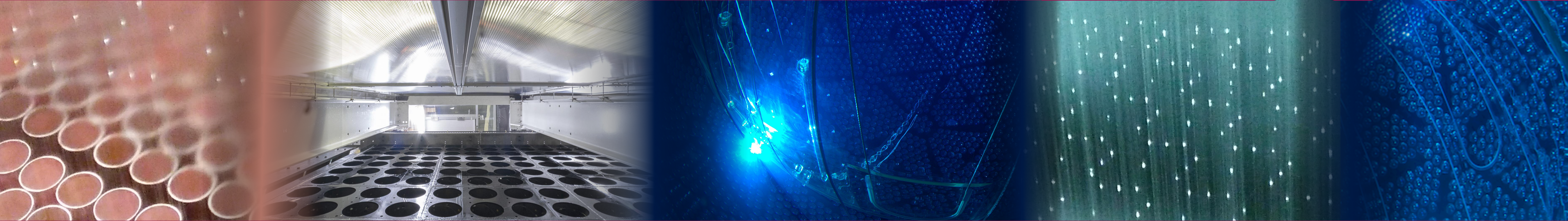


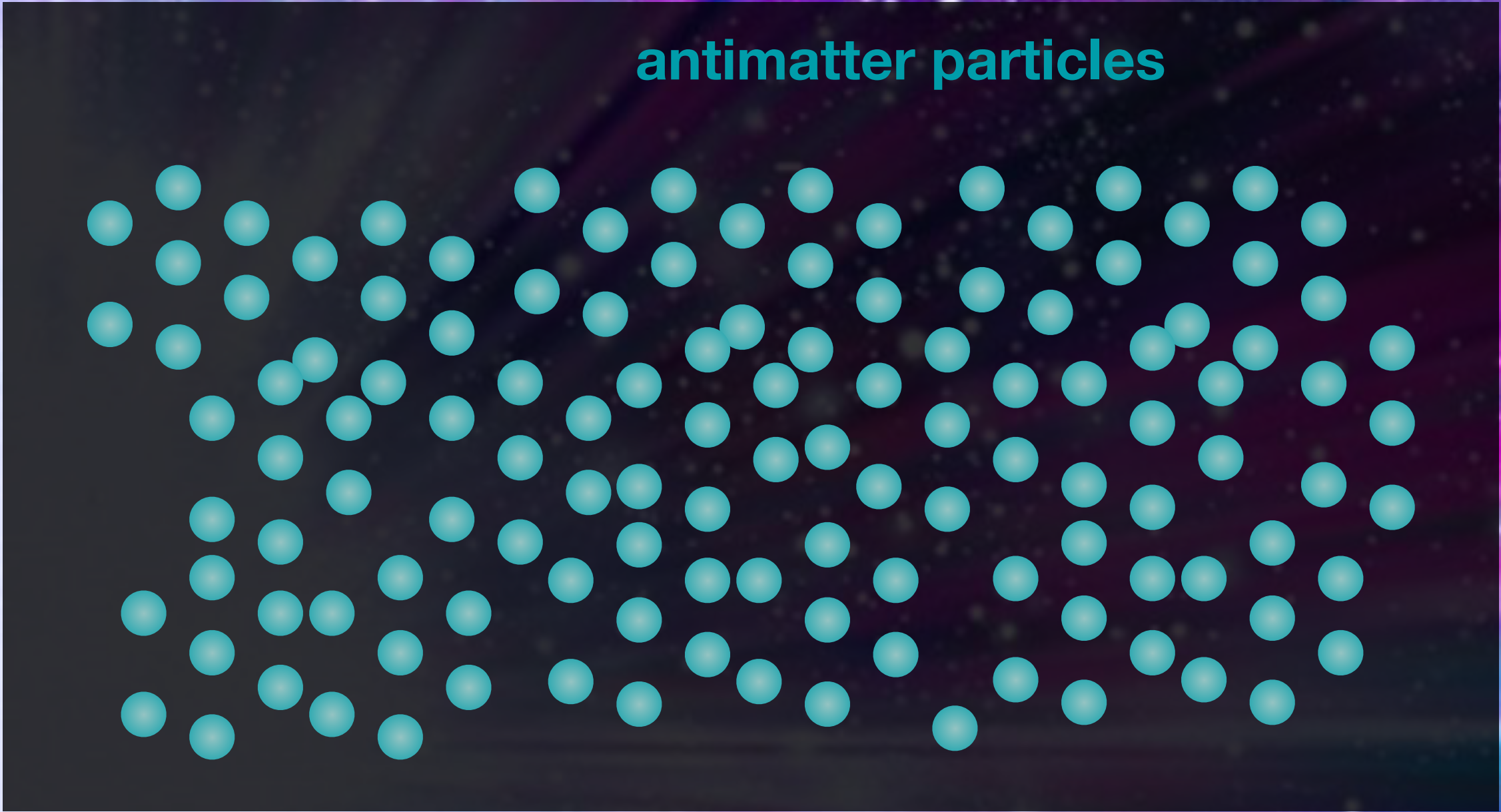
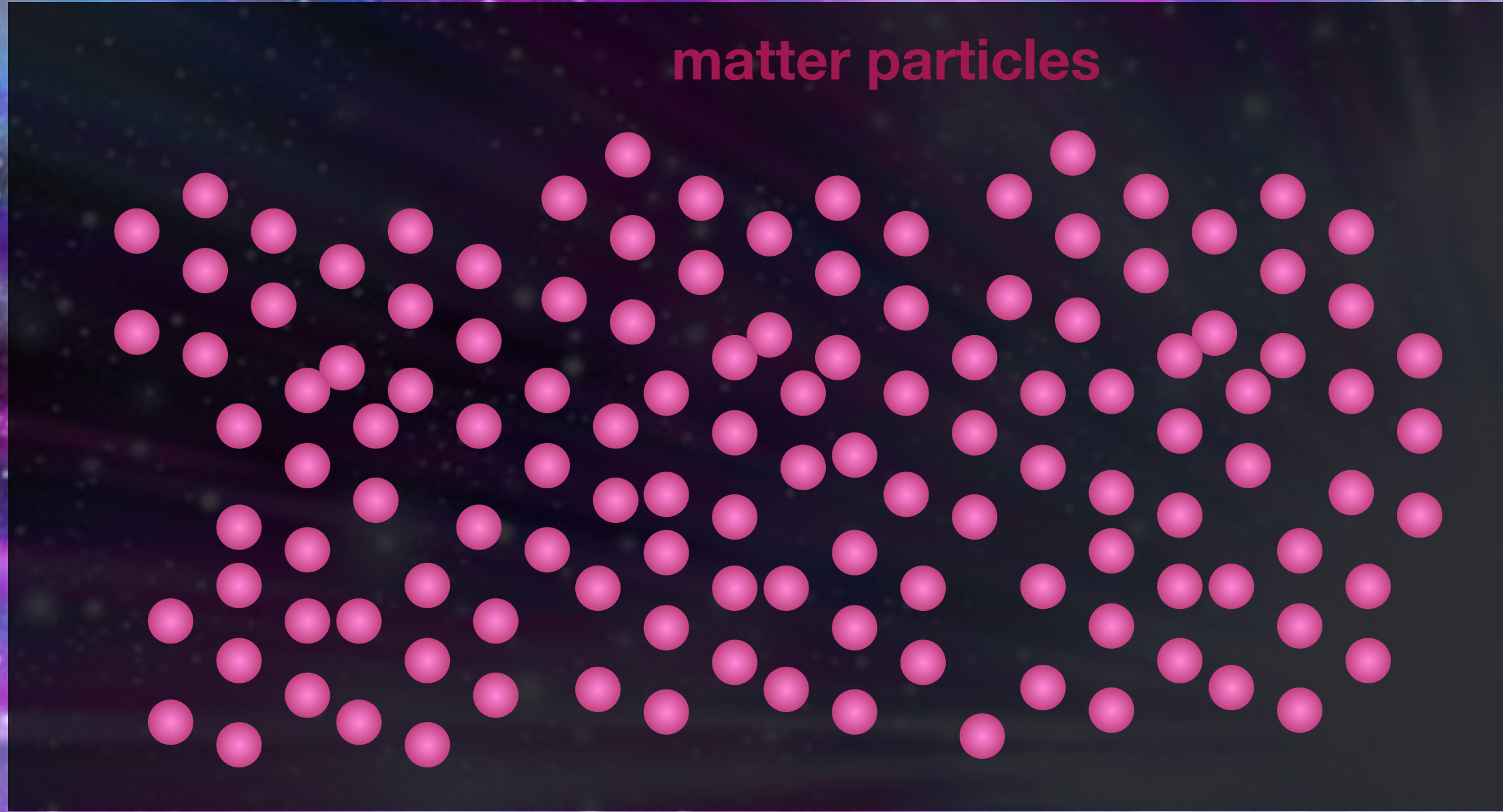
# Neutrinoless $\beta\beta$ decay: physics that matters



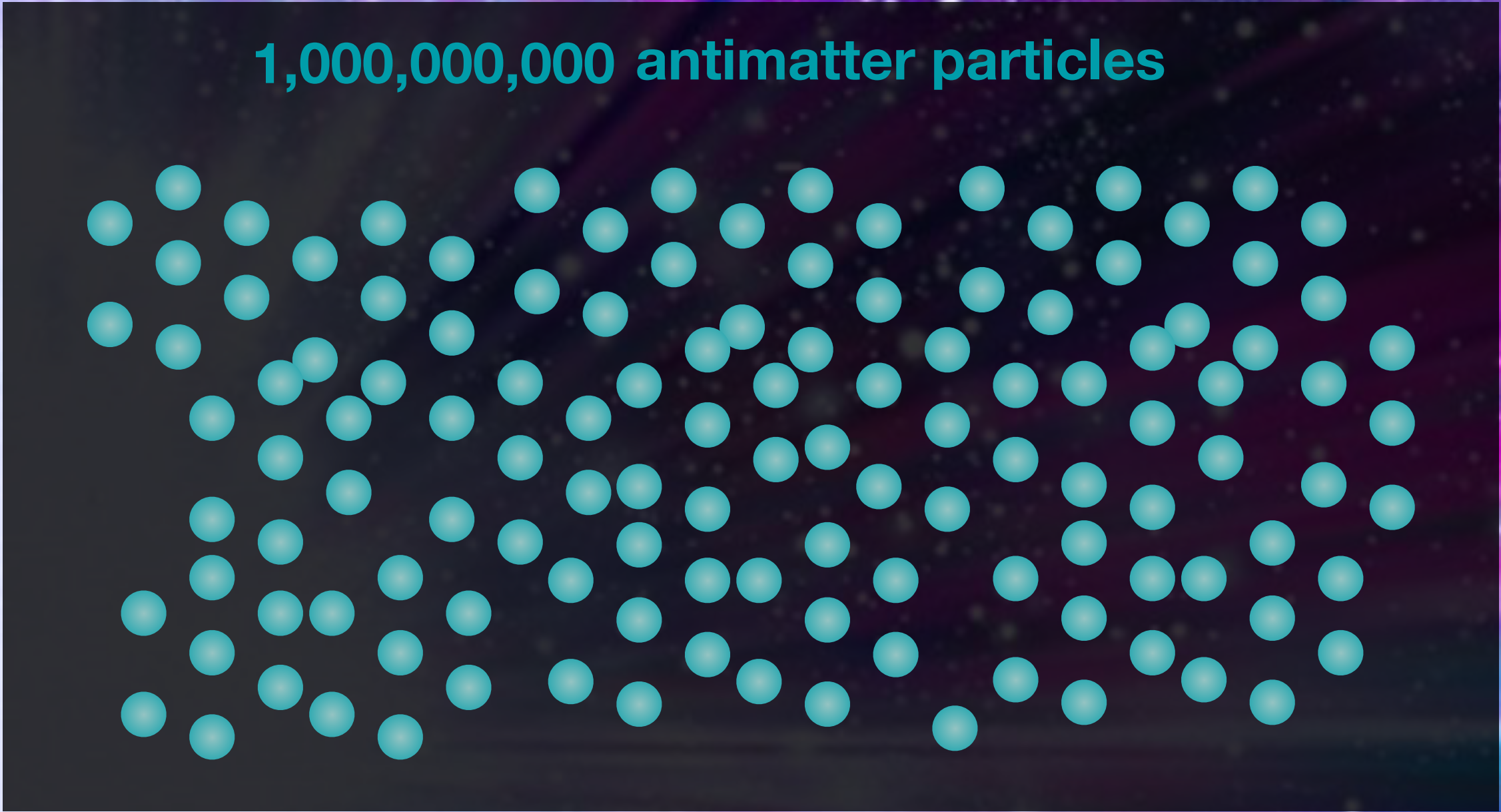
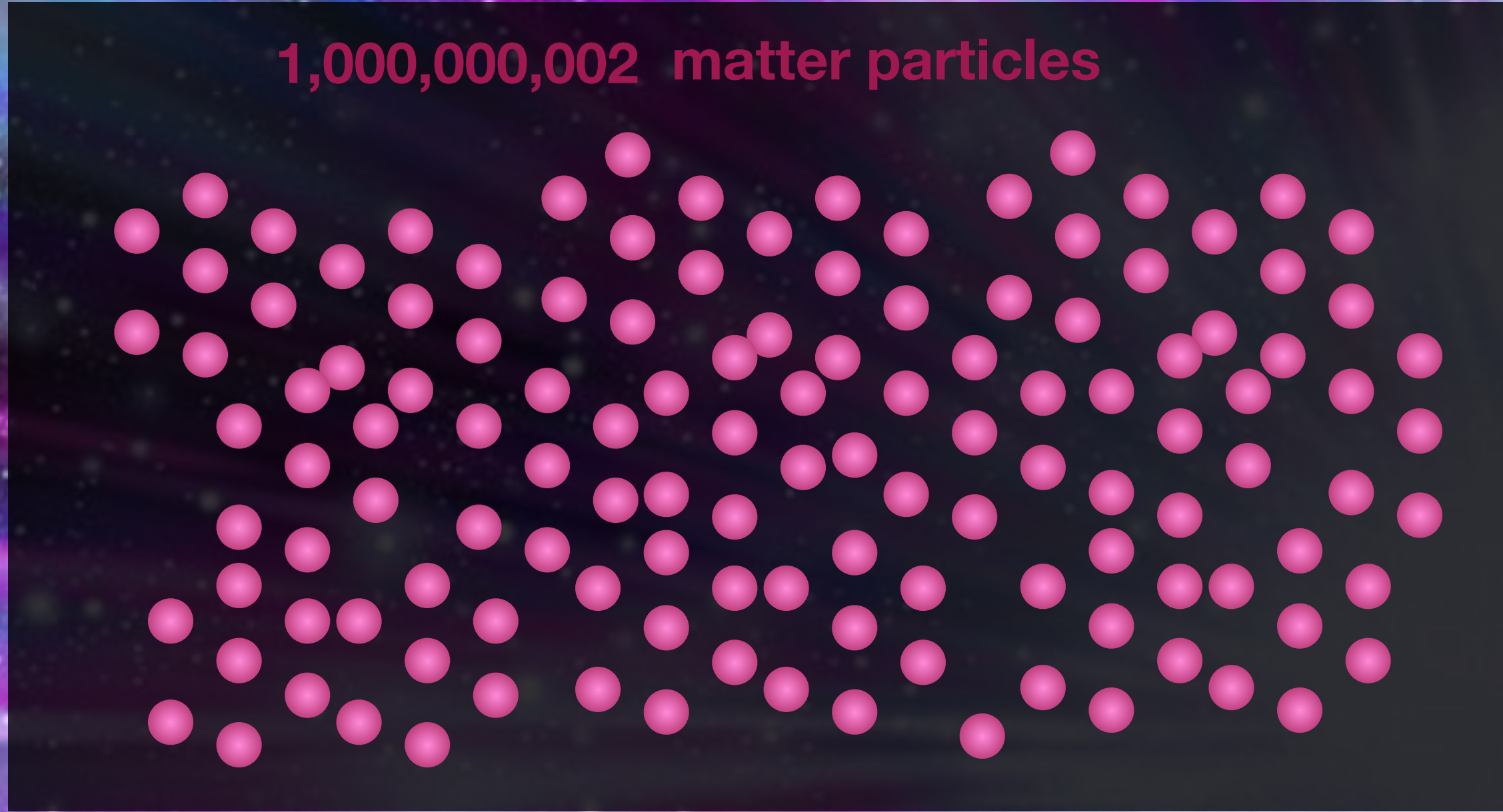
(...but doesn't antimatter)

**Cheryl Patrick, UCL**  
Joint IoP APP/HEPP Conference 2019

# In the beginning was...

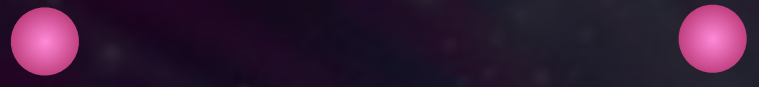


# In the beginning was...

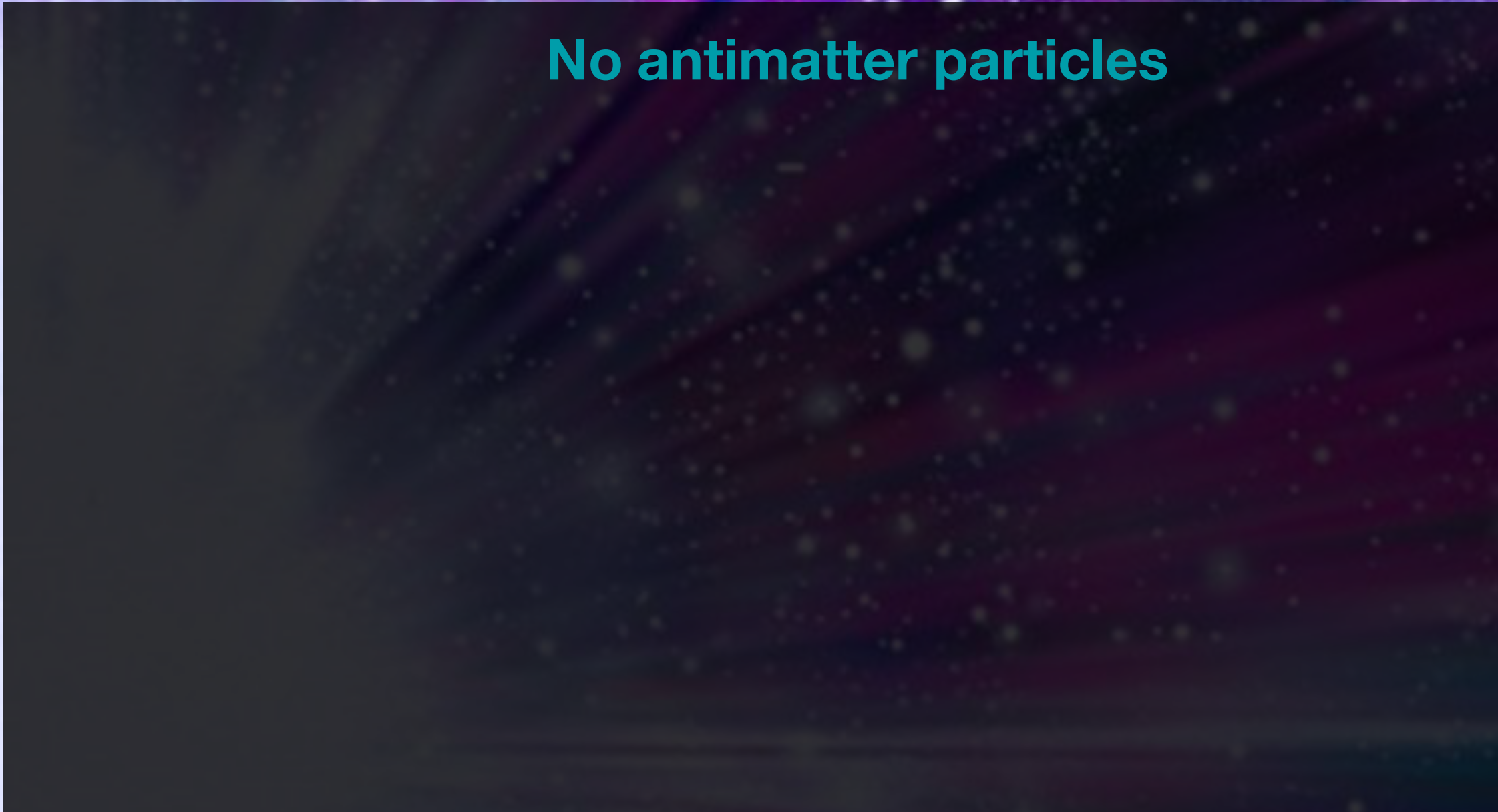


# In the beginning was...

2 matter particles



No antimatter particles



# In the beginning was...

2 matter particles



No antimatter particles

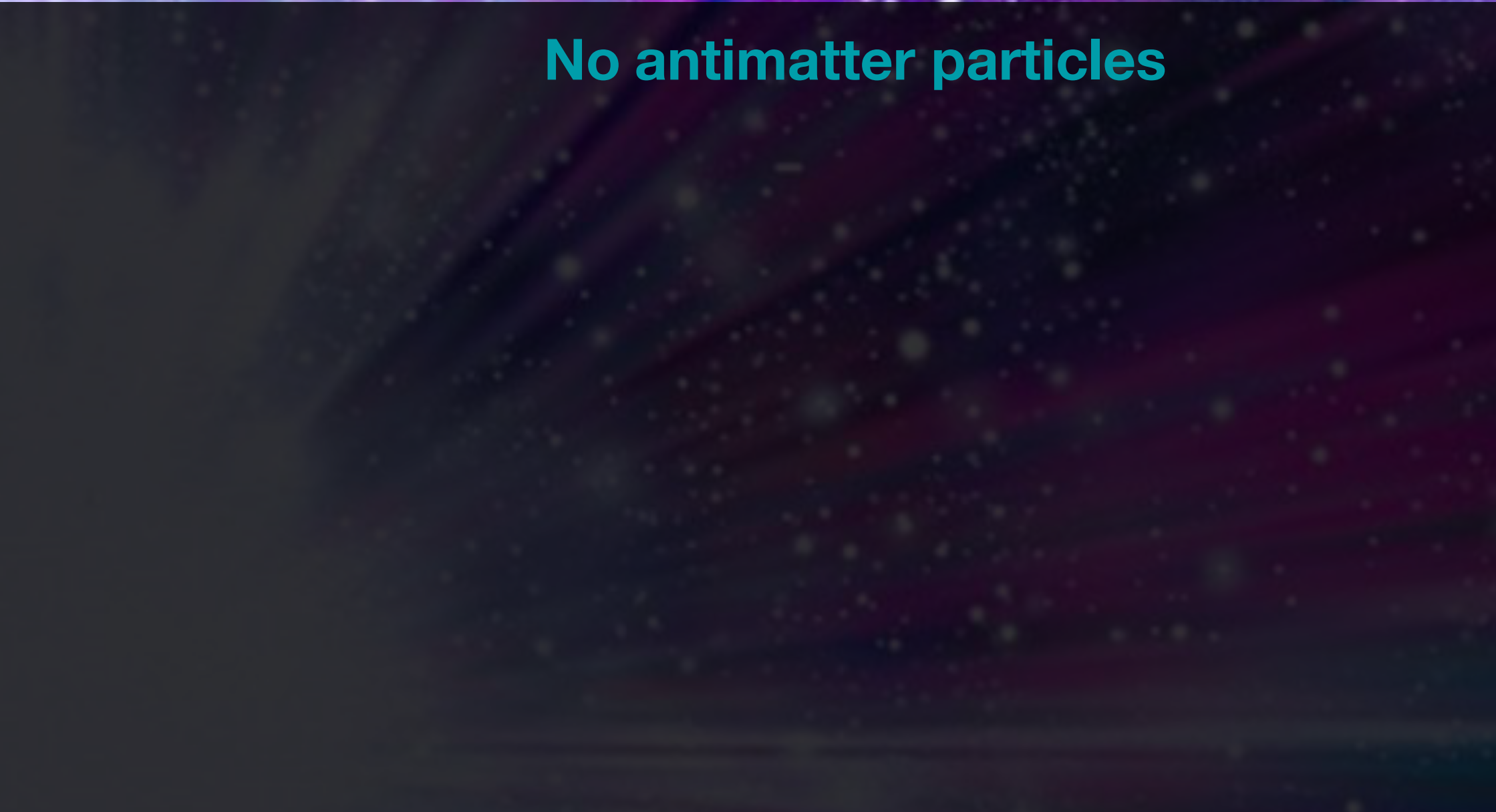
- How did we get this imbalance?

# In the beginning was...

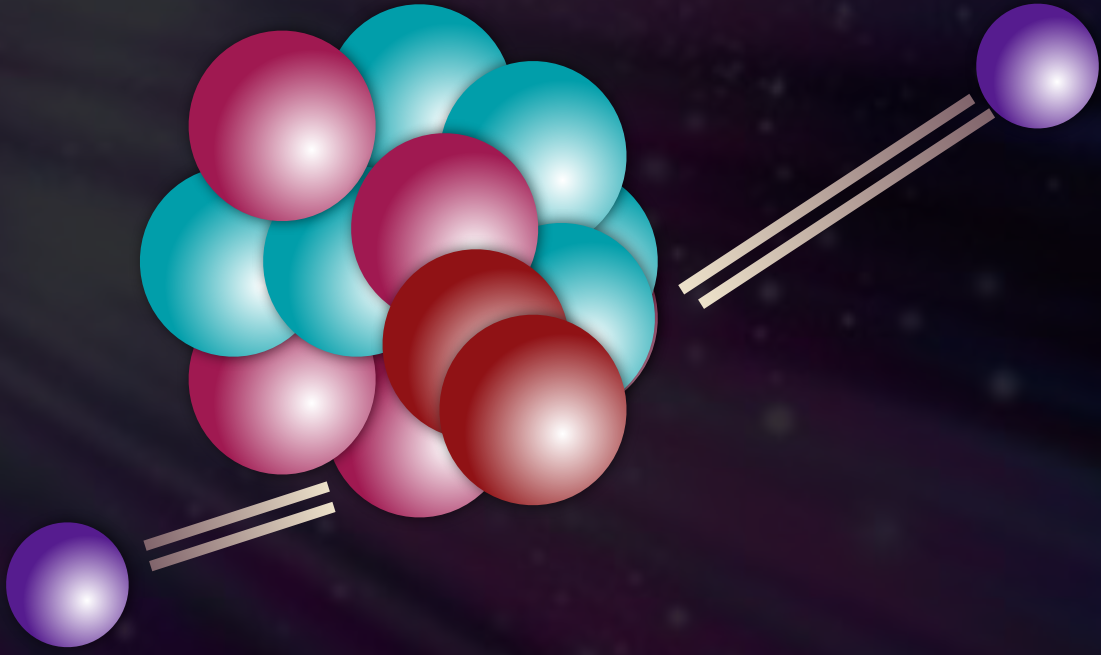
2 matter particles



No antimatter particles



- How did we get this imbalance?
- **Neutrinoless double-beta decay** would CREATE matter...



# In the beginning was...

2 matter particles



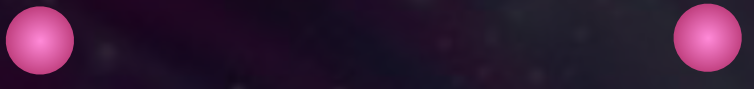
No antimatter particles

- How did we get this imbalance?
- **Neutrinoless double-beta decay** would CREATE matter...
- ... tells us about the **nature of the neutrino** (that isn't there)...



# In the beginning was...

2 matter particles



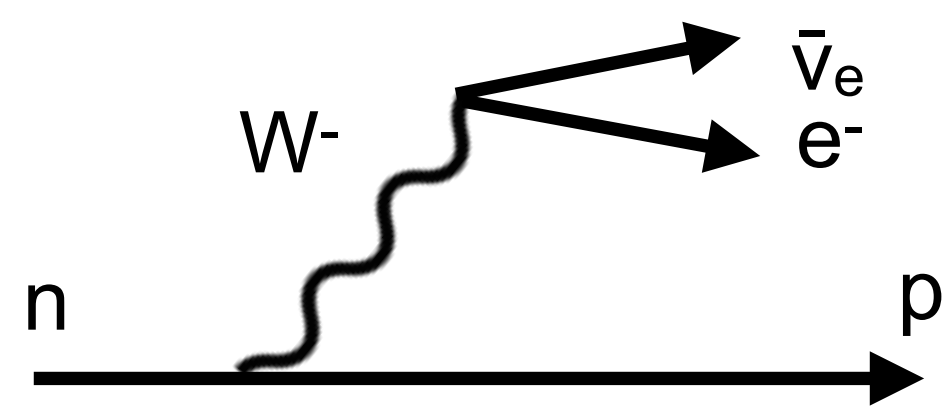
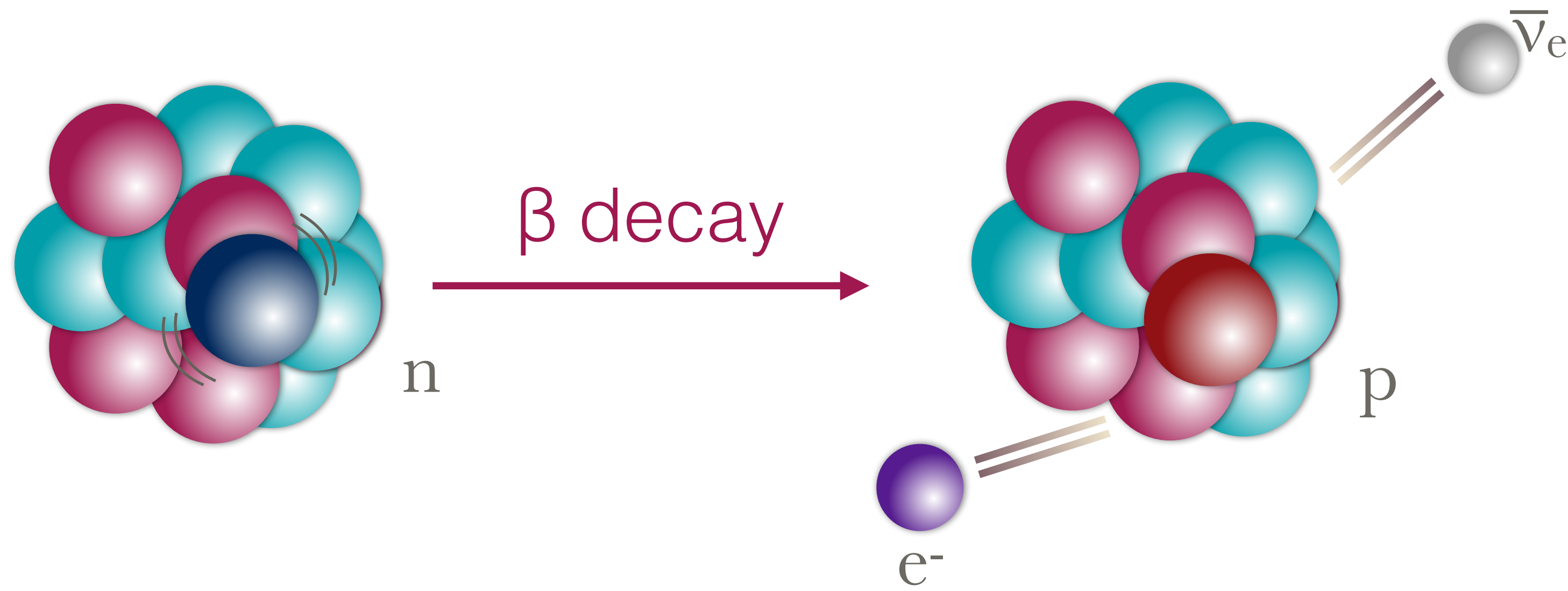
No antimatter particles

- How did we get this imbalance?
- **Neutrinoless double-beta decay** would CREATE matter...
- ... tells us about the **nature of the neutrino** (that isn't there)...
- ... and could help us measure its **mass**

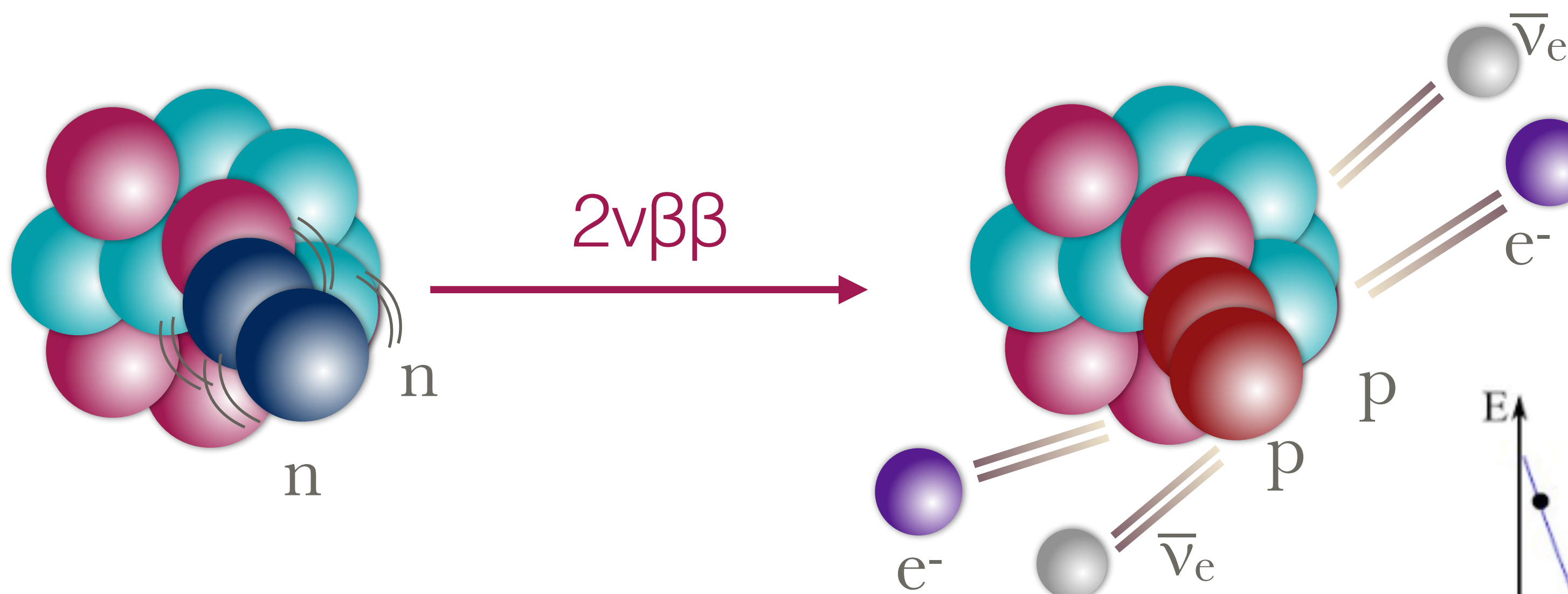




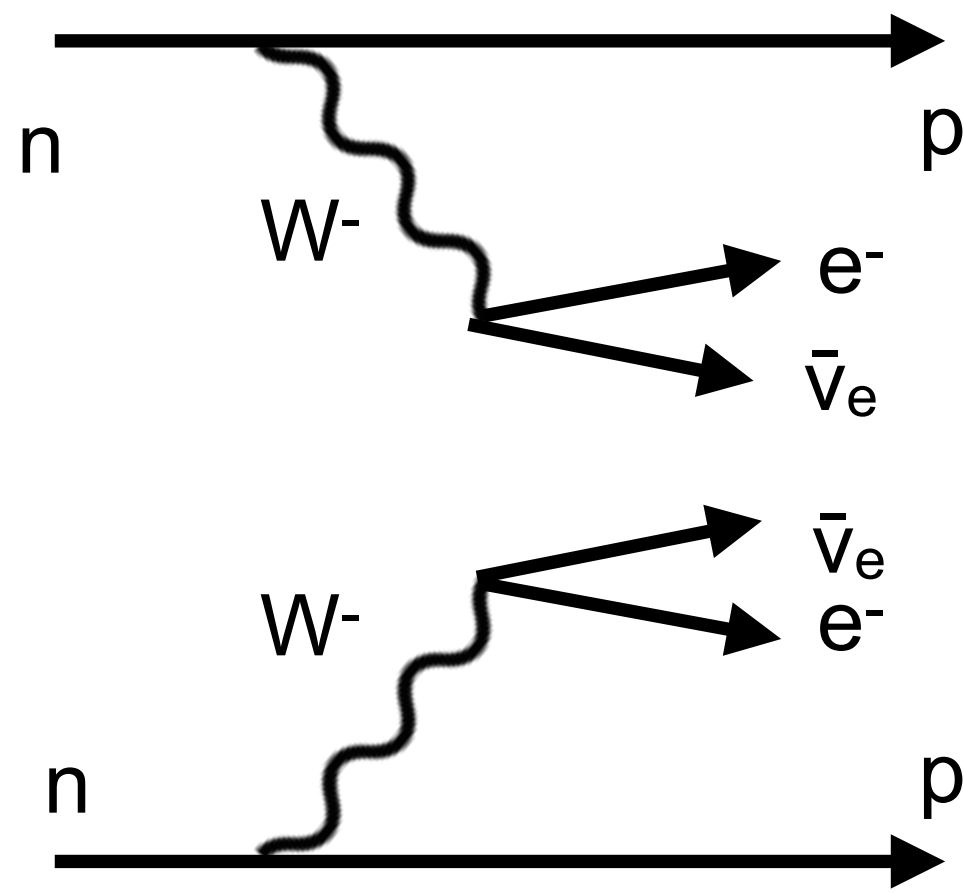
# beta decay



# double-beta decay



1 H 1.008	2 He 4.0026											13 Al 26.982	14 Si 28.085	15 P 30.974	16 S 32.06	17 Cl 35.45	18 Ar 39.948
3 Li 6.94	4 Be 9.0122											5 B 10.81	6 C 12.011	7 N 14.007	8 O 15.999	9 F 18.998	10 Ne 20.180
11 Na 22.990	12 Mg 24.305	3 Sc 44.956	4 Ti 47.867	5 V 50.942	6 Cr 51.996	7 Mn 54.938	8 Fe 55.845	9 Co 58.933	10 Ni 58.693	11 Cu 63.546	12 Zn 65.38	31 Ga 69.723	32 Ge 72.63	33 As 74.922	34 Se 78.96	35 Br 79.904	36 Kr 83.798
19 K 39.098	20 Ca 40.078	21 Sc 44.956	22 Ti 47.867	23 V 50.942	24 Cr 51.996	25 Mn 54.938	26 Fe 55.845	27 Co 58.933	28 Ni 58.693	29 Cu 63.546	30 Zn 65.38	49 In 114.82	50 Sn 118.71	51 Sb 121.76	52 Te 127.6	53 I 126.90	54 Xe 131.29
37 Rb 85.468	38 Sr 87.62	39 Y 88.906	40 Zr 91.224	41 Nb 92.906	42 Mo 95.94	43 Tc (98)	44 Ru 101.07	45 Rh 102.91	46 Pd 106.42	47 Ag 107.87	48 Cd 112.41	76 Ge 72.63	77 As 74.922	78 Se 78.96	79 Br 79.904	80 Kr 83.798	81 Xe 131.29
55 Cs 132.91	56 Ba 137.33	57-71 * Lanthanide series	72 Hf 178.49	73 Ta 180.95	74 W 183.84	75 Re 186.21	76 Os 190.23	77 Ir 192.22	78 Pt 195.08	79 Au 196.97	80 Hg 200.59	81 Tl 204.38	82 Pb 207.2	83 Bi 208.98	84 Po (209)	85 At (210)	86 Rn (222)
87 Fr (223)	88 Ra (226)	89-103 * Actinide series	104 Rf (261)	105 Db (262)	106 Sg (263)	107 Bh (264)	108 Hs (265)	109 Mt (266)	110 Ds (271)	111 Rg (272)	112 Cn (285)	113 Nh (286)	114 Fl (289)	115 Mc (290)	116 Lv (293)	117 Ts (294)	118 Og (294)
			57 La 138.91	58 Ce 140.12	59 Pr 140.91	60 Nd 144.24	61 Pm (145)	62 Sm 150.36	63 Eu 151.96	64 Gd 157.25	65 Tb 158.93	66 Dy 162.50	67 Ho 164.93	68 Er 167.26	69 Tm 168.93	70 Yb 173.05	71 Lu 174.97
			89 Ac (227)	90 Th 232.04	91 Pa 231.04	92 U 238.03	93 Np (237)	94 Pu (244)	95 Am (243)	96 Cm (247)	97 Bk (247)	98 Cf (251)	99 Es (252)	100 Fm (257)	101 Md (258)	102 No (259)	103 Lr (262)



Lepton number conserved  
 $\Delta L = 0$   
 $T_{1/2} \sim 10^{20}$  years

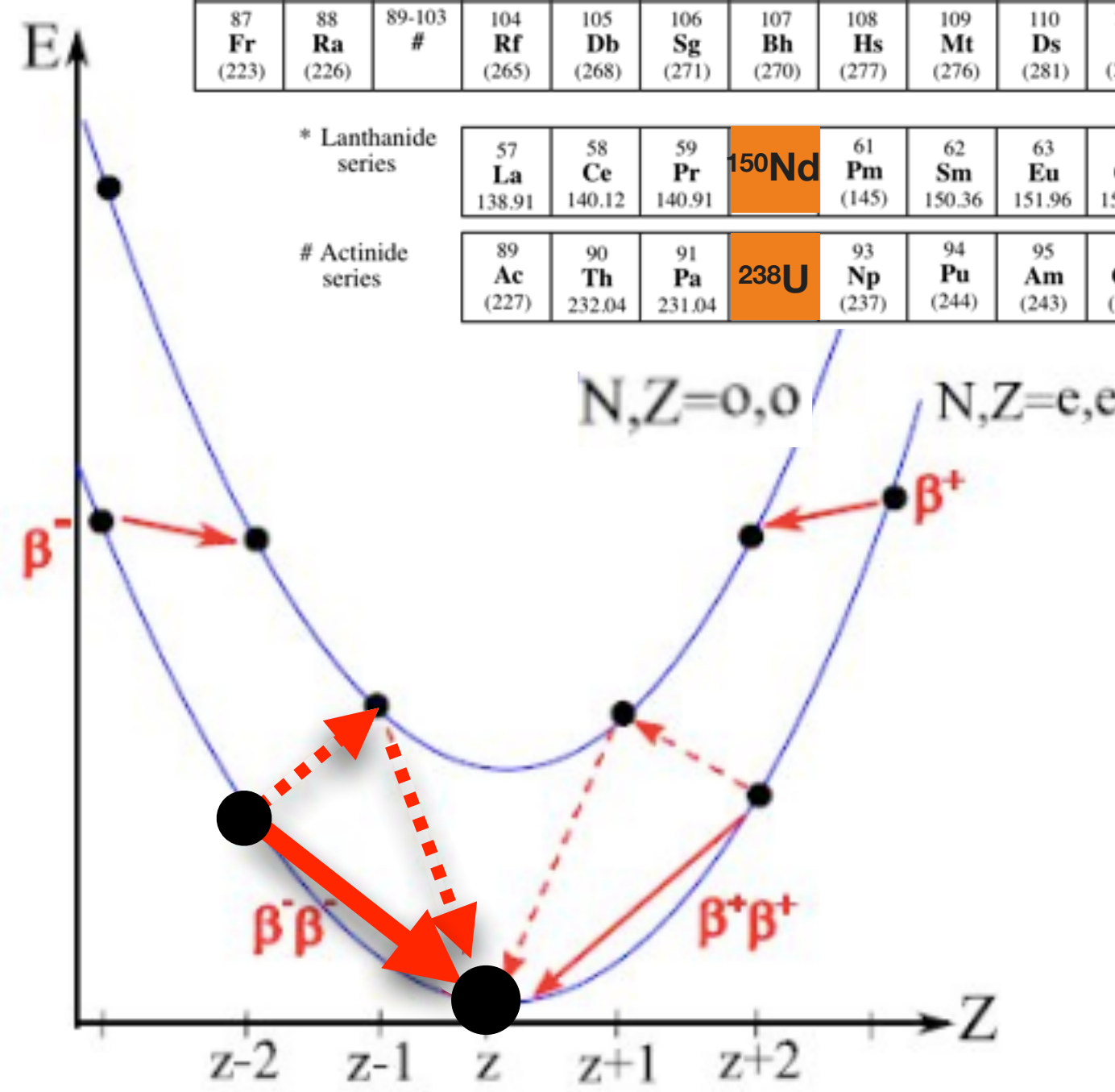
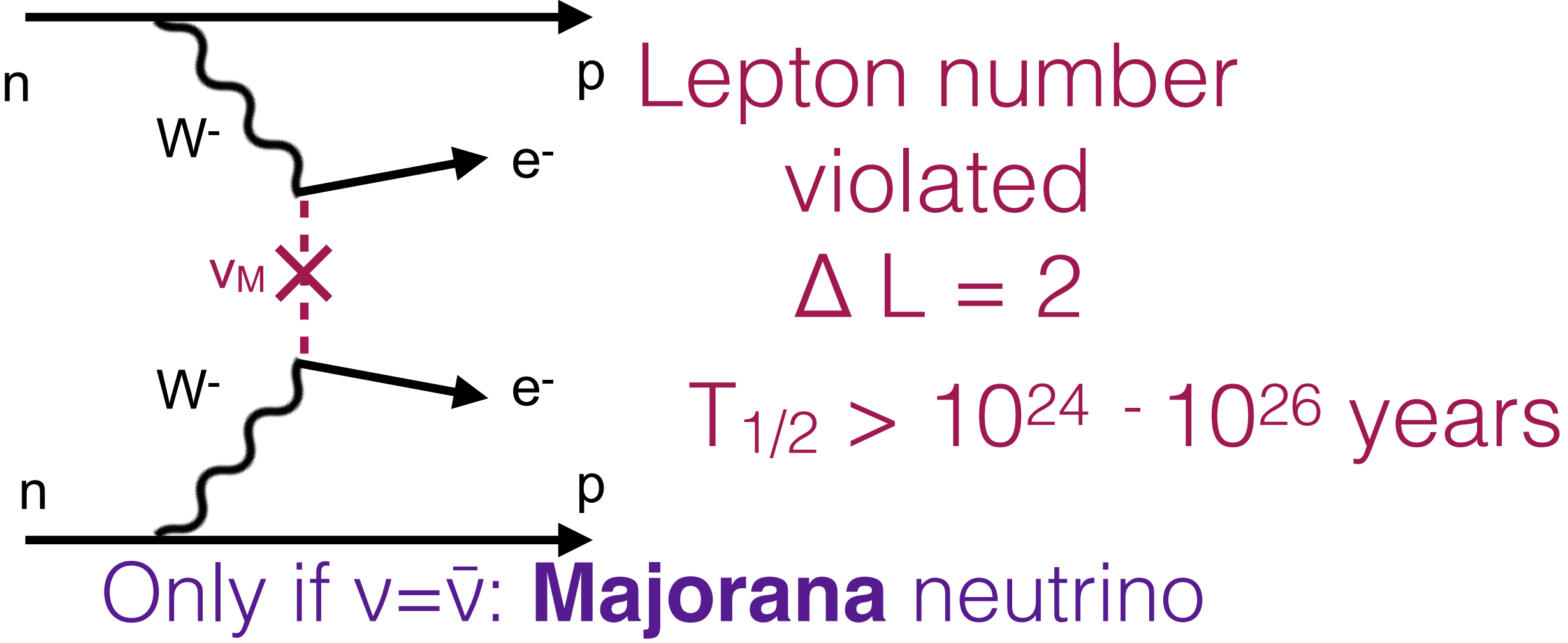
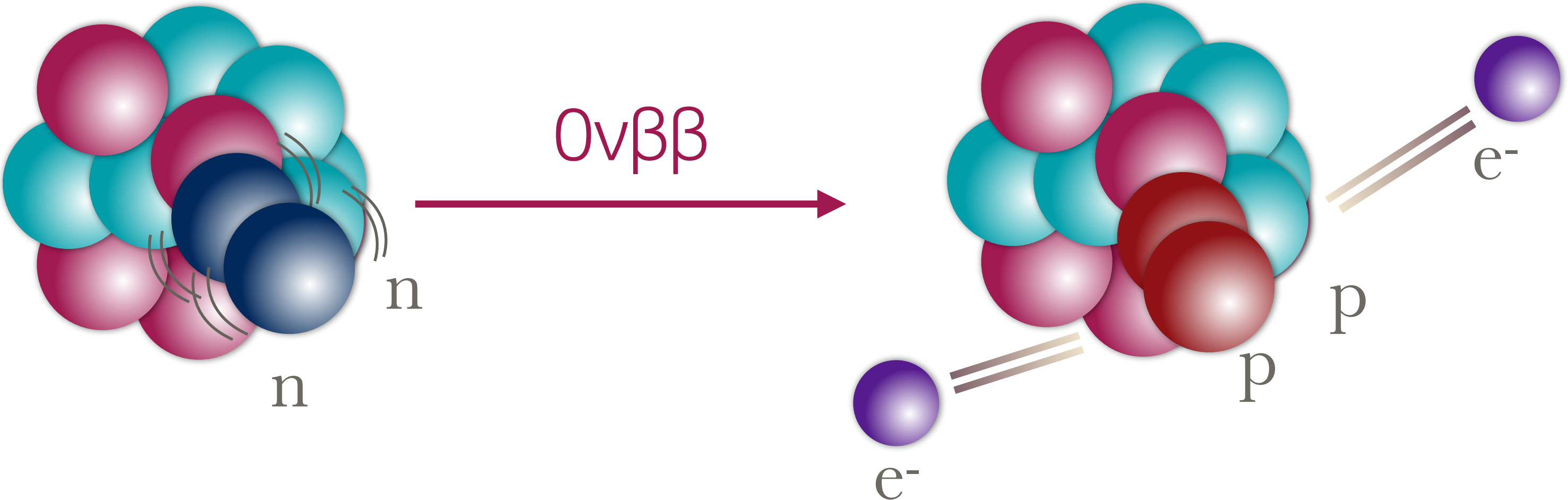


Diagram - E. Falk

# Neutrinoless double-beta decay : the smoking gun for Majorana



# $0\nu\beta\beta$ could tell us about neutrino mass

$$0\nu\beta\beta \text{ rate} = \frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

Phase space factor

Nuclear matrix element  
(hard to calculate)

Effective neutrino mass

## $0\nu\beta\beta$ could tell us about neutrino mass

$$0\nu\beta\beta \text{ rate} = \frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$
$$m_{\beta\beta} = c_{12}^2 c_{13}^2 m_{\nu_1} + s_{12}^2 c_{13}^2 m_{\nu_2} e^{i\phi_{12}} + s_{13}^2 m_{\nu_3} e^{i\phi_{13}}$$

# $0\nu\beta\beta$ could tell us about neutrino mass

$$0\nu\beta\beta \text{ rate} = \frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

**PMNS mixing angles/phases**

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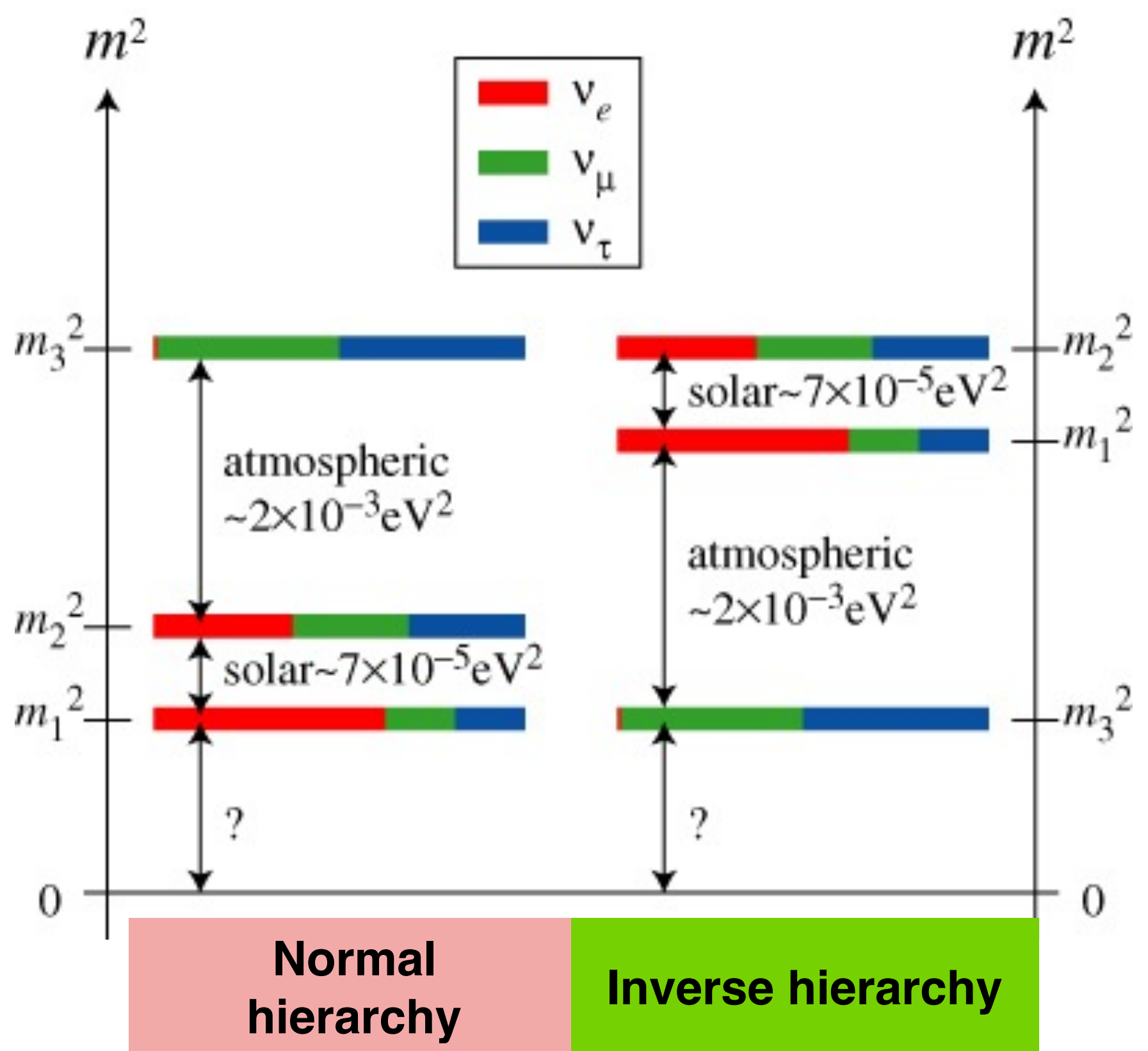
$$0\nu\beta\beta \text{ rate} = \frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

$$m_{\beta\beta} = c_{12}^2 c_{13}^2 m_{\nu_1} + s_{12}^2 c_{13}^2 m_{\nu_2} e^{i\phi_{12}} + s_{13}^2 m_{\nu_3} e^{i\phi_{13}}$$

PMNS mixing angles/phases  
Individual neutrino masses

# $0\nu\beta\beta$ could tell us about neutrino mass

$$0\nu\beta\beta \text{ rate} = \frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$



$$m_{\beta\beta} = c_{12}^2 c_{13}^2 m_{\nu_1} + s_{12}^2 c_{13}^2 m_{\nu_2} e^{i\phi_{12}} + s_{13}^2 m_{\nu_3} e^{i\phi_{13}}$$

PMNS mixing angles/phases

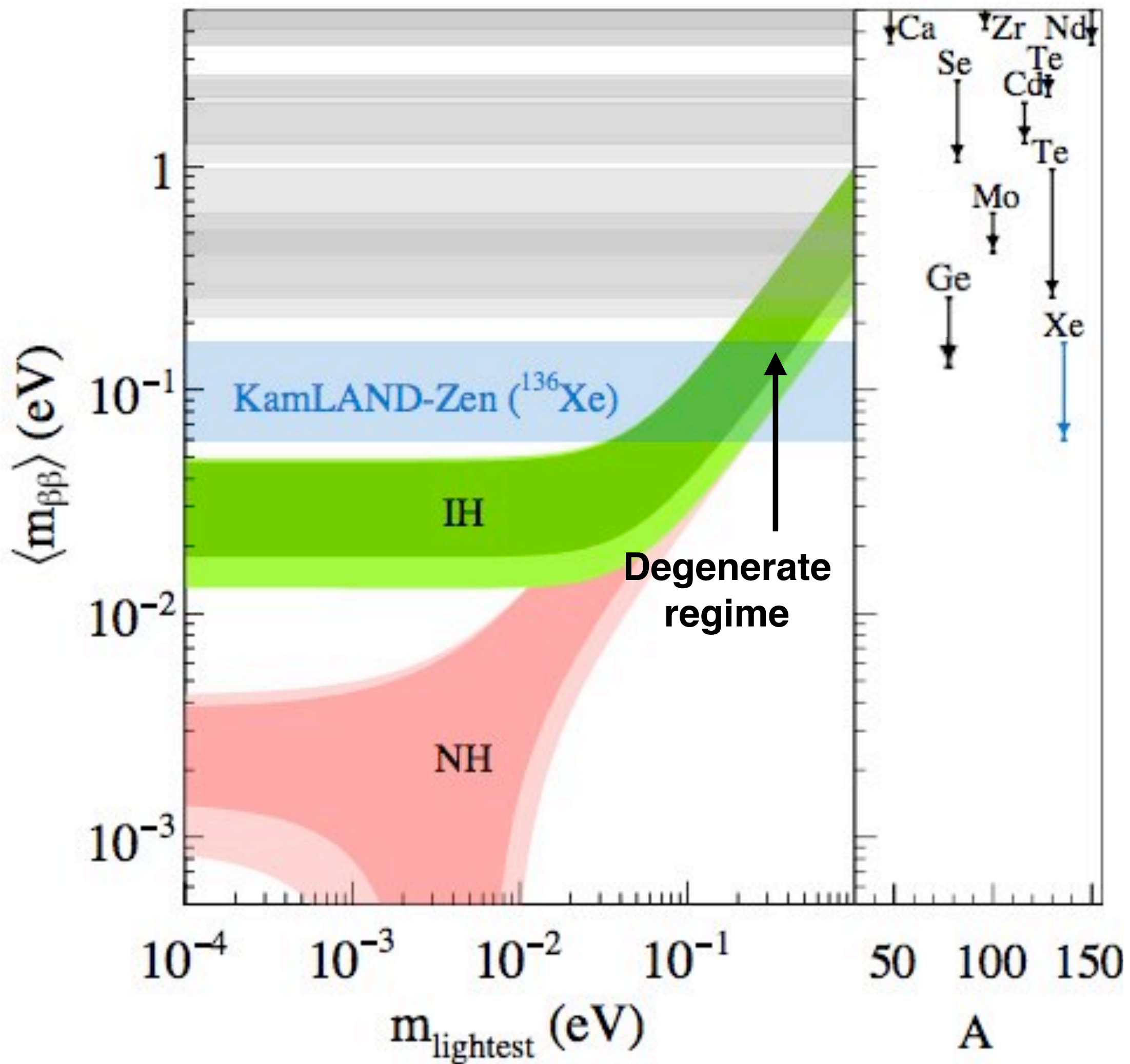
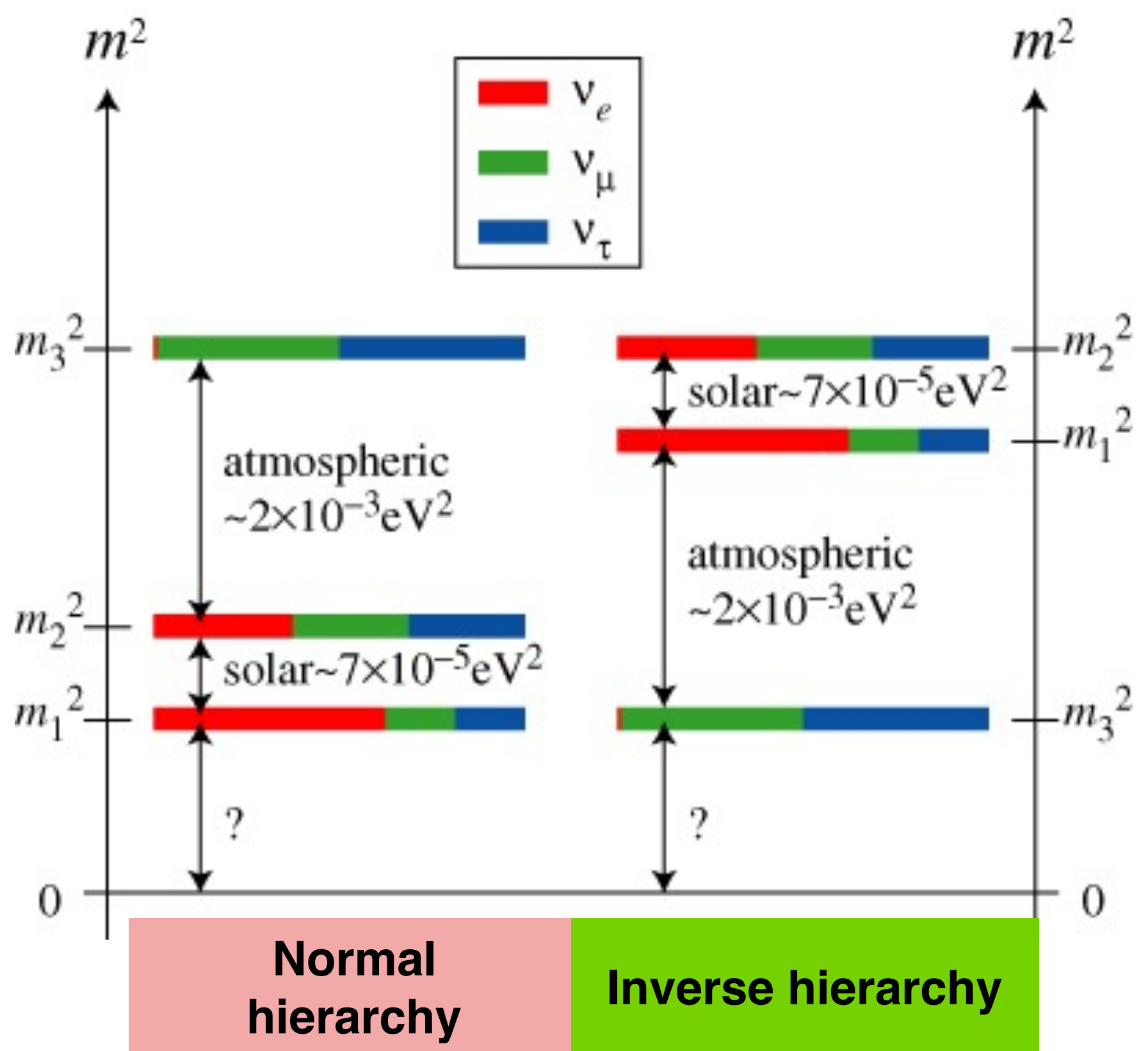
Individual neutrino masses



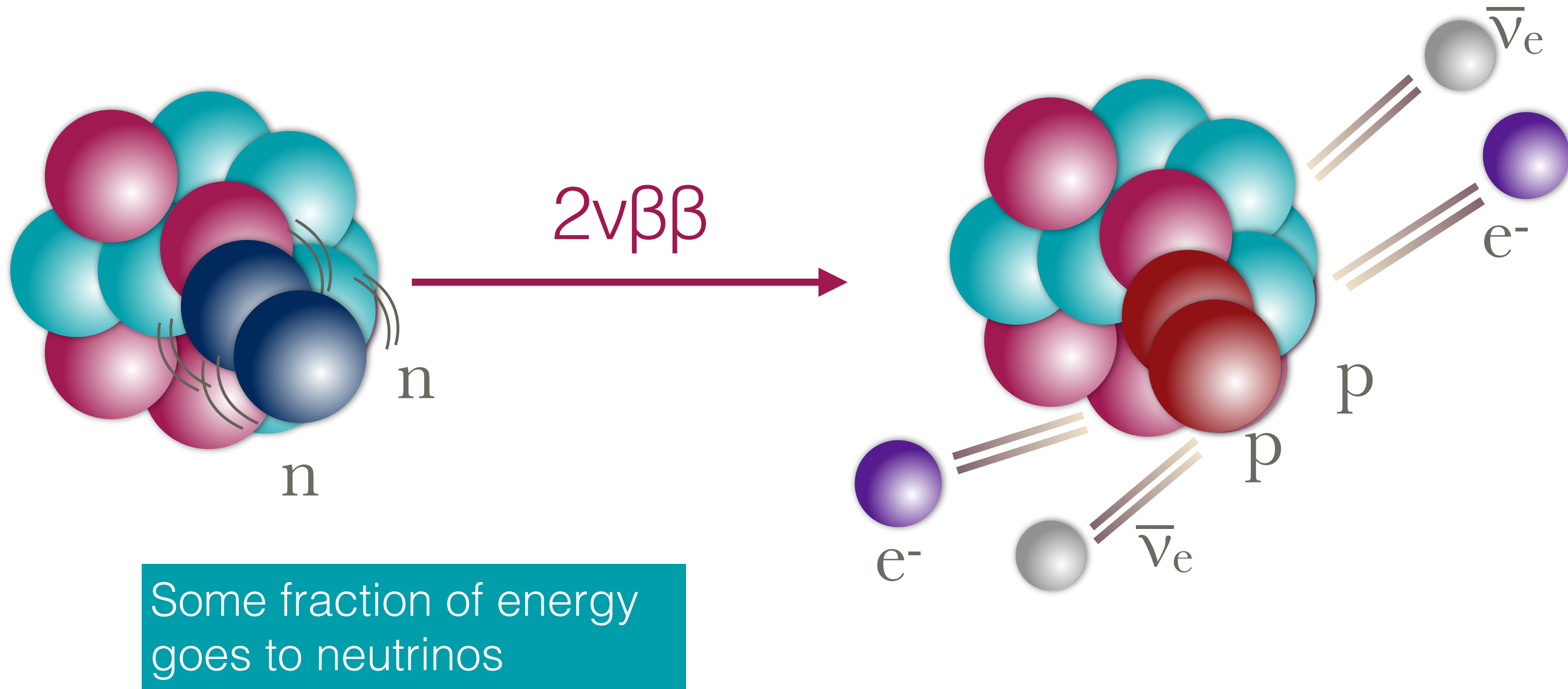
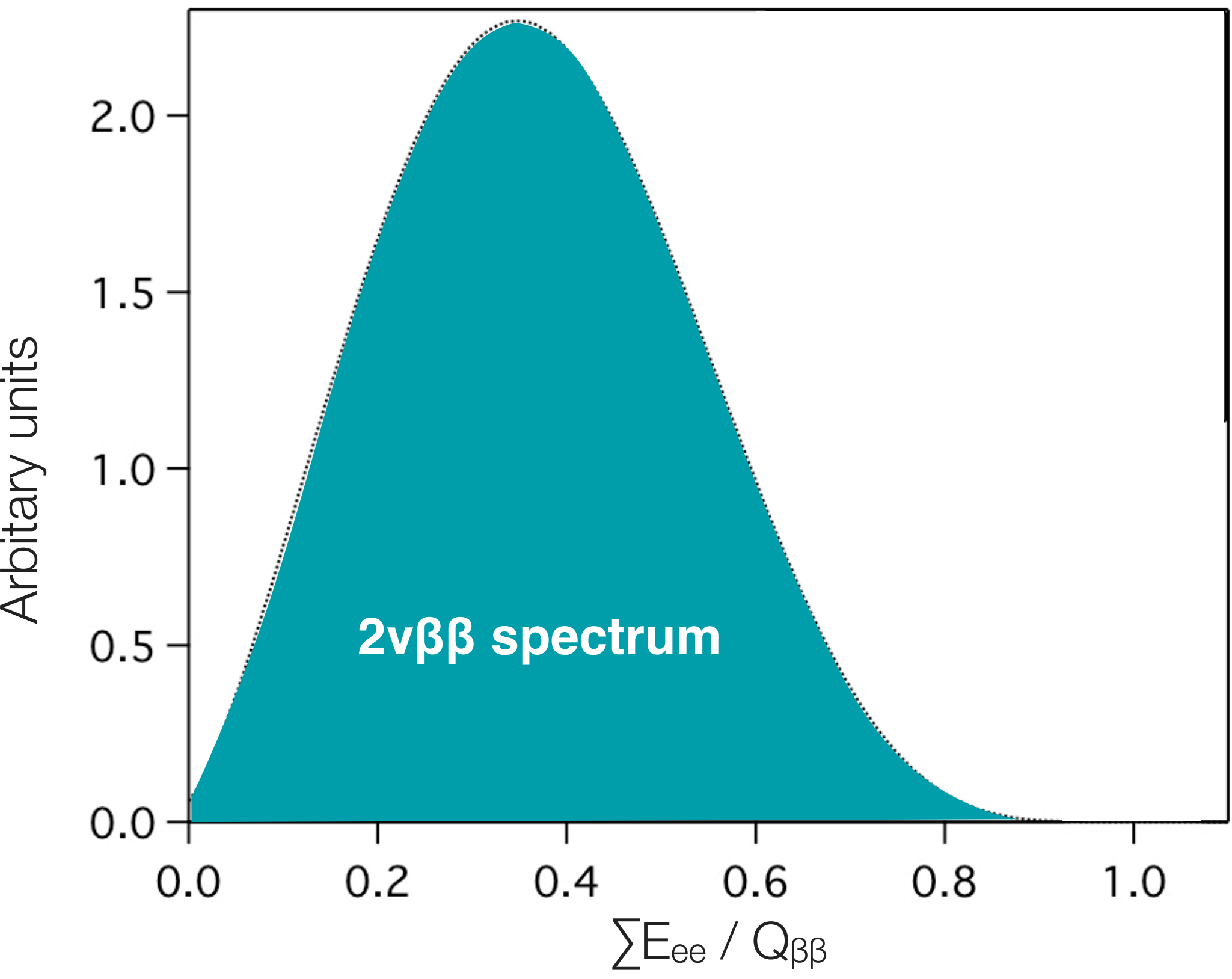
# $0\nu\beta\beta$ could tell us about neutrino mass

$$0\nu\beta\beta \text{ rate} = \frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

Adapted from PRL 117, 082503 (2016)

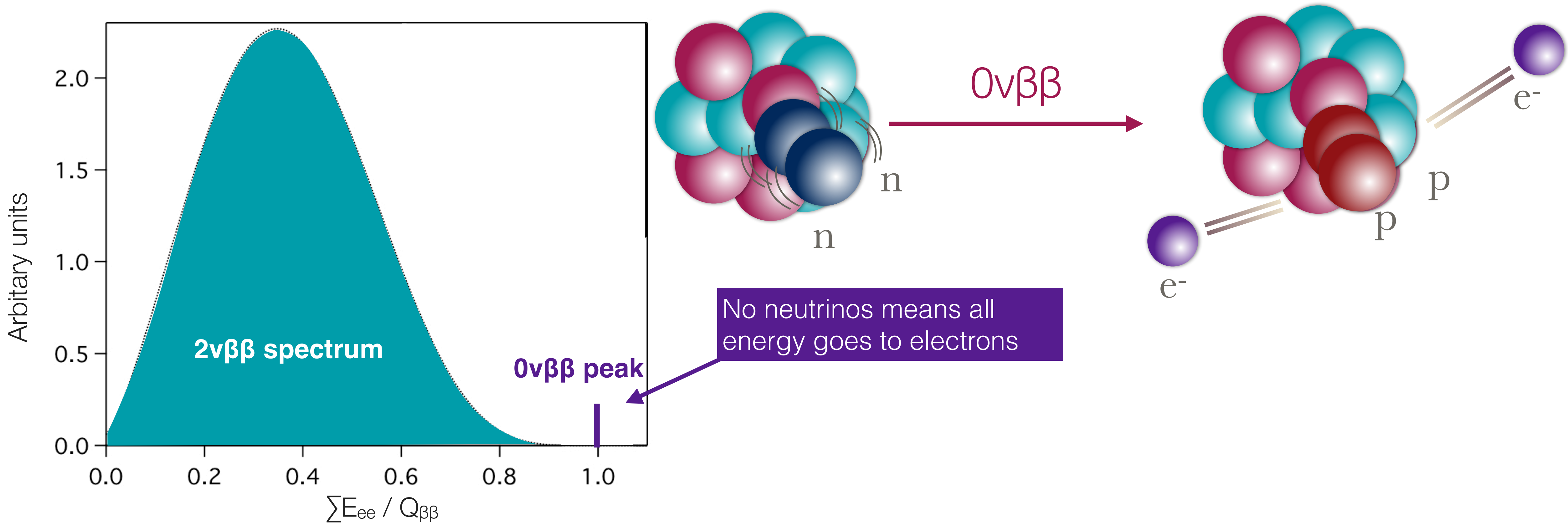


# Looking for $0\nu\beta\beta$



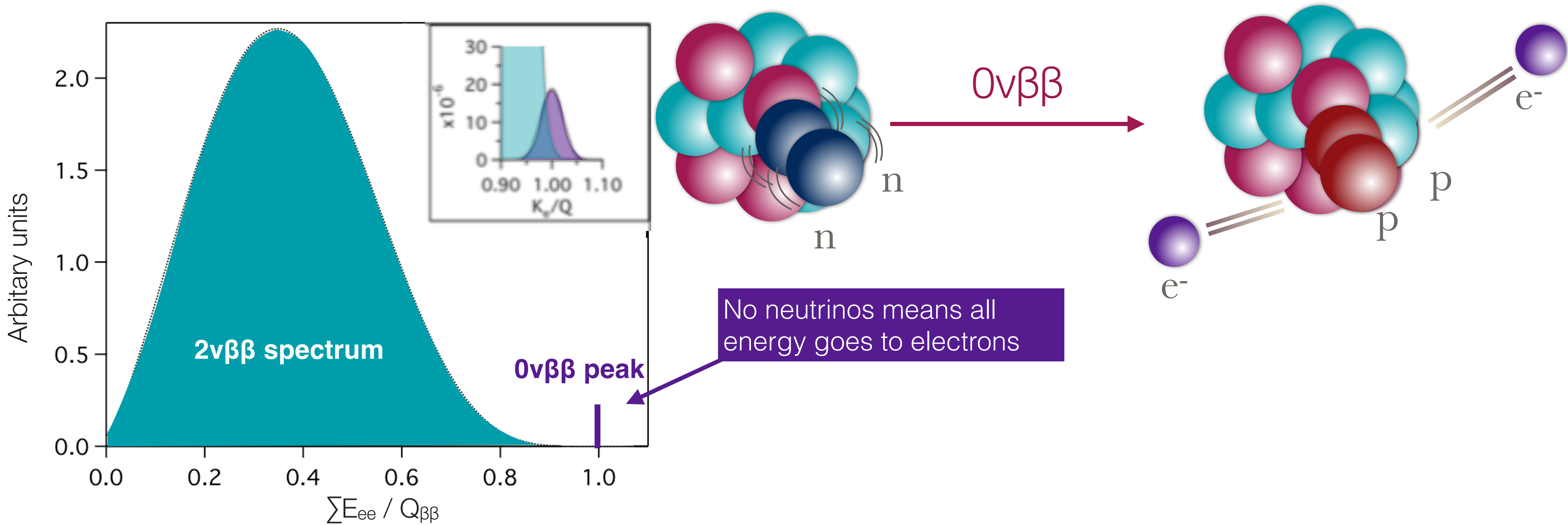
**Sum** of the **2 electron energies**, as fraction of  $\beta\beta$  decay energy

# Looking for $0\nu\beta\beta$



**Sum** of the **2 electron energies**, as fraction of  $\beta\beta$  decay energy

# Looking for $0\nu\beta\beta$



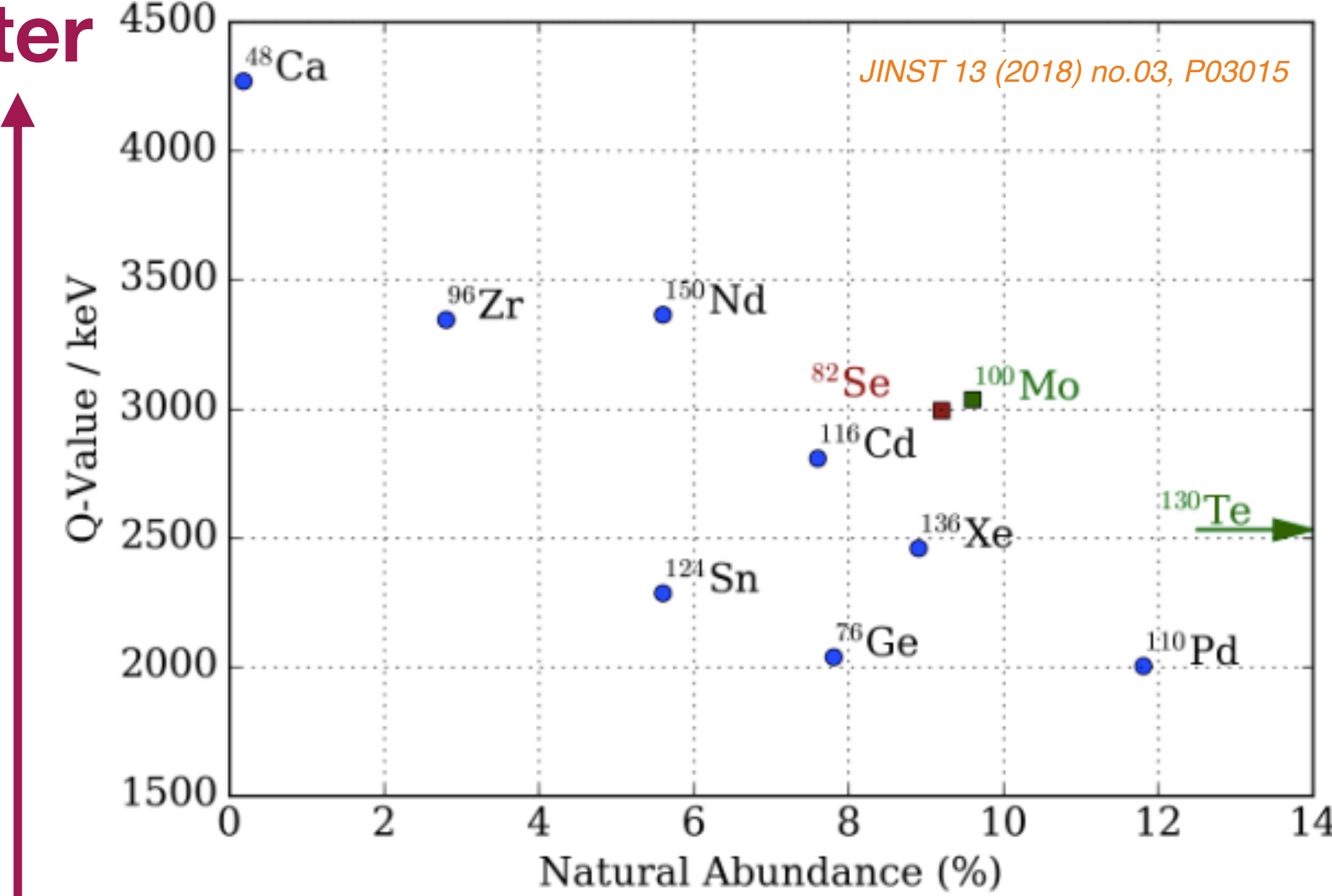
**Sum** of the **2 electron energies**, as fraction of  $\beta\beta$  decay energy

# Choosing an isotope

- “Short” half-life

$$\frac{1}{T_{1/2}^{0\nu\beta\beta}} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

