

Project 8 and Neutrino Mass

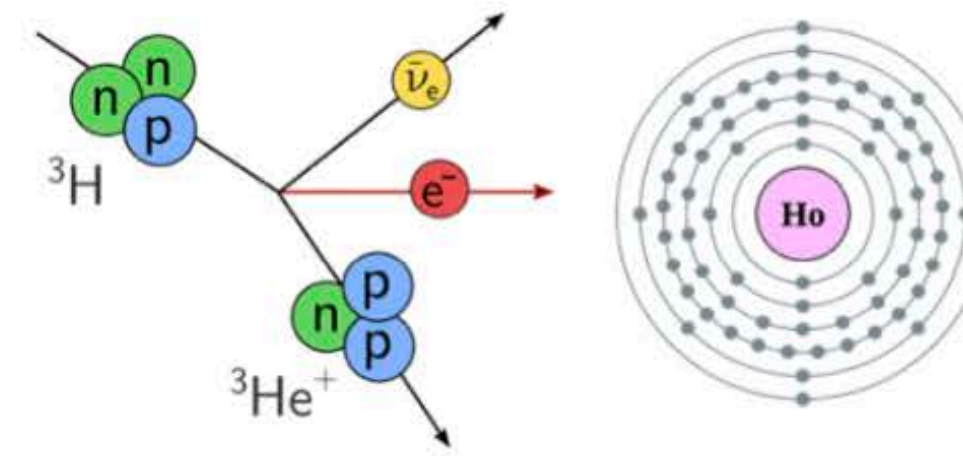
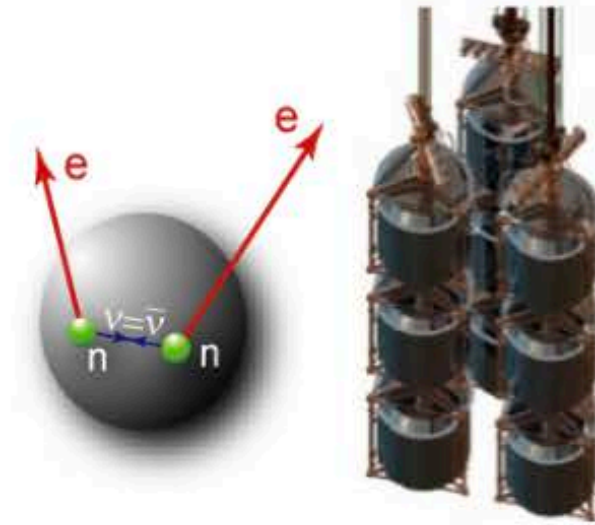
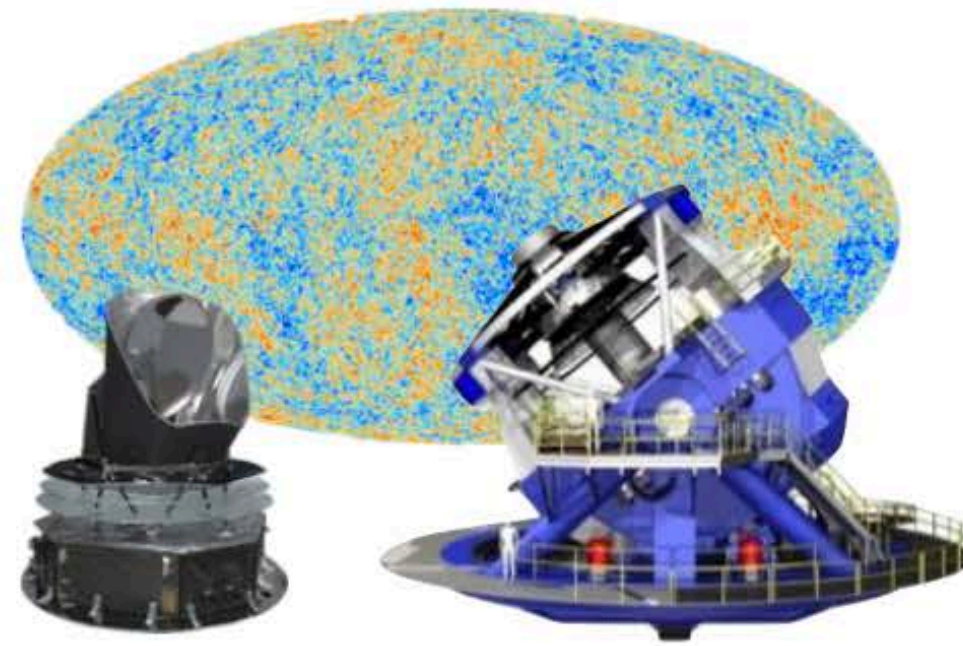
A frequency based approach to measure the neutrino mass



Karsten M. Heeger
Yale University

CERN October 2019

Paths to the Neutrino Mass Scale



	Cosmology	Search for $0\nu\beta\beta$	β -decay & electron capture
Observable	$M_\nu = \sum_i m_i$	$m_{\beta\beta}^2 = \left \sum_i U_{ei}^2 m_i \right ^2$	$m_\beta^2 = \sum_i U_{ei} ^2 m_i^2$
Present upper limit	~0.1 – 0.6 eV	~0.1 – 0.4 eV	2 eV
Potential: near-term (long-term)	60 meV (15 meV)	50 – 200 meV (20 – 40 meV)	200 meV (40 – 100 meV)
Model dependence	Multi-parameter cosmological model	<ul style="list-style-type: none"> - Majorana nature of ν, lepton number violation - BSM contributions other than $m(\nu)$? - Nuclear matrix elements 	Direct , only kinematics; no cancellations in incoherent sum

