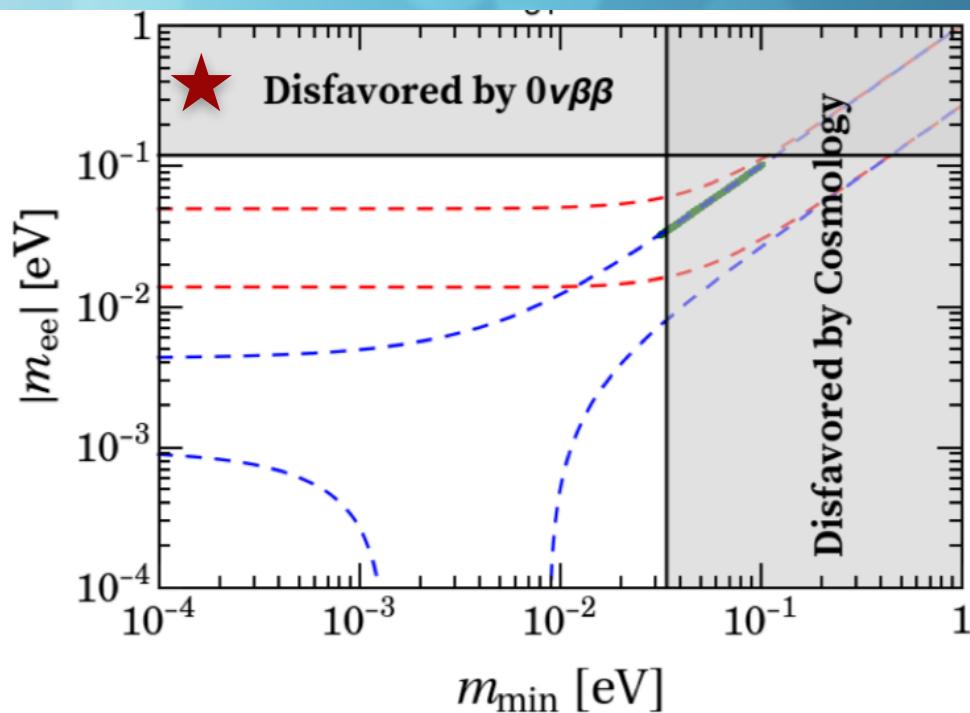
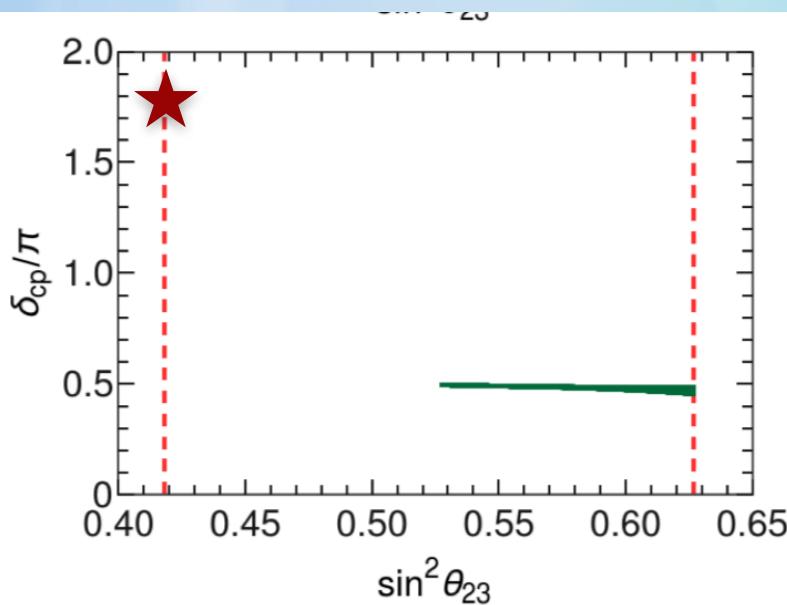
Seesaw