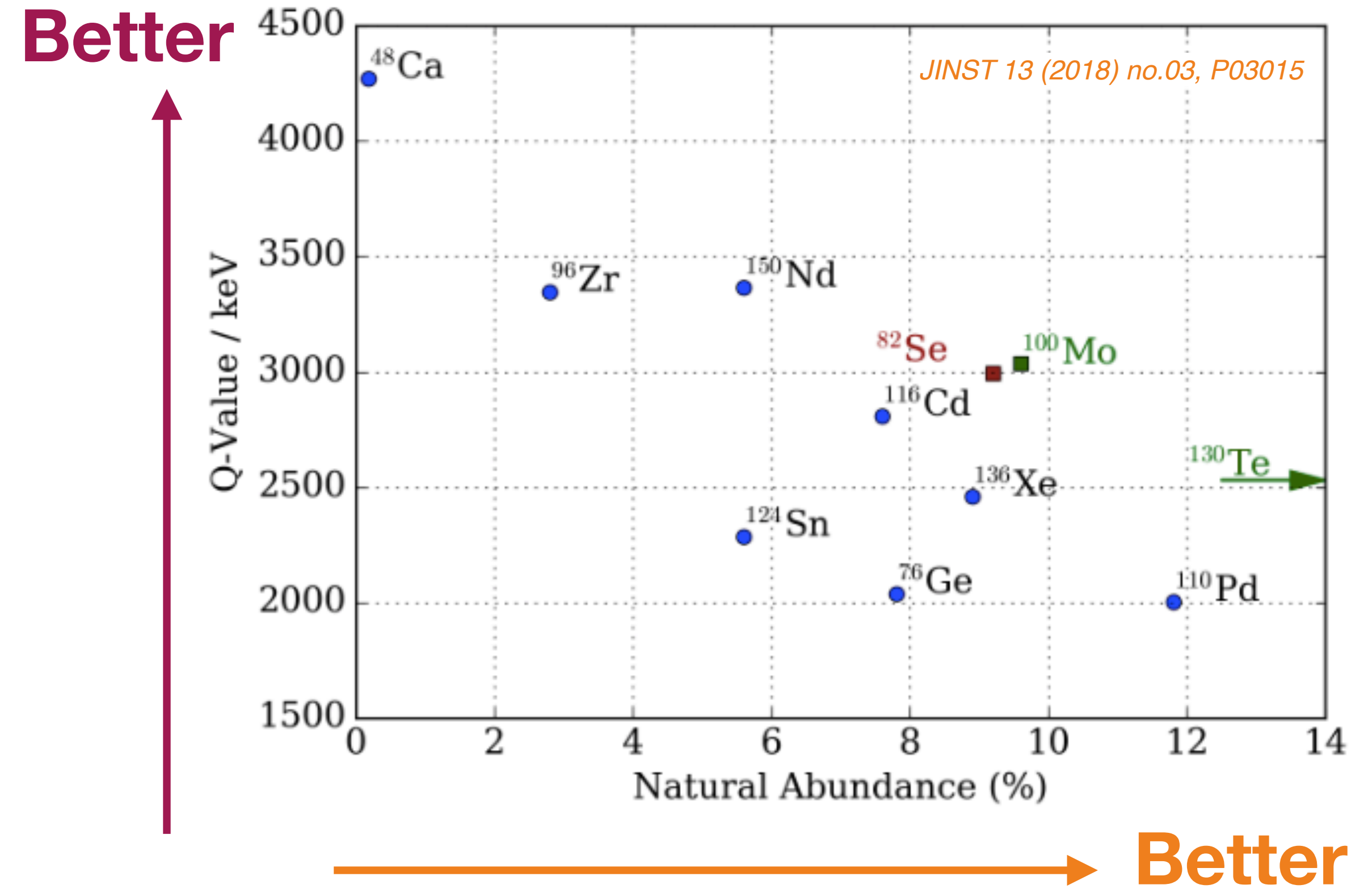
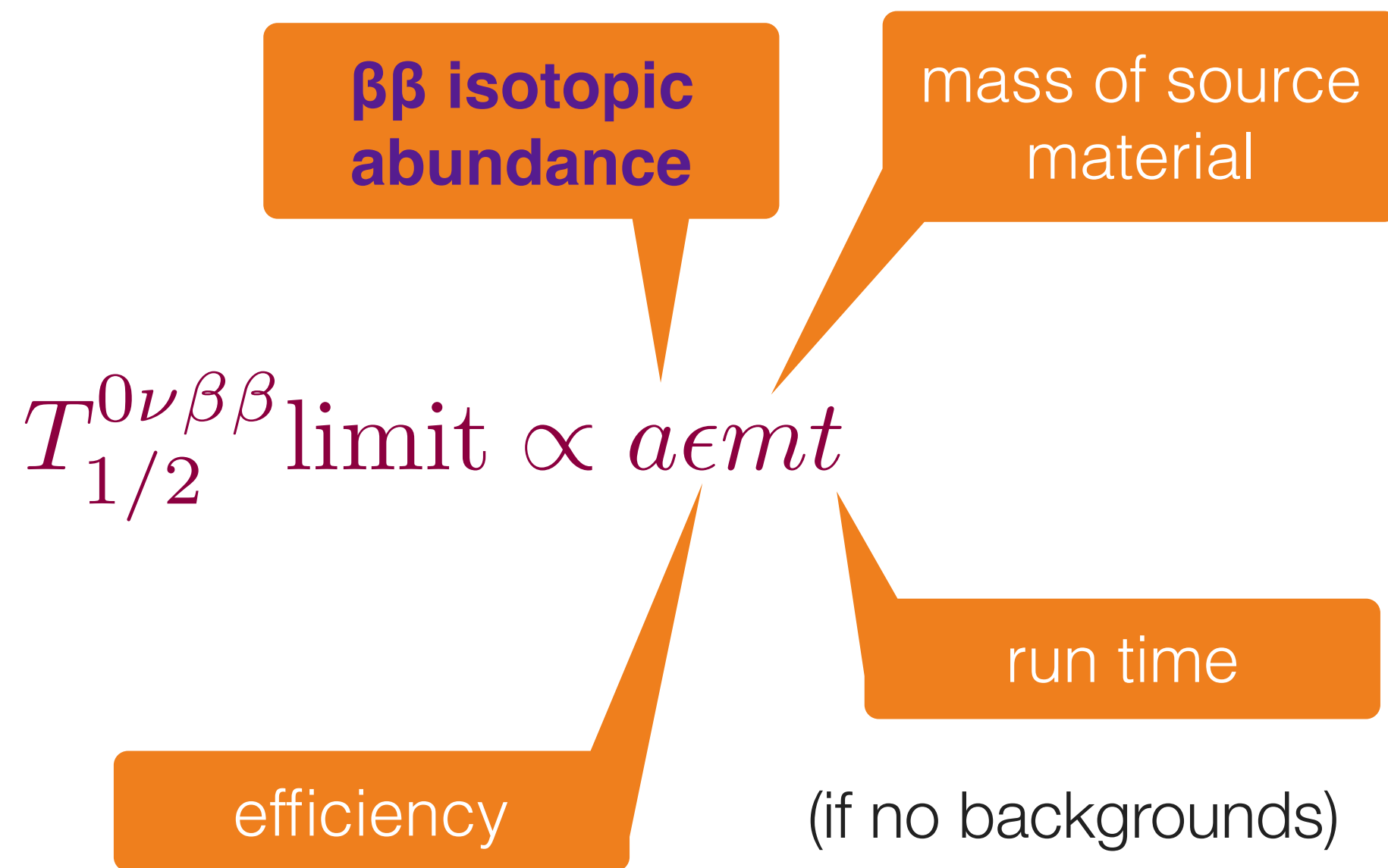
Better



$$T_{1/2}^{0\nu} \propto Q^{-5}$$

# Choosing an isotope

- “Short” half-life
- Lots of isotope



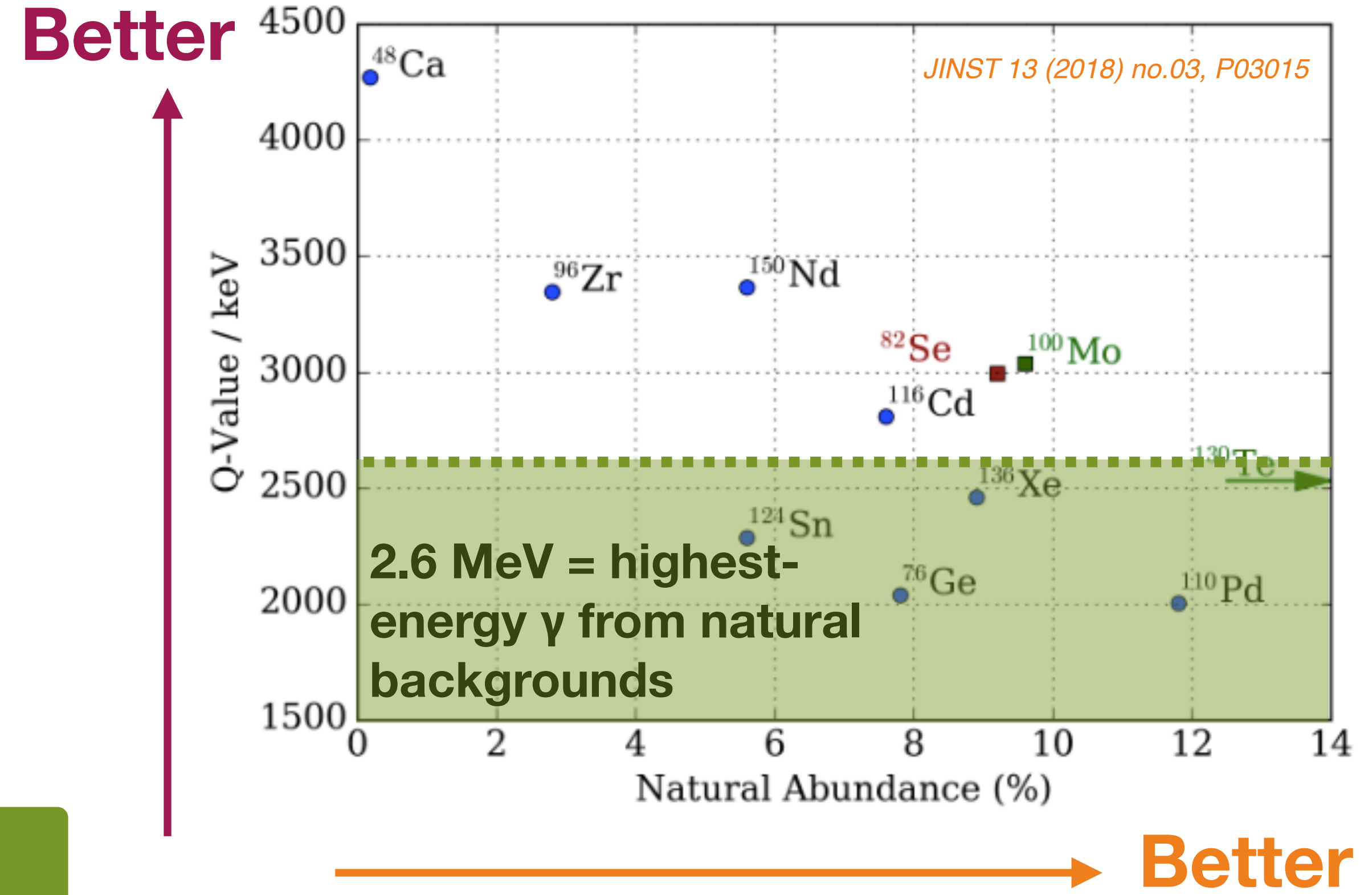
# Choosing an isotope

- “Short” half-life
- Lots of isotope
- (Ultra-) Low backgrounds

$$T_{1/2}^{0\nu} \text{ limit} \propto a \epsilon \sqrt{\frac{mt}{b \Delta E}}$$

background events

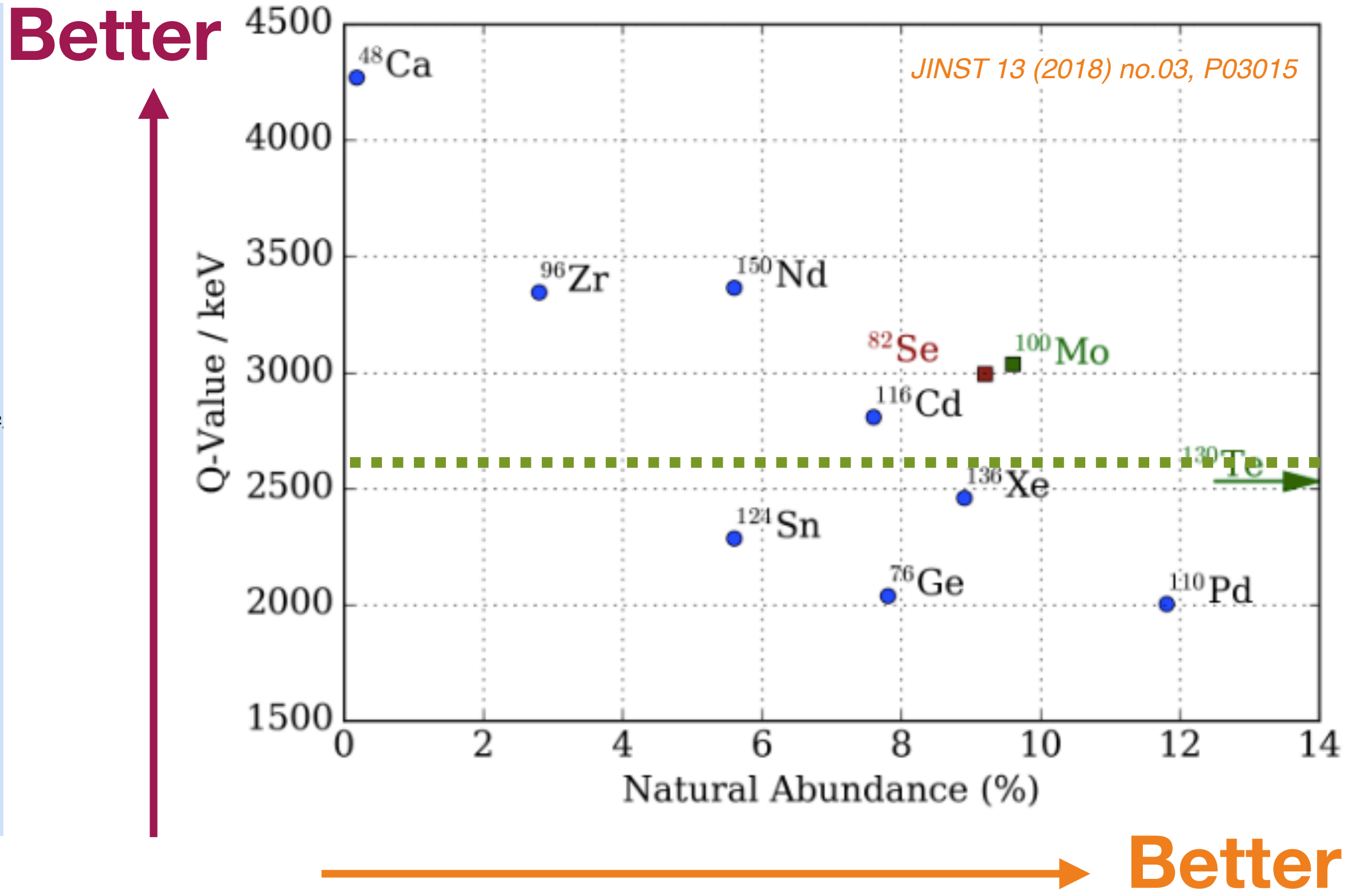
energy resolution



Backgrounds include

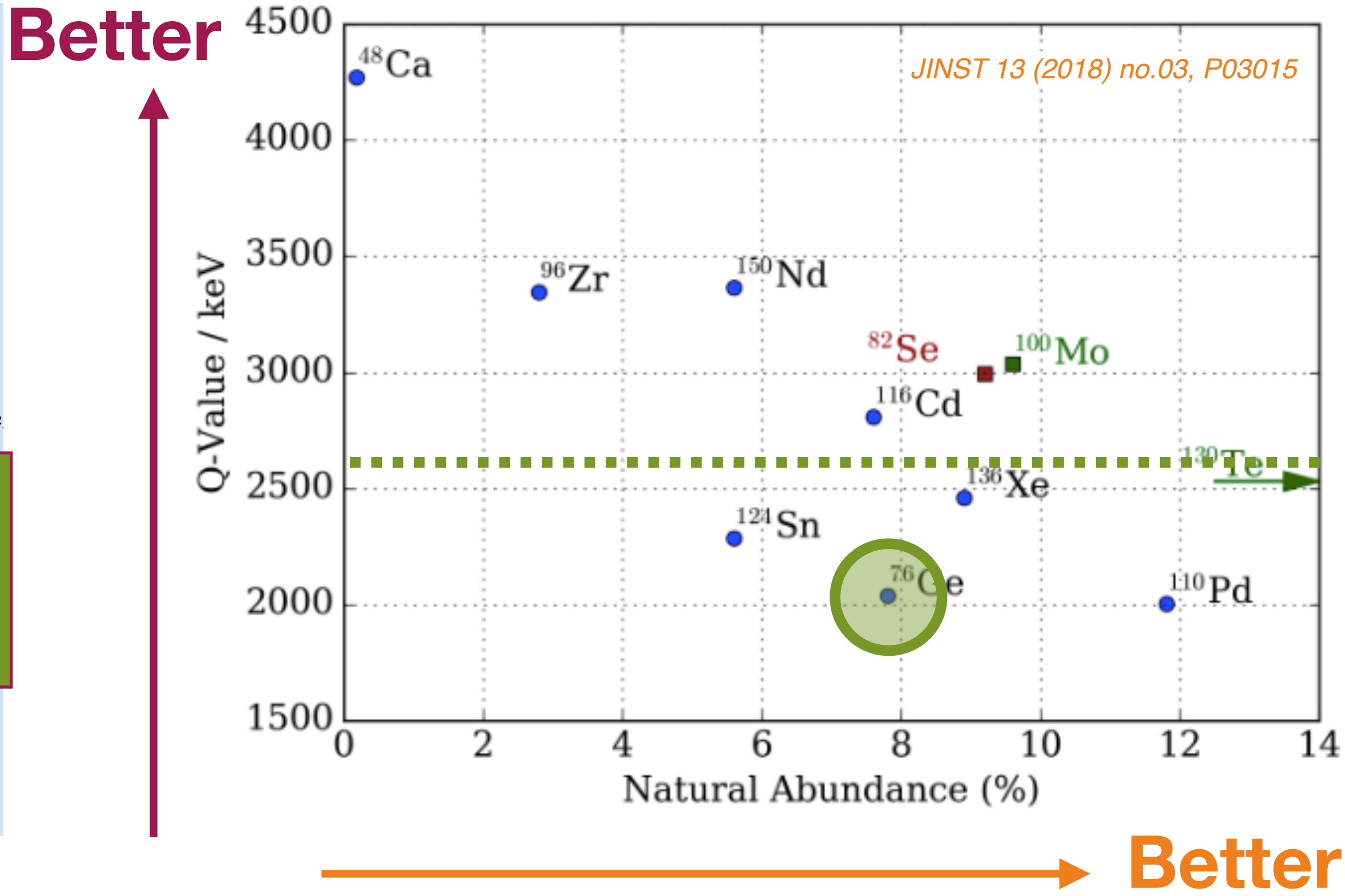
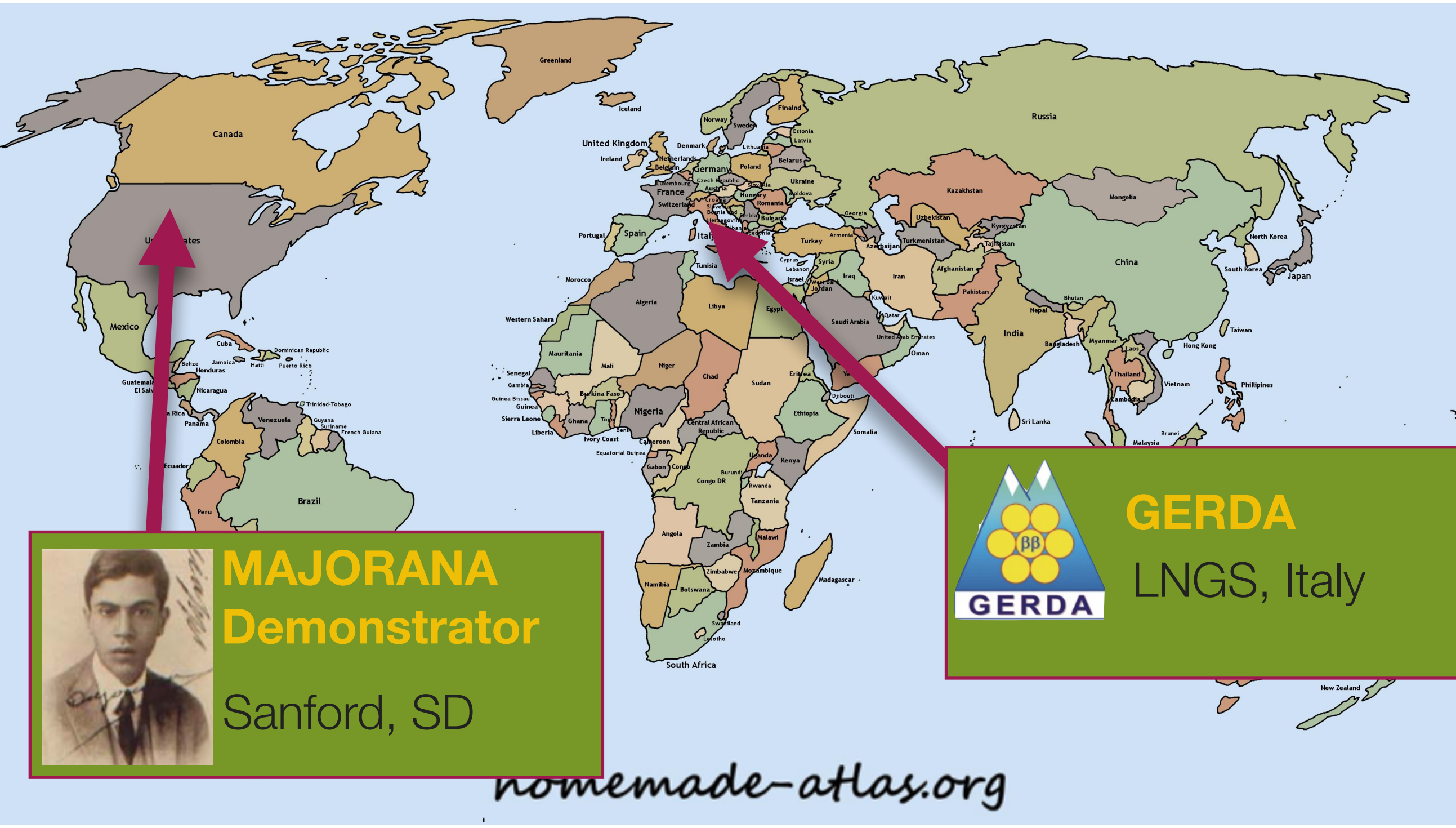
- <sup>214</sup>Bi and <sup>208</sup>Tl - β-emitting daughters of U & Th
- Particular danger from **radon**
- Irreducible **2νββ** background

# Choosing an isotope



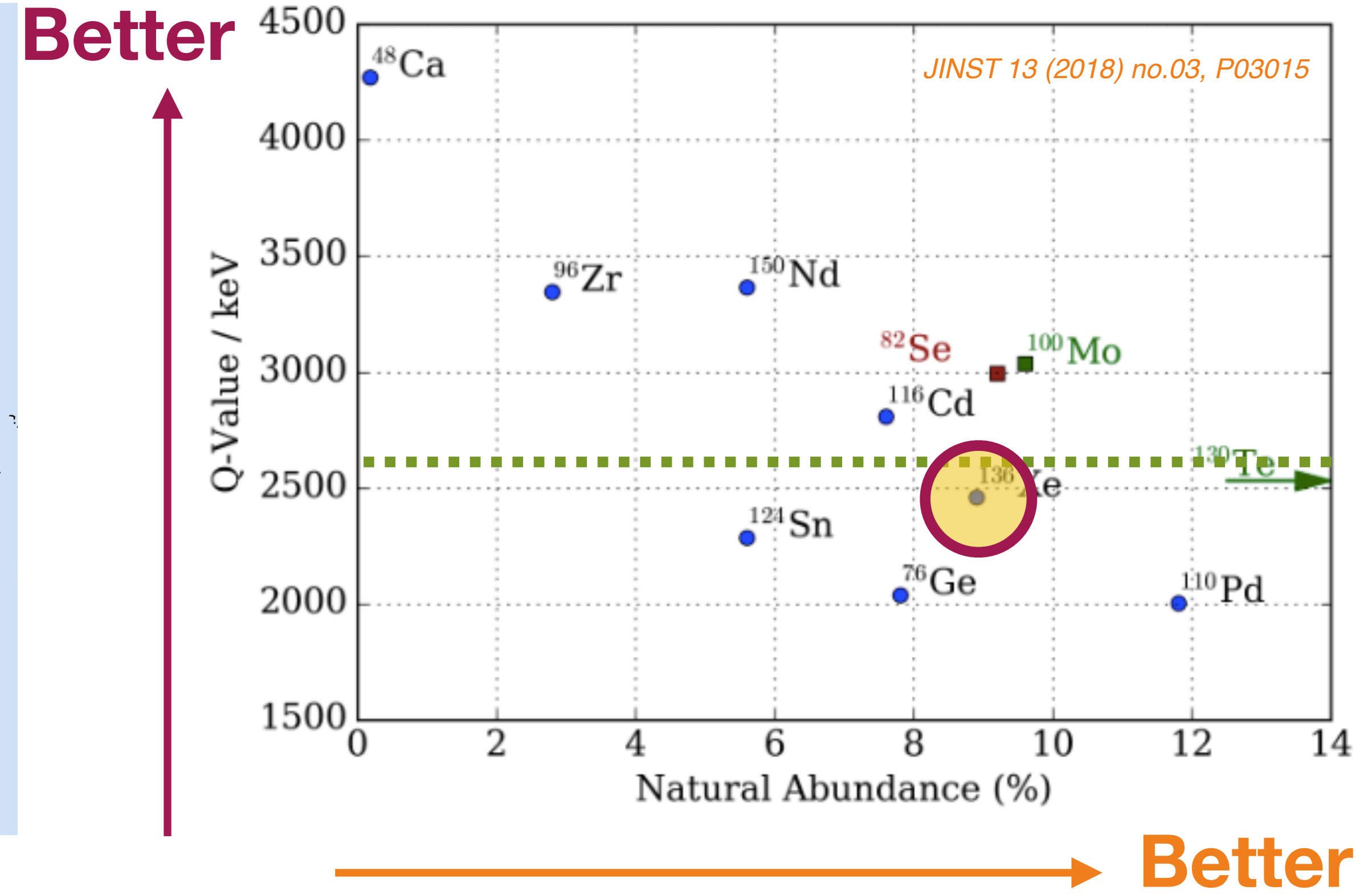
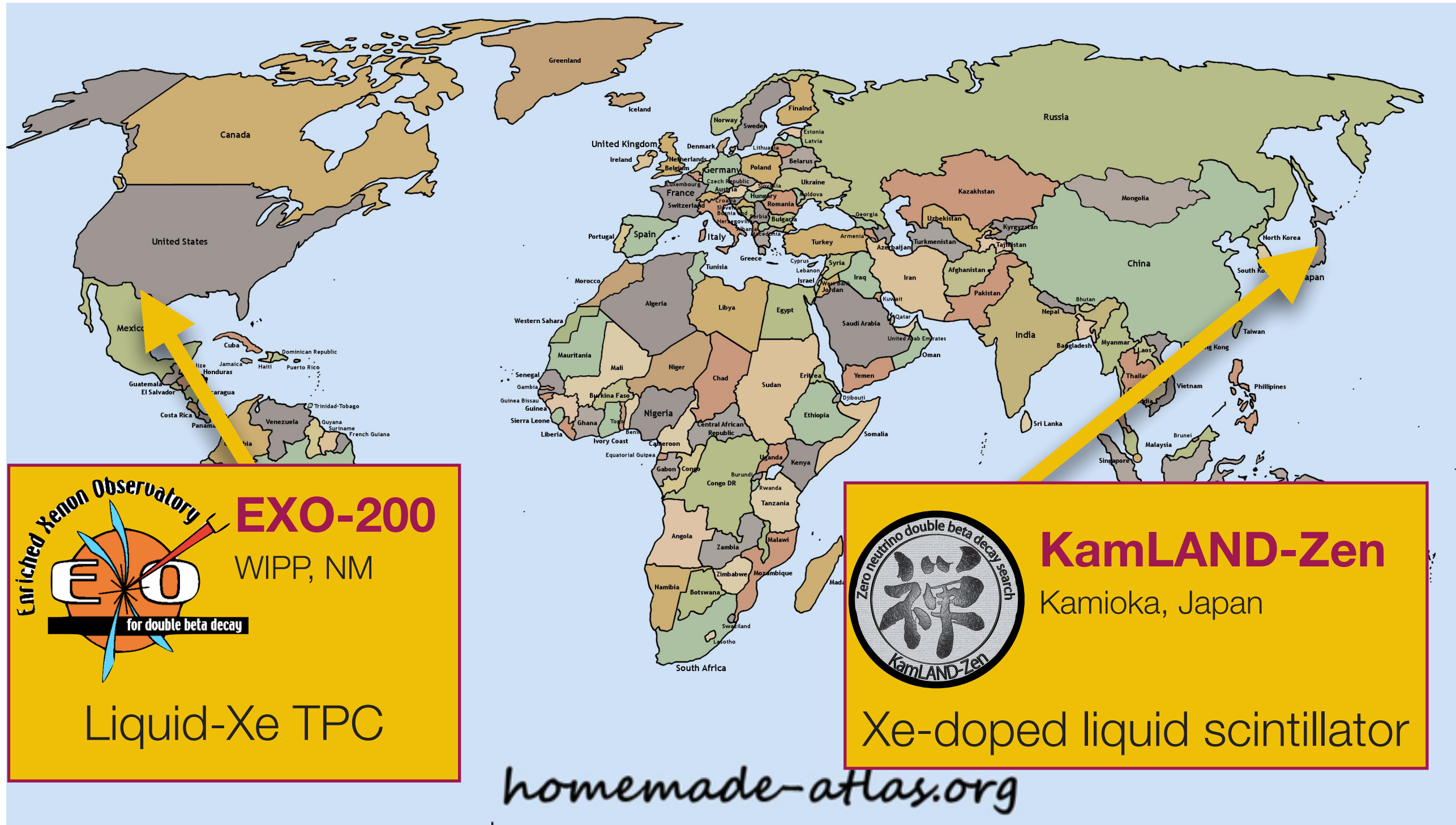


# Choosing an isotope



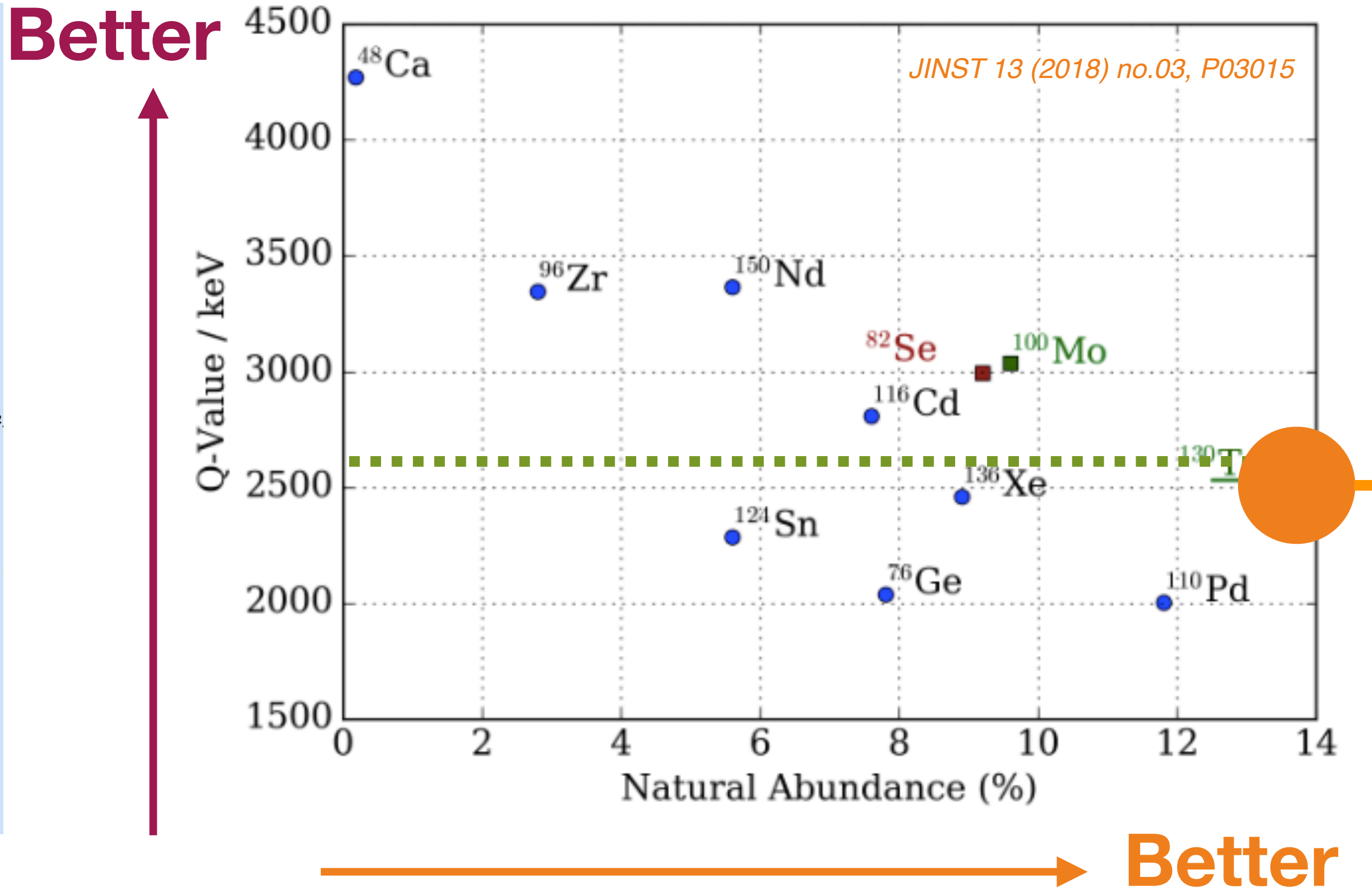
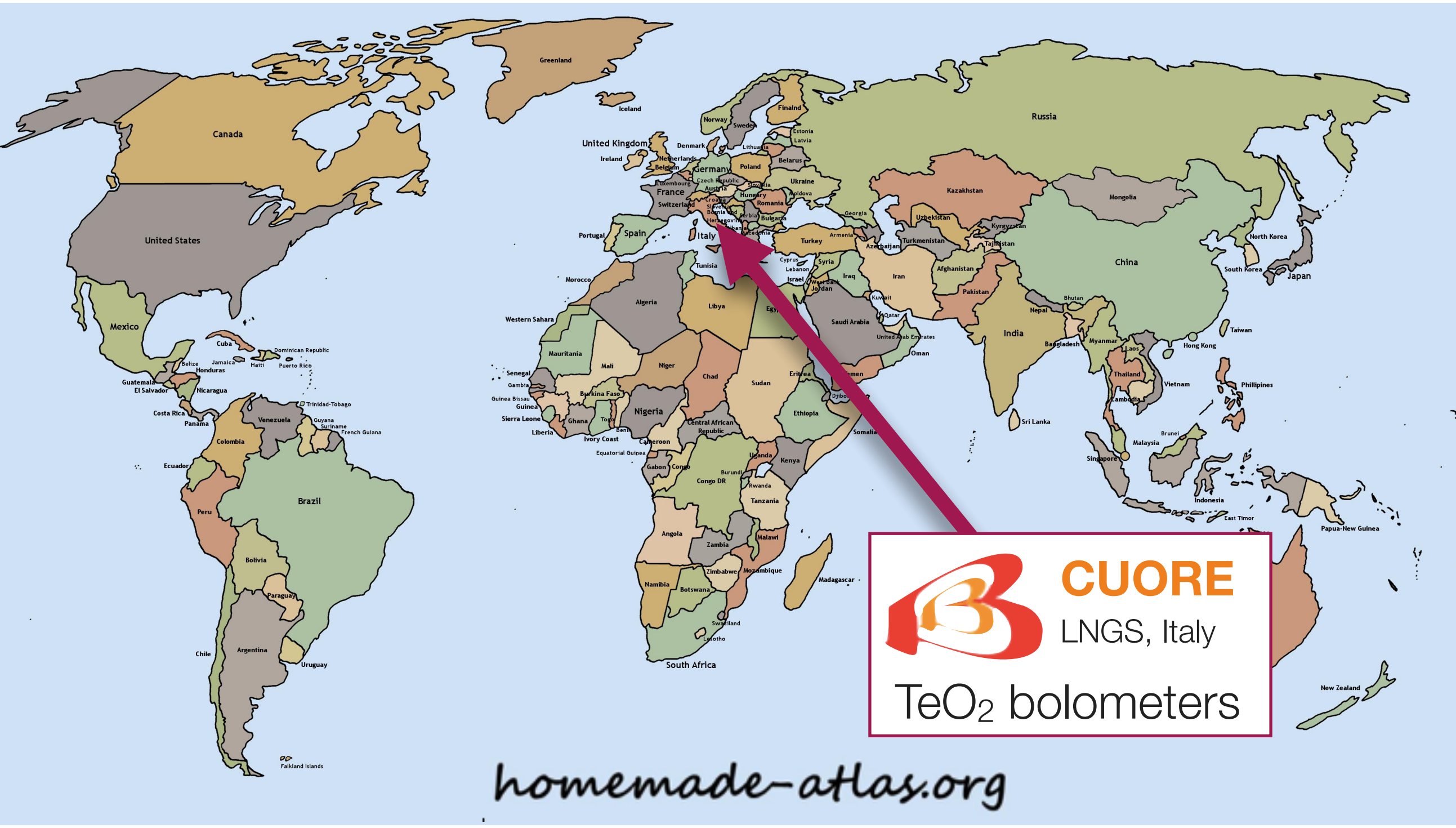
- High-purity germanium **detector** array is **also  $\beta\beta$  source** (<sup>76</sup>Ge)
- Excellent **efficiency** and **resolution**: zero background in  $0\nu\beta\beta$  region of interest
- Future detector: LEGEND

# Choosing an isotope



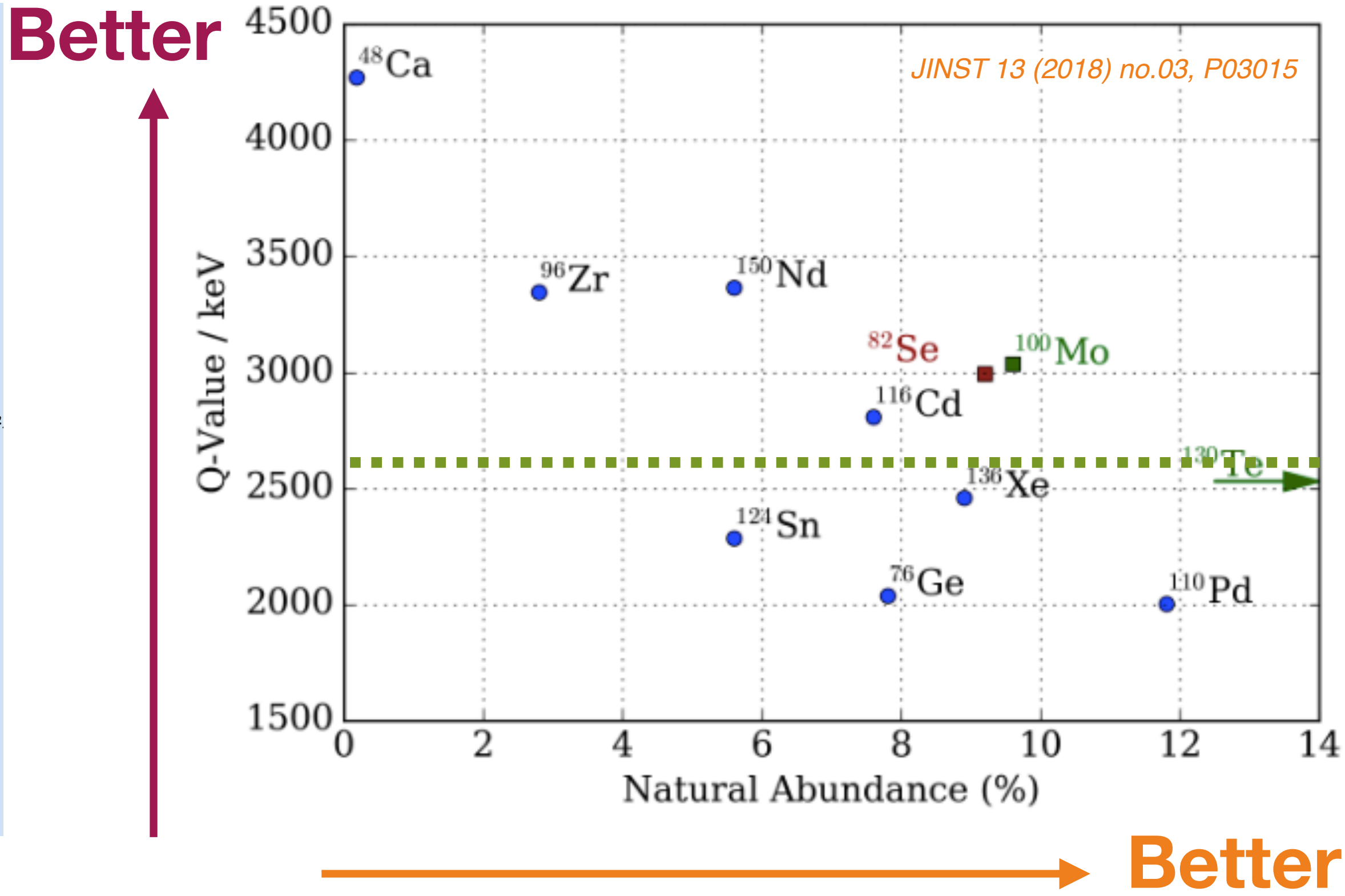
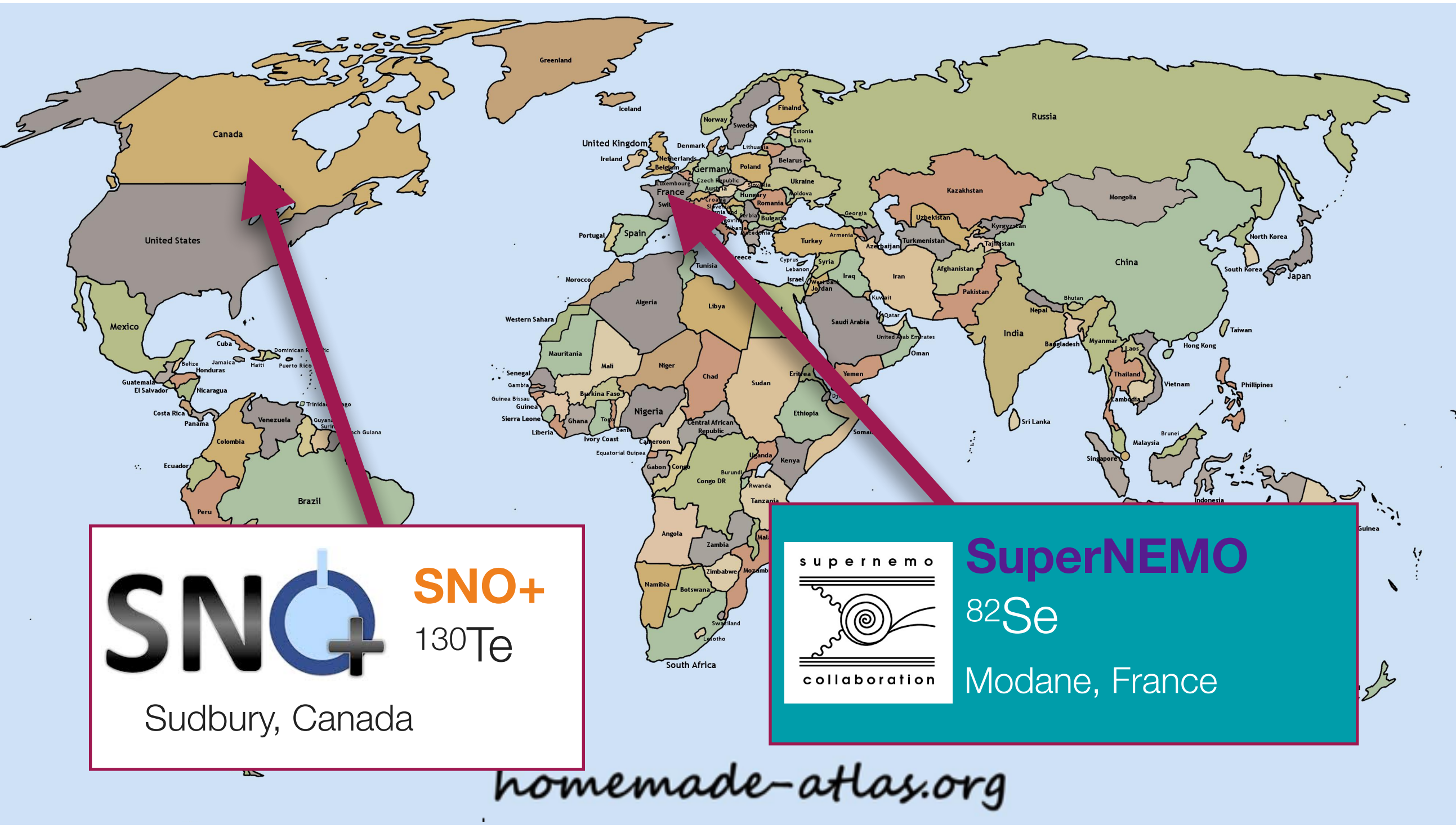
- Large detectors with **hundreds of kg** of isotope (<sup>136</sup>Xe)
- KamLAND-Zen has current best  $0\nu\beta\beta$  half-life /  $m_{\beta\beta}$  mass limit
- $T_{1/2} > 1.07 \times 10^{26}$  years (  $\langle m_{\beta\beta} \rangle < 61-165$  meV )
- Future detectors - nEXO, KamLAND2 Zen

# Choosing an isotope



- <sup>130</sup>Te has 34% natural **abundance**
- TeO<sub>2</sub> crystals at 10mK heat up when decay occurs
- 0.2% energy **resolution**
- **CUPID** adds particle ID

# Choosing an isotope



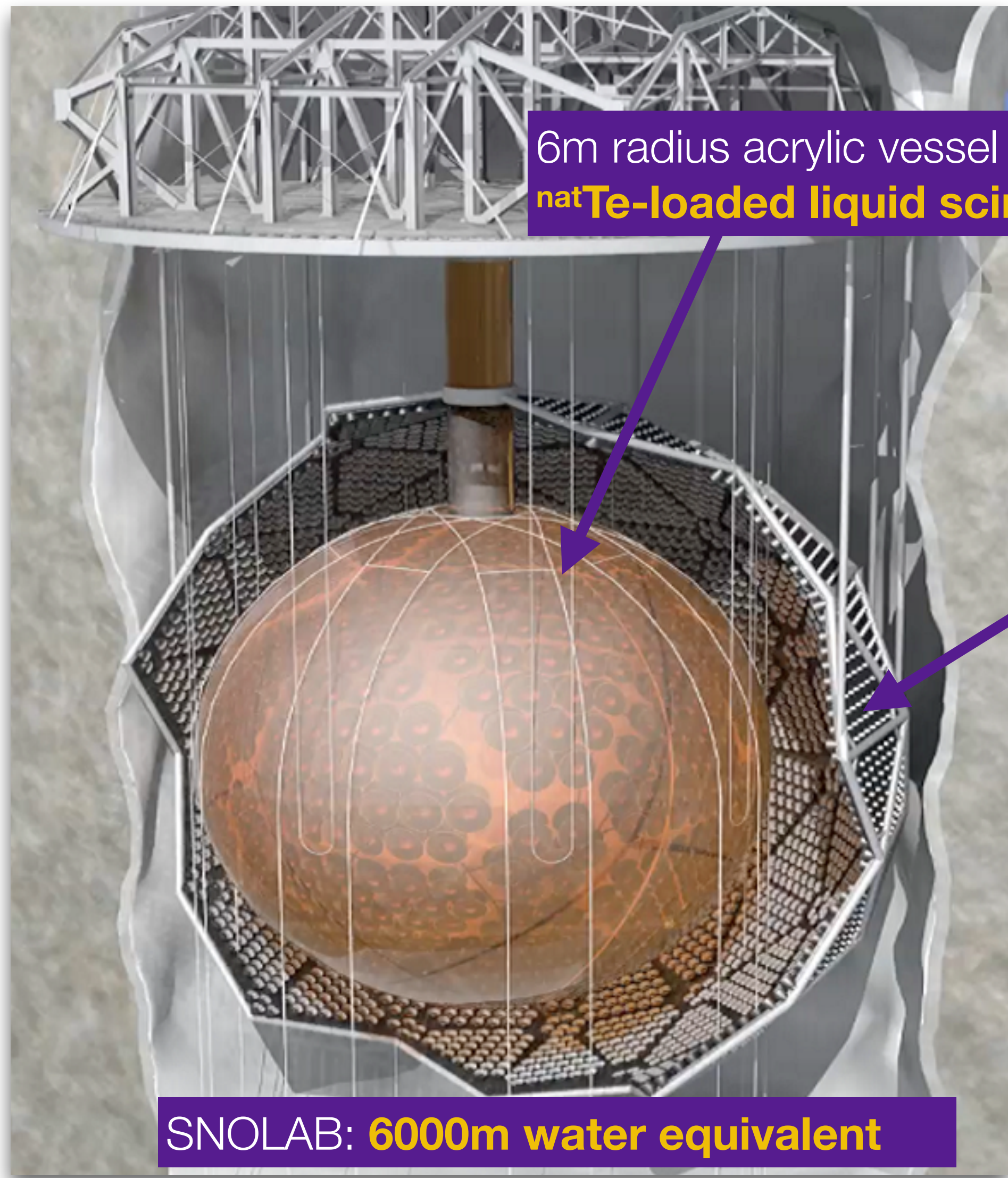
**Significant UK involvement**

# SNO+ at Sudbury, Canada



Thanks to Steve Biller & Esther Turner for slide content

# SNO+ - a new approach for $0\nu\beta\beta$

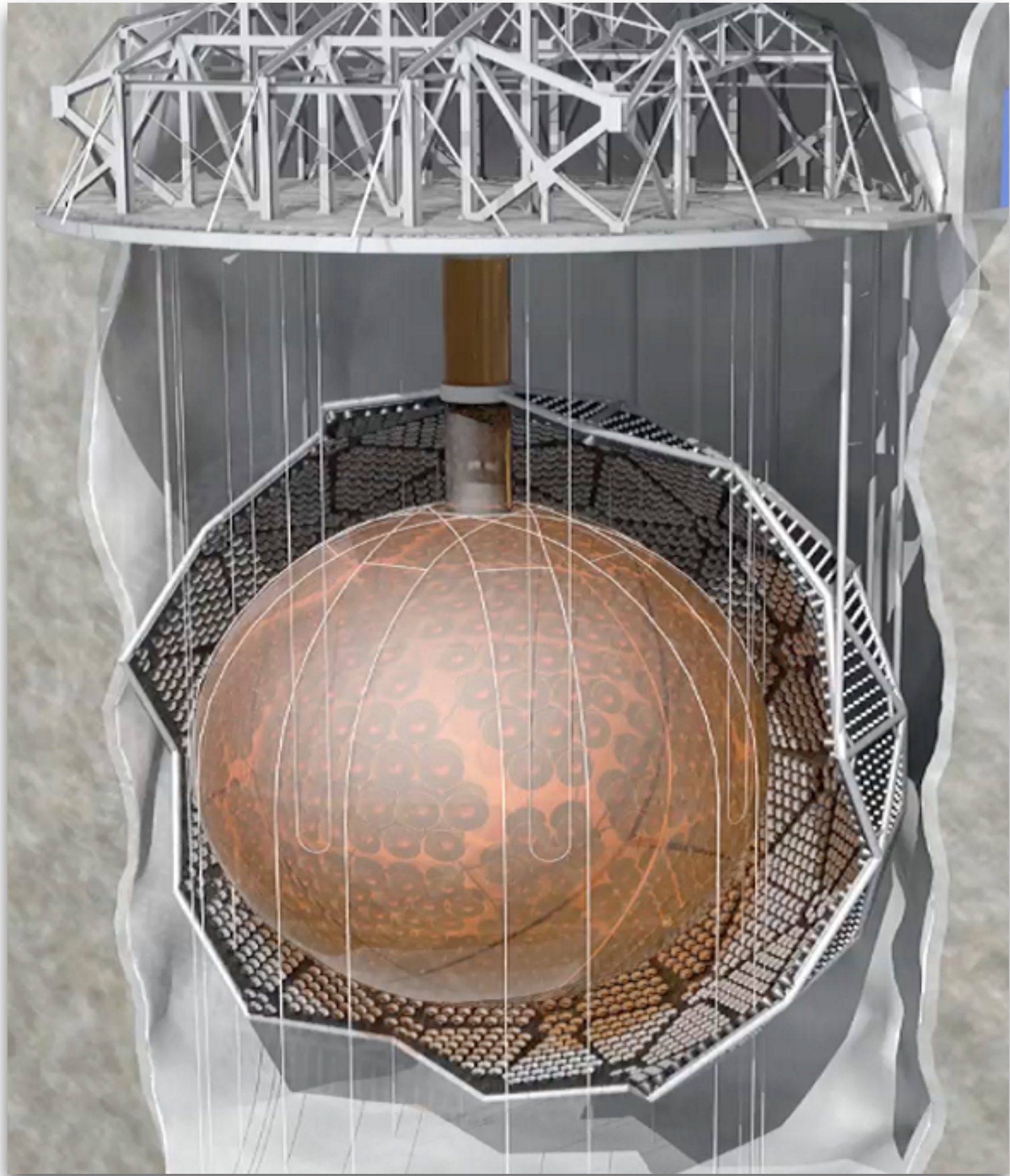


6m radius acrylic vessel filled with **natTe-loaded liquid scintillator**

Surrounded by ~**9300 PMTs** in a water-filled cavity

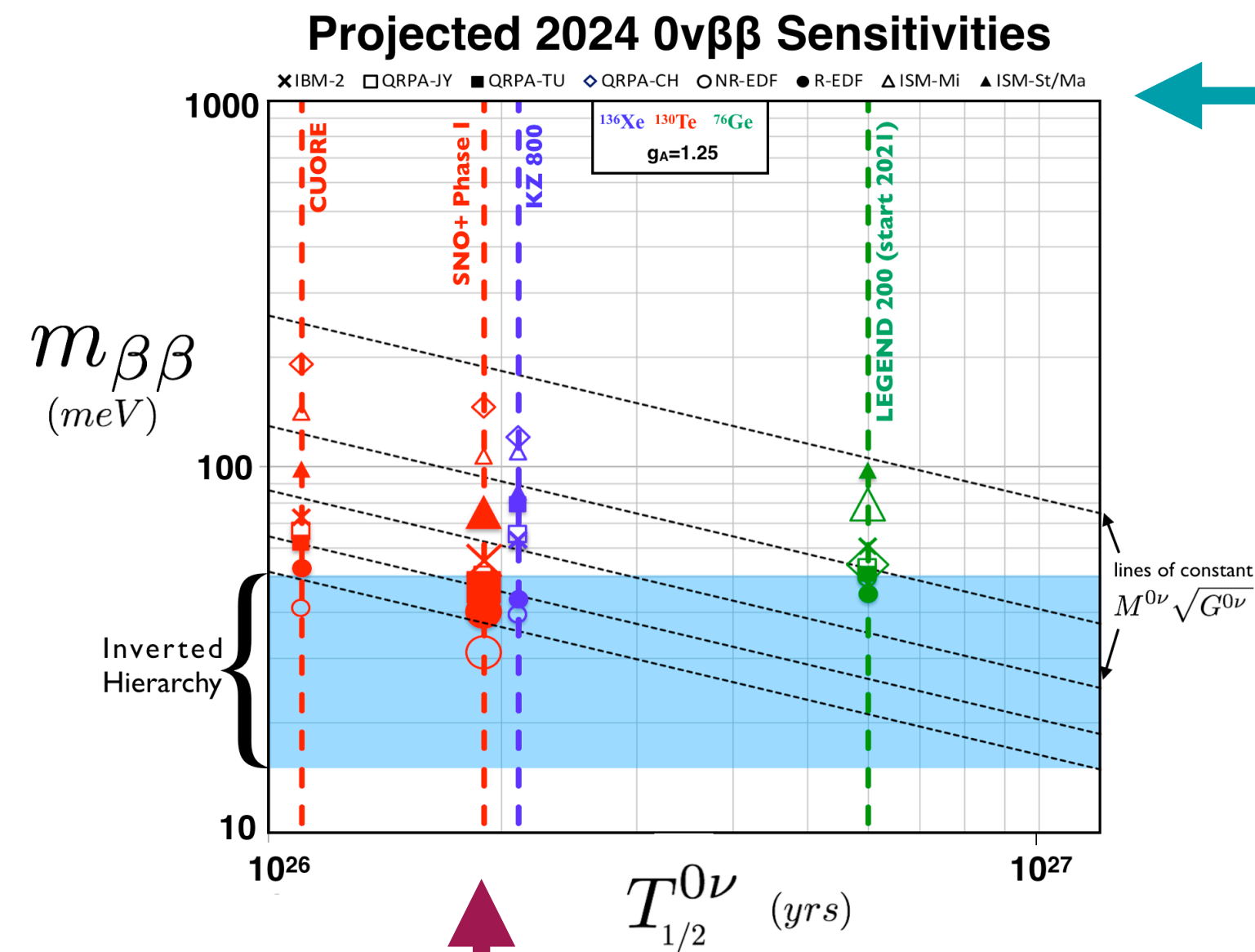
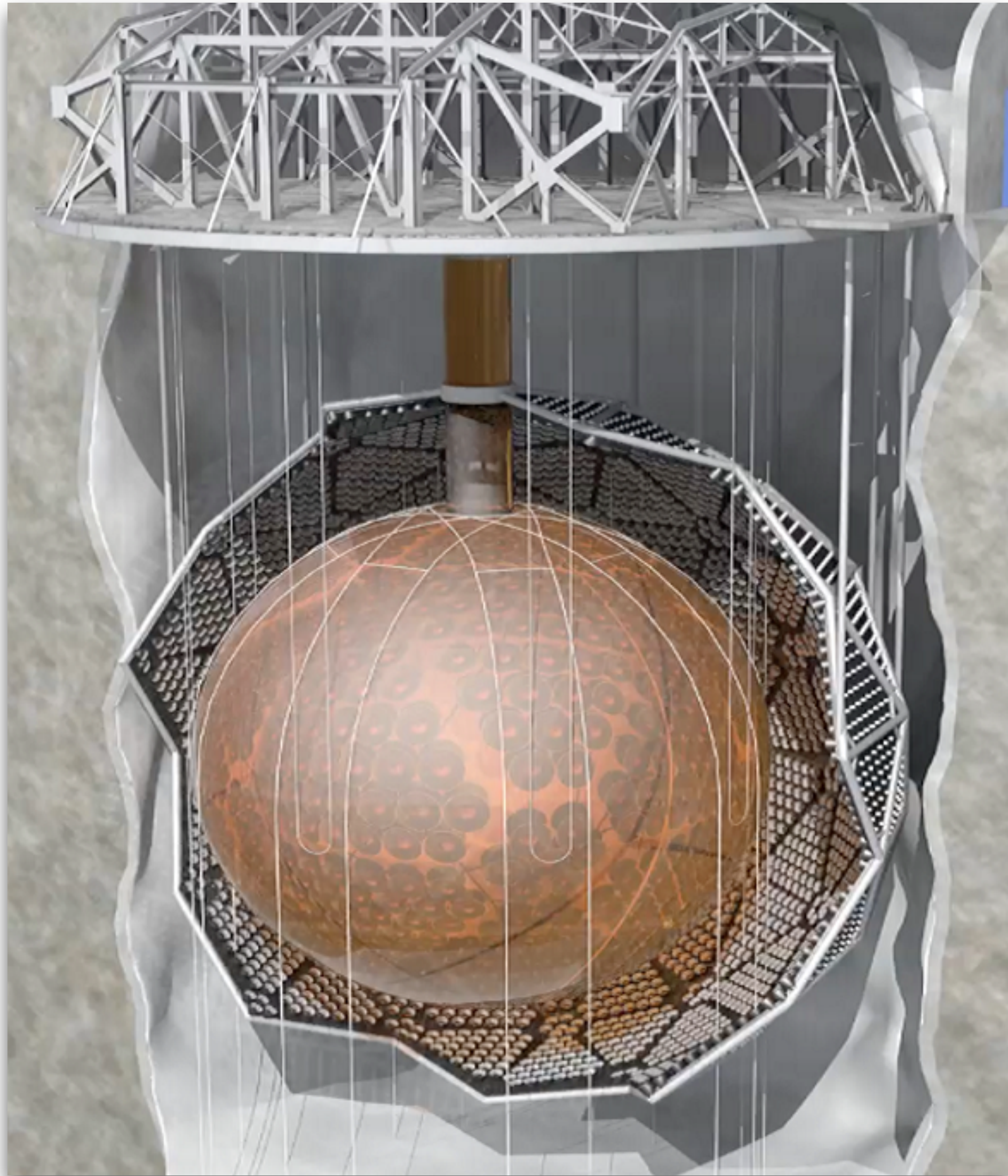
SNOLAB: **6000m water equivalent**

# SNO+ - a new approach for $0\nu\beta\beta$



- Highly **economical**
  - $^{130}\text{Te}$  is the most economically scalable **isotope** (high natural abundance);
  - Liquid scintillator also very economically scalable **detector technology!**
- Potential for dramatic **scale-up**

# SNO+ - a new approach for $0\nu\beta\beta$

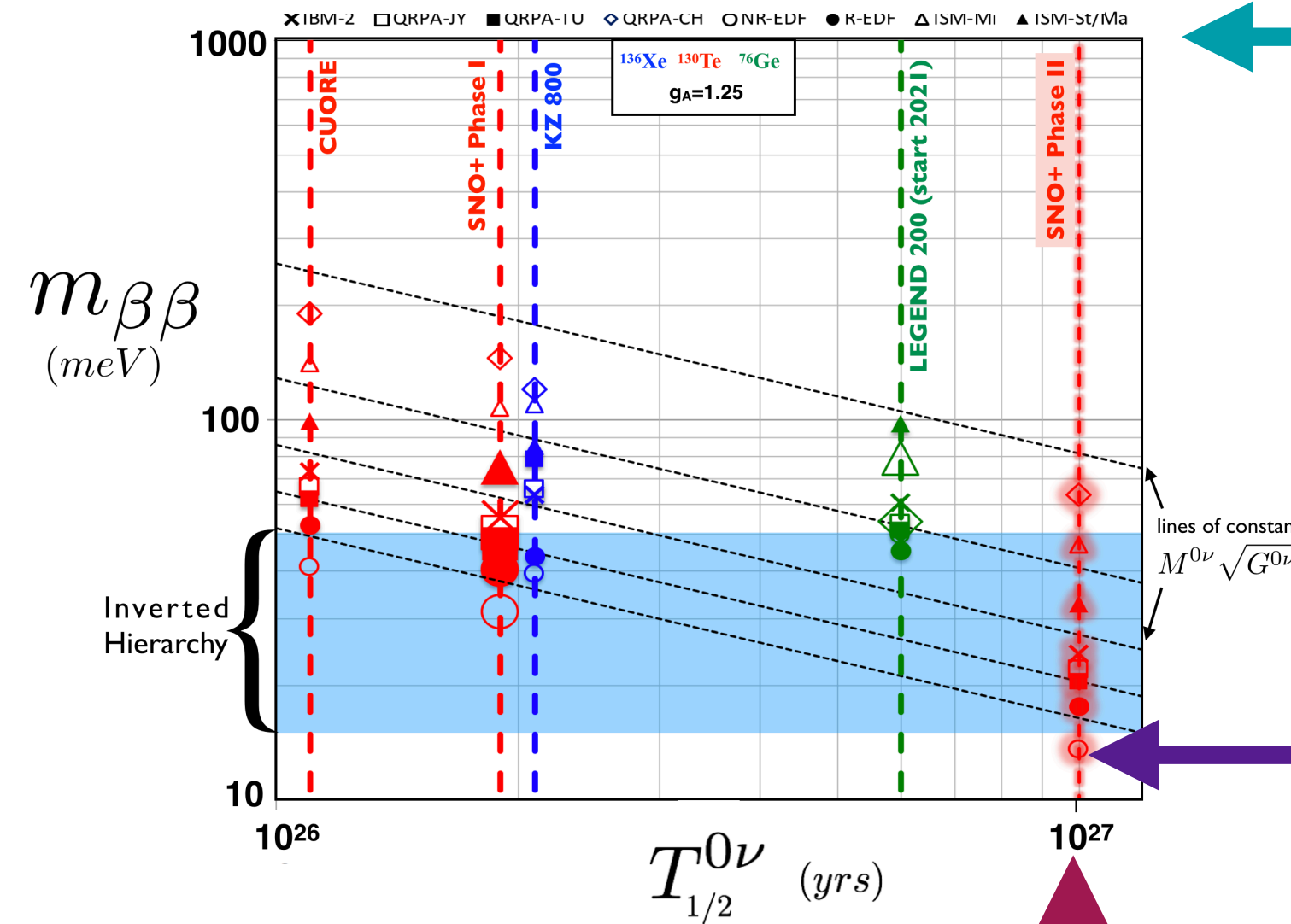
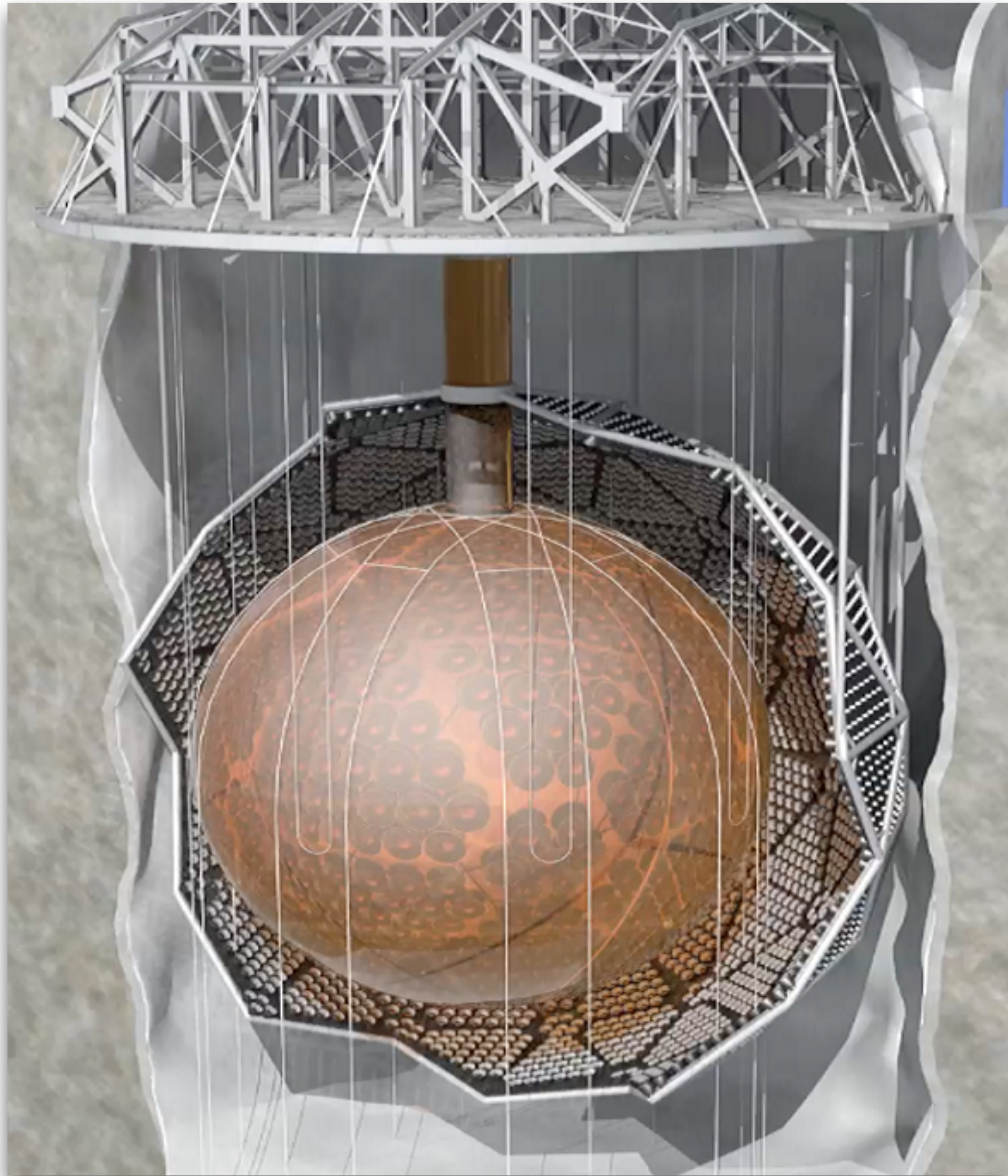


Different matrix element calculations give different masses for the same half-life

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  - Liquid scintillator also very economically scalable **detector technology!**
- Potential for dramatic **scale-up**
- Allows **sensitivity** above current leading measurement:
  - $T_{1/2}^{0\nu\beta\beta} > 2.1 \times 10^{26}$  years (  $m_{\beta\beta} < 37-89$  meV ) after 5 years of running



# SNO+ - a new approach for $0\nu\beta\beta$

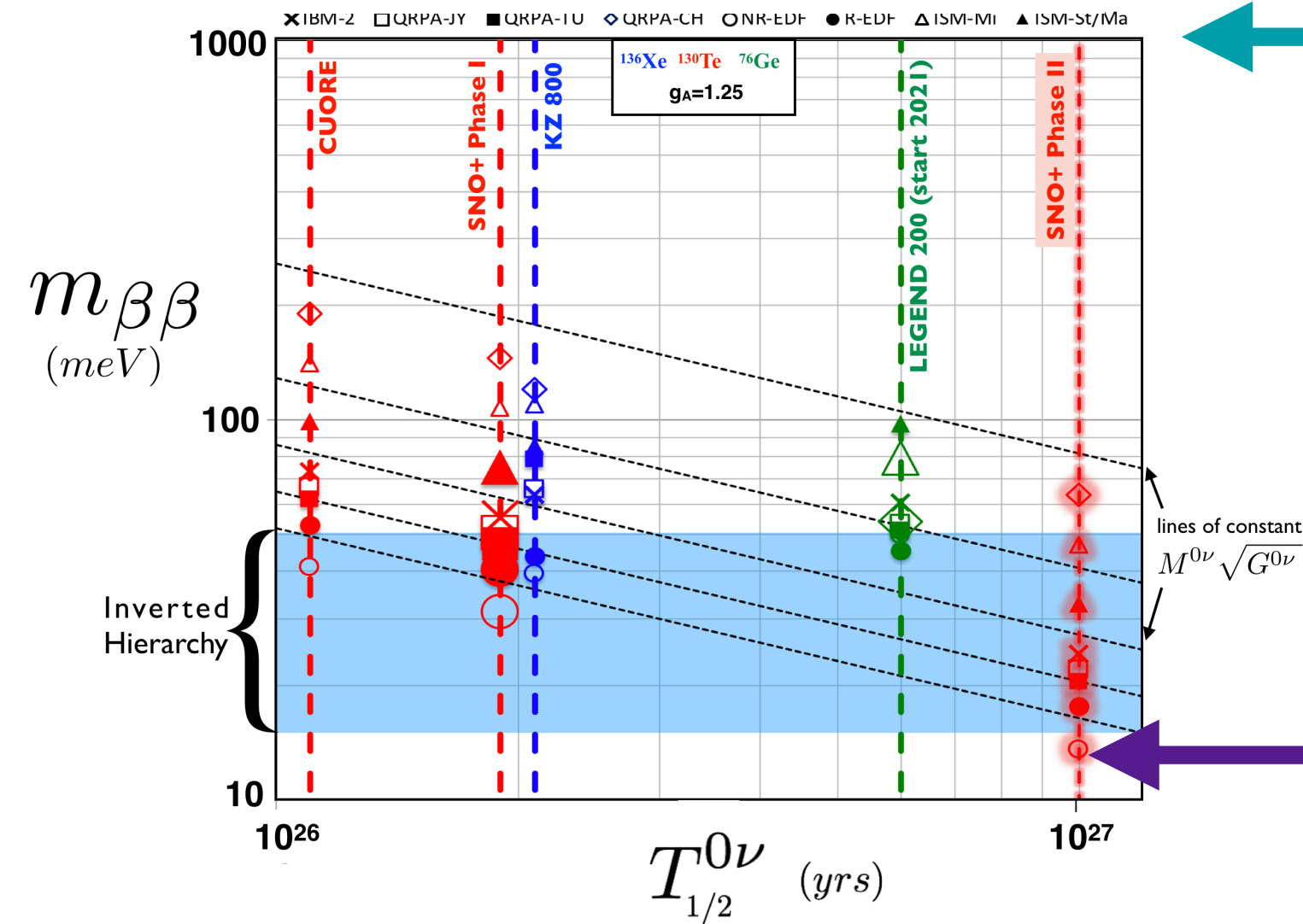
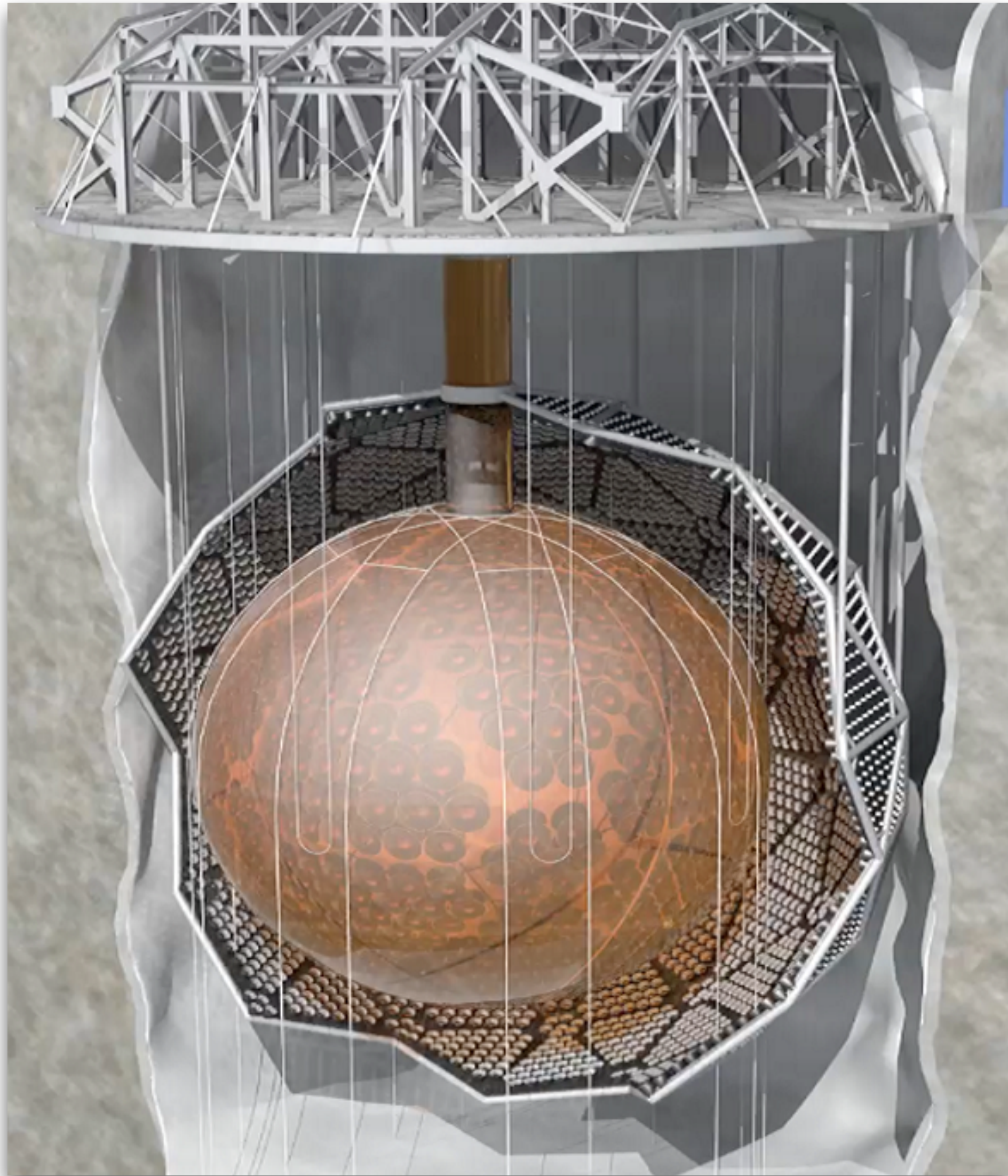


Different matrix element calculations give different masses for the same half-life

New chemistry developments may make it possible to simply increase the loading in the current instrument to achieve this

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# SNO+ - a new approach for $0\nu\beta\beta$




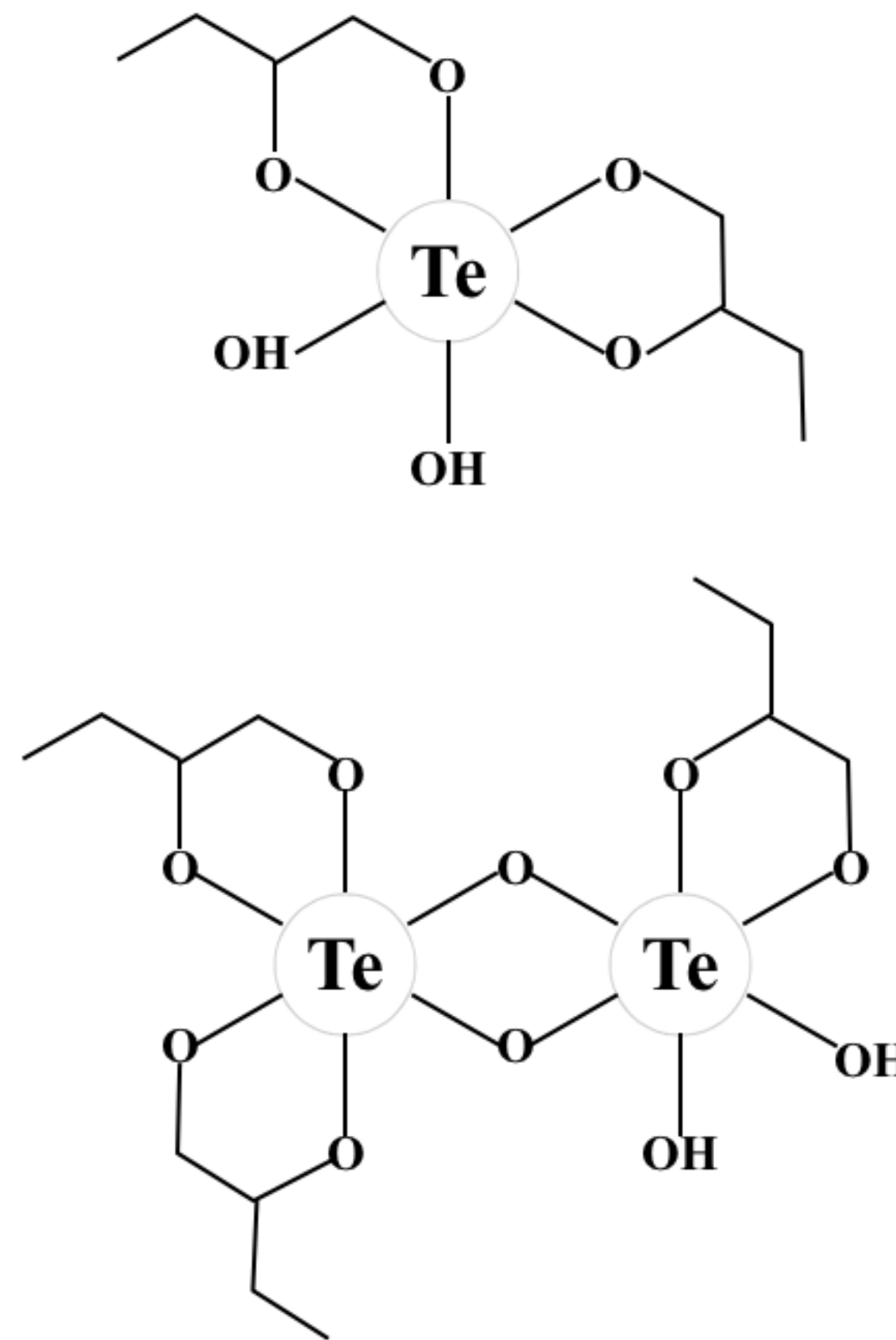
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- **Phase II** could reach  $10^{27}$  years with the same detector but **higher loading**
- Concept originated in **UK** (Biller & Chen, 2012)

# New loading method: Te-butanediol complex dissolves in liquid scintillator

- Simple **synthesis**
- Single **safe**, distillable chemical
- Low **radioactivity** levels
- Minimal optical **absorption**
- High **light levels** at 0.5% <sup>nat</sup>Te loading
- Developed in UK! 



Natural tellurium is 34% <sup>130</sup>Te



See Tereza Kroupova's slides!

- Operating with **water** from 2017

- Invisible nucleon decay
- Solar neutrinos
- Supernova neutrinos

PHYSICAL REVIEW D **99**, 012012 (2019)

### Measurement of the $^8\text{B}$ solar neutrino flux in SNO+ with very low backgrounds

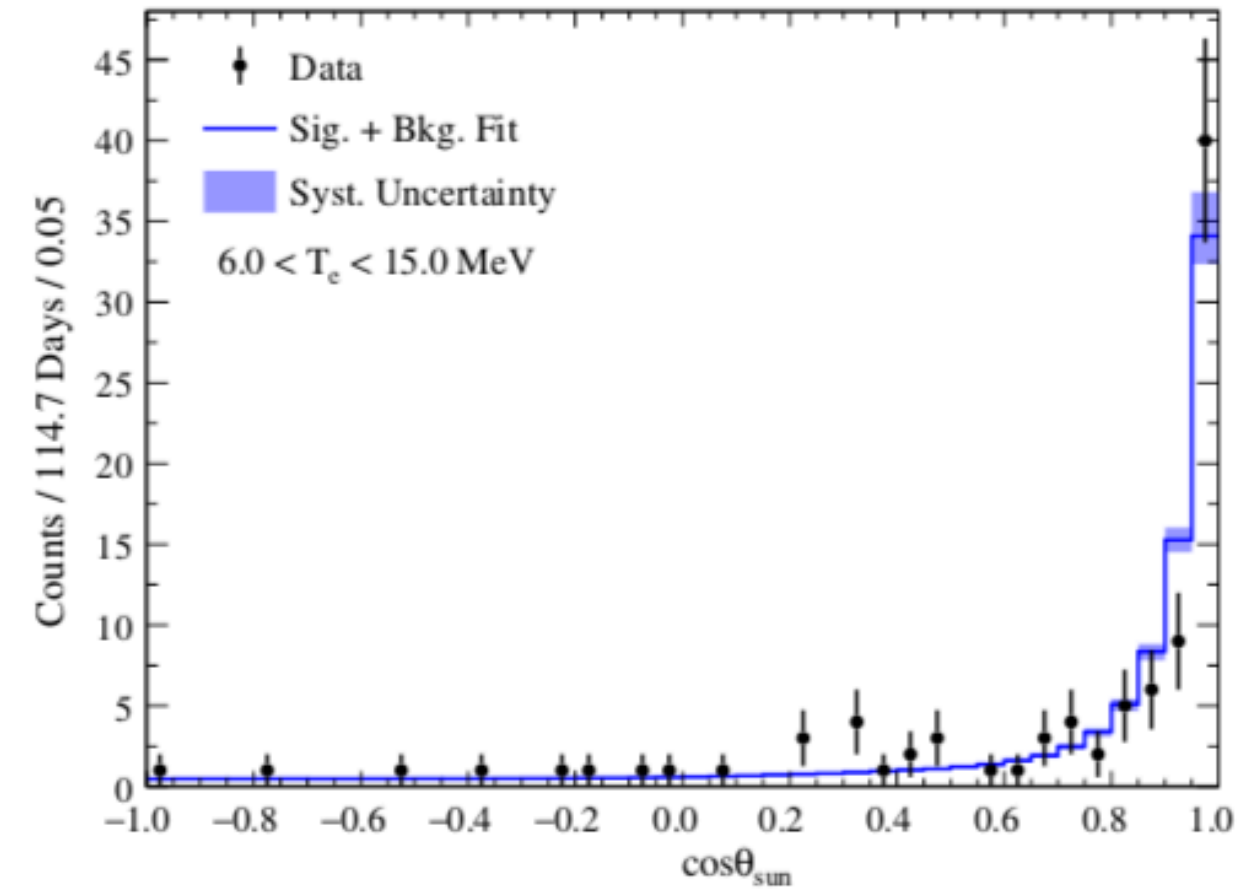


FIG. 4. Distribution of event directions with respect to solar direction for events with energy in the range 6.0–15.0 MeV.

↑  
Largest background to  $0\nu\beta\beta$

PHYSICAL REVIEW D **99**, 032008 (2019)

### Search for invisible modes of nucleon decay in water with the SNO+ detector

TABLE VI. Lifetime limits at 90% C.I. for the spectral and counting analysis, including statistical and systematic uncertainties alongside the existing limits.

	Spectral analysis	Counting analysis	Existing limits
$n$	$2.5 \times 10^{29}$ y	$2.6 \times 10^{29}$ y	$5.8 \times 10^{29}$ y [9]
$p$	$3.6 \times 10^{29}$ y	$3.4 \times 10^{29}$ y	$2.1 \times 10^{29}$ y [10]
$pp$	$4.7 \times 10^{28}$ y	$4.1 \times 10^{28}$ y	$5.0 \times 10^{25}$ y [11]
$pn$	$2.6 \times 10^{28}$ y	$2.3 \times 10^{28}$ y	$2.1 \times 10^{25}$ y [13]
$nn$	$1.3 \times 10^{28}$ y	$0.6 \times 10^{28}$ y	$1.4 \times 10^{30}$ y [9]

See Martti Nirkko's slides!

