





Update on UFO Dynamics Studies

Simulations and Run 2 events analysis

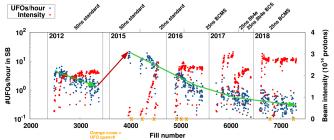
P. Bélanger, B. Lindstrom, R. Schmidt, C. Wiesner, D. Wollmann





A word on UFOs

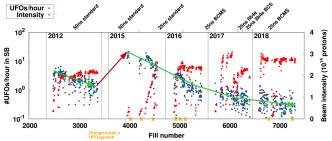
From A. Lechner: "Update about (non-16L2) UFOs" (2018)





A word on UFOs

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Why do we care?

- Beam dumps? Yes, 115 during Run 2
- Magnet quenches? Yes, 8 during Run 2
- Intensity drop? No, negligible during p-p physics (not an electron machine)



Data Collection

Instruments

- Beam Loss Monitors (BLM)
- Diamond BLMs (dBLM)



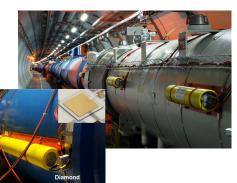




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Databases

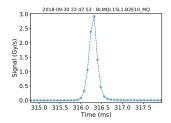
- UFO Buster :
 - Beam parameters
 - Triggers Capture Buffer : 80 μ s resolution BLM signal
- Post-Mortem database : 40 μ s resolution BLM signal
- dBLM database : 1.6 ns resolution diamond BLMs signal

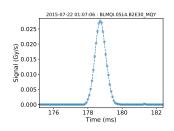




Time profile examples

Some events with signal

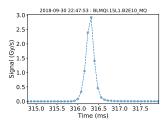




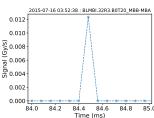


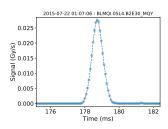
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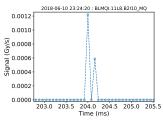
Some events with signal



And some with not much information...











Number of events
337,217



Source	Number of events
UFO Buster	337,217
Matching Capture Buffer	57,262



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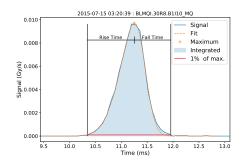
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Followi

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Following slides:

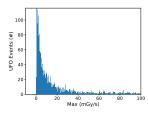


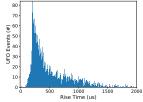


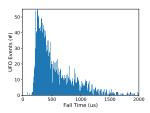


Measurements overview

Parameter		Min.	-	Max.
Maximum Losses	(Gy/s)	9.05×10^{-5}	-	3.16
Integrated Signal	(Gy)	2.90×10^{-8}	-	2.02×10^{-3}
Rise Time	(μs)	91	-	4241
Fall Time	(μs)	109	-	3876
Full length	(μs)	317	-	7118







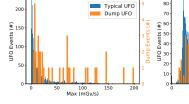


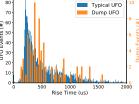
Normal UFOs v.s. Dump UFOs

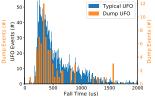
Dump UFOs seem to have :

- Higher losses
- · Longer rise time
- Shorter fall time

However, not a lot of statistics... In-depth analysis of the dump UFOs is still to be done.











Dynamics Simulation Tool



Tool developed (starting in 2010, F. Zimmermann, B. Auchmann *et al.*) to simulate UFO dynamics :

- 1. UFO begins to fall and/or be attracted toward the proton beam
- 2. UFO-beam interaction knocks-off electrons in the UFO (ionization) and lead to proton losses
- 3. The charged UFO is repelled by the beam

Recently translated in Python by R. Schmidt, and under continuous development



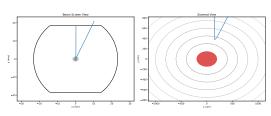
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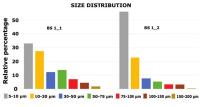


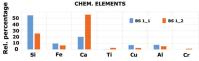


Simulations - Parameter scan

- Range of parameters is based on past studies:
 - FLUKA simulations (A. Lechner)
 - Dust collection in the beam pipe (L. Grob)

From L. Grob: "Dust Analysis From LHC Vacuum System to Identify the Source of Macro Particle-Beam-Interactions" (2019)





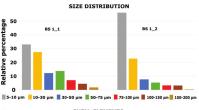


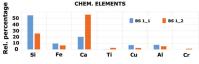
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Parameter	Values
UFO Charge	$0 - 10^8 \cdot (-e)$
UFO Position	Top of beam screen
UFO Radius	$1~\mu$ m $-$ 50 $~\mu$ m
UFO Material	Cu, Si, Al
Beam Energy	6.5 TeV
Beam Intensity	$3 \times 10^{10} - 3 \times 10^{14}$
Beam σ_x	50 μ m $-$ 500 μ m
Beam σ_v	$50~\mu\mathrm{m}-500~\mu\mathrm{m}$