 $p = 5$ 

NO

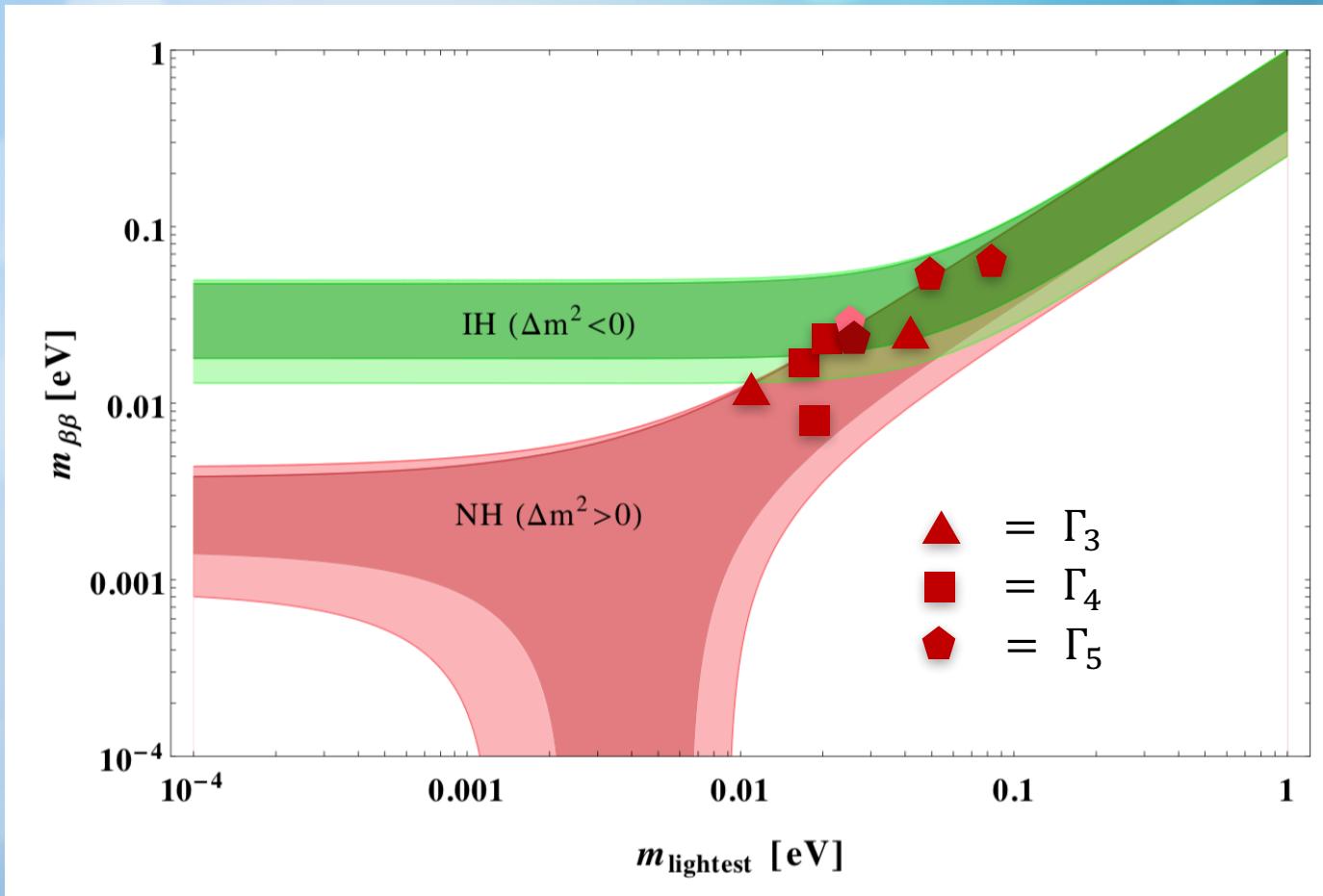
$m_1$	83 meV	49 meV	26 meV	26 meV
$m_2$	83 meV	50 meV	27 meV	27 meV
$m_3$	97 meV	70 meV	57 meV	57 meV
$ m_{ee} $	46 meV	50 meV	24 meV	23 meV
$\delta_{CP}$	$\pm 1.7 \pi$	$\pm 1.5 \pi$	$\pm 1.2 \pi$	$\pm 1.0 \pi$

[P. P. Novichkov, J. T. Penedo,  
S. T. Petcov and A. V. Titov 1812.02158  
G. J. Ding, S. F. King and X. G. Liu 1903.12588]



# Some trend

most of the solutions with NO prefer a nearly degenerate spectrum  
 $m_1 > 10 \text{ meV}$  and  $|m_{ee}|$  on the high side of allowed range



[ $0\nu\beta\beta$  region from: S. Dell'Oro, S. Marcocci and F. Vissani, 1404.2616]

# Some trend

most of the solutions with NO prefer a nearly degenerate spectrum  
 $m_1 > 10 \text{ meV}$  and  $|m_{ee}|$  on the high side of allowed range

higher N  $\rightarrow$  more solutions

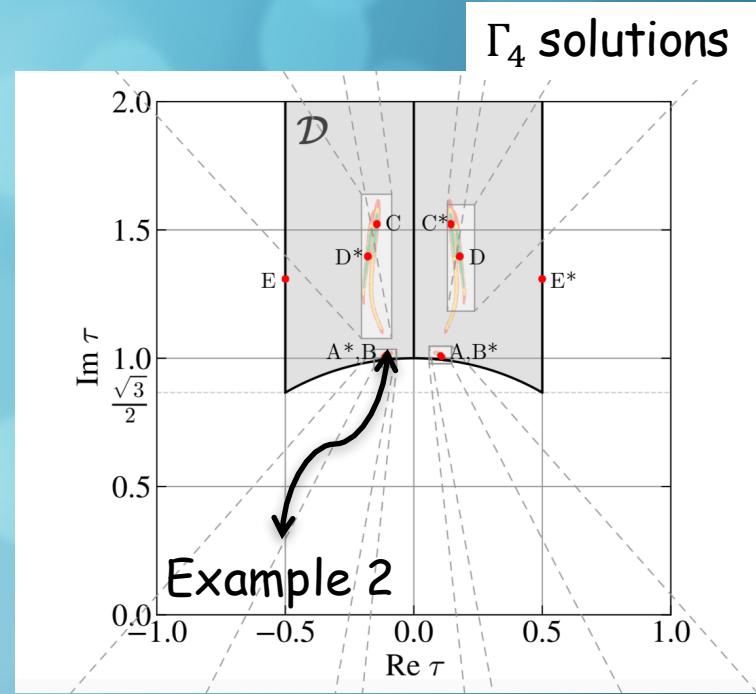
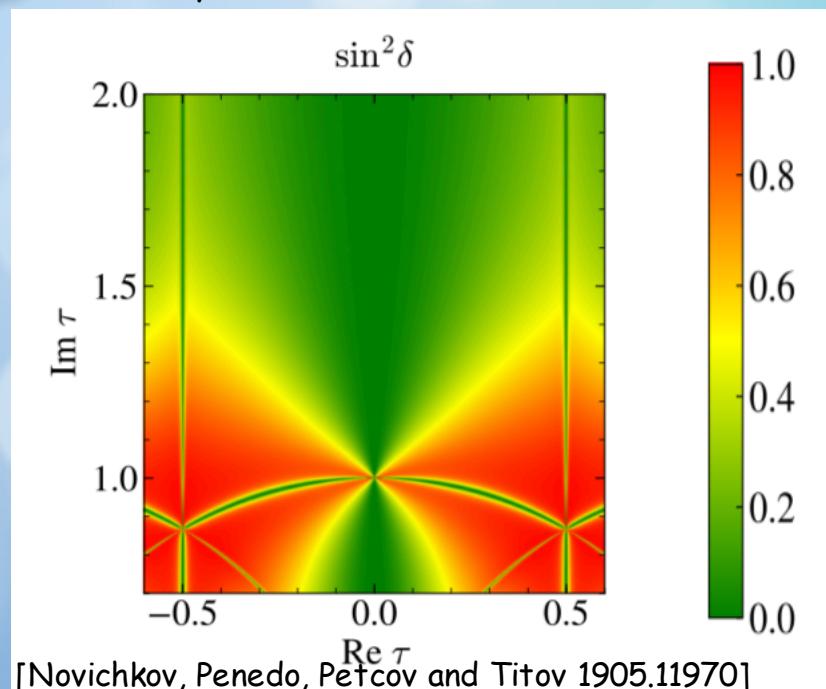
any preferred value of  $\tau$  ?