Plus other analyses underway

# Detector progress

- Operating with **water** from 2017
- Transition to **scintillator** happening now

- Scintillator purification plant commissioned
- **LAB** successfully distilled underground
- **PPO** prep underway
- N<sub>2</sub>/steam stripping tested



- Invisible nucleon decay
- Solar neutrinos
- Supernova neutrinos
- **Reactor neutrinos ( $\Delta m^2_{12}$ )**
- **Geo-neutrinos**



# Detector progress

- Operating with **water** from 2017
- Transition to **scintillator** happening now
- **Tellurium** loading for  $\beta\beta$  due in 2019-20 (1330 kg  $^{130}\text{Te}$ )

- Invisible nucleon decay
- Solar neutrinos
- Supernova neutrinos
- Reactor neutrinos ( $\Delta m^2_{12}$ )
- Geo-neutrinos
- **Neutrinoless double-beta decay**

Te needed for Phase I **all underground**



Te **purification system** almost complete

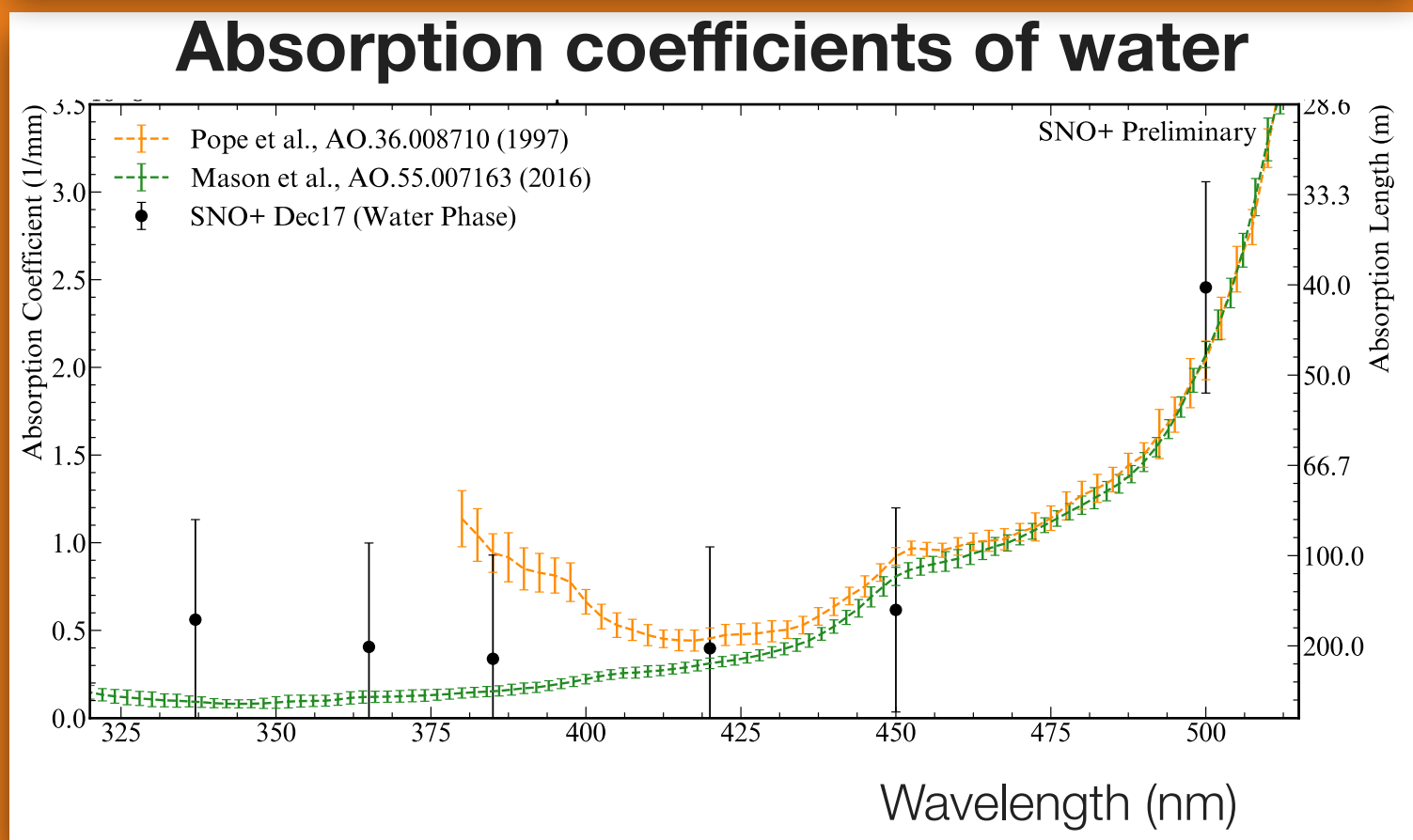
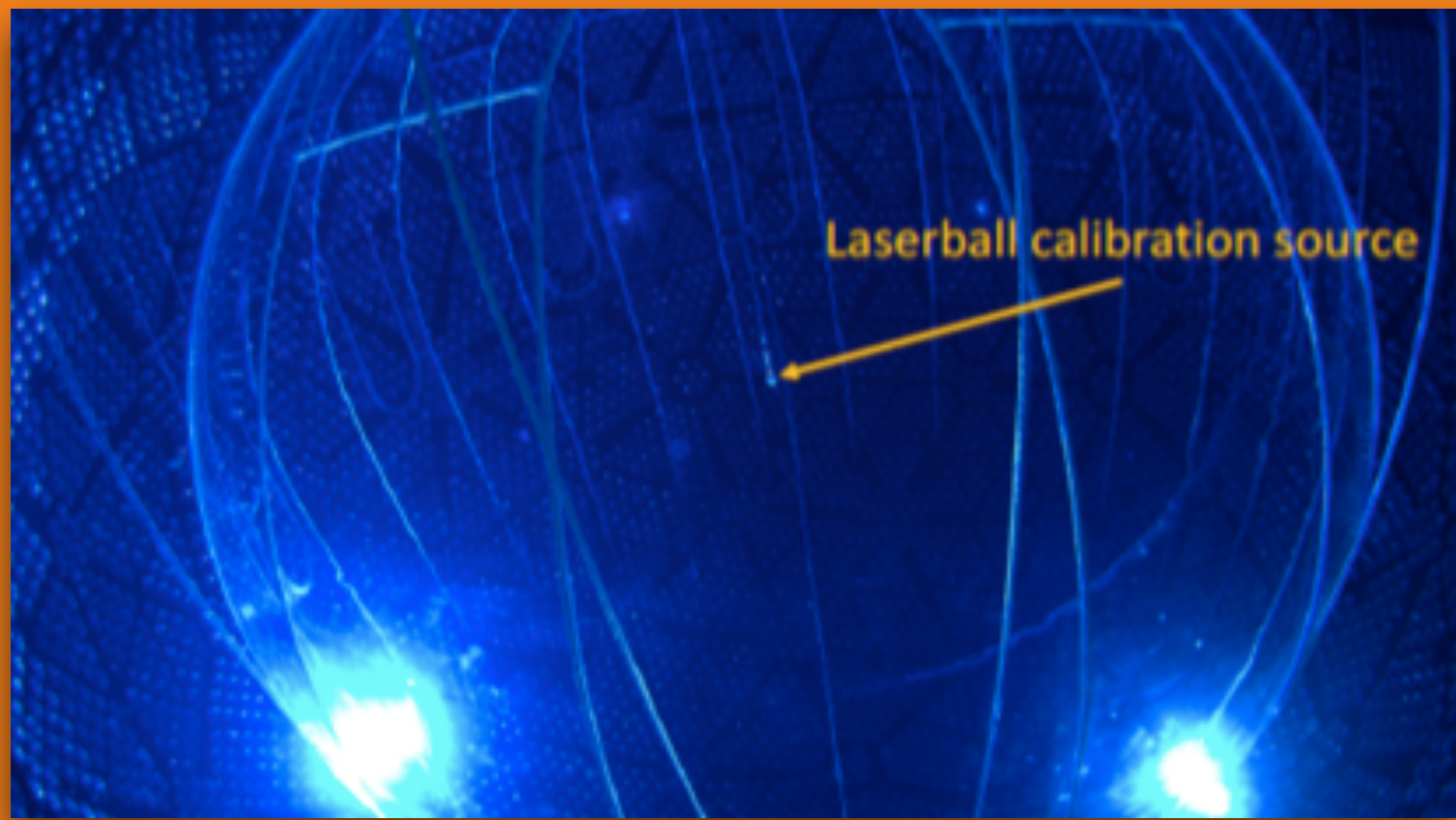


**Te-diol synthesis plant** construction is well advanced (synthesised from telluric acid)





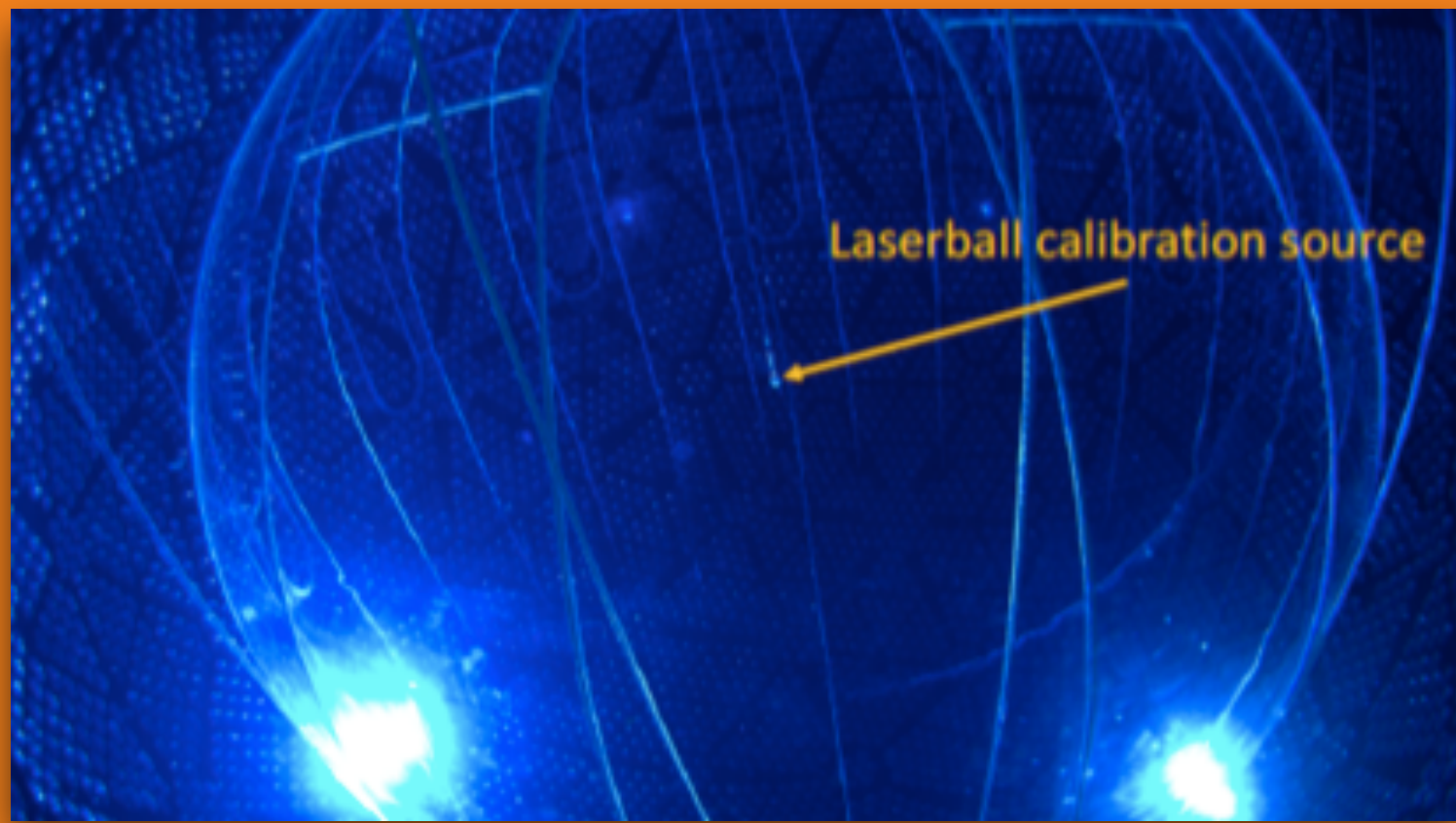
Water phase: measure **absorption** coefficient with light-diffusing “**laserball**” and underwater camera



# Detector calibrations

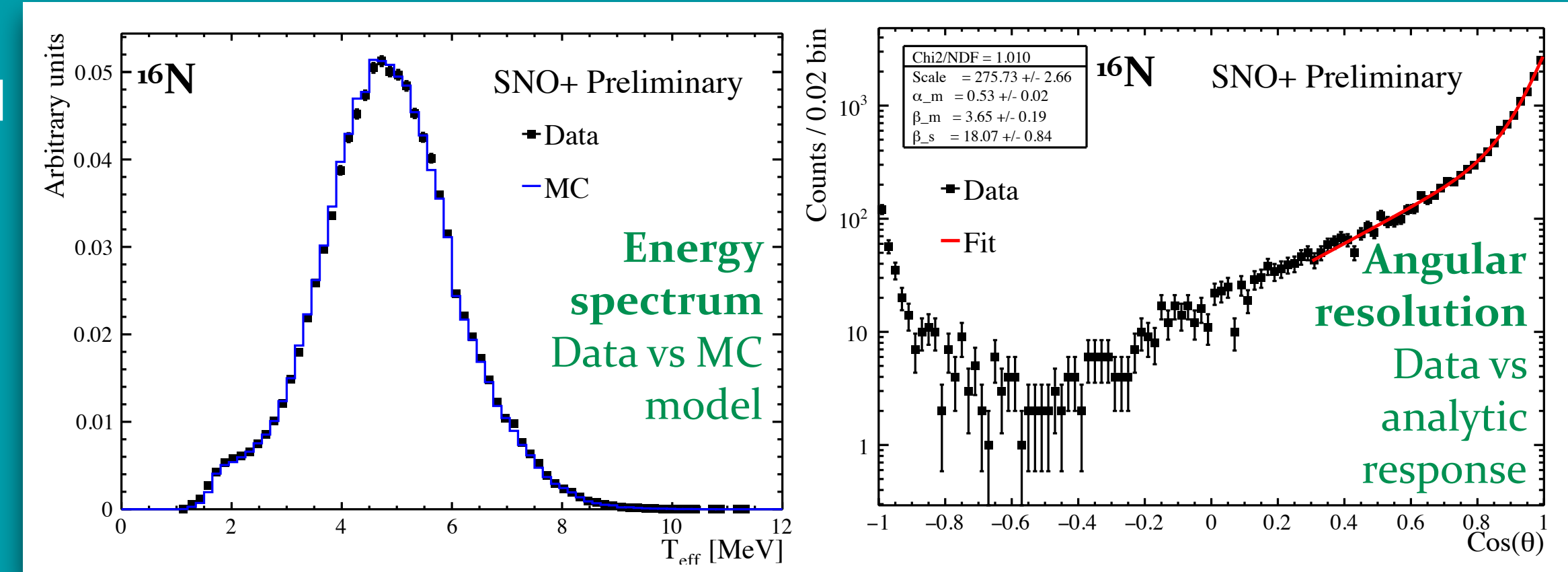


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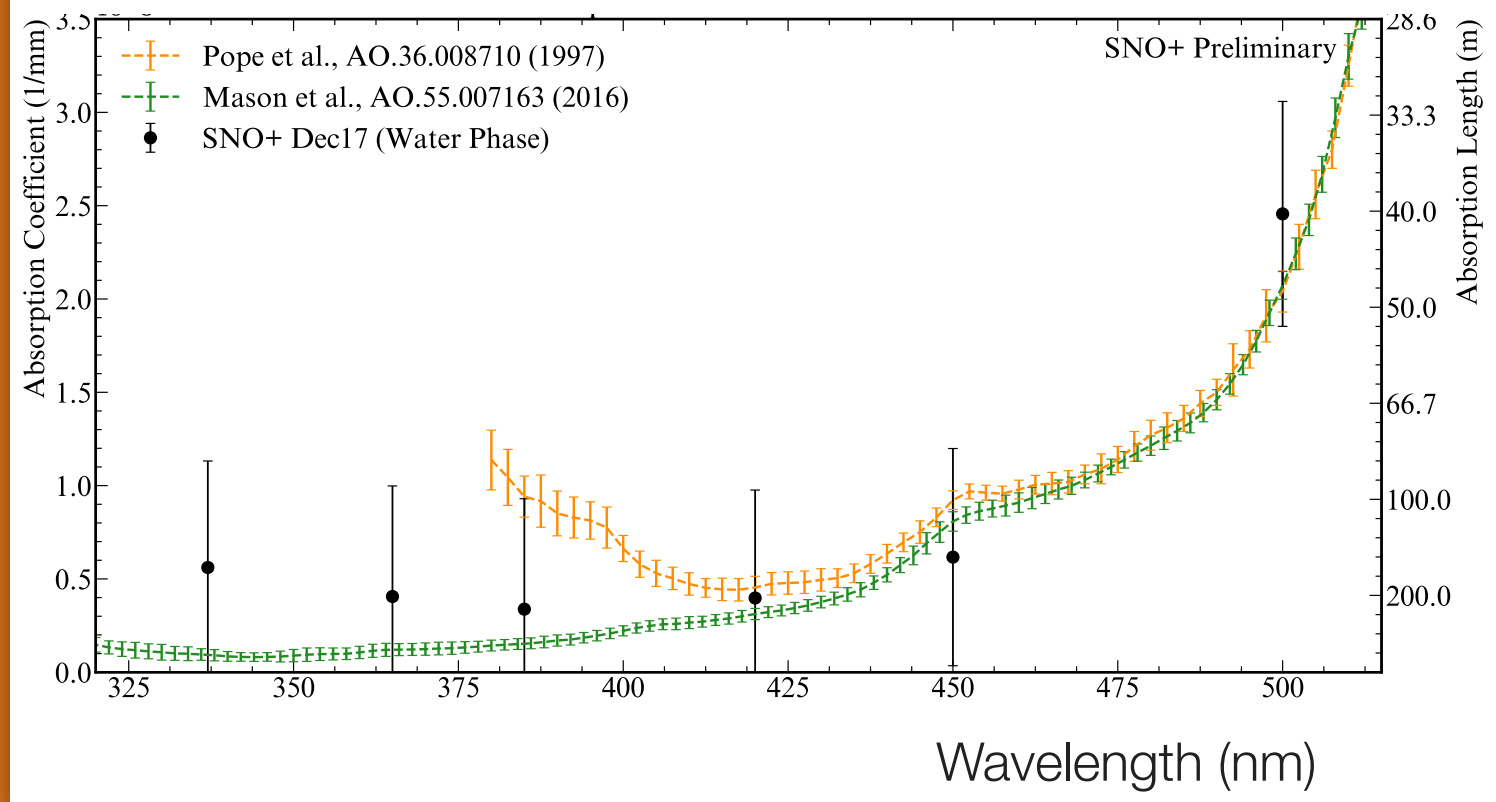


**Detector response** assessed vs models with  $^{16}\text{N}$  source along 3 axes:

- Energy scale, resolution
- Vertex shift, scale, resolution
- Angular resolution



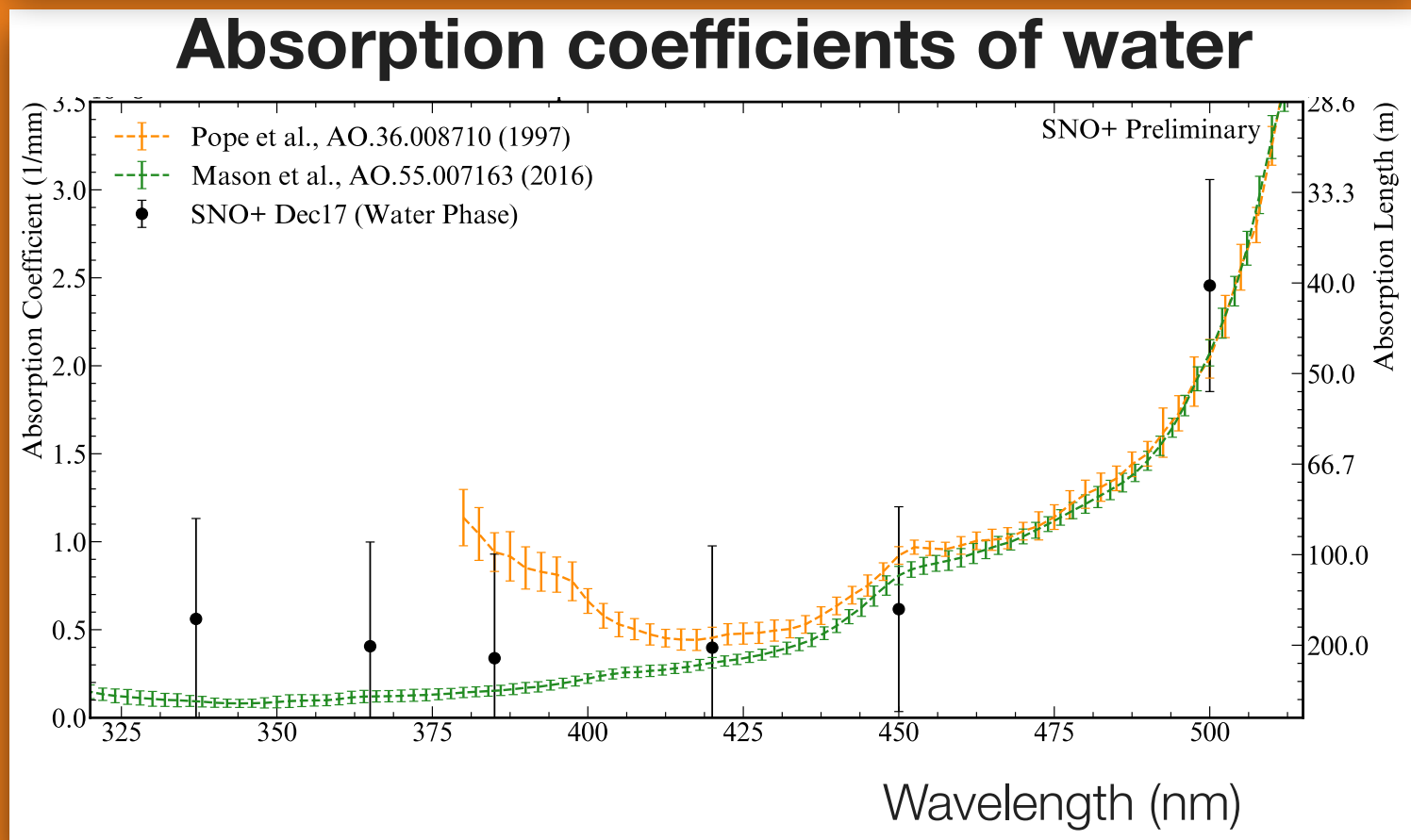
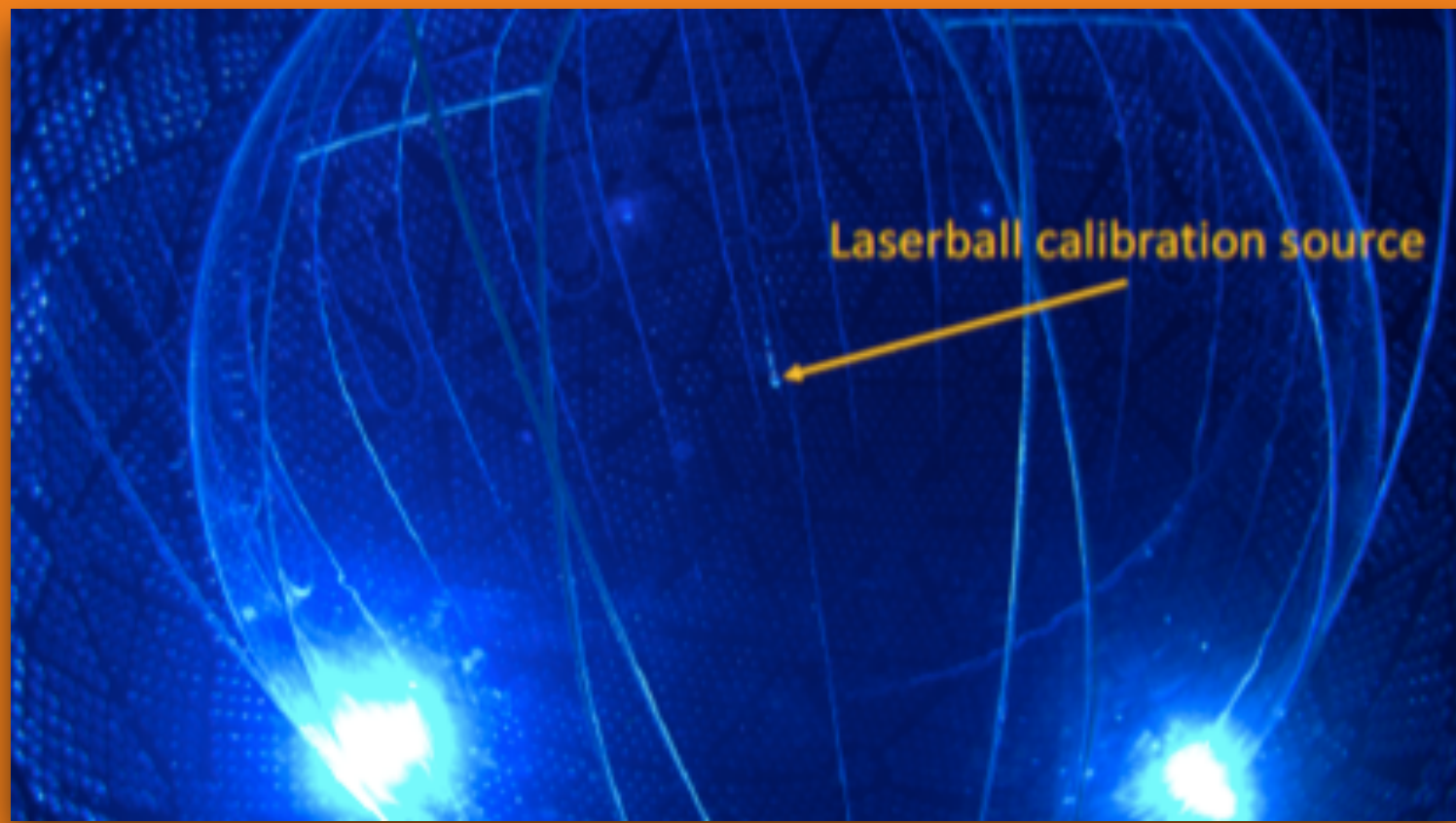
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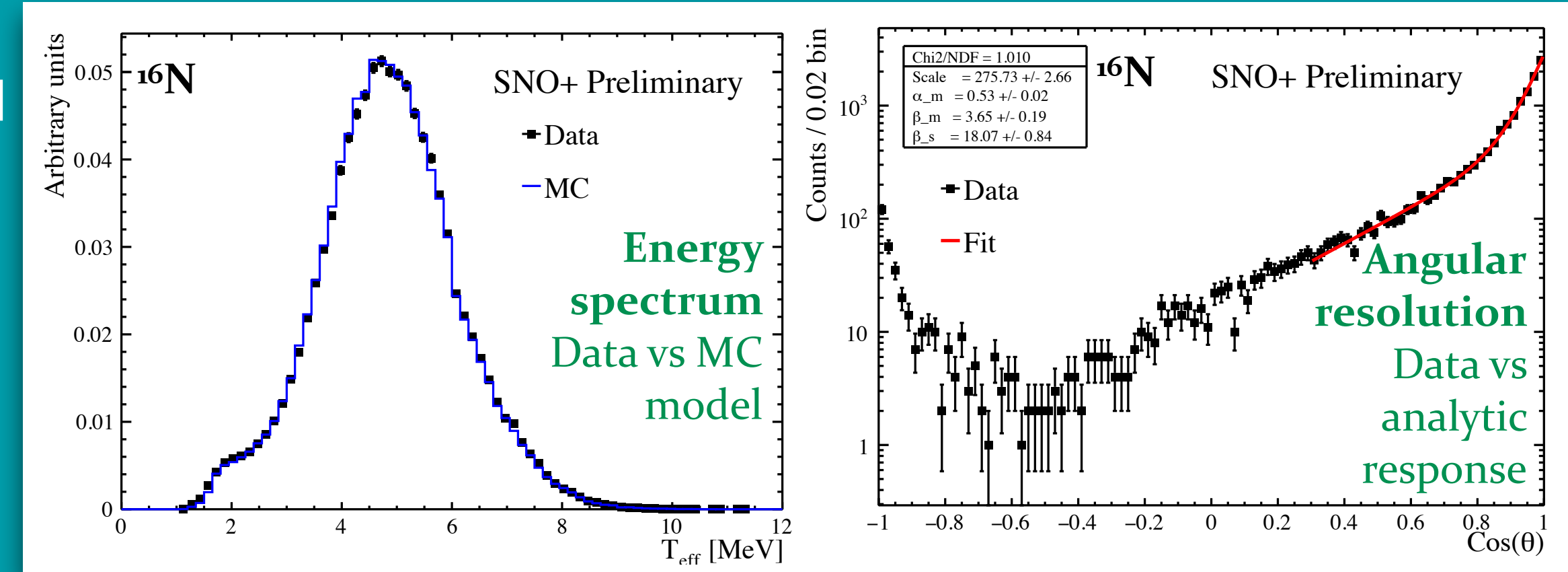


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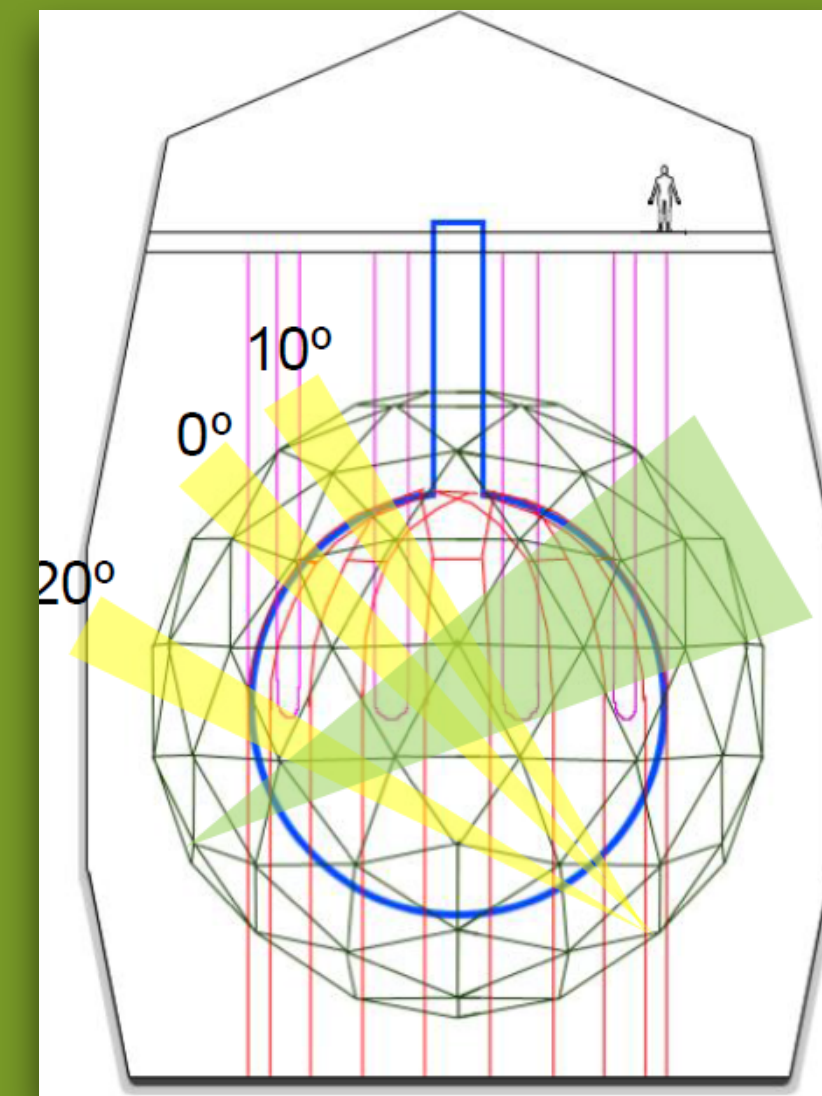
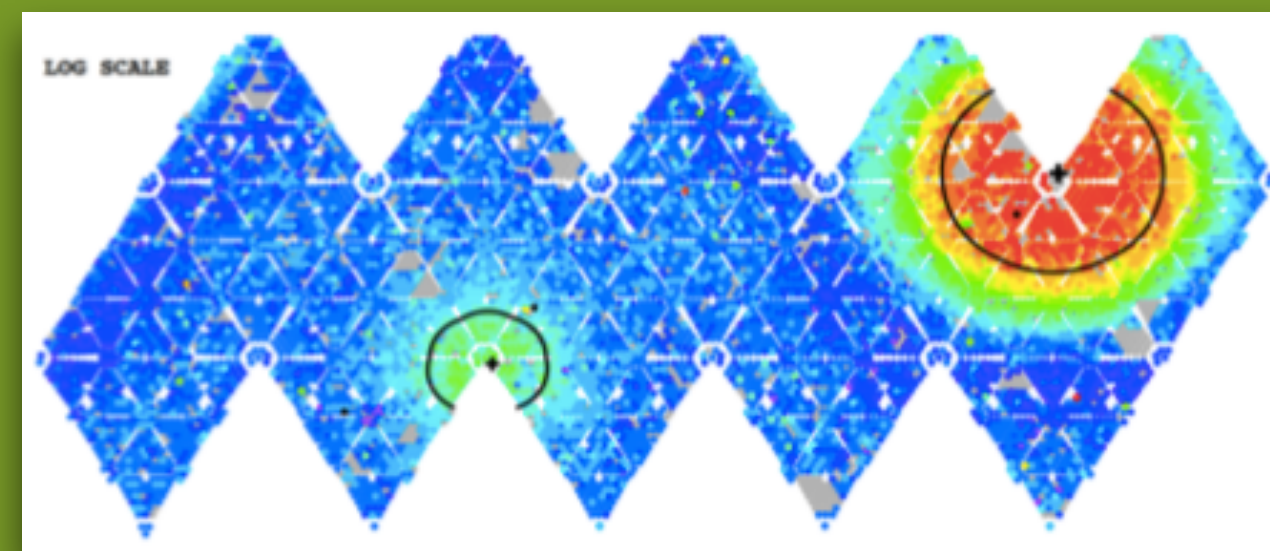
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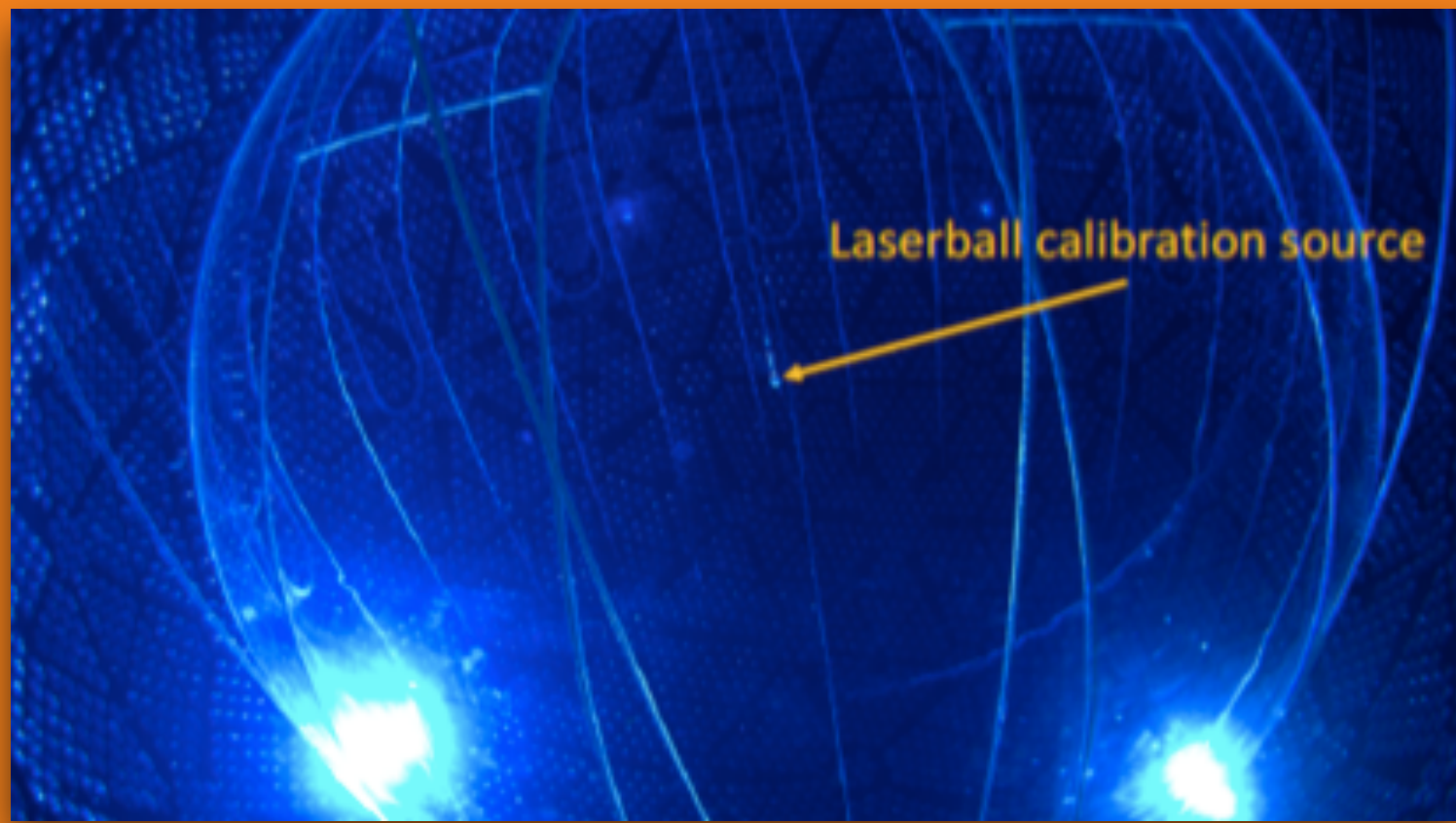
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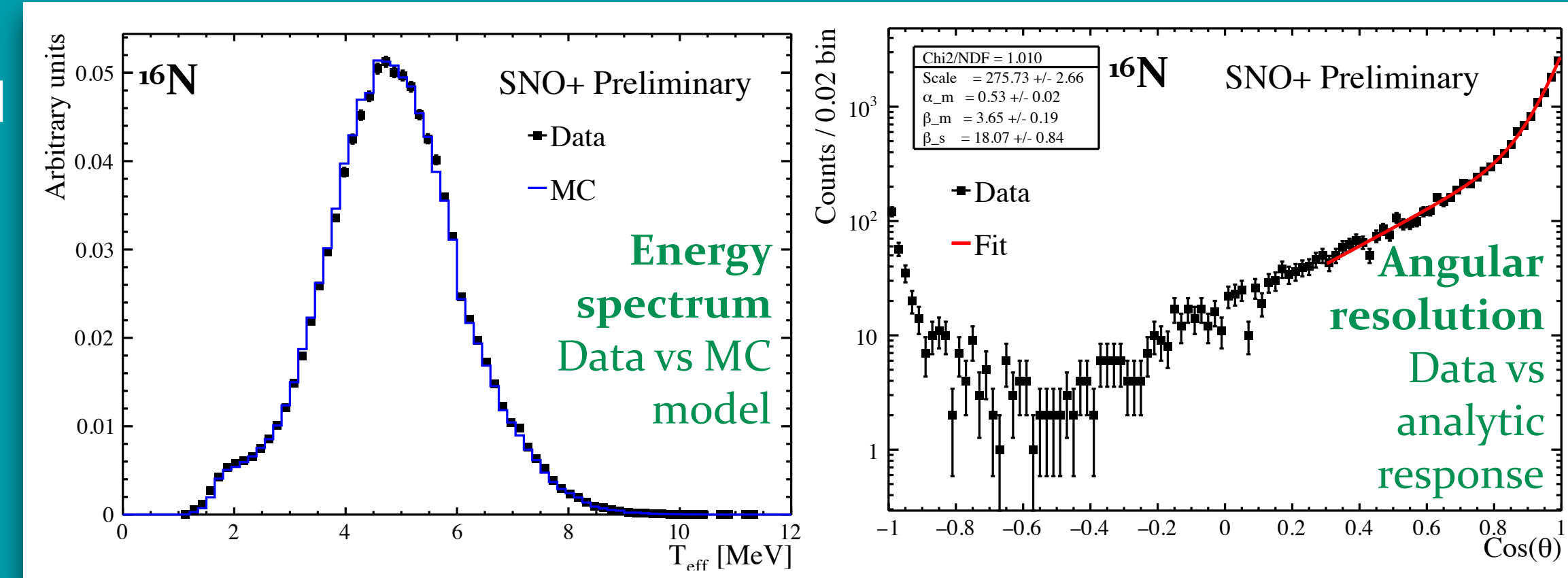


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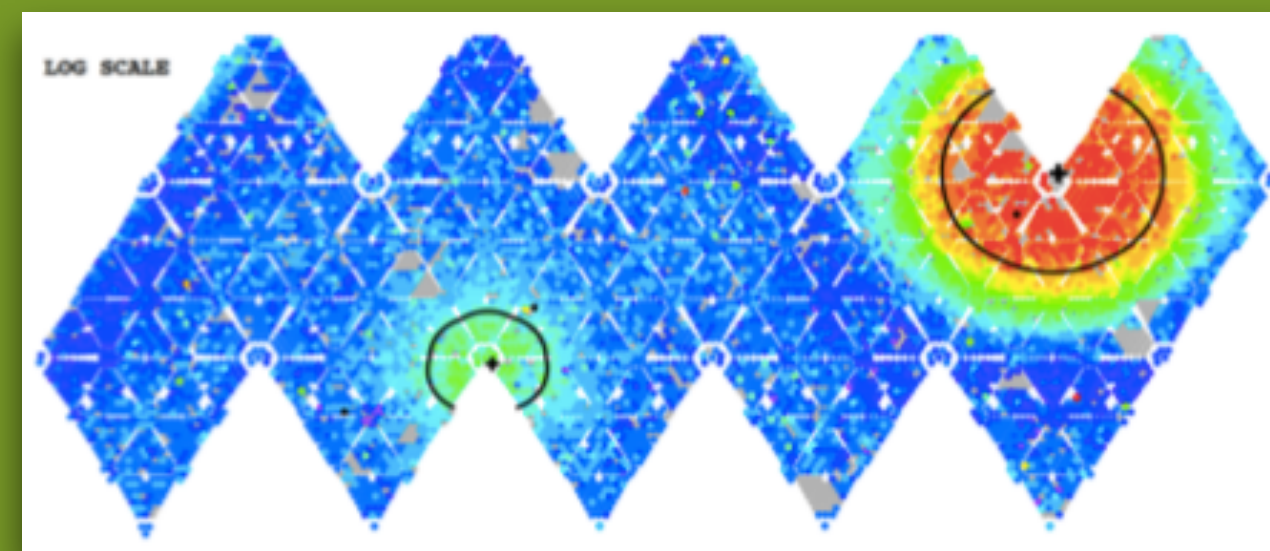
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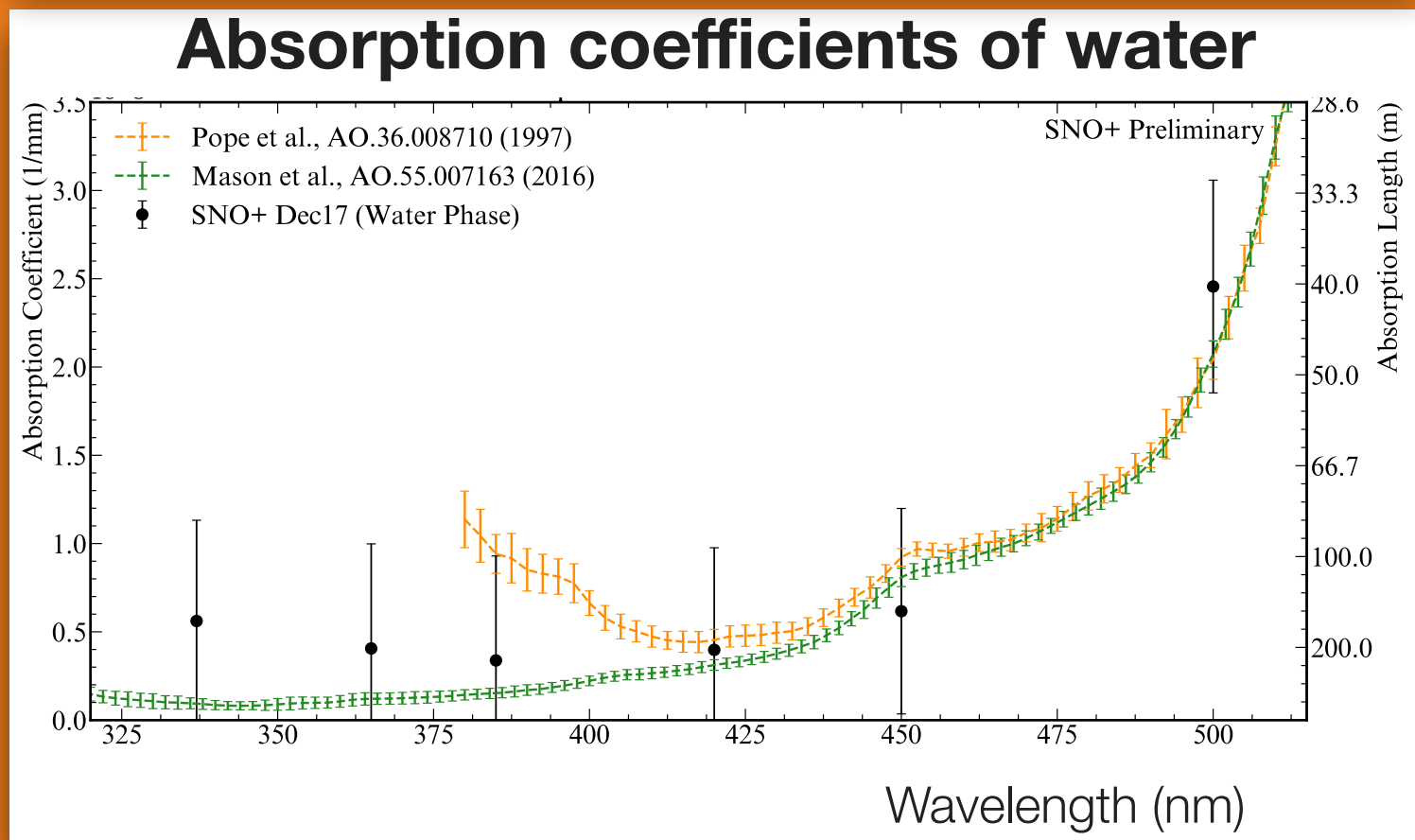
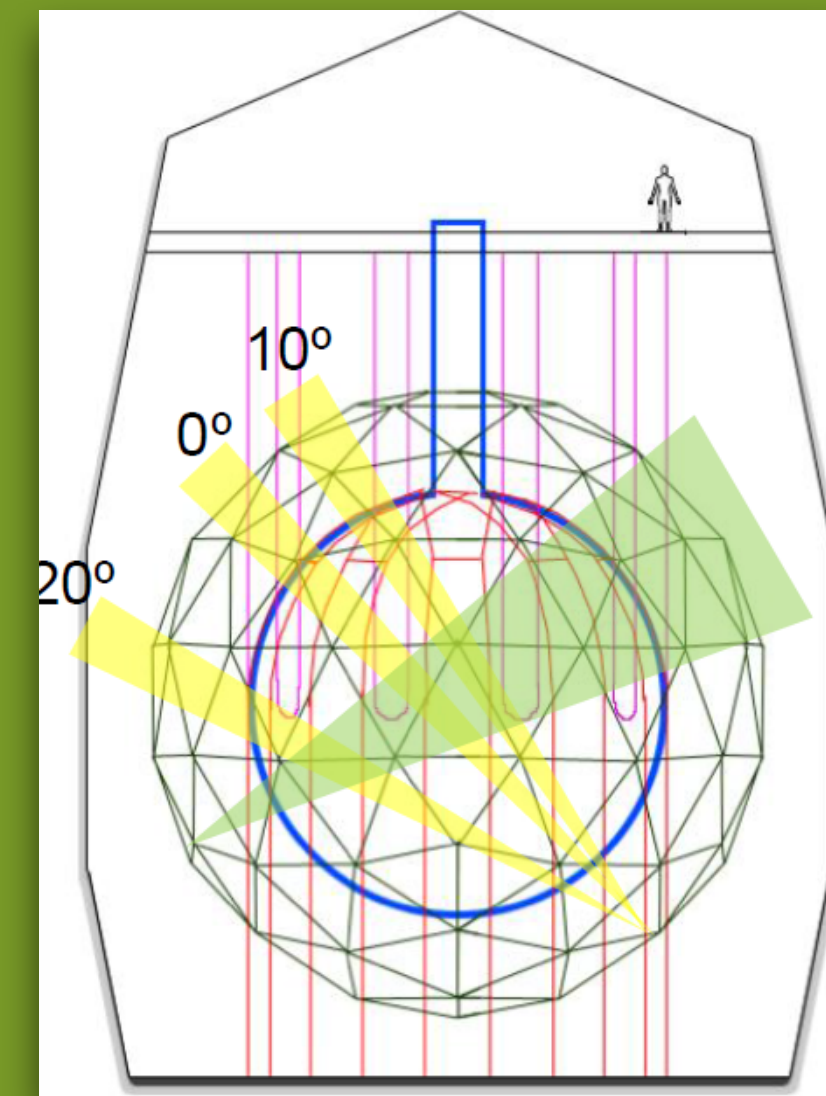


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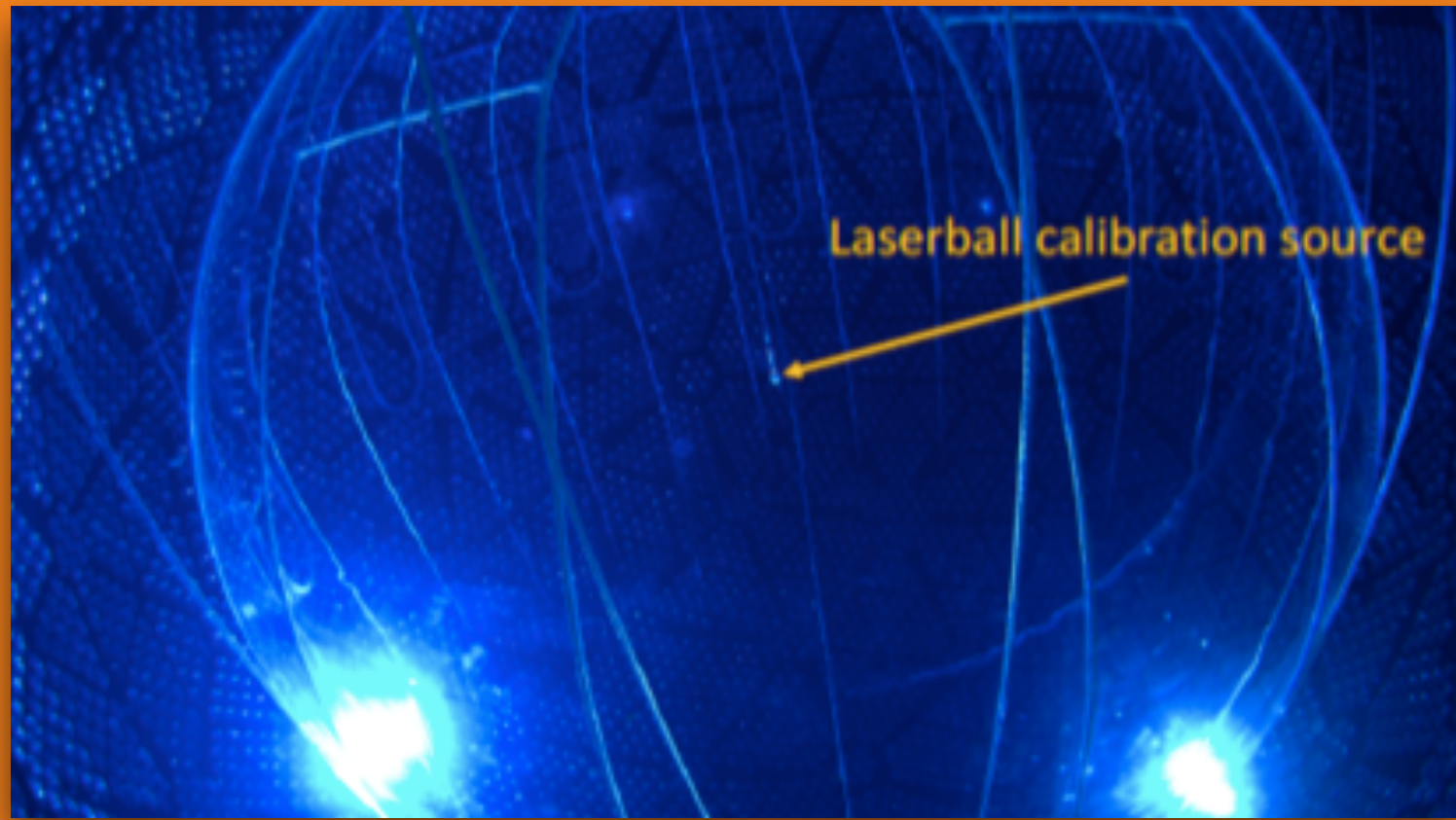


Timing and Monitoring:



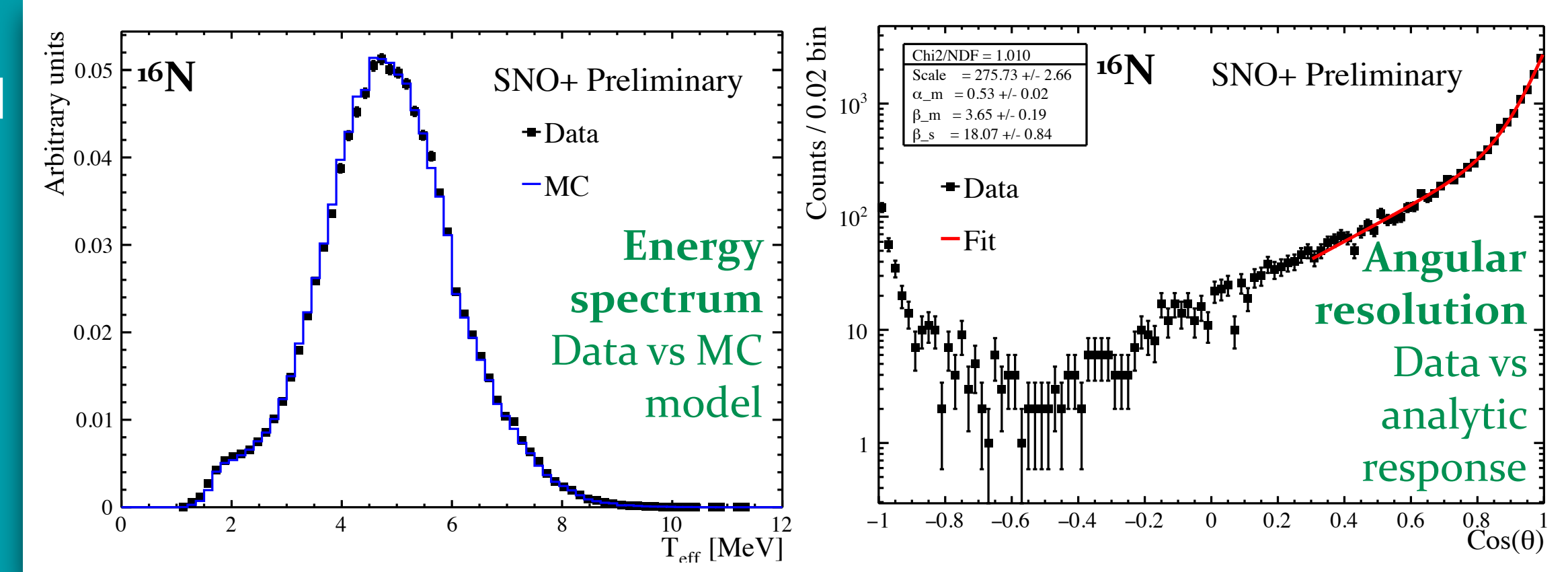


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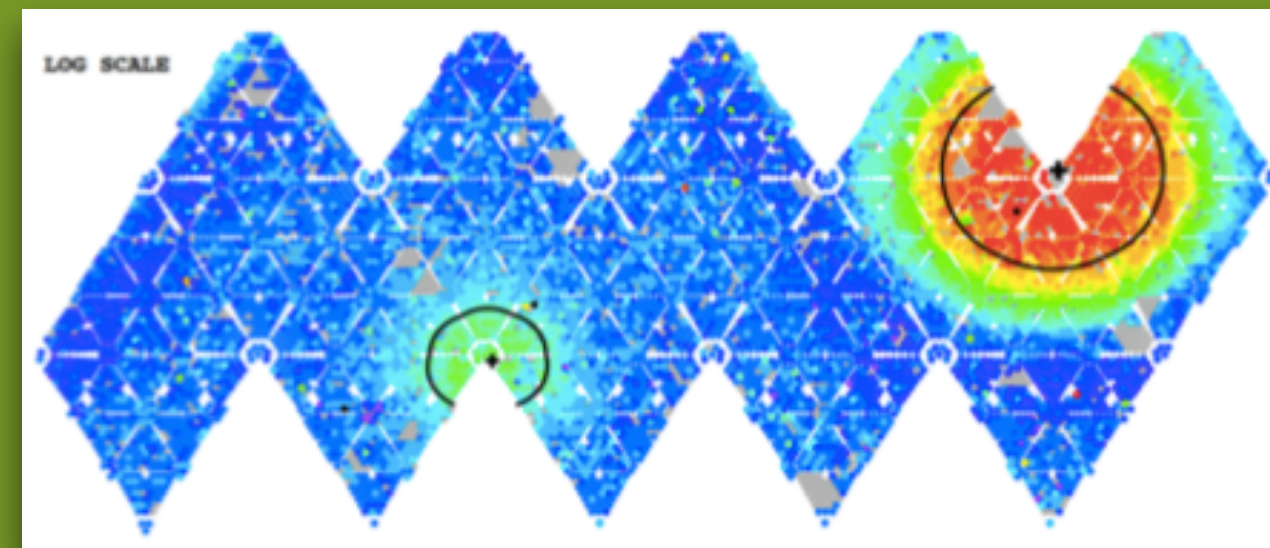
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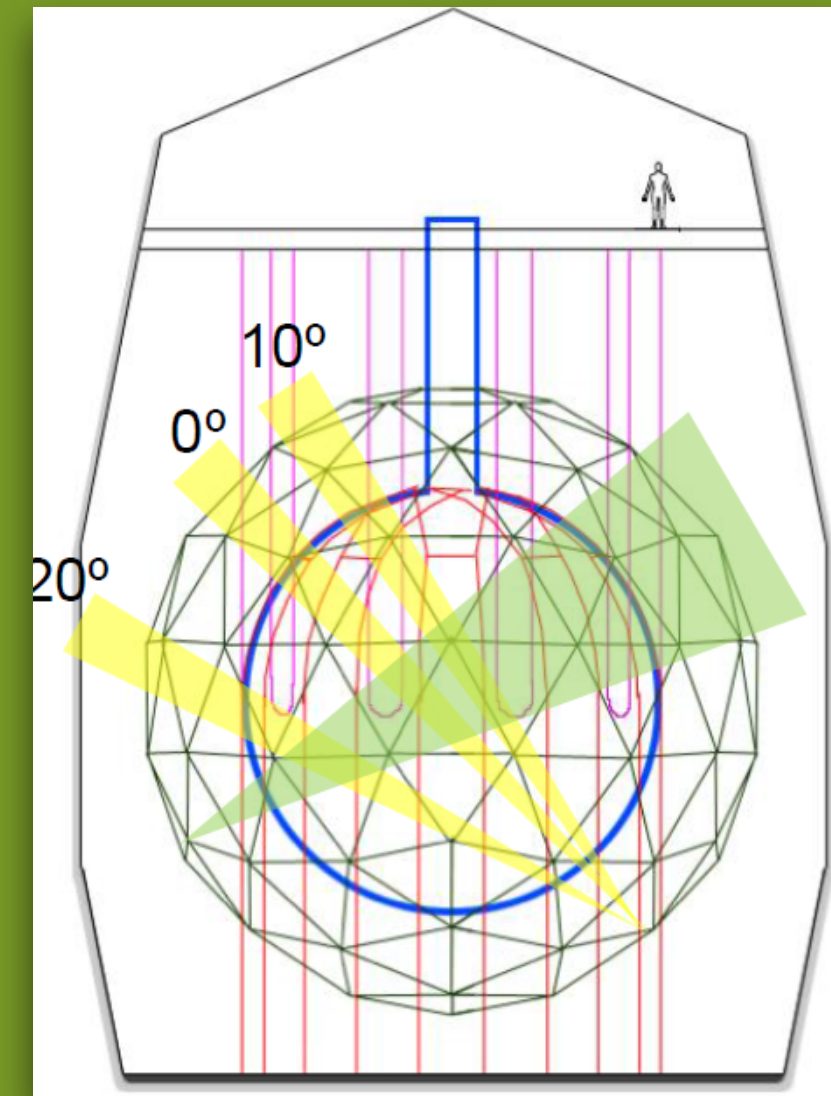


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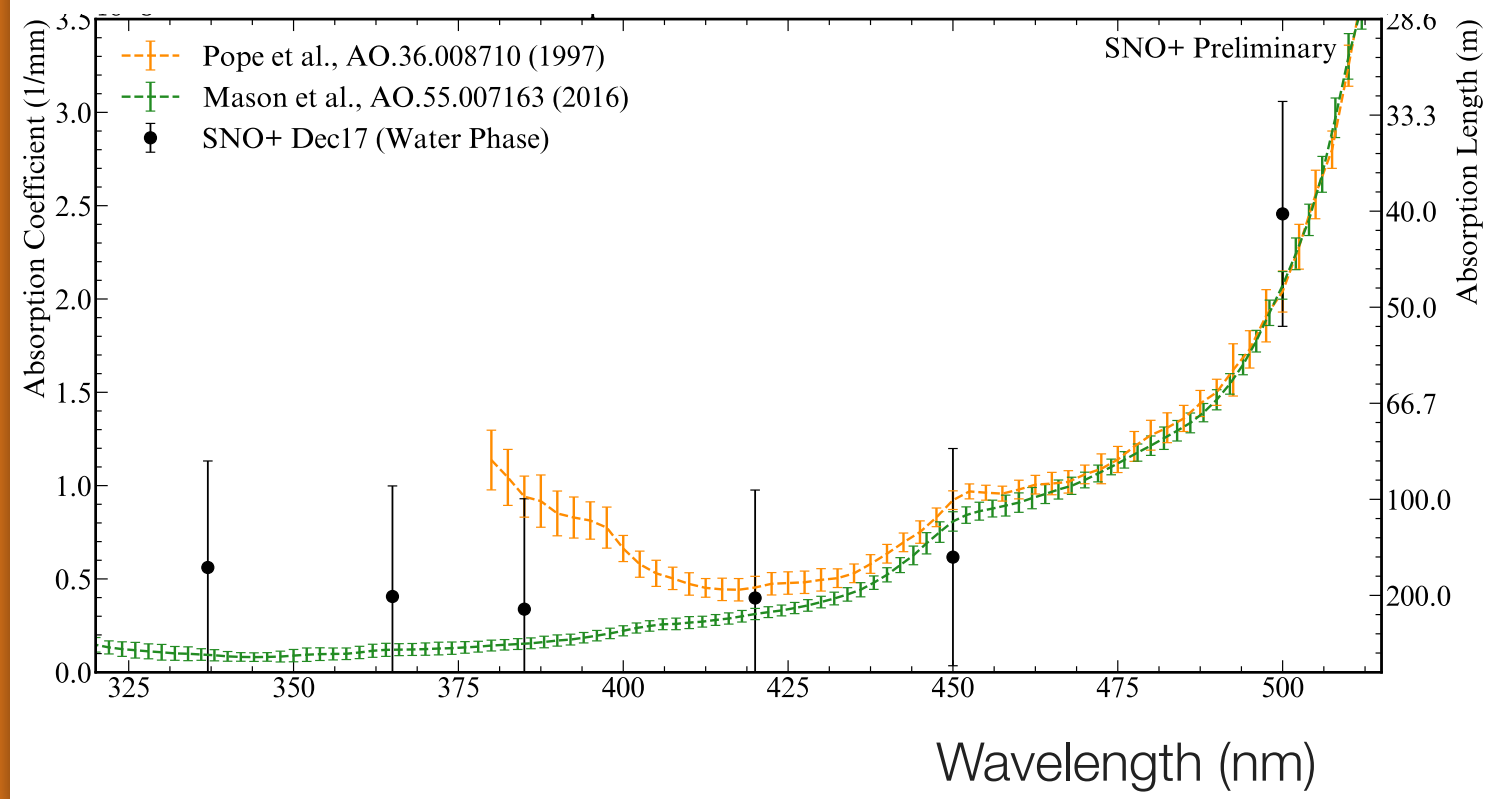
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Timing and Monitoring: **TELLIE**

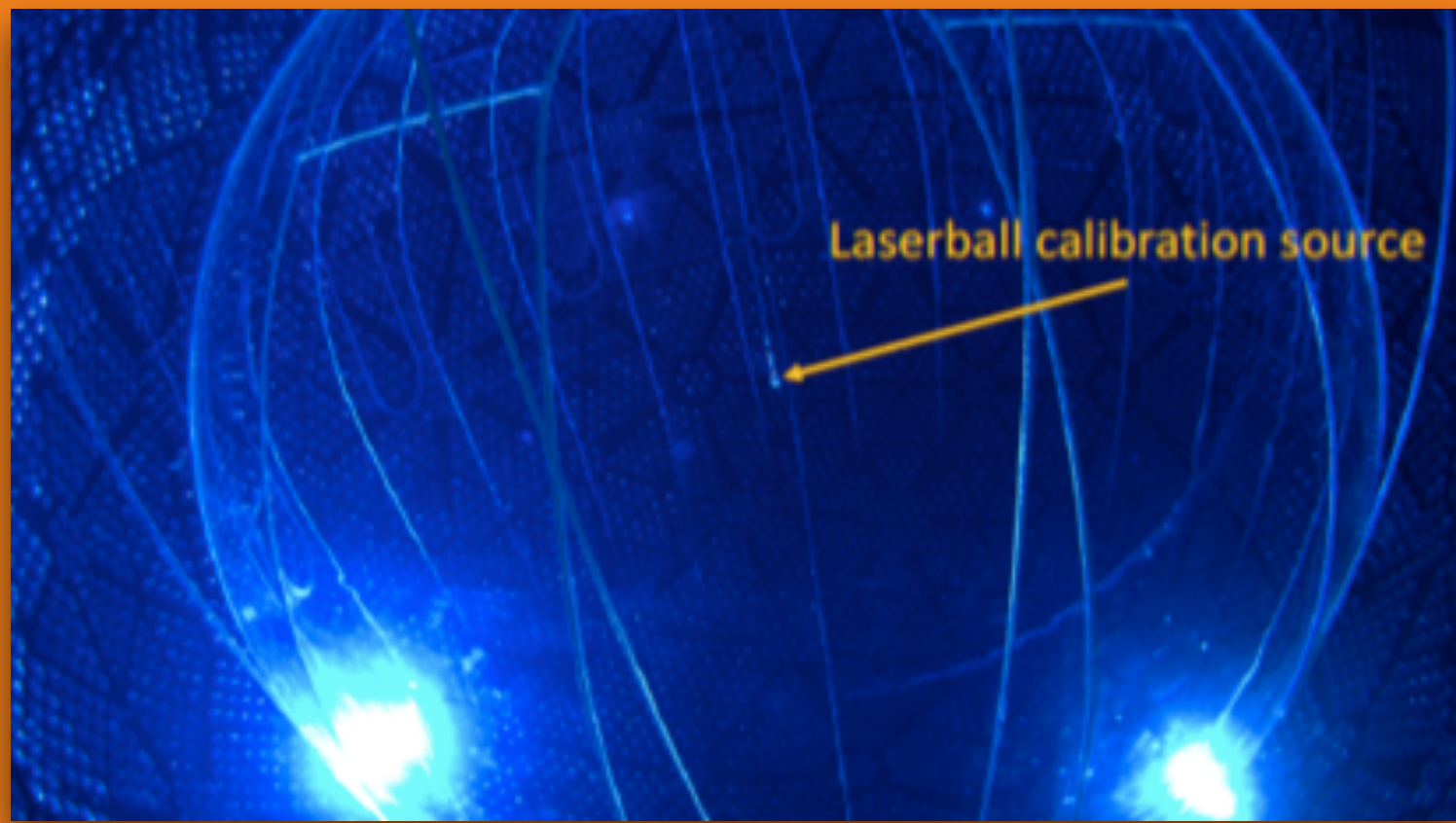


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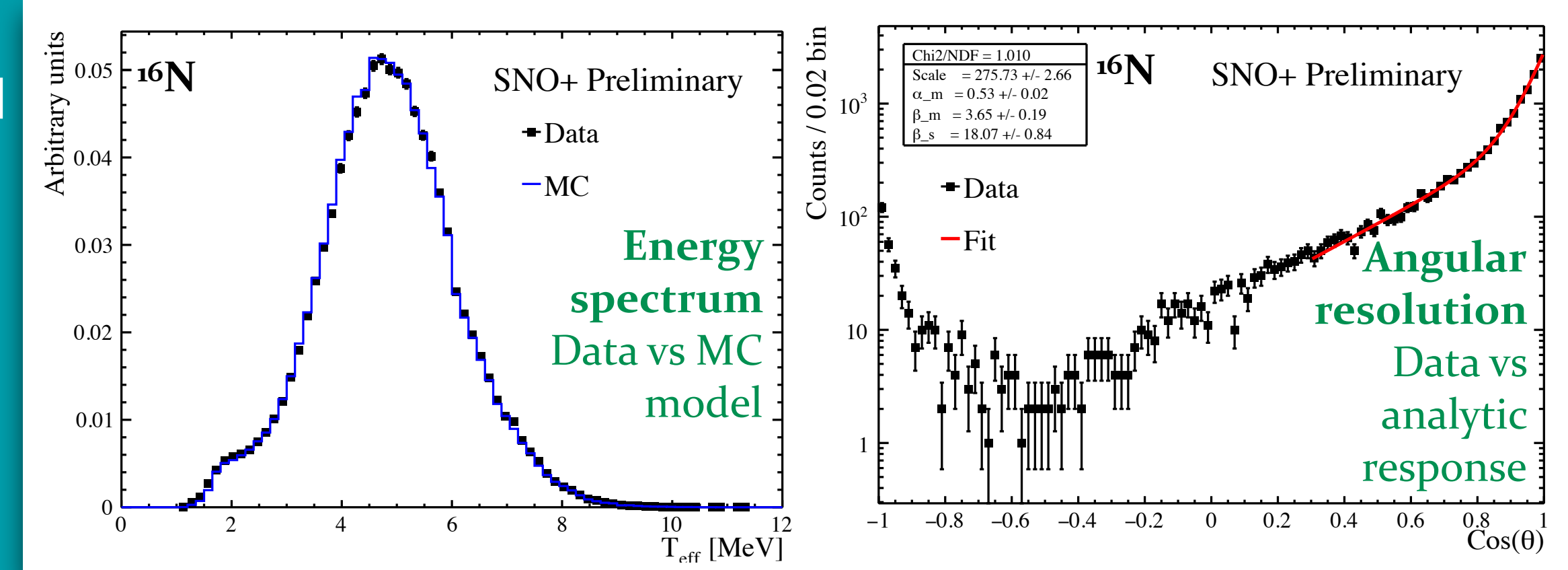


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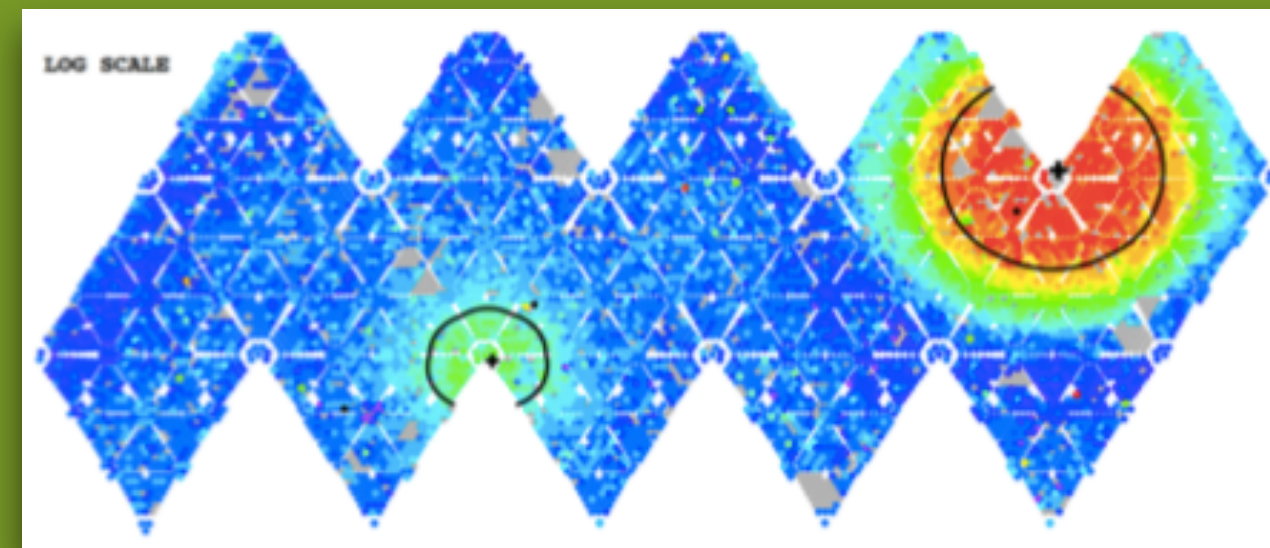
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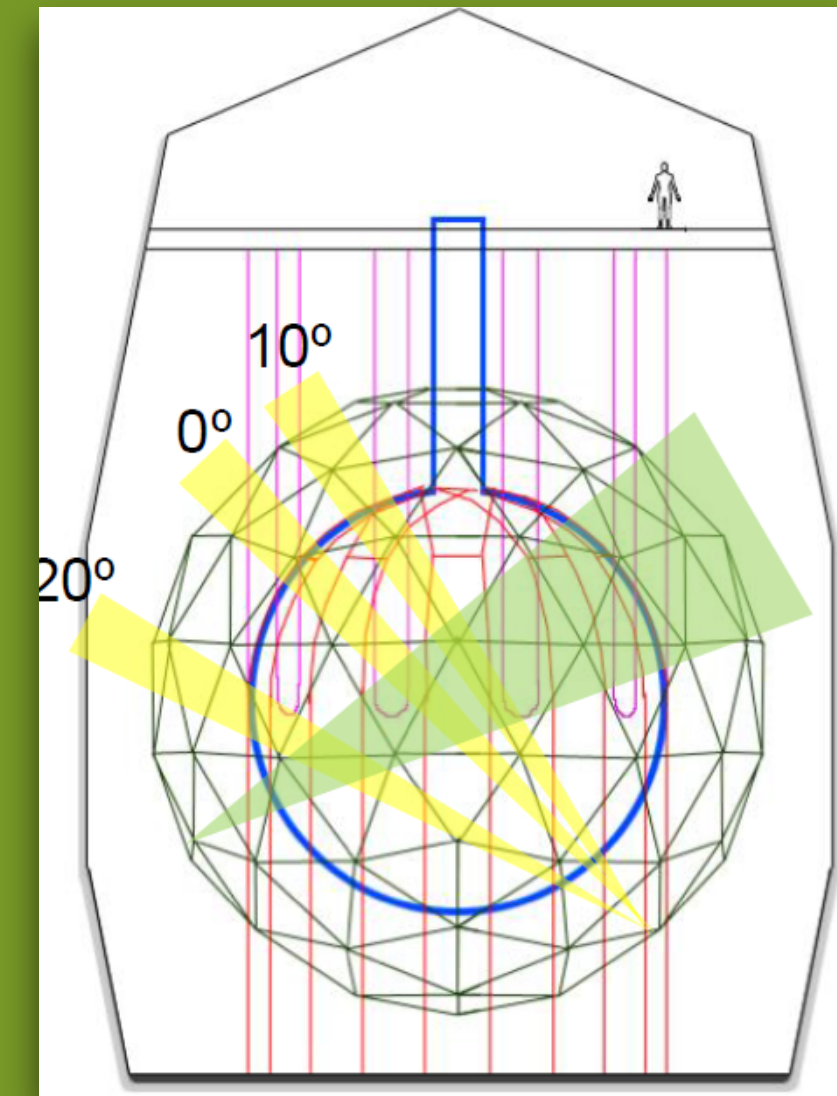


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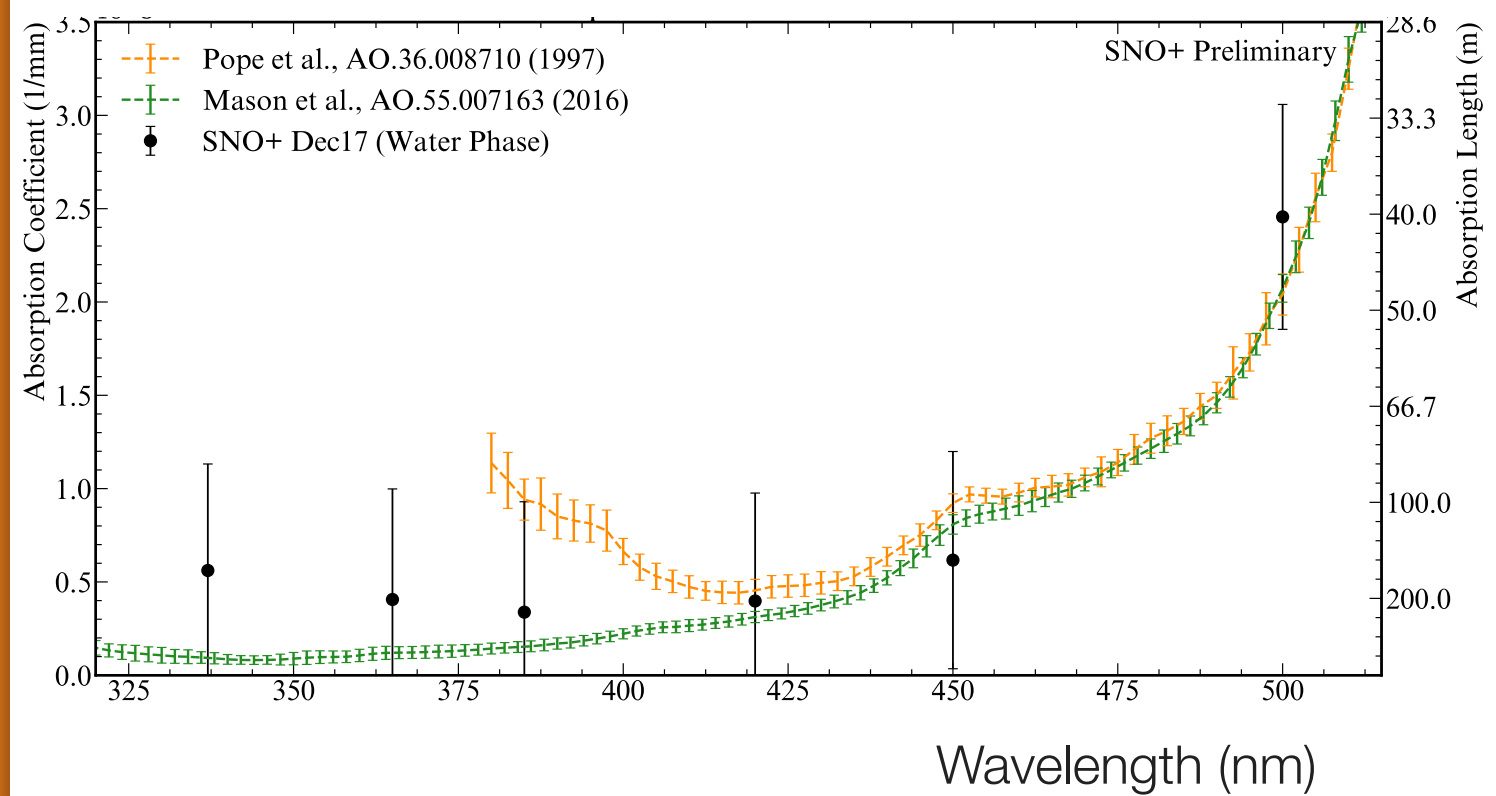
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Timing and Monitoring: **TELLIE**  
Attenuation Module:

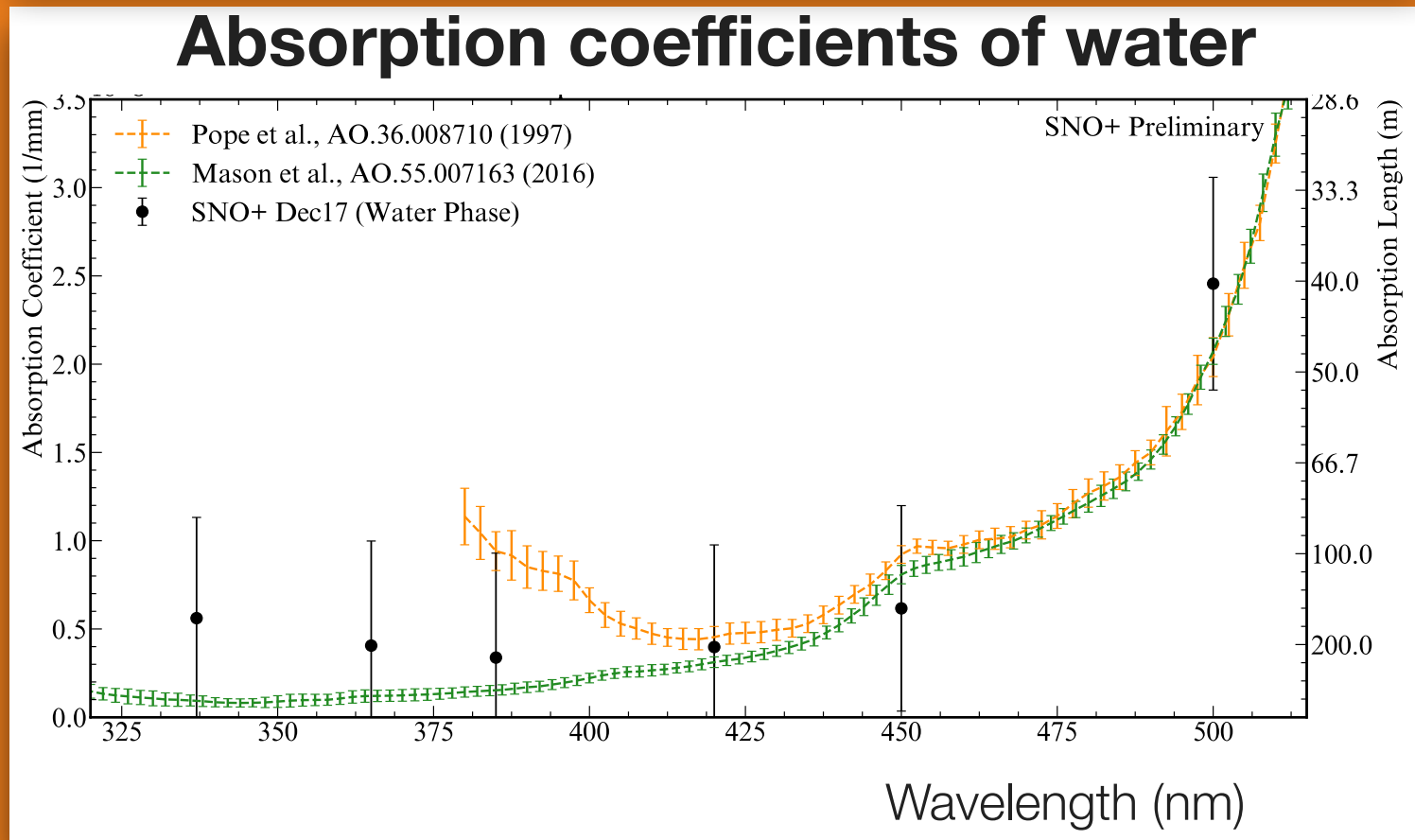
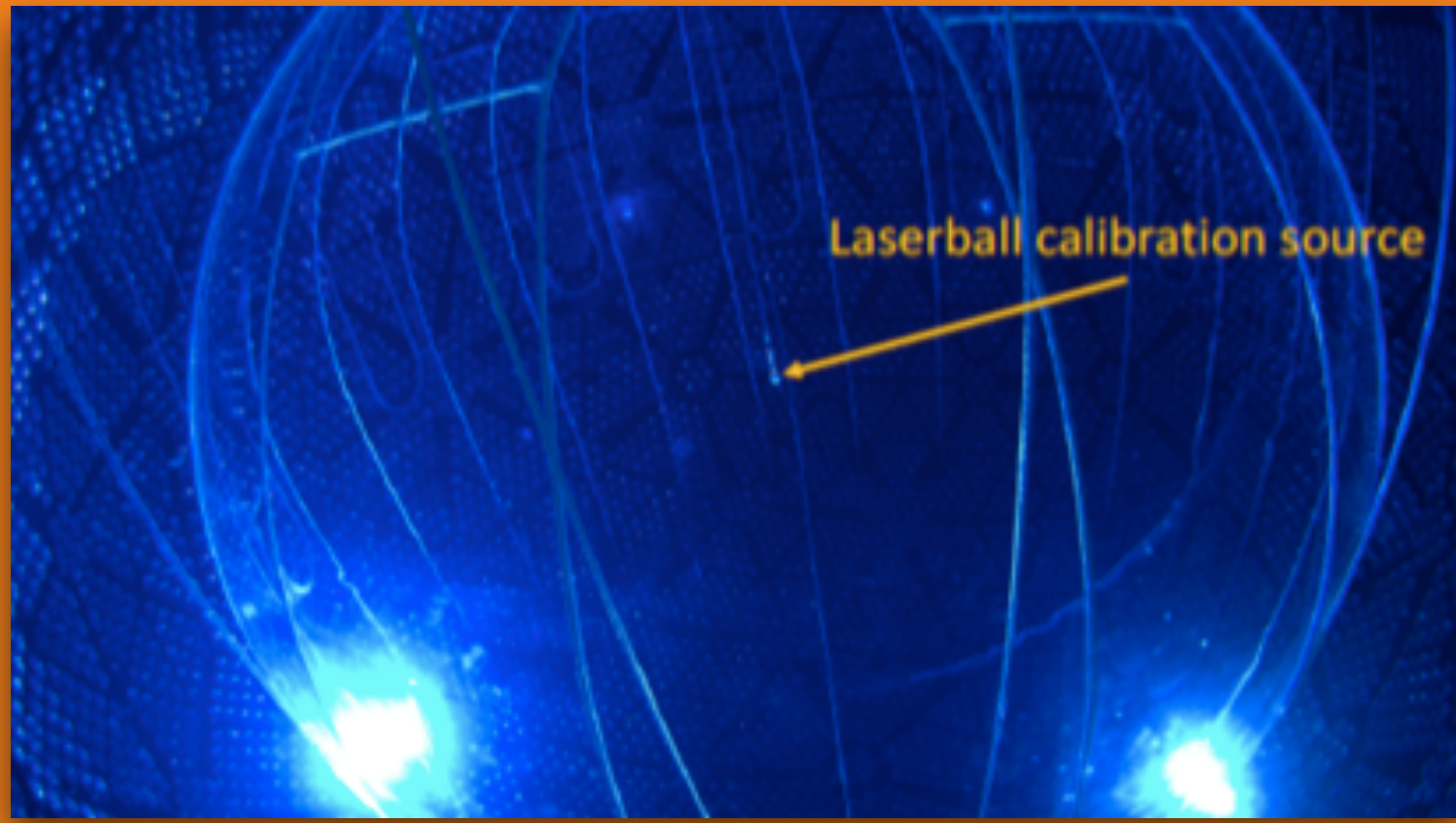


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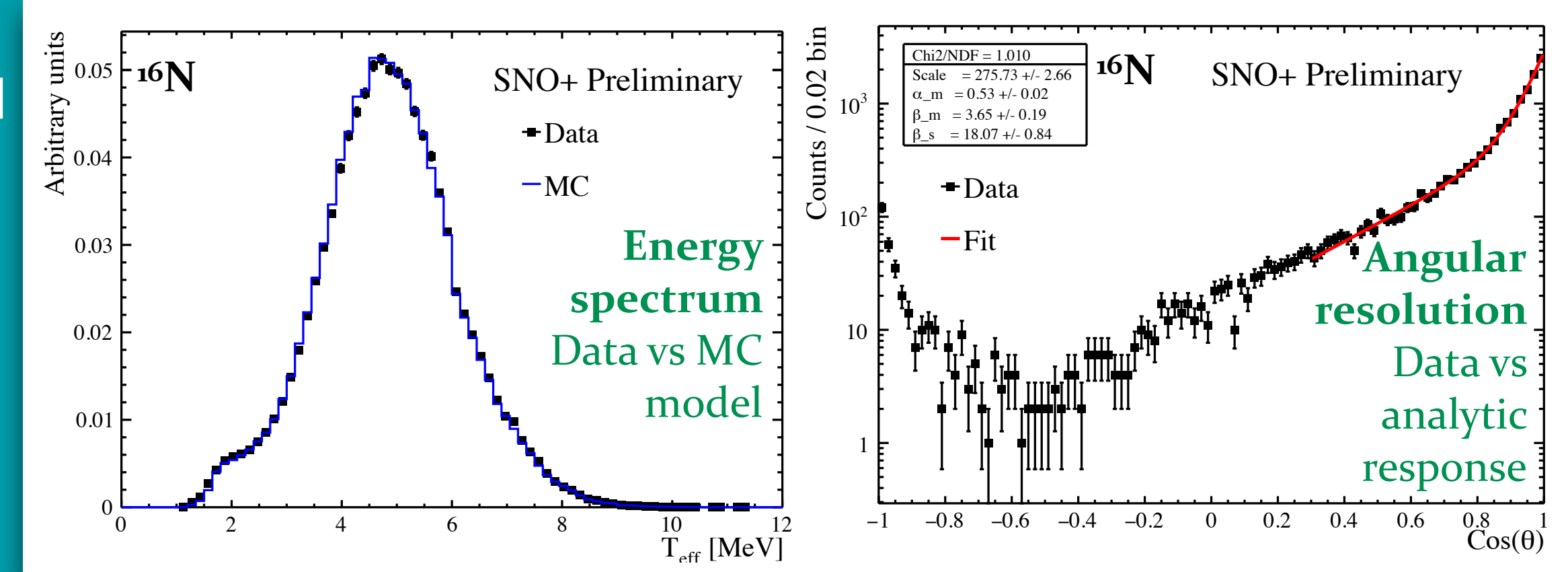


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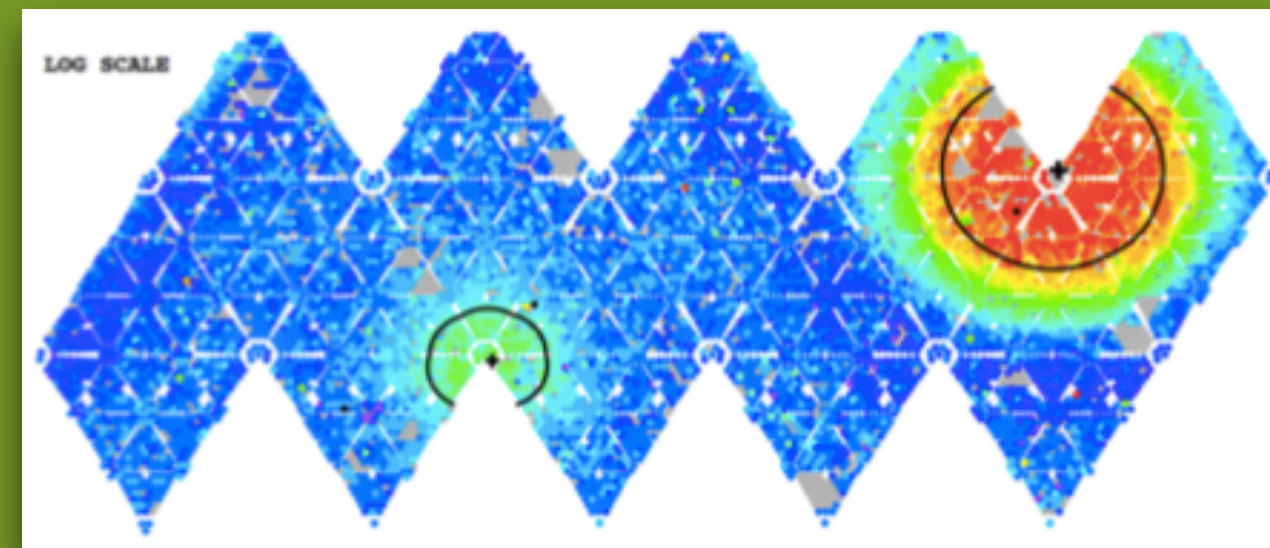
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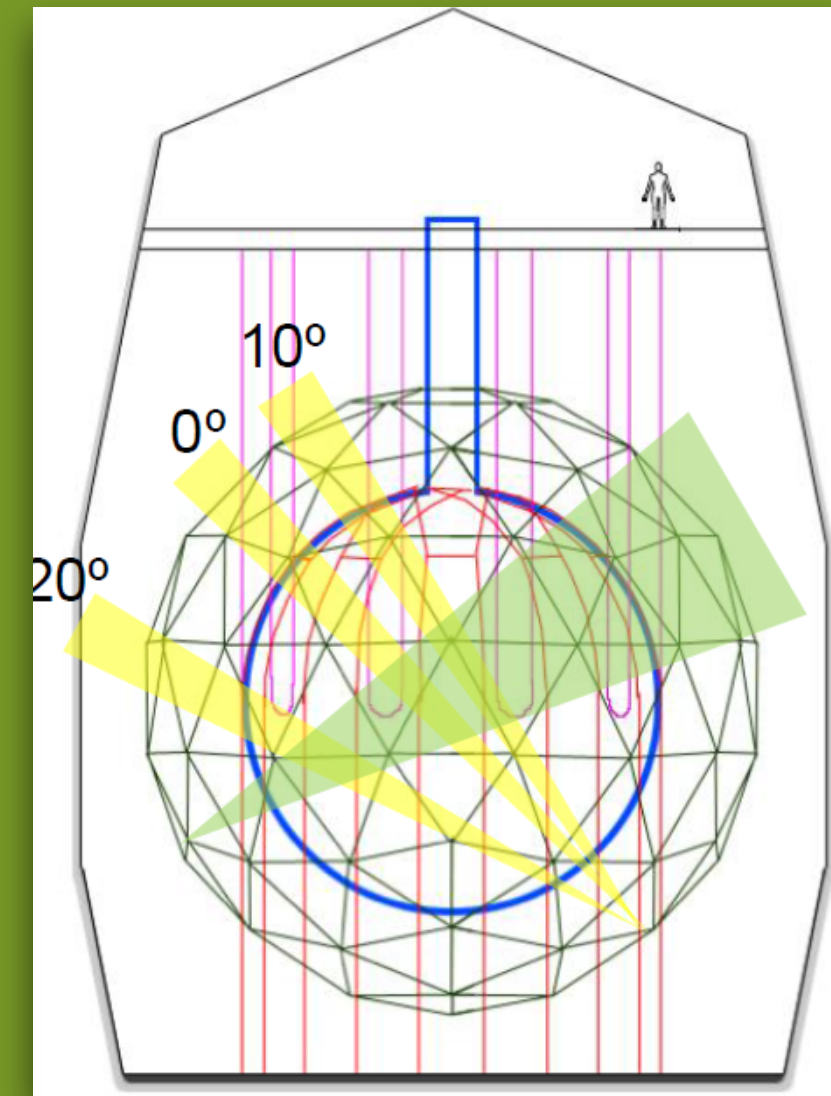