Valerius

Neutrino Mass Constraints

Cosmology measures

$$\sum_i m_i$$

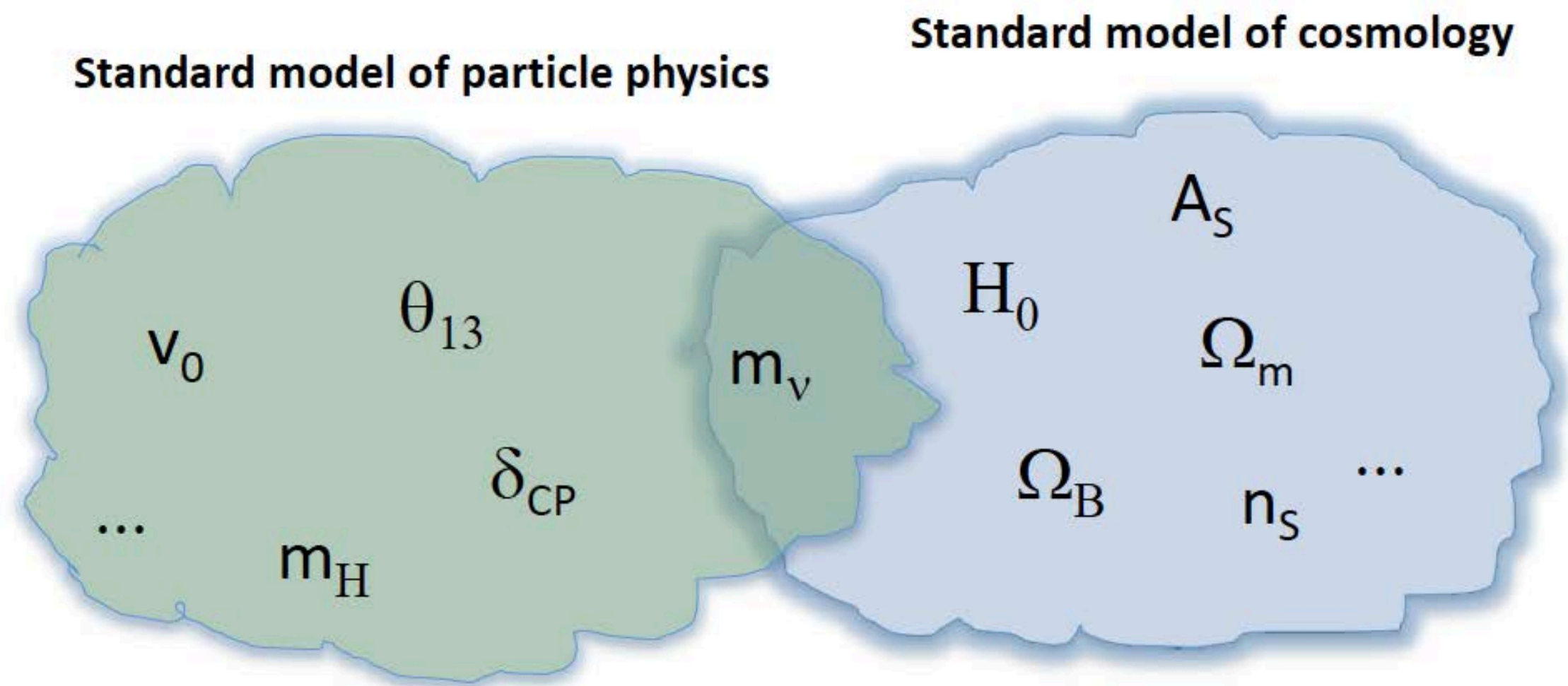
Double beta decay measures

$$\left| \sum_i U_{ei}^2 m_i \right|$$

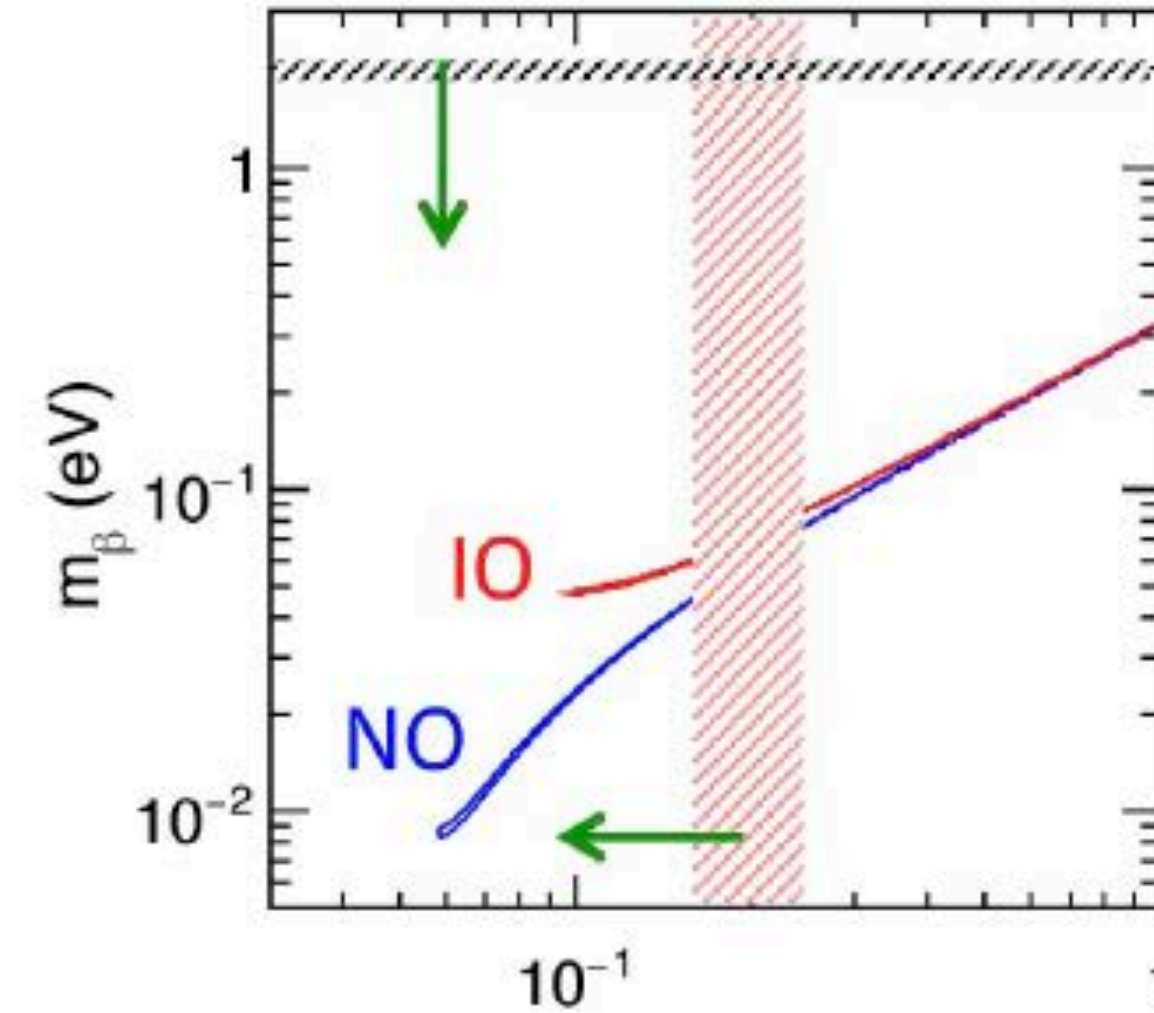
Direct searches measure

$$\left(\sum_i |U_{ei}^2| m_i^2 \right)^{1/2}$$

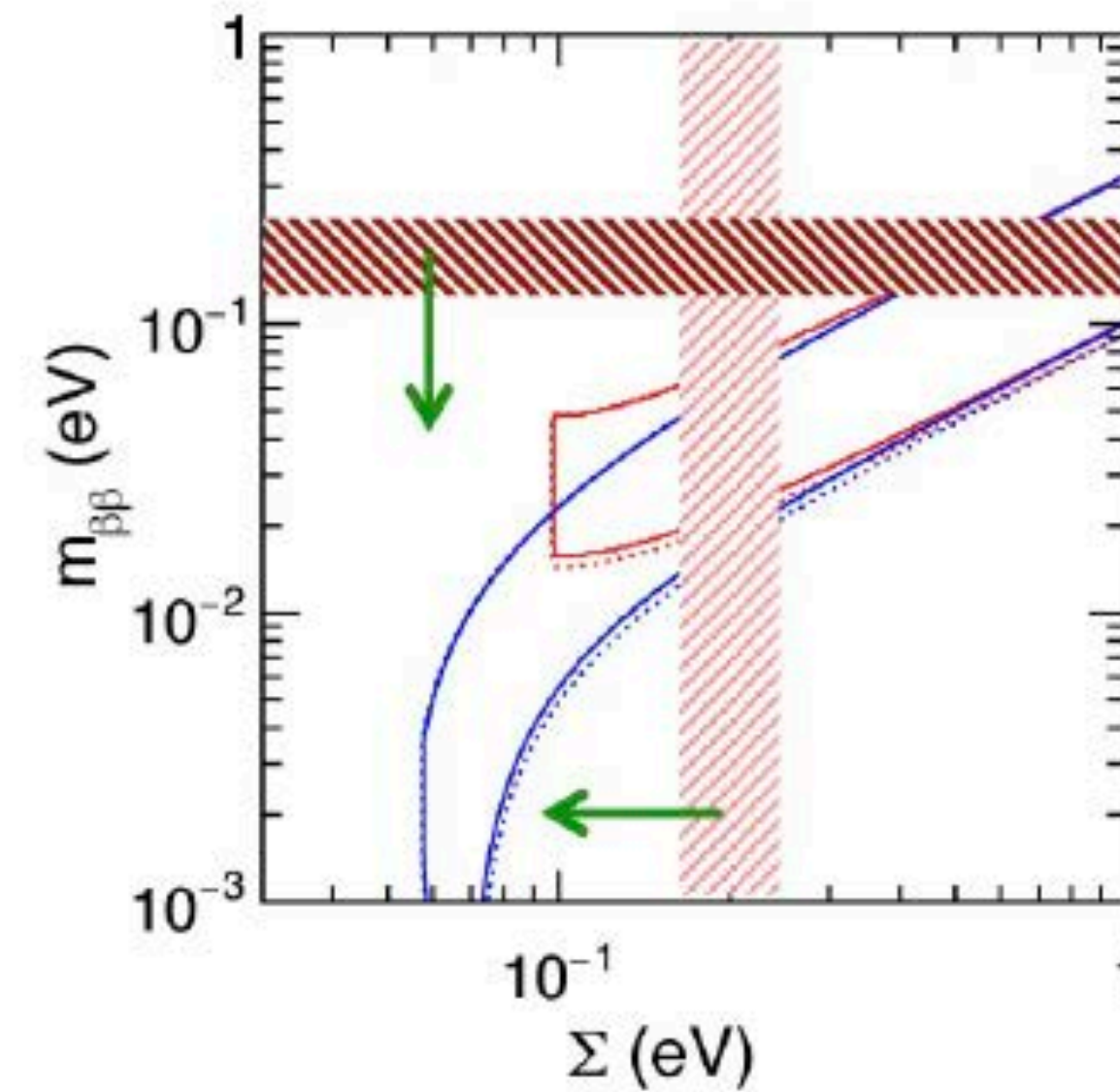
m_ν measurable both by laboratory experiments and cosmology
 a critical test of consistency