extrema of  $V(\tau)$  at the border of the fundamental region and along the  $\text{Im}(\tau)$  axis ?

[Cvetic, Font, Ibanez, Lust and Quevedo,  
*Nucl.Phys.B* 361 (1991) 194]

[P. P. Novichkov, J. T. Penedo,  
S. T. Petcov and A. V. Titov,  
1811.04933 and 1812.02158]

CP is conserved here



# Corrections from SUSY breaking

unknown breaking mechanism. Here:

F-component of a chiral supermultiplet, gauge and modular invariant

$$X = \vartheta^2 F$$

messenger scale  $M$

SUSY-breaking scale

$$m_{SUSY} = \frac{F}{M}$$

most general correction term to lepton masses and mixing angles

$$\delta\mathcal{S} = \frac{1}{M^2} \int d^4x d^2\theta d^2\bar{\theta} X^\dagger f(\Phi, \bar{\Phi}) + h.c.$$

$f(\Phi, \bar{\Phi})$  has dimension 3, determined by gauge invariance and lepton number conservation (treating  $\Lambda$  as spurion with  $L=+2$ )



$$\delta\mathcal{W}/\mathcal{W} \approx \delta\mathcal{Y}/\mathcal{Y} \approx \frac{m_{SUSY}}{M}$$

tiny, if sufficient gap between  $m_{SUSY}$  and  $M$

$10^{-10}$  for

$$m_{SUSY} = 10^8 \text{ GeV}$$
$$M = 10^{18} \text{ GeV}$$

# Conclusions

neutrino data set a new standard in model building

- accurate predictions
- falsifiable models

modular invariance as flavour symmetry can determine the functional dependence of Yukawa couplings on a modulus field

$y(\tau)$

- no/less flavons,
- less parameters
- no corrections from higher dimensional operators
- stability against SUSY breaking corrections

can be implemented in a bottom-up approach:

absolute masses,  $m_{ee}$  and phases are predicted

new directions:

- + role of modular symmetry in quark sector [Kobayashi, Shimizu, Takagi, Tanimoto, Tatsuishi and Uchida, 1812.11072  
Okada and Tanimoto, 1812.09677]
- + in GUTs [de Anda, King and Perdomo, 1812.05620]
- + more moduli [de Medeiros Varzielas, King and Zhou, 1906.02208]

## Backup Slides

# Modular forms of level 3 [1706.08749]

dimension of linear space  $\mathcal{M}_k(\Gamma(3))$  is  $(k+1)$ ,  $k > 0$  even integer

3 linearly independent modular forms of level 3 and minimal weight  $k_I = 2$

$$\begin{aligned} Y_1(\tau) &= \frac{i}{2\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} - \frac{27\eta'(3\tau)}{\eta(3\tau)} \right] \\ Y_2(\tau) &= \frac{-i}{\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \omega^2 \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \omega \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} \right] \\ Y_3(\tau) &= \frac{-i}{\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \omega \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \omega^2 \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} \right]. \end{aligned}$$

can be expressed in terms of the Dedekind eta function

$$\eta(\tau) = q^{1/24} \prod_{n=1}^{\infty} (1 - q^n) \quad q \equiv e^{i2\pi\tau}$$

they transform in a triplet 3 of  $\Gamma_3$

$$Y(-1/\tau) = \tau^2 \rho(S) Y(\tau)$$

$$\rho(S) = \frac{1}{3} \begin{pmatrix} -1 & 2 & 2 \\ 2 & -1 & 2 \\ 2 & 2 & -1 \end{pmatrix}$$

$$Y(\tau + 1) = \rho(T) Y(\tau)$$

$$\rho(T) = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \omega & 0 \\ 0 & 0 & \omega^2 \end{pmatrix}$$

$$\omega = -\frac{1}{2} + \frac{\sqrt{3}}{2}i$$

they generate the whole ring  $\mathcal{M}(\Gamma(3))$

any modular form of level 3 and weight  $2k$  can be written as an homogeneous polynomial in  $Y_i$  of degree  $k$