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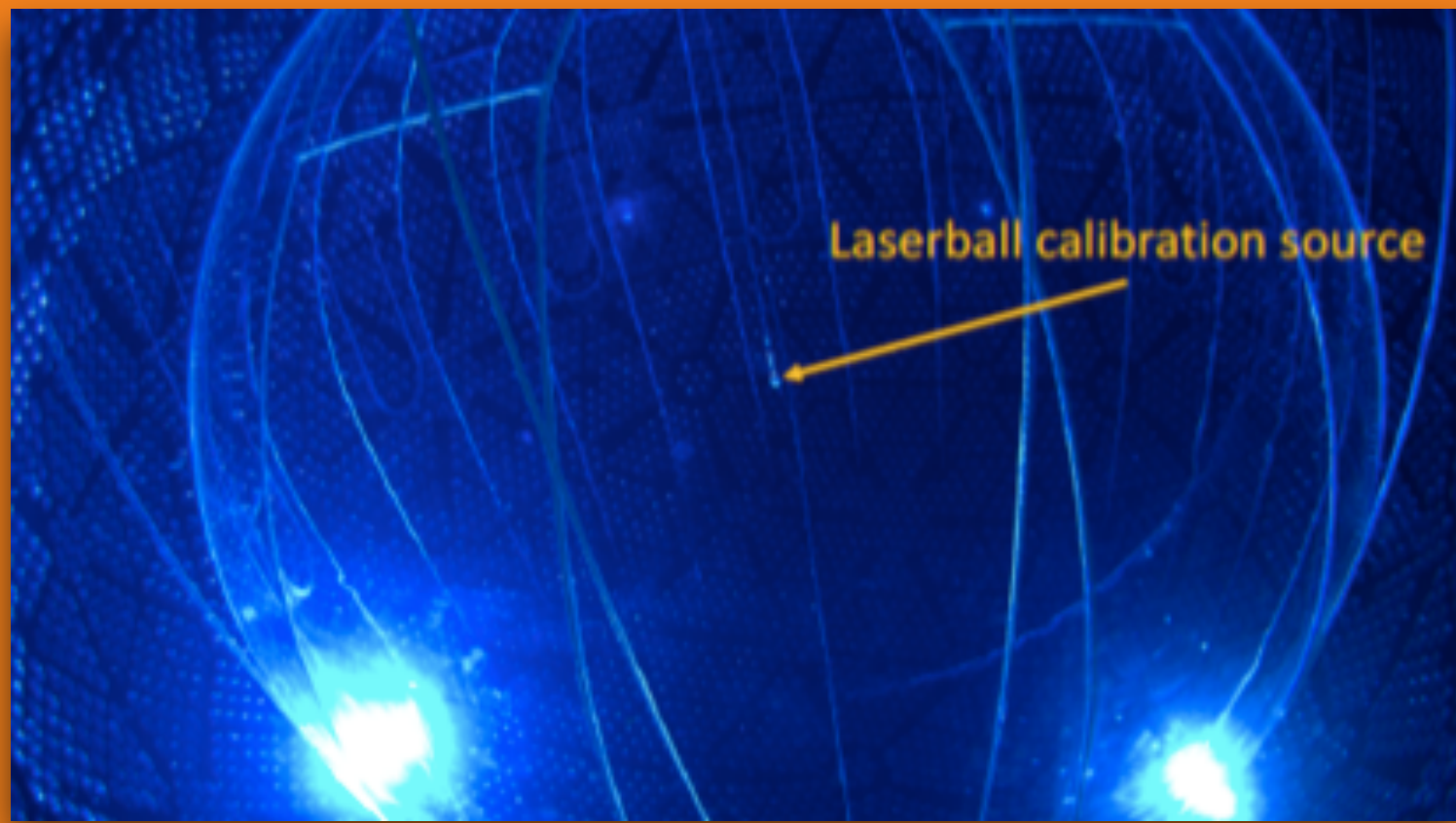


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 Attenuation Module: **AMELLIE**



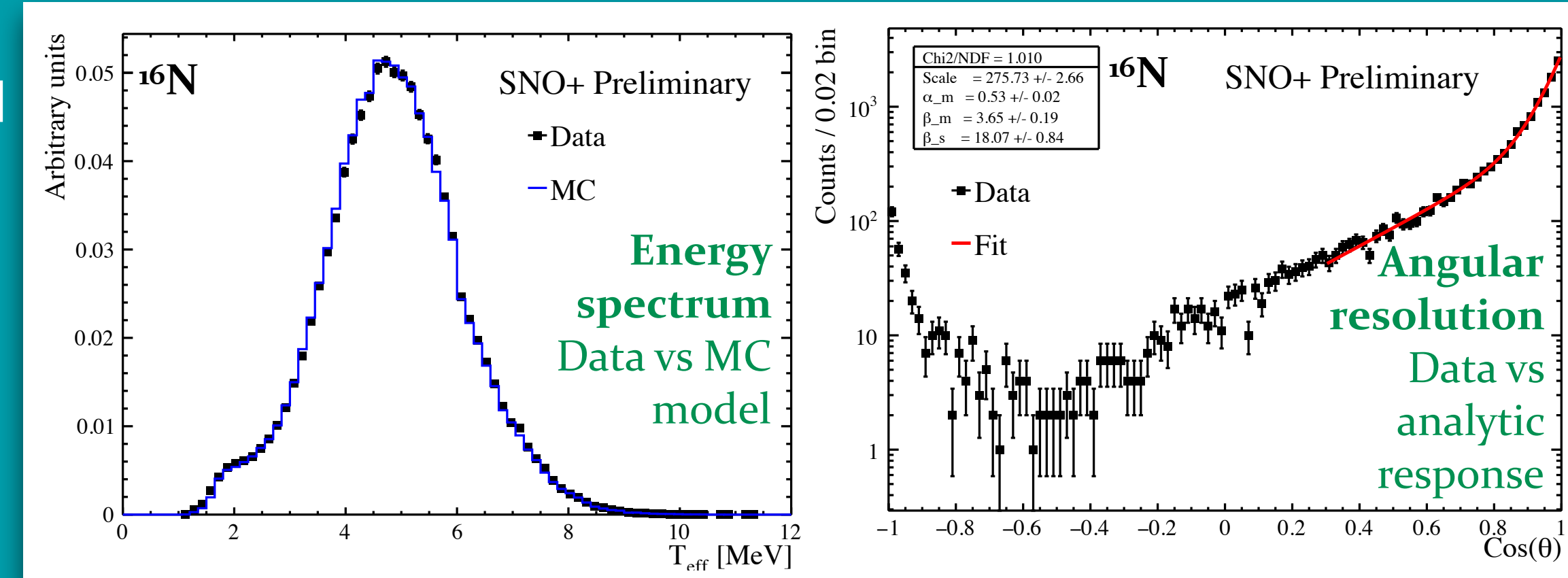


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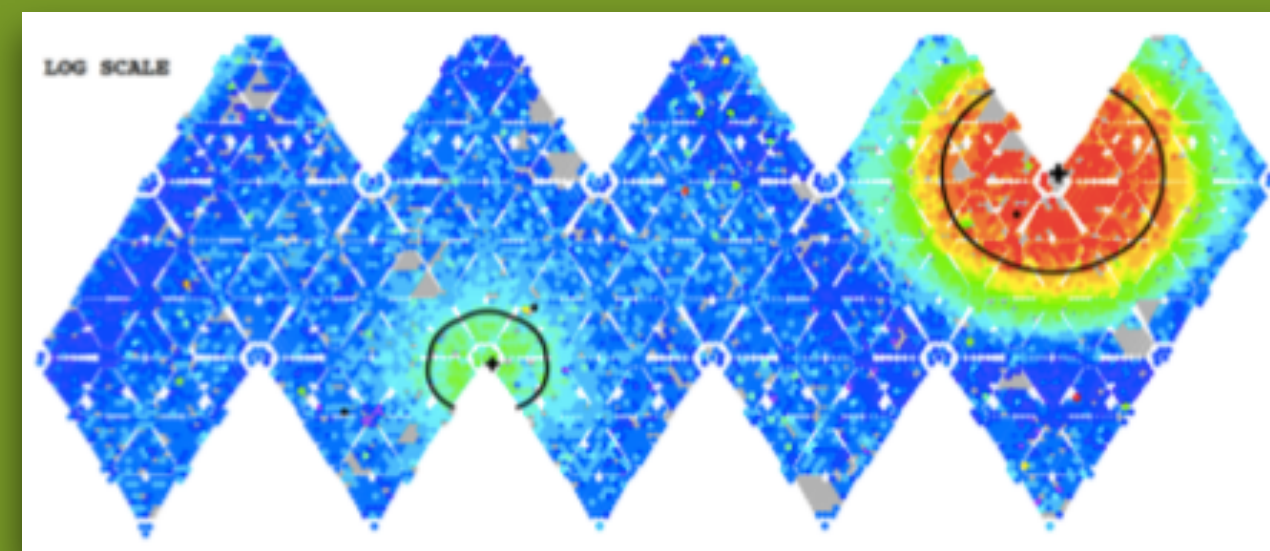
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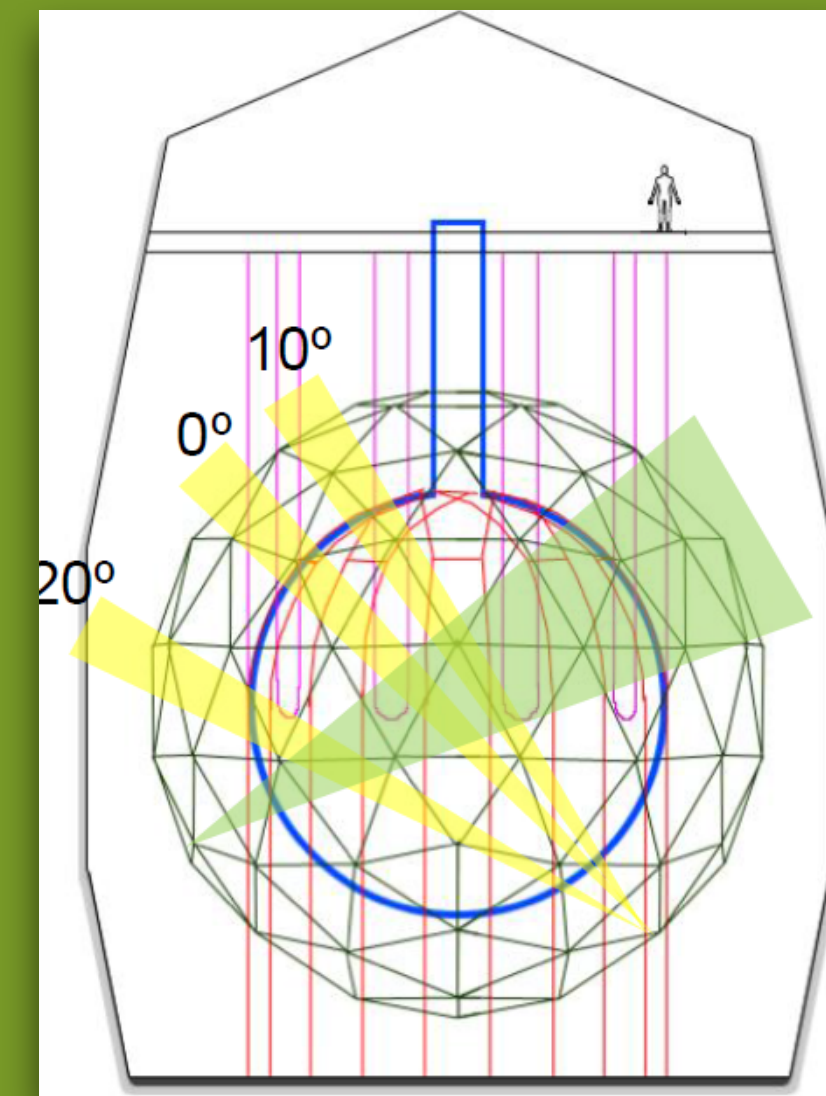


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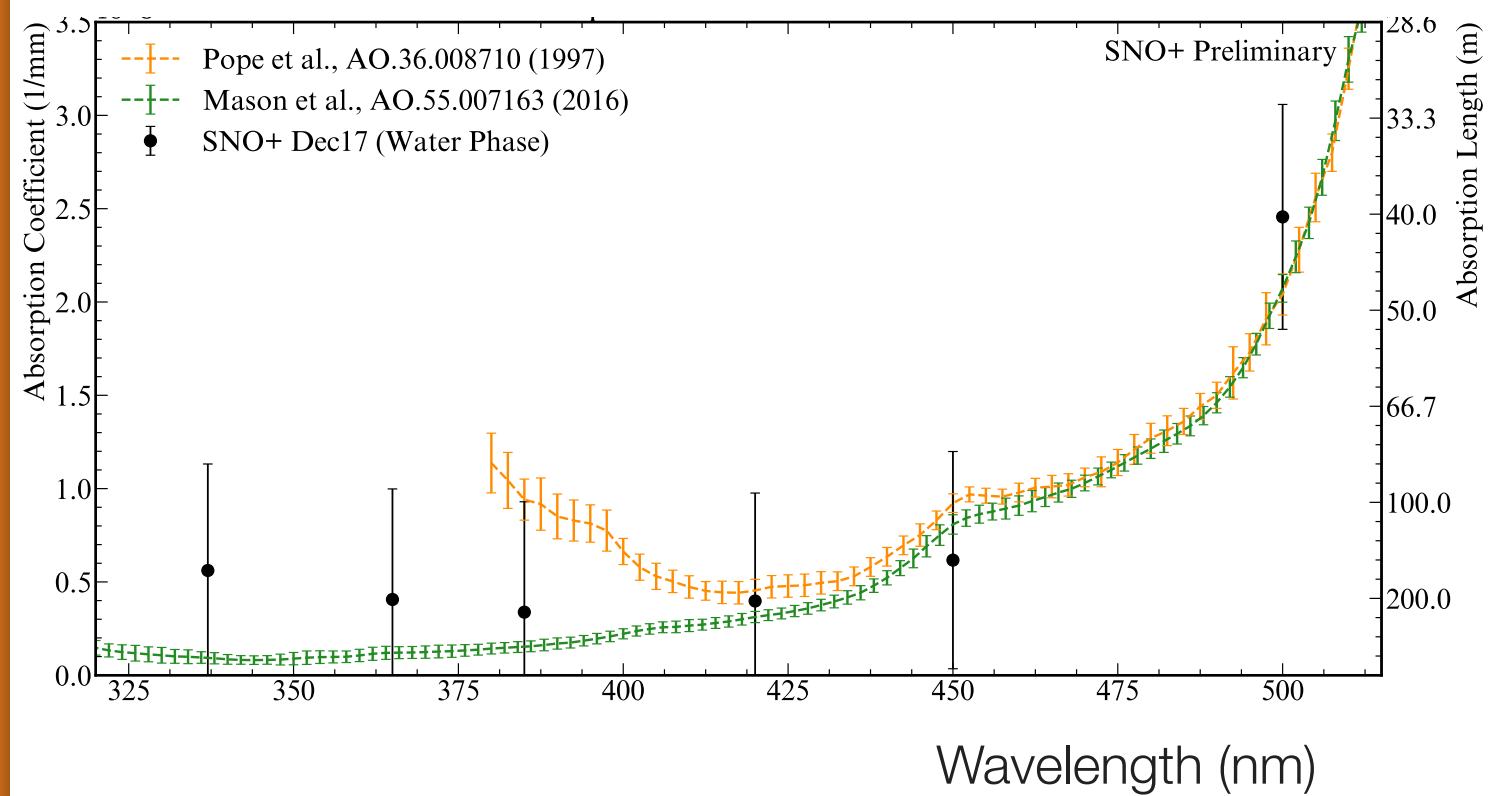
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Timing and Monitoring: **TELLIE**  
 Attenuation Module: **AMELLIE**  
 Scattering Module:

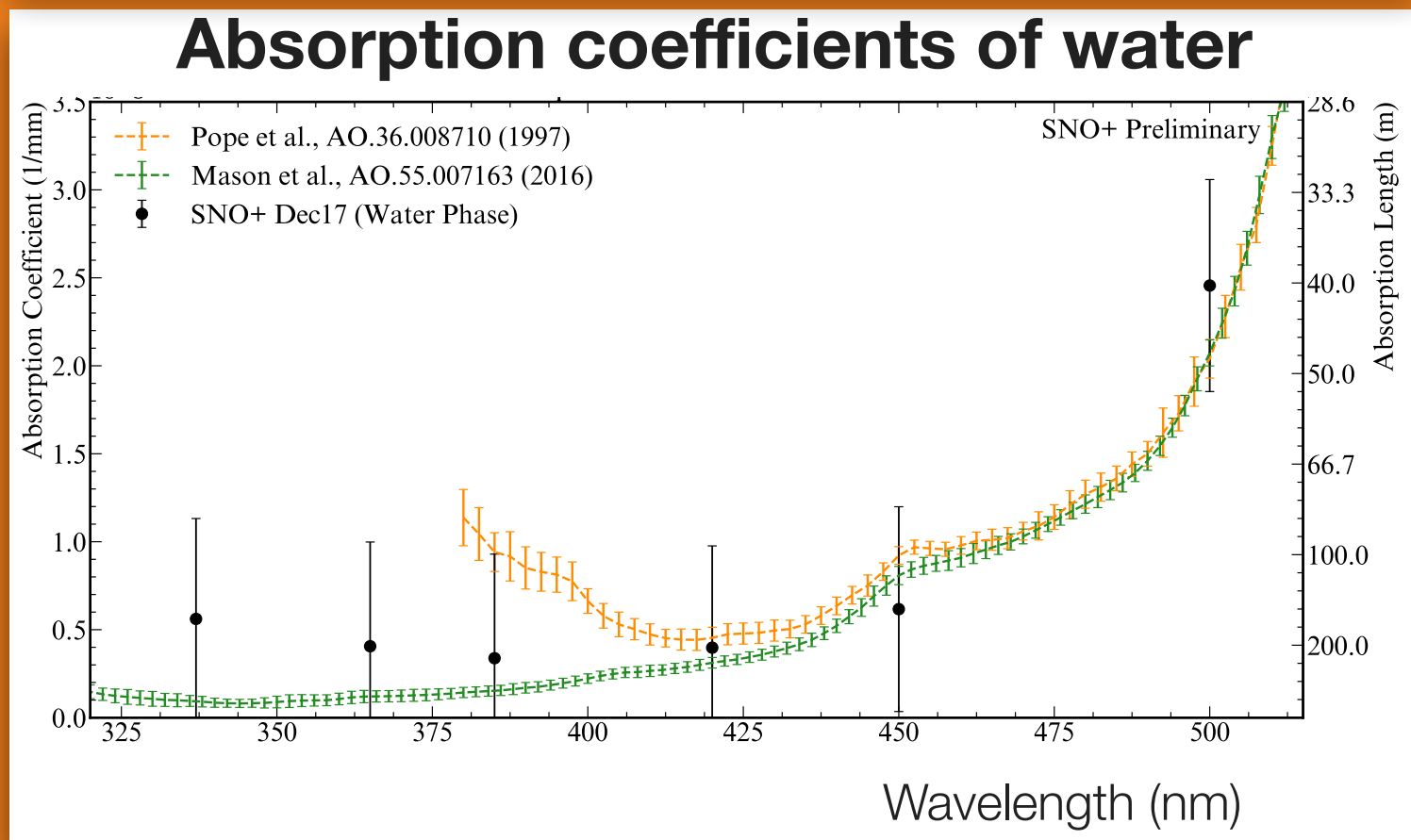
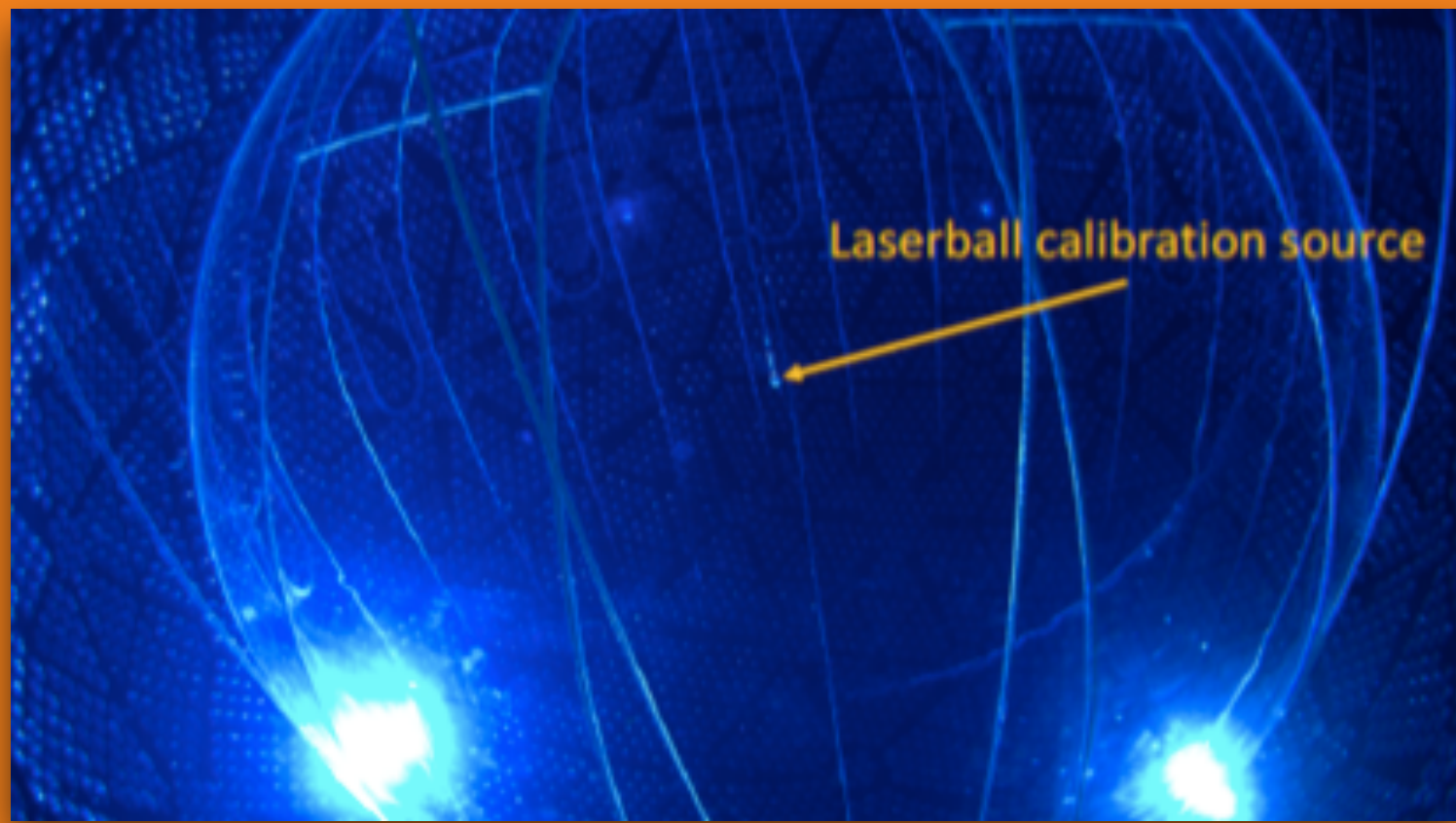


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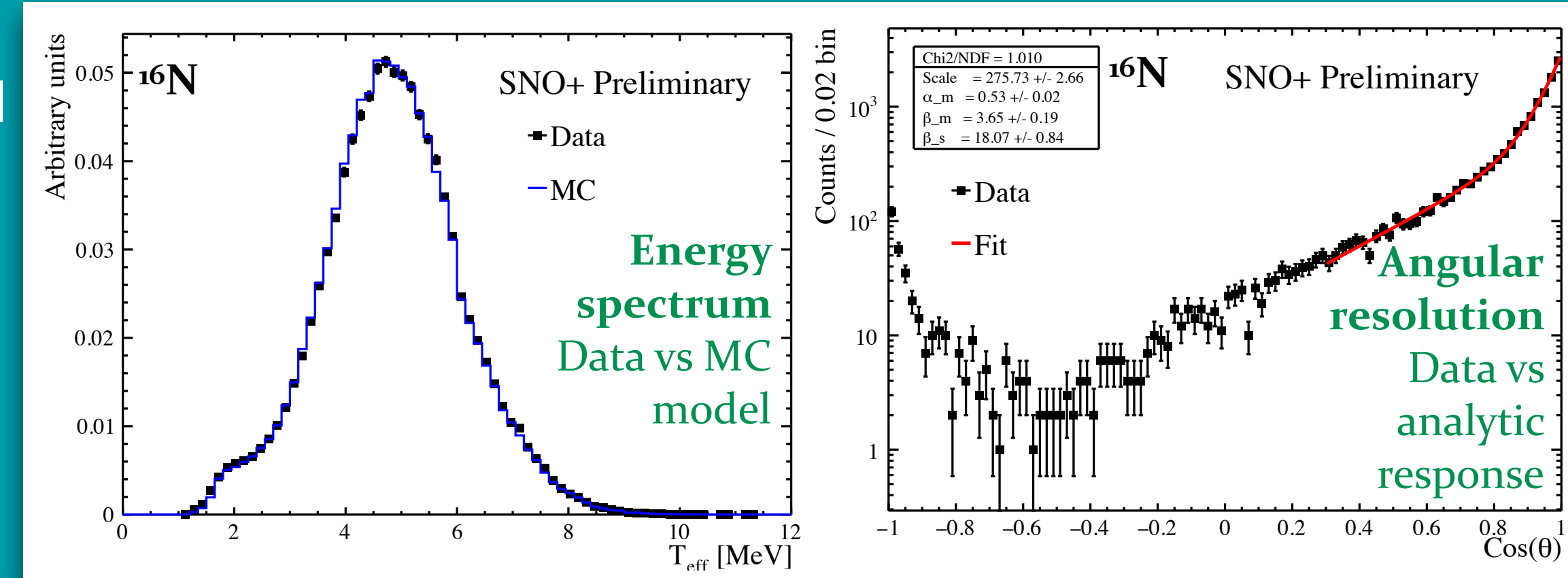


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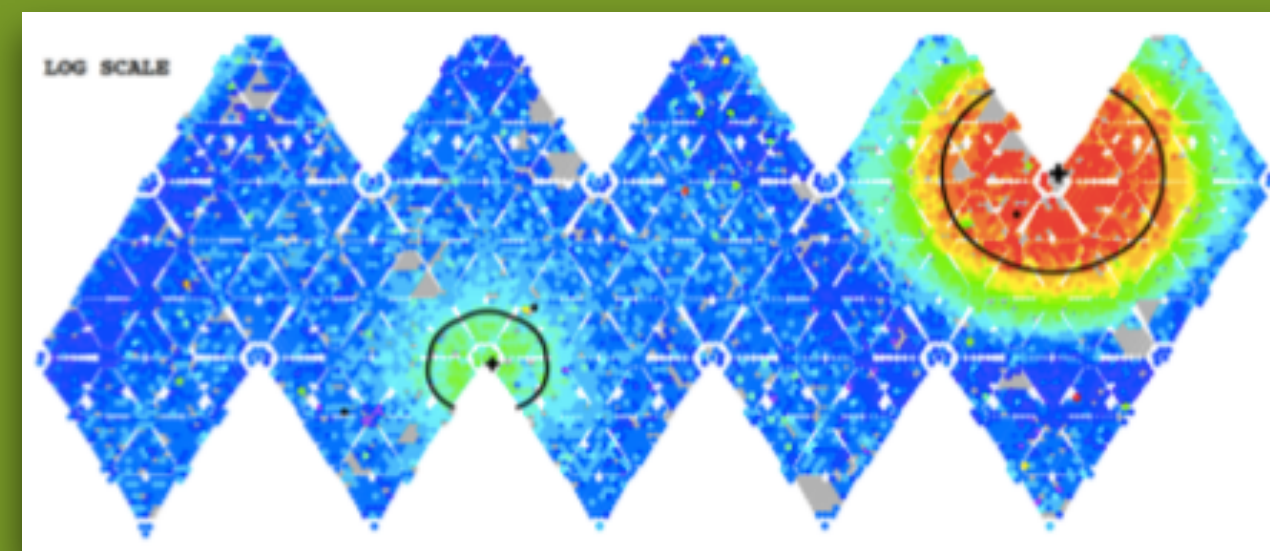
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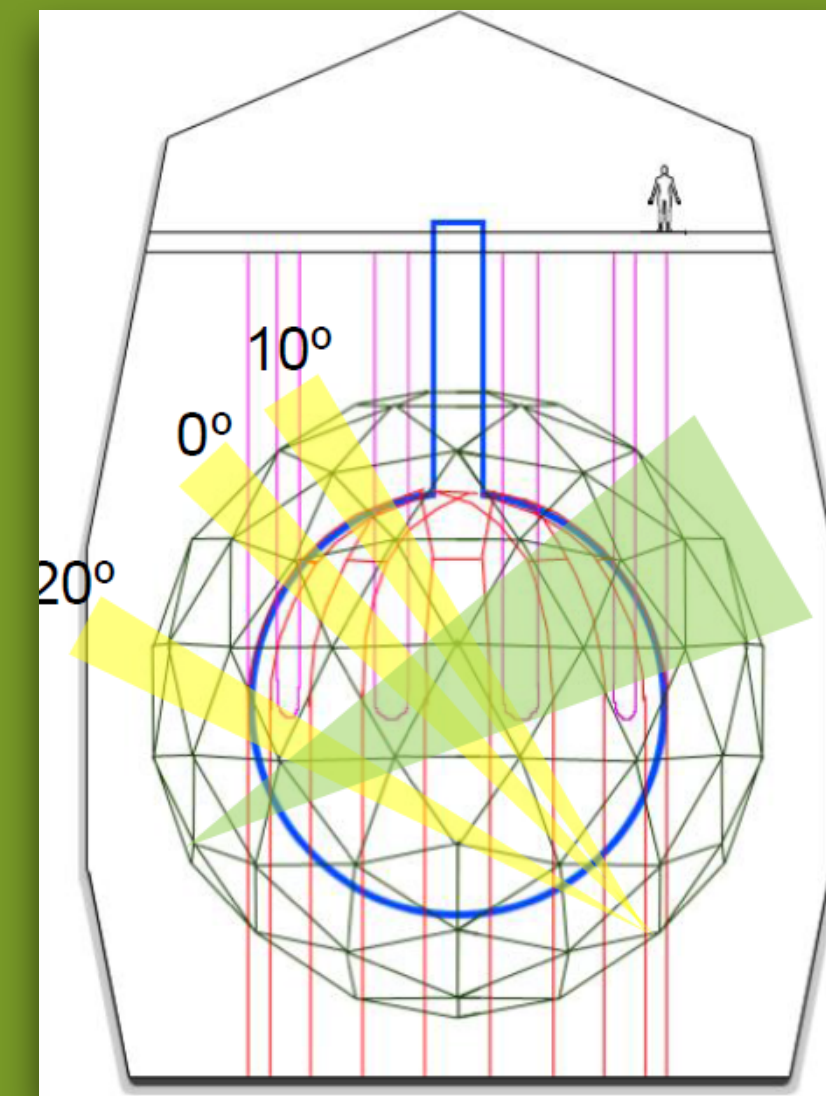


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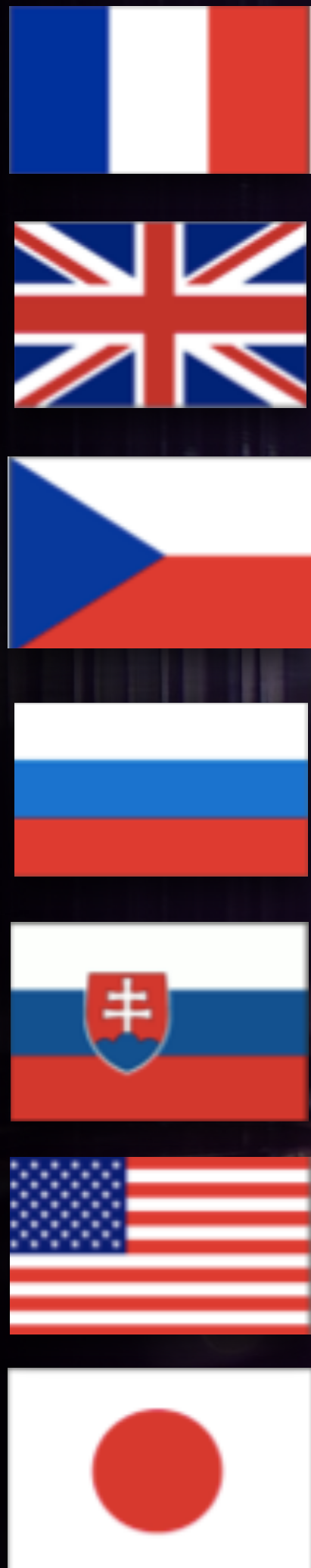
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Timing and Monitoring: **TELLIE**  
 Attenuation Module: **AMELLIE**  
 Scattering Module: **SMELLIE**



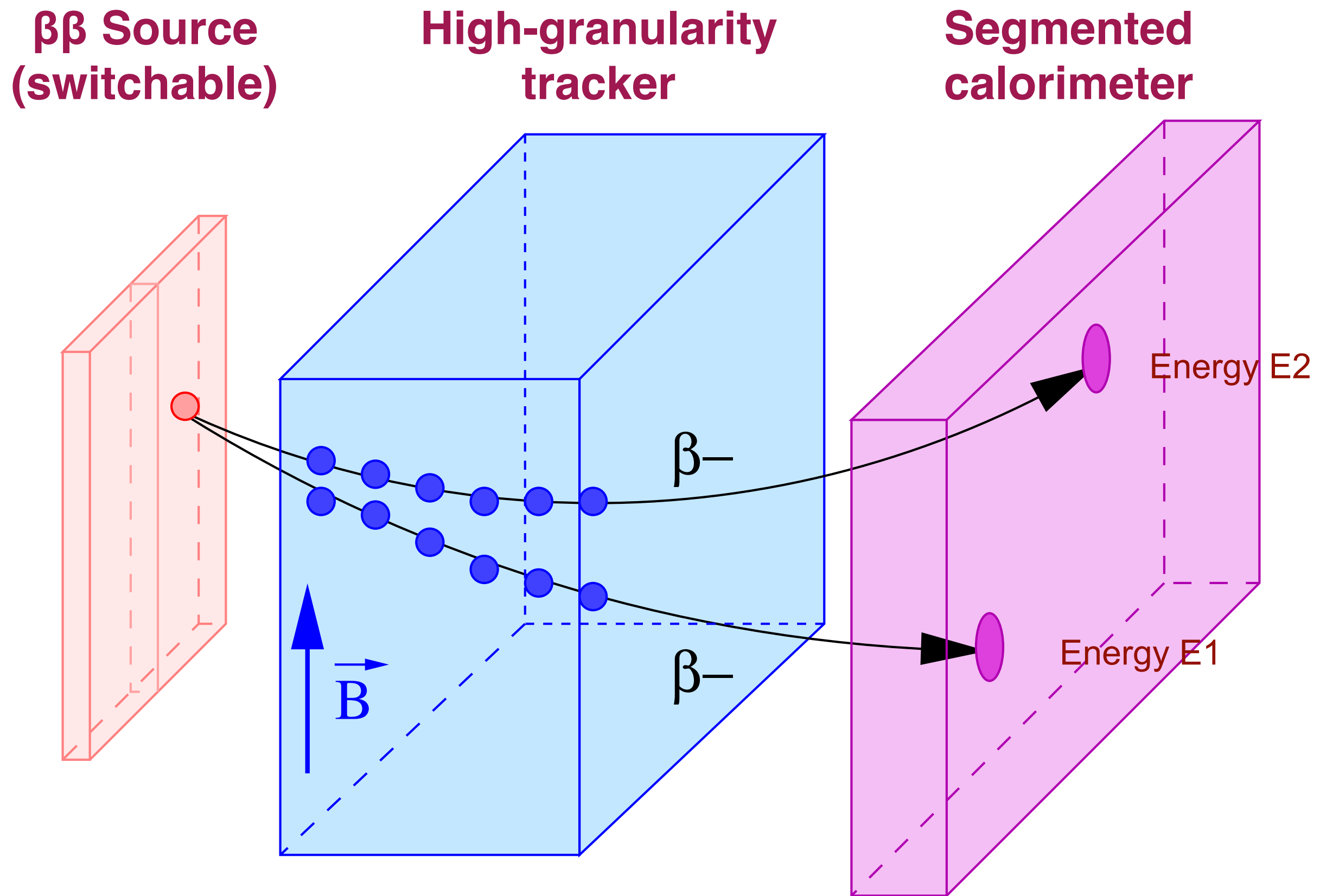
# SuperNEMO and NEMO-3, at LSM, France



A vertical stack of four university logos. At the top is the UCL logo, featuring a small building icon and the letters "UCL" in a large, bold, black font on a yellow background. Below it is the Manchester University logo, with the word "MANCHESTER" in white serif font and "1824" in yellow serif font on a purple background, with the text "The University of Manchester" in a smaller white font below. Next is the Warwick University logo, featuring a stylized "W" made of two triangles (one yellow, one blue) above the word "WARWICK" in purple serif font and "THE UNIVERSITY OF WARWICK" in a smaller purple font below. At the bottom is the Imperial College London logo, with the text "Imperial College London" in white serif font on a dark blue background.



# The NEMO principle



## Strengths



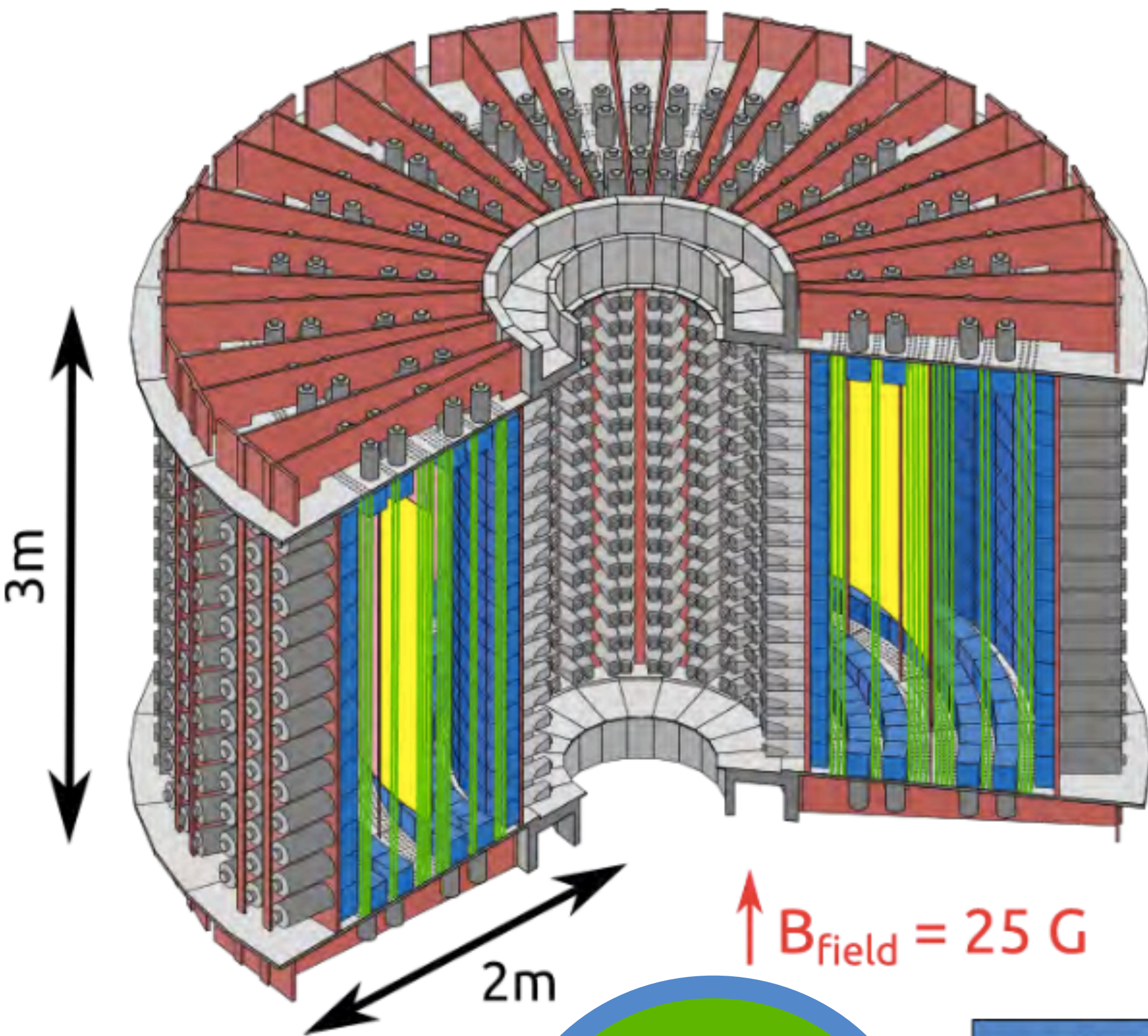
- Source decoupled from detector - use **any solid  $\beta\beta$  source** isotope
- Track reconstruction gives **particle identification**
- Combine with timings to identify topologies for ultra-high **background rejection**
- Tracking info (angle between tracks) & individual energy distributions can distinguish between  **$\beta\beta$  mechanisms**

## Weaknesses

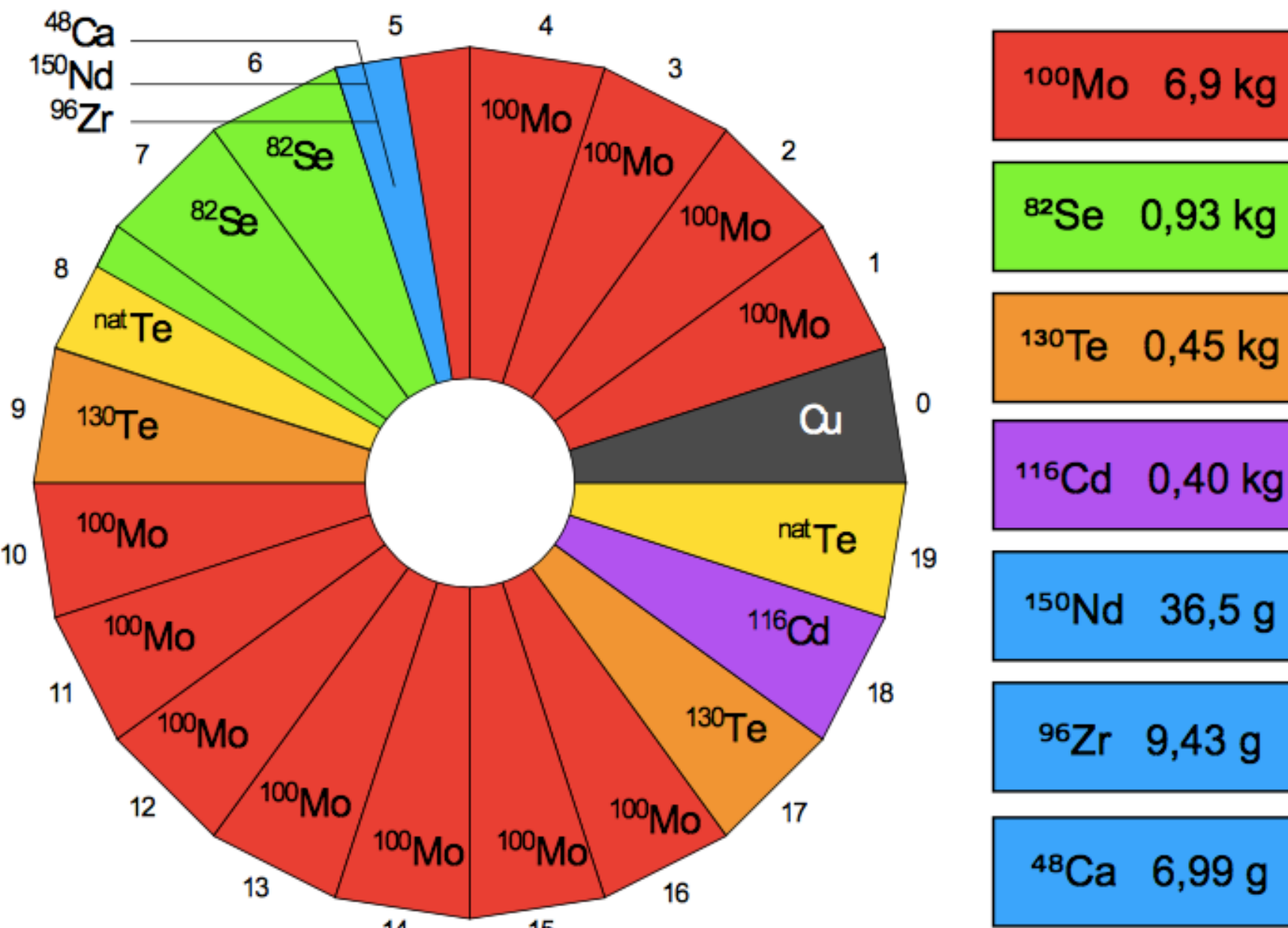


- **Energy resolution** poorer than for most homogenous detectors
- Doesn't scale as well as some other designs

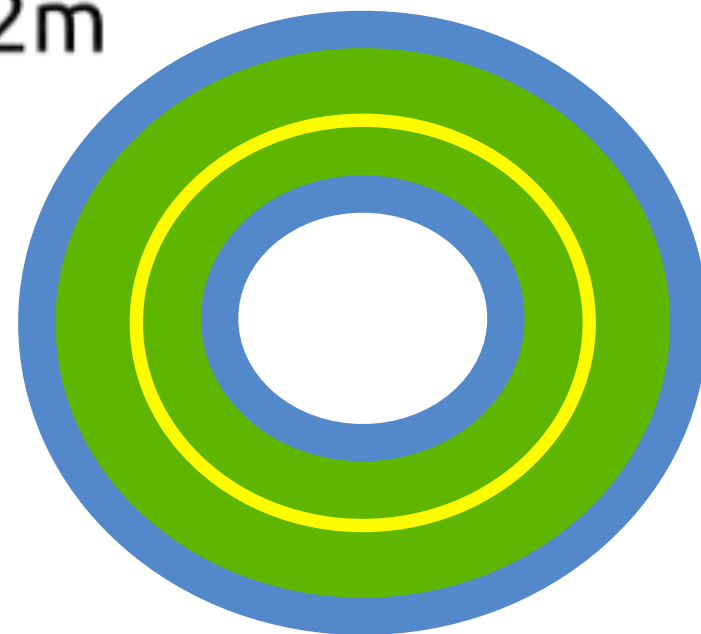
# NEMO-3 (2003-2011)



NEMO-3 "camembert" (source top view)



$\uparrow B_{\text{field}} = 25 \text{ G}$



**calorimeter**  
 1940 optical modules :  
 polystyren scintillators  
 + 3" and 5" PMTs  
 $\text{FWHM}_E \sim 15\% / \sqrt{E_{\text{MeV}}}$   
 $\sigma_t \sim 250 \text{ ps}$

**tracker**  
 6180 Geiger cells  
 vertex resolution :  
 $\sigma_{xy} \sim 3 \text{ mm}$   $\sigma_z \sim 10 \text{ mm}$

**sources**  
 60 mg/cm<sup>2</sup> foils  
 10 kg of  $\beta\beta$  isotopes

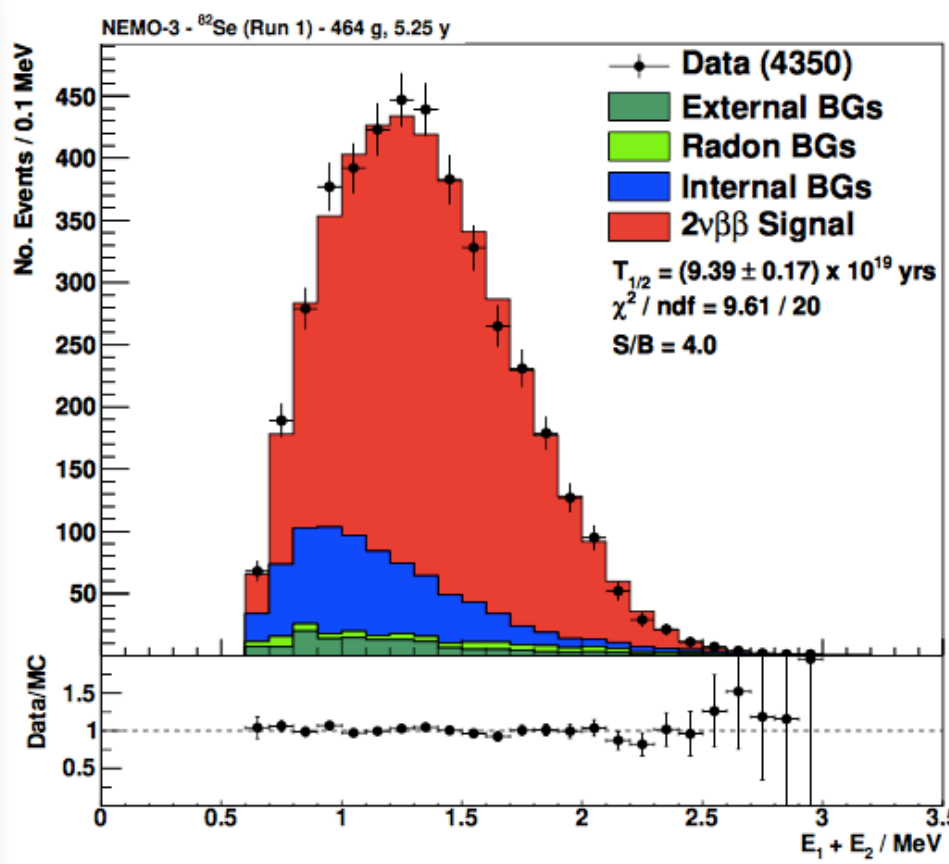


**NEW**

# results from NEMO-3

2νββ measurements and 0νββ limit

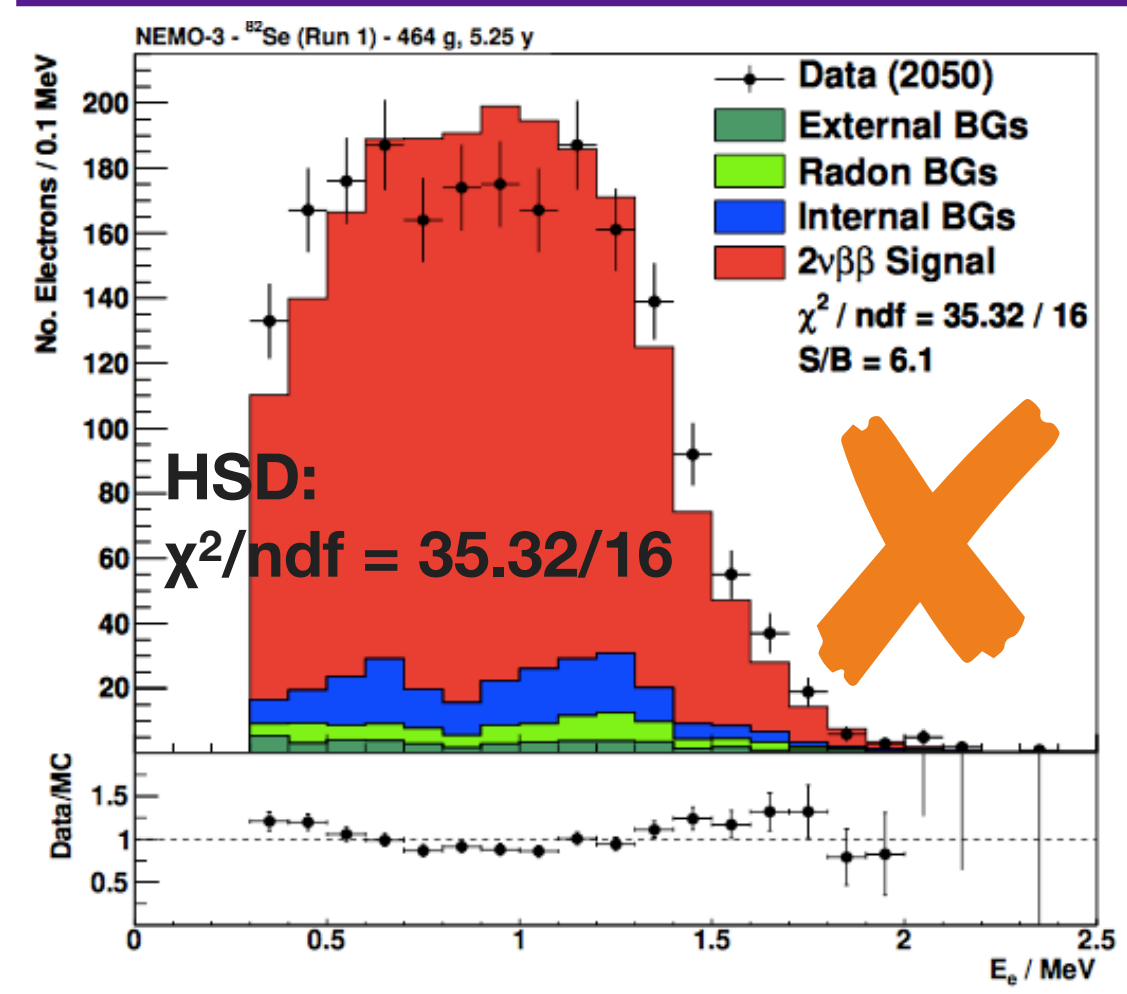
- **<sup>82</sup>Se** (Eur. Phys. J. C (2018) 78: 821)



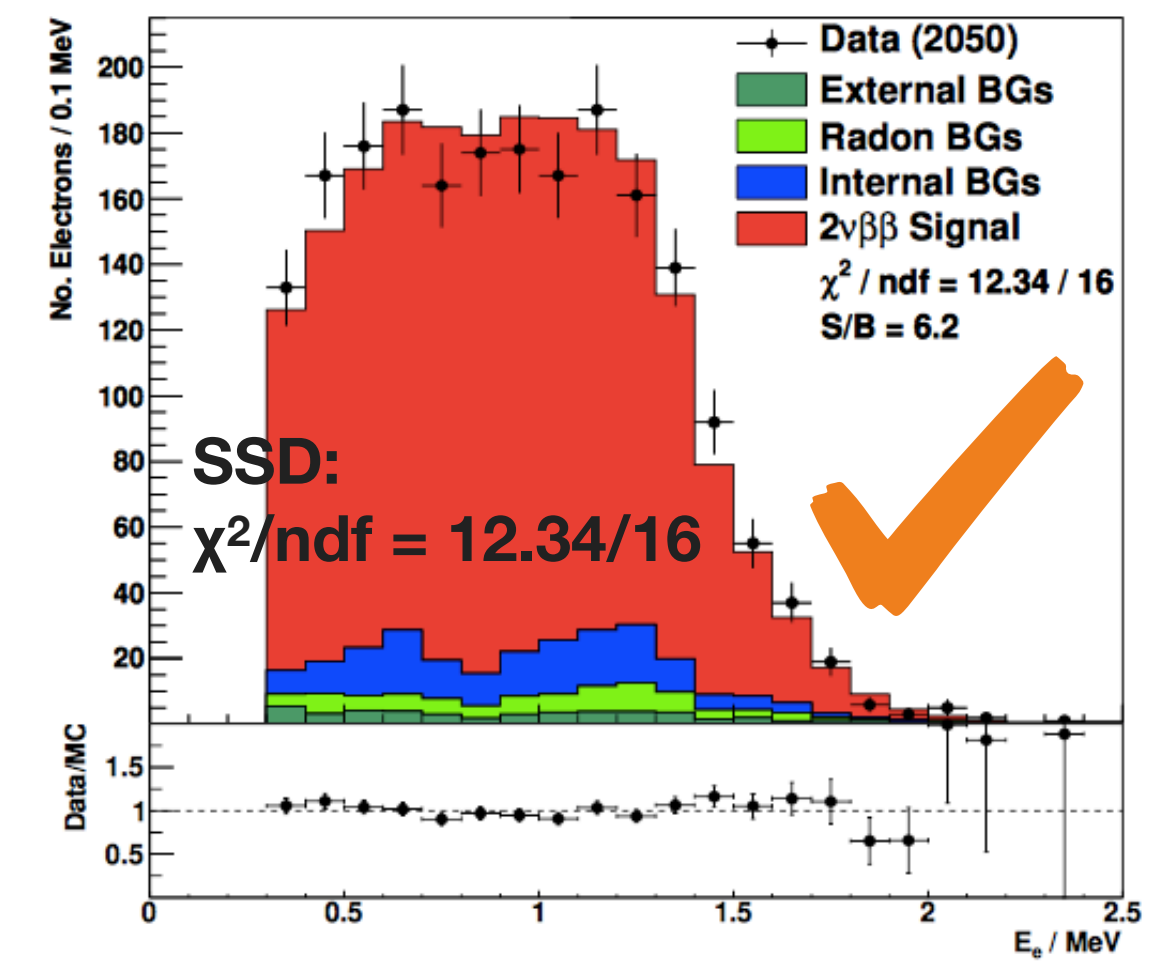
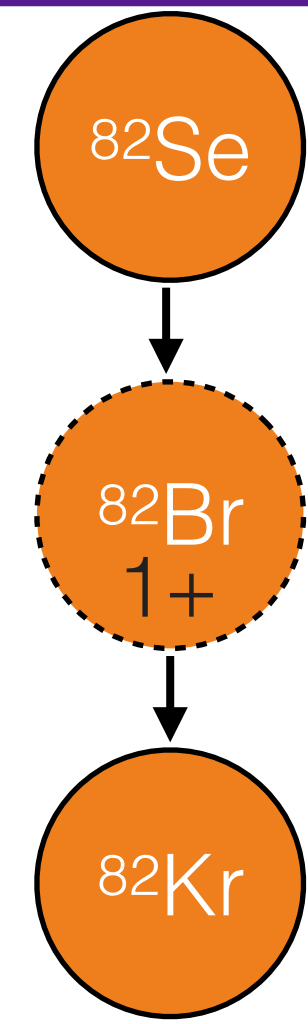
Summed 2-electron spectrum

**2νββ:**  
 $T_{1/2} = 9.39 \pm 0.17$  (stat)  $\pm 0.58$  (sys)  $\times 10^{19}$  years  
 (SSD hypothesis)

**0νββ:**  
 $T_{1/2} > 2.5 \times 10^{23}$  years (90% C.L.)



Higher state dominated - many excited states



Single state dominated - mostly one intermediate state

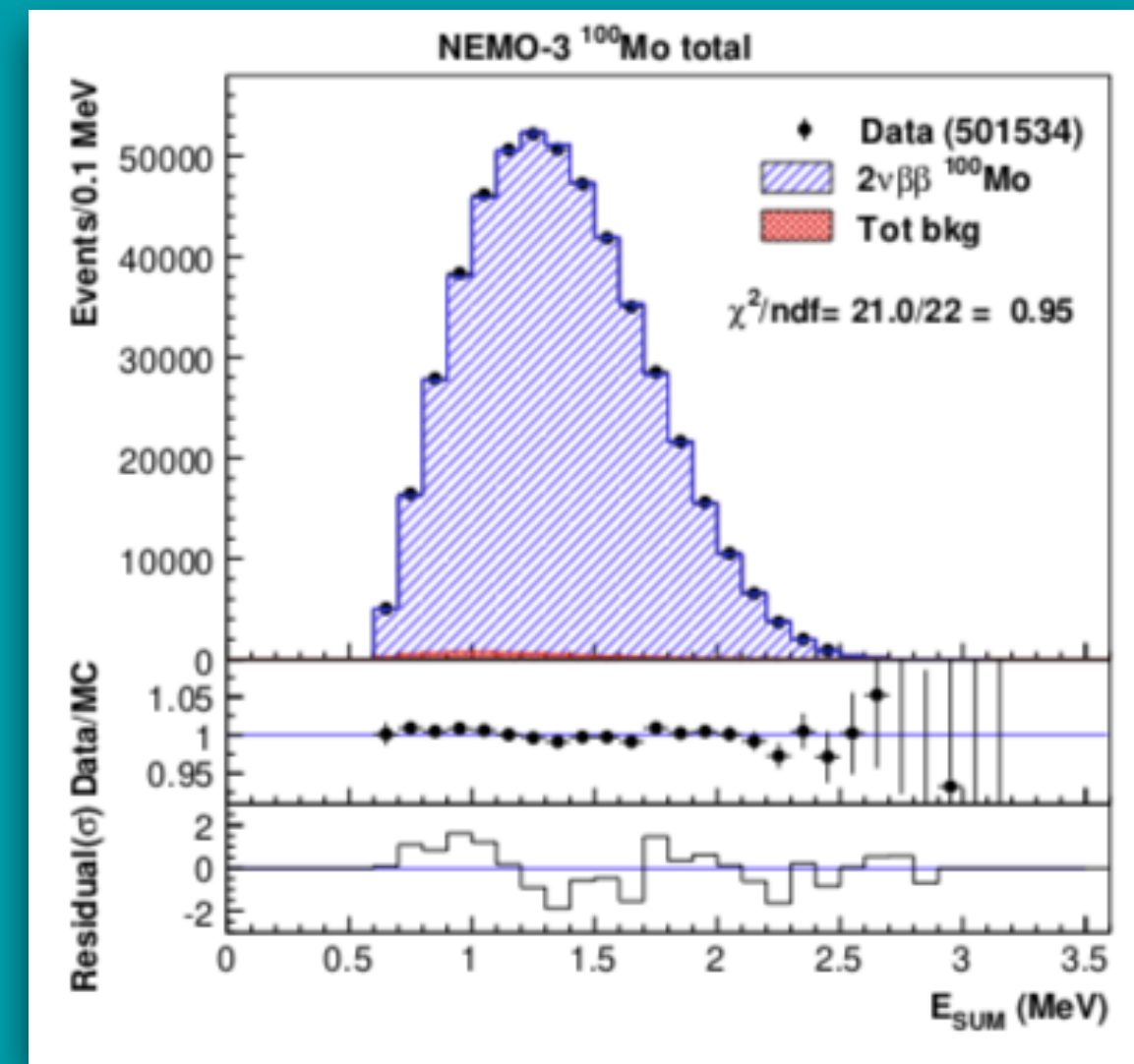
**Individual electron spectrum tells us about intermediate 1+ states**

**NEW**

# results from NEMO-3

2νββ measurements and 0νββ limit

- <sup>82</sup>Se (Eur. Phys. J. C (2018) 78: 821)
- <sup>100</sup>Mo [arXiv 903.08084 \[nucl-ex\]](https://arxiv.org/abs/1903.08084)



**2νββ**

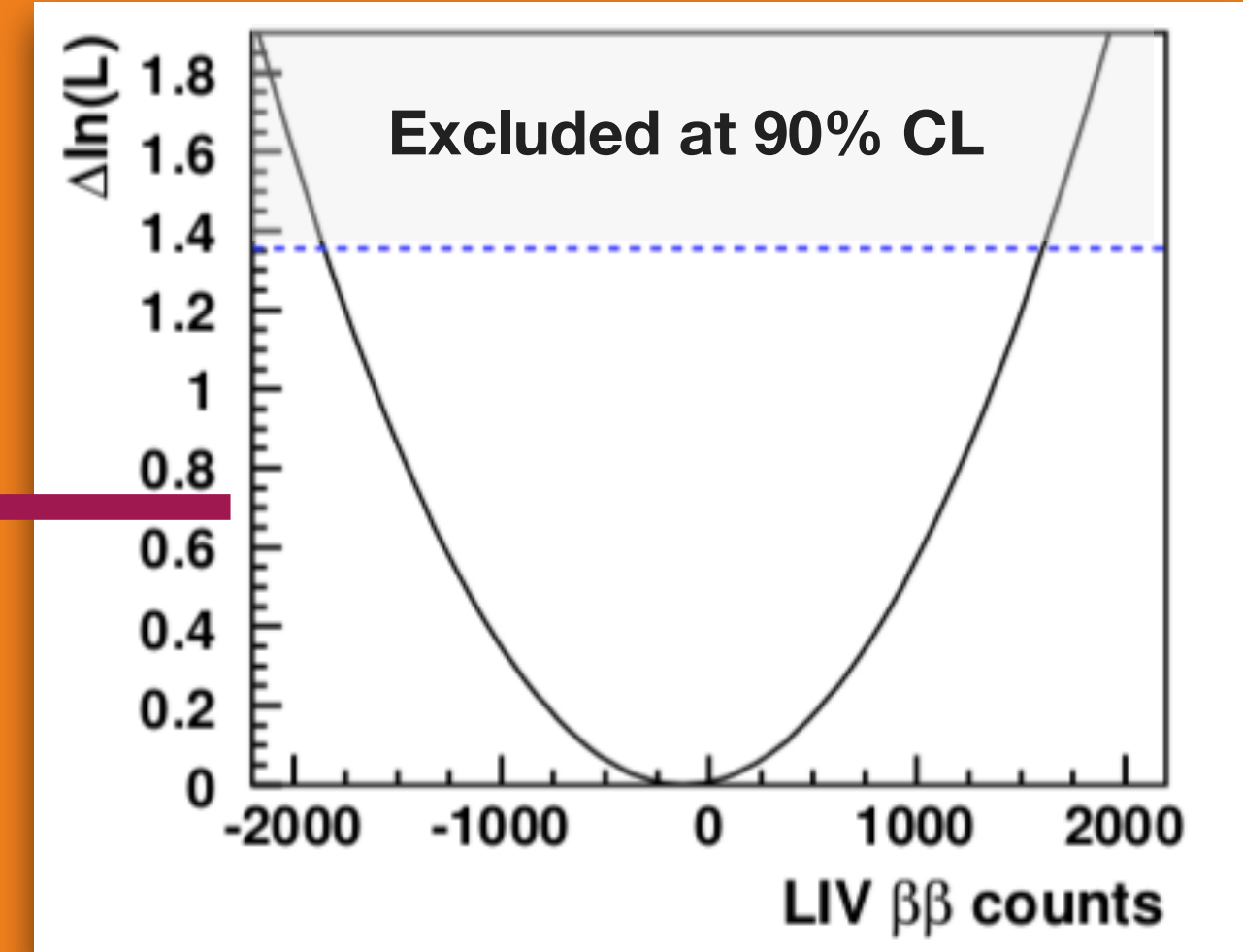
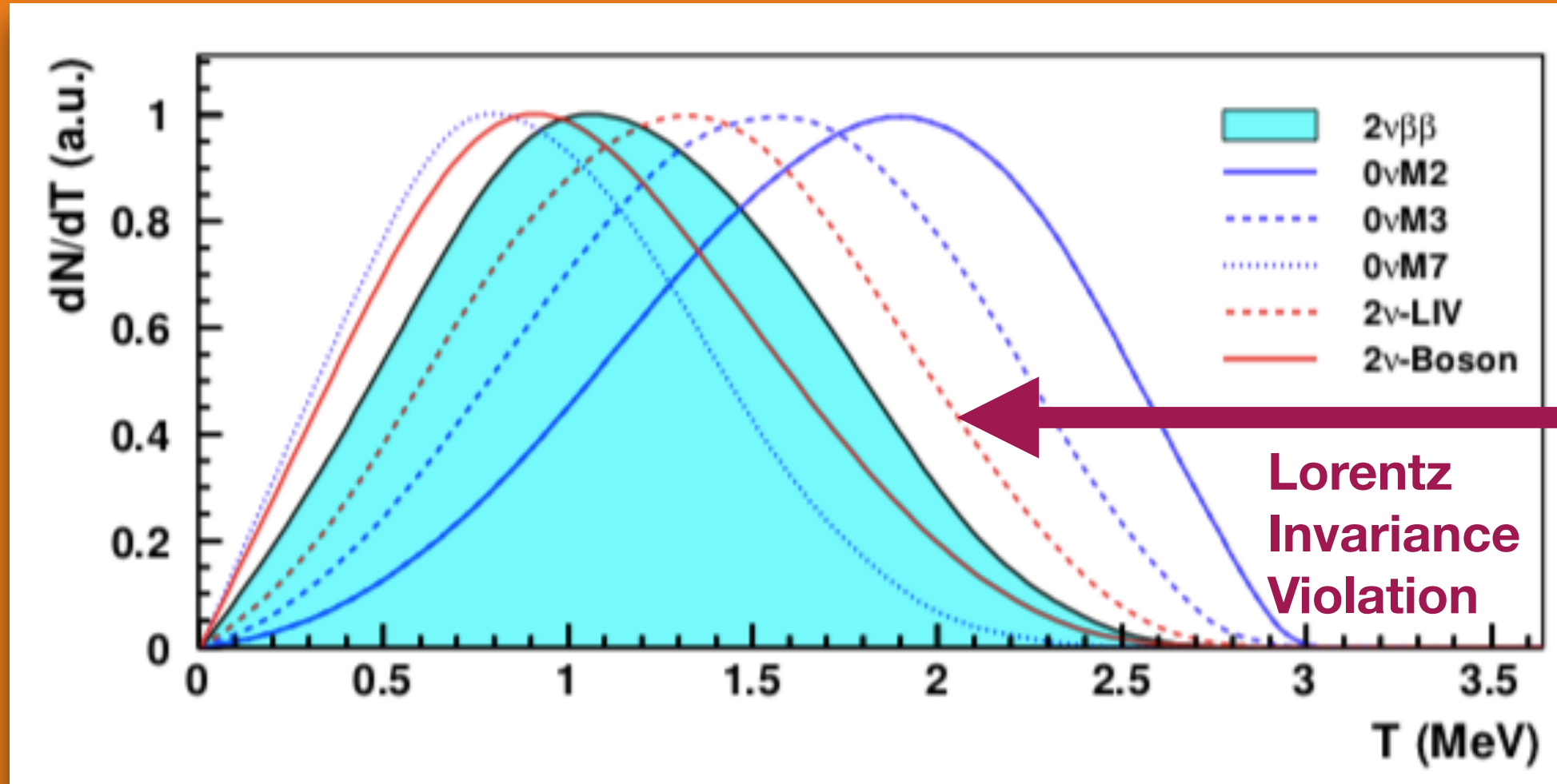
$T_{1/2} = 6.81 \pm 0.01$  (stat)  $\pm 0.46$  (sys)  $\times 10^{18}$  years (SSD model)



- Over  $5 \times 10^5$  events with S/B  $\approx 80$
- **Lorentz Invariance Violation** and exotic 0νββ mechanisms would modify energy spectrum
- Limit set on contribution from Lorentz-Invariance violating events

$$-4.2 \times 10^{-7} \text{ GeV} < \dot{a}_{of}^{(3)} < 3.5 \times 10^{-7} \text{ GeV (90\% C.L.)}$$

**Best published result!**

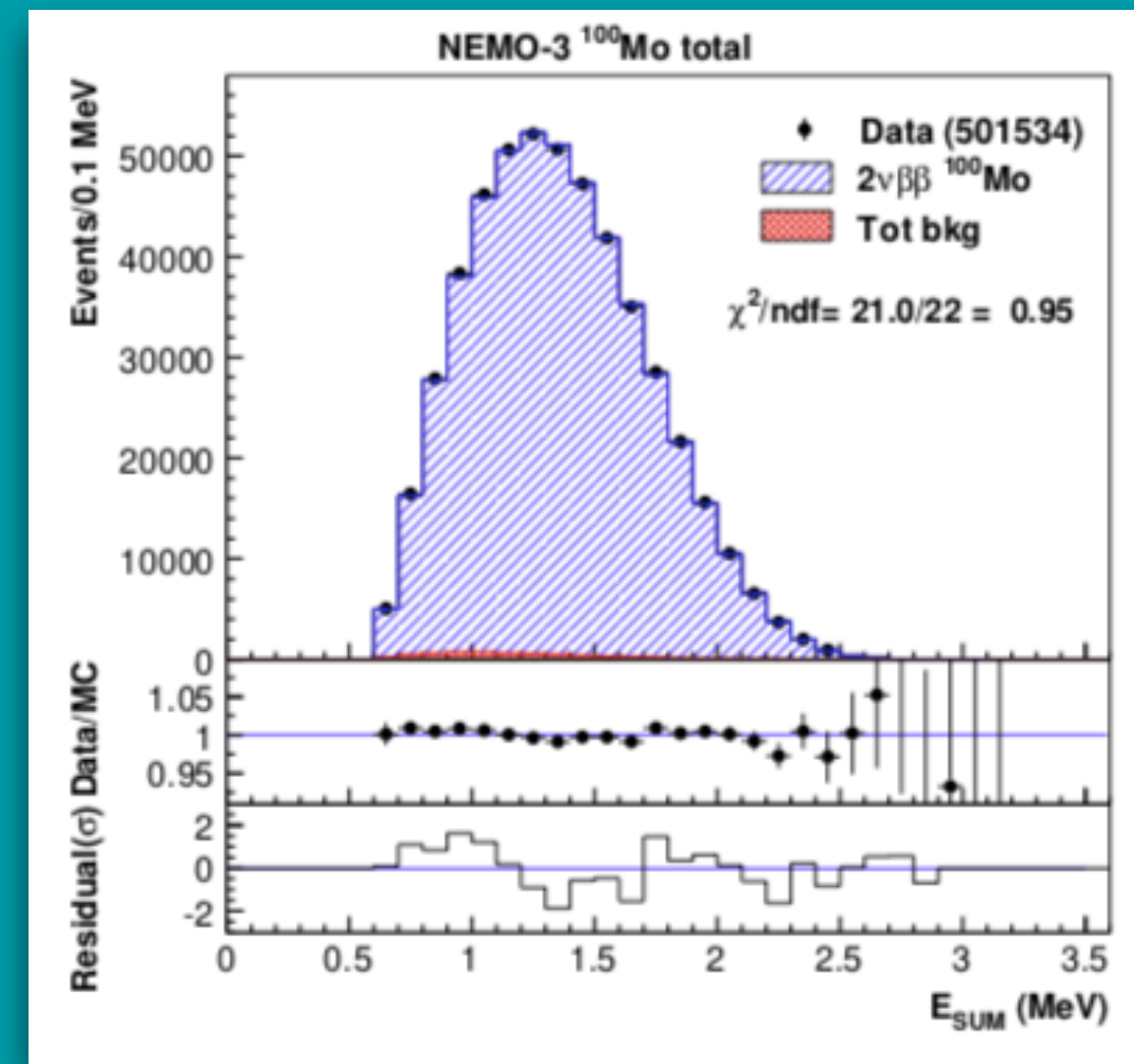


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- **<sup>48</sup>Ca** (Phys. Rev. D 93, 112008)
- **<sup>150</sup>Nd** (Phys. Rev. D 94, 072003)
- **<sup>116</sup>Cd** (Phys. Rev. D 95, 012007)
- **<sup>130</sup>Te** (Phys. Rev. Lett. 107, 062504)
- **<sup>96</sup>Zr** (Nucl.Phys.A847:168-179)



**2νββ**

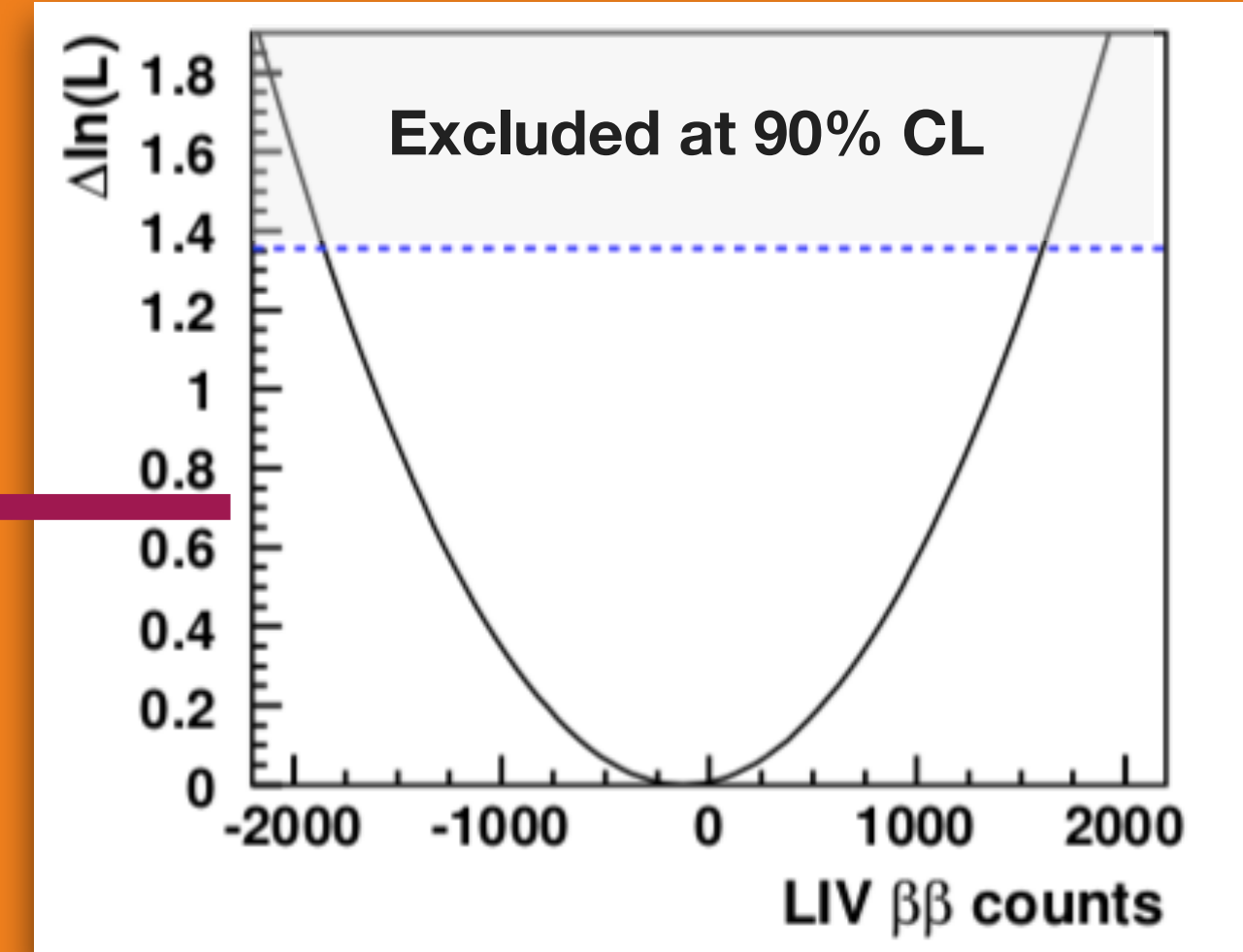
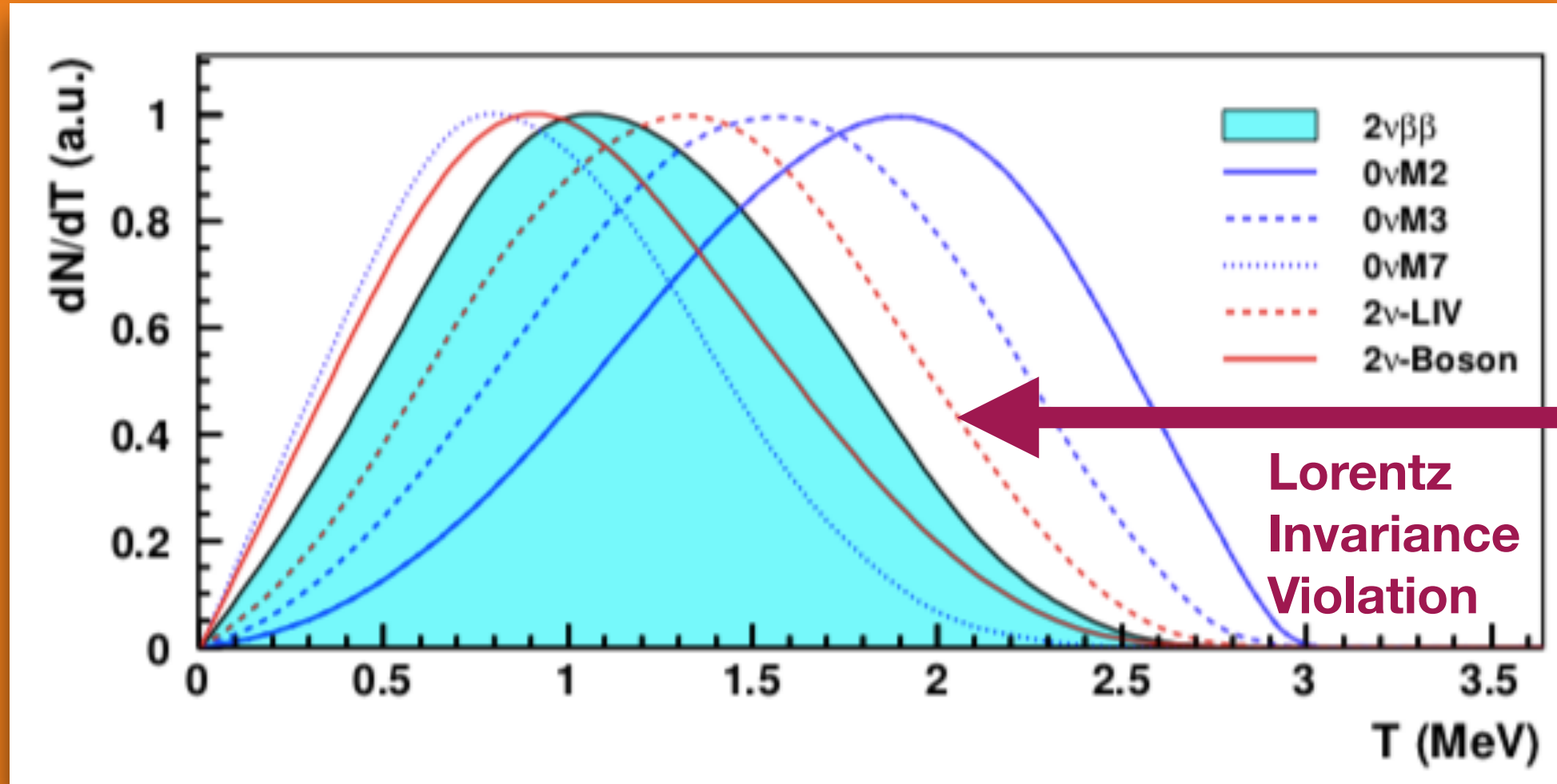
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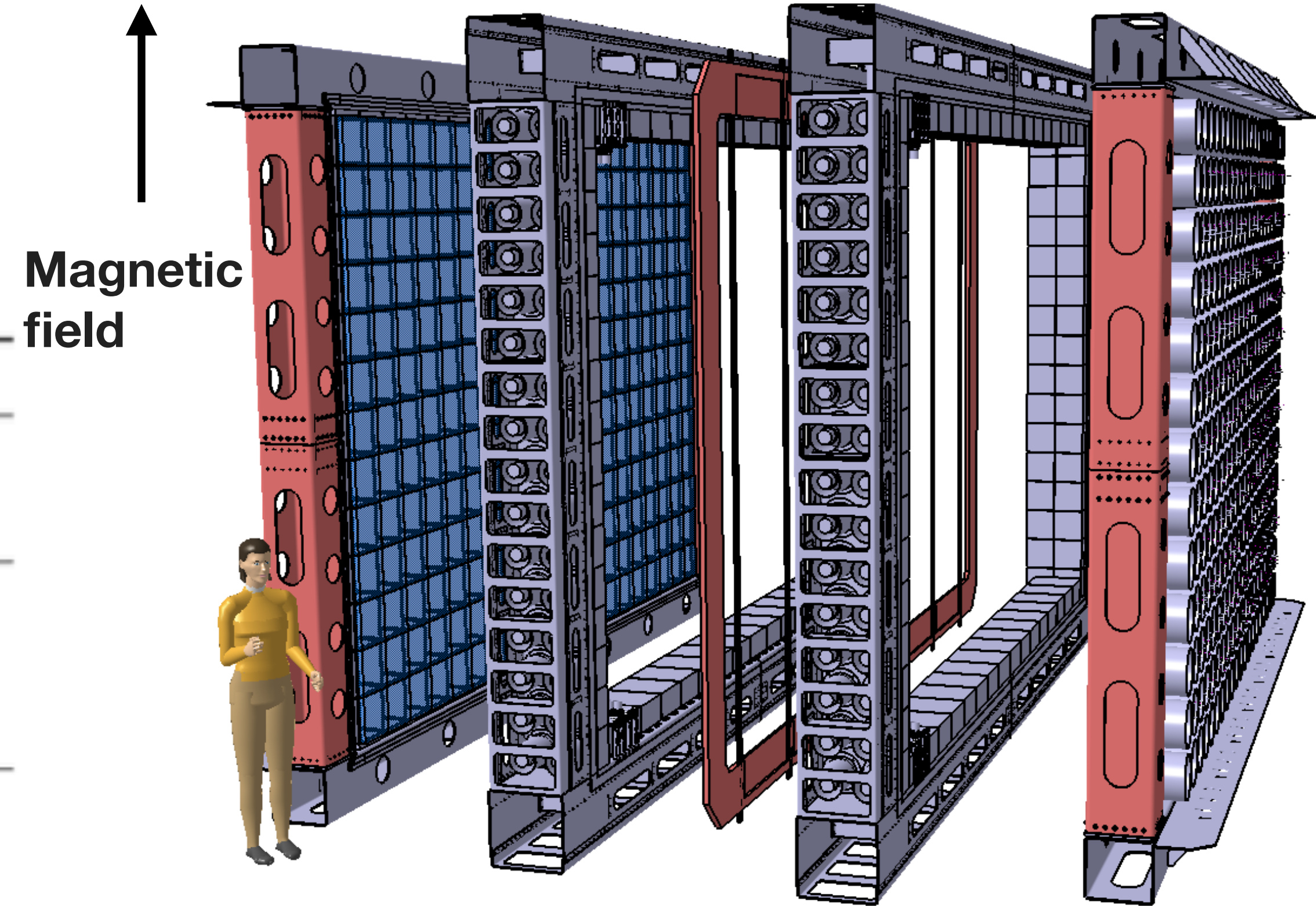
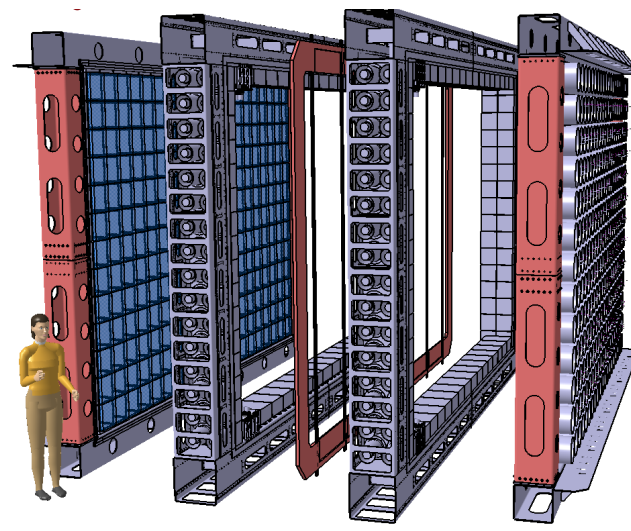
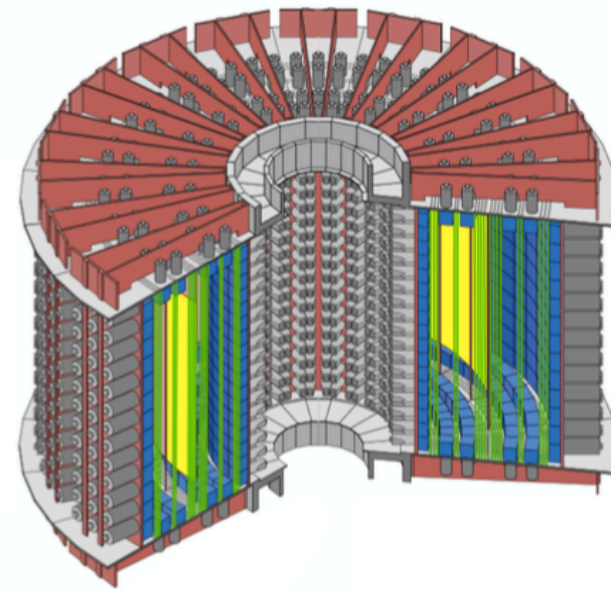
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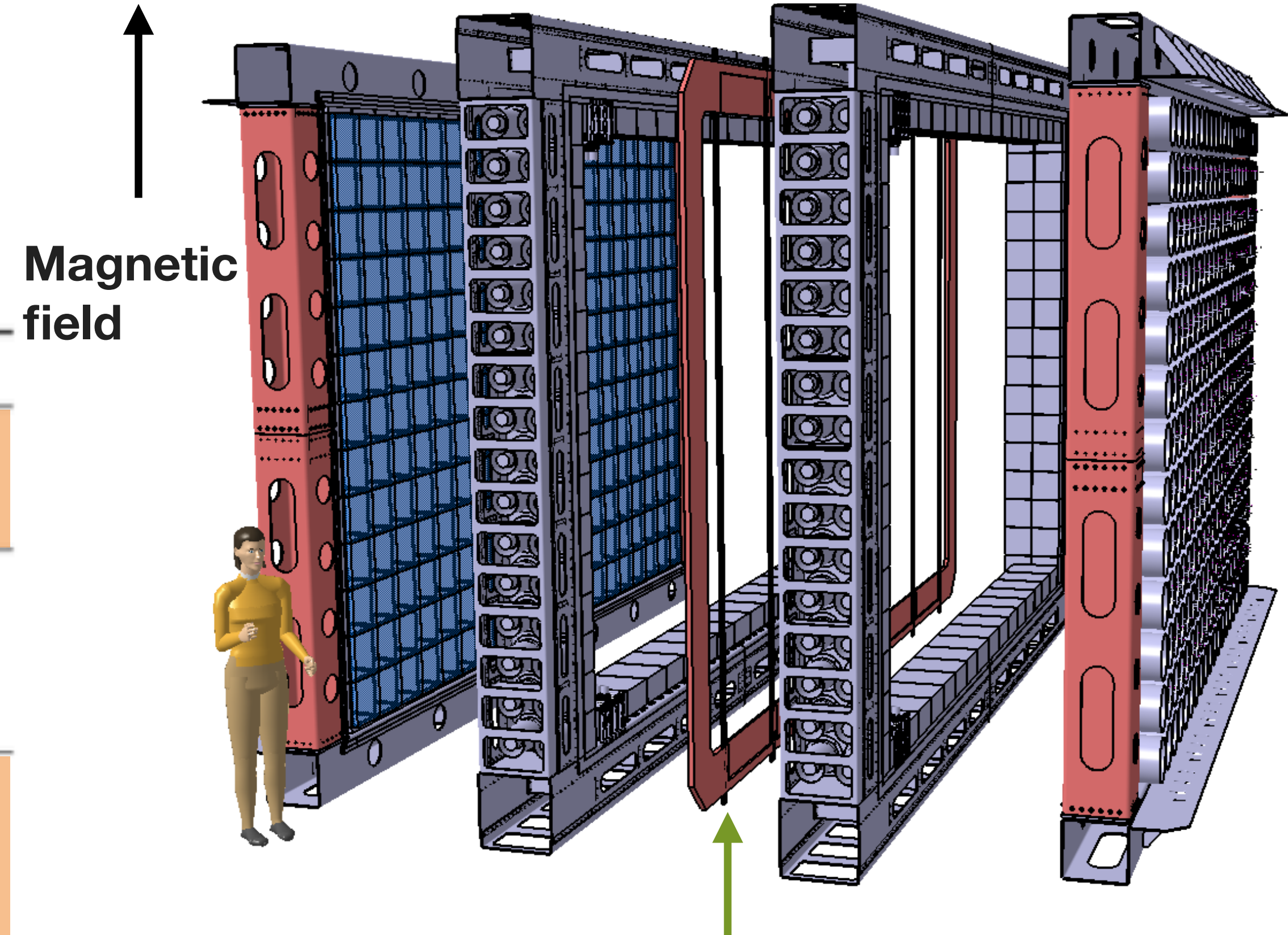
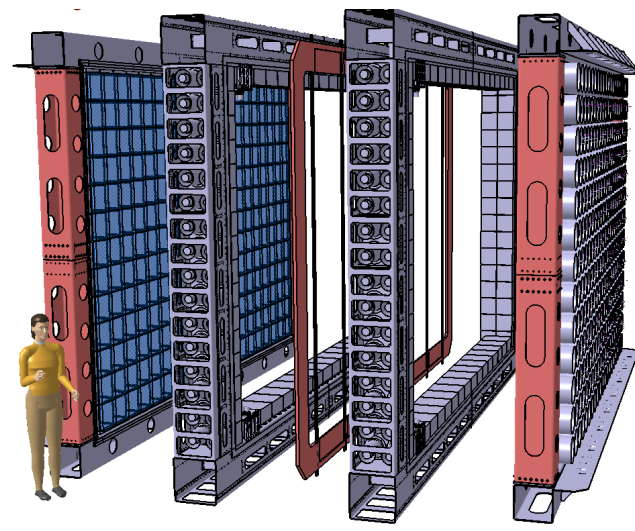
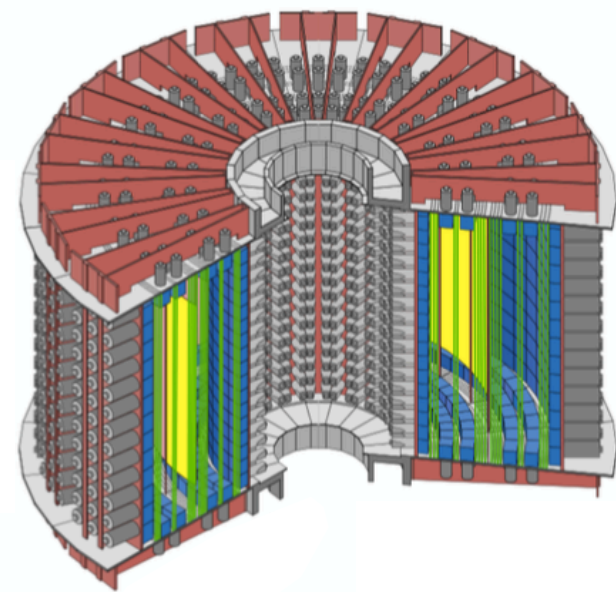
# Currently at LSM: SuperNEMO Demonstrator