Direct mass searches



Double beta decay



Cosmology

Mezetto

Neutrino Mass Constraints

Cosmology measures

$$\sum_i m_i$$

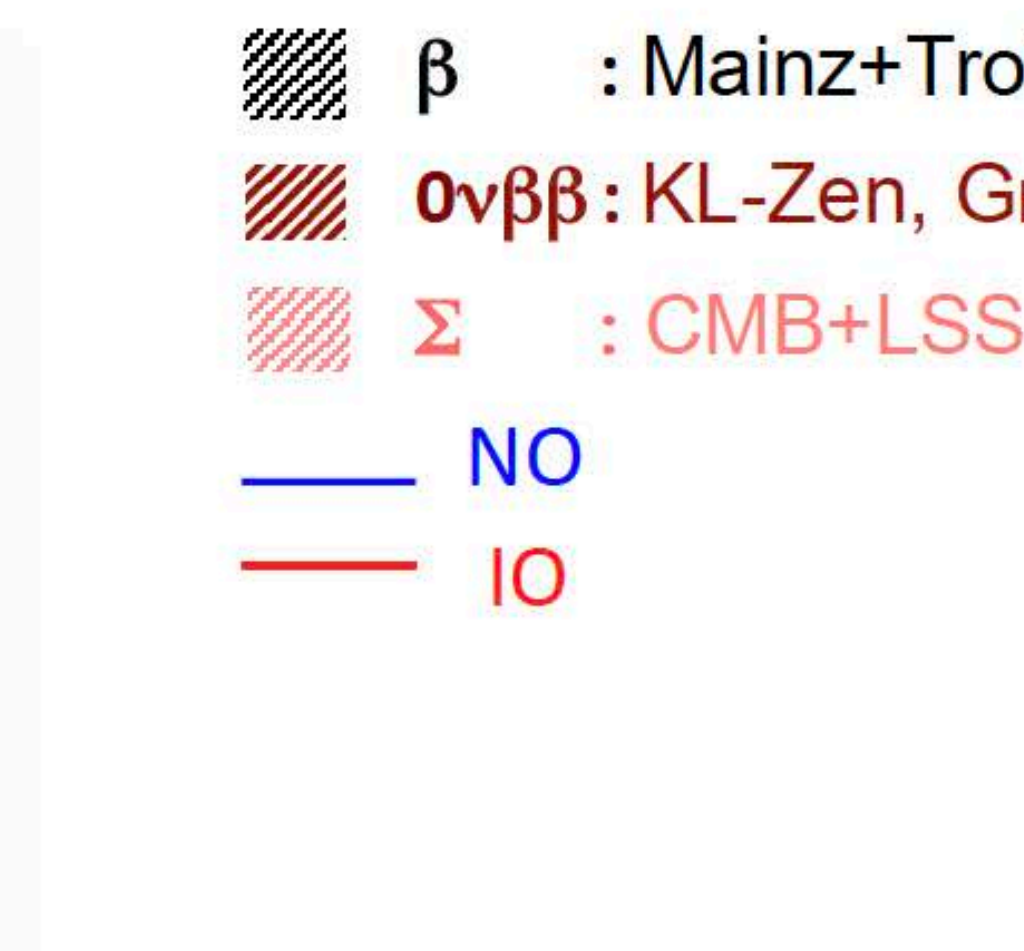
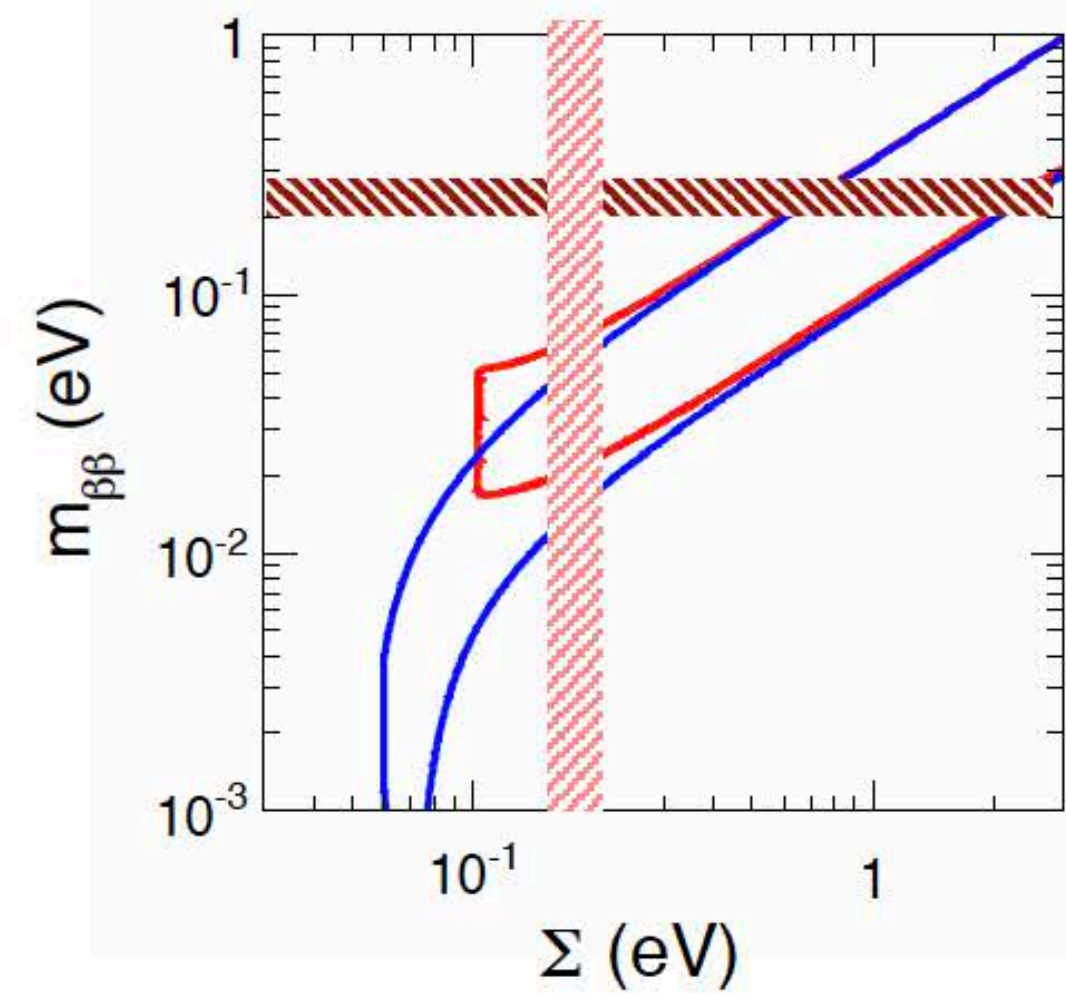
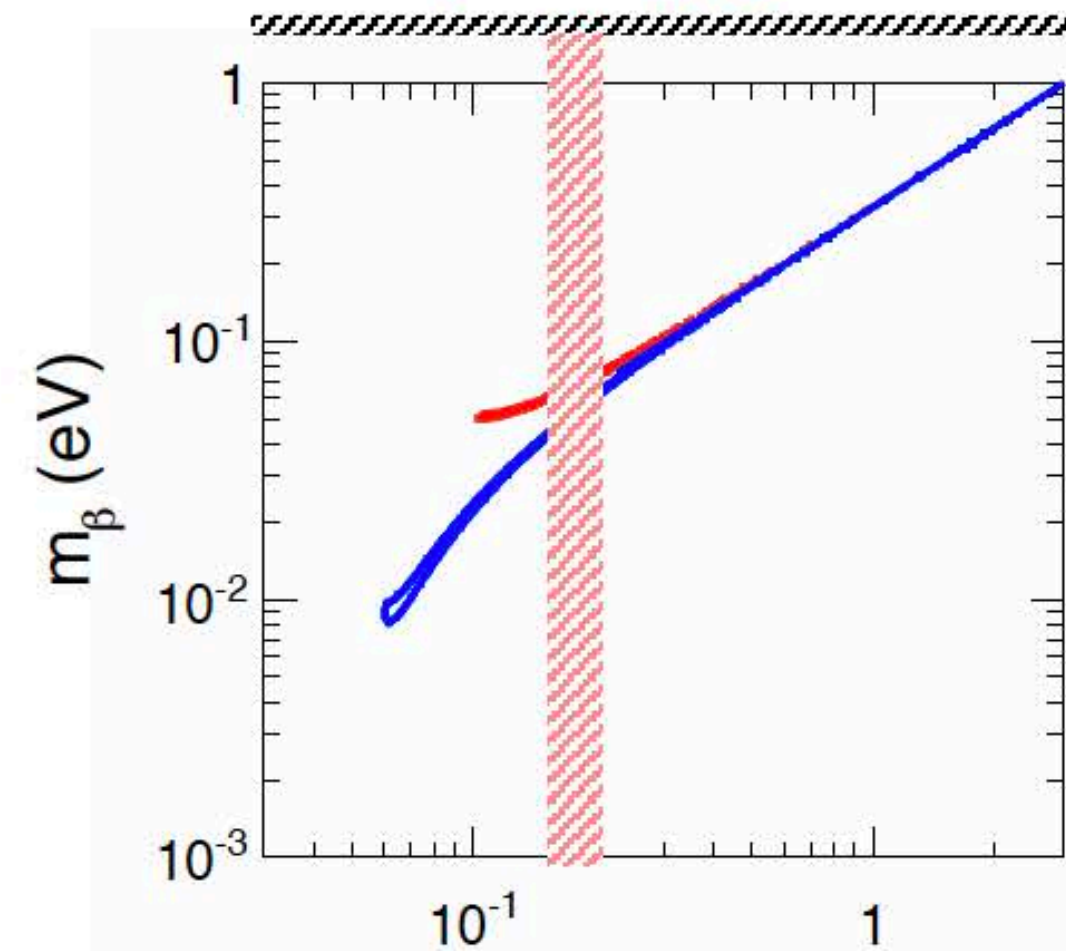
Double beta decay measures

$$\left| \sum_i U_{ei}^2 m_i \right|$$

Direct searches measure

$$\left(\sum_i |U_{ei}^2| m_i^2 \right)^{1/2}$$

m_ν measurable both by laboratory experiments and cosmology
 a critical test of consistency



- β : Mainz+Troitsk
- $0\nu\beta\beta$: KL-Zen, GERDA, EXO, Cuore...
- Σ : CMB+LSS
- NO
- IO