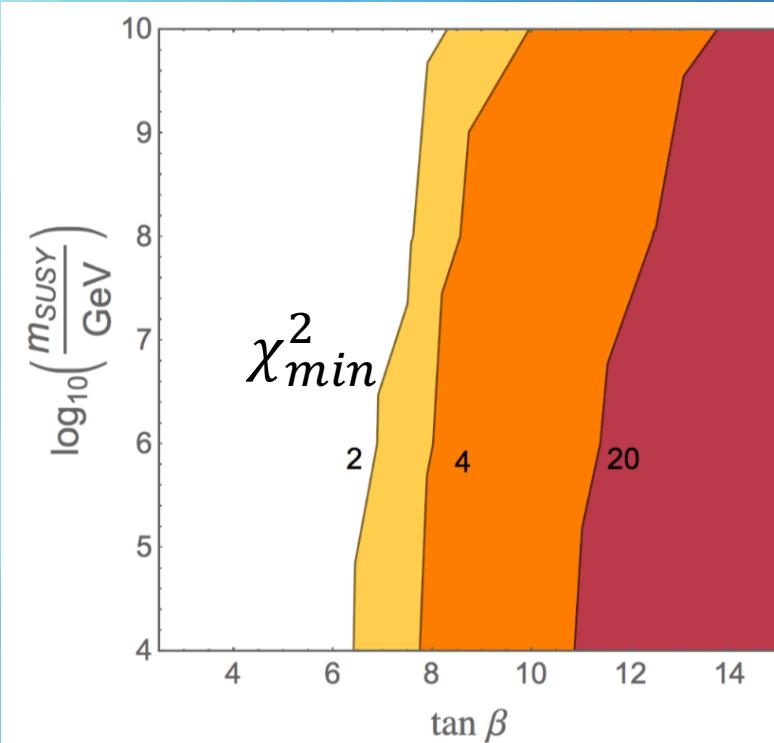
# Corrections from RGE

## Model 1 (IO)

- $r$  and  $\sin^2 \theta_{12}$  mostly affected, at large  $\tan\beta$

$\Lambda = 10^{15} \text{ GeV}$

$m_{\text{SUSY}}$	Quantity	$\tan\beta = 2.5$	$\tan\beta = 10$	$\tan\beta = 15$
$10^4 \text{ GeV}$	$r$	0.0302	0.0292	0.0288
	$\sin^2 \theta_{12}$	0.304	0.345	0.418
	$\chi^2_{\min}$	0.4	12.2	82.0
$10^8 \text{ GeV}$	$r$	0.0302	0.0294	0.0286
	$\sin^2 \theta_{12}$	0.303	0.335	0.389
	$\chi^2_{\min}$	0.4	7.0	47.7



## Model 2 (NO)

negligible corrections for  $\tan\beta$  up to 25 and  $m_{\text{SUSY}}$  as low as  $10^4 \text{ GeV}$

# $\mathcal{N}=1$ SUSY modular invariant theories

known since late 1980s

S. Ferrara, D. Lust, A. D. Shapere and S. Theisen, Phys. Lett. B **225** (1989) 363.

S. Ferrara, D. Lust and S. Theisen, Phys. Lett. B **233** (1989) 147.

focus on Yukawa interactions and  $\mathcal{N}=1$  global SUSY

$$\mathcal{S} = \int d^4x d^2\theta d^2\bar{\theta} K(\Phi, \bar{\Phi}) + \int d^4x d^2\theta w(\Phi) + h.c.$$

$$\Phi = (\tau, \varphi)$$

Kahler potential,  
kinetic terms

superpotential, holomorphic function of  $\Phi$   
Yukawa interactions

$\mathcal{S}$  invariant if

$$\begin{cases} w(\Phi) \rightarrow w(\Phi) \\ K(\Phi, \bar{\Phi}) \rightarrow K(\Phi, \bar{\Phi}) + f(\Phi) + f(\bar{\Phi}) \end{cases}$$

invariance of the Kahler potential easy to achieve. For example:

$$K(\Phi, \bar{\Phi}) = -h \log(-i\tau + i\bar{\tau}) + \sum_I (-i\tau + i\bar{\tau})^{-k_I} |\varphi^{(I)}|^2$$

minimal K

extension to  $\mathcal{N}=1$  SUGRA straightforward: ask invariance of  $G=K+\log|w|^2$

# Few facts about (level-N) Modular Forms

■ transformation property under the modular group

$$f_i(\gamma\tau) = (c\tau + d)^k \rho(\gamma)_{ij} f_j(\tau)$$

unitary representation of the finite modular group  $\Gamma_N \equiv \bar{\Gamma}/\bar{\Gamma}(N)$

■  $q$ -expansion

$$f(\tau + N) = f(\tau)$$

$$f(\tau) = \sum_{n=0}^{\infty} a_n q_N^n \quad q_N = e^{\frac{i2\pi\tau}{N}}$$

$k < 0$



$$f(\tau) = 0$$

$k = 0$



$$f(\tau) = \text{constant}$$

$k > 0$  (even integer)



$$f(\tau) \in \mathcal{M}_k(\Gamma(N))$$

finite-dimensional linear space

■ ring of modular forms generated by few elements

$$\mathcal{M}(\Gamma(N)) = \bigoplus_{k=0}^{\infty} \mathcal{M}_{2k}(\Gamma(N))$$

an explicit example in a moment

# Level N modular forms

$N$	$g$	$d_{2k}(\Gamma(N))$	$\mu_N$	$\Gamma_N$
2	0	$k + 1$	6	$S_3$
3	0	$2k + 1$	12	$A_4$
4	0	$4k + 1$	24	$S_4$
5	0	$10k + 1$	60	$A_5$
6	1	$12k$	72	
7	3	$28k - 2$	168	

Table 1: Some properties of modular forms:  $g$  is the genus of the space  $\mathcal{H}/\Gamma(N)$  after compactification,  $d_{2k}(\Gamma(N))$  the dimension of the linear space  $\mathcal{M}_{2k}(\Gamma(N))$ ,  $\mu_N$  is the dimension of the quotient group  $\Gamma_N \equiv \overline{\Gamma}/\overline{\Gamma}(N)$ , which, for  $N \leq 5$ , is isomorphic to a permutation group.