Magnetic field

	NEMO-3	SuperNEMO demonstrator
Mass [kg] (main isotopes)	7 ( $^{100}\text{Mo}$ )	6.3 ( $^{82}\text{Se}$ )
$T_{1/2}^{2\nu}$ [y]	$6.8 \times 10^{18}$	$9.4 \times 10^{19}$
Energy resolution FWHM at 1 MeV	15 %	8 %
FWHM at 3 MeV	8 %	4 %
Source radiopurity $A(^{208}\text{Tl})$	$\sim 100 \mu\text{Bq/kg}$	$< 2 \mu\text{Bq/kg}$
$A(^{214}\text{Bi})$	$< 300 \mu\text{Bq/kg}$	$< 10 \mu\text{Bq/kg}$
Level of radon $A(^{222}\text{Rn})$	$\sim 5.0 \text{ mBq/m}^3$	$< 0.15 \text{ mBq/m}^3$
Sensitivity after 5 (2.5) y data taking	$T_{1/2}^{0\nu} > 10^{24} \text{ y}$	$T_{1/2}^{0\nu} > 6 \times 10^{24} \text{ y}$

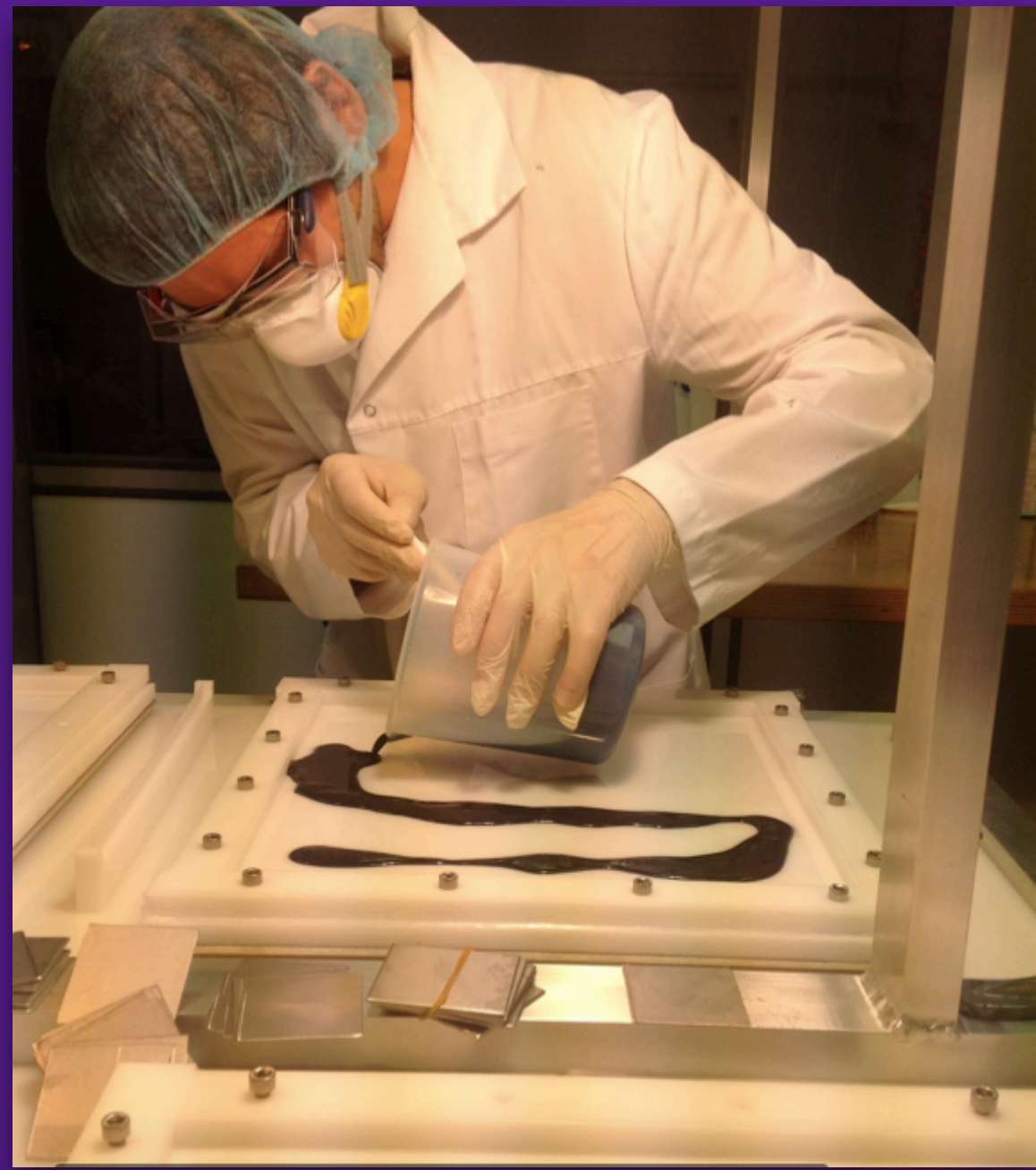
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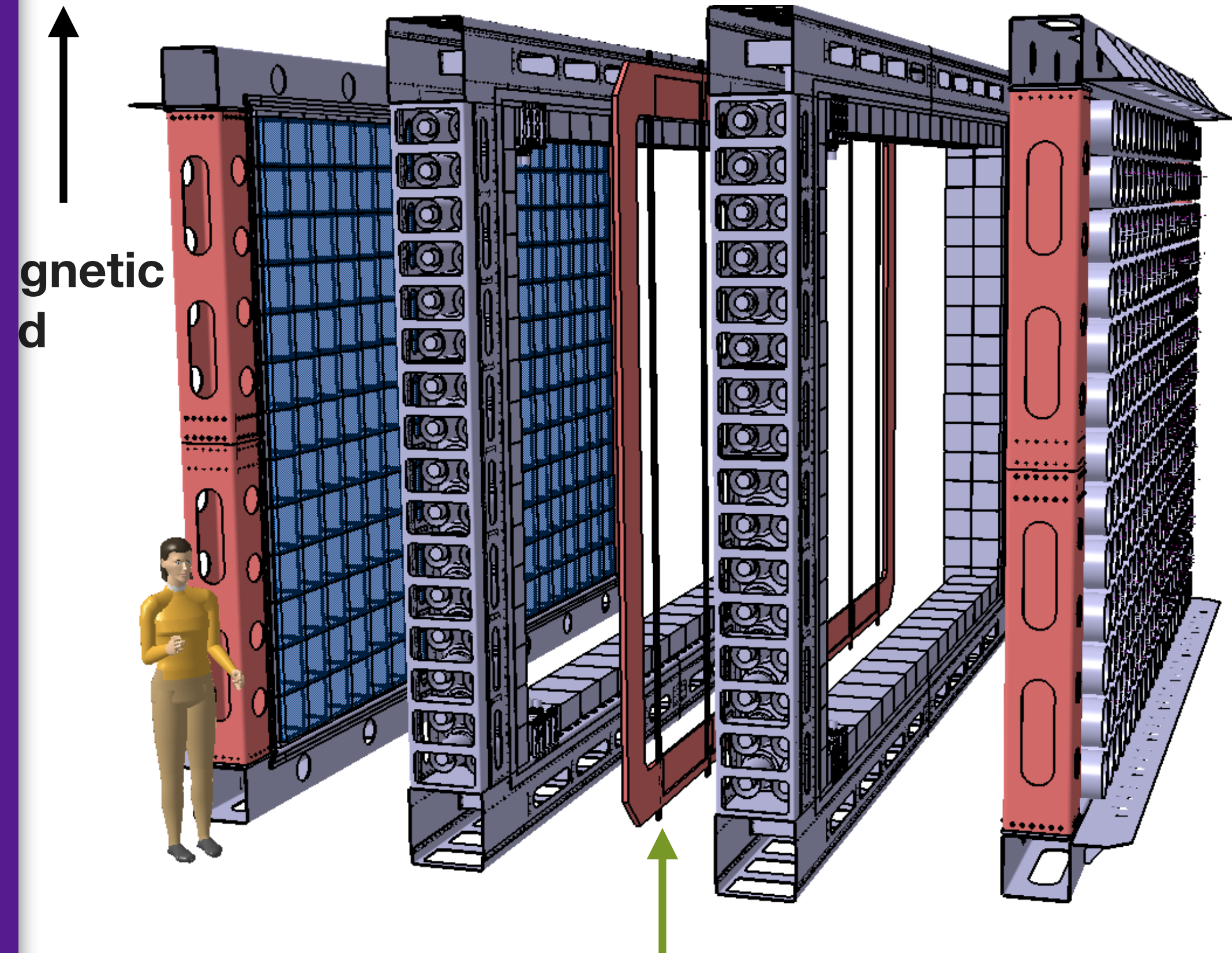
Source frame holding 6.3kg of  $\beta\beta$  emitter ( $^{82}\text{Se}$ )

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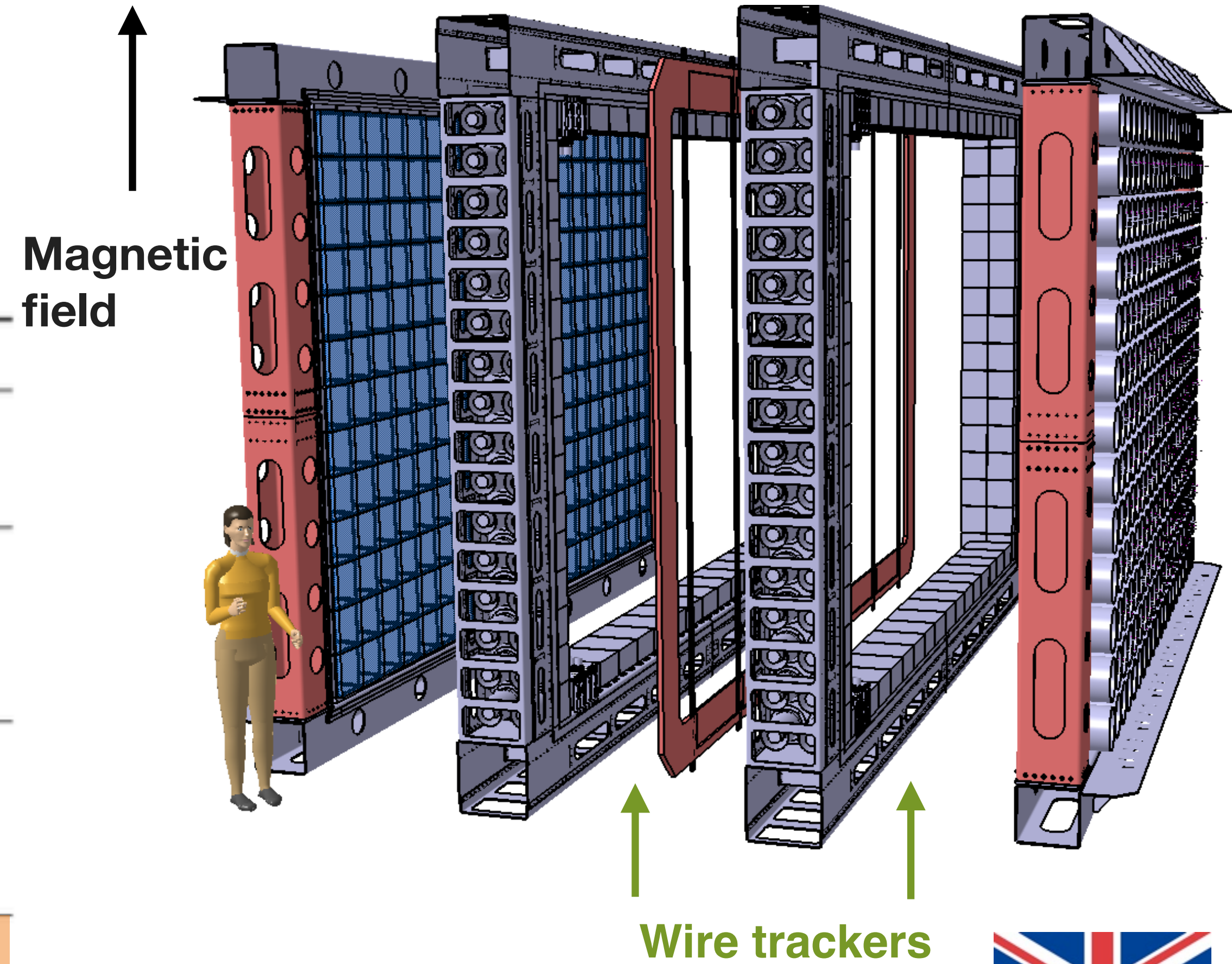
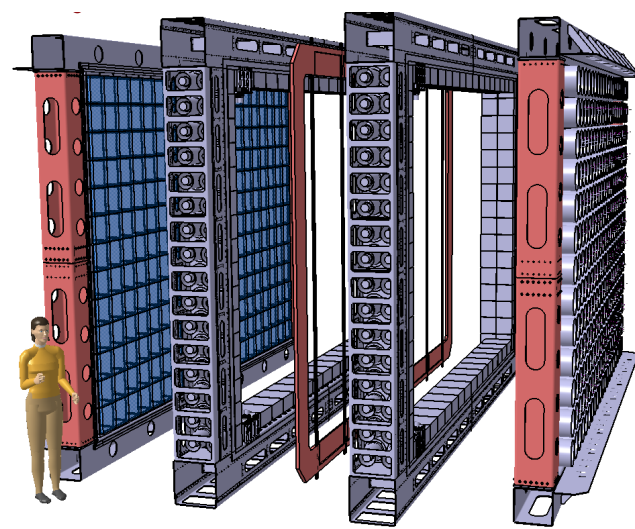
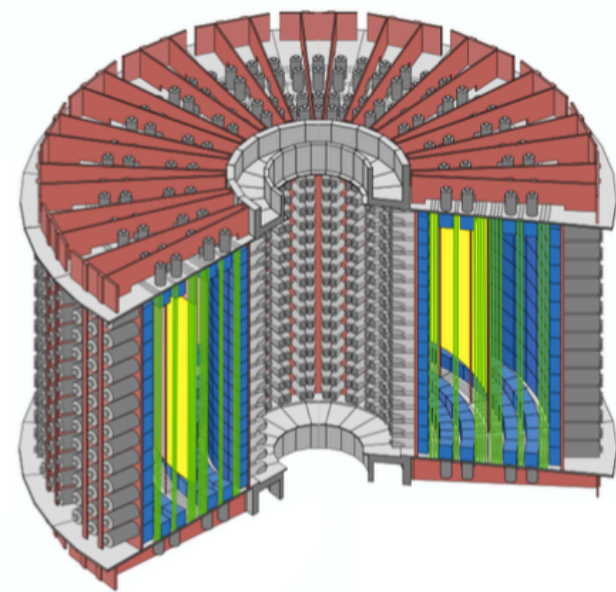
- 34 foils
- Enriched Se powder mixed with PVA
- Increased radio purity through distillation / chromatography / chemical precipitation



Source frame holding 6.3kg of  $\beta\beta$  emitter ( $^{82}\text{Se}$ )

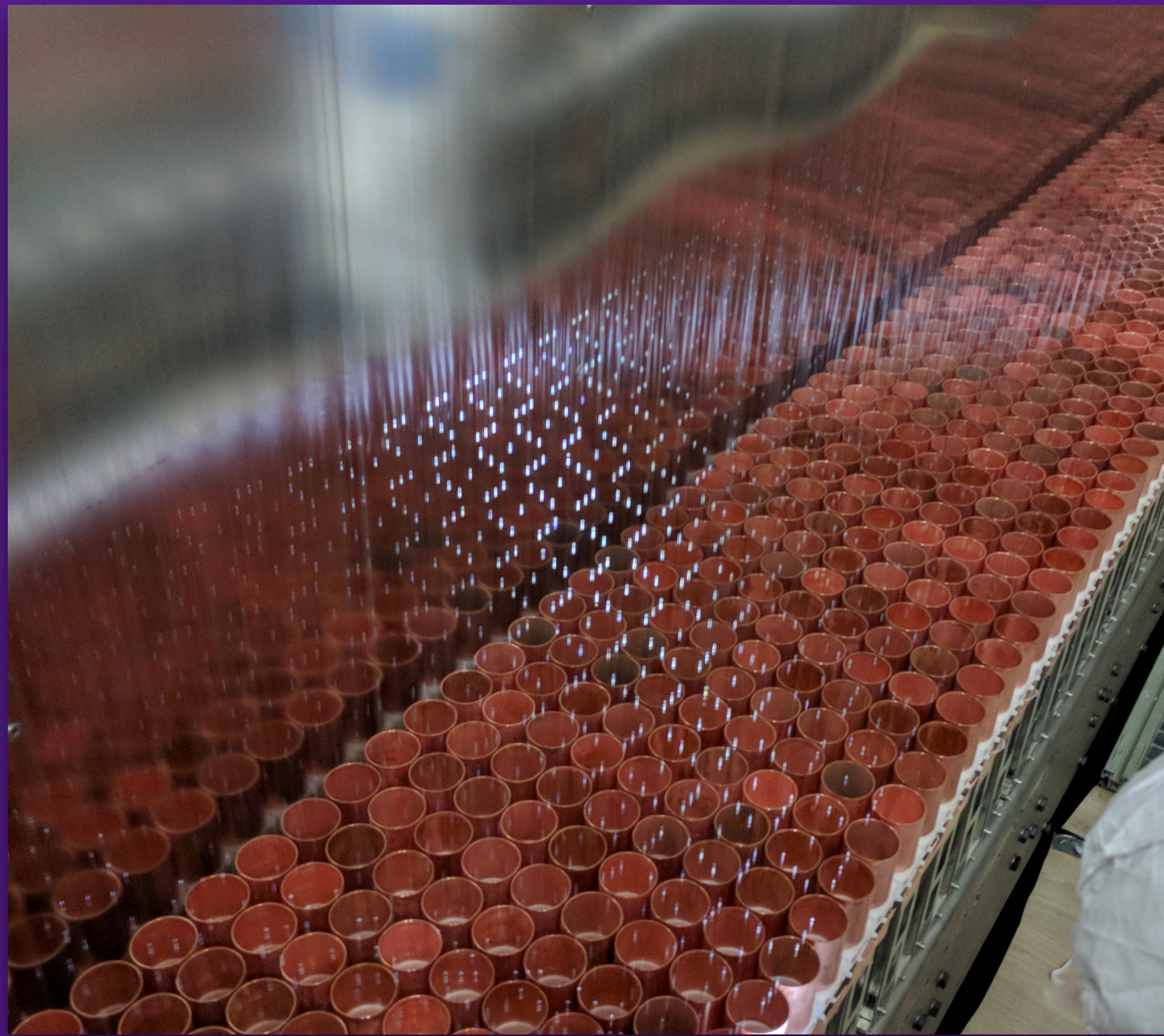


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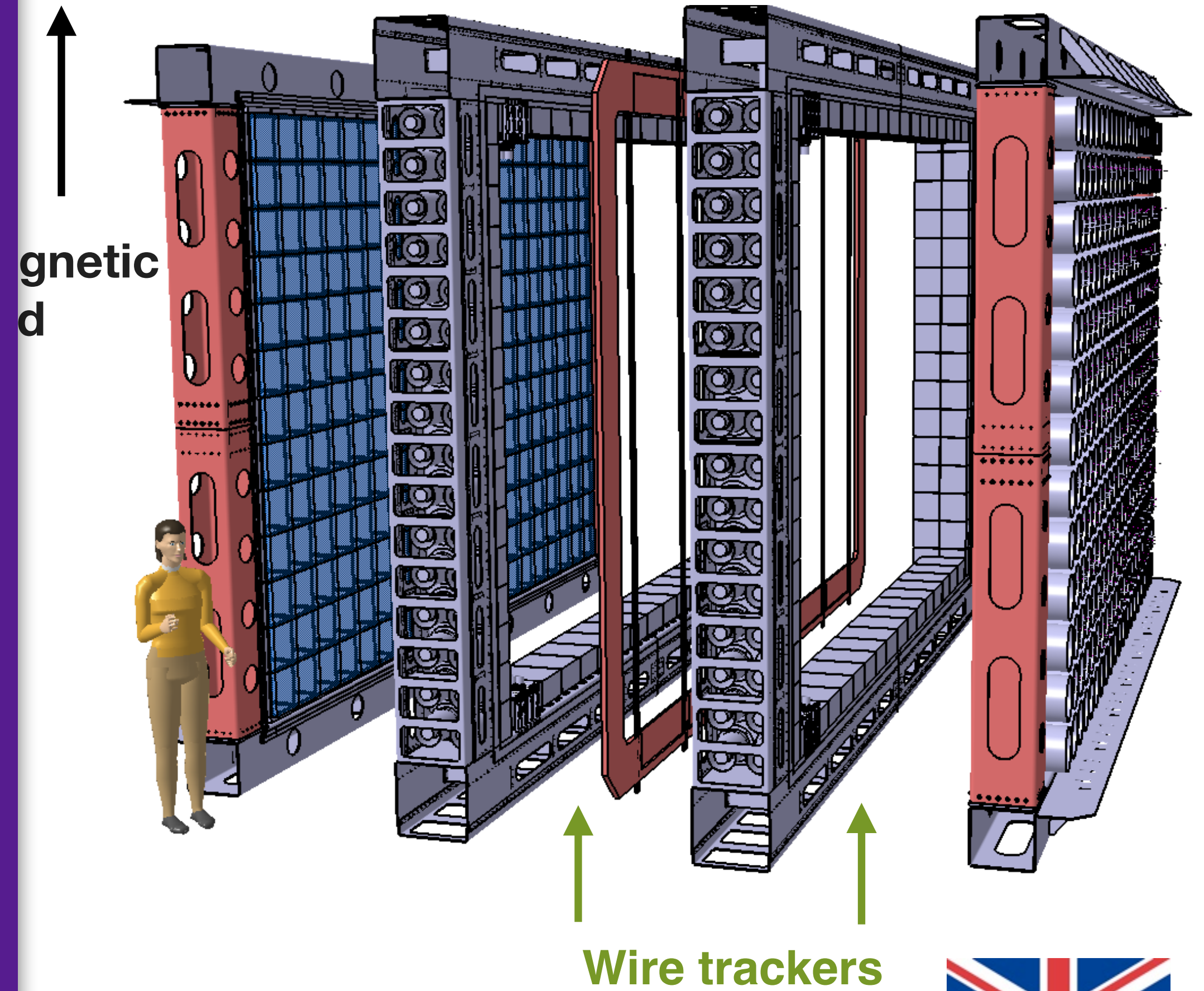


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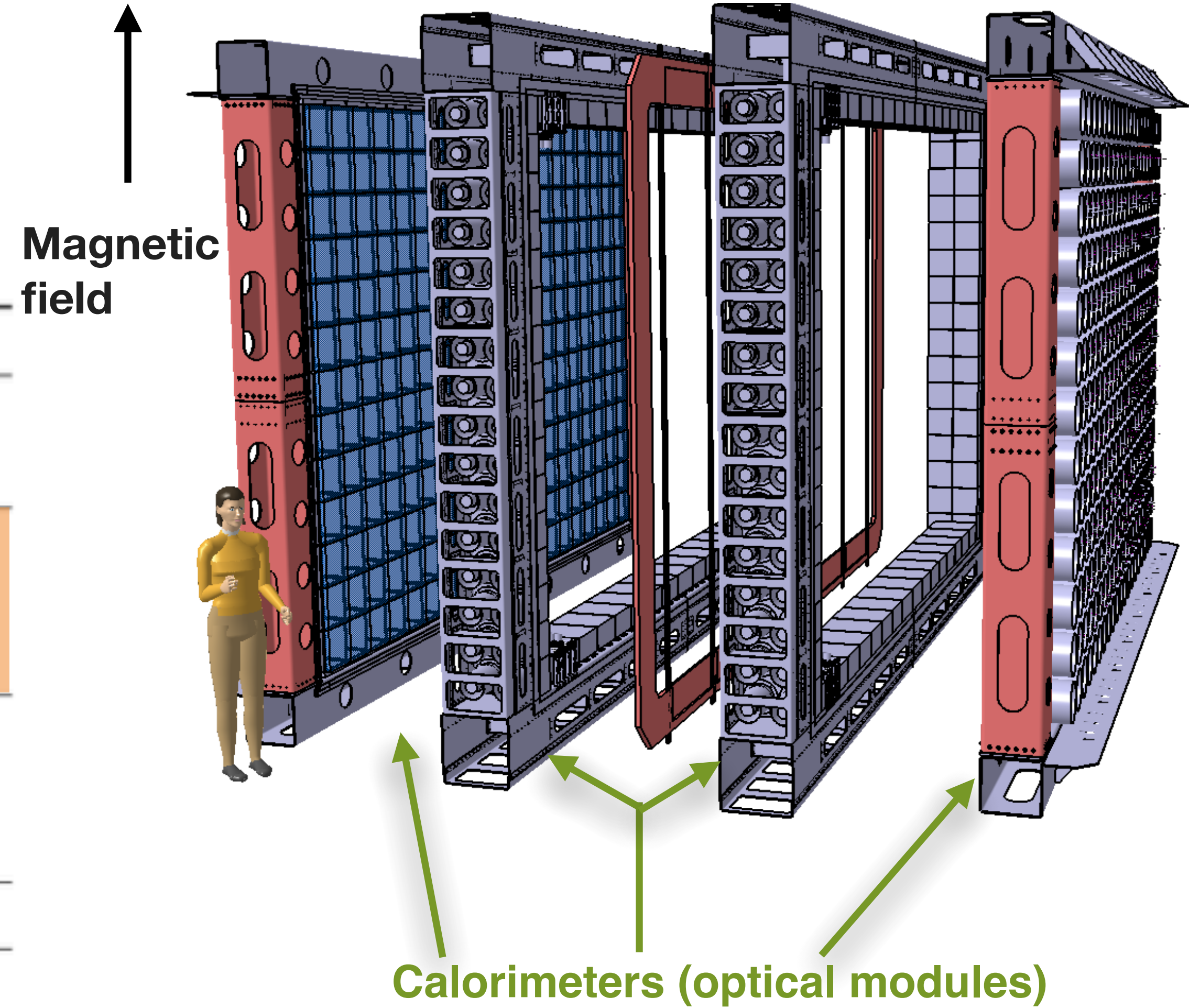
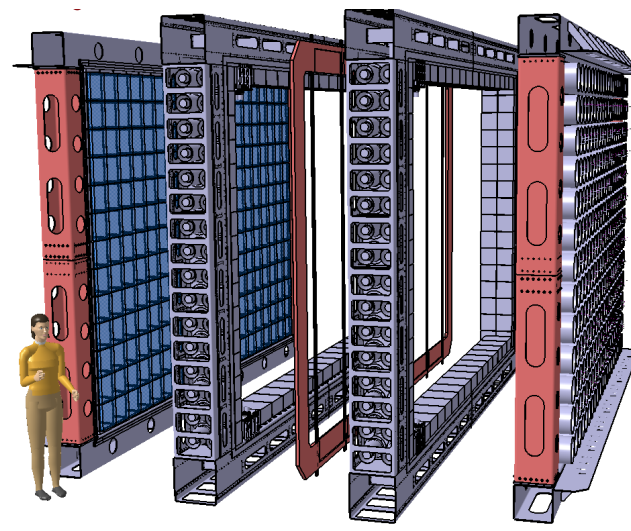
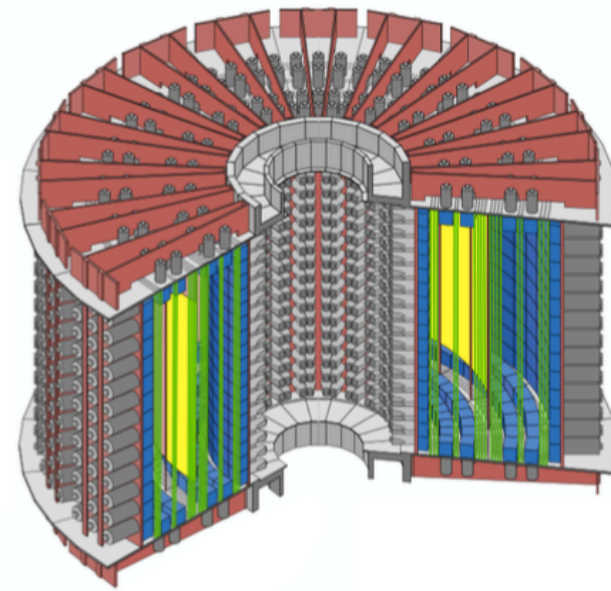
# Currently at LSM: SuperNEMO Demonstrator



- 2034 drift cells (13,000 wires!)
- Built and installed by UK team
- UK radon reduction / measurement programme also used by dark matter experiments

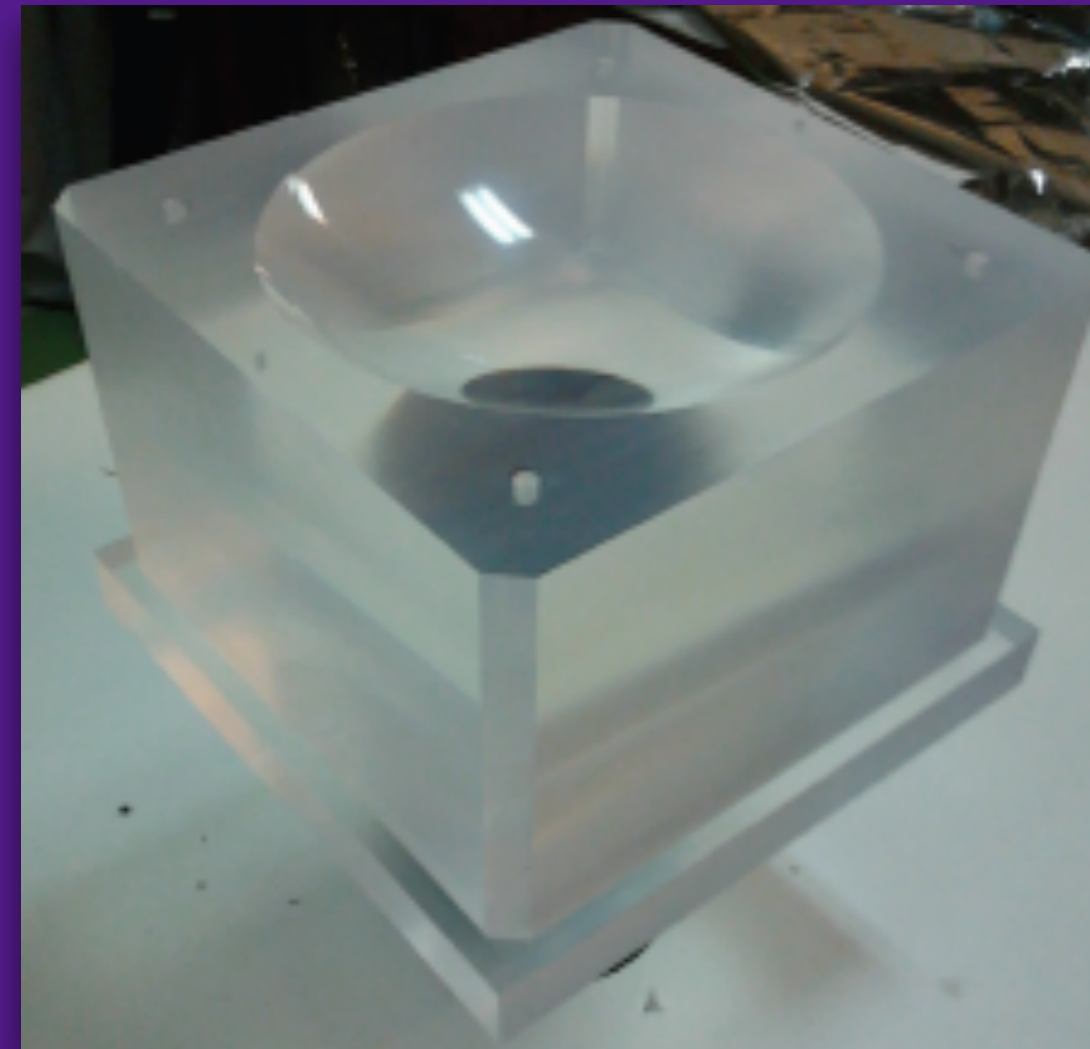
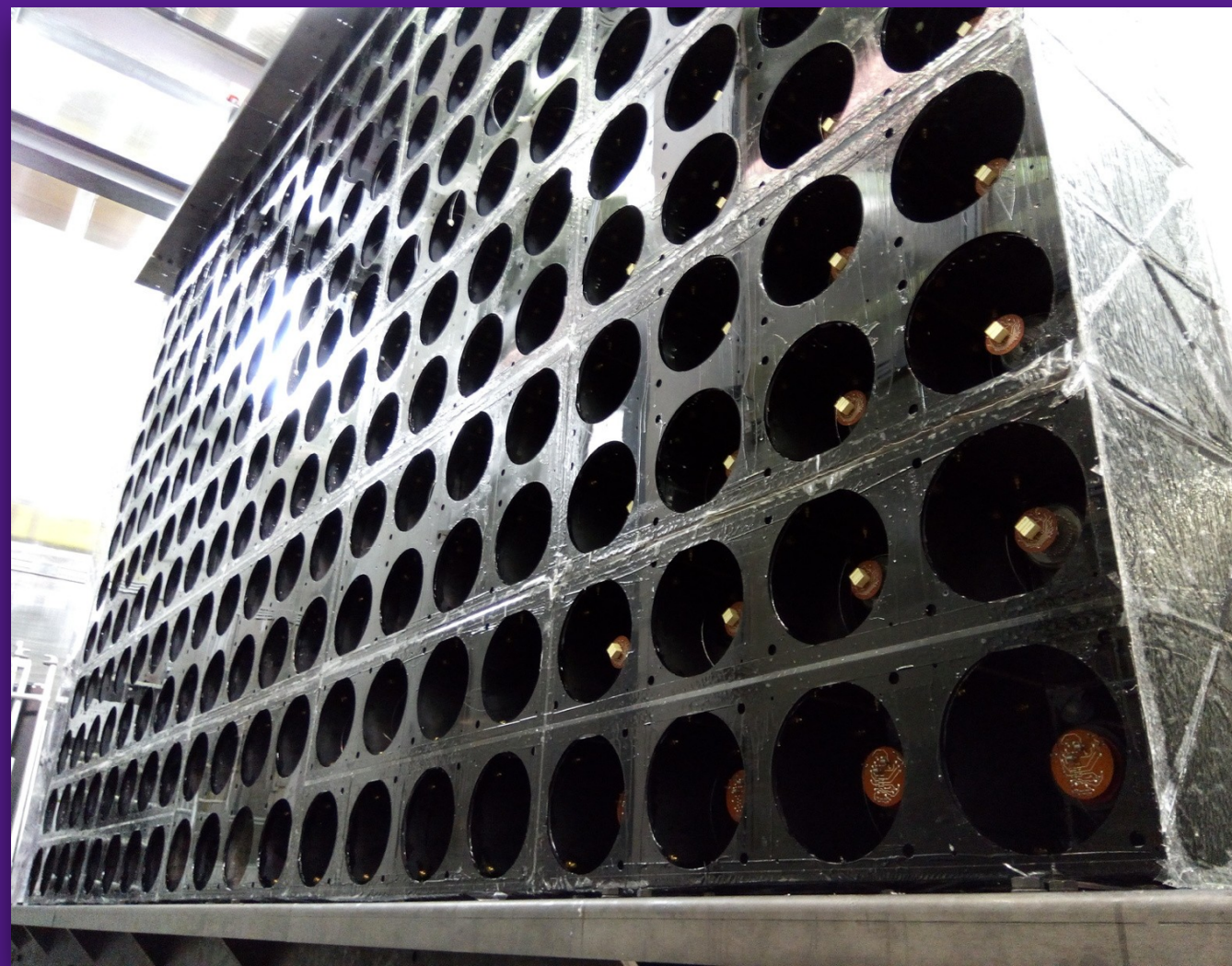


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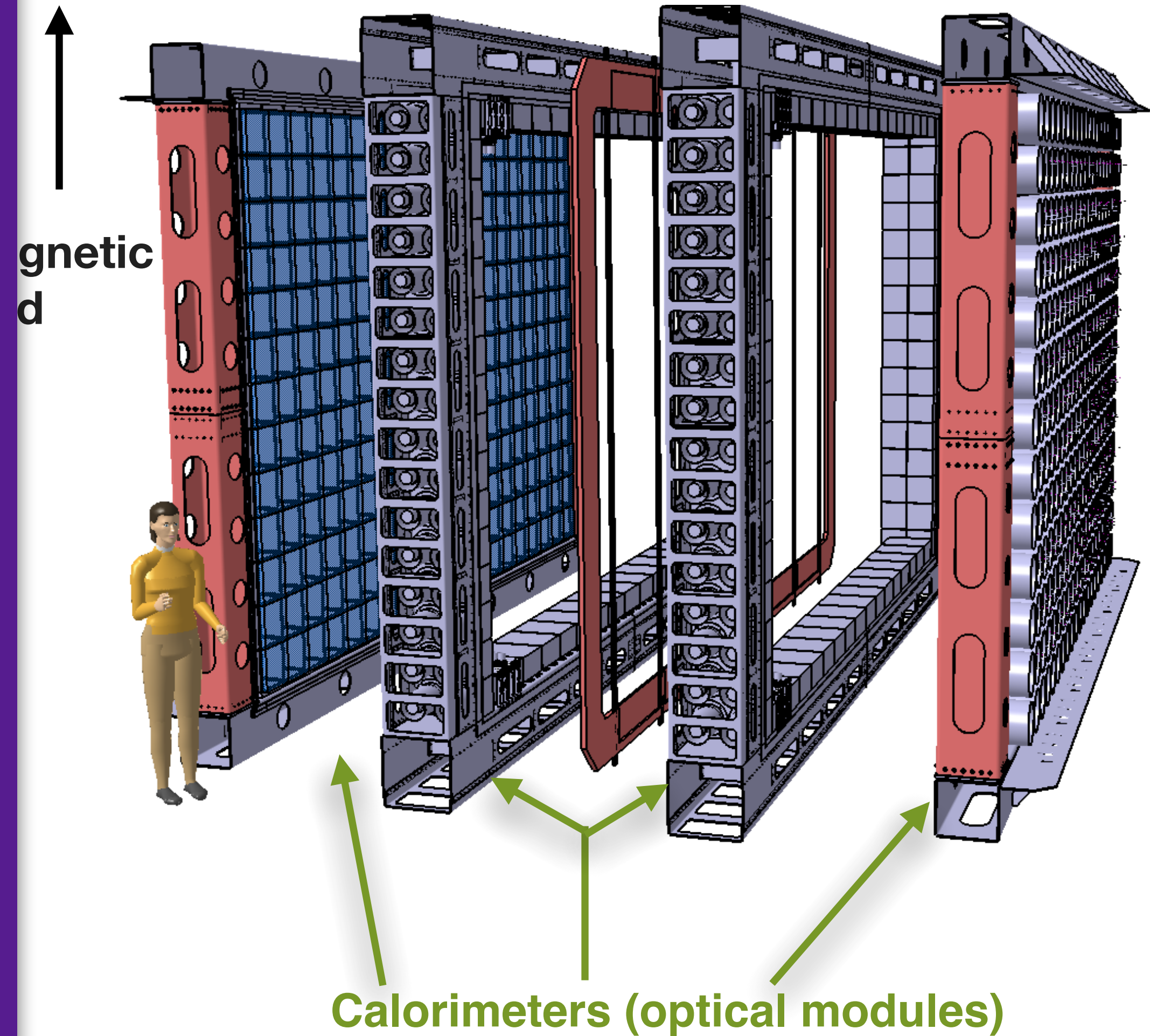
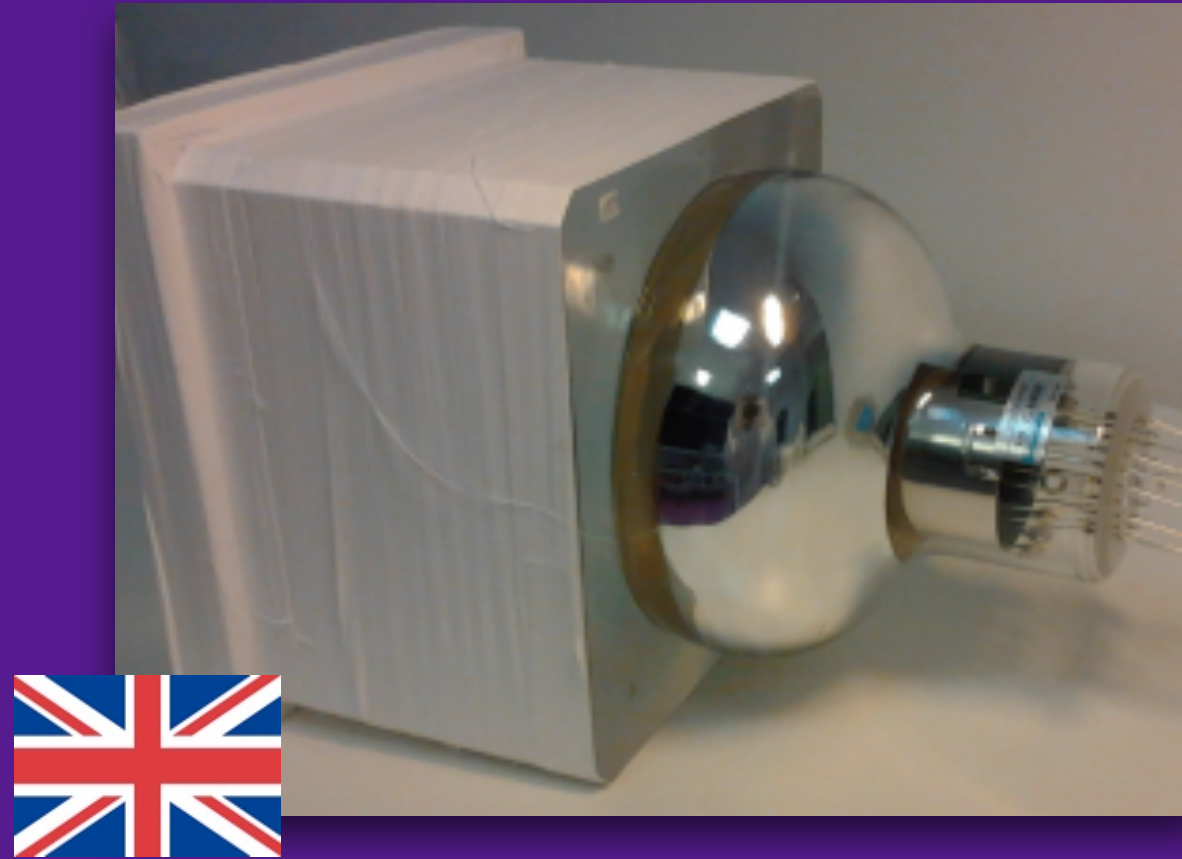


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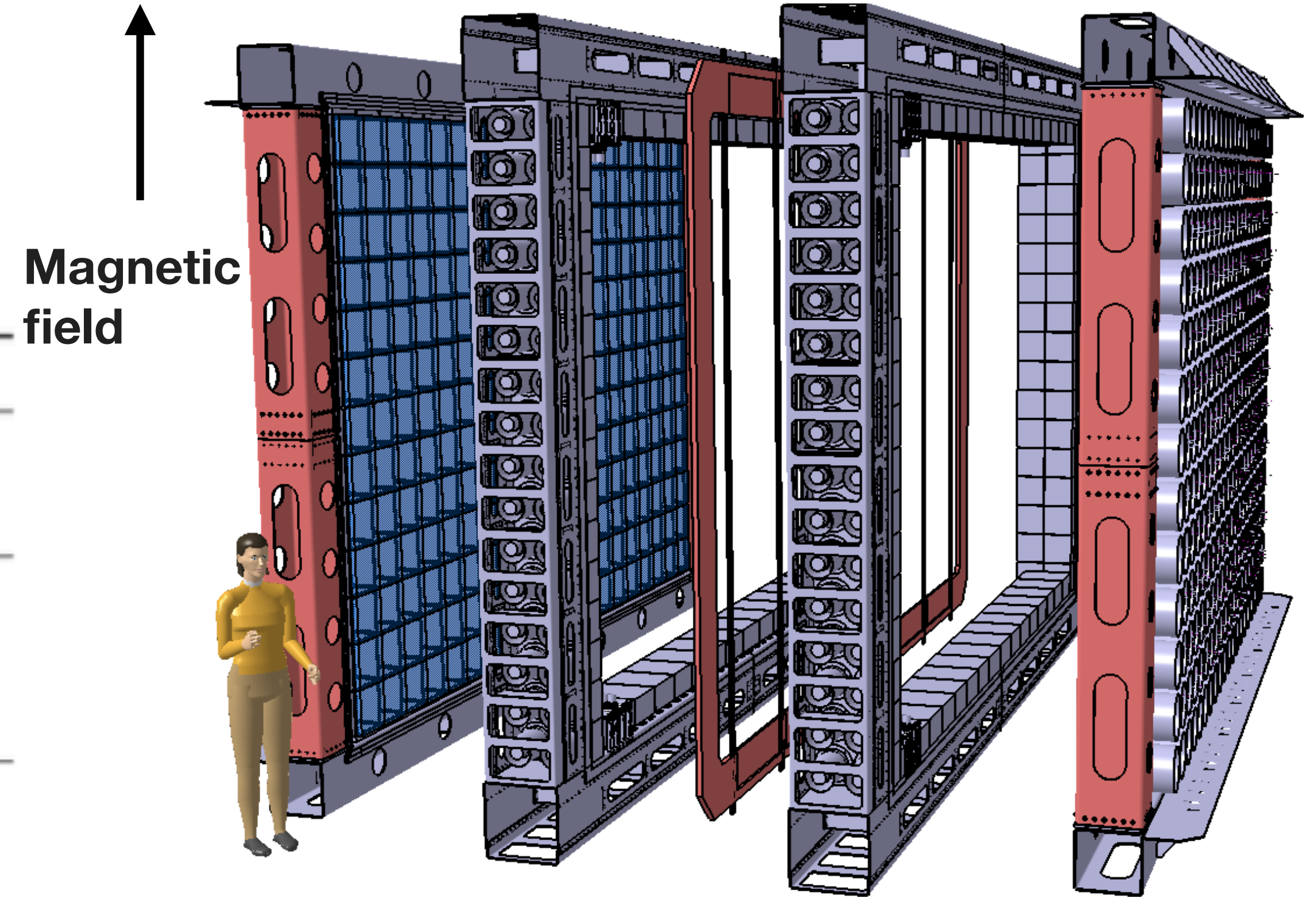
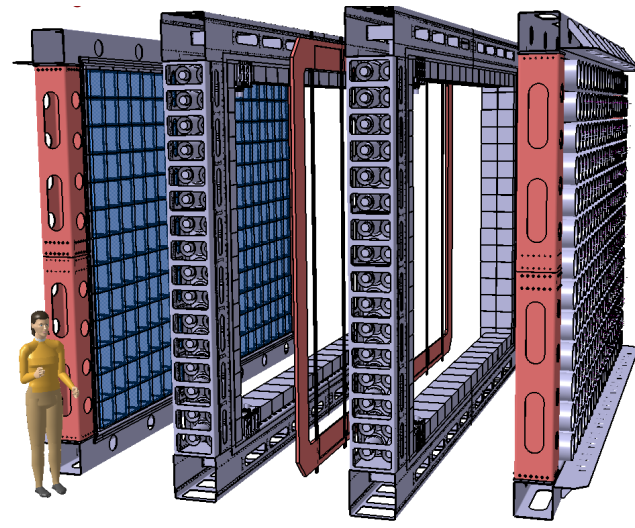
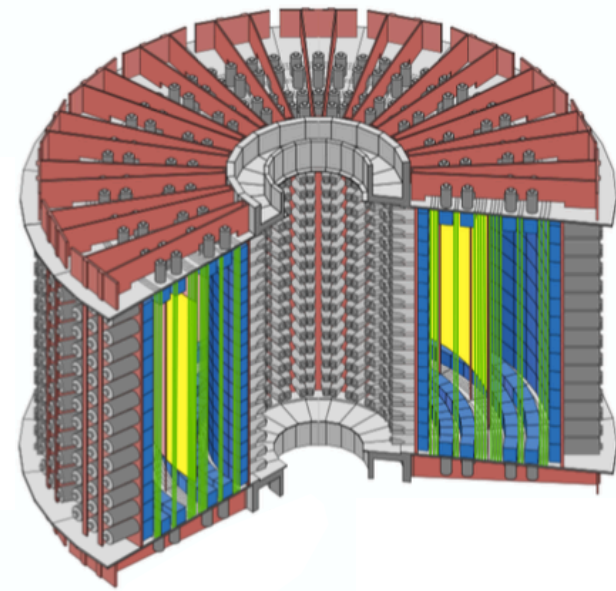
# Currently at LSM: SuperNEMO Demonstrator



- 440 8" radiopure PMTs (plus 5" NEMO-3 PMTs)
- Improved photocathode quantum efficiency
- Directly coupled to polystyrene scintillator (no light guide)
- UK involved in upgraded design (*Nucl.Inst.Meth. A 868 98-108*)



# Currently at LSM: SuperNEMO Demonstrator

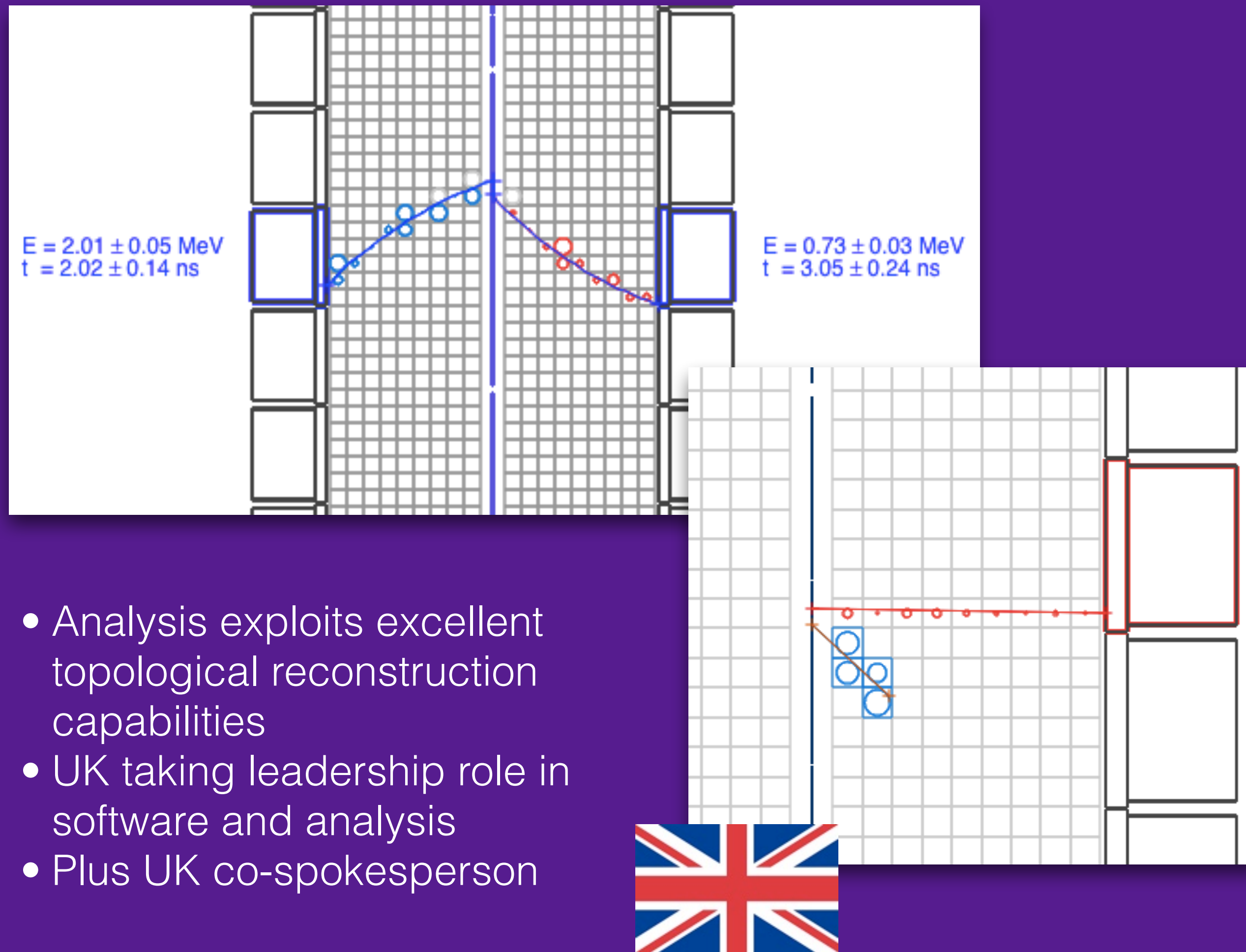


Magnetic field

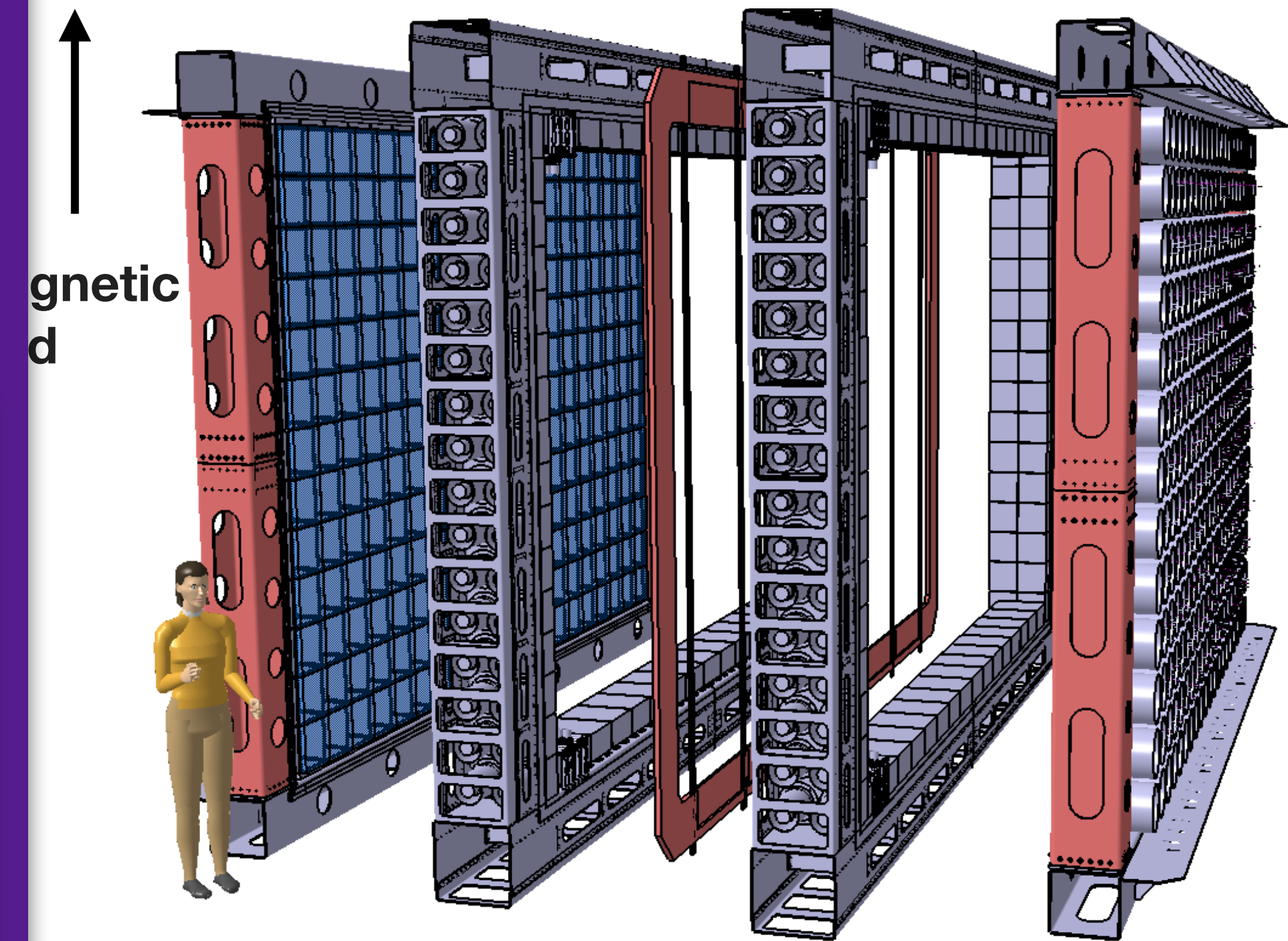
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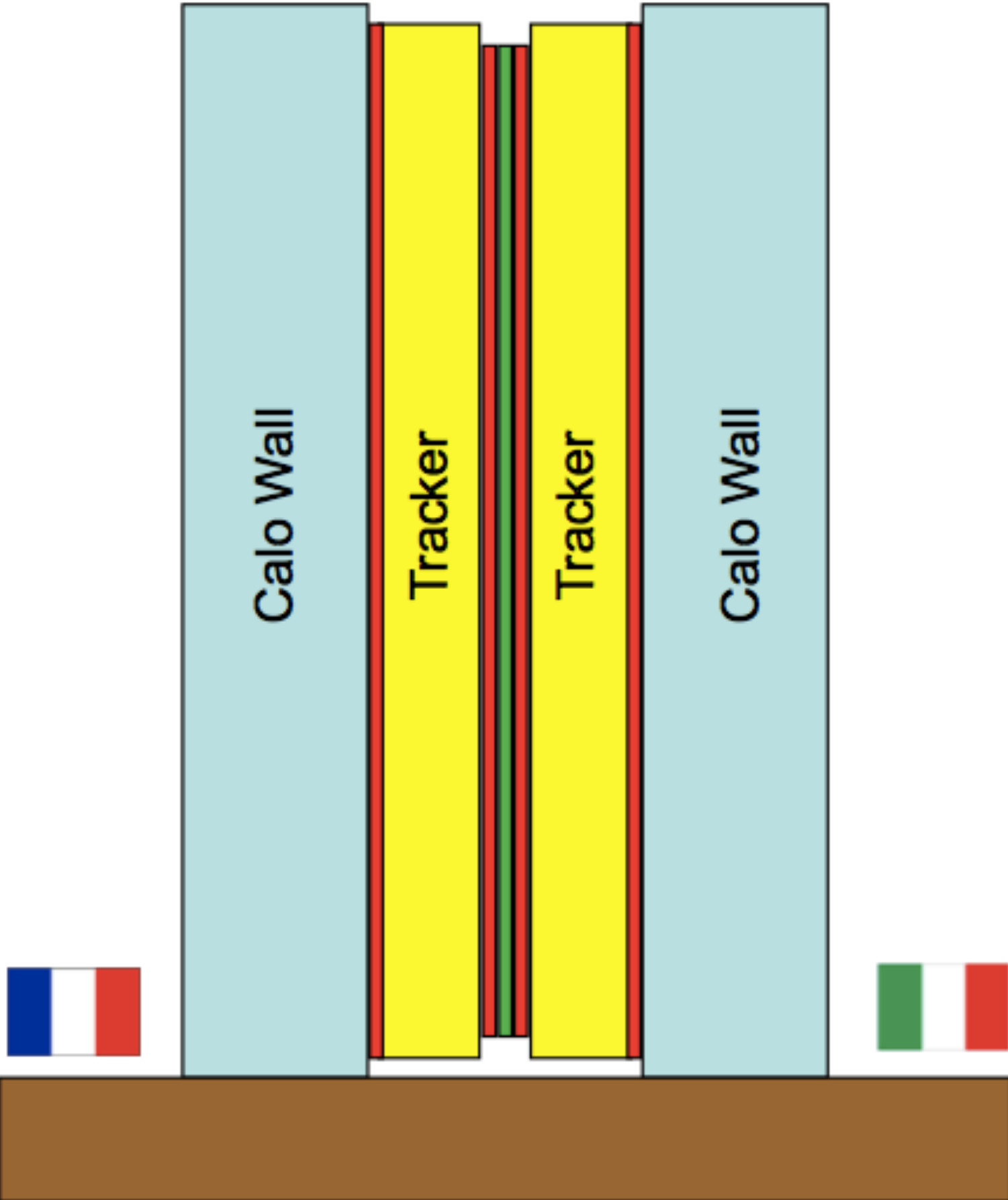


- Analysis exploits excellent topological reconstruction capabilities
- UK taking leadership role in software and analysis
- Plus UK co-spokesperson



$T_{1/2}^{0\nu} > 6 \times 10^{24} \text{ years}$

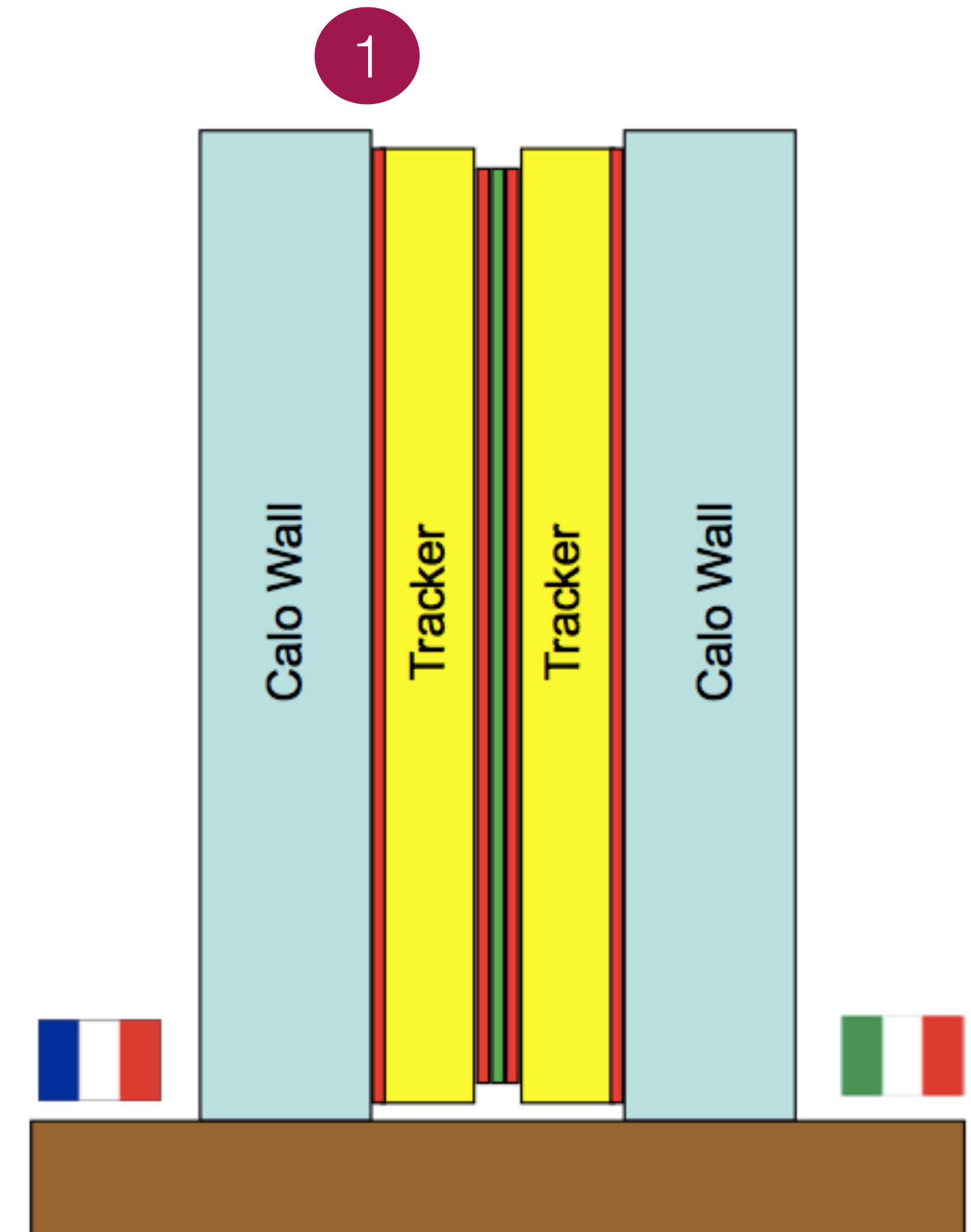
# Installation progress since last year



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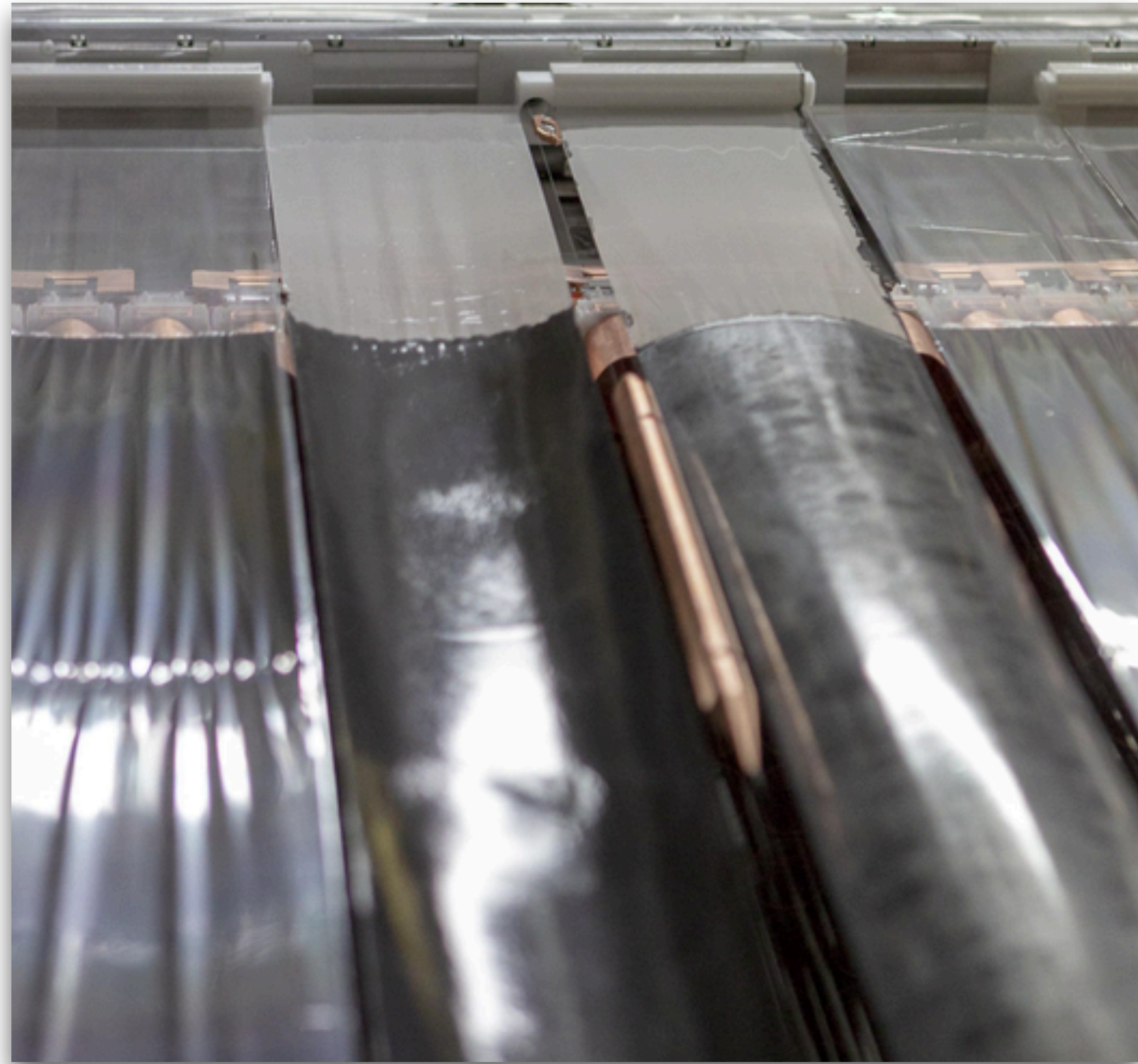


1. Trackers joined to calorimeter wall



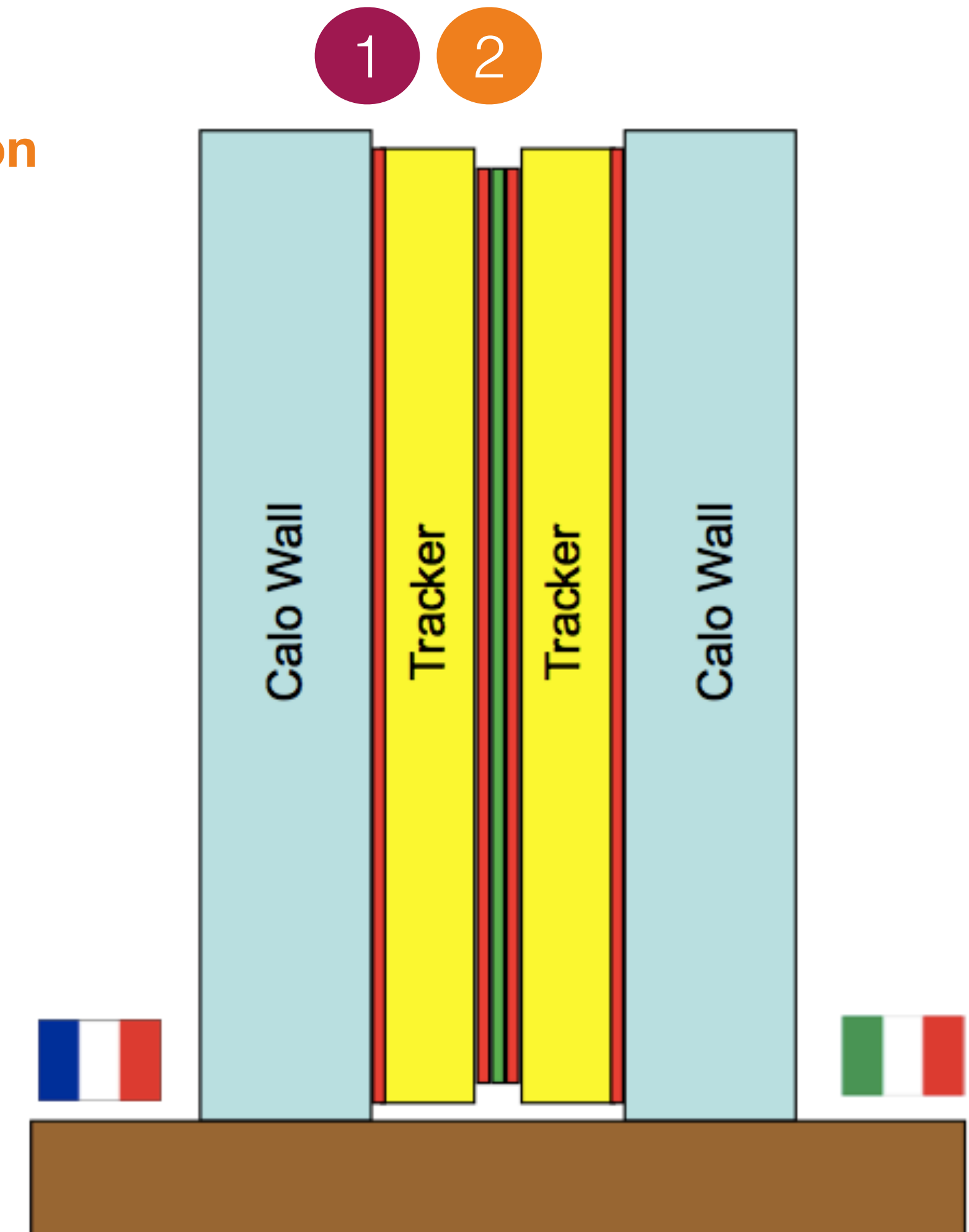


# Installation progress since last year



1. Trackers joined to calorimeter wall

2. Source foils and  $^{207}\text{Bi}$  calibration system installed



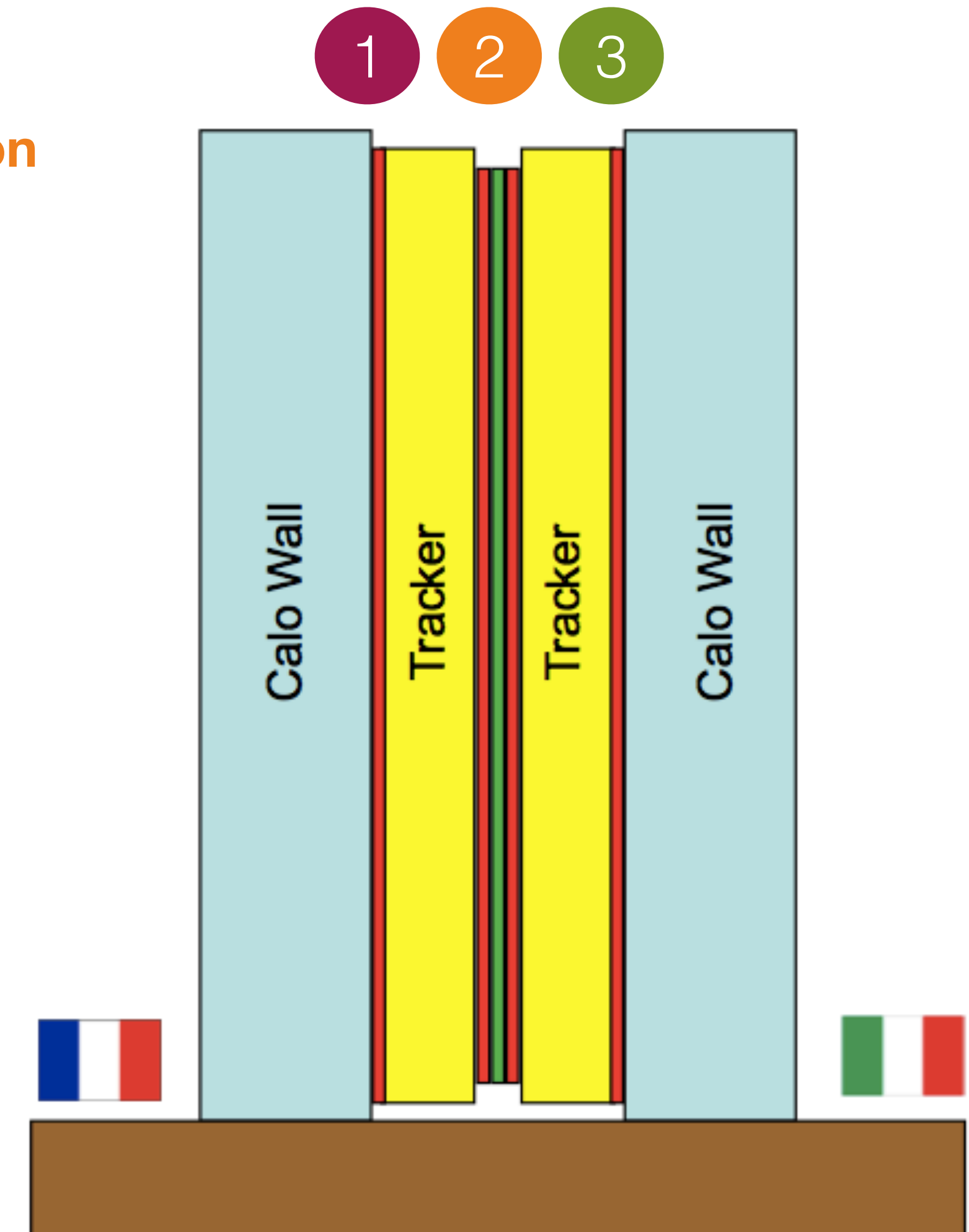
# Installation progress since last year



1. Trackers joined to calorimeter wall

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3. Detector closed



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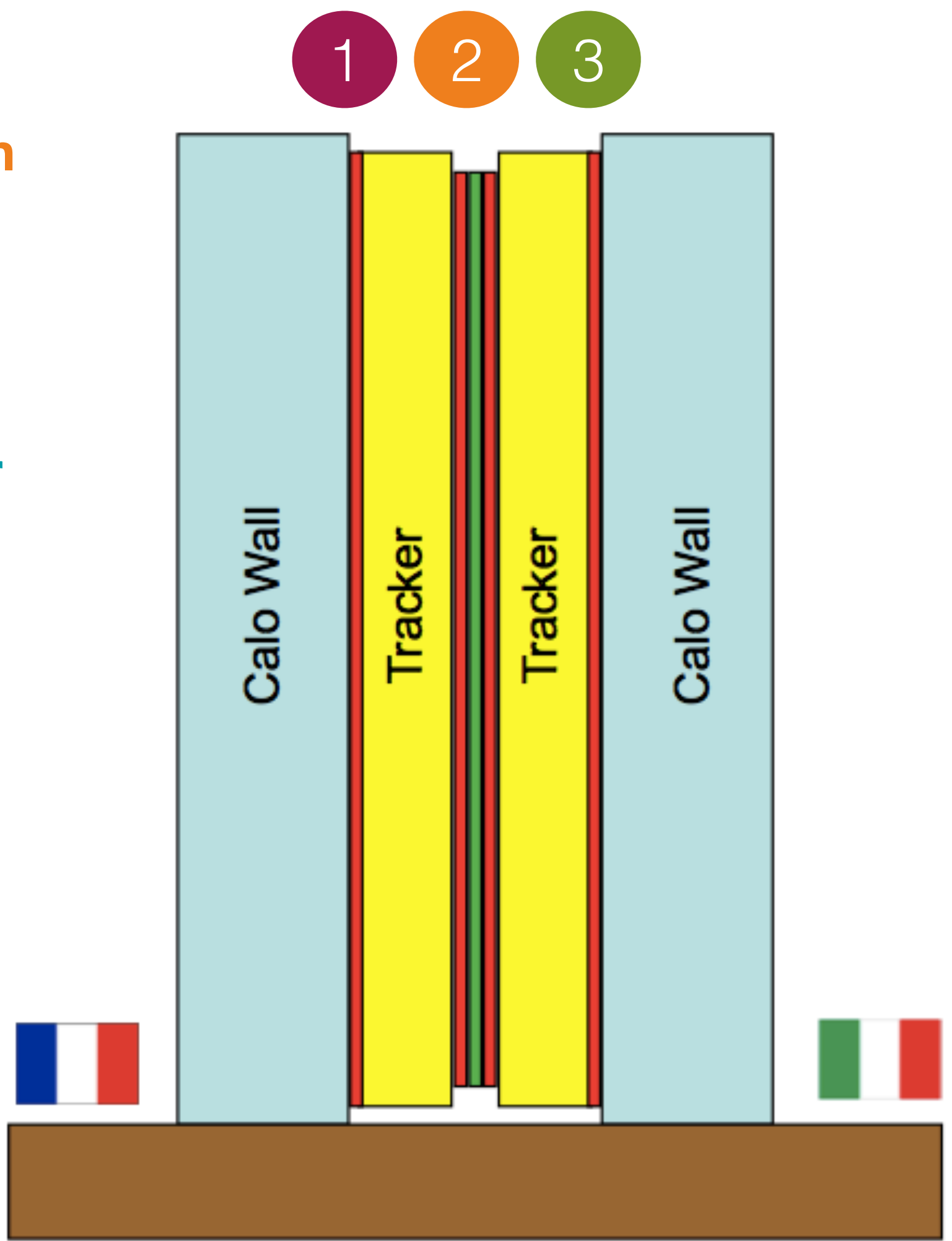
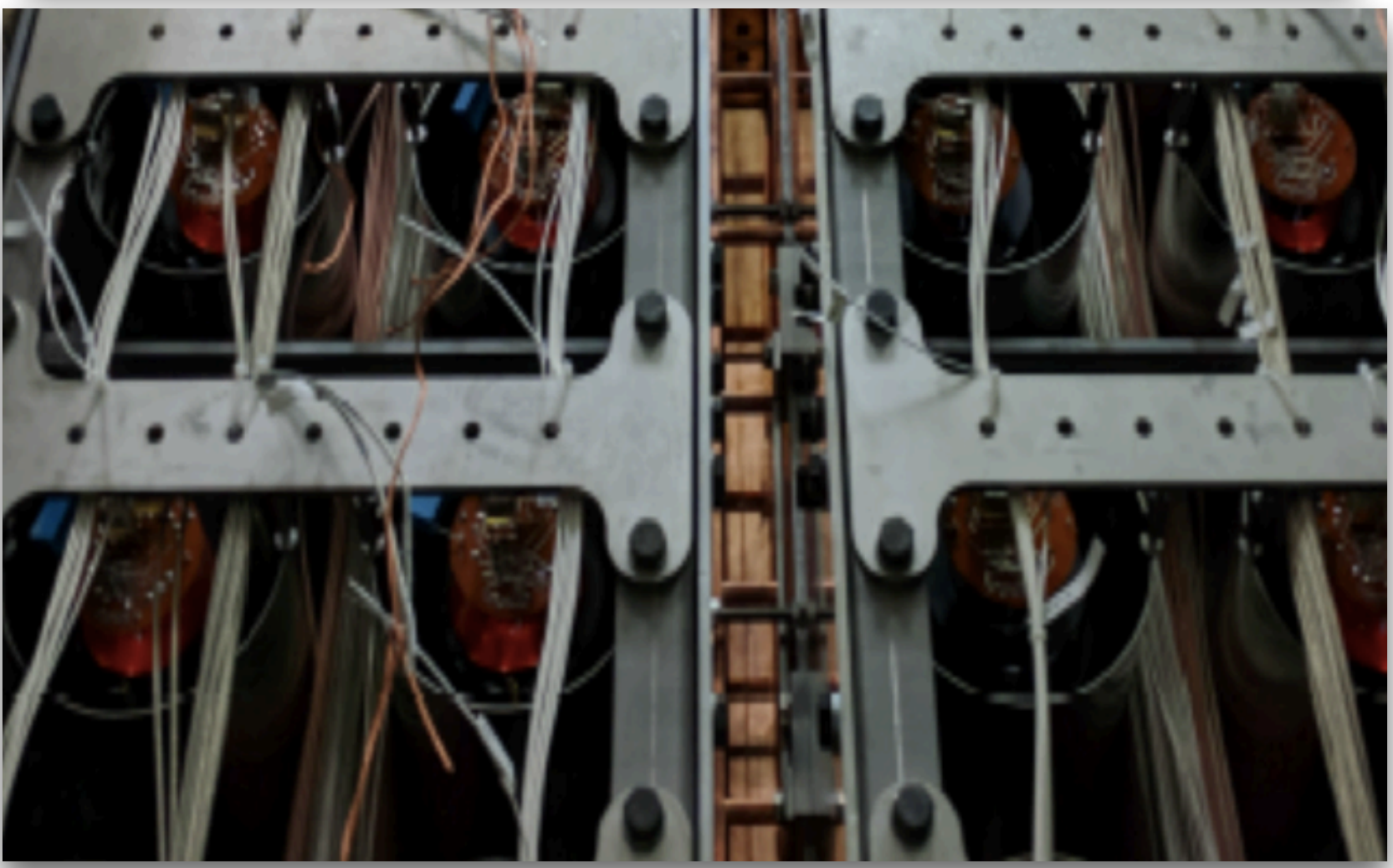


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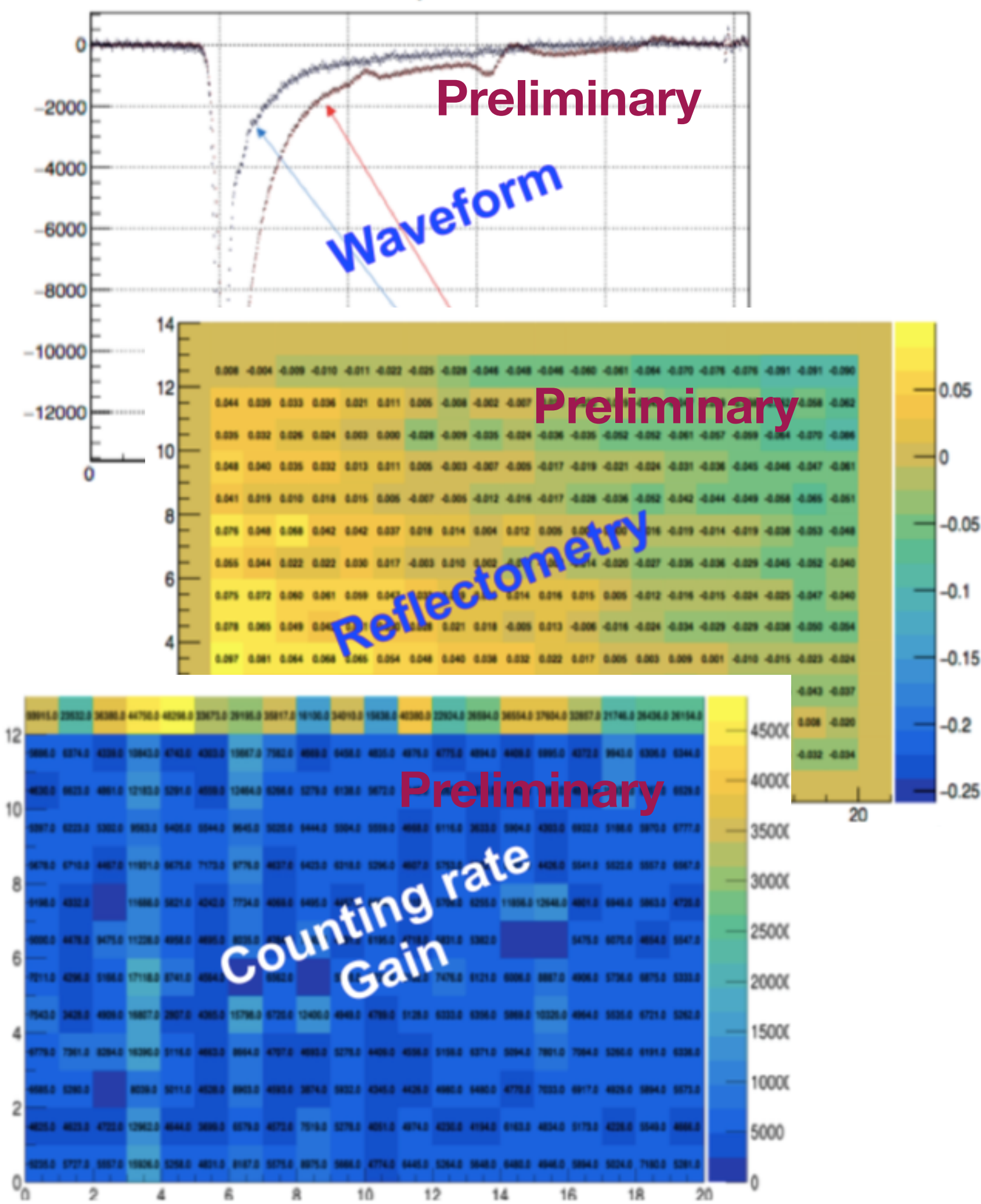
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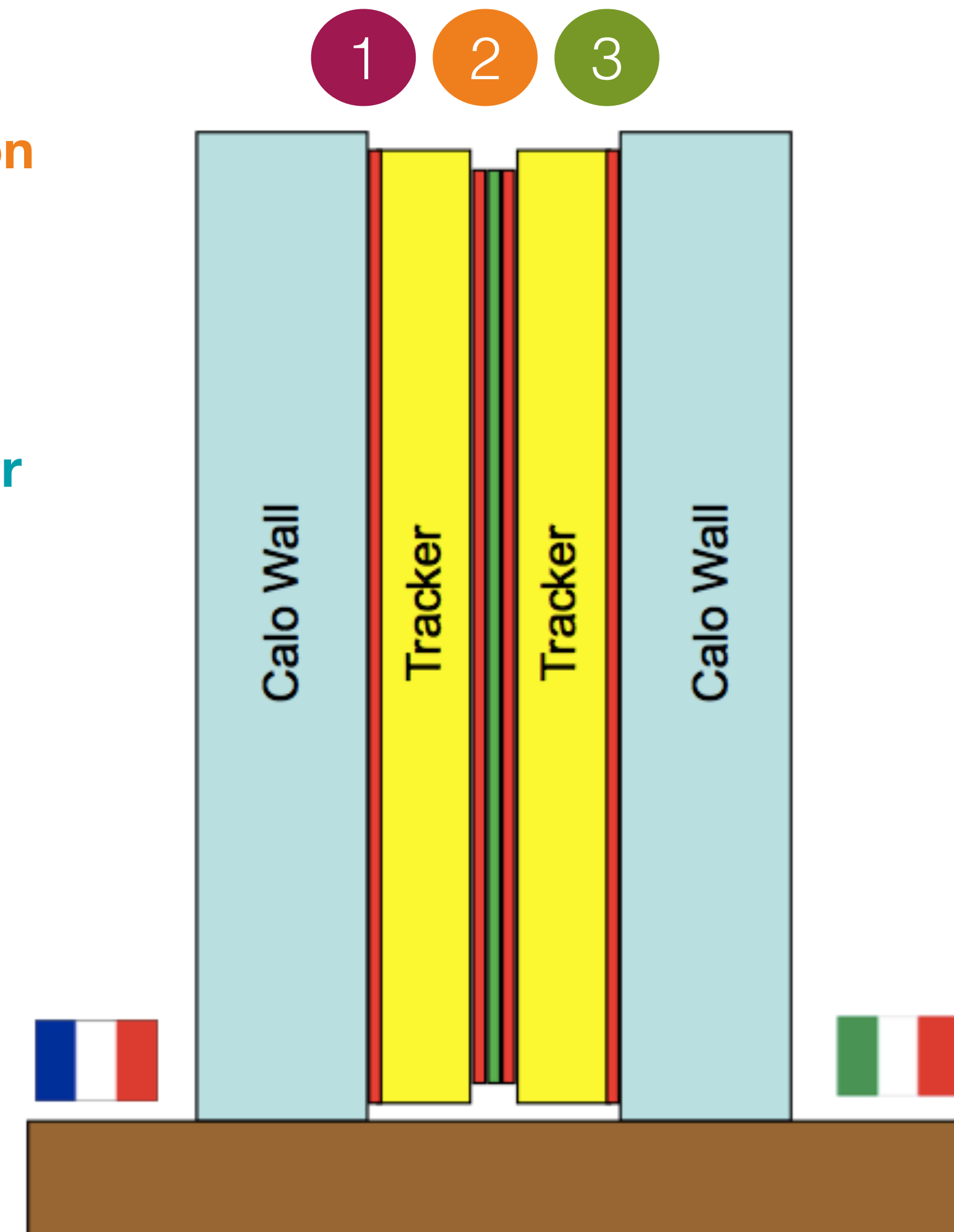
4. Cabling calorimeter and tracker



# Installation progress since last year



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3. Detector closed
4. Cabling calorimeter and tracker
5. First commissioning data!