Mezetto

Neutrino Mass Constraints

Cosmology measures

$$\sum_i m_i$$

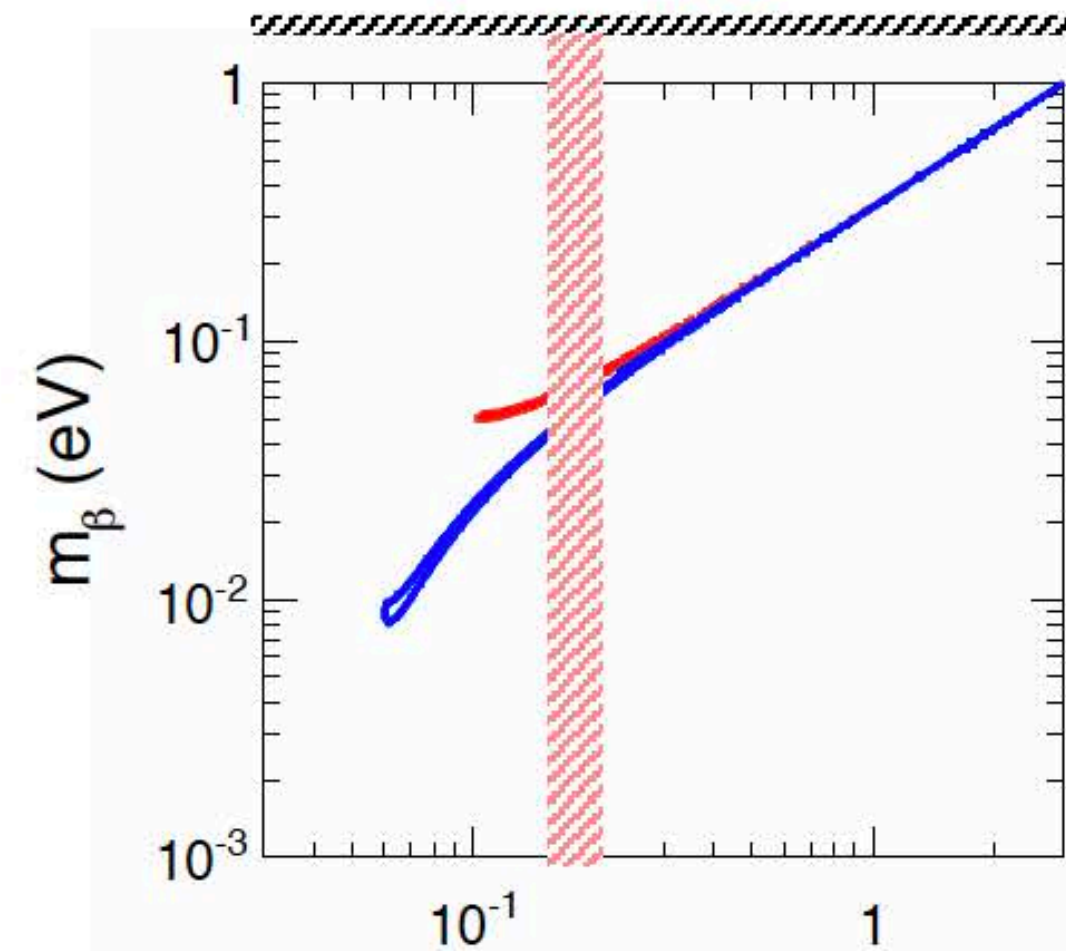
Double beta decay measures

$$\left| \sum_i U_{ei}^2 m_i \right|$$

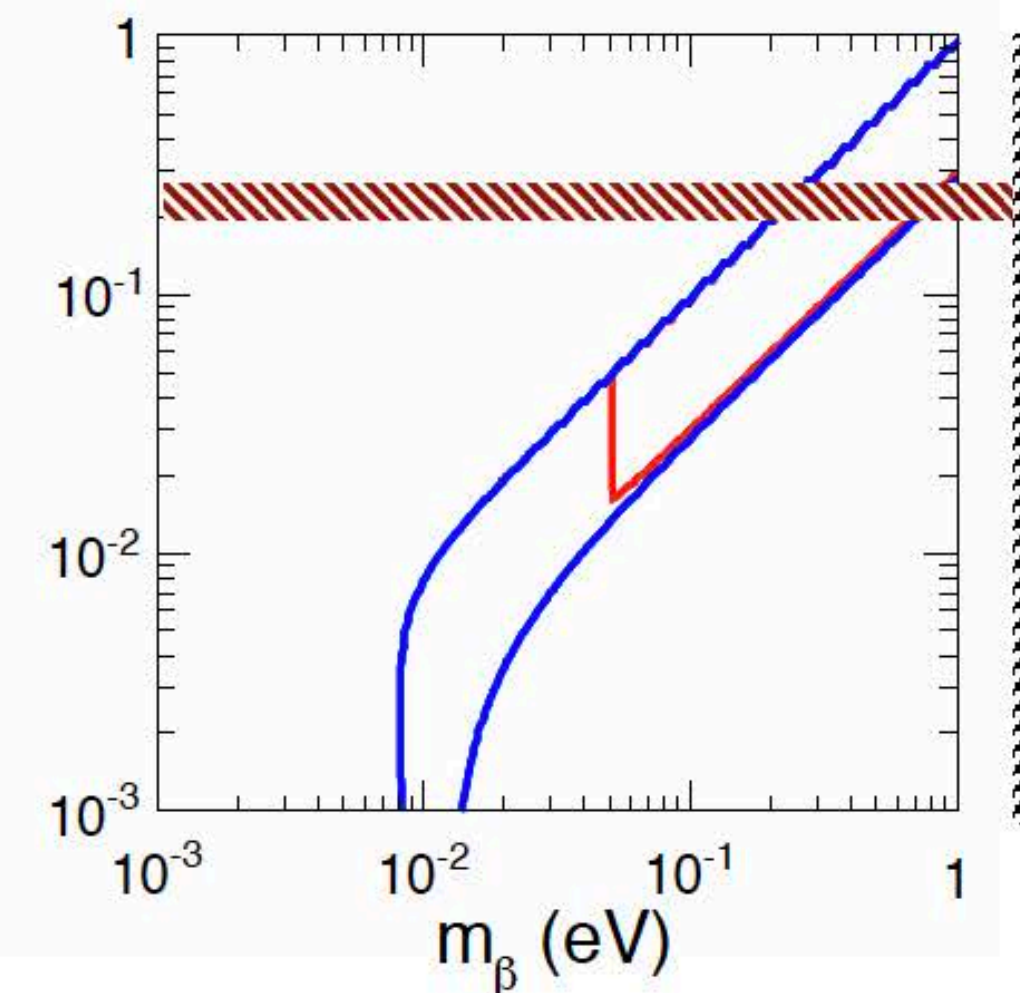
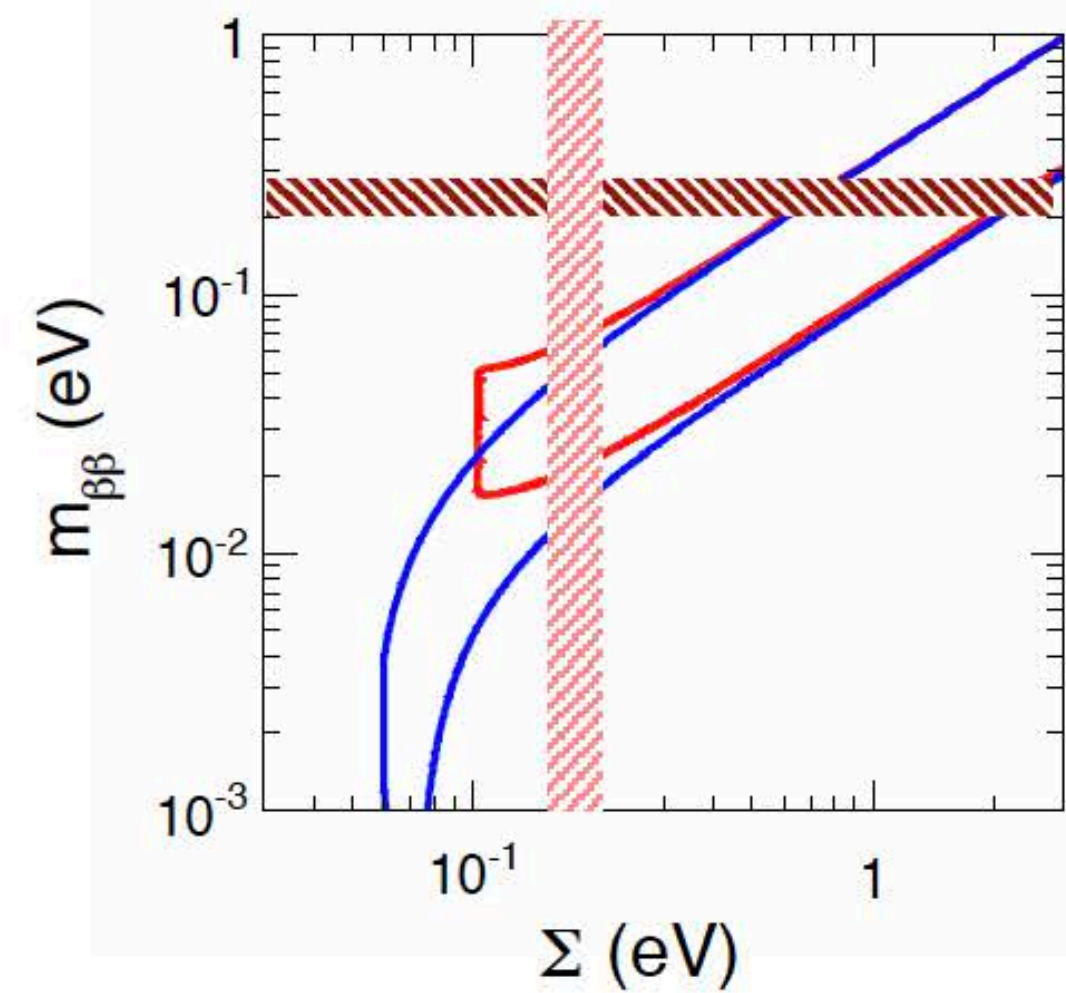
Direct searches measure

$$\left(\sum_i |U_{ei}^2| m_i^2 \right)^{1/2}$$

m_ν measurable both by laboratory experiments and cosmology
 a critical test of consistency



- β : Mainz+Troitsk
- $0\nu\beta\beta$: KL-Zen, GERDA, EXO, Cuore...
- Σ : CMB+LSS
- NO
- IO



"consistency would be spectacular confirmation!"