# Modular forms of level 3 [1706.08749]

dimension of linear space  $\mathcal{M}_k(\Gamma(3))$  is  $(k+1)$ ,  $k > 0$  even integer

3 linearly independent modular forms of level 3 and minimal weight  $k_I = 2$

$$\begin{aligned} Y_1(\tau) &= \frac{i}{2\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} - \frac{27\eta'(3\tau)}{\eta(3\tau)} \right] \\ Y_2(\tau) &= \frac{-i}{\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \omega^2 \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \omega \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} \right] \\ Y_3(\tau) &= \frac{-i}{\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \omega \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \omega^2 \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} \right]. \end{aligned}$$

can be expressed in terms of the Dedekind eta function

$$\eta(\tau) = q^{1/24} \prod_{n=1}^{\infty} (1 - q^n) \quad q \equiv e^{i2\pi\tau}$$

they transform in a triplet 3 of  $\Gamma_3$

$$Y(-1/\tau) = \tau^2 \rho(S) Y(\tau)$$

$$\rho(S) = \frac{1}{3} \begin{pmatrix} -1 & 2 & 2 \\ 2 & -1 & 2 \\ 2 & 2 & -1 \end{pmatrix}$$

$$Y(\tau + 1) = \rho(T) Y(\tau)$$

$$\rho(T) = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \omega & 0 \\ 0 & 0 & \omega^2 \end{pmatrix}$$

$$\omega = -\frac{1}{2} + \frac{\sqrt{3}}{2}i$$

they generate the whole ring  $\mathcal{M}(\Gamma(3))$

any modular form of level 3 and weight  $2k$  can be written as an homogeneous polynomial in  $Y_i$  of degree  $k$

# Fundamental domain of $\Gamma(3)$

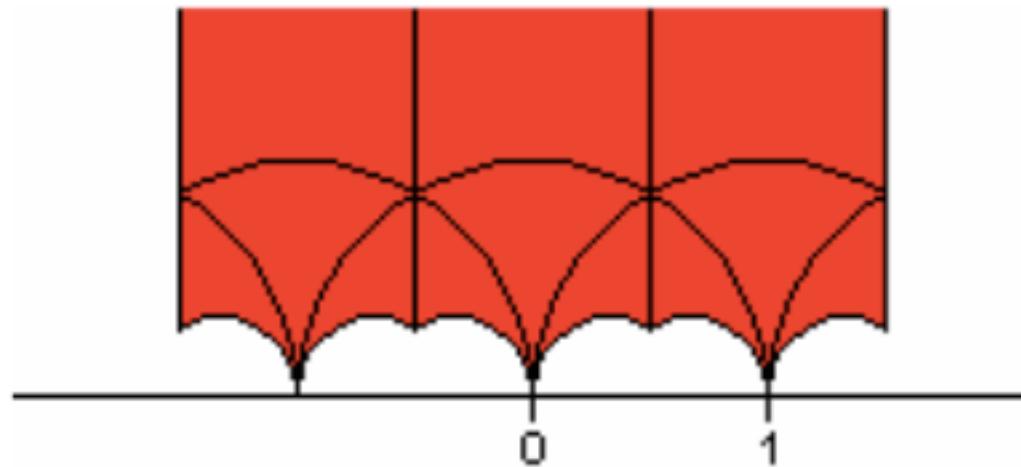


Figure 1: Fundamental domain for  $\Gamma(3)$ .

# *Q-expansion*

$$Y_1(\tau) = 1 + 12q + 36q^2 + 12q^3 + \dots$$

$$Y_2(\tau) = -6q^{1/3}(1 + 7q + 8q^2 + \dots)$$

$$Y_3(\tau) = -18q^{2/3}(1 + 2q + 5q^2 + \dots) \quad .$$

$$Y_2^2 + 2Y_1Y_3 = 0$$

*some VEVs*

$$(Y_1, Y_2, Y_3)|_{\tau=i\infty} = (1, 0, 0)$$

$$(Y_1, Y_2, Y_3)|_{\tau=i} = Y_1(i)(1, 1 - \sqrt{3}, -2 + \sqrt{3})$$

$$Y(-1/\tau)|_{\tau=i} = -\rho(S)Y(\tau)|_{\tau=i}$$

# Ring of level-3 modular forms

$$Y_2^2 + 2Y_1Y_3 = 0$$

As discussed explicitly in Appendix D, the constraint (30) is essential to recover the correct dimension of the linear space  $\mathcal{M}_{2k}(\Gamma(3))$ . On the one side from table 1 we see that this space has dimension  $2k + 1$ . On the other hand the number of independent homogeneous polynomial  $Y_{i_1}Y_{i_2} \cdots Y_{i_k}$  of degree  $k$  that we can form with  $Y_i$  is  $(k+1)(k+2)/2$ . These polynomials are modular forms of weight  $2k$  and, to match the correct dimension,  $k(k-1)/2$  among them should vanish. Indeed this happens as a consequence of eq. (30). Therefore the ring  $\mathcal{M}(\Gamma(3))$  is generated by the modular forms  $Y_i(\tau)$  ( $i = 1, 2, 3$ ).