# Physics goals

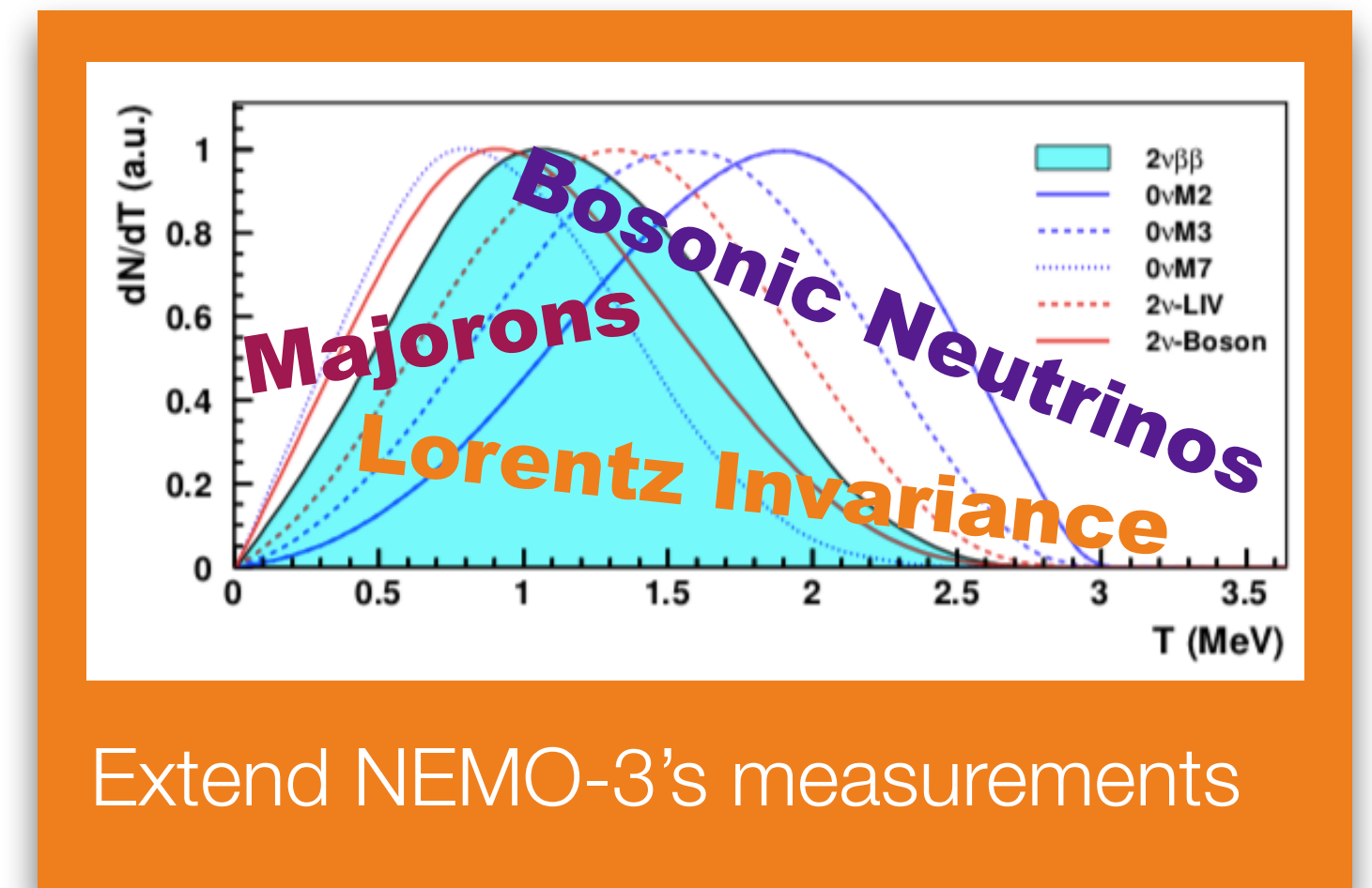
$0\nu\beta\beta$ :  $T_{1/2} > 6 \times 10^{24}$  years;  $\langle m_\nu \rangle < 160\text{-}400$  meV

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Exotic  $0\nu\beta\beta$  mechanisms

Lorentz invariance violation test



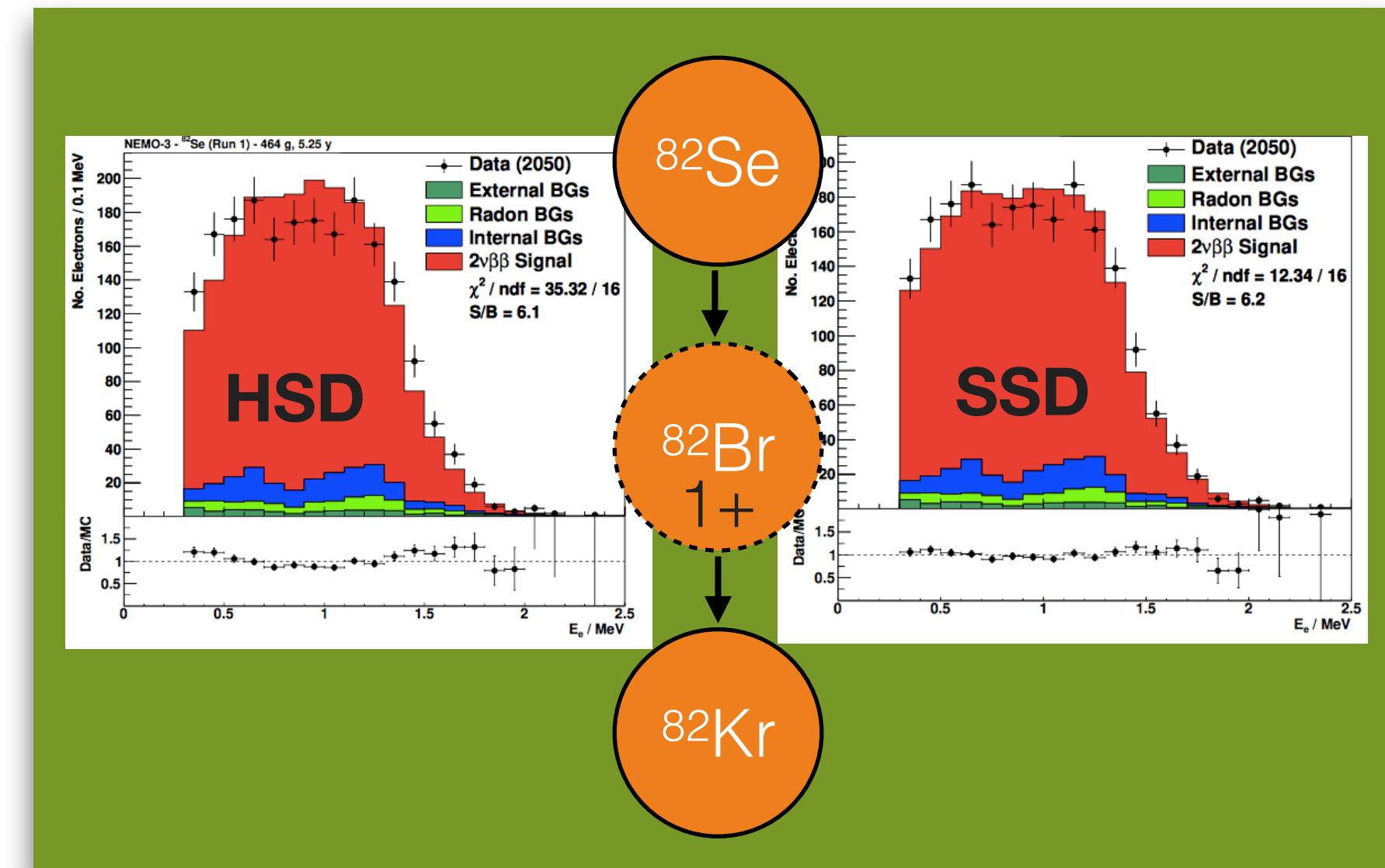
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$2\nu\beta\beta$ : SSD/HSD discrimination at  $5\sigma$  level



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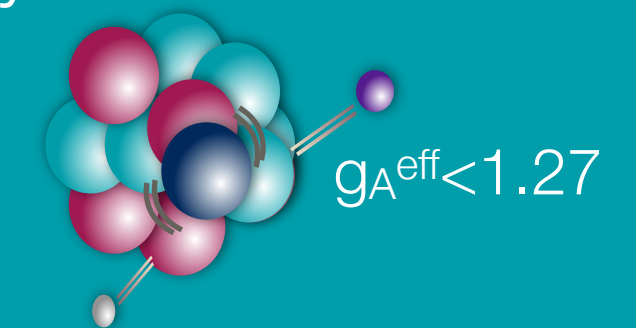
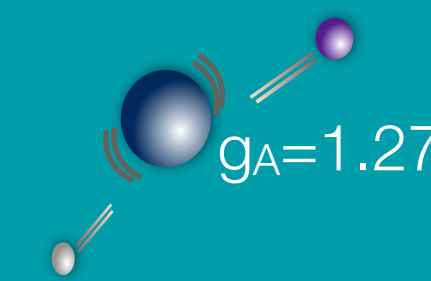
Exotic  $0\nu\beta\beta$  mechanisms

Lorentz invariance violation test

$2\nu\beta\beta$ : SSD/HSD discrimination at  $5\sigma$  level

Probe nuclear physics by investigating  $g_A$

- Axial-vector coupling constant  $g_A$  is **quenched** in heavy nuclei



- $2\nu\beta\beta$  rate proportional to  $g_A^4$   
$$\left(T_{1/2}^{2\nu}\right)^{-1} = \left(g_A^{\text{eff}}\right)^4 \left|M_{GT}^{2\nu}\right|^2 G^{2\nu}$$
- New KamLAND-Zen paper investigates this quenching <https://arxiv.org/pdf/1901.03871.pdf>

Precision measurement of the  $^{136}\text{Xe}$  two-neutrino  $\beta\beta$  spectrum in KamLAND-Zen and its impact on the quenching of nuclear matrix elements

Gando,<sup>1</sup> Y. Gando,<sup>1</sup> T. Hachiya,<sup>1</sup> M. Ha Minh,<sup>1</sup> S. Hayashida,<sup>1</sup> Y. Honda,<sup>1</sup> K. Hosokawa,<sup>1</sup> H. Ikeda,<sup>1</sup> K. Inoue,<sup>1</sup> K. Ishidoshiro,<sup>1</sup> Y. Kamei,<sup>1</sup> K. Kamizawa,<sup>1</sup> T. Kinoshita,<sup>1</sup> M. Koga,<sup>1,2</sup> S. Matsuda,<sup>1</sup> T. Mitsui,<sup>1</sup> K. Nakamura,<sup>1,2</sup> Ono,<sup>1</sup> N. Ota,<sup>1</sup> S. Otsuka,<sup>1</sup> H. Ozaki,<sup>1</sup> Y. Shibukawa,<sup>1</sup> I. Shimizu,<sup>1</sup> Y. Shirahata,<sup>1</sup> J. Shirai,<sup>1</sup> T. Sato,<sup>1</sup> K. Soma,

- NEMO's topological capabilities mean it could do even **better!**



# Physics goals

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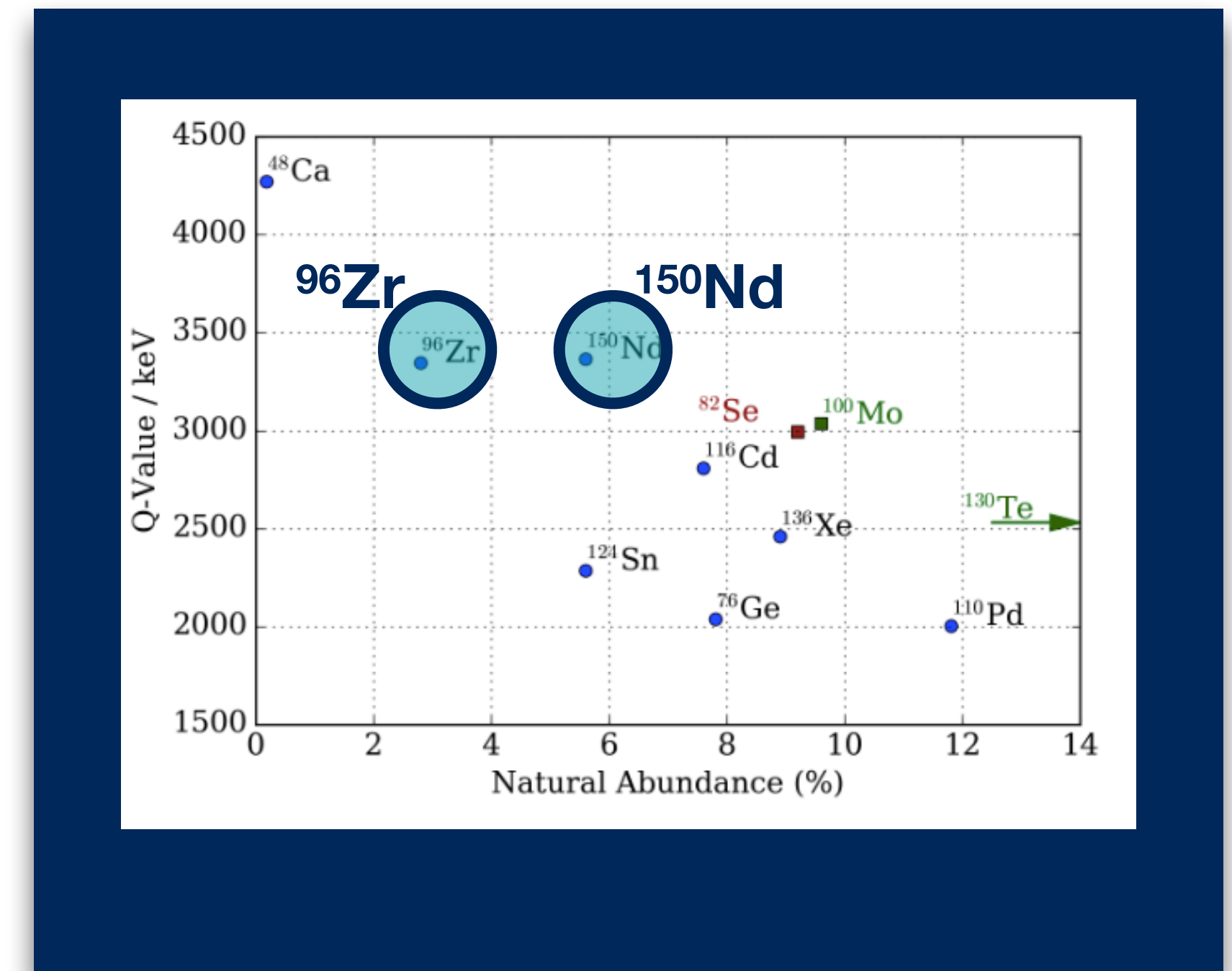
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Alternative isotopes:  $^{150}\text{Nd}$  and  $^{96}\text{Zr}$



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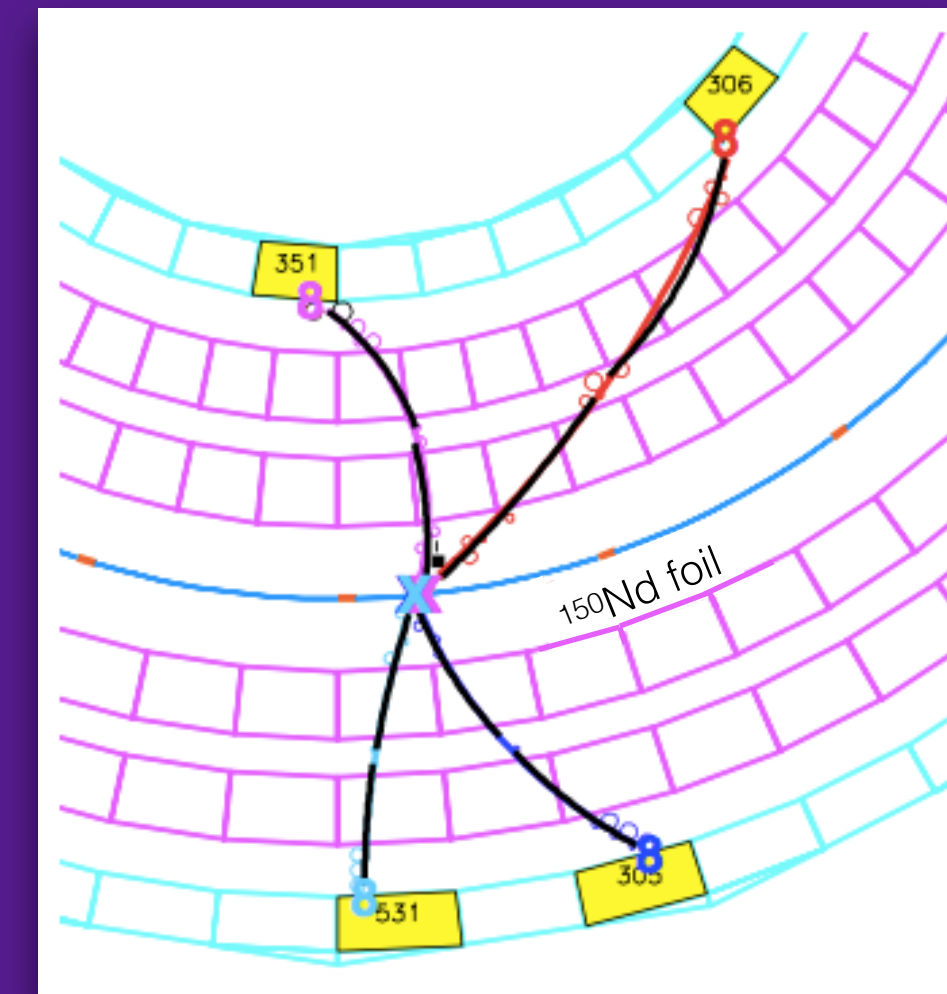
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Probe nuclear physics by investigating  $g_A$

Alternative isotopes:  $^{150}\text{Nd}$  and  $^{96}\text{Zr}$

$0\nu 4\beta$ : for  $^{150}\text{Nd}$



NEMO-3 placed limit on lepton number-violating process, which could affect even Dirac neutrinos  
Phys. Rev. Lett. 119, 041801

# Physics goals

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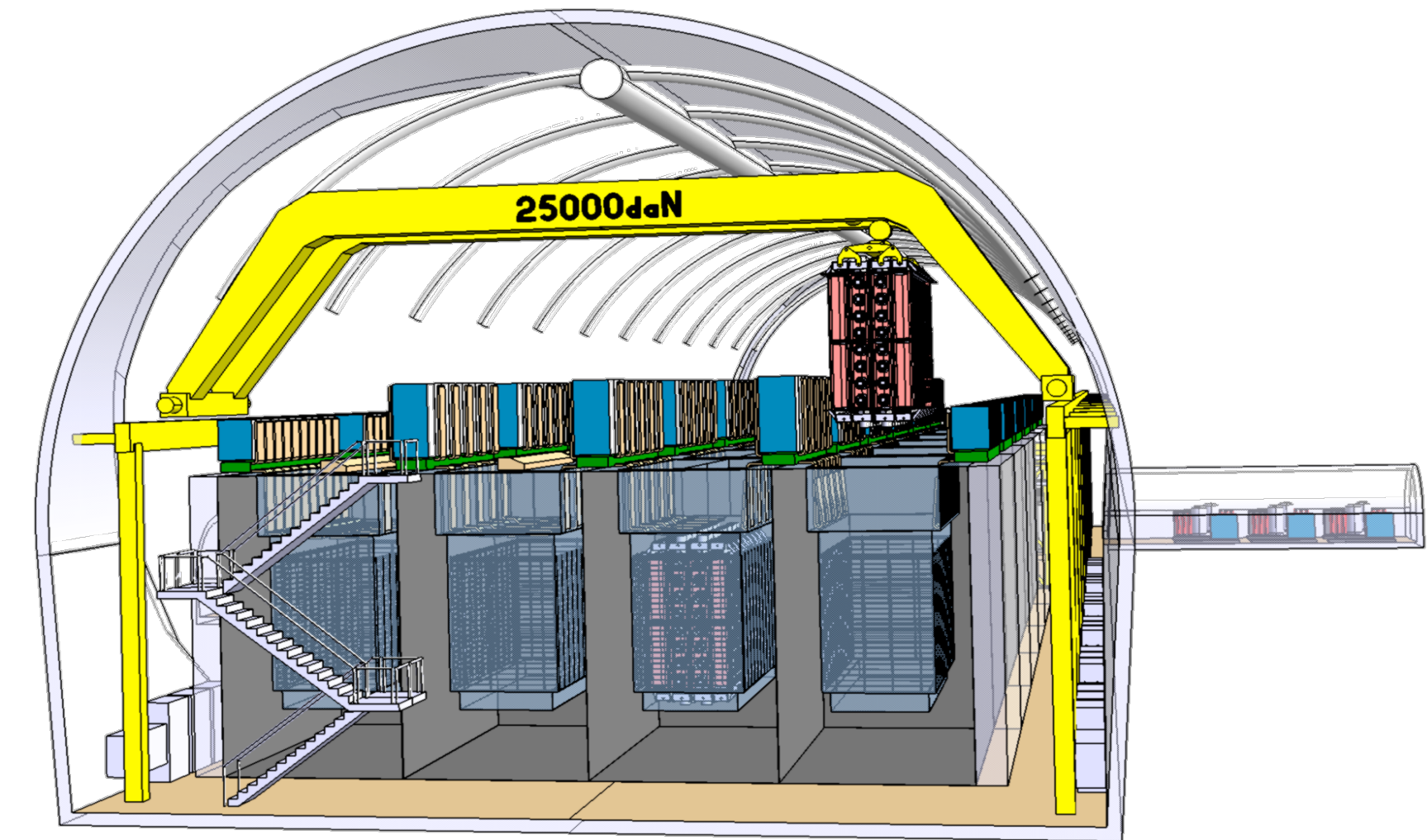
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$0\nu 4\beta$ : for  $^{150}\text{Nd}$

plus proof of concept for...



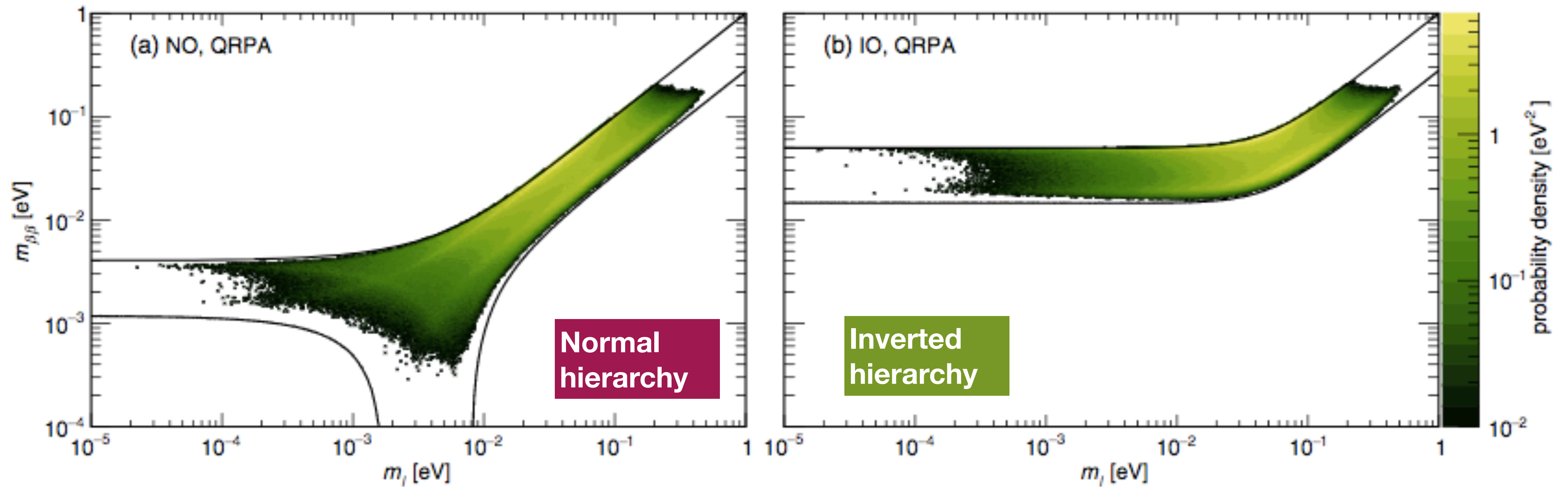
## Full SuperNEMO

- **Modular** design allows easy scaling up
- 20 modules x 5 years (500 kg year) gives sensitivity comparable or better than current **leading experiments**
- Best technique to understand more about  **$0\nu\beta\beta$  mechanism** in the event of discovery

Look to the future...

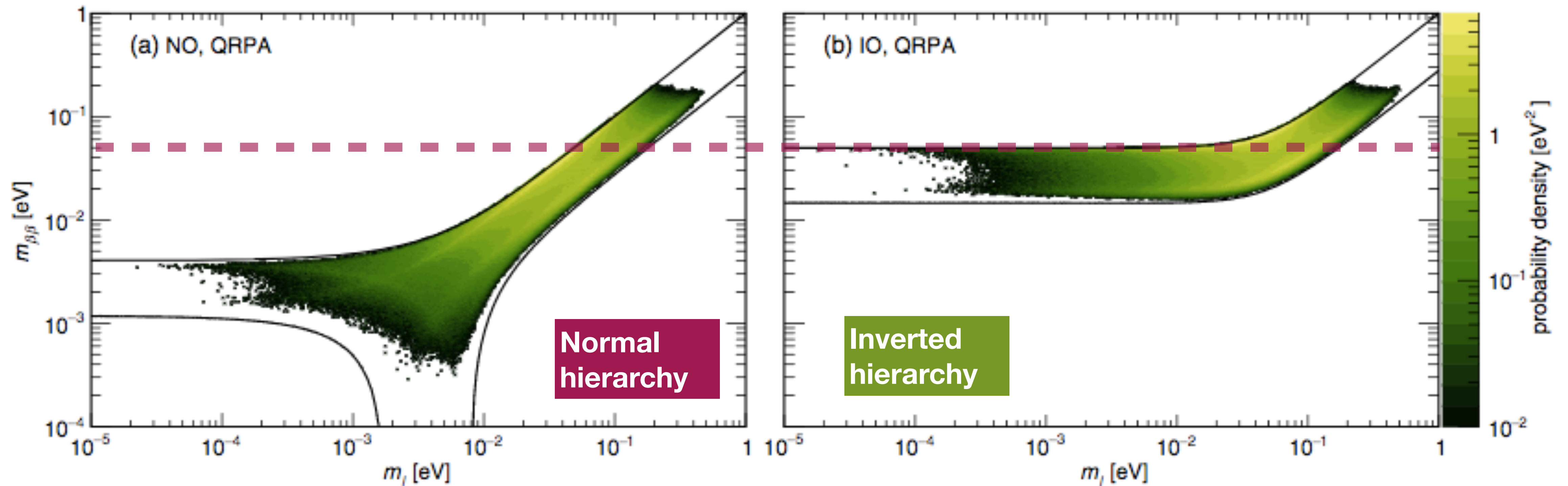
# Next-generation $0\nu\beta\beta$ searches

Bayesian probability density fit by D'Agostini, Benato & Detwiler: *Phys. Rev. D* 96, 053001 (2017)



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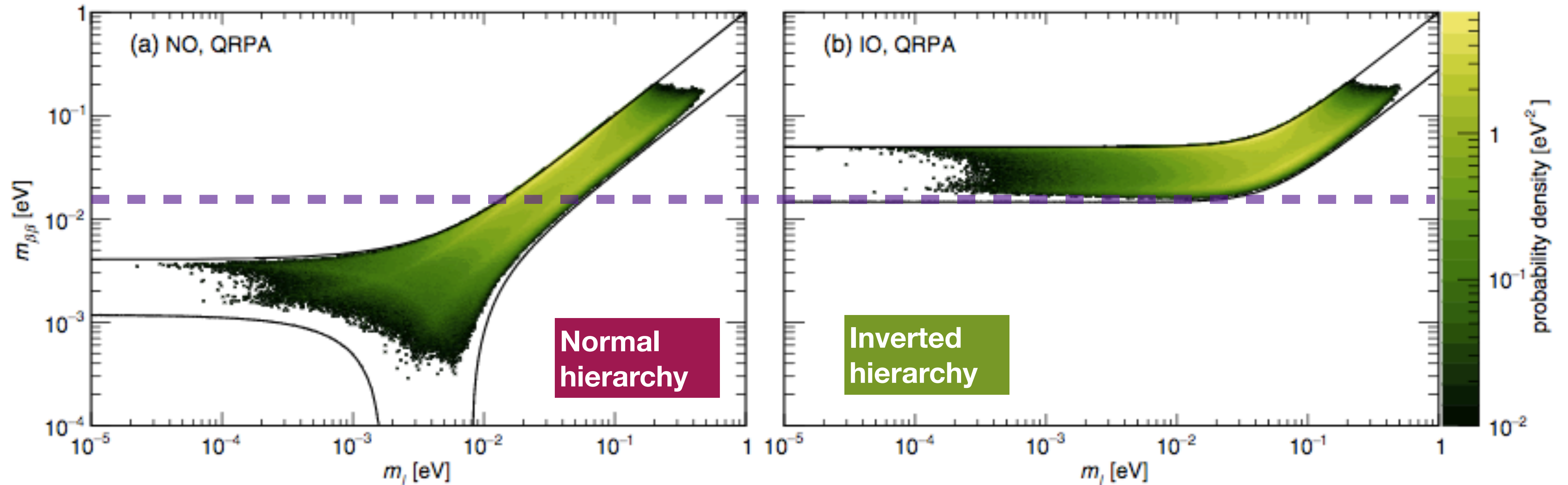
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- Current experiments probe the **degenerate** regime

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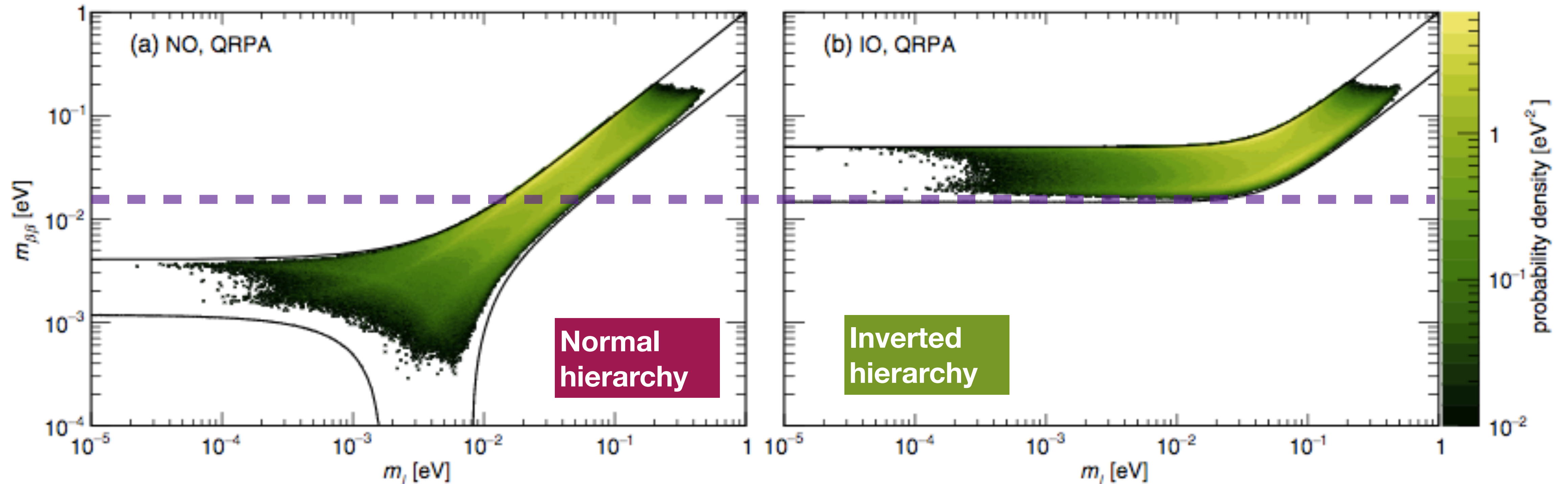
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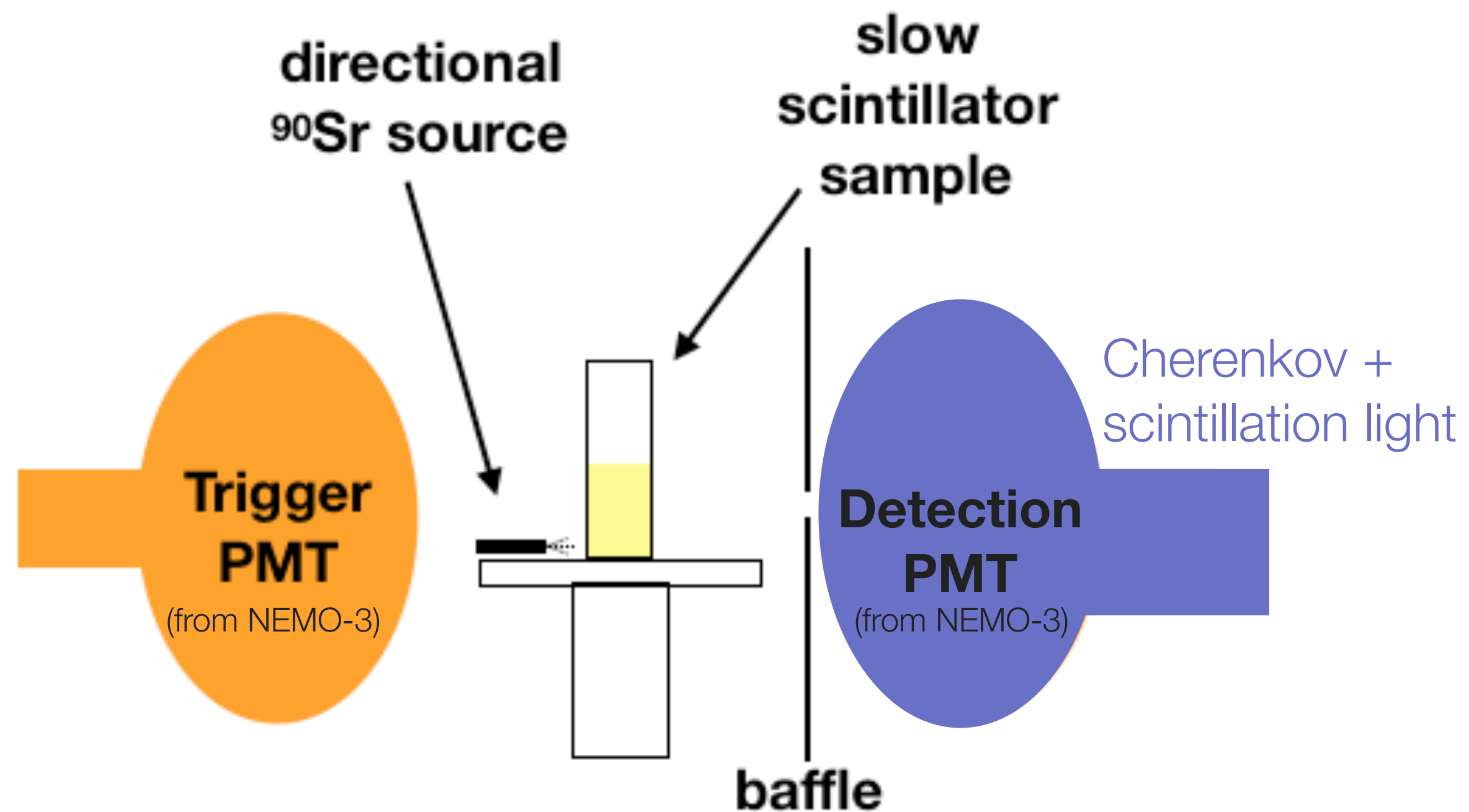


- Current experiments probe the **degenerate** regime
- Next-generation will cover **full inverted hierarchy** region
- When likelihood density is considered, this mass range also covers more than **50% of normal hierarchy** probability

# UK R&D: slow liquid scintillator

Several slow liquid scintillator mixtures developed at Oxford provide:

- excellent **time separation** of Cherenkov light to help reconstruct event **topology**
- high **scintillation light** yield for high **energy resolution**



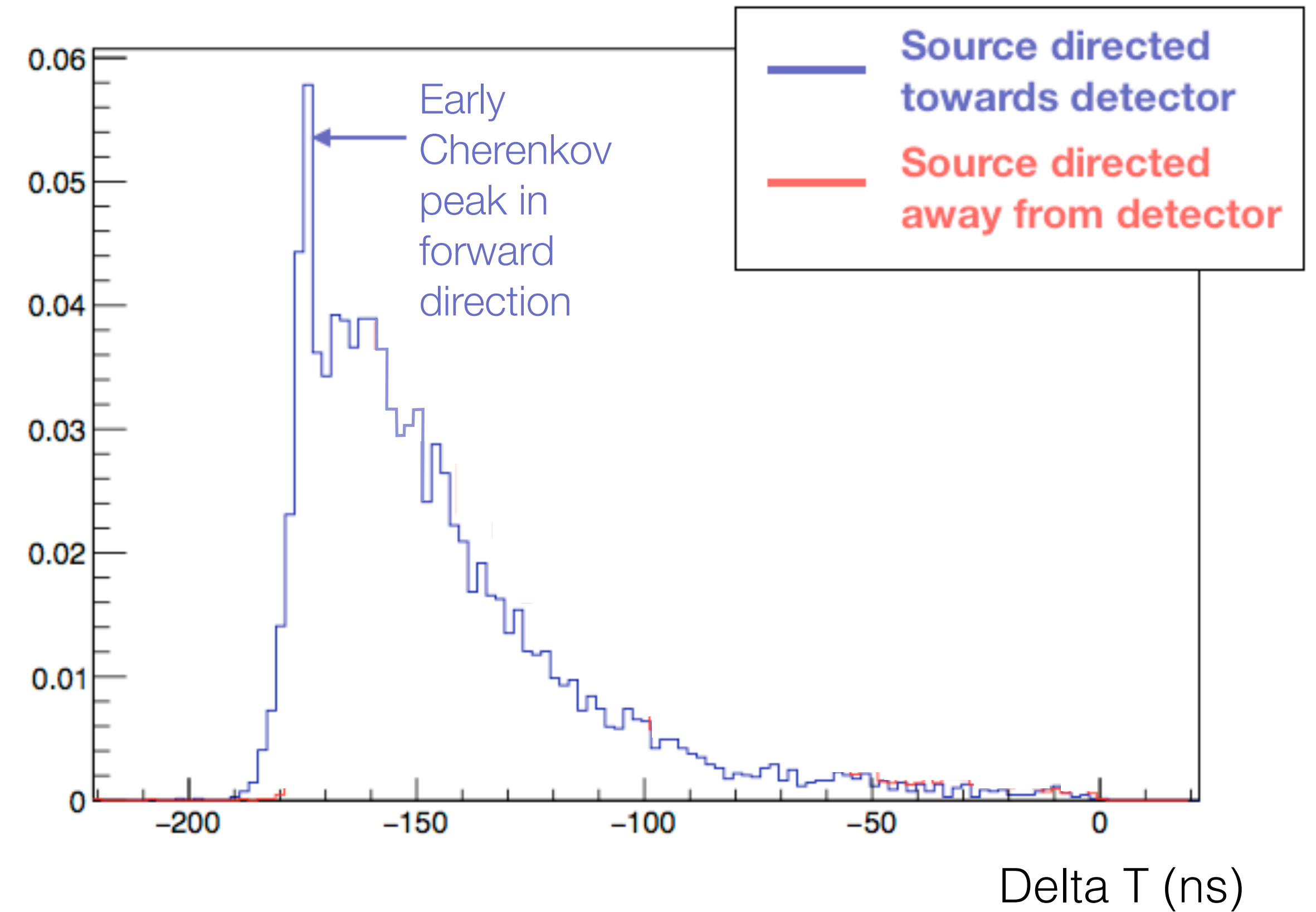
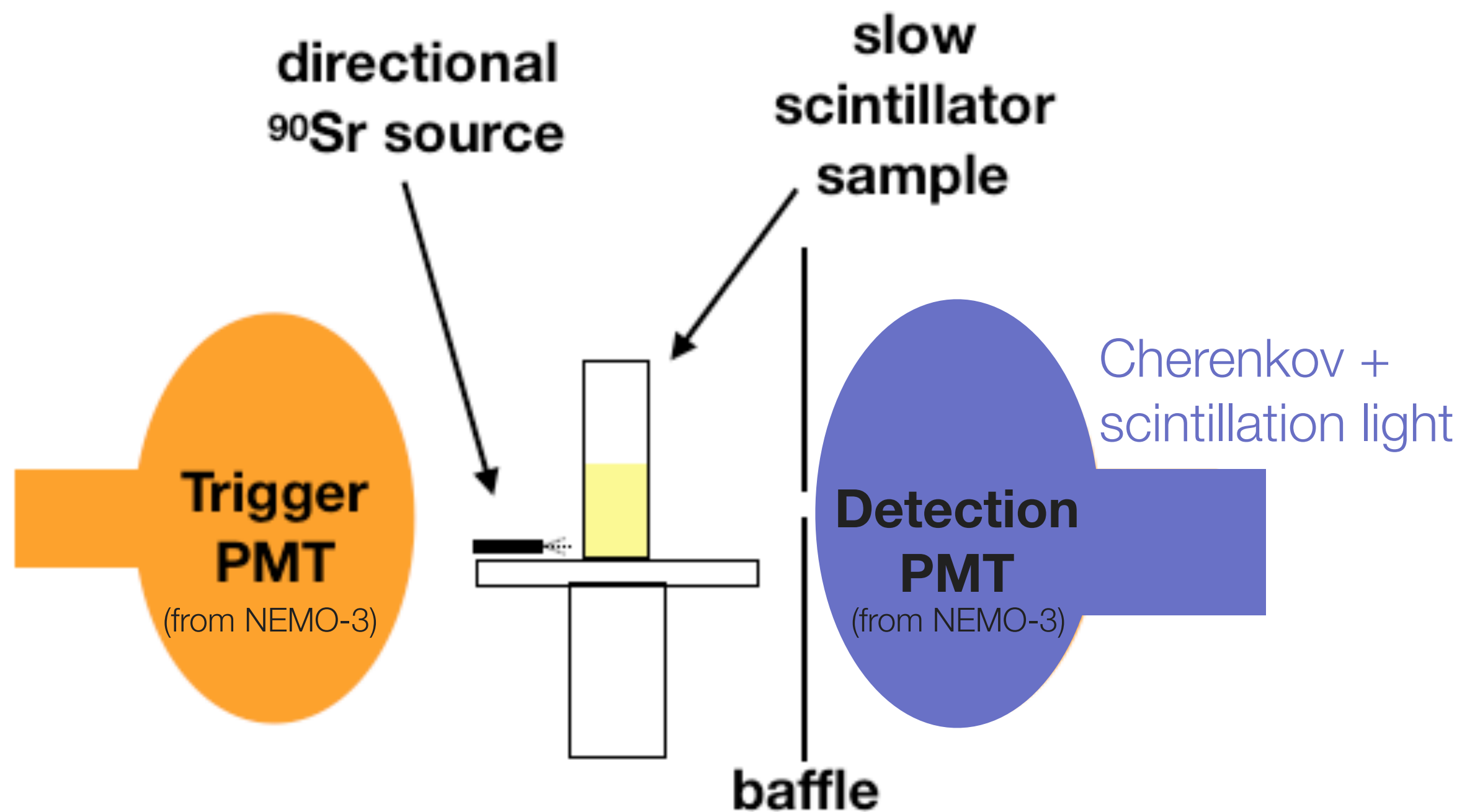
Thanks to Steve Biller for slide content



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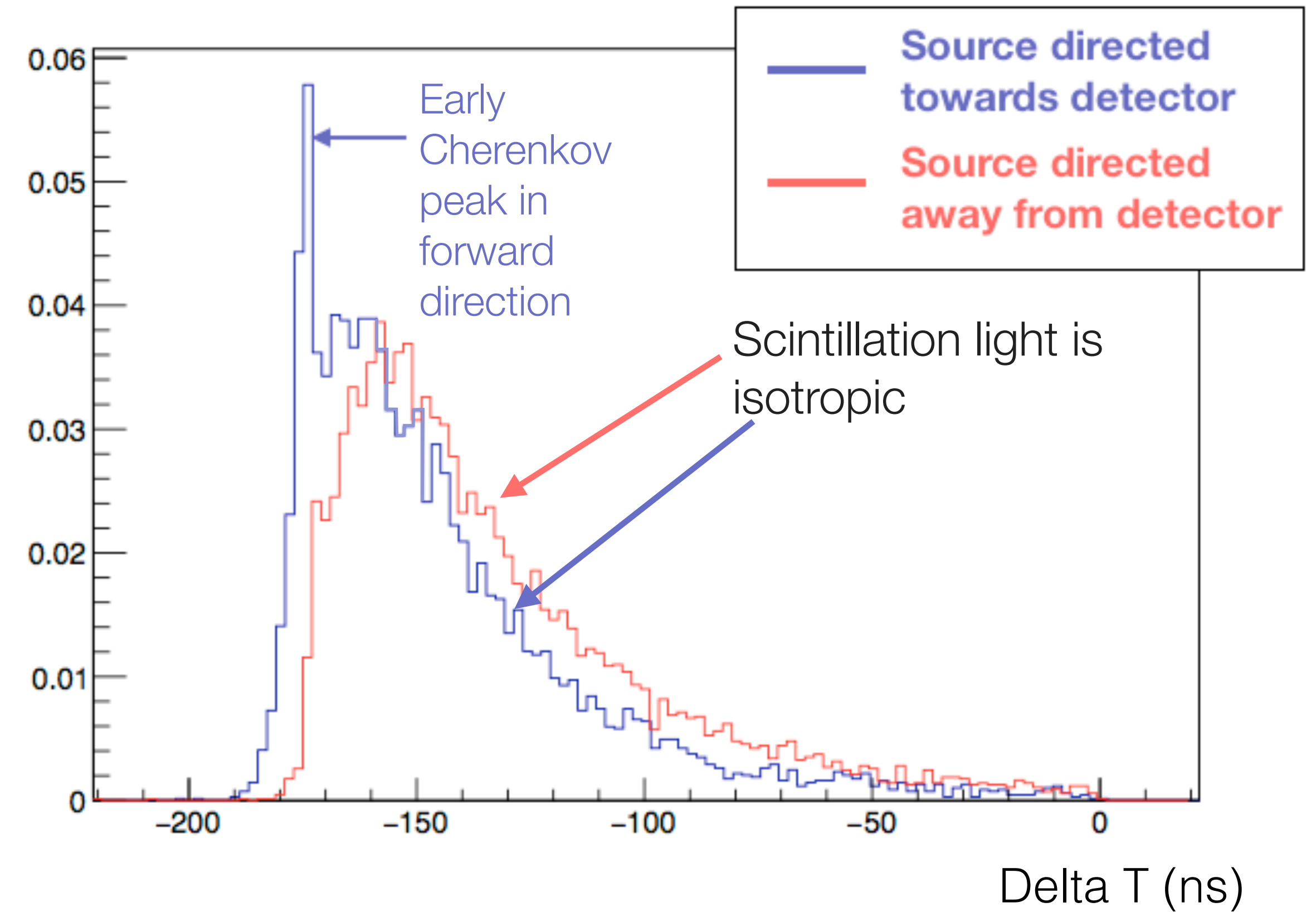
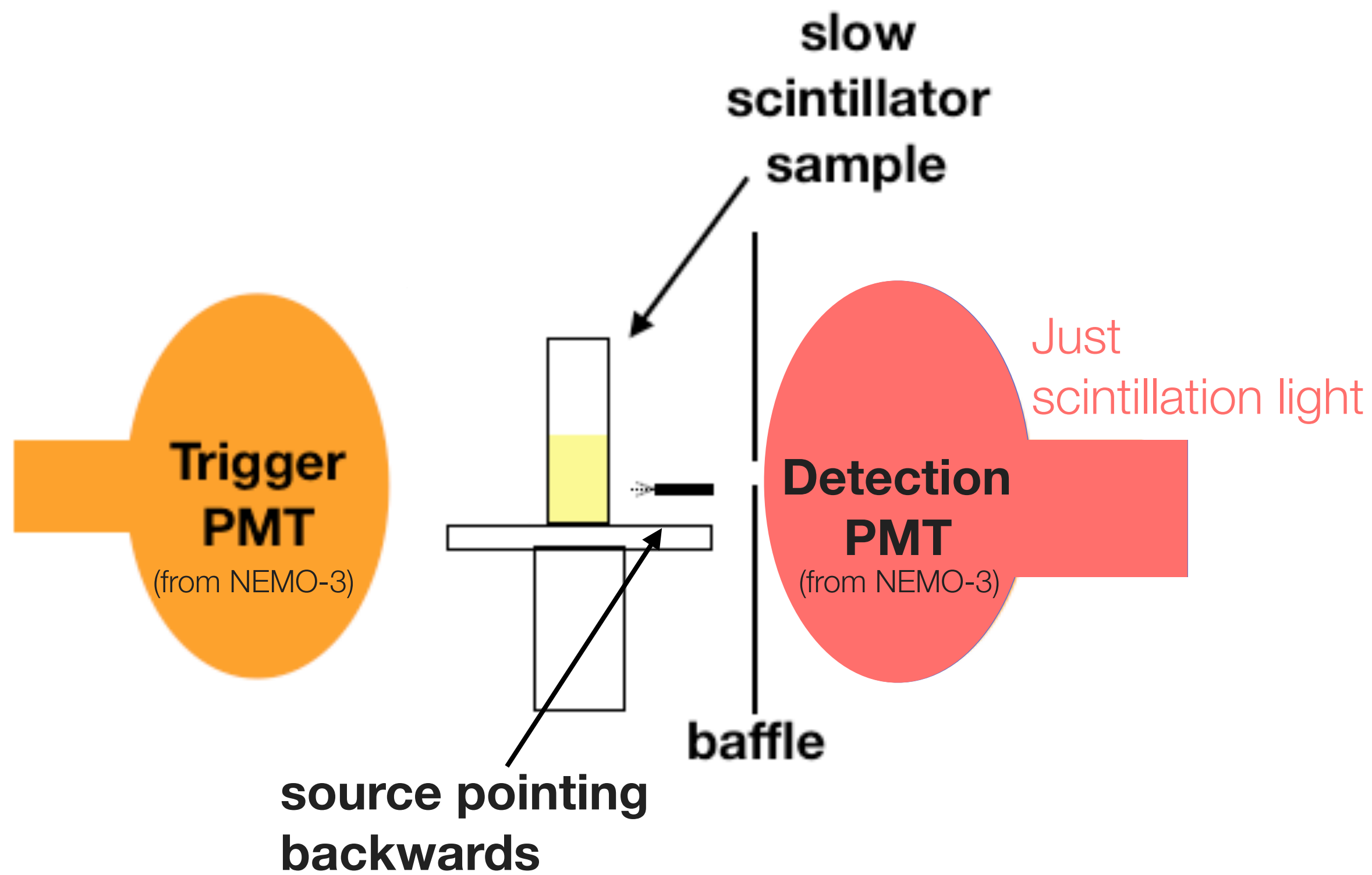


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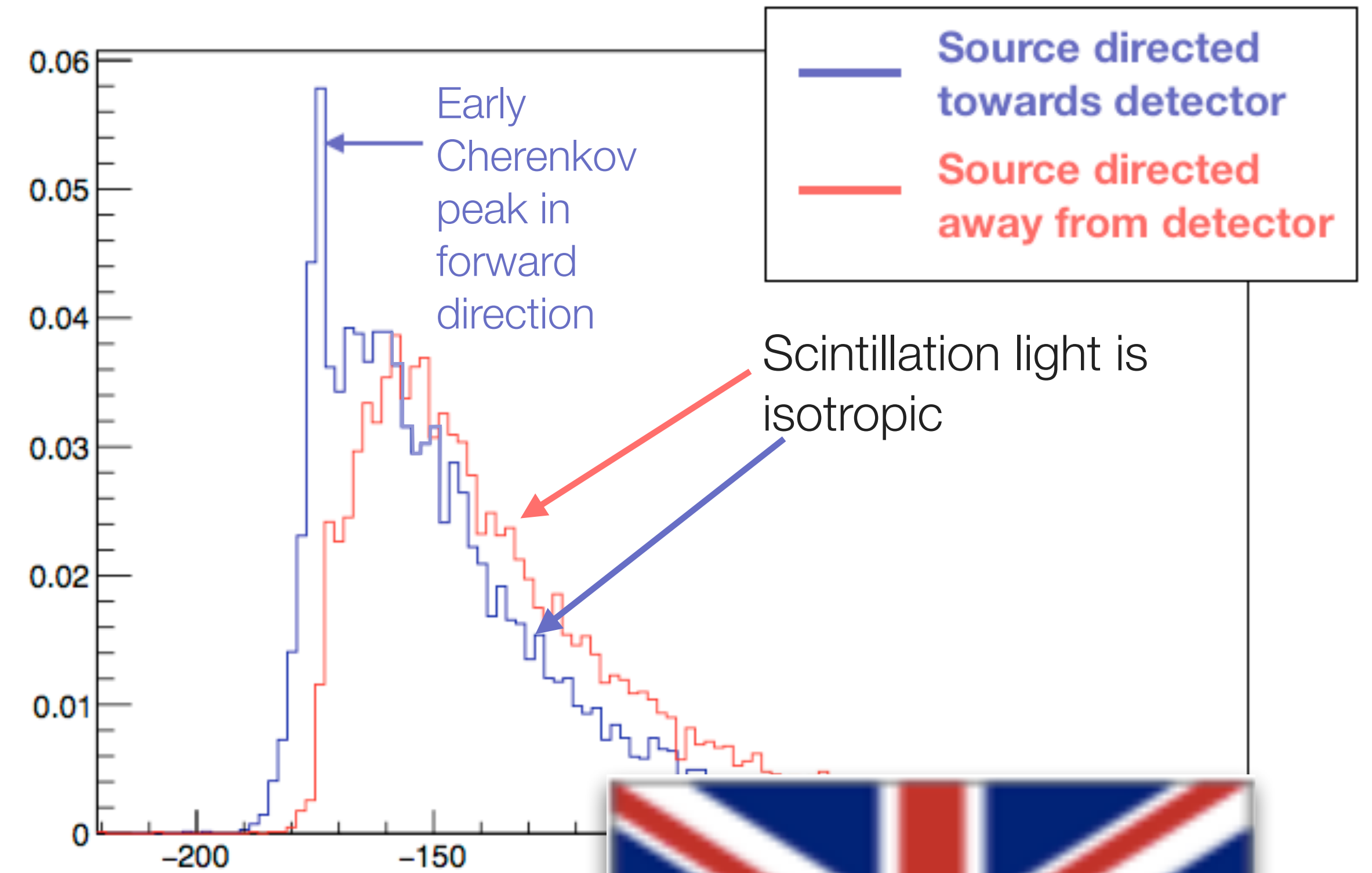
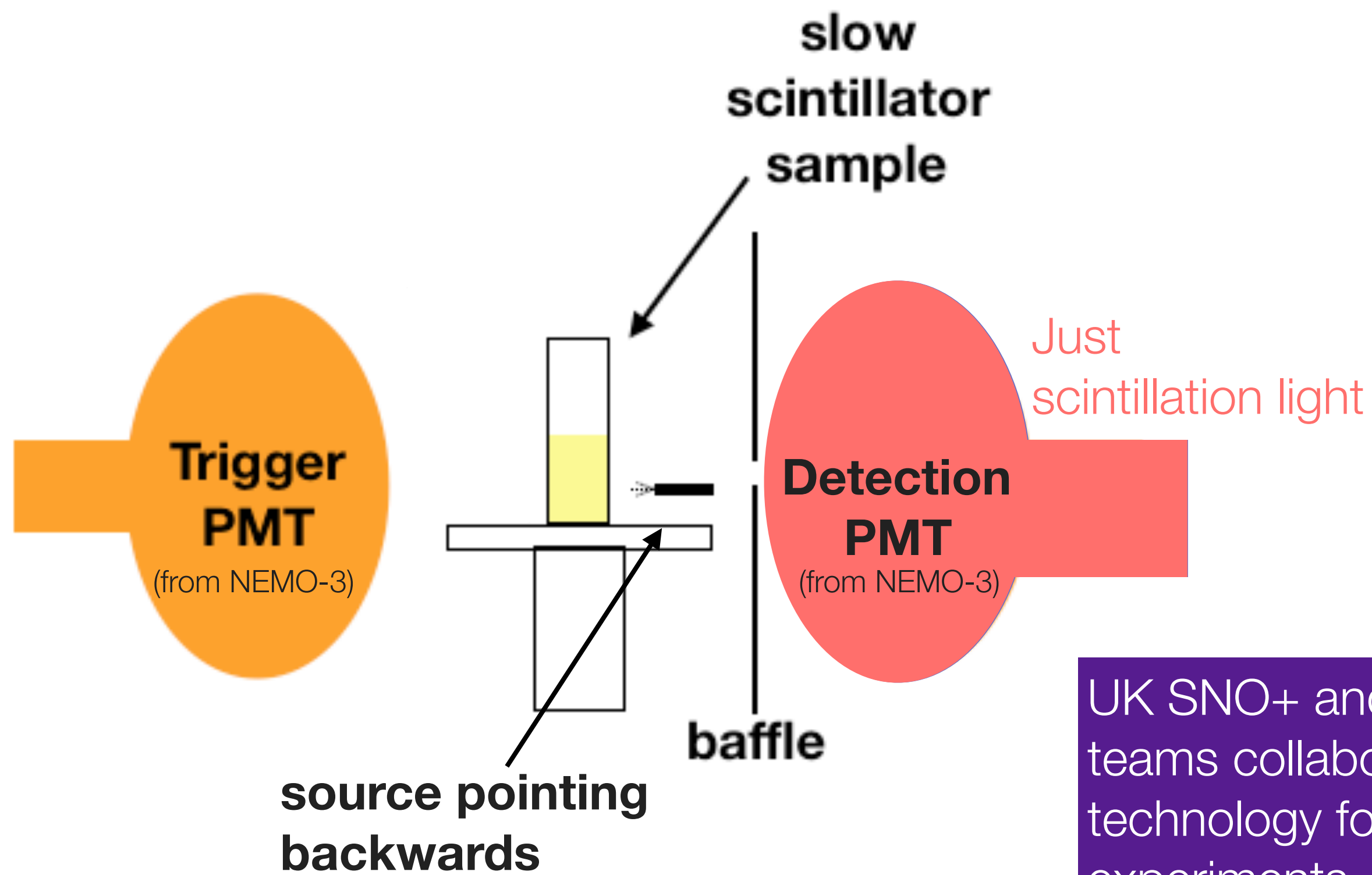


Thanks to Steve Biller for slide content

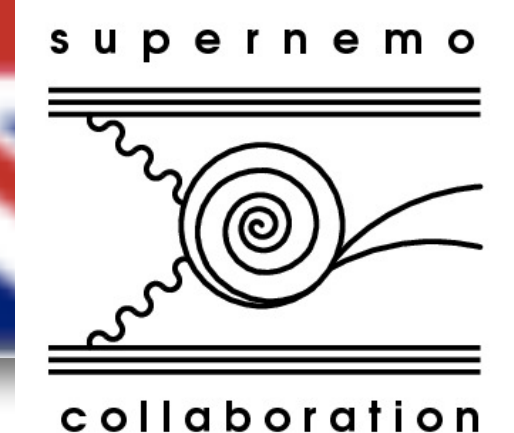
# UK R&D: slow liquid scintillator

Several slow liquid scintillator mixtures developed at Oxford provide:

- excellent **time separation** of Cherenkov light to help reconstruct event **topology**
- high **scintillation light** yield for high **energy resolution**



UK SNO+ and SuperNEMO teams collaborating to research technology for future  $0\nu\beta\beta$  experiments

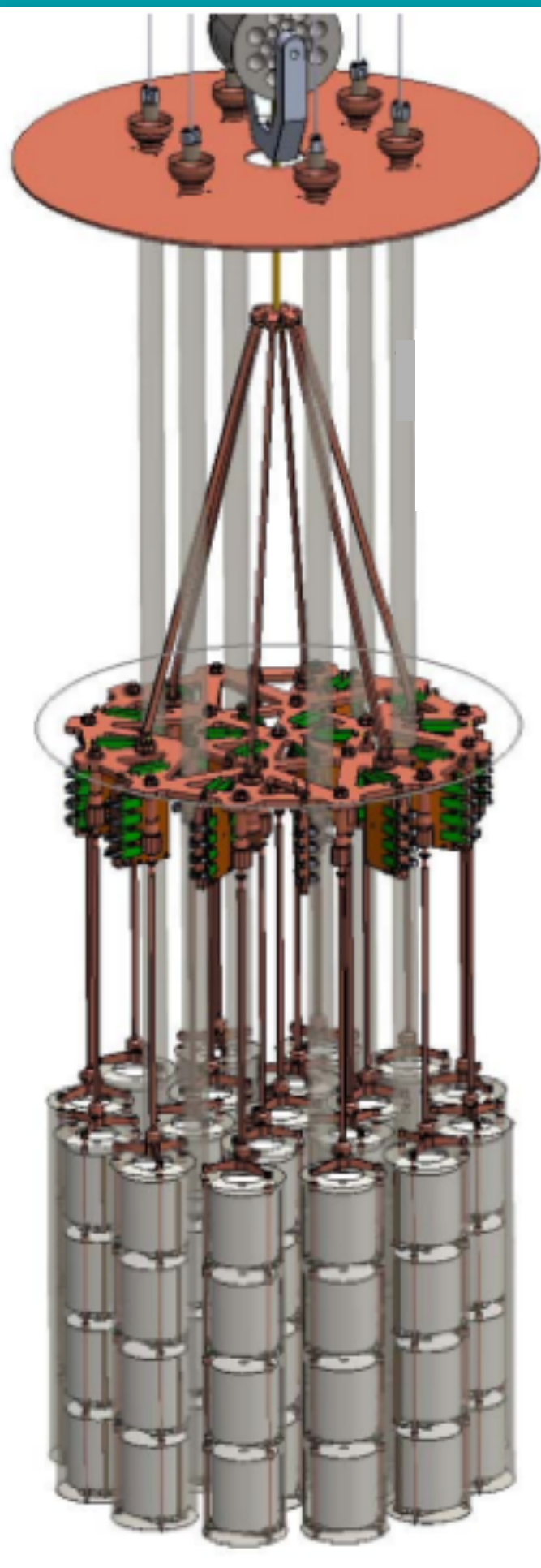


Thanks to Steve Biller for slide content

# UK future involvement - LEGEND HP<sup>76</sup>Ge detector

## LEGEND-200kg @ LNGS

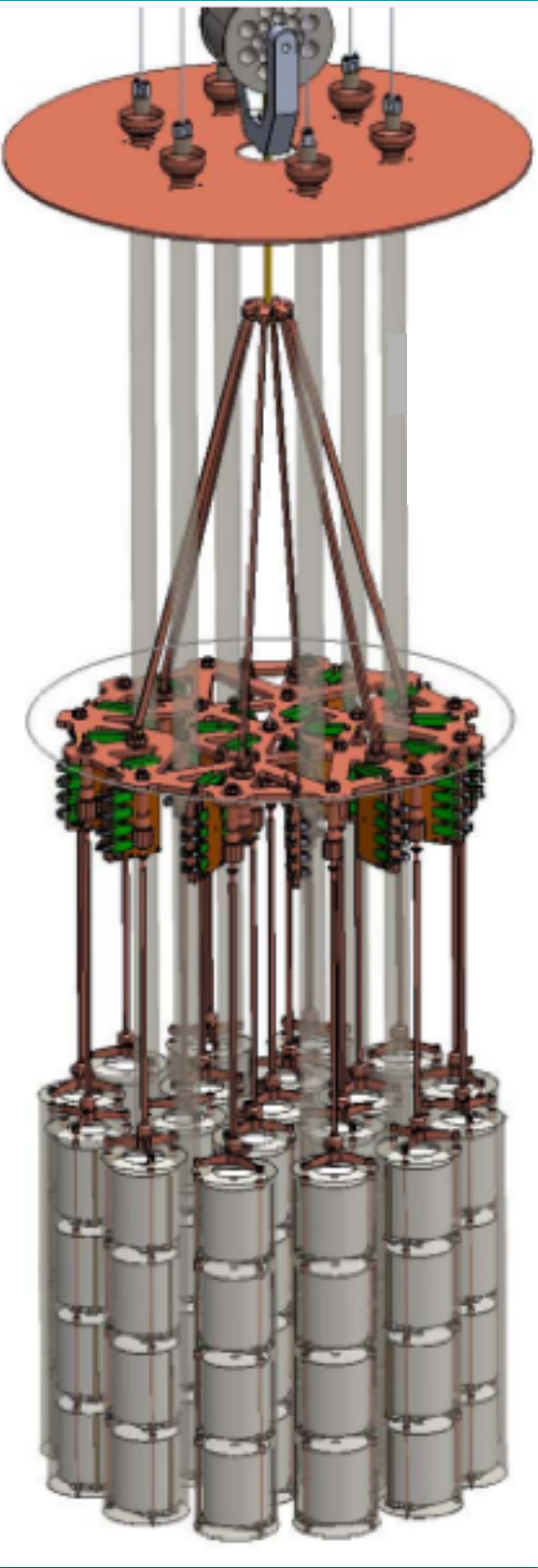
- **10<sup>27</sup> yrs** : 1 order of magnitude more sensitive than current leading experiments. Neutrino mass discovery reach **50 meV**.
- Start running **2021**, run for 5-7 years.



# UK future involvement - LEGEND HP<sup>76</sup>Ge detector

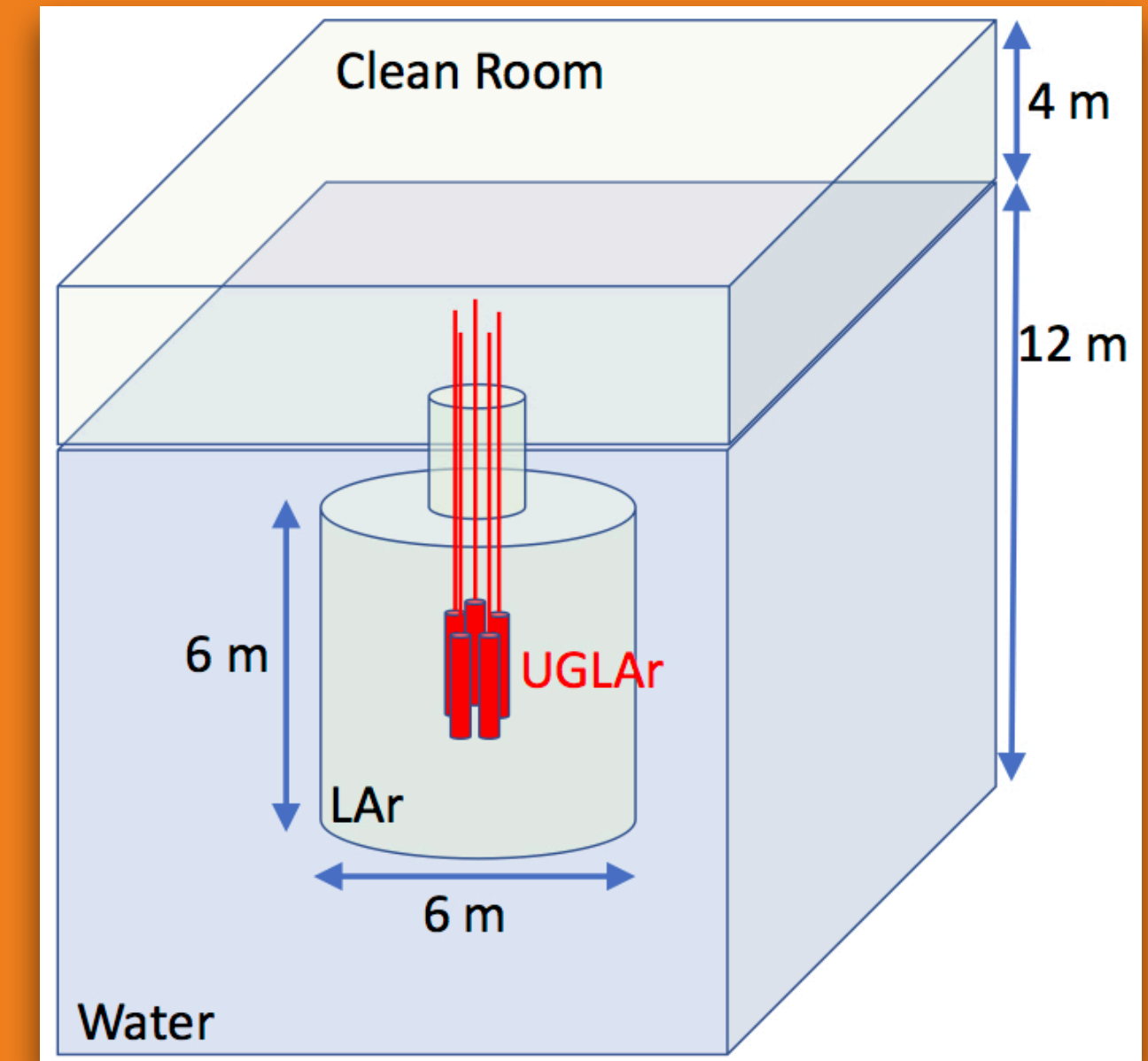
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## LEGEND-1T

- **10<sup>28</sup> yrs** : Neutrino mass discovery reach **18 meV** even for pessimistic NMEs.
- Turn on in **2025** with 1-tonne of isotope.



Thanks to Dave Waters for slide content

# UK future involvement - LEGEND HP<sup>76</sup>Ge detector

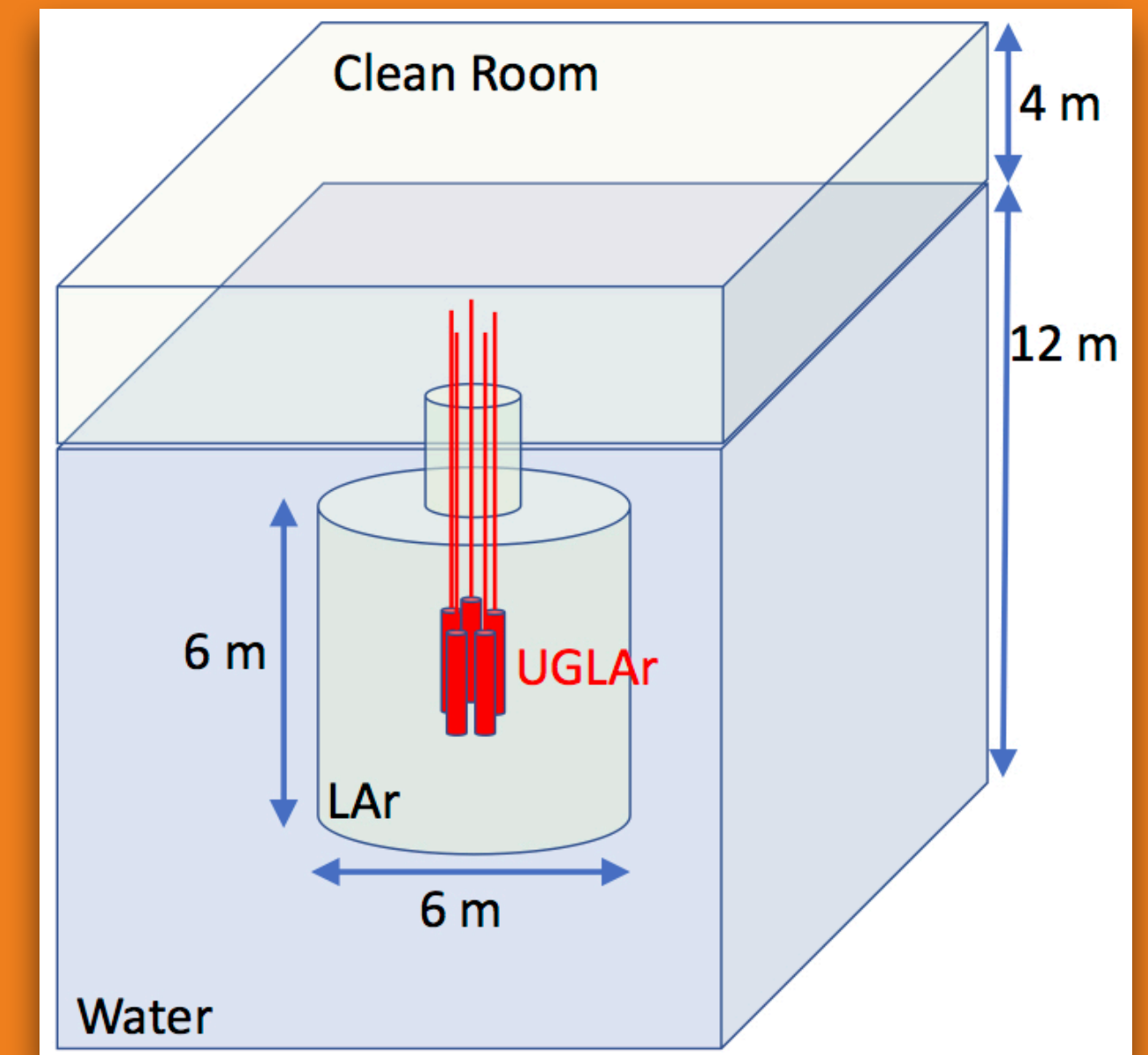
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Thanks to Dave Waters for slide content

Lancaster University



UNIVERSITY OF LIVERPOOL

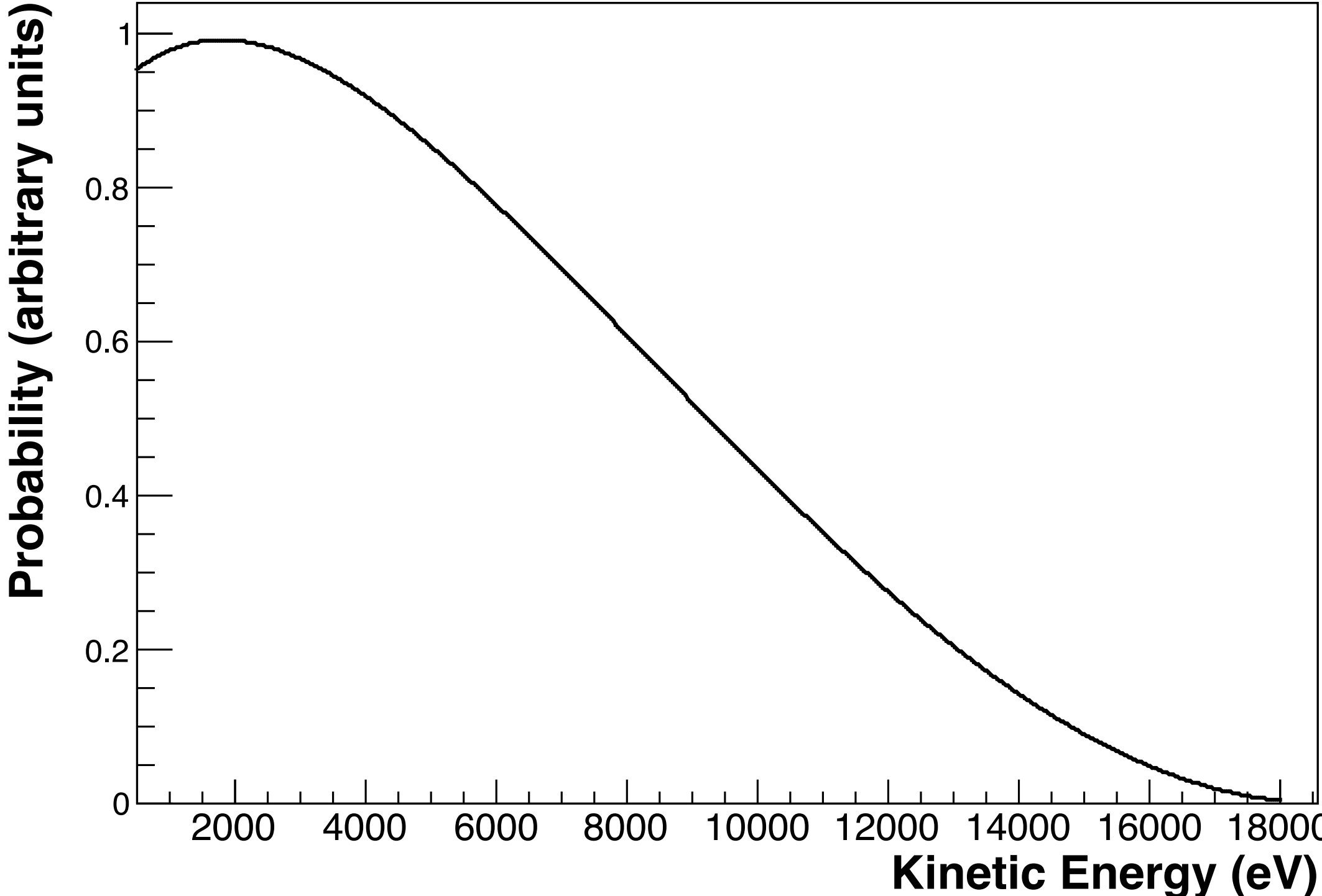
THE UNIVERSITY OF WARWICK



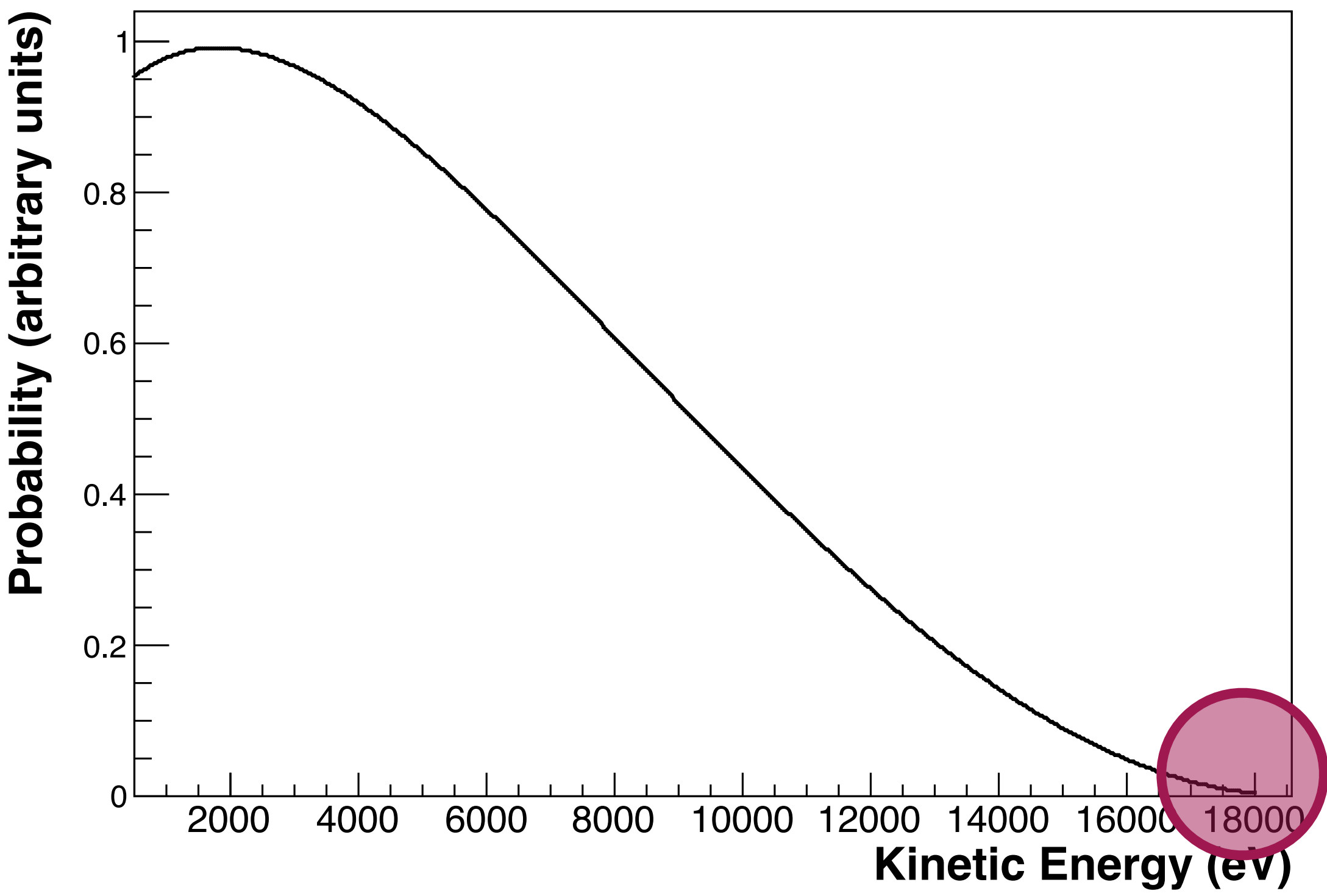
### UK Initial Participation :

- Combination of **Particle** Physics and **Nuclear** Physics groups.
- Builds on world-renowned expertise in **HPGe** detector development, **low-background** techniques and **software/analysis** expertise.