Mezetto

Neutrino Mass Constraints

Cosmology measures

$$\sum_i m_i$$

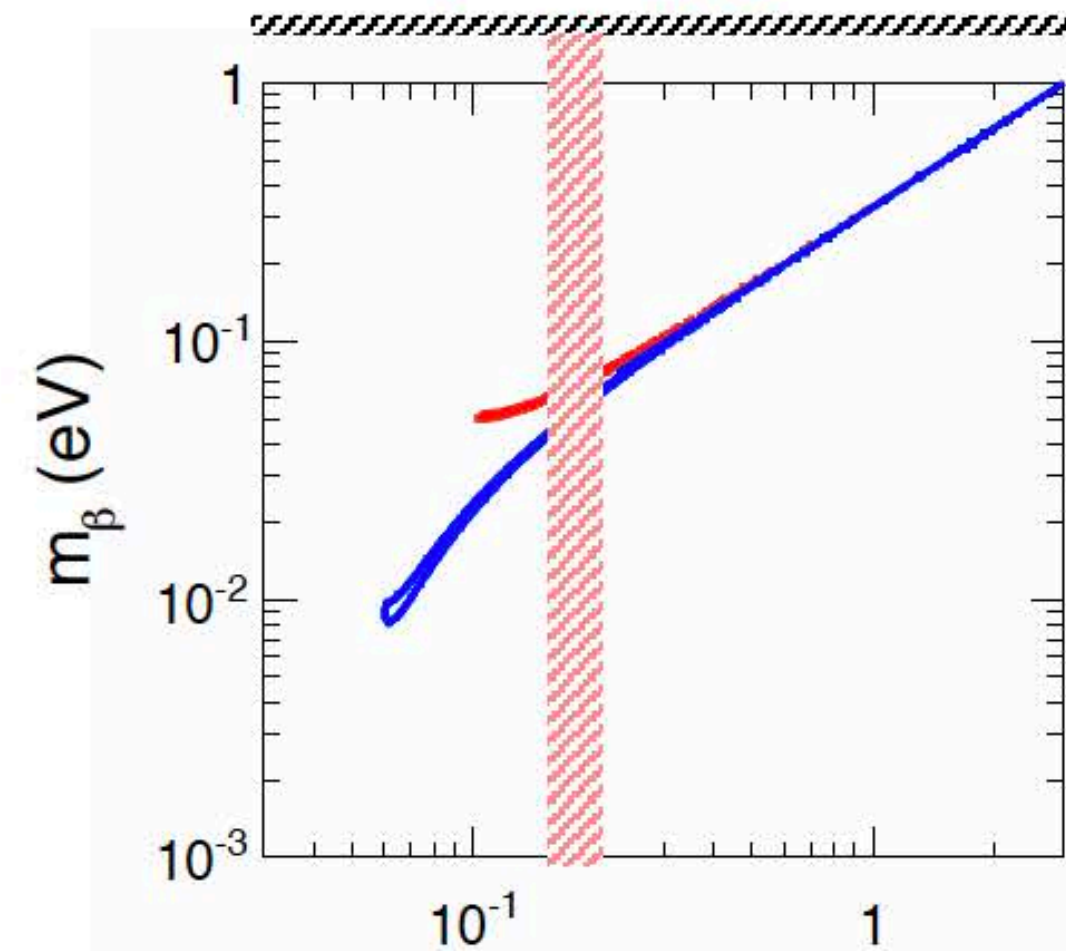
Double beta decay measures

$$\left| \sum_i U_{ei}^2 m_i \right|$$

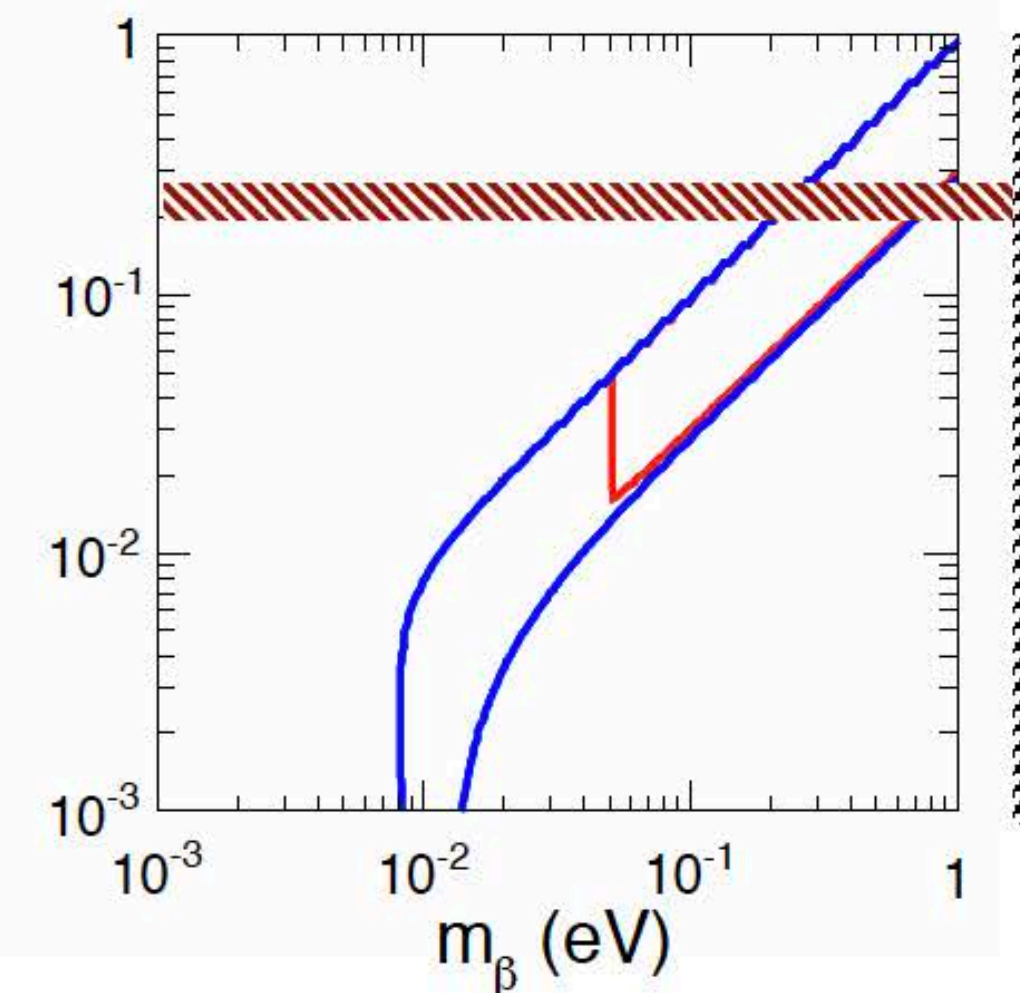
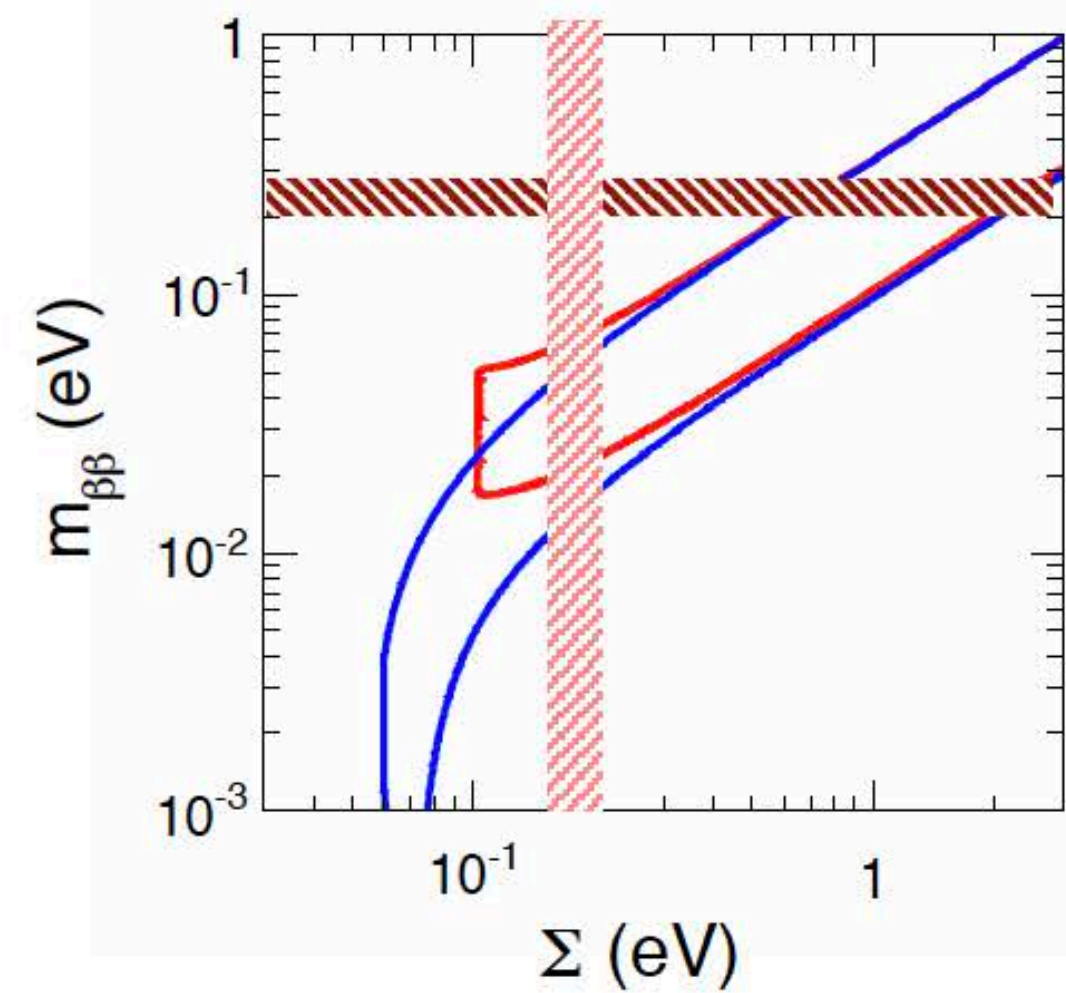
Direct searches measure

$$\left(\sum_i |U_{ei}^2| m_i^2 \right)^{1/2}$$

m_ν measurable both by laboratory experiments and cosmology
 a critical test of consistency



- β : Mainz+Troitsk
- $0\nu\beta\beta$: KL-Zen, GERDA, EXO, Cuore...
- Σ : CMB+LSS
- NO
- IO



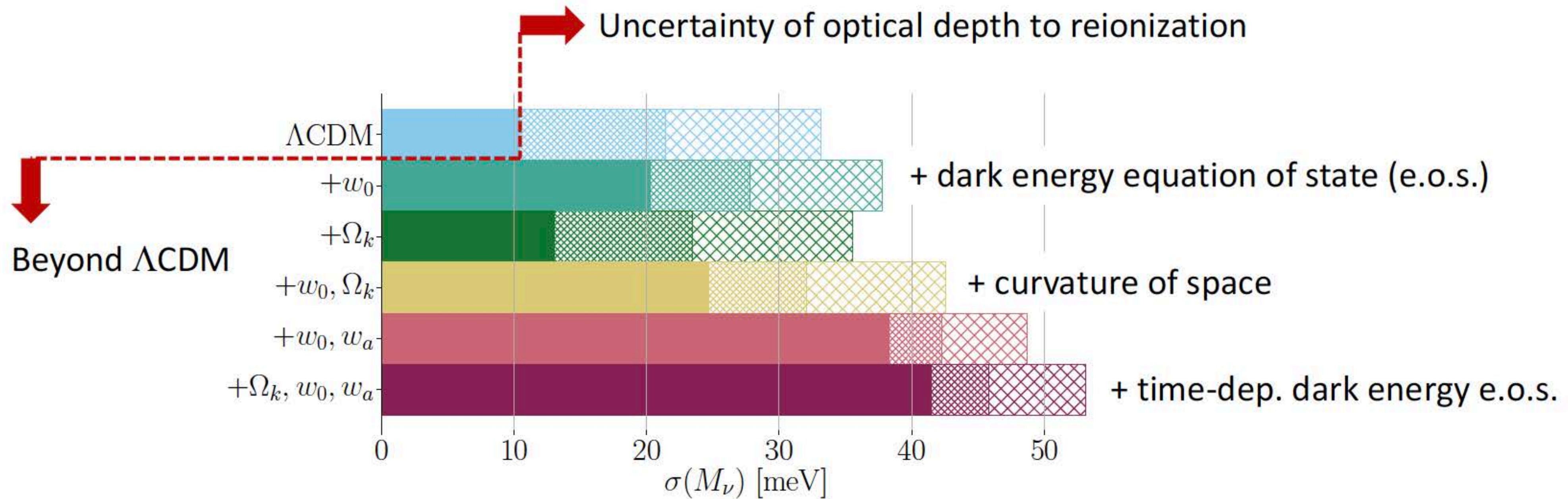
“inconsistency would be major discovery”

Mezetto

Future of Cosmological Neutrino Bounds

Example Forecast:

EUCLID + Planck + Simon's Observatory



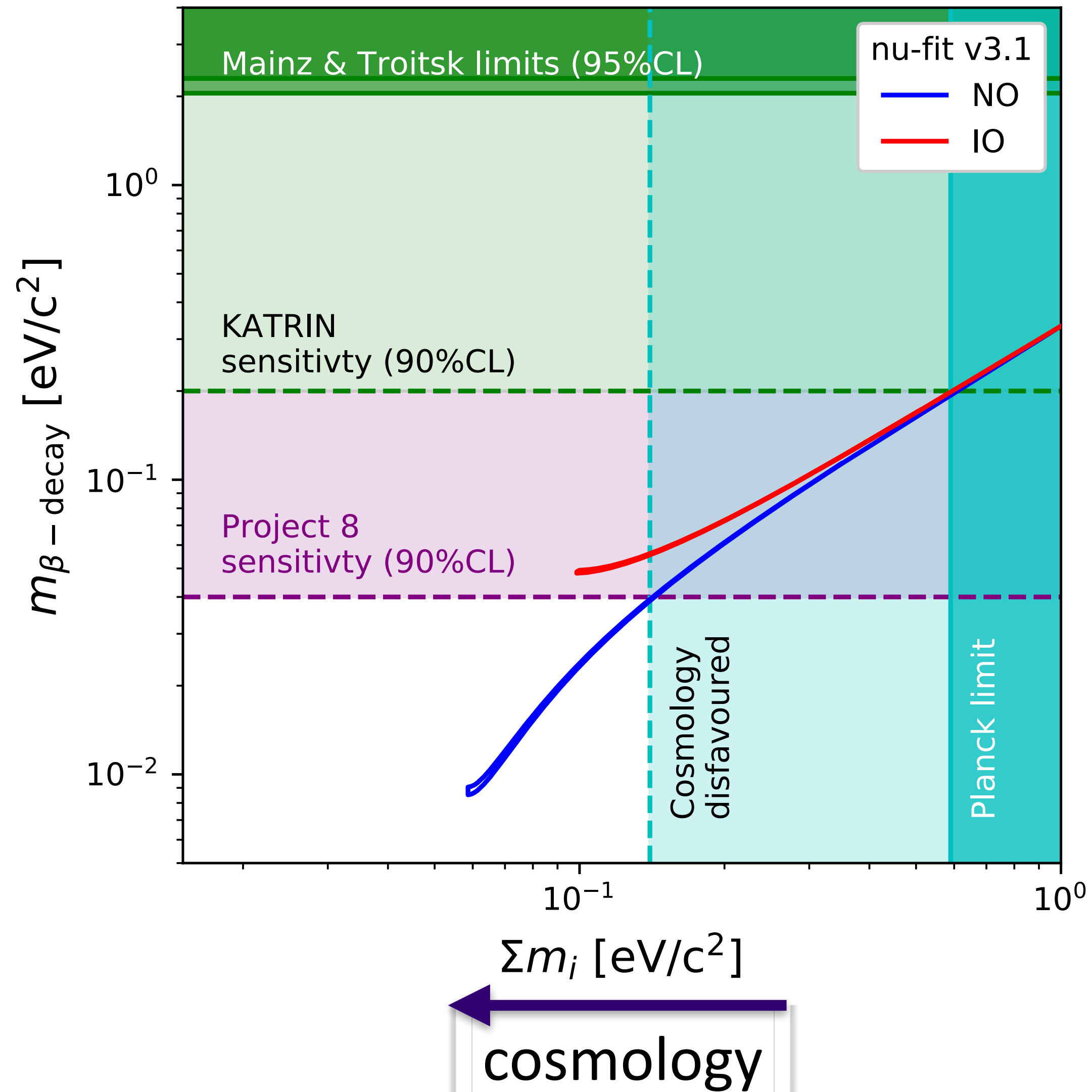
A. Boylea, E. Komatsua, JCAP 2018

Mertens

Complementary Neutrino Mass Limits

inverted mass ordering

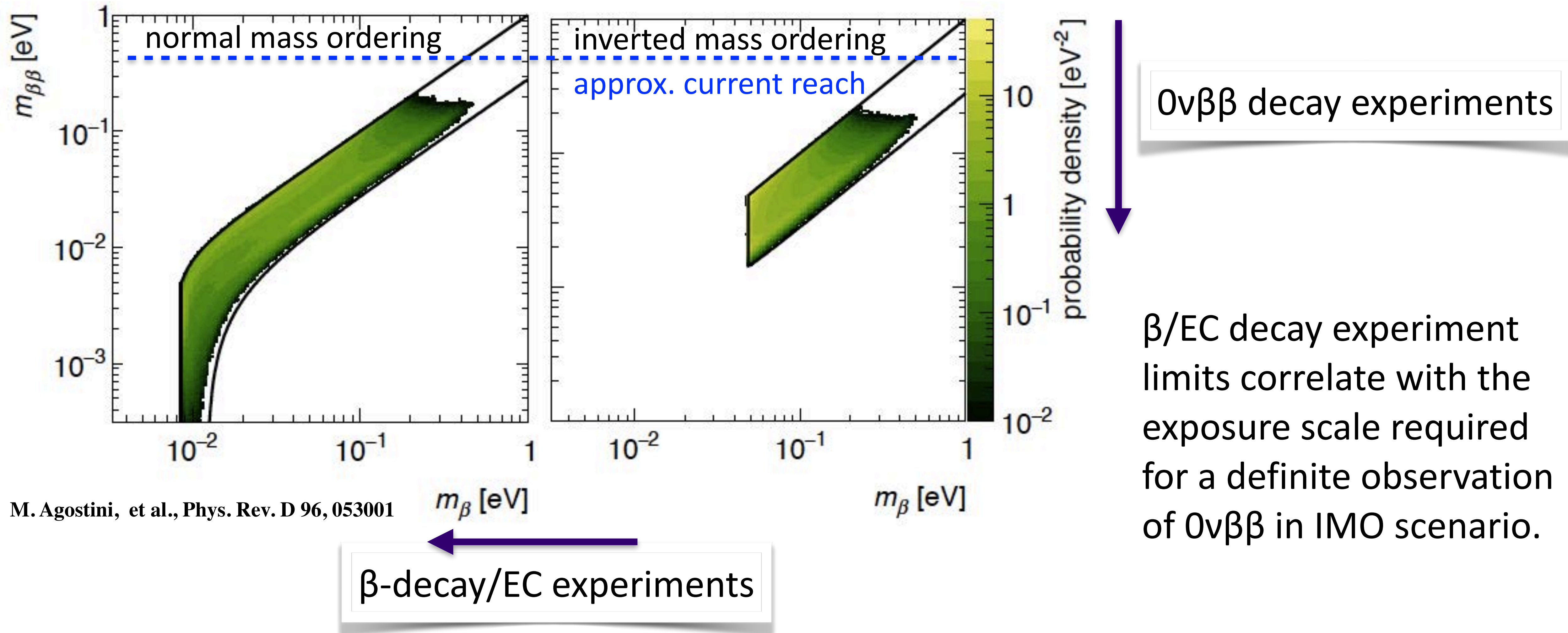
normal mass ordering



β /EC decay experiments

S. Boeser with input from www.nu-fit.org

Complementary Neutrino Mass Limits



Neutrinos and Beta Spectrum

first evidence of neutrino from continuous beta spectrum

Chadwick, 1914

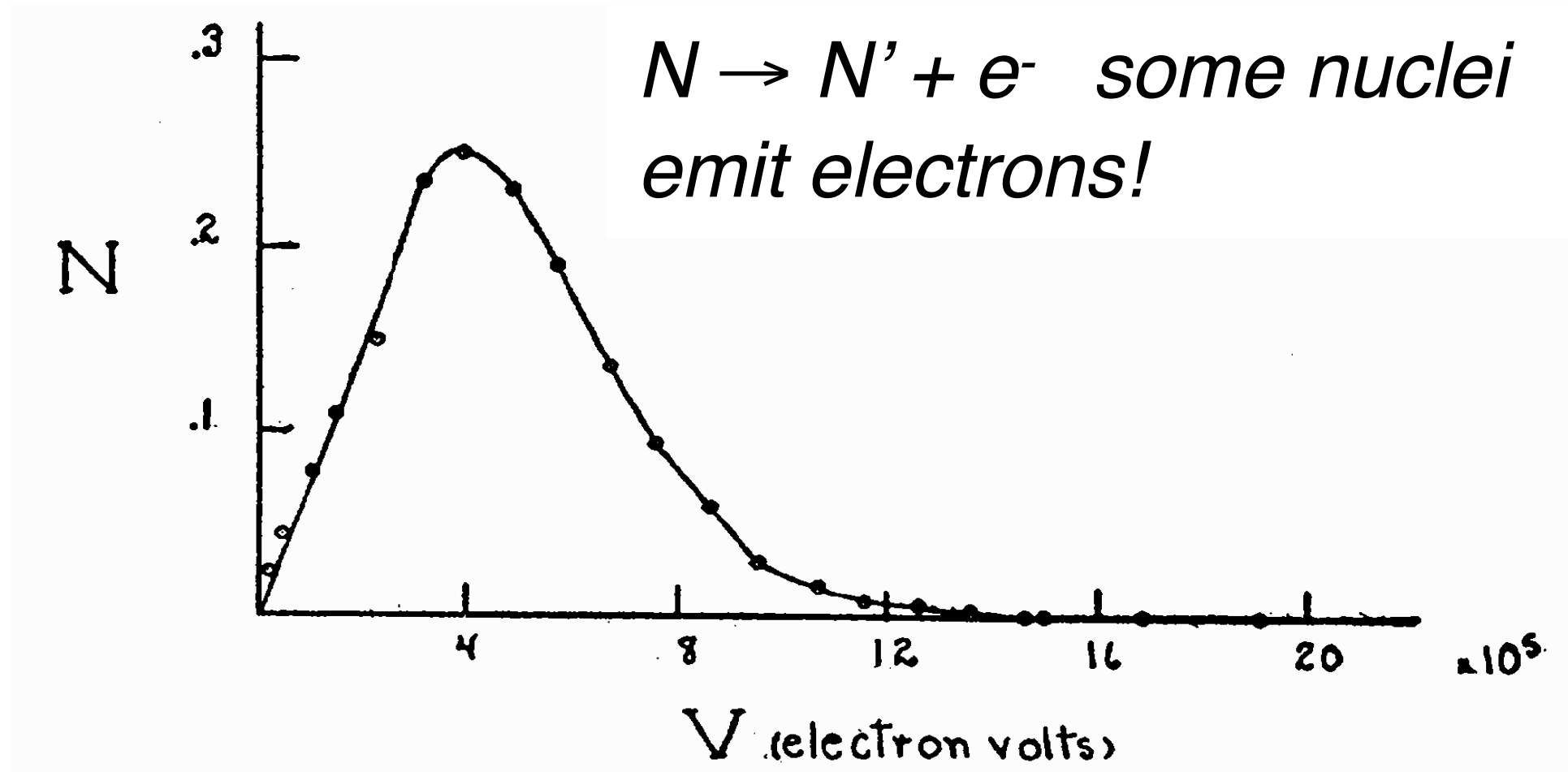
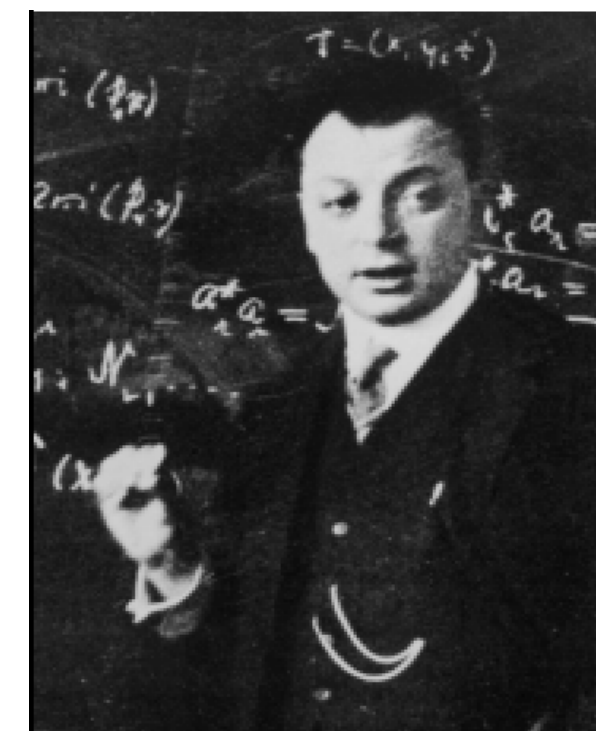


FIG. 5. Energy distribution curve of the beta-rays.