# Models 1 and 2 are based on $\Gamma_3$

Why  $\Gamma_3$ ?  $\Gamma_3$  is isomorphic to  $A_4$ , smallest group of the  $\Gamma_N$  series possessing a 3-dimensional irreducible representation

[Ma, Rajasekaran, 0106291  
Babu, Ma, Valle 0206292]

[recent extensions to  $\Gamma_2$  and  $\Gamma_4$  in Kobayashi, Tanaka, Tatsuishi, 1803.10391;  
Penedo, Petcov 1806.11040]

	$(E_1^c, E_2^c, E_3^c)$	$N^c$	$L$	$H_d$	$H_u$	$\varphi$
$SU(2)_L \times U(1)_Y$	$(1, +1)$	$(1, 0)$	$(2, -1/2)$	$(2, -1/2)$	$(2, +1/2)$	$(1, 0)$
$\Gamma_3 \equiv A_4$	$(1, 1'', 1')$	3	3	1	1	3
$k_I$	$(k_{E_1}, k_{E_2}, k_{E_3})$	$k_N$	$k_L$	$k_d$	$k_u$	$k_\varphi$

**Table 1:** Chiral supermultiplets, transformation properties and weights. Model 1 has no gauge singlets  $N^c$ .

	$k_{E_i}$	$k_N$	$k_L$	$k_d$	$k_u$	$k_\varphi$
Model 1	-2	-	-1	0	0	+3
Model 2	-4	-1	+1	0	0	+3

modular invariance  
broken by

$\tau$

$\varphi = (1, 0, \varphi_3)$

real

**Table 2:** Weights of chiral multiplets. Model 1 has no gauge singlets  $N^c$ .

if we go minimal

we get

	$L$	$H_u$	$Y$
$SU(2) \times U(1)$	(2, -1/2)	(2, +1/2)	(1, 0)
$\Gamma_3 \equiv A_4$	3	1	3
$k_I$	+1	0	+2

$$m_\nu = \begin{pmatrix} 2Y_1 & -Y_3 & -Y_2 \\ -Y_3 & 2Y_2 & -Y_1 \\ -Y_2 & -Y_1 & 2Y_3 \end{pmatrix} \frac{v_u^2}{\Lambda}$$

by scanning  $\tau$  VEVs the best agreement is obtained for

$$\tau = 0.0111 + 0.9946i$$

	$\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}$	$\sin^2 \vartheta_{12}$	$\sin^2 \vartheta_{13}$	$\sin^2 \vartheta_{23}$	$\frac{\delta_{CP}}{\pi}$	$\frac{\alpha_{21}}{\pi}$	$\frac{\alpha_{31}}{\pi}$
<i>Exp</i>	0.0292	0.297	0.0215	0.5	1.4	–	–
$1\sigma$	0.0008	0.017	0.0007	0.1	0.2	–	–
<i>prediction</i>	0.0292	0.295	0.0447	0.651	1.55	0.22	1.80

many  
 $\sigma$  away

2-parameter fit to 5 physical quantities

the operator

$$w_\nu = \frac{1}{\Lambda} (H_u H_u \ LL \ Y)_1$$

is completely specified up to an overall constant

a familiar matrix but now  $Y_i$  are determined by the choice of  $\tau$

8 dimensionless physical quantities independent on any coupling constant!

# Variants

neutrino masses from see-saw mechanism

$$w_\nu = g (N^c H_u L)_1 + \Lambda (N^c N^c Y)_1$$

assignement

	$L$	$N^c$	$H_u$	$Y$
$SU(2) \times U(1)$	(2, -1/2)	(1, 0)	(2, +1/2)	(1, 0)
$\Gamma_3 \equiv A_4$	3	3	1	3
$k_I$	$k_L$	+1	$k_u$	+2