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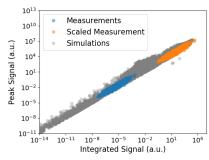








Simulations - Results



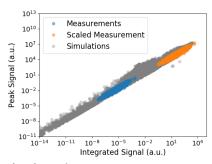
Log-Log axis

- Expected "calibration" constant (BLM response to incoming particles)
- · Can be calculated with FLUKA
- Scaled measurements move along the simulated line: compatible with the need for this BLM calibration



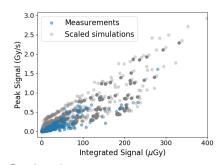


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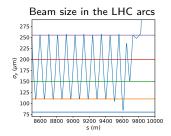
Regular axis

- Range of σ_y influences the thickness of the simulation band (left figure) and the slope (right figure)
- Coverage is compatible with measurements for same σ_y range





Simulations - Beam size

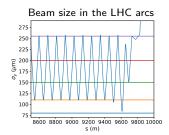




Simulations - Beam size

Let's look at the ratio (Integrated signal)/(Peak Signal)

- Units of seconds
- Independent of the BLM calibration
- Equivalent to time spent in the beam by the UFO
- From previous slide : highly dependant on σ_y

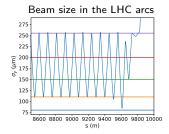


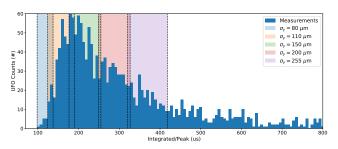


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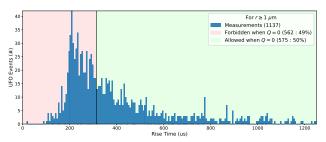


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- We can verify this hypothesis by fixing Q=0 (no UFO charge) and scanning through all the other parameters





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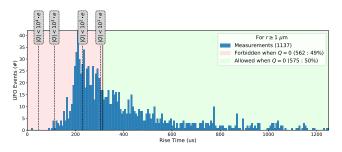
UFO dynamics driven by gravity alone is incompatible with measurements!



 Next logical step: what is the minimum charge which allows to explain the measurements?



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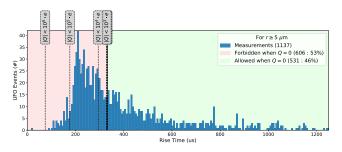


UFOs are generally charged with $|Q|>1000\cdot e \quad (-1.6\times 10^{-16} \text{ C})$





 With increasing UFO radius, more and more charges are needed to explain the measurements

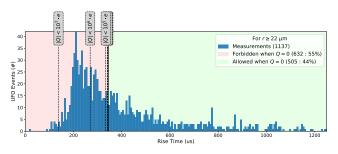


UFOs with $r > 5 \mu m$ are generally charged with $|Q| > 10^6 \cdot e \quad (-1.6 \times 10^{-13} \text{ C})$





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UFOs with $r \geq$ 22 μ m are generally charged with $|\it{Q}| > 10^7 \cdot e ~(-1.6 \times 10^{-12}~\rm{C})$





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- Established a lower bound of 10^6 electrons as the initial UFO charge for radius above 5 μ m in order to explain all measurements.
- Overall: gathered important statistics which could help understand release mechanism!



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- Monte-Carlo simulations with appropriate distributions for the input parameters of the dynamics simulation is needed to further our understanding.





Summary of work done

Since January 2019:

- Literature review of previous UFO studies.
- Development of a more accurate description of the electric field around the LHC proton beam.
- Upgrade of the UFO Dynamics Simulation Tool
- Development of new methods of analysis for studying UFO bunch-by-bunch signals.
- Gathering of all available data from standard beam operation between 2015 and 2018. First look and direct conclusions.
- Started an in-depth analysis of the data mentioned above.



E-field around the LHC proton beam

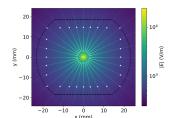
 Original approximation considered the Houssais E-field in free space (i.e. no beam screen)

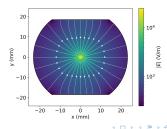
 A new method to solve Laplace's equation for complex boundary conditions was developed and allowed to find a more accurate description of the E-field

Bélanger, P. (2019). Generalizing the Method of Images for Complex

Boundary Conditions: Application on the LHC Beam Screen.

https://arxiv.org/abs/1905.03405

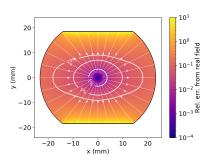






E-field around the LHC proton beam

Conclusion: ignoring the beam screen accounts for an error of 1% in the E-field at 1σ from the center of the beam and 10% at 30σ



The method developed can also be used for off-centered beam, at really low computing cost.