Pauli, 1930

Mythos. Photographie auf 24. 12. 1933
Abschrift/15.12.56 PM

die Gruppe der Radioaktiven bei der
zu Tübingen.

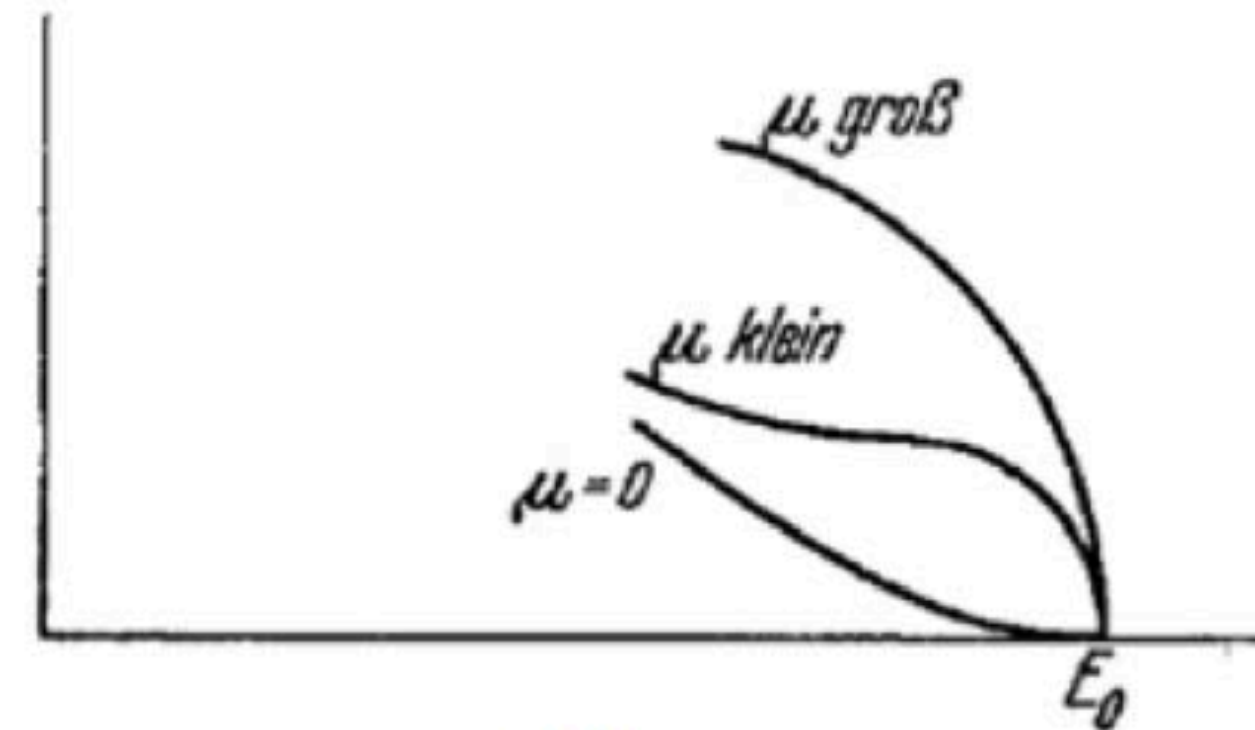
stitut
hen Hochschule

Zürich, 4. Des. 1930
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich halbvollst anhören bitte, Ihnen das näheren auseinandersetzen wird, bin ich angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen beta-Spektrums auf einen verzweifelten Ausweg verfallen um den "Wechselatz" (1) der Statistik und den Energiesatz zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale Teilchen, die ich Neutronen nennen will, in den Kernen existieren, welche den Spin 1/2 haben und das Ausschliessungsprinzip befolgen und sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen dürfte von derselben Grossenordnung wie die Elektronenmasse sein, jedenfalls nicht grösser als 0,01 Protonenmasse. Das kontinuierliche beta-Spektrum wäre dann verständlich unter der Annahme, dass beim beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert wird, derart, dass die Summe der Energien von Neutron und Elektron konstant ist.