$$1+k_L+k_u=0$$

we get the best agreement at

$$\tau = -0.195 + 1.0636i$$

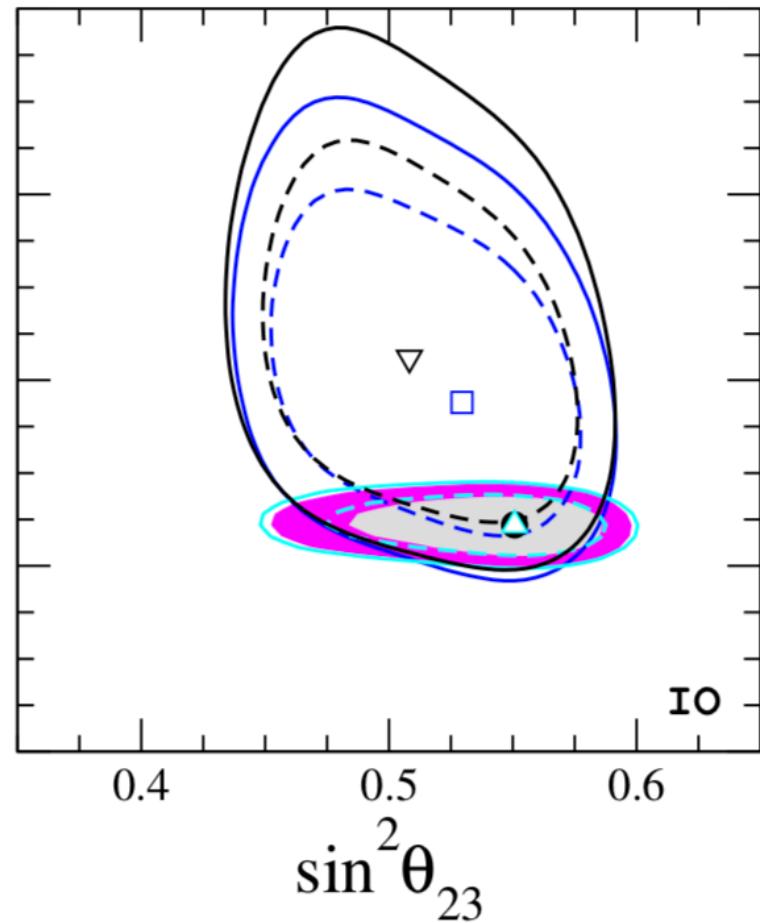
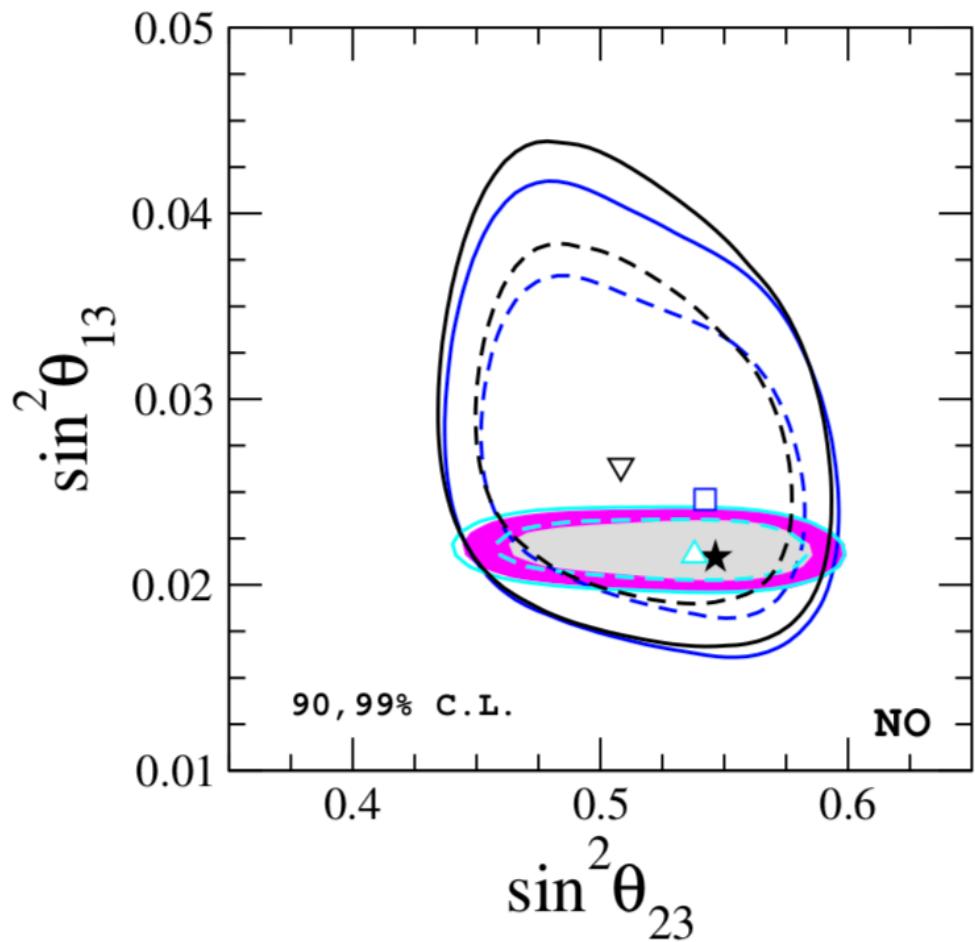
	$\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}$	$\sin^2 \vartheta_{12}$	$\sin^2 \vartheta_{13}$	$\sin^2 \vartheta_{23}$	$\frac{\delta_{CP}}{\pi}$	$\frac{\alpha_{21}}{\pi}$	$\frac{\alpha_{31}}{\pi}$
<i>Exp</i>	0.0292	0.297	0.0215	0.5	1.4	-	-
$1\sigma$	0.0008	0.017	0.0007	0.1	0.2	-	-
<i>prediction</i>	0.0280	0.291	0.0486	0.331	1.47	1.83	1.26

Normal mass ordering is predicted

$$m_1 = 1.096 \times 10^{-2} \text{ eV}$$

$$m_2 = 1.387 \times 10^{-2} \text{ eV}$$

$$m_3 = 5.231 \times 10^{-2} \text{ eV}$$



**Status of neutrino oscillations 2018: 3  
 $\sigma$  hint for normal mass ordering and improved CP sensitivity**

P.F. de Salas (Valencia U., IFIC), D.V. Forero (Campinas State U. & Virginia Tech.), C.A. Ternes, M. Tortola, J.W.F. Valle (Valencia U., IFIC). Aug 3, 2017. 8 pp.  
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# Charged Lepton Sector

$$\mathcal{Y}_e = \begin{pmatrix} a \varphi_1 & a \varphi_3 & a \varphi_2 \\ b \varphi_2 & b \varphi_1 & b \varphi_3 \\ c \varphi_3 & c \varphi_2 & c \varphi_1 \end{pmatrix}$$

$$U_e = \begin{pmatrix} 1 & \varphi_3 & 0 \\ 0 & -\varphi_3 & 1 \\ -\varphi_3 & 1 & \varphi_3 \end{pmatrix} + \dots$$

where dots stand for terms of order  $\varphi_3^2$ ,  $(m_e^2/m_\mu^2)\varphi_3$  and  $(m_\mu^2/m_\tau^2)\varphi_3$ .