# Direct mass measurement: single $\beta$ -decay endpoint

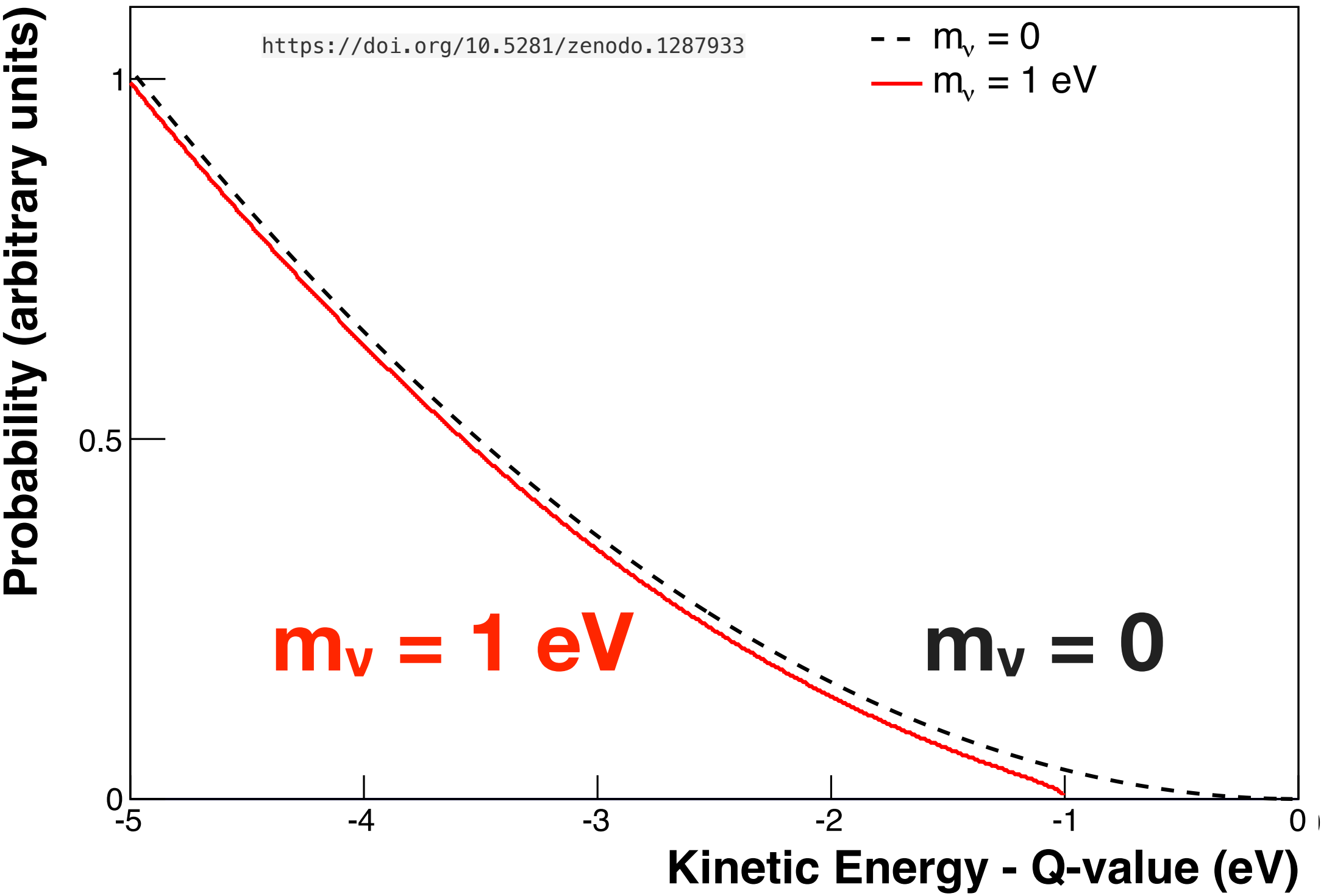


# Direct mass measurement: single $\beta$ -decay endpoint

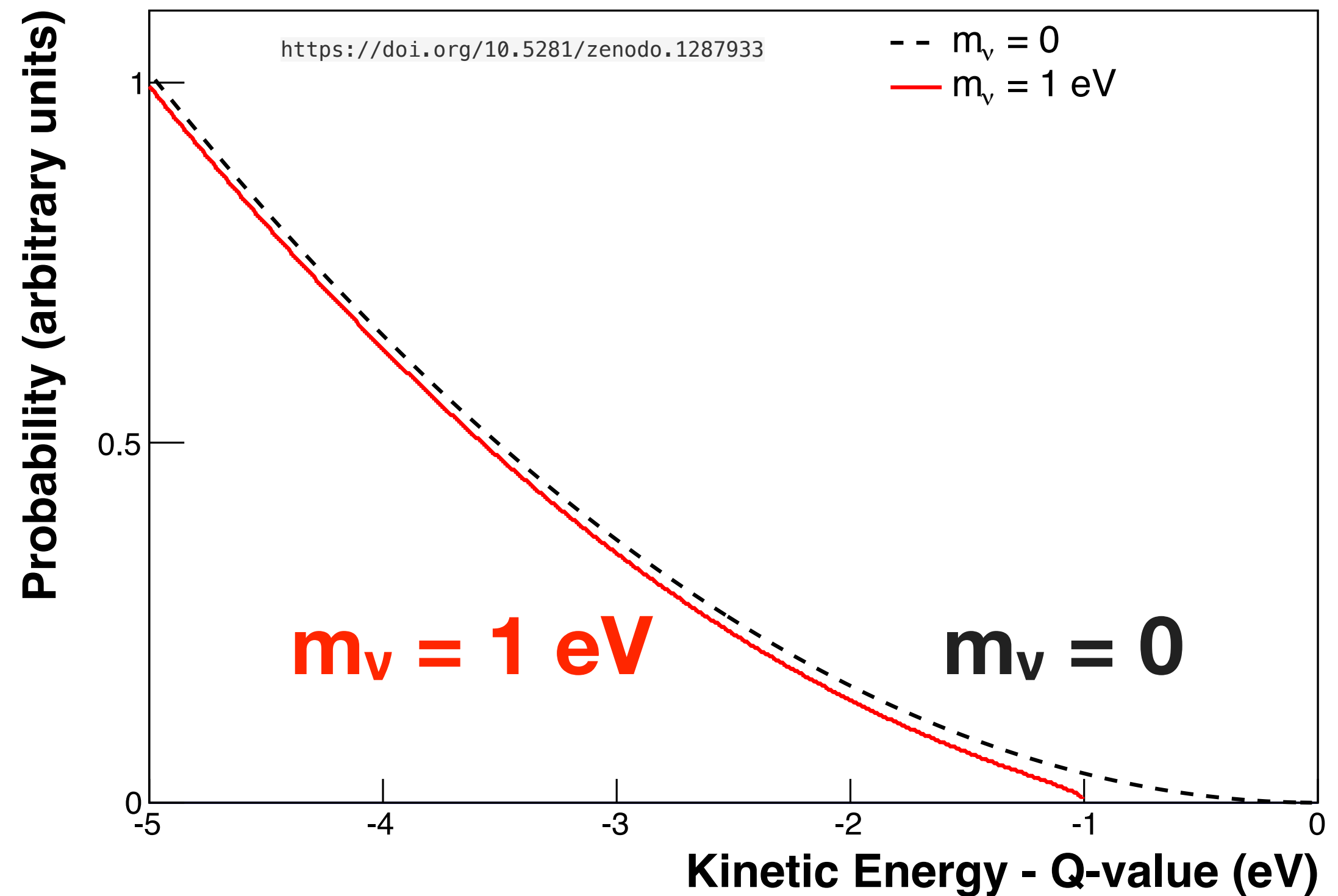




# Direct mass measurement: single $\beta$ -decay endpoint



# Direct mass measurement: single $\beta$ -decay endpoint




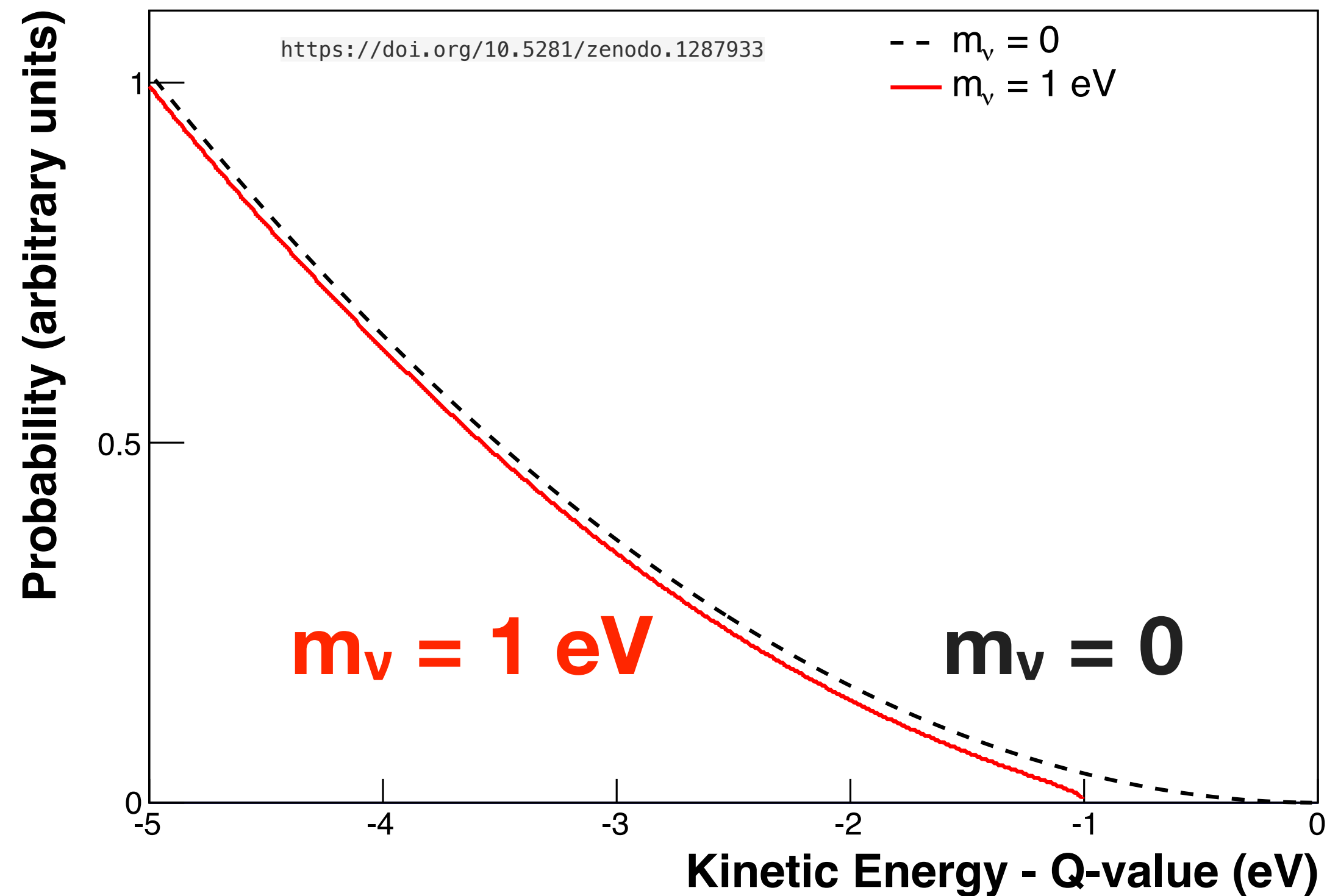
## KATRIN (Karlsruhe, Germany)

- Launched **June 2018**
- $\beta$  decay of **tritium molecules**
- Spectrometer uses collimator/filter to measure highest energy electrons
- Sensitivity  $m_\nu < 240 \text{ meV}$  is the best achievable with this technique




Photo: KATRIN

# Direct mass measurement: single $\beta$ -decay endpoint



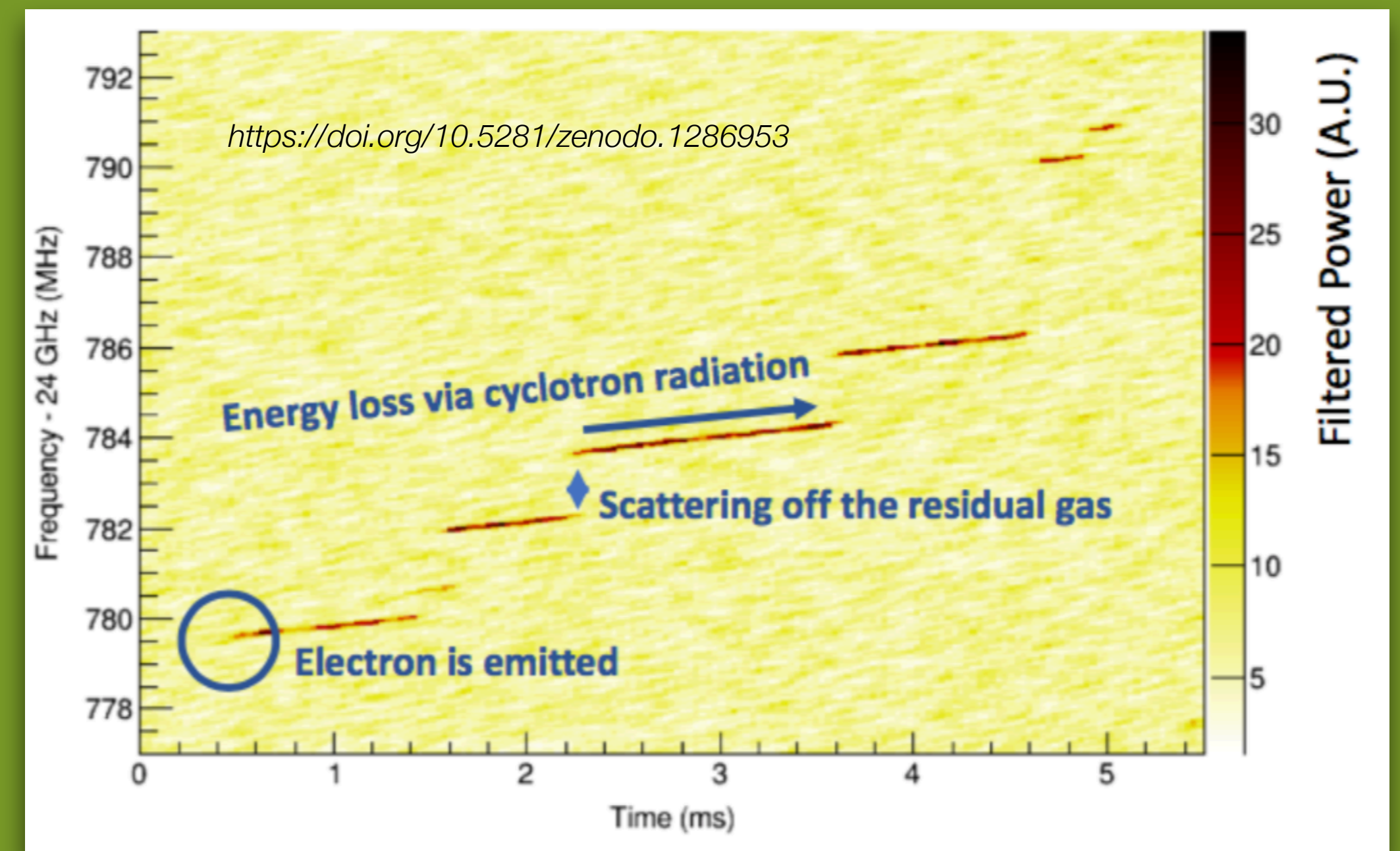
Quantum Sensors for  
Fundamental Physics

See Jon Coleman's slides!

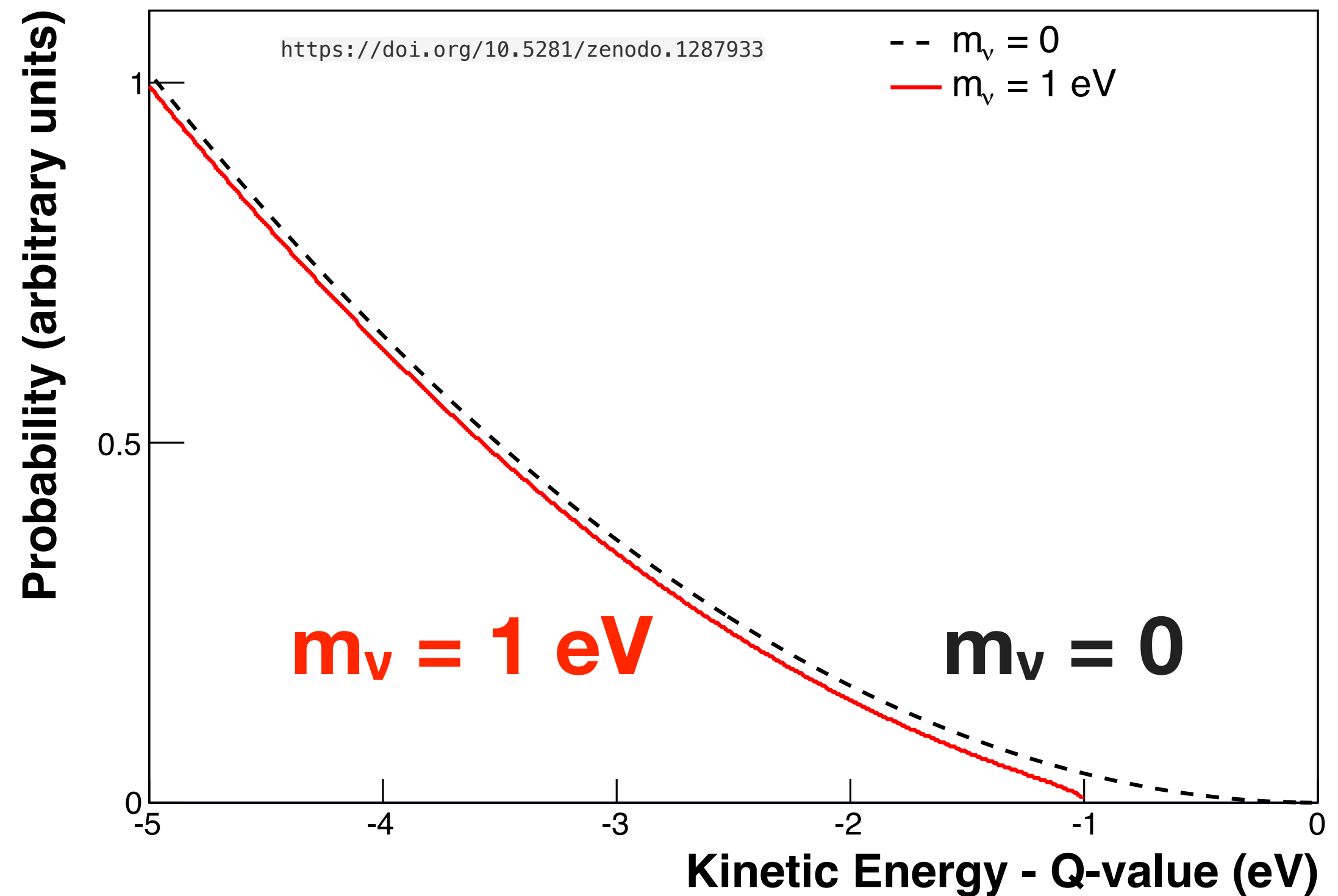


## Project 8 (design phase)

- Electron energy measured using **cyclotron radiation**: frequency related to kinetic energy
- **Atomic tritium** improves sensitivity to **40meV**

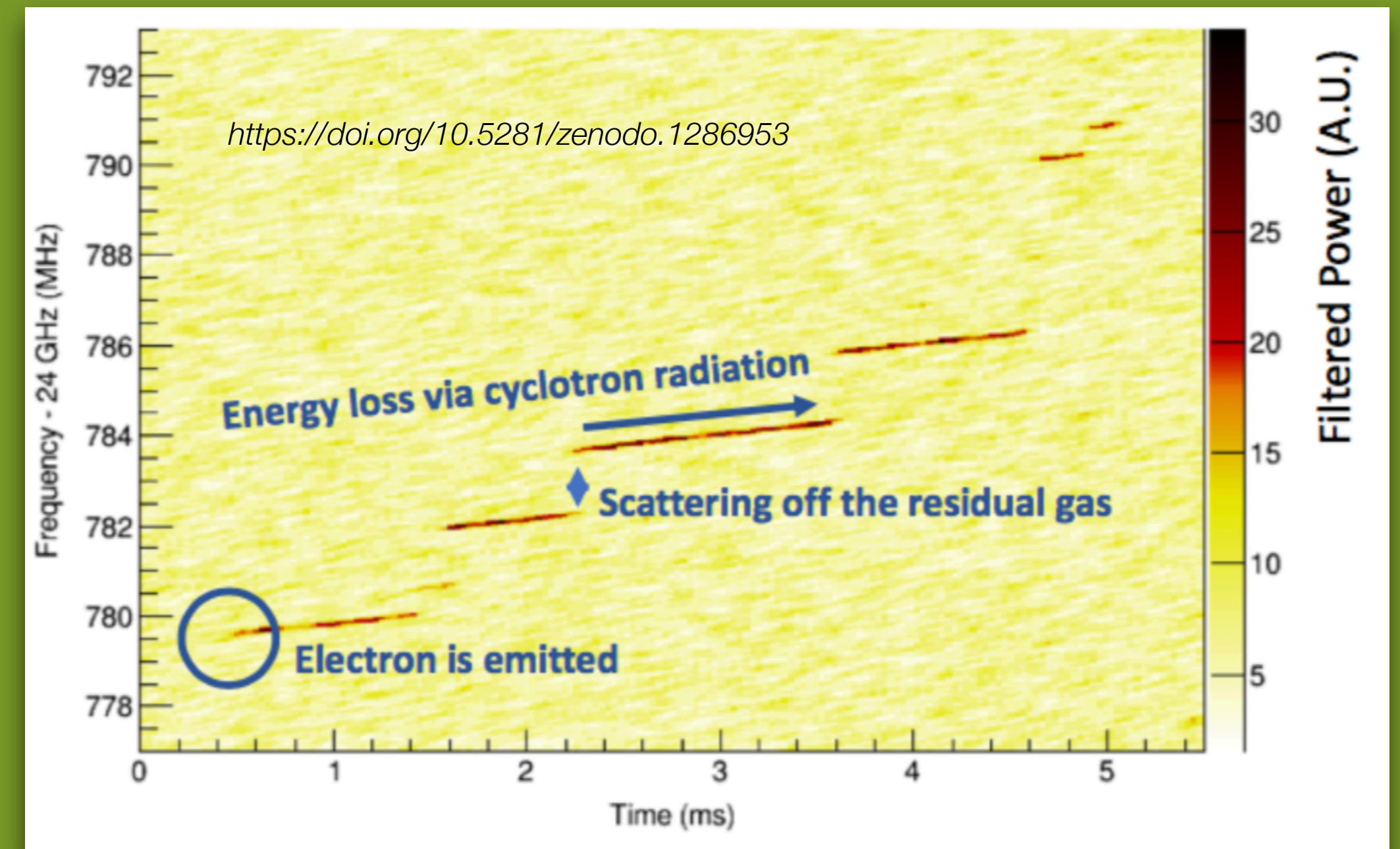


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**QS FP** Quantum Sensors for Fundamental Physics

See Jon Coleman's slides!

Also for future  $^{163}\text{Ho}$  electron-capture endpoint experiments

**HVLMES**

**ECHo**

# $\beta\beta$ decay in the UK



Entering exploitation phase: scintillator filling at SNO+, commissioning data at SuperNEMO

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Slow scintillator

LEGEND

Low-background techniques

Quantum sensors

SNO+ phase II

Full SuperNEMO

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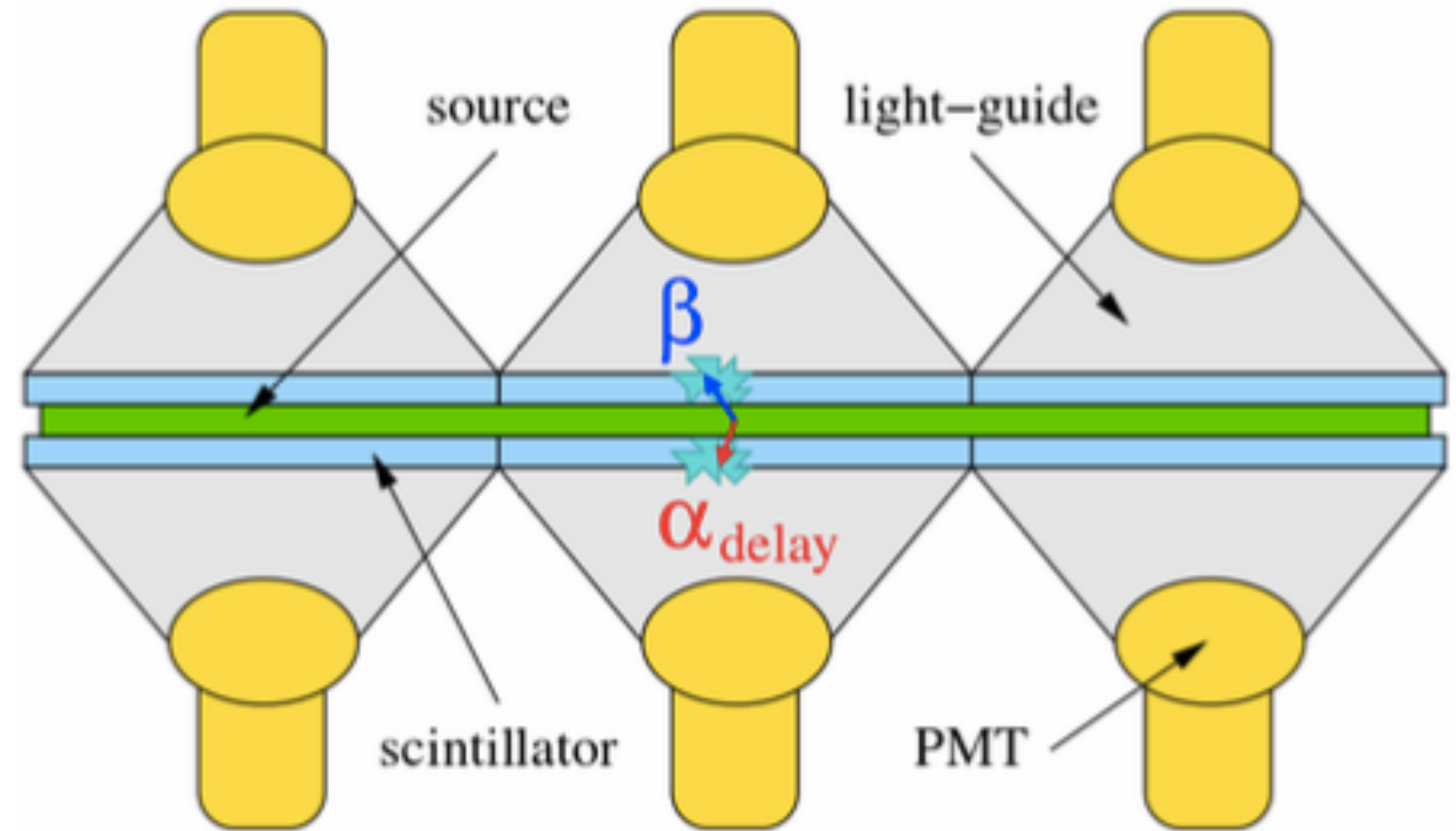
Backup Slides



# Source foil contamination measured at the BiPo-3 detector



JINST 12 (2017) P06002



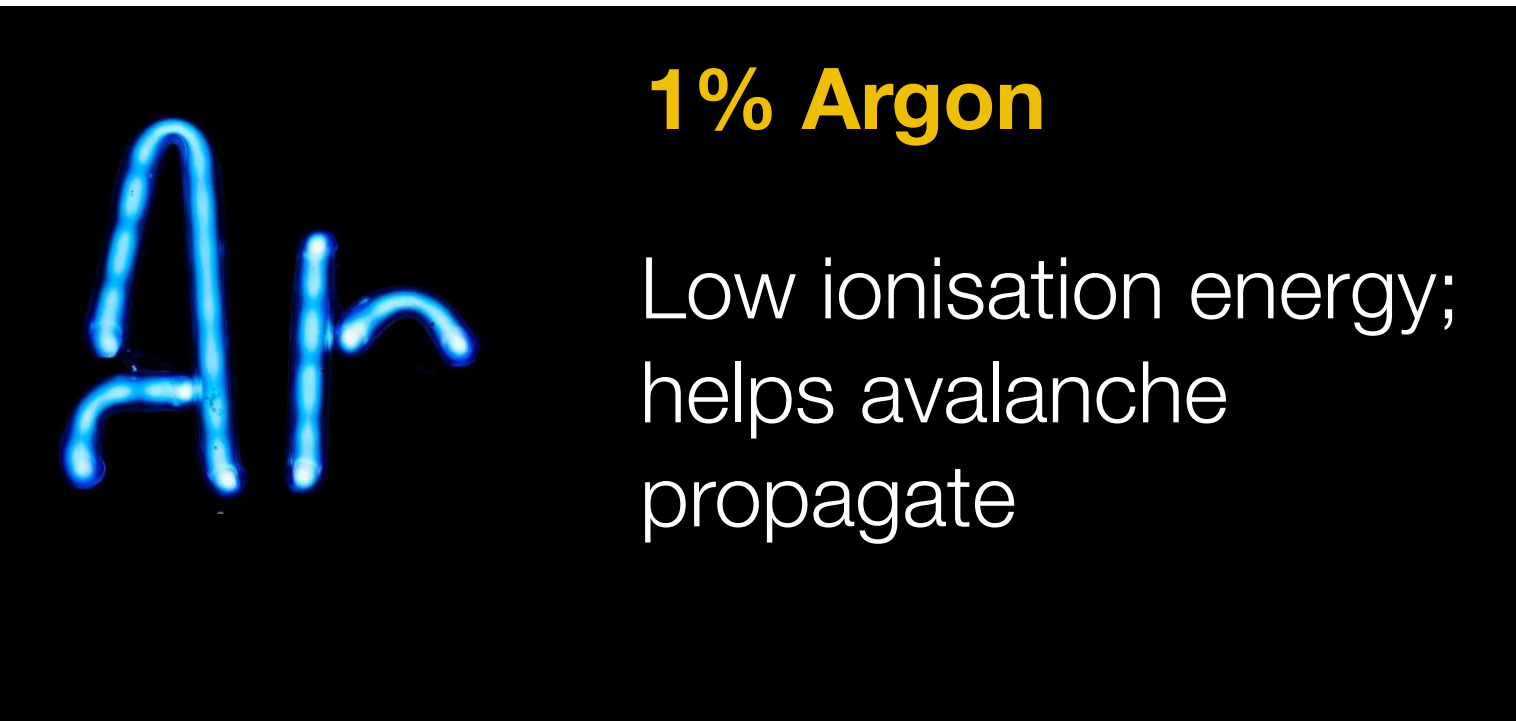
- Dedicated detector at Canfranc, Spain
- Designed to measure very **low activities**
- Looks for characteristic signature of Bi  $\beta$  decay followed by  $\alpha$  decay of Po daughter (U and Th decay chains)
- Targets  **$10\mu\text{Bq}/\text{kg}$  ( $^{214}\text{Bi}$ ),  $2\mu\text{Bq}/\text{kg}$  ( $^{208}\text{Tl}$ )**
- Not very sensitive to  $^{214}\text{Bi}$  - final measurements will be taken *in situ*

# Tracker gas system



## 95% Helium

Low atomic mass; prevents multiple scattering and energy loss



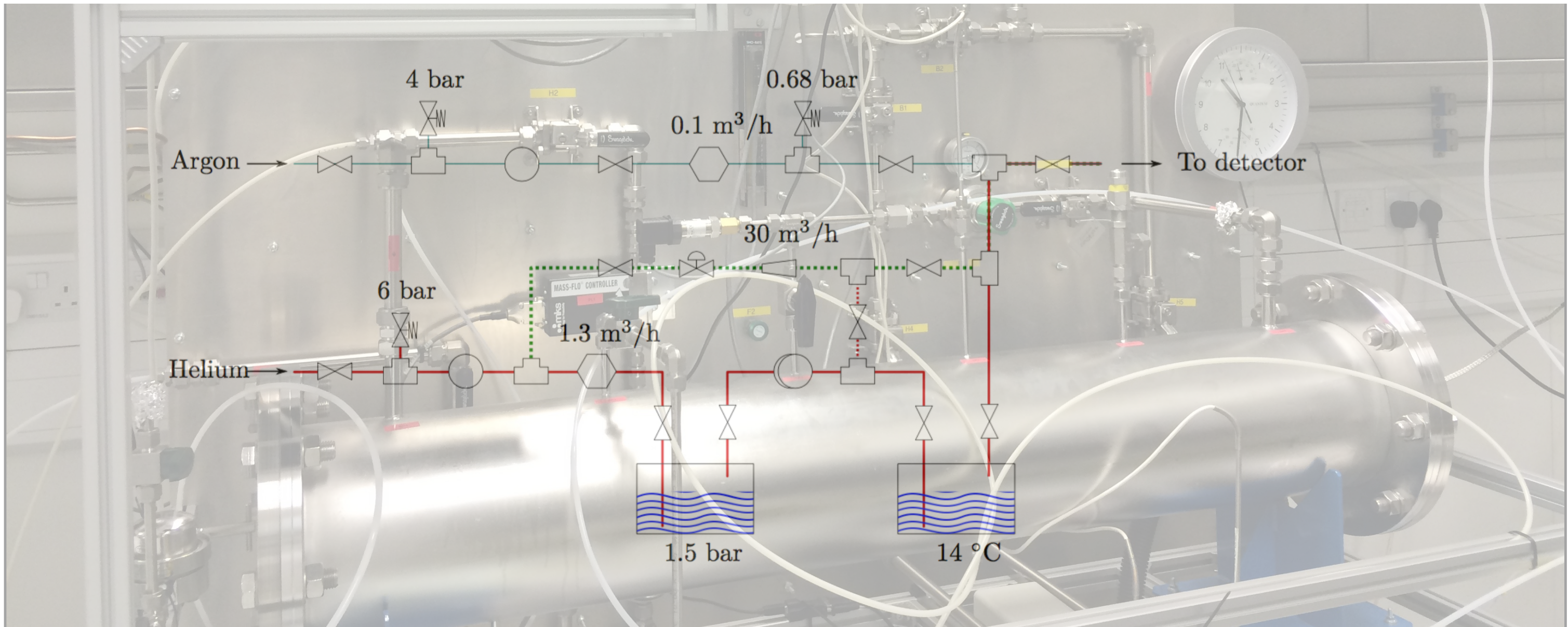
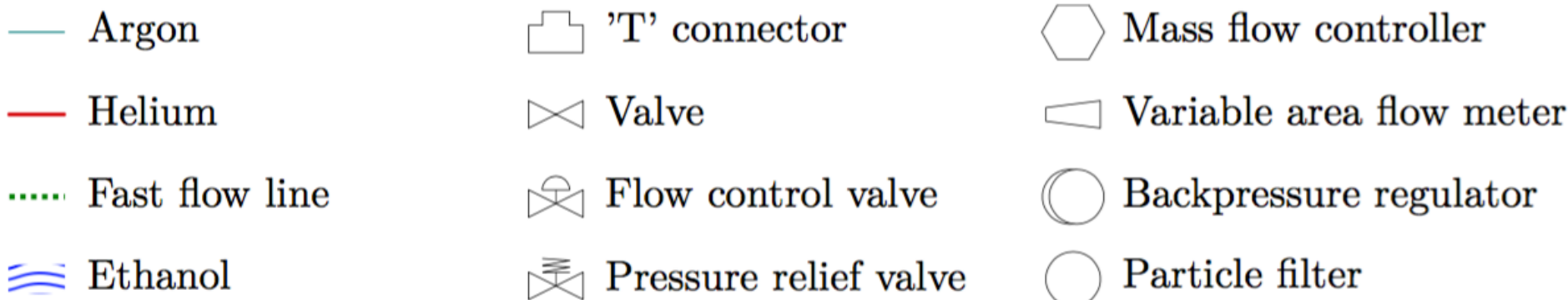
## 1% Argon

Low ionisation energy; helps avalanche propagate



## 4% Ethanol

Quenches avalanche; prevents re-firing



Gas system controlled by Raspberry Pi to monitor and control temperature, pressure, flow rate  
 2°C temperature change → 0.5% change in ethanol fraction → tracker efficiency

# Event count targets in SuperNEMO demonstrator

## Aiming at zero background

Events in window $E_{\text{SUM}} \in [2.8, 3.2] \text{ MeV}$	NEMO-3 Phase 2 (29 kg.yr)	Demonstrator Module (29 kg.yr)	Comments
External Bkgnd	<0.16	<0.16	(conservative)
Bi214 from Rn222	$2.5 \pm 0.2$	0.07	radon reduction
Bi214 internal	$0.80 \pm 0.08$	0.07	internal contamination reduction
Tl208 internal	$2.7 \pm 0.2$	0.05	
$2\nu\beta\beta$	$7.16 \pm 0.05$	0.20	Mo100 to Se82 8% to 4% resolution
Total expected	$13.1 \pm 0.3$	0.39	
Data	12	N/A (yet)	

NEMO-3  
sensitivity in  
**4.5 months !**

# NEMO-3 results summary

Isotope	Mass (g)	$Q_{\beta\beta}$ (keV)	$T(2\nu)$ ( $\times 10^{19}$ yrs)	S/B	Comment	Reference
Se82	932	2997.9	$9.4 \pm 0.6$	4	World's best	Eur. Phys. J. C (2018) 78: 821
Cd116	405	2813.5	$2.74 \pm 0.18$	10	World's best*	Phys. Rev. D 95 (2017) 012007
Nd150	37	3371.4	$0.93 \pm 0.06$	2.7	World's best	Phys. Rev. D 94 (2016) 072003
Zr96	9.4	3355.8	$2.35 \pm 0.21$	1	World's best	Nucl.Phys.A 847(2010) 168
Ca48	7	4268	$6.4 \pm 1.2$	6.8 (h.e.)	World's best	Phys. Rev. D 93 (2016) 112008
Mo100	6914	3034	$0.68 \pm 0.05$	80	World's best	Neutrino 2018
Te130	454	25227.5	$70 \pm 14$	0.5	First direct detection	Phys. Rev. Lett. 107, 062504 (2011)

**NEW!**

**UPDATED**

\* Together with Aurora

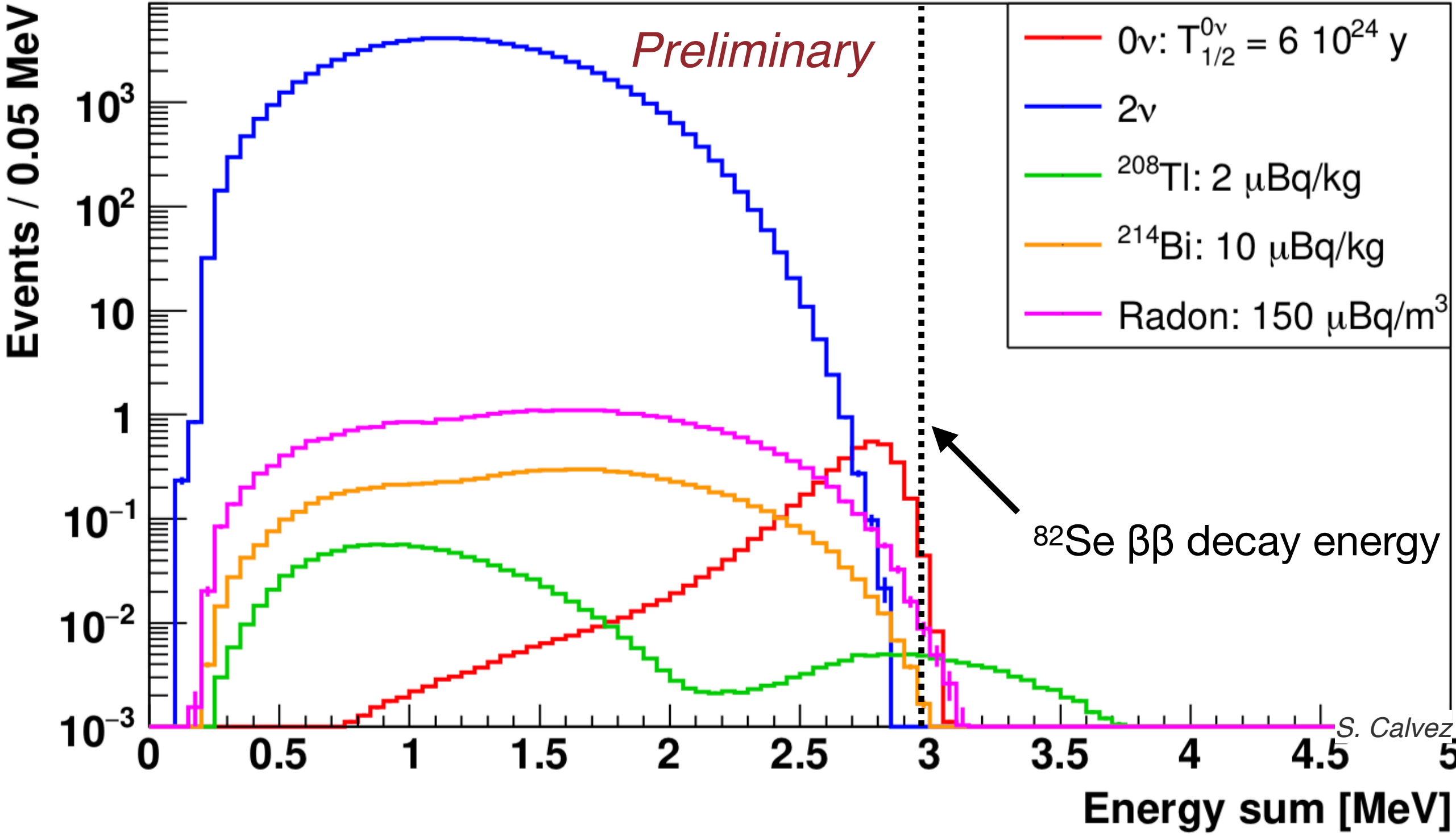
**Crucial experimental input for**

- 1) NME calculations
- 2) Ultimate background characterisation for  $0\nu$
- 3) Sensitive to exotic BSM physics (e.g. Lorentz violation,  $G_f$  time dependence, bosonic neutrinos etc)

*Taken from R Saakyan, NDM2018*

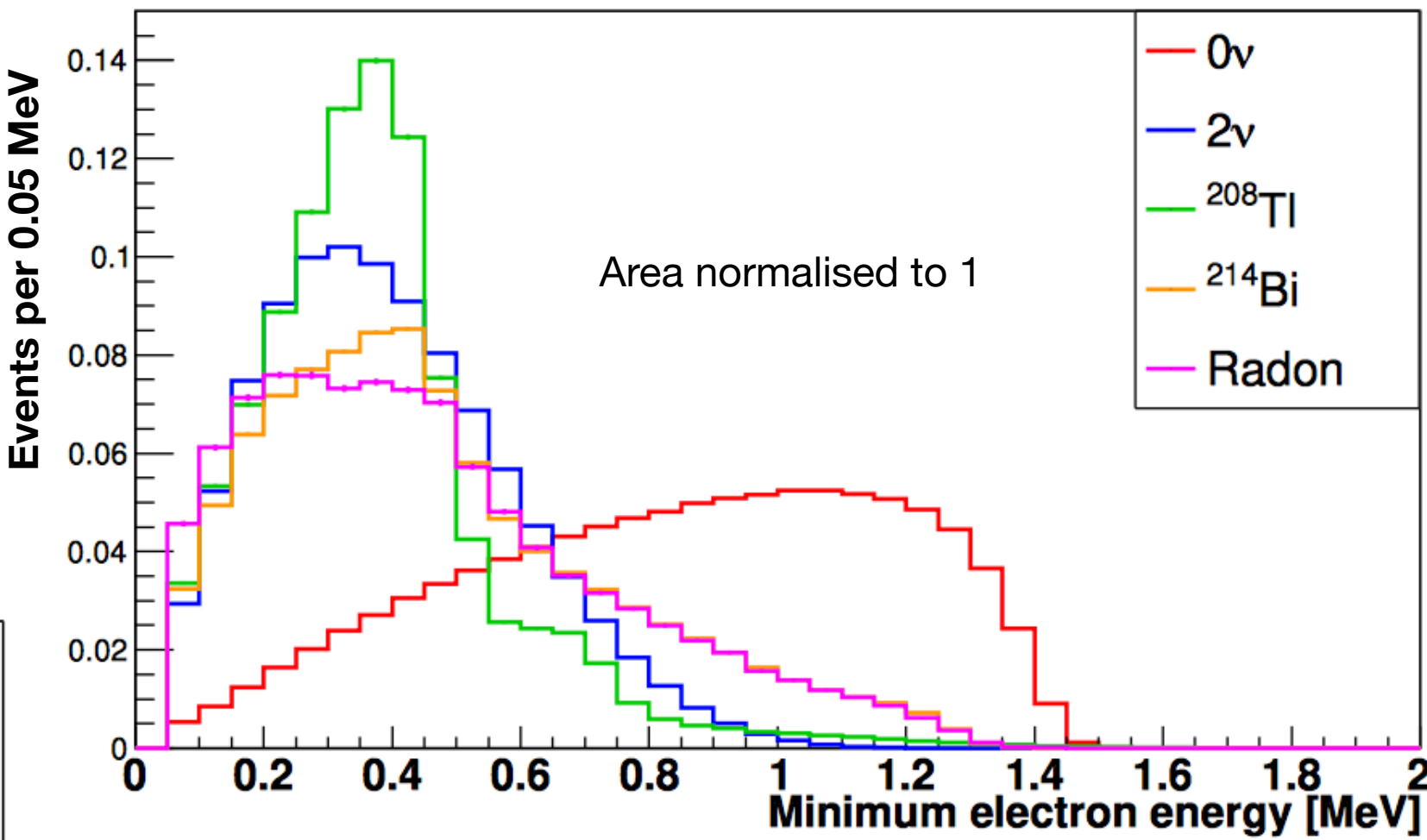
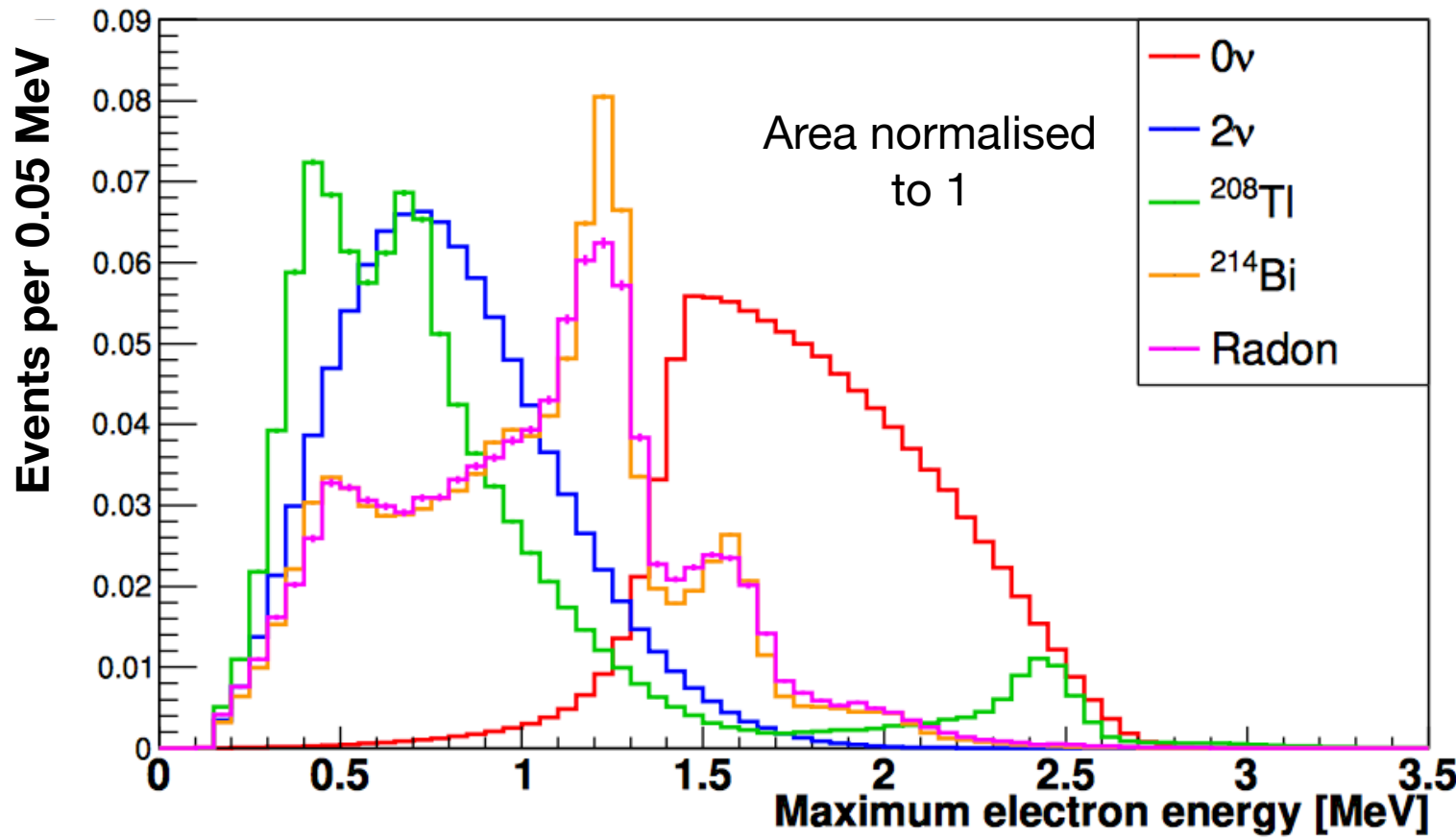
32

# Sensitivity to $0\nu\beta\beta$



**Summed 2-electron energy** is best distribution to separate signal from background

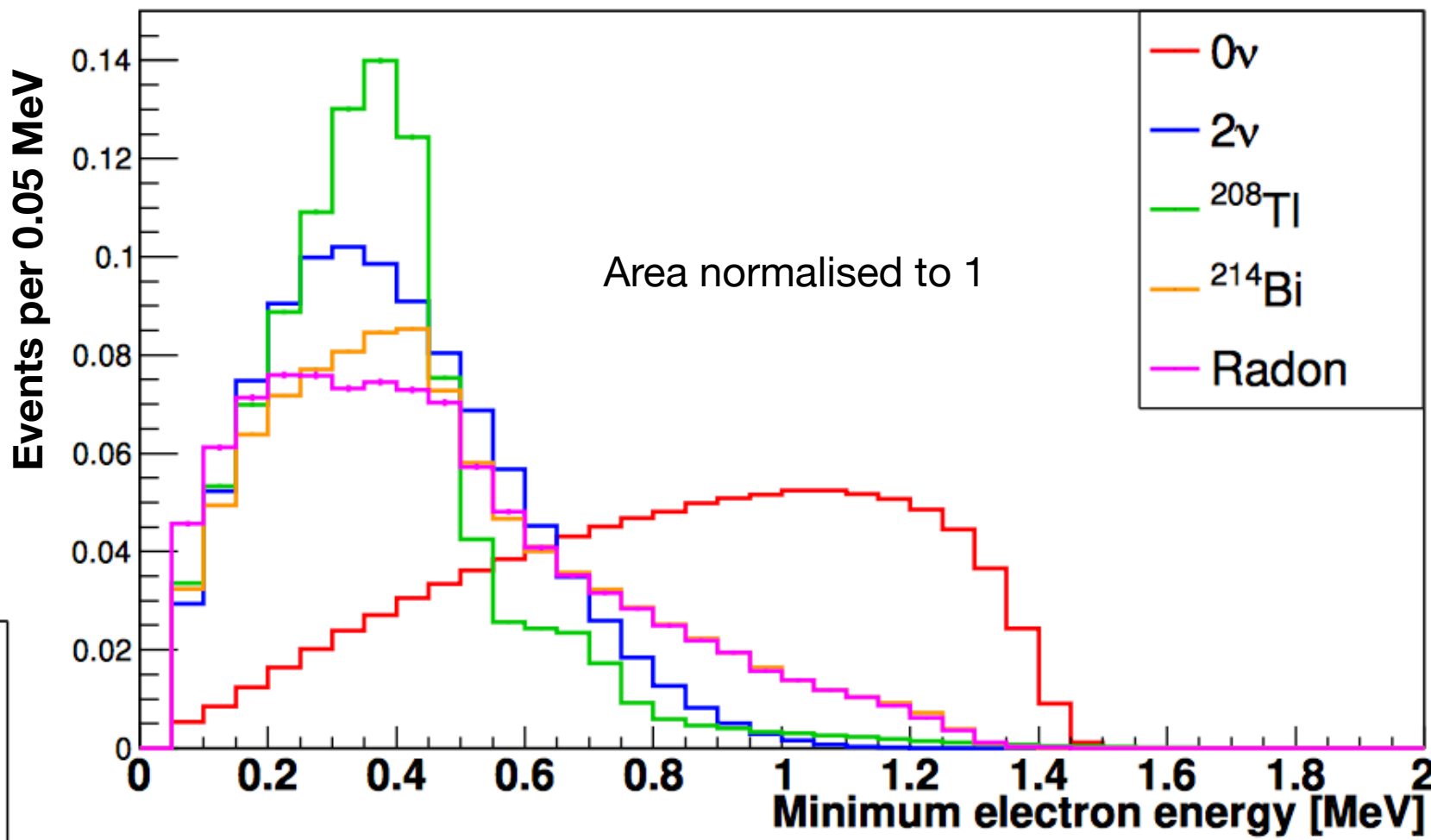
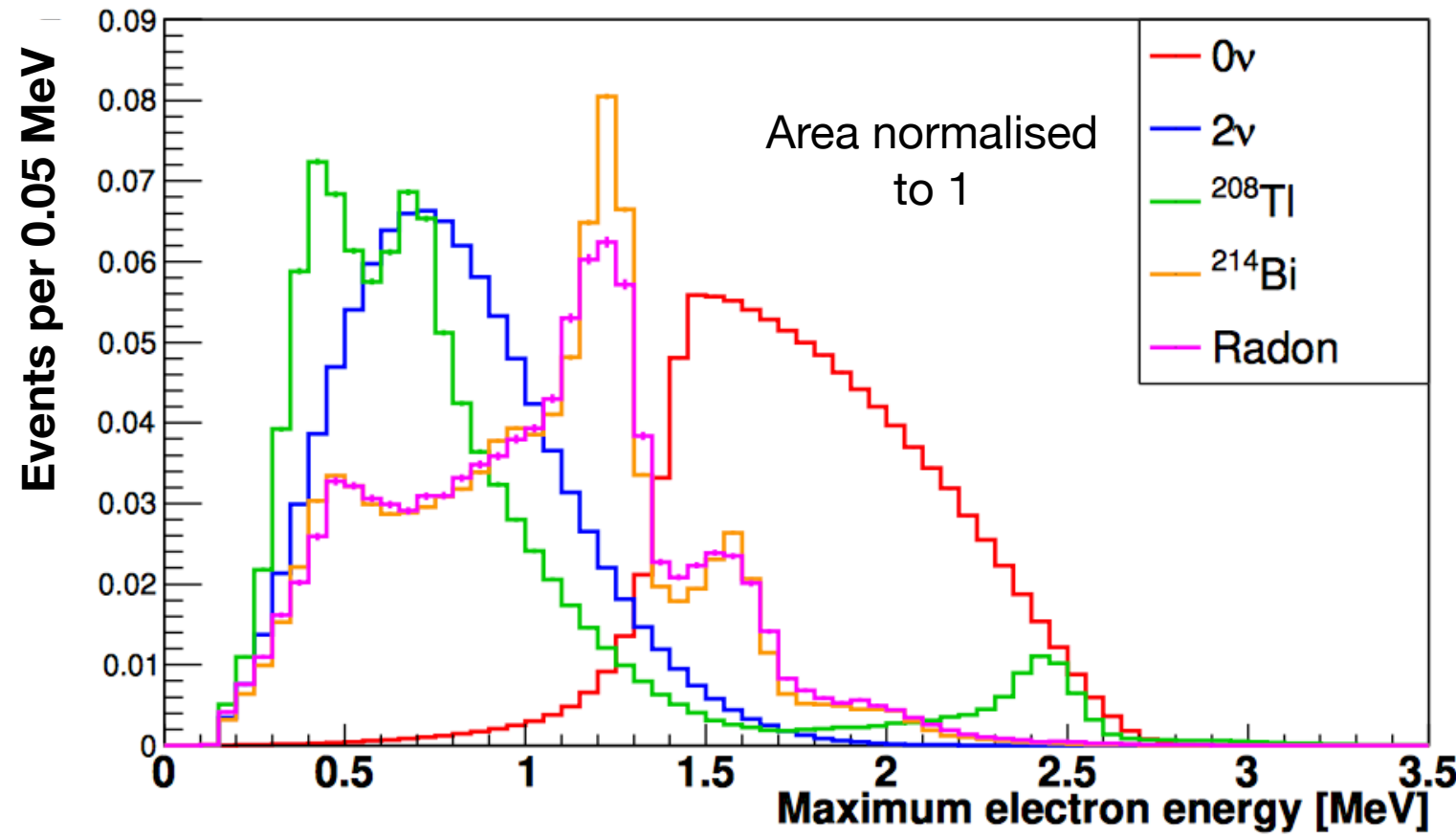
# Sensitivity to $0\nu\beta\beta$



**Summed 2-electron energy** is best distribution to separate signal from background

Using a **boosted decision tree**, we can **improve sensitivity** by including **other variables** (angle between tracks, individual electron energies, internal/external probability, vertex separation...) *(approx 10% improvement)*

# Sensitivity to $0\nu\beta\beta$



**Summed 2-electron energy** is best distribution to separate signal from background

Using a **boosted decision tree**, we can **improve sensitivity** by including **other variables** (angle between tracks, individual electron energies, internal/external probability, vertex separation...) (approx 10% improvement)

$T_{1/2} > 5.85 \times 10^{24}$  years (90% C.L.)  
 For 7kg of  $^{82}\text{Se}$  (demonstrator) and 2.5 years' exposure

# ELLIE - Embedded Laser/LED Light-Injection Entity

- **UK**/Lisbon system providing a wealth of detector info
- Aim is to **minimise radon** ingress when source is deployed
- Now **deployed and operational!**

## TELLIE: Timing and Monitoring

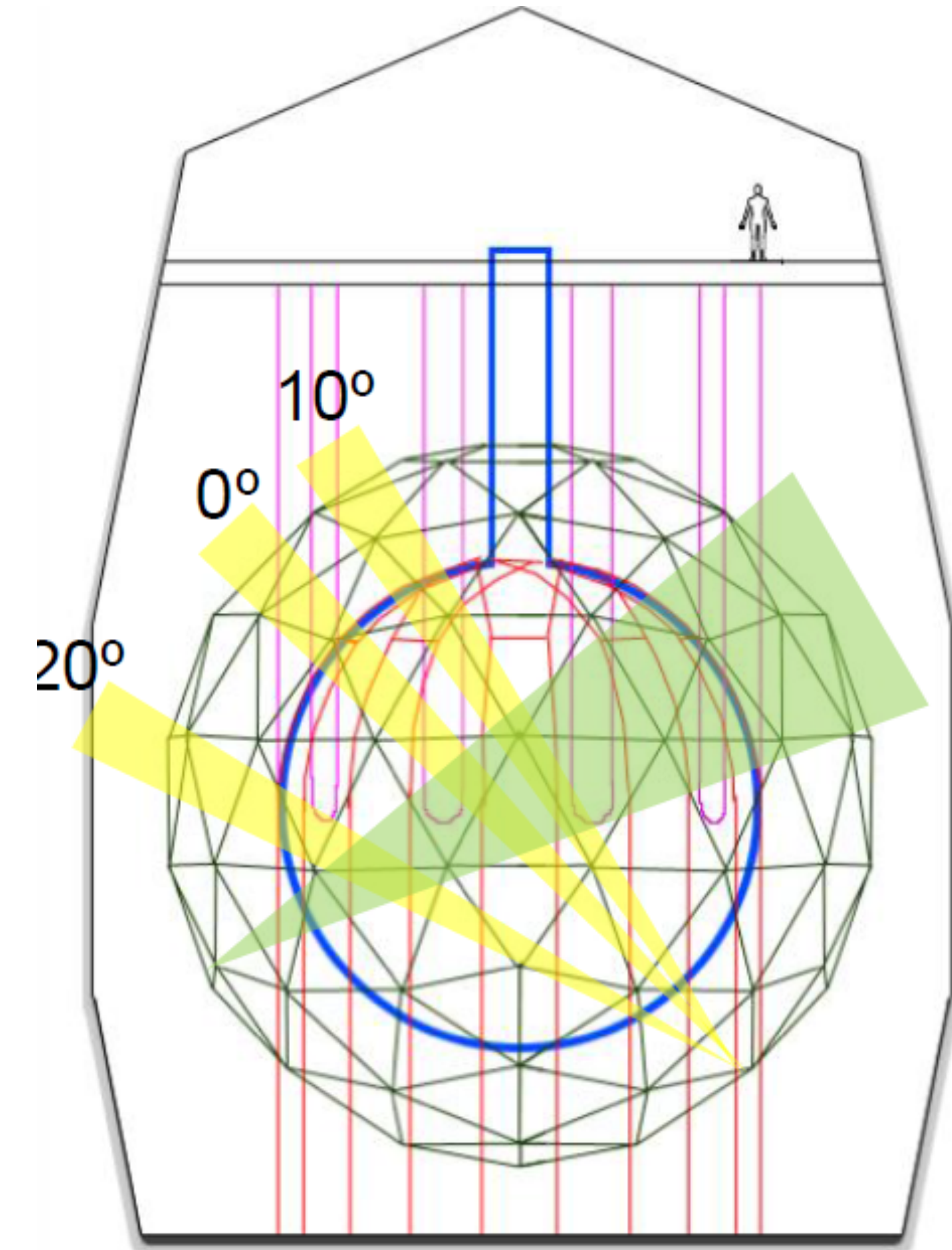
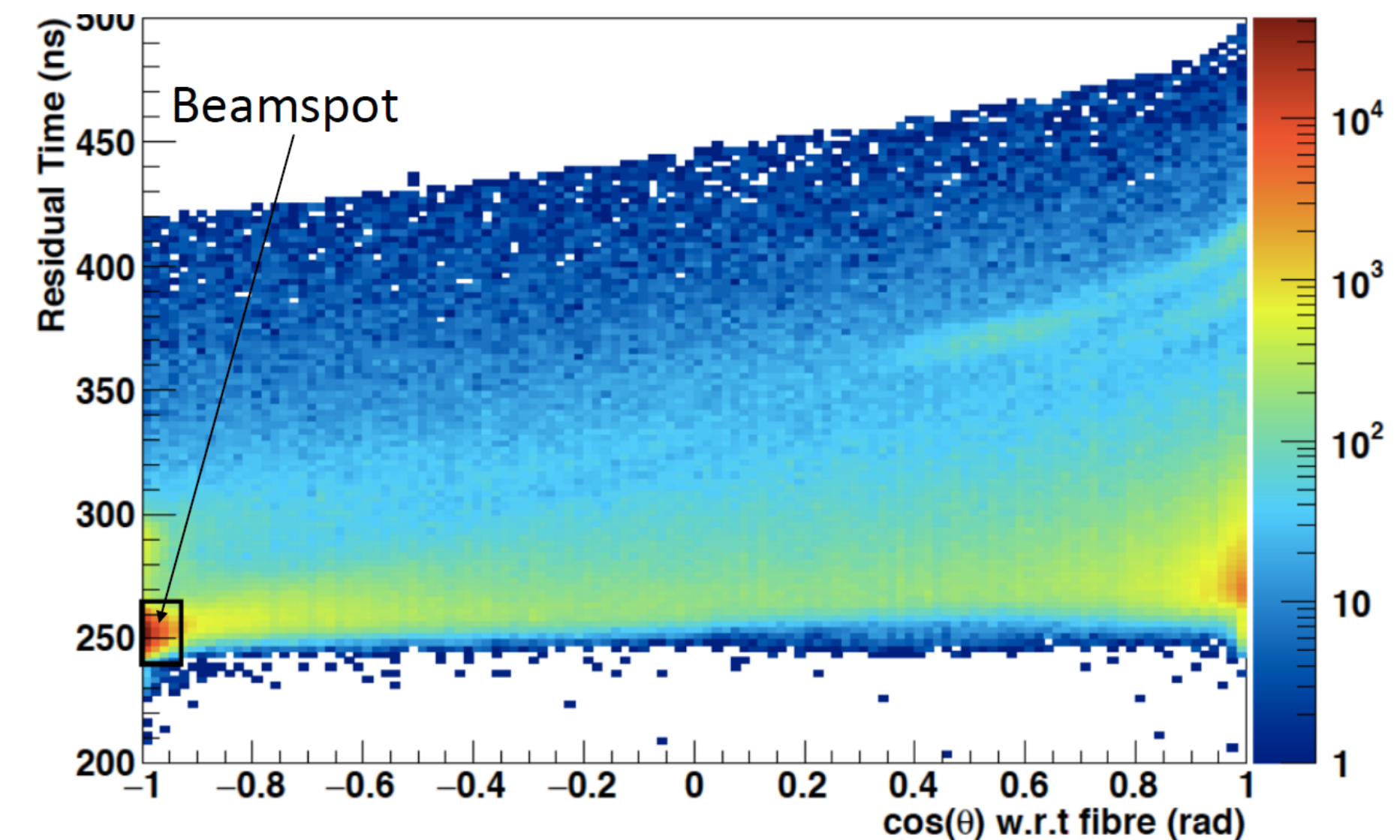
90 wide LED beams @520nm, aimed at the centre

## SMELLIE: Scattering module

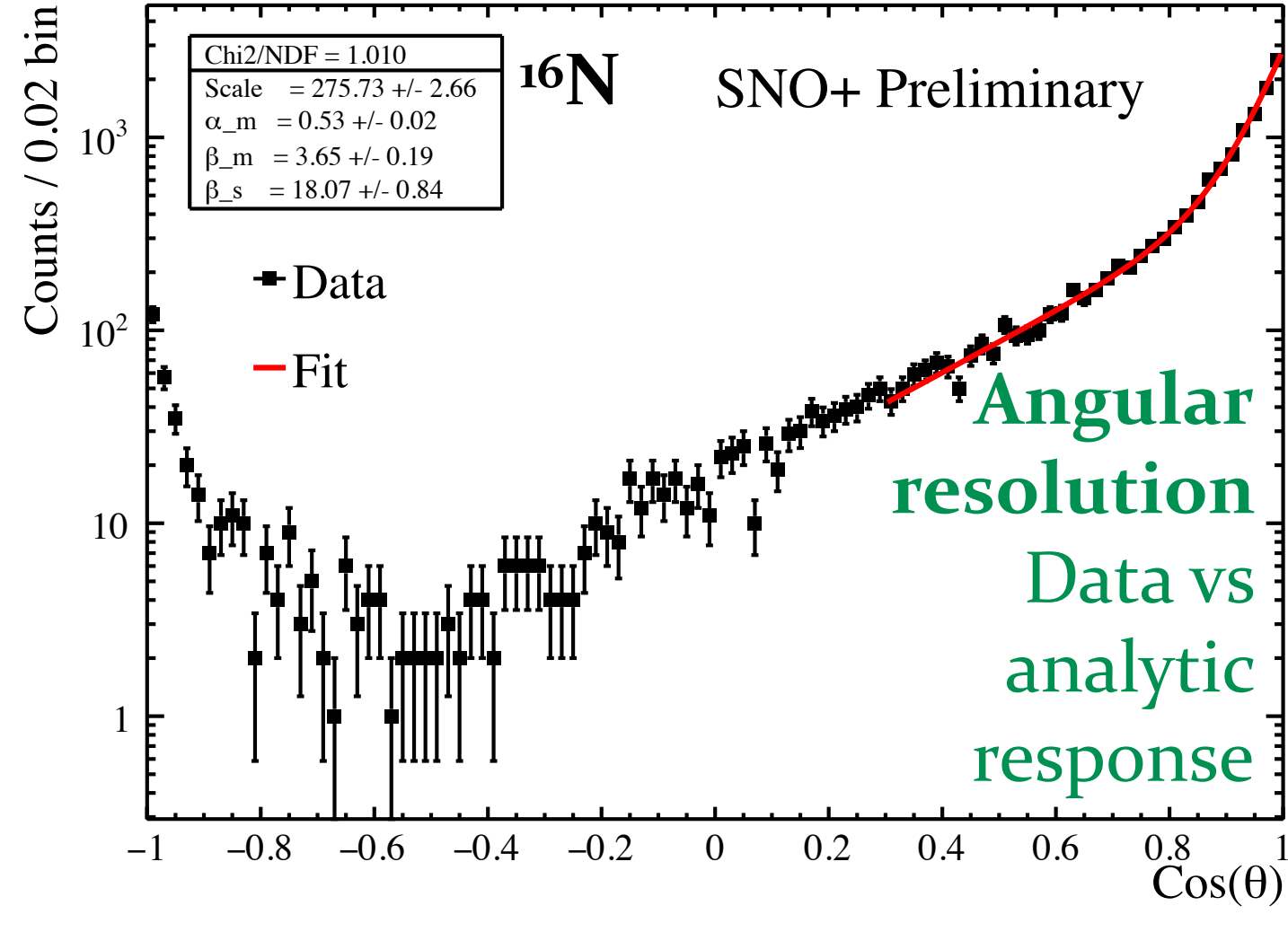
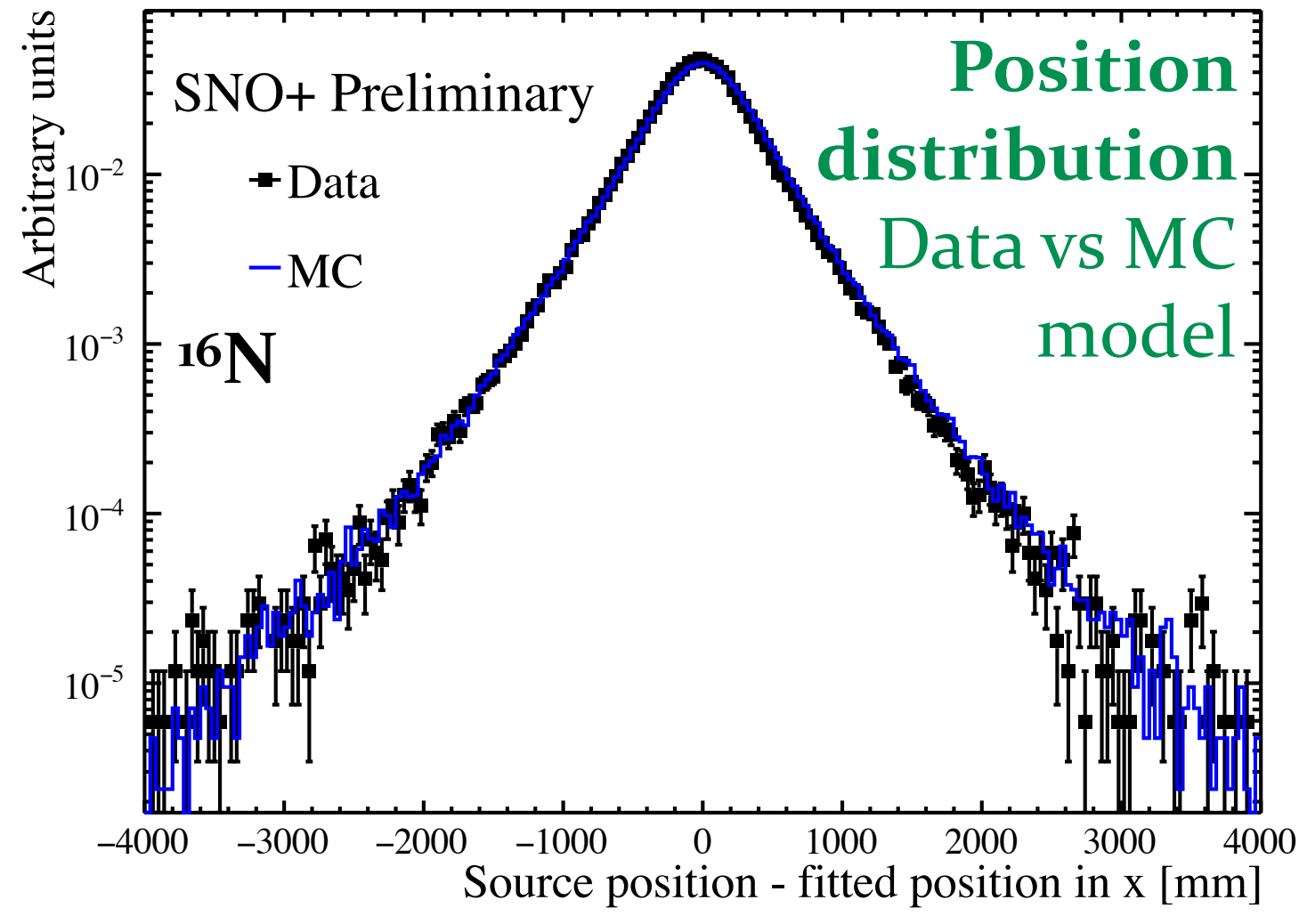
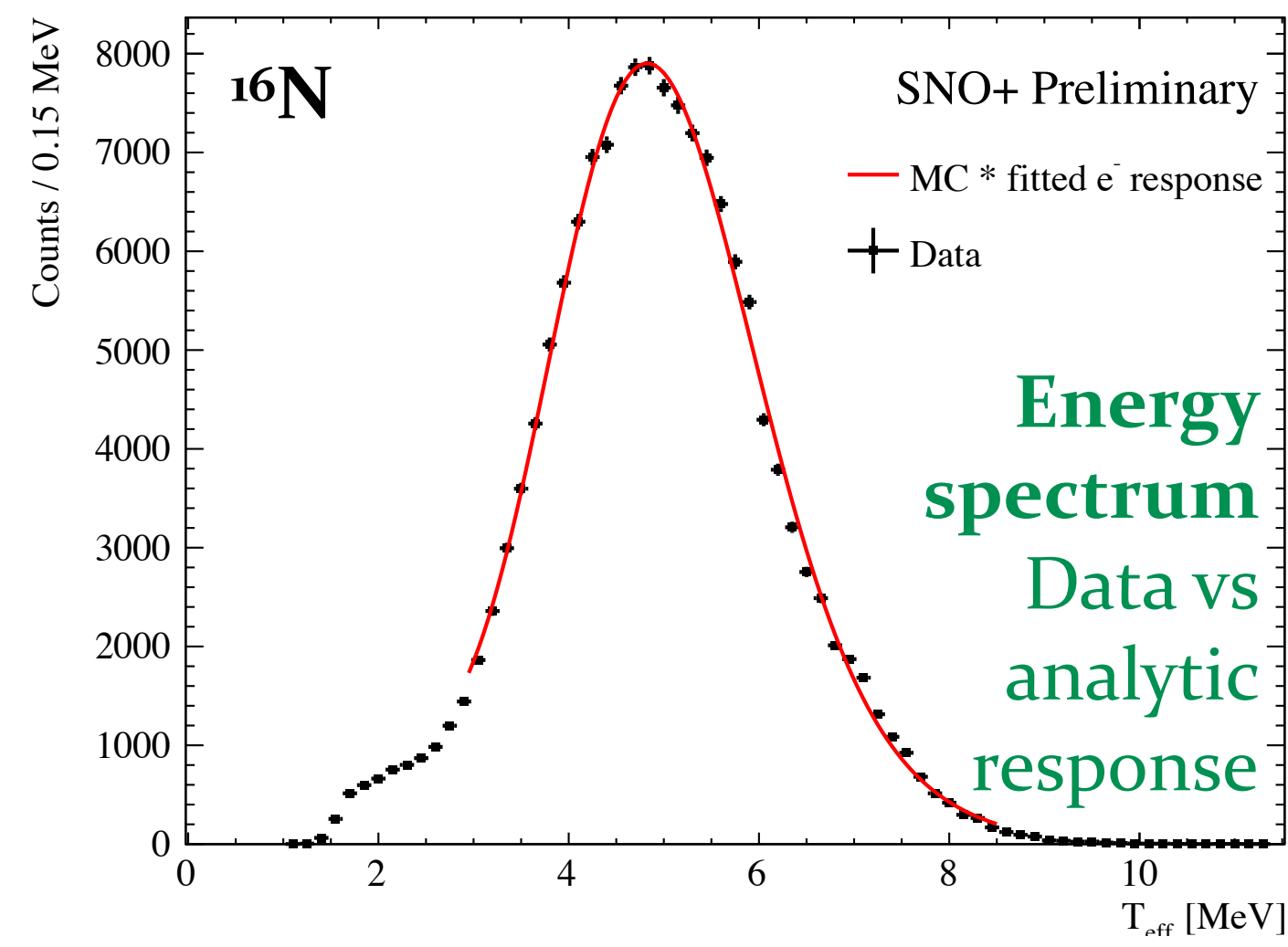
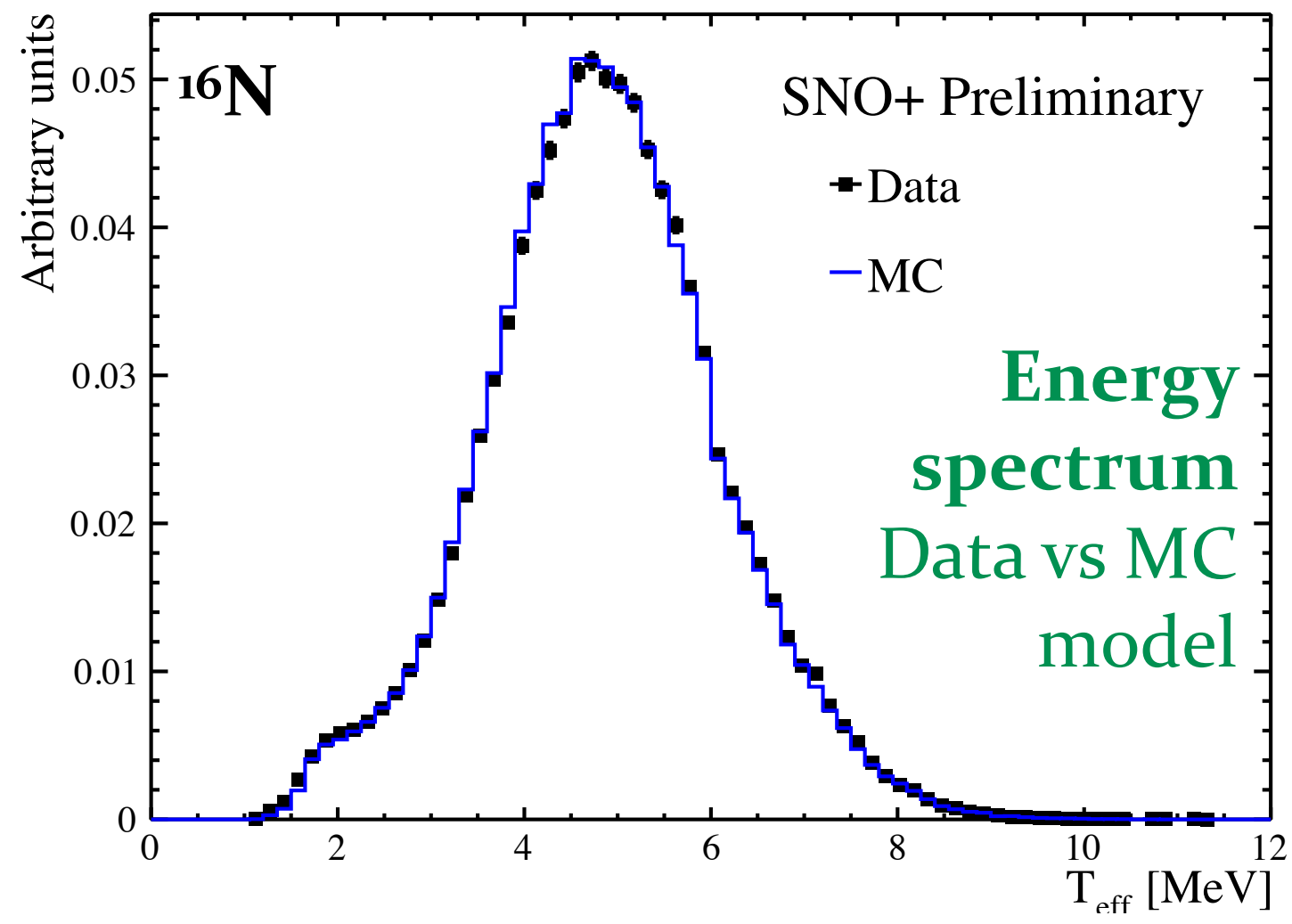
5 narrow laser beams injection points, 3 angles @ 375, 405, 495nm and 400-700nm

## AMELLIE: Attenuation module

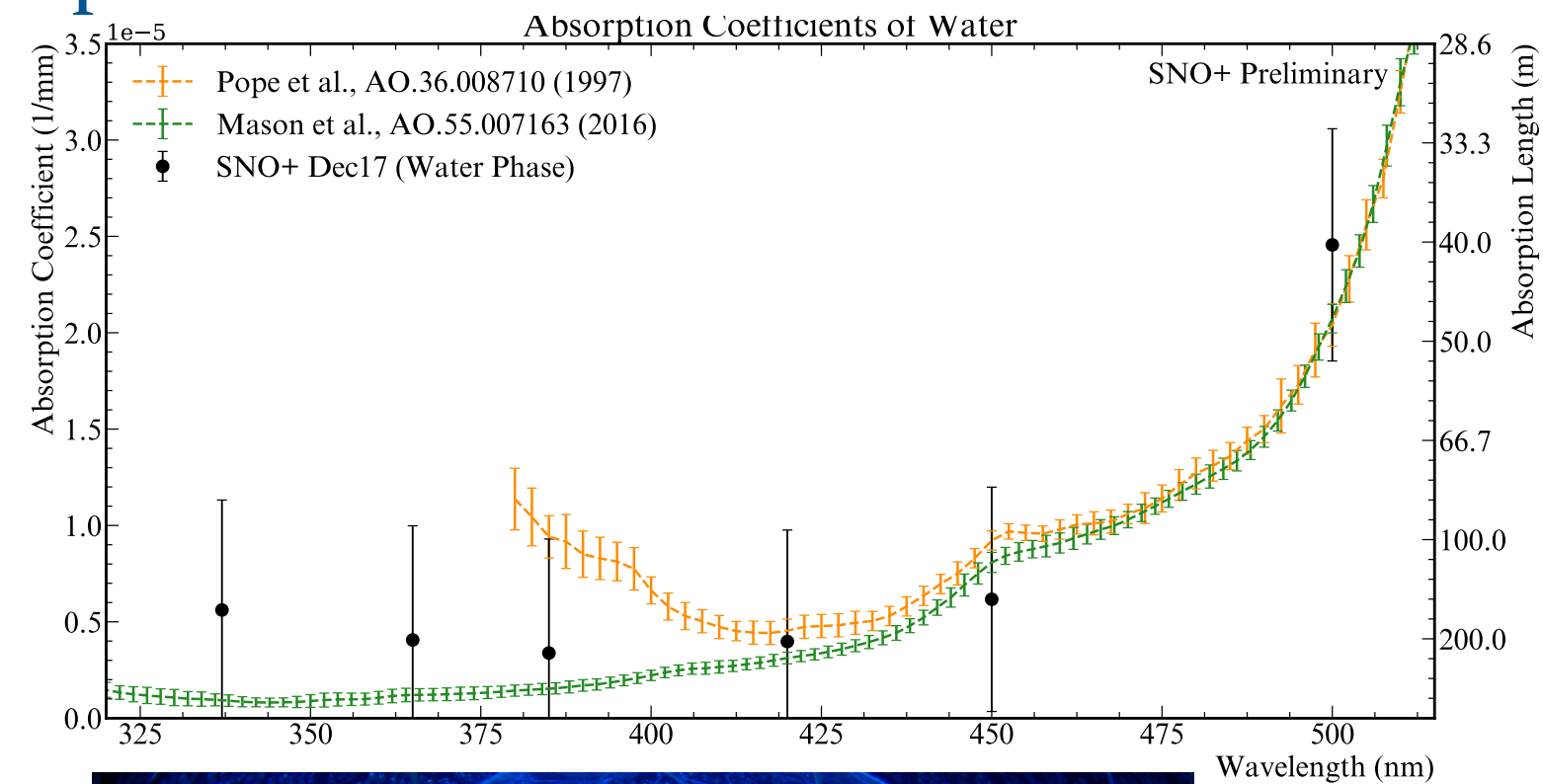
4 narrow LED beams at 0 and 20 degrees (wavelength TBD)