# Fit to Model 1

	best value	pull
$r \equiv  \Delta m_{sol}^2 / \Delta m_{atm}^2 $	0.0302(11)	+0.13
$m_3/m_2$	0.0150(5)	—
$\sin^2 \theta_{12}$	0.304(17)	+0.08
$\sin^2 \theta_{13}$	0.0217(8)	-0.13
$\sin^2 \theta_{23}$	0.577(4)	+0.67
$\delta/\pi$	1.529(3)	+0.07
$\alpha_{21}/\pi$	0.135(6)	—
$\alpha_{31}/\pi$	1.728(18)	—

Inverted mass Ordering

- no SUSY breaking effects
- no RGE corrections

best fit parameters

$\tau$	$0.0117 + i 0.9948$
$\varphi_3$	-0.086

close to  
the self-dual  
critical point

$$\chi^2_{min} = 0.4$$

8 dimensionless physical  
quantities independent on  
any coupling constant!

$$m_1 = 4.90(3) \times 10^{-2} \text{eV} \quad , \quad m_2 = 4.98(2) \times 10^{-2} \text{eV} \quad , \quad m_3 = 7.5(3) \times 10^{-4} \text{eV}$$

$|m_{ee}| = 4.73(4) \times 10^{-2} \text{eV}$  by reproducing individually  
 $\Delta m_{sol}^2$  and  $\Delta m_{atm}^2$

# Fit to Yukawa couplings

Model 1

$a \cos \beta$	$2.806923 \times 10^{-6}$
$b \cos \beta$	$9.992488 \times 10^{-3}$
$c \cos \beta$	$5.899778 \times 10^{-4}$

Model 2

$a \cos \beta$	$2.809569 \times 10^{-6}$
$b \cos \beta$	$9.961316 \times 10^{-3}$
$c \cos \beta$	$5.899455 \times 10^{-4}$

$y_e(m_Z)$	$2.794745 \times 10^{-6}$	0.0
$y_\mu(m_Z)$	$5.899864 \times 10^{-4}$	+0.05
$y_\tau(m_Z)$	$1.002950 \times 10^{-2}$	0.0

$y_e(m_Z)$	$2.794745 \times 10^{-6}$	0.0
$y_\mu(m_Z)$	$5.899863 \times 10^{-4}$	0.0
$y_\tau(m_Z)$	$1.002950 \times 10^{-2}$	0.0

# $1\sigma$ parameter space

Intervals where  $\chi^2 \leq \chi^2_{\min} + 1$ :

	IO	NO
$\text{Re}(\tau)$	$[0.0113, 0.0120]$	$[-0.2023, -0.1987]$
$\text{Im}(\tau)$	$[0.9944, 0.9951]$	$[1.0522, 1.0633]$
$\text{Re}(\varphi_3)$	$[-0.090, -0.082]$	$[0.113, 0.121]$

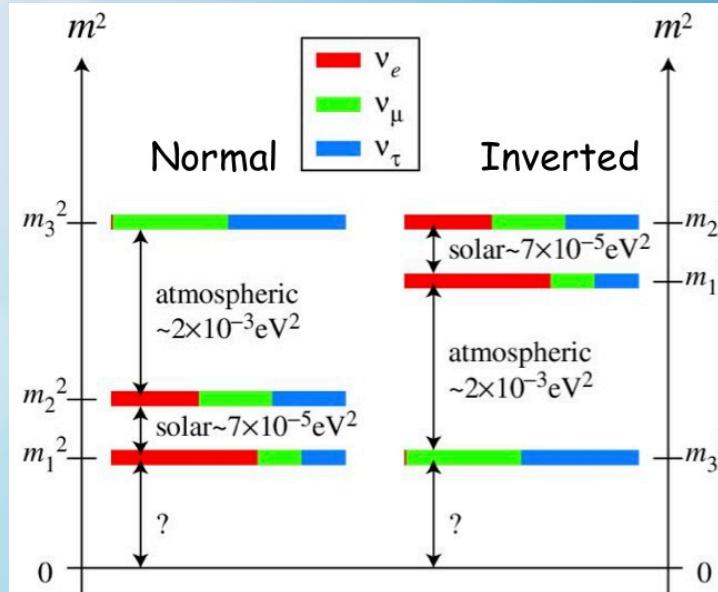
# relevant parameters

$$m_1 < m_2 \quad [\Delta m_{ij}^2 \equiv m_i^2 - m_j^2]$$

$$\Delta m_{21}^2 < |\Delta m_{32}^2|, |\Delta m_{31}^2|$$

i.e. 1 and 2 are, by definition, the closest levels

two possibilities: NO and IO



## Mixing matrix $U_{PMNS}$ (Pontecorvo,Maki,Nakagawa,Sakata)

$$L_{CC} = -\frac{g}{\sqrt{2}} W_\mu^- \bar{e}_L \gamma^\mu U_{PMNS} \nu_L$$

standard parametrization

$$U_{PMNS} = \begin{pmatrix} \nu_e & \nu_1 & \nu_2 & \nu_3 \\ \nu_\mu & c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{-i\delta} \\ \nu_\tau & -s_{12} c_{23} - c_{12} s_{13} s_{23} e^{i\delta} & c_{12} c_{23} - s_{12} s_{13} s_{23} e^{i\delta} & c_{13} s_{23} \\ \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{\frac{i\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{\frac{i\alpha_{31}}{2}} \end{pmatrix}$$

$$0 \leq \vartheta_{ij} \leq \pi / 2$$

$$0 \leq \delta < 2\pi$$

Majorana phases