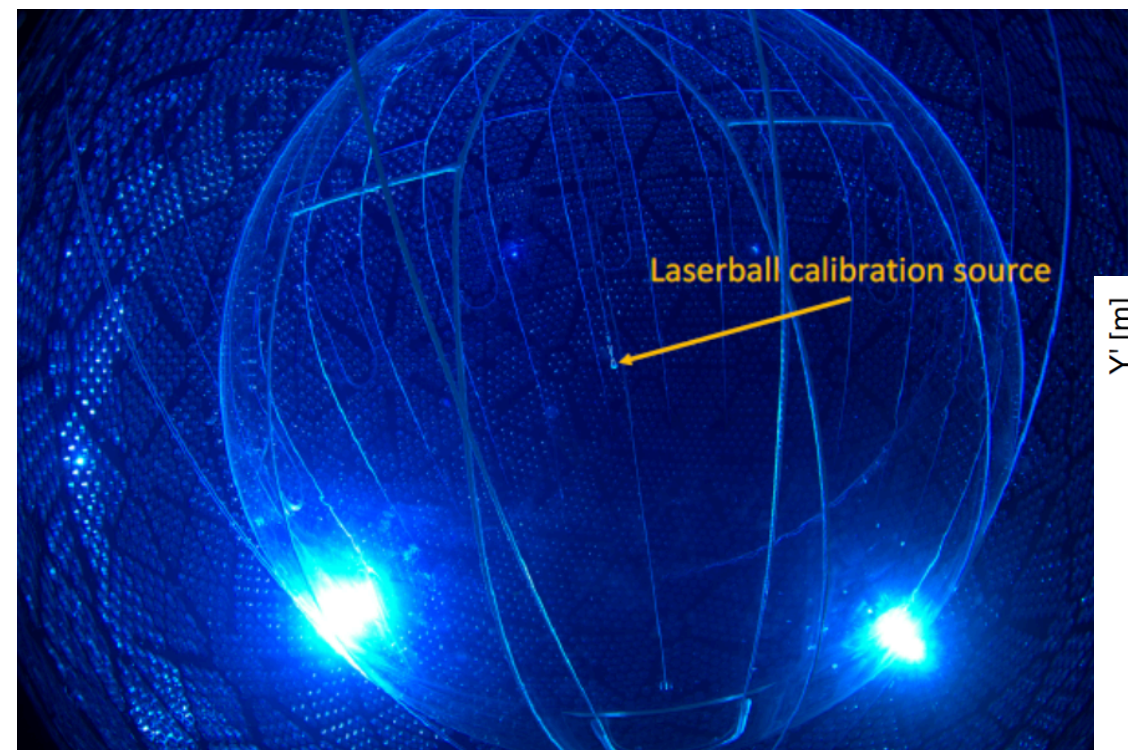
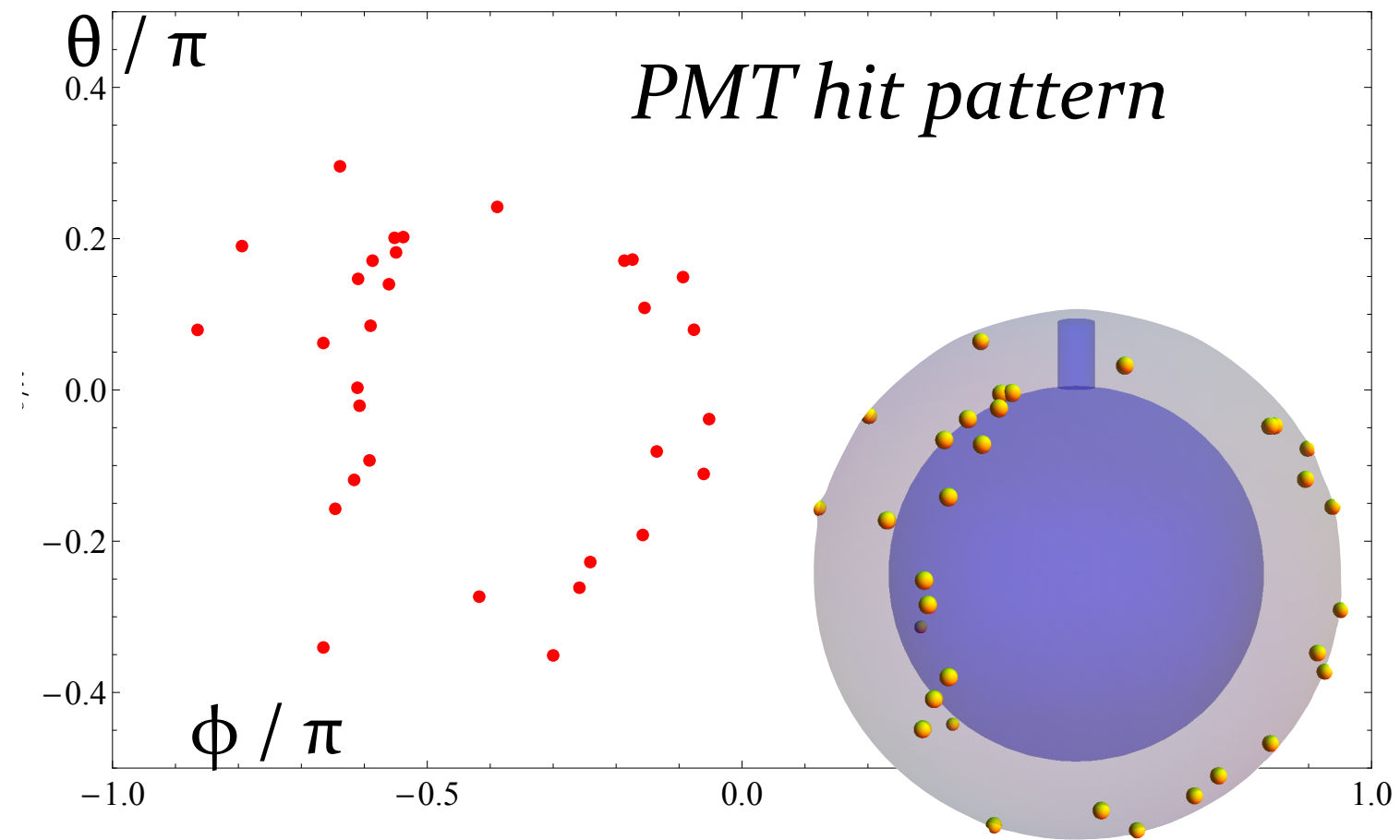




## Full internal scan with laserball, photos from underwater cameras

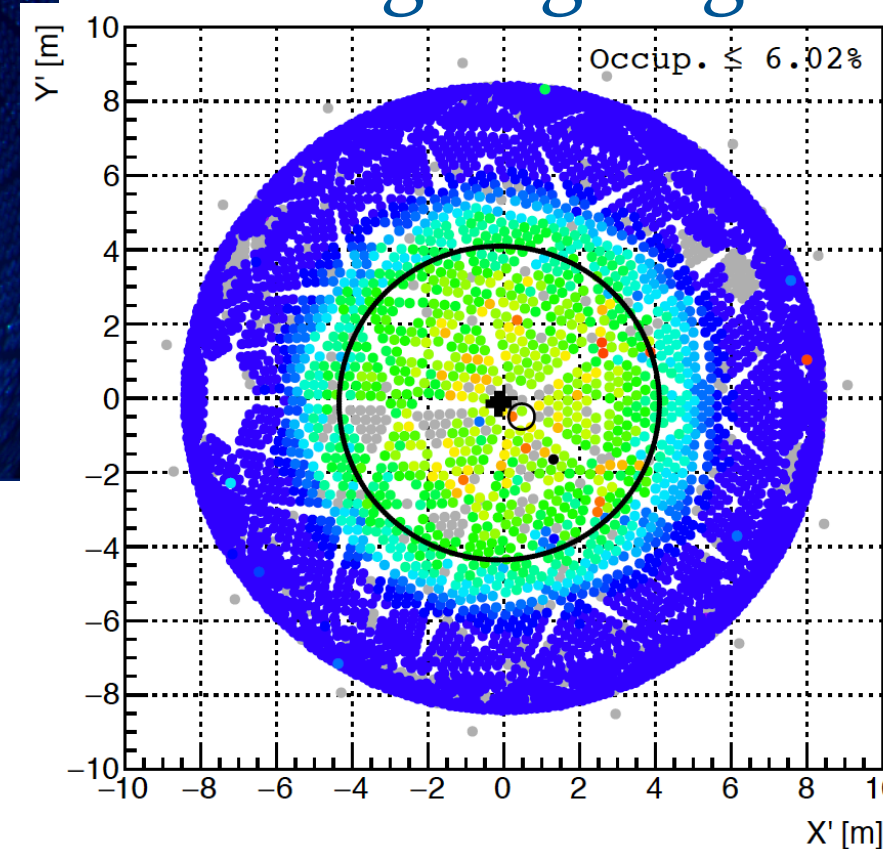


## Cherenkov source deployed

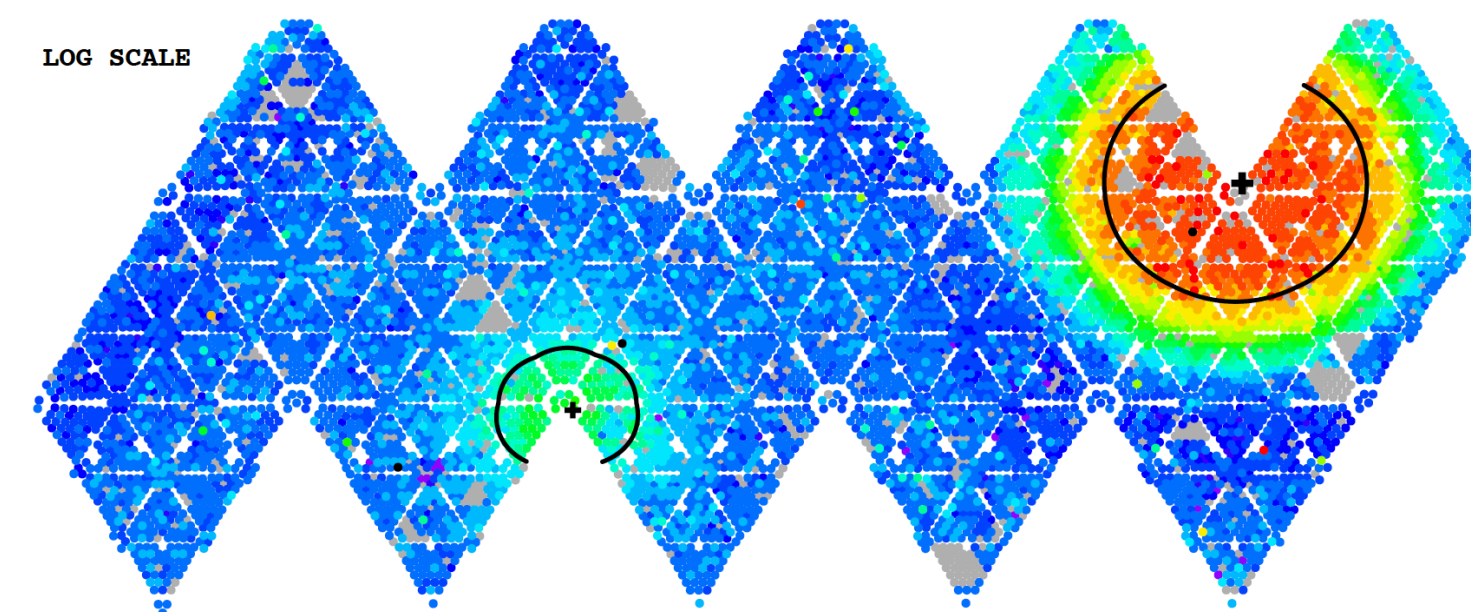


Poster #73  
*Calibration of the SNO+ Detector with a Light Diffusing Source in the Water Phase*  
 A. Inacio

## Fibre data taking ongoing

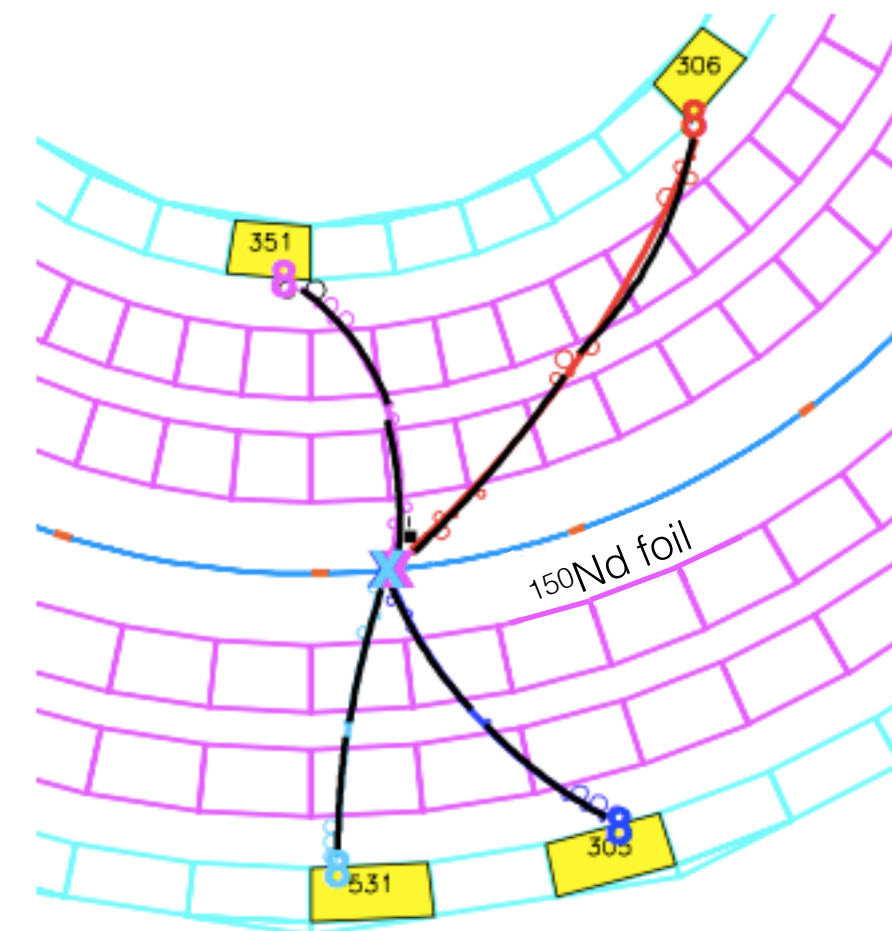


Poster #83  
*SMELLIE: A Laser Calibration System for SNO+*  
 E. Turner



# NEMO-3 - quadruple beta decay

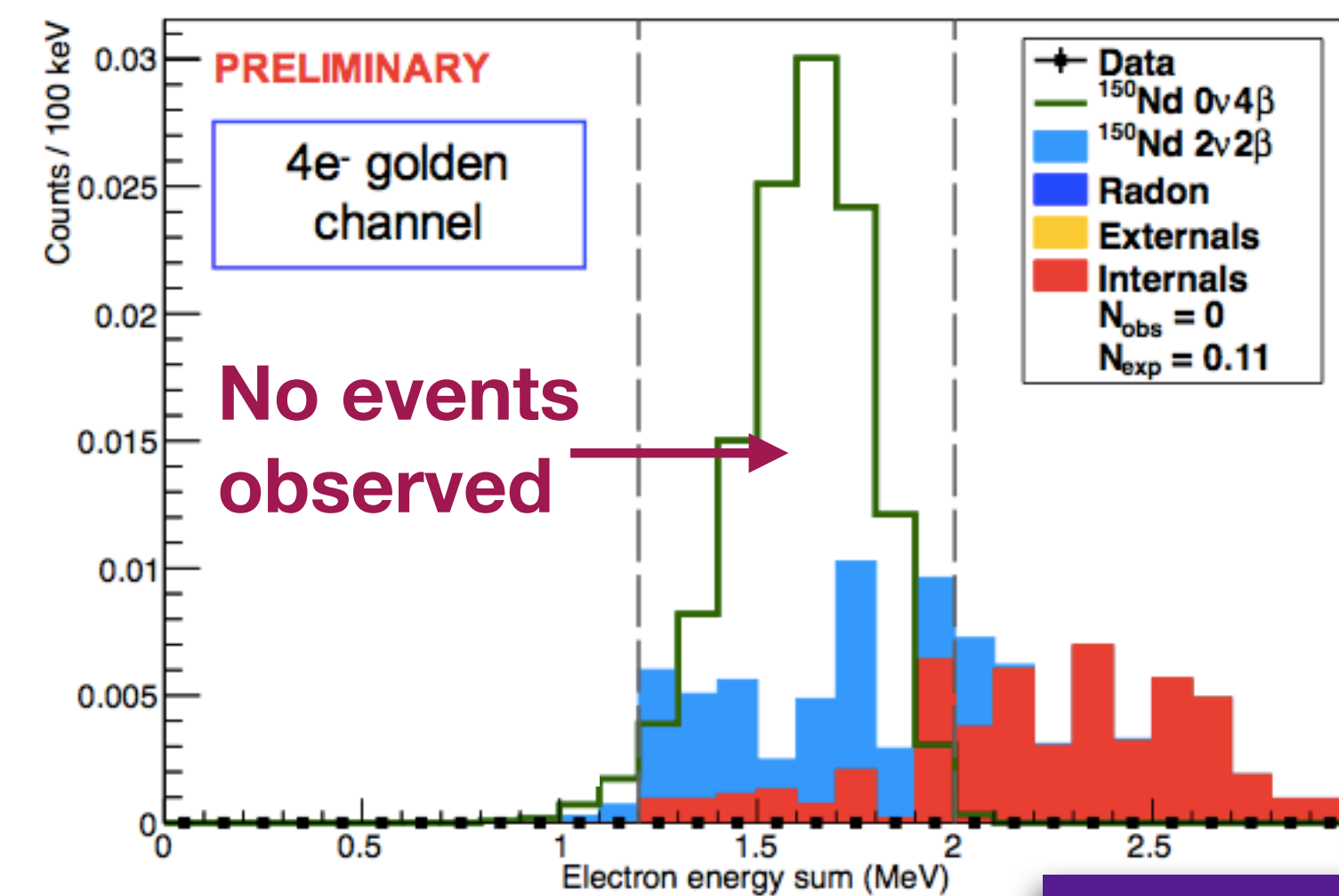
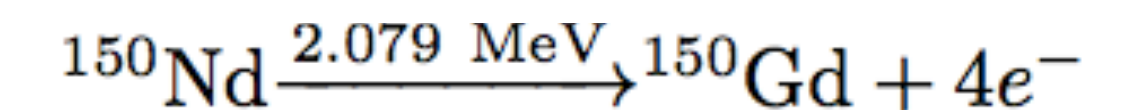
- $2\nu\beta\beta$  measurements and  $0\nu\beta\beta$  limits for several isotopes
  - $^{100}\text{Mo}$  (Phys. Rev. Lett. 95, 182302)
  - $^{48}\text{Ca}$  (Phys. Rev. D 93, 112008)
  - $^{82}\text{Se}$  (Eur. Phys. J. C (2018) 78: 821)
  - $^{150}\text{Nd}$  (Phys. Rev. D 94, 072003)
  - $^{116}\text{Cd}$  (Phys. Rev. D 95, 012007)
  - $^{130}\text{Te}$  (Phys. Rev. Lett. 107, 062504)
  - $^{96}\text{Zr}$  (Nucl.Phys.A847:168-179)
- **Quadruple  $\beta$  decay** (Phys. Rev. Lett. 119, 041801)



$0\nu 4\beta$  decays would violate lepton number, but could occur even if neutrinos are **Dirac** fermions

*Heeck and Rodejohann 2013*

Allowed for 3 isotopes, including



NEMO architecture helps us identify these topologies

$T_{1/2} > 2.6 \times 10^{21} \text{ yr}$  (90%CL)

# Low background strategy: **reduce**, **remove**, **reject**

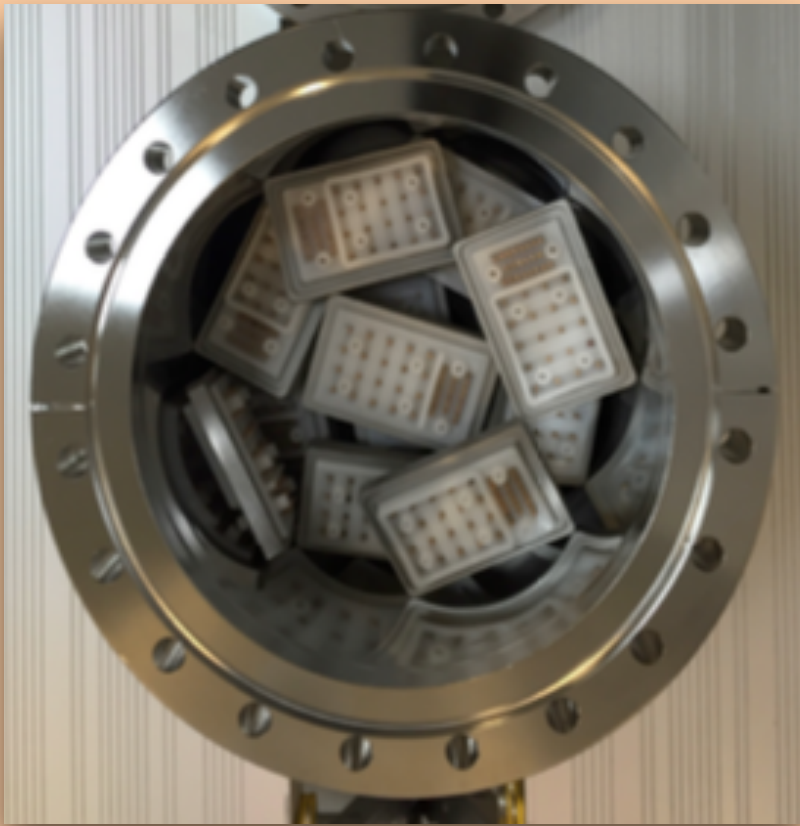
**Radon 222** (from U decay chain): target activity **150  $\mu\text{Bq} / \text{m}^3$**

~ 30 times lower than NEMO-3

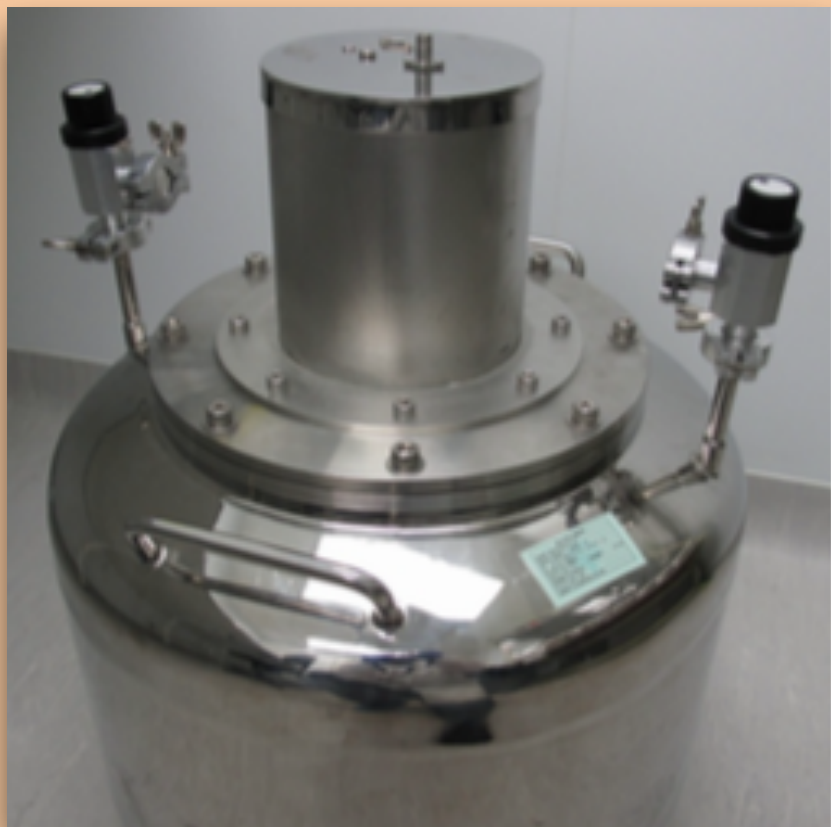
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## **Reduce** radon contamination with radio-pure components



**Emanation chamber** lets us measure activity of **tracker components** and materials: select only the most **radio-pure**

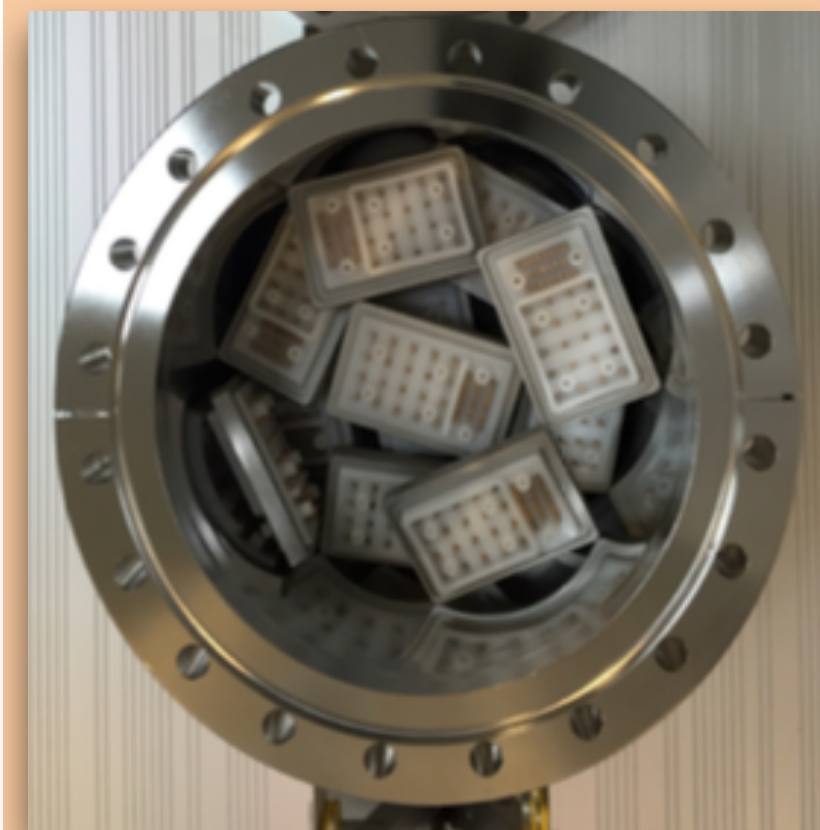


70 litre **electrostatic detector** sensitive down to 0.09mBq

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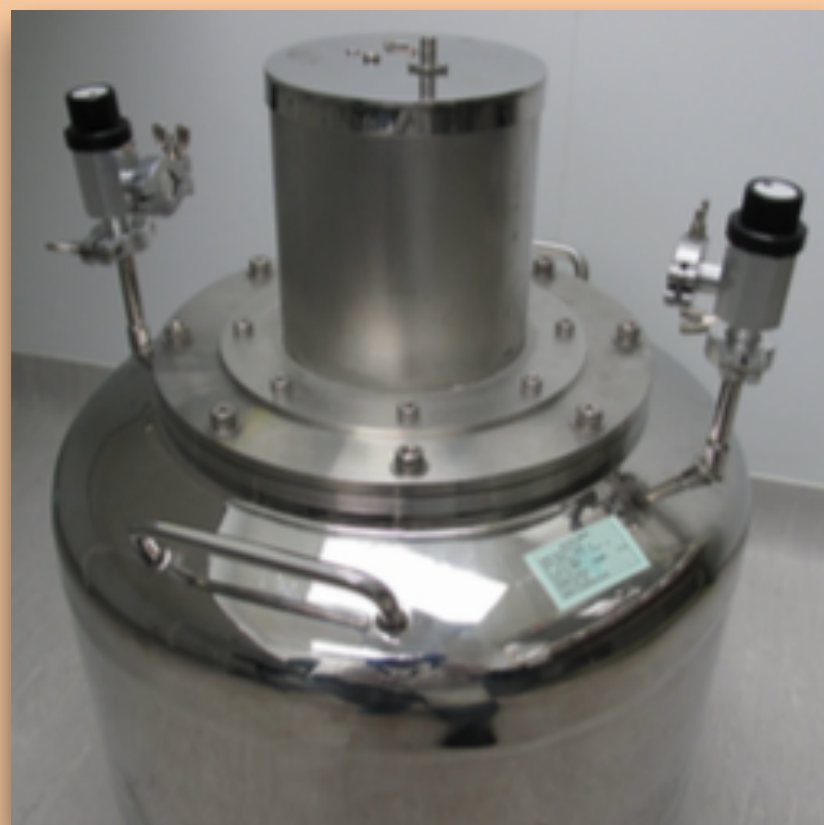
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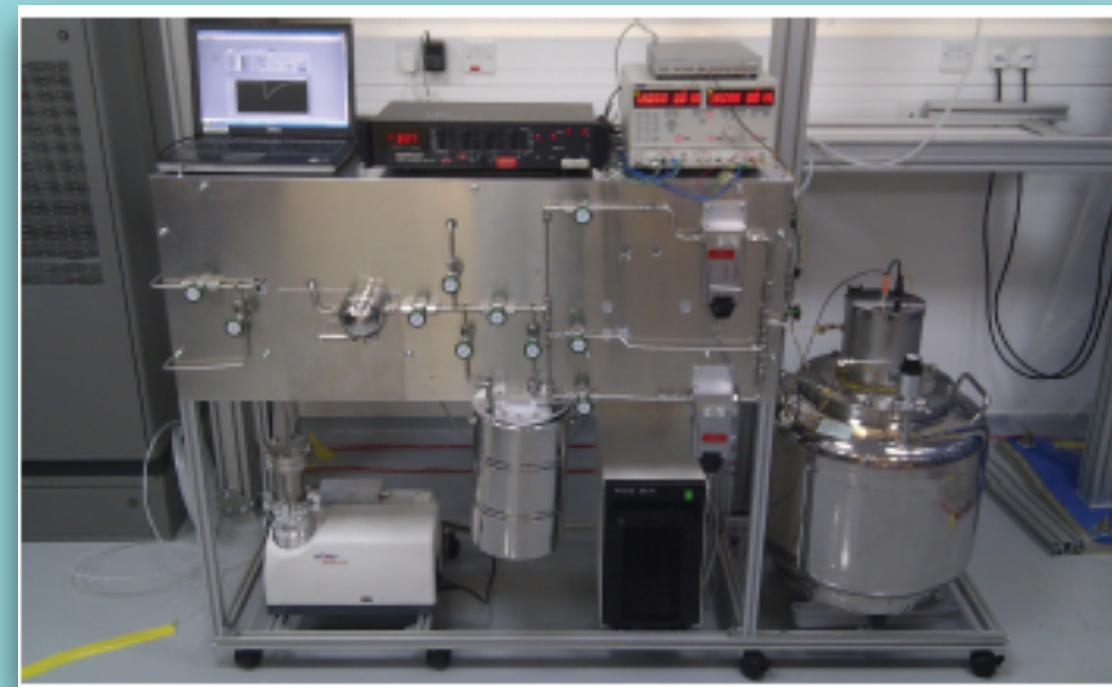
70 litre **electrostatic detector** sensitive down to 0.09mBq



## **Remove** radon from tracker gas (95% helium, 1% argon, 4% ethanol)

Remove Rn with **cold carbon trap**

- **He:**  $10^{10}$  x suppression - completely **clean**
- **N<sub>2</sub>:** 20x purification - 20  $\mu\text{Bq}/\text{m}^3$



**Radon concentration line** lets us measure the low activities in the tracker

**Measured** activity:  $2.7 \pm 0.3 \text{ mBq} / \text{m}^3$

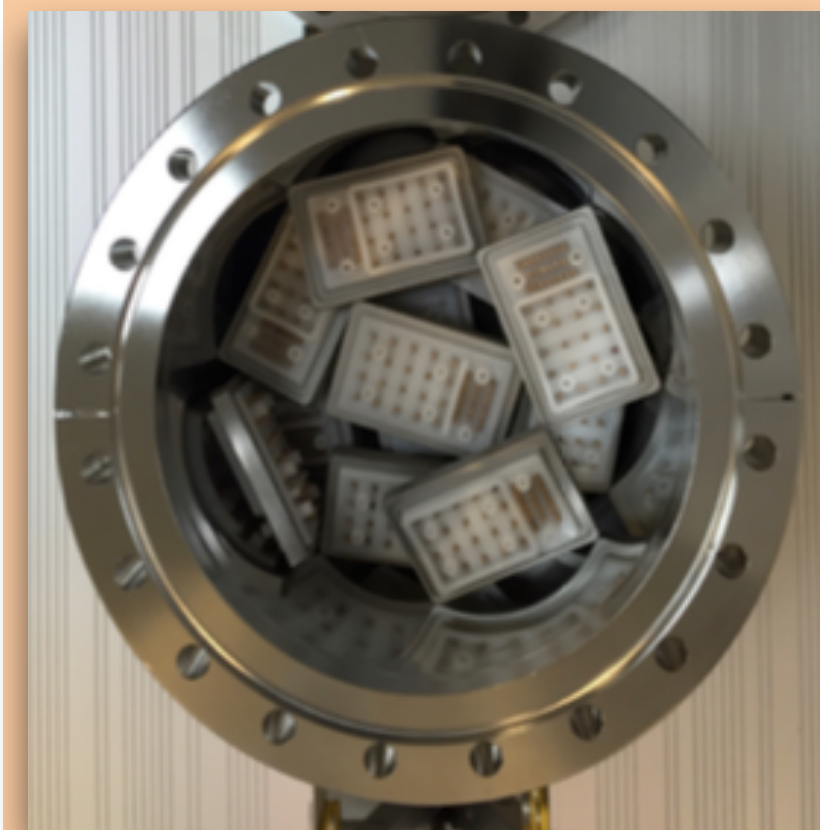
**Flush** with He:  $2 \text{ m}^3 / \text{hour}$

**Resulting** activity:  $0.15 \text{ mBq} / \text{m}^3$

# Low background strategy: reduce, remove, reject

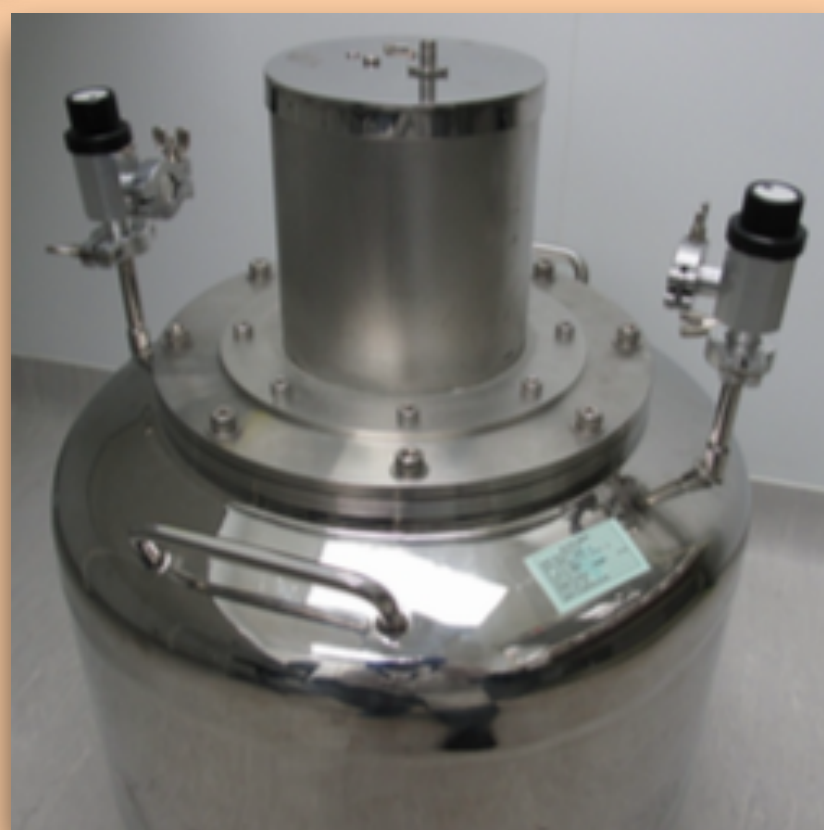
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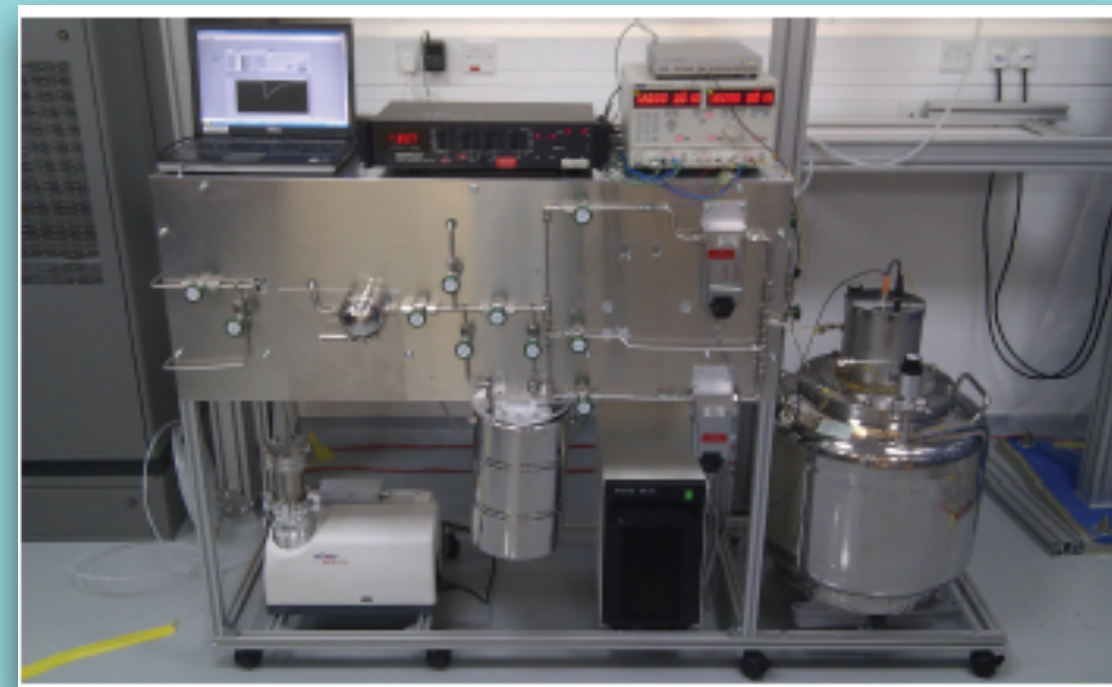
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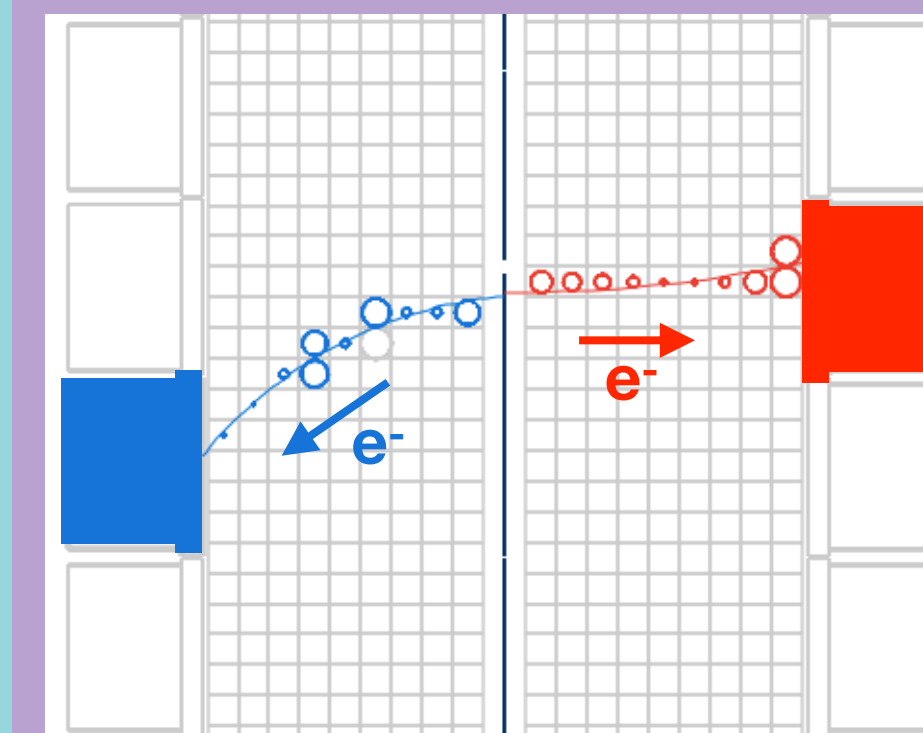
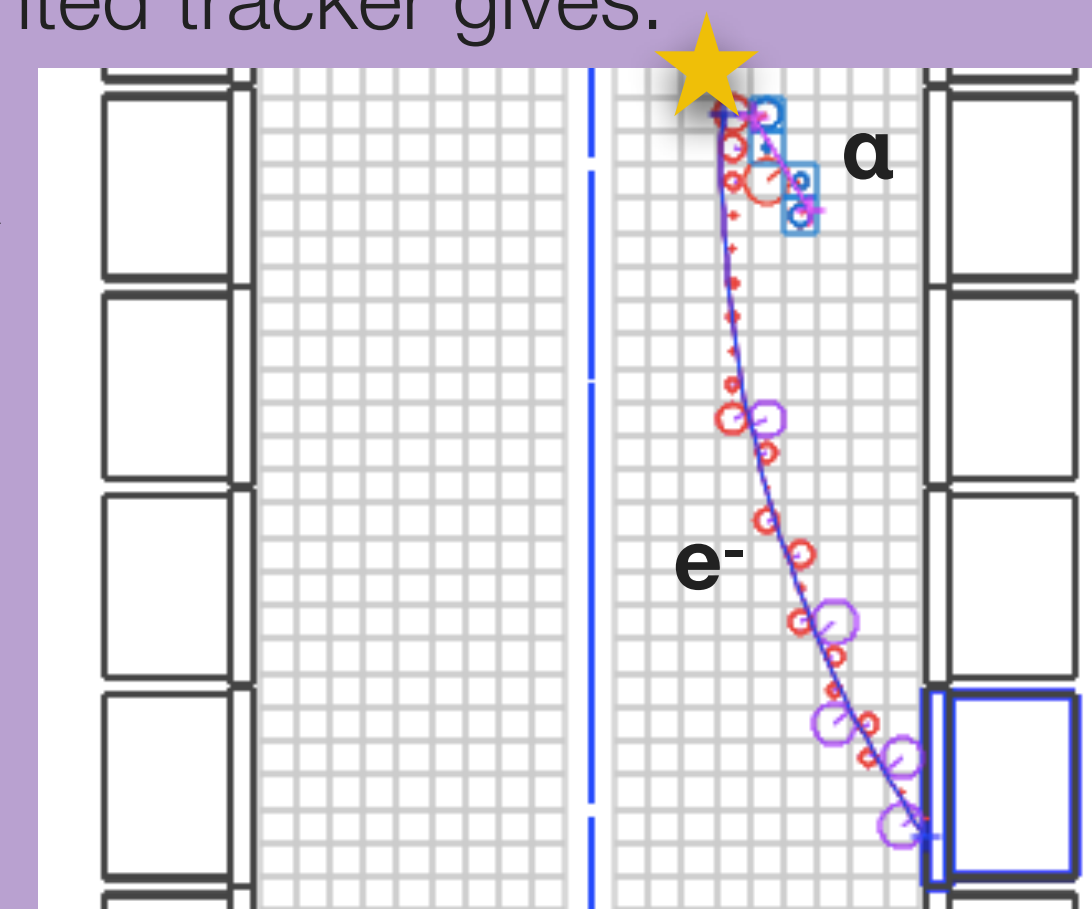
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**Flush** with He:  $2 \text{ m}^3 / \text{hour}$   
**Resulting** activity:  $0.15 \text{ mBq} / \text{m}^3$

## Reject background events with topological and timing cuts

Fully-instrumented tracker gives:

- Event vertex
- Particle ID
- Timings  $\rightarrow$  direction of travel

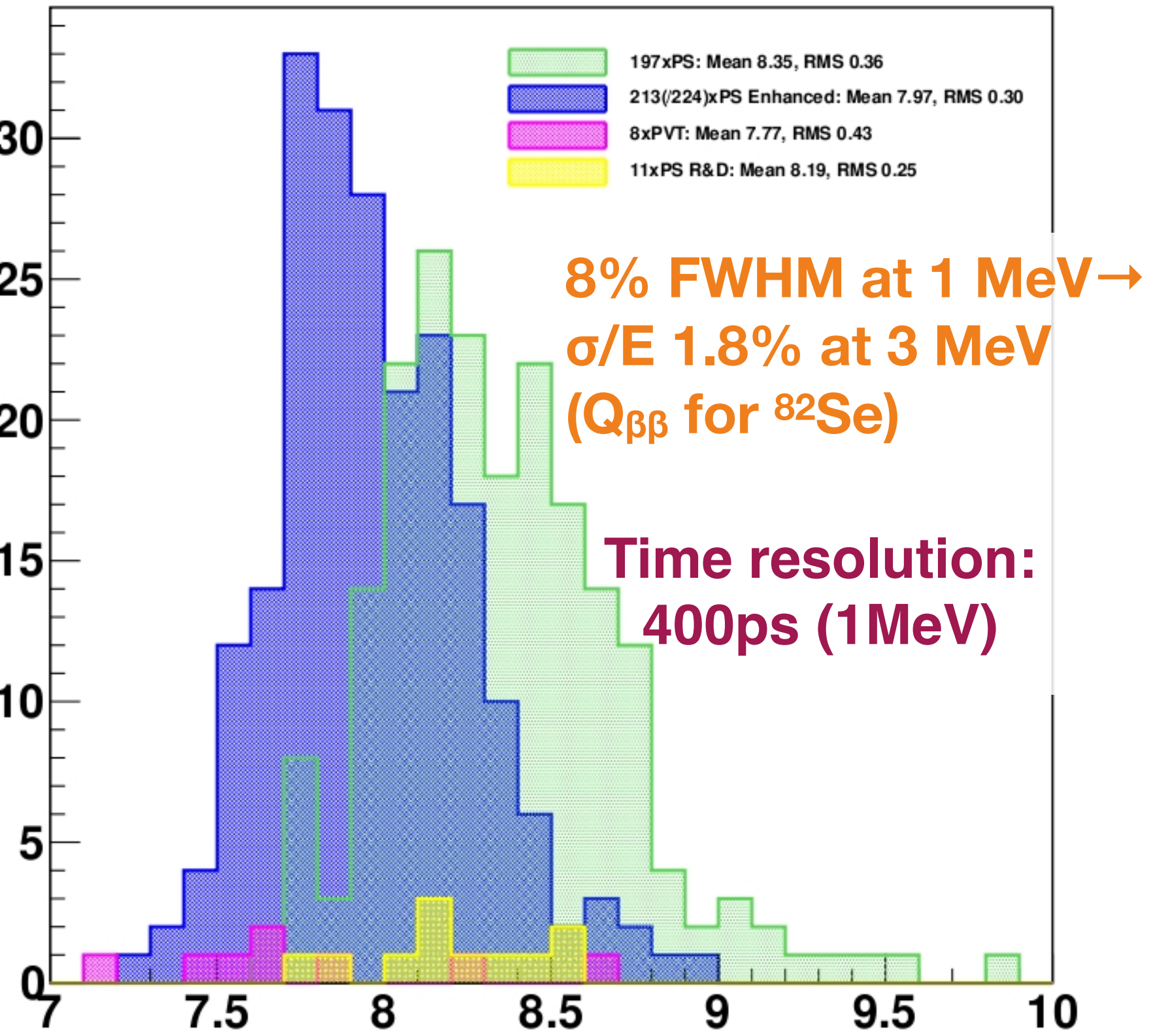


Reject non- $\beta\beta$  topologies at analysis time

# Calorimeter development

Main calorimeter walls: 520 optical modules  
With side, top and bottoms: 712 modules total

*Nucl. Inst. Meth. A 868, 98-108 (2017)*

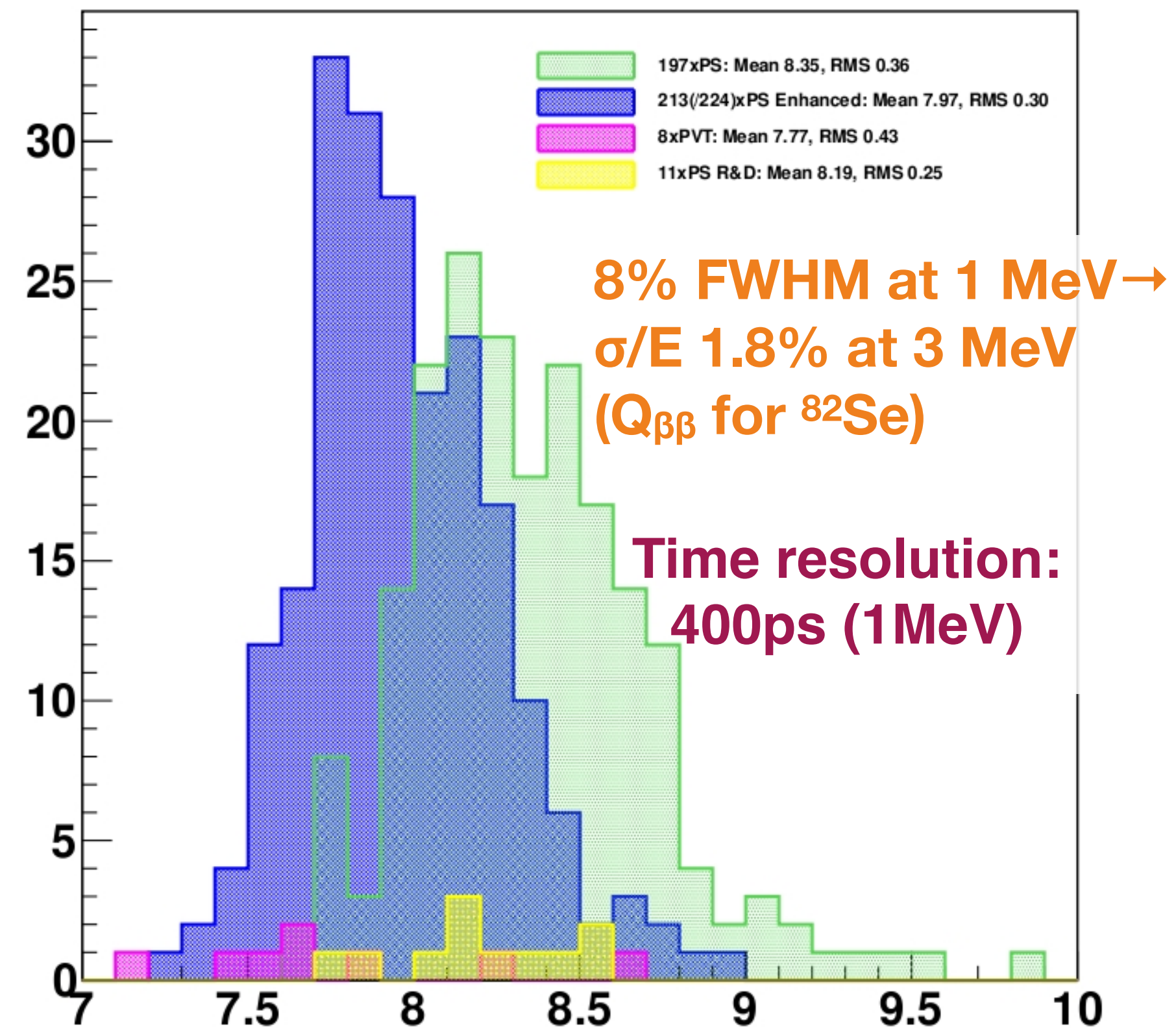




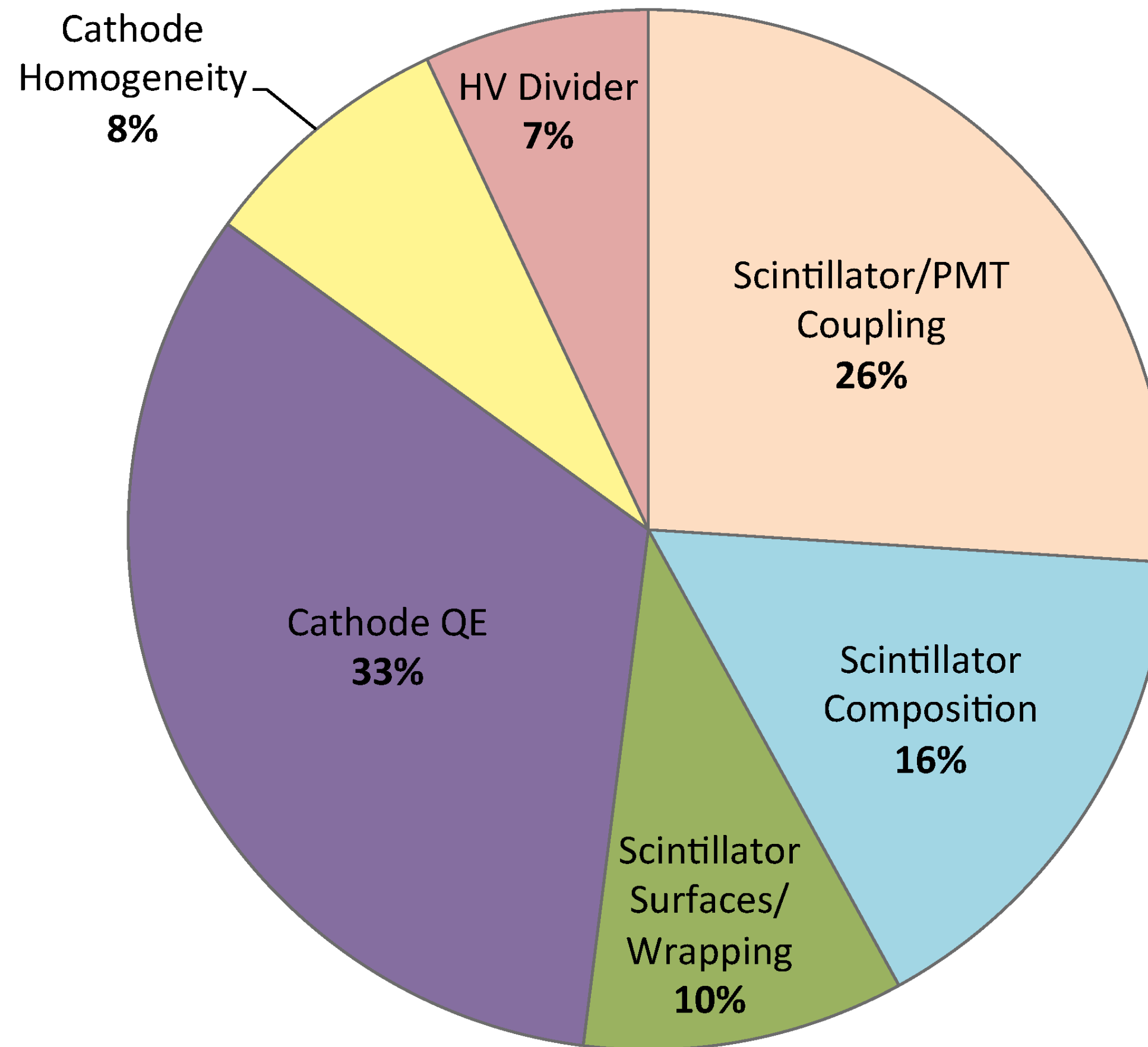
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 With side, top and bottoms: 712 modules total

*Nucl. Inst. Meth. A 868, 98-108 (2017)*



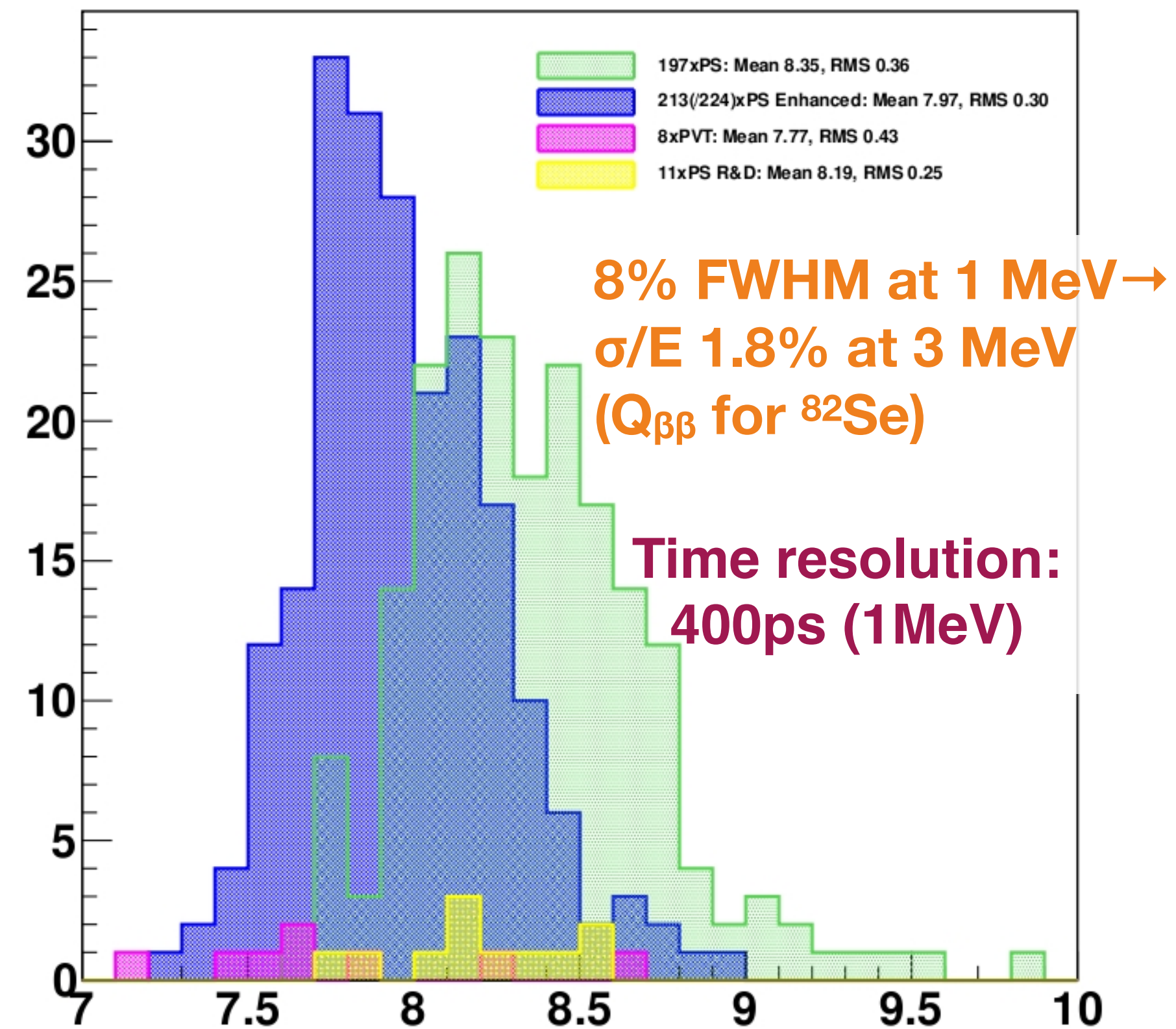
## Contributions to improved resolution



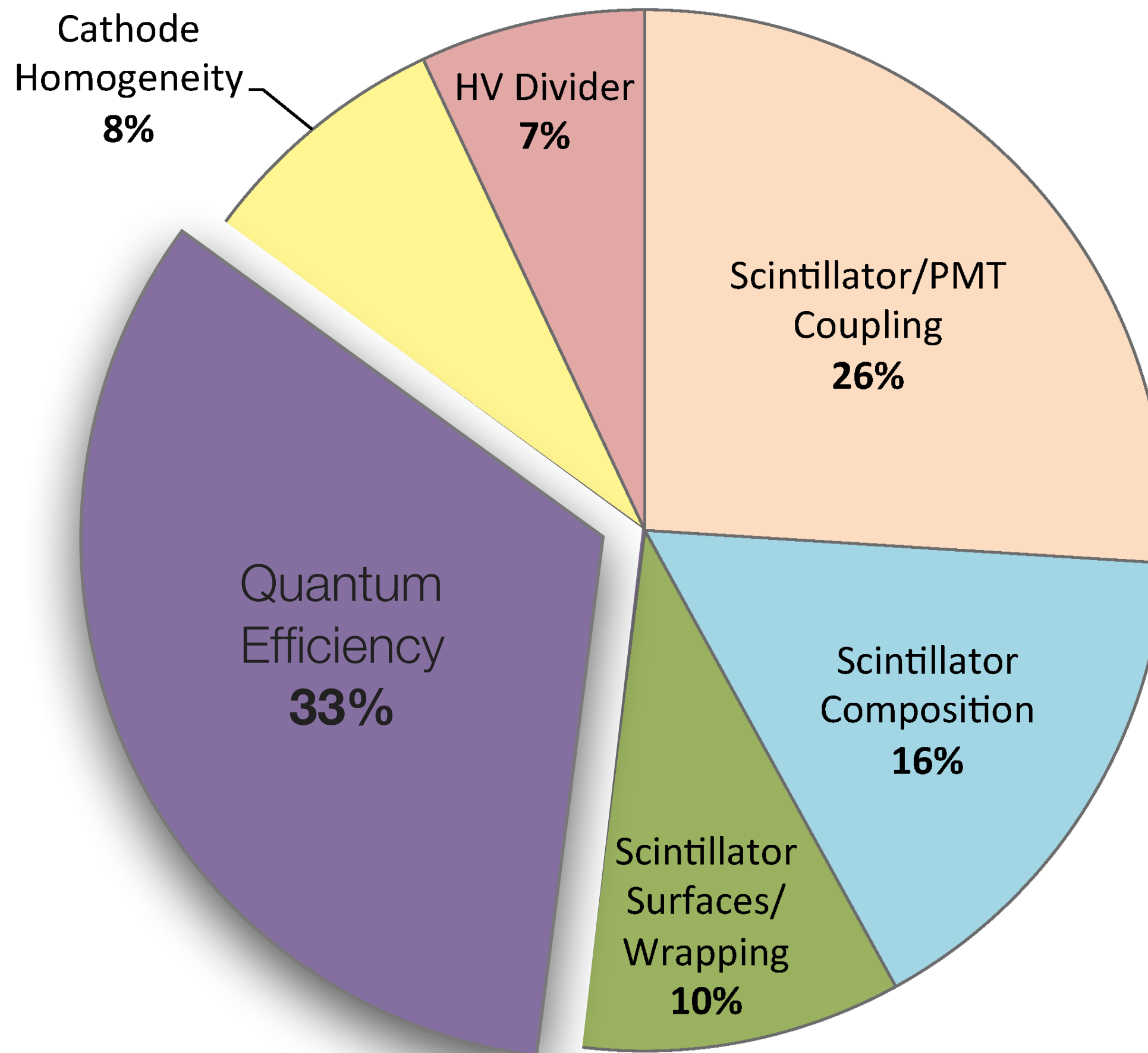
# Calorimeter development

Main calorimeter walls: **520** optical modules  
 With side, top and bottoms: **712** modules total

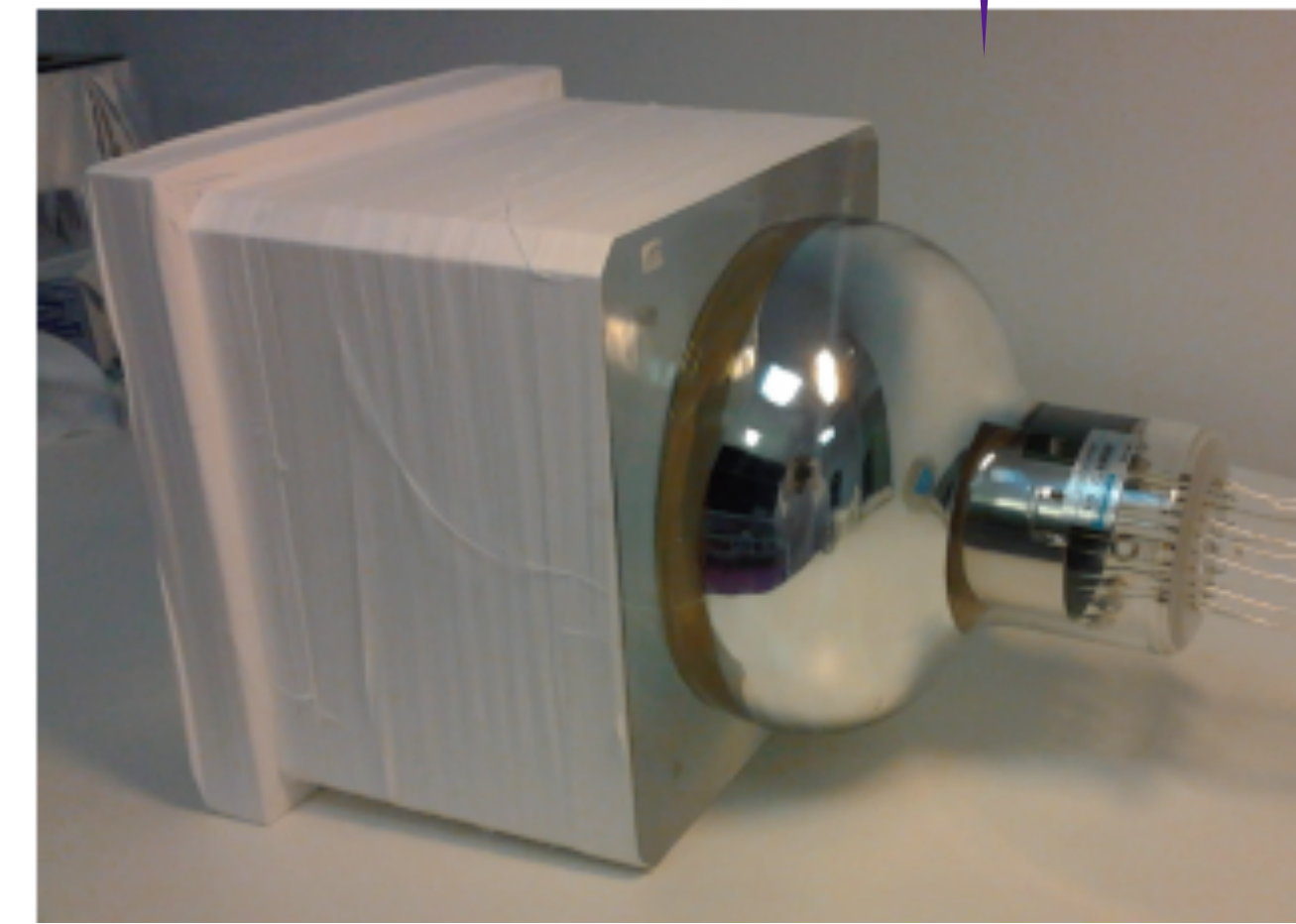
*Nucl. Inst. Meth. A 868, 98-108 (2017)*



## Contributions to improved resolution



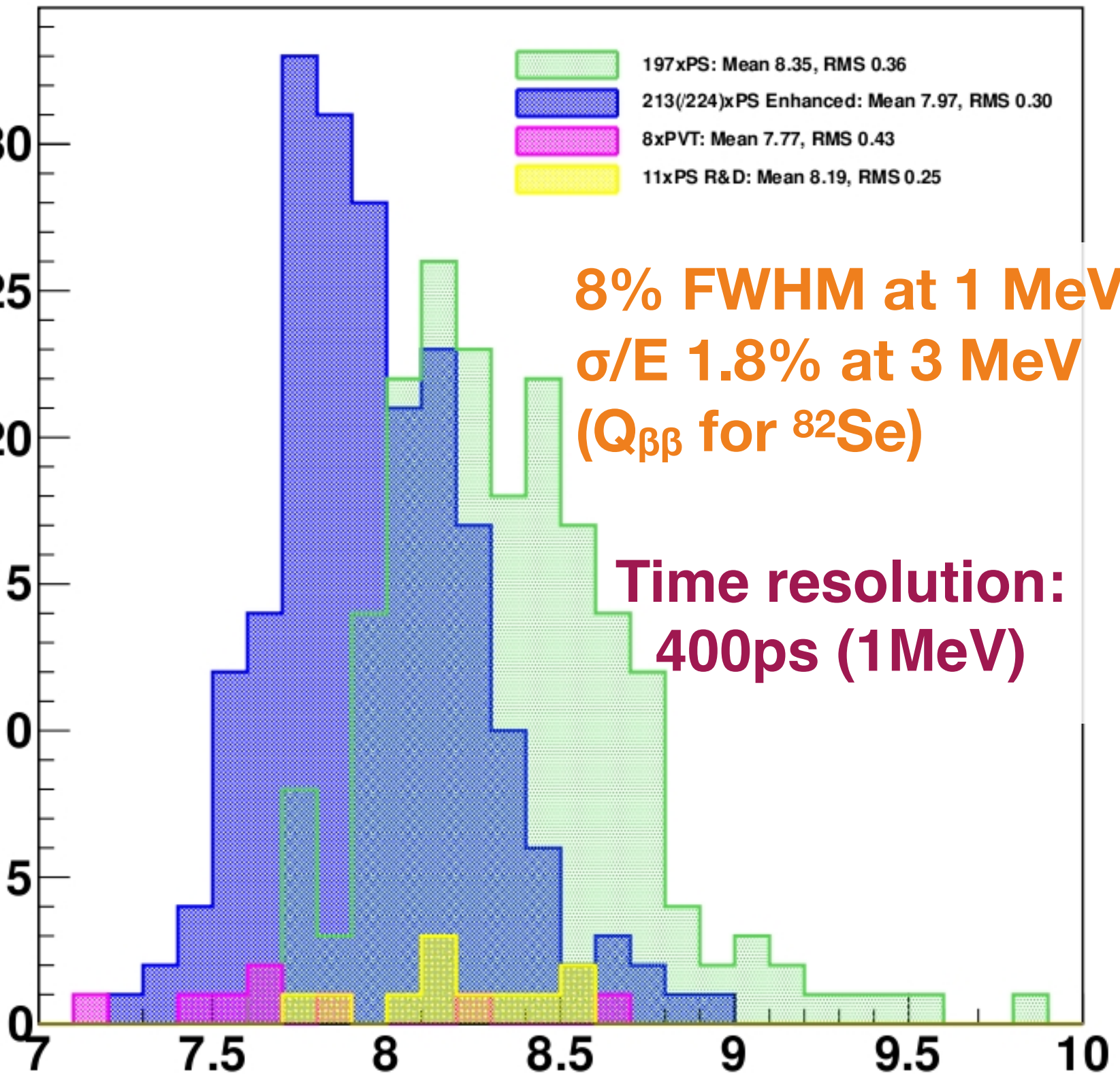
**440 8" radiopure PMTs**  
 with improved photocathode quantum efficiency  
 (5" PMTs for outer rows and columns. side, top and bottom)



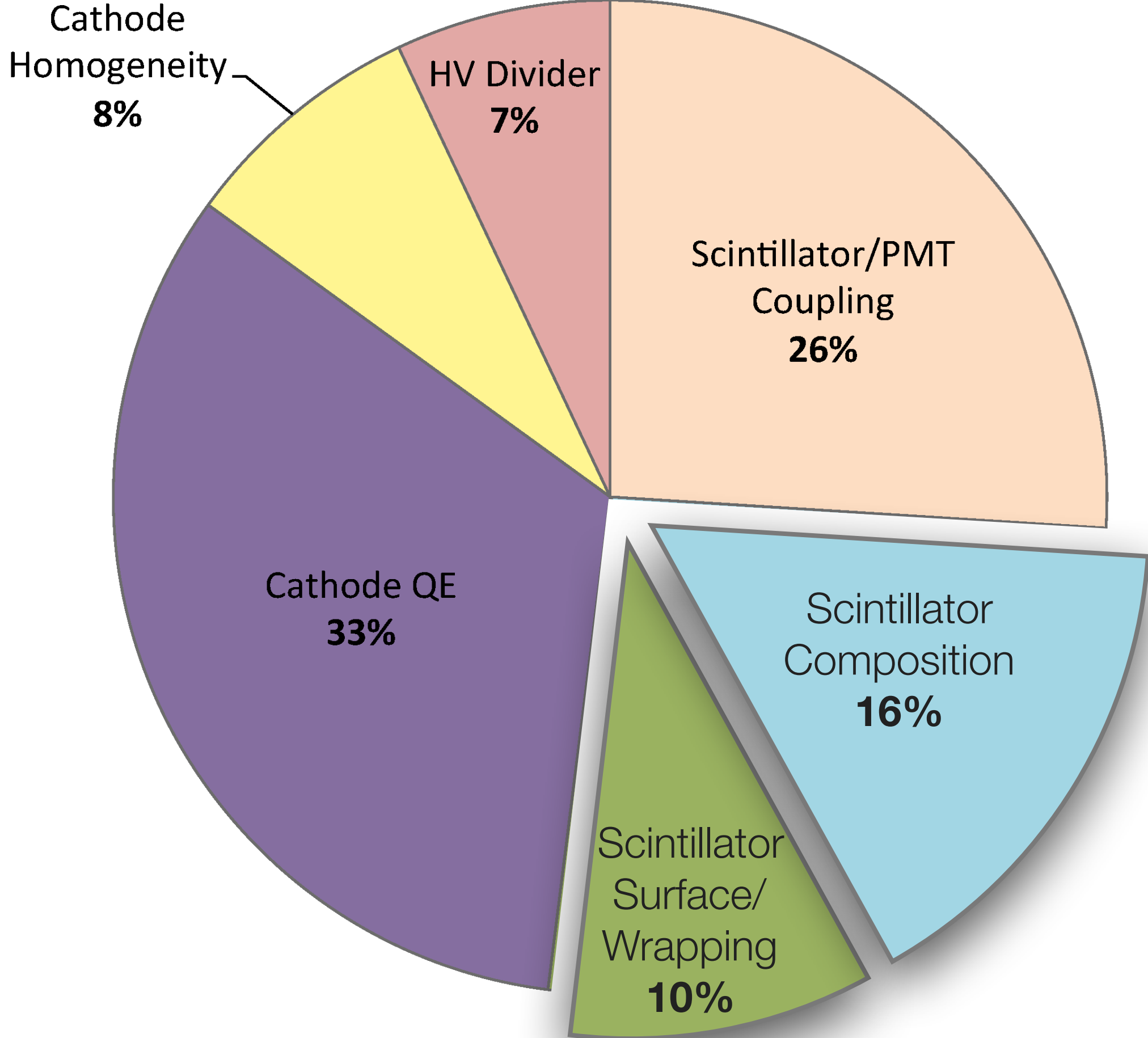
# Calorimeter development

Main calorimeter walls: 520 optical modules  
 With side, top and bottoms: 712 modules total

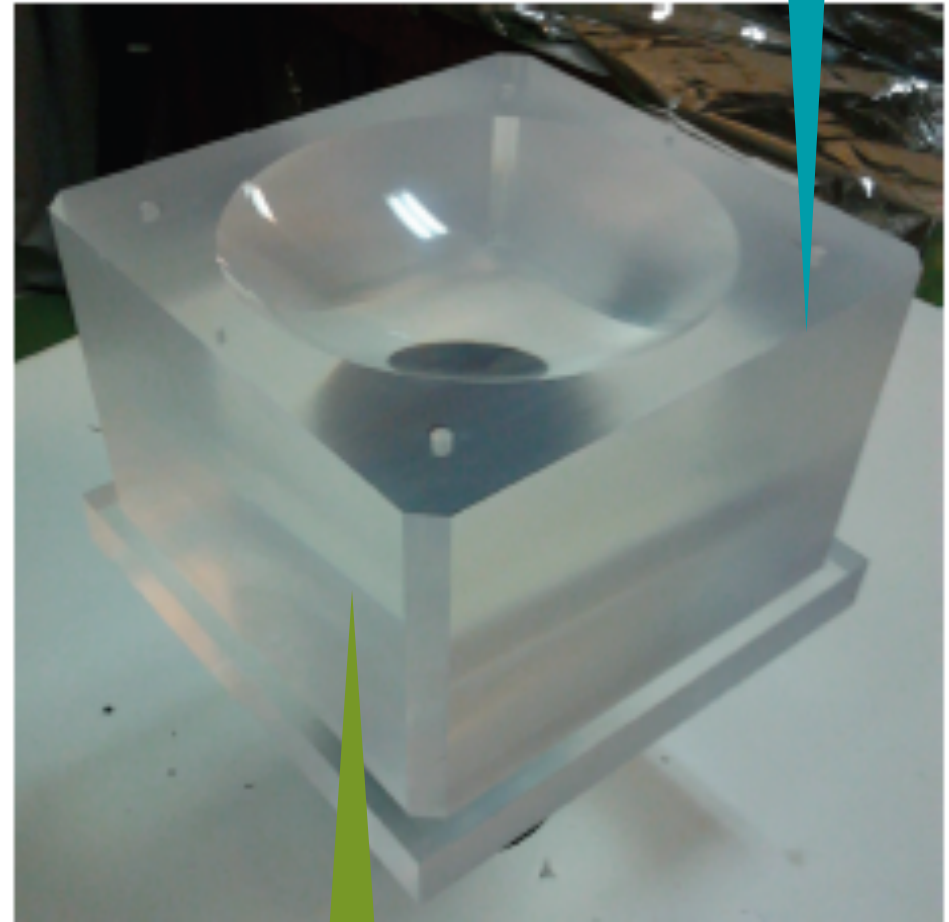
*Nucl. Inst. Meth. A 868, 98-108 (2017)*



## Contributions to improved resolution



256 x 256 x 194 mm  
**Polystyrene scintillator block**

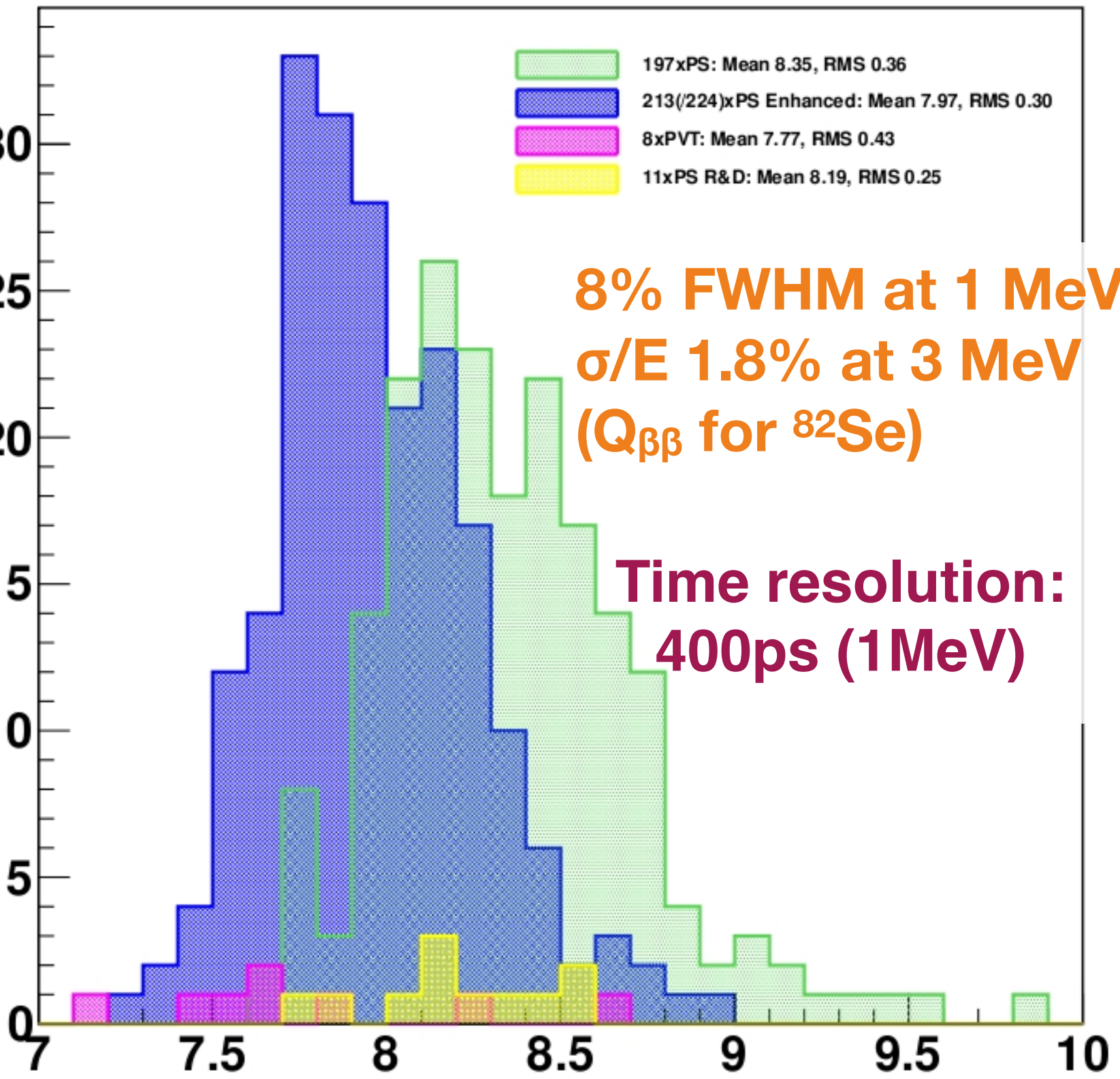


Optimised surface finish

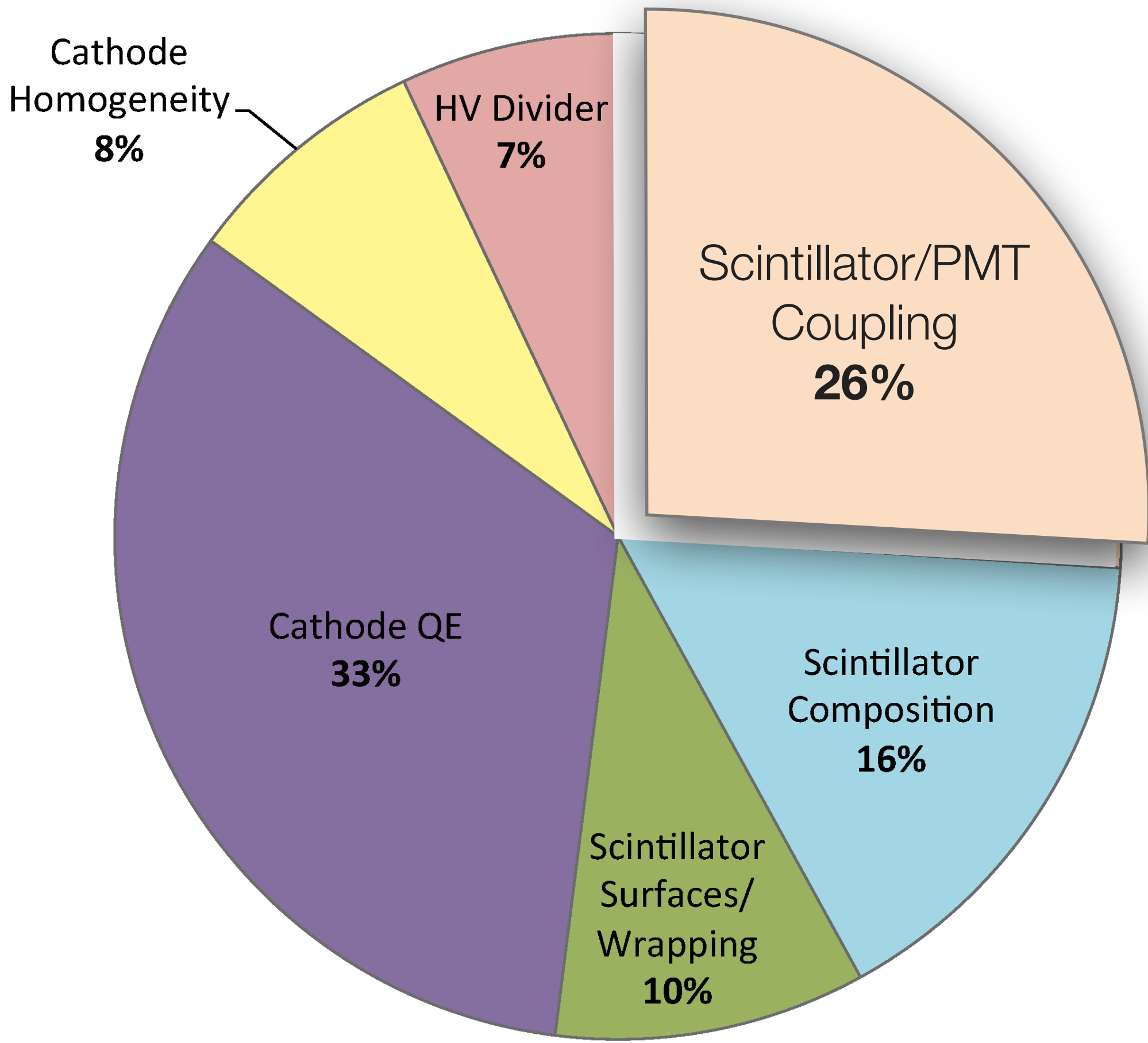
# Calorimeter development

Main calorimeter walls: 520 optical modules  
 With side, top and bottoms: 712 modules total

*Nucl. Inst. Meth. A 868, 98-108 (2017)*



## Contributions to improved resolution



Directly coupled - no light-guide