a hypothesis



β Energy

E. Fermi, Zeitschrift für Physik 1934

Neutrino Mass and Tritium Beta Decay Endpoint

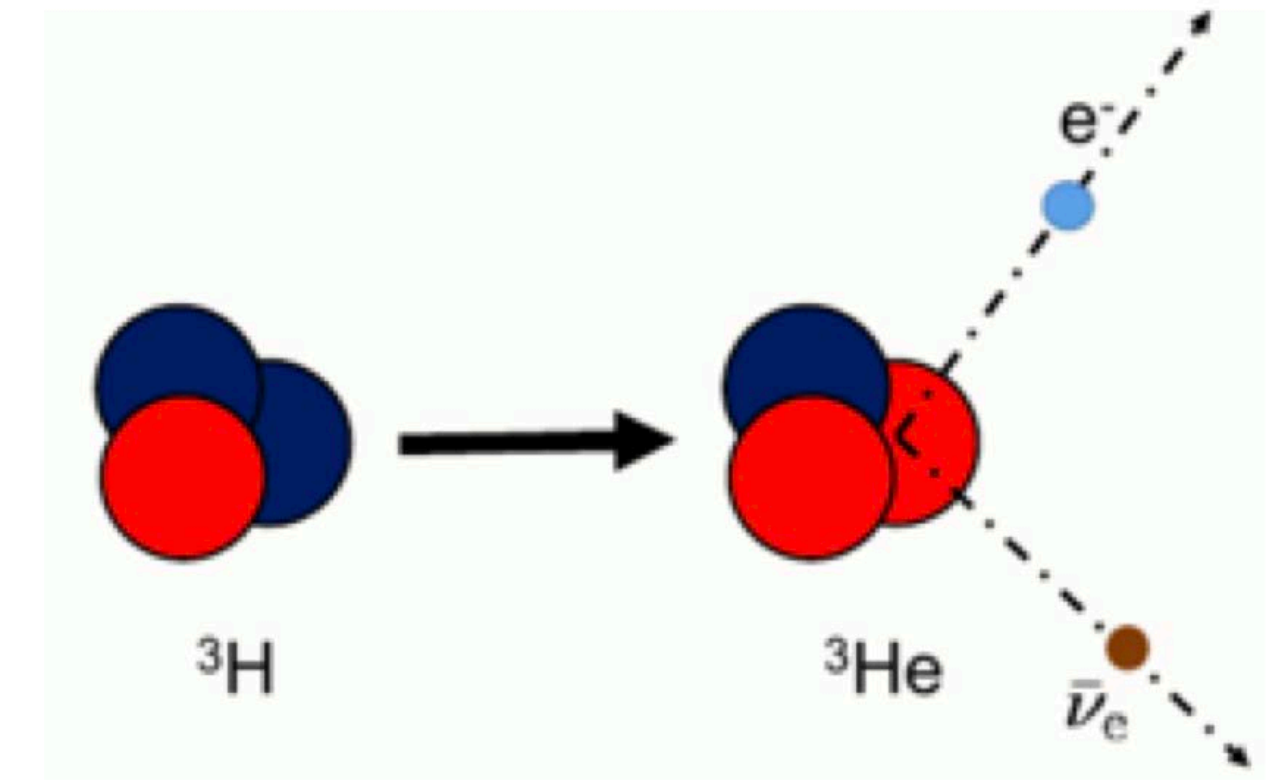
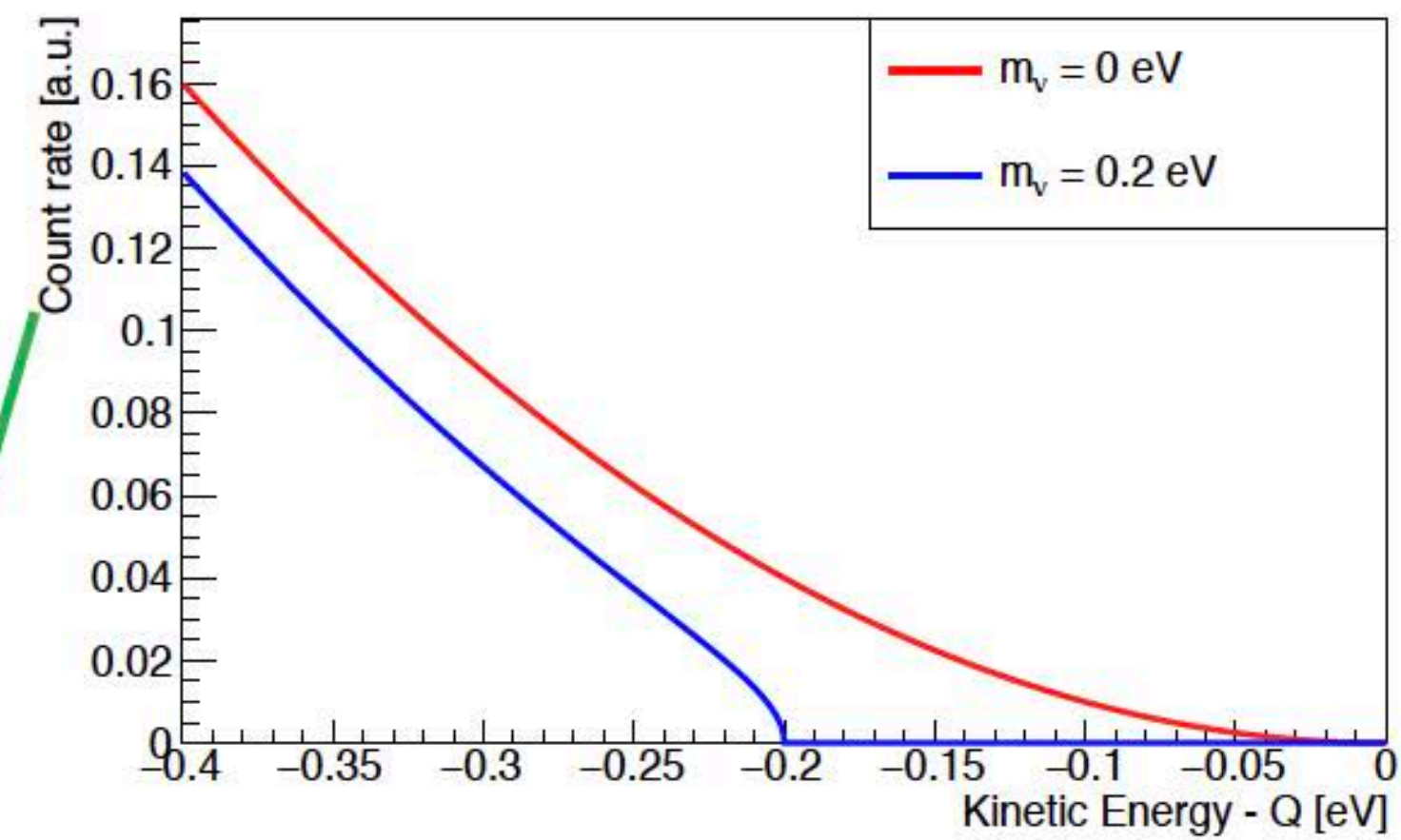
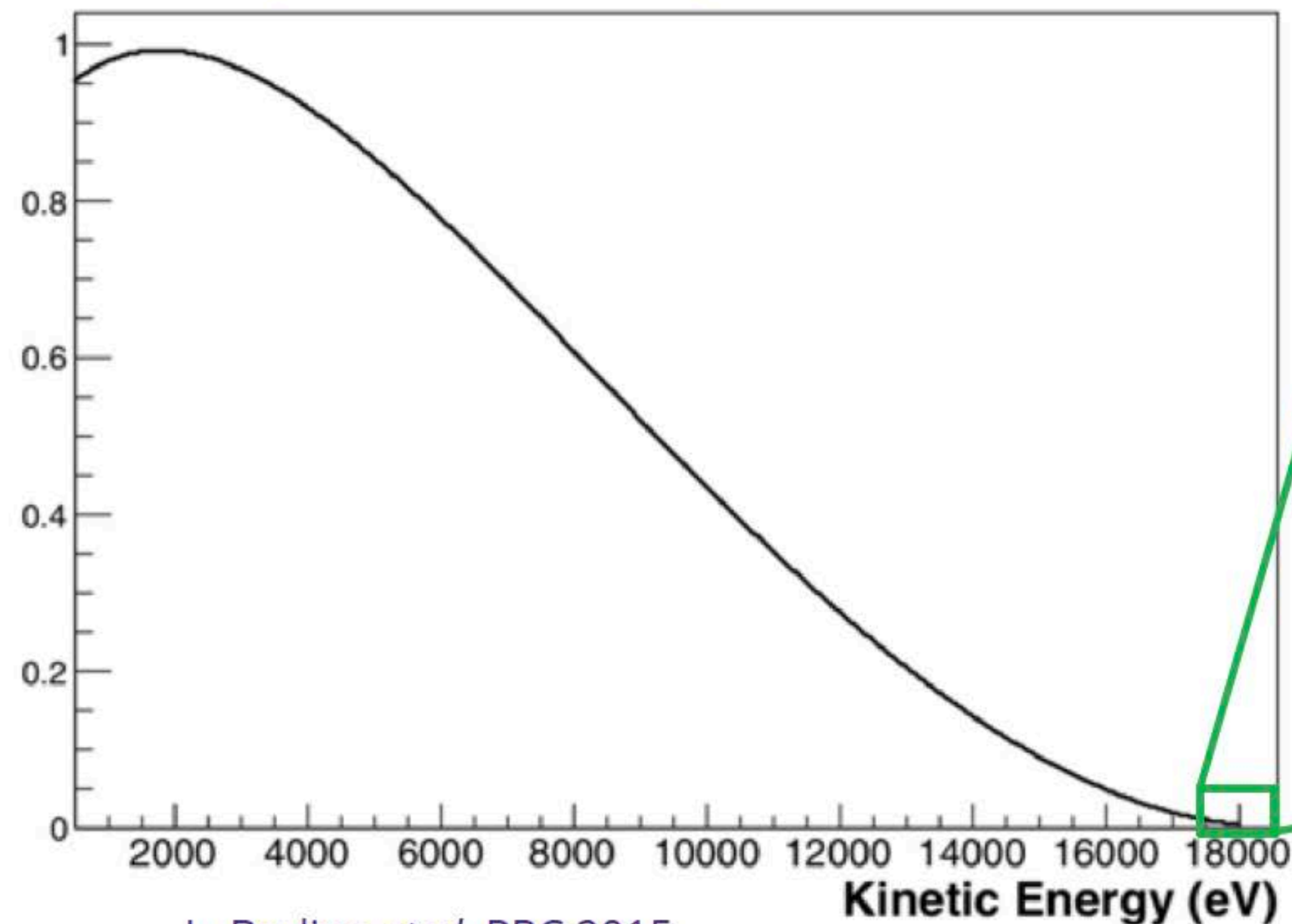
Finite neutrino mass modifies the decay electron spectrum!

With **Coulomb distortion** for T_{nuc} and **neutrino mixing**:

Idealized situation:

1. Super-allowed β -decay of *isolated* atom
2. Single neutrino mass state

$$\frac{dN}{dE_e} = \frac{G_F^2 m_e^5 \cos^2 \theta_C}{2\pi^3 \hbar^7} |M_{\text{nuc}}|^2 F(Z, E_e) p_e (E_e + m_e) \sum_i |U_{ei}|^2 (E_{\text{max}} - E_e) \times \sqrt{(E_{\text{max}} - E_e)^2 - m_{\nu i}^2} \cdot \Theta(E_{\text{max}} - E_e - m_{\nu i})$$



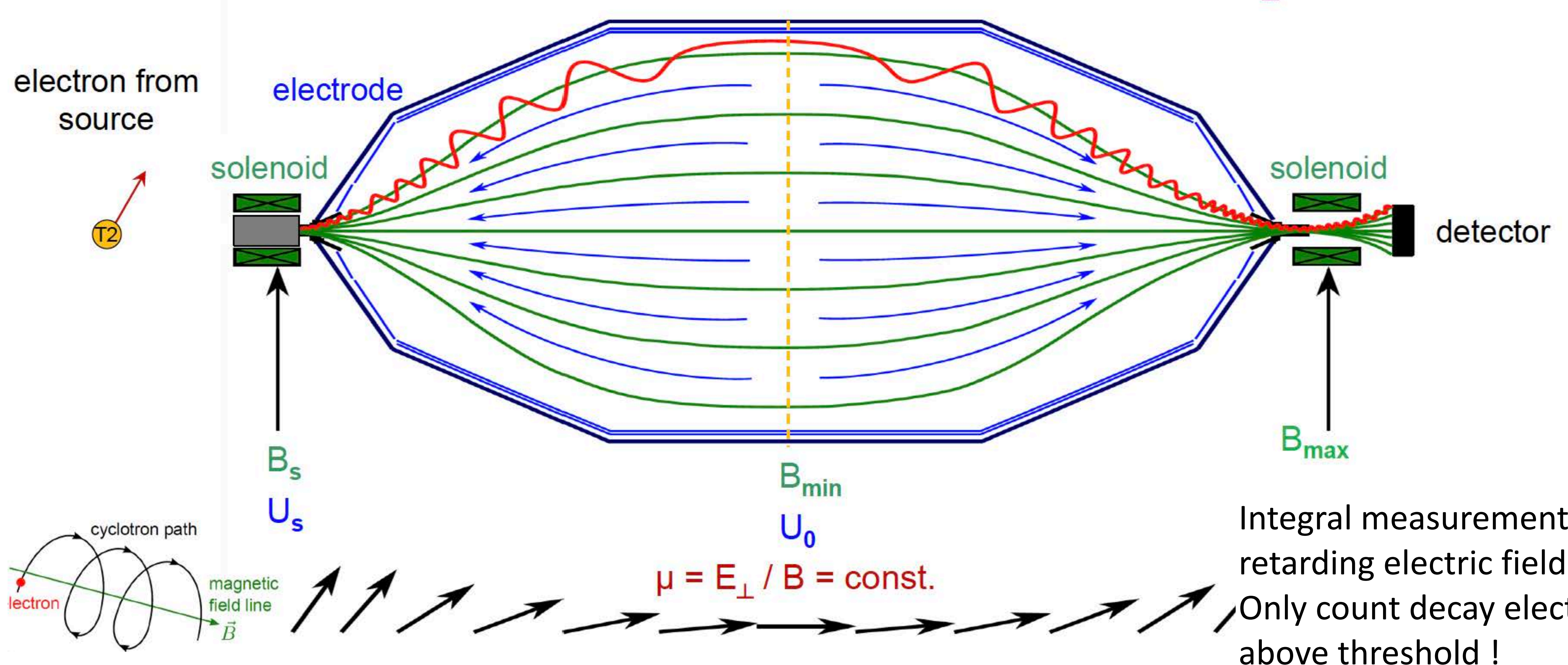
State-of-the-art Technology for T2: MAC-E filter

Magnetic Adiabatic Collimation with Electrostatic filter

A. Picard et al., Nucl. Instr. Meth. B 63 (1992)

Karlsruhe Institute of Technology

■ **Magnetic Adiabatic Collimation & Electrostatic Filter:** adiabatic conversion $E_{\perp} \rightarrow E_{\parallel}$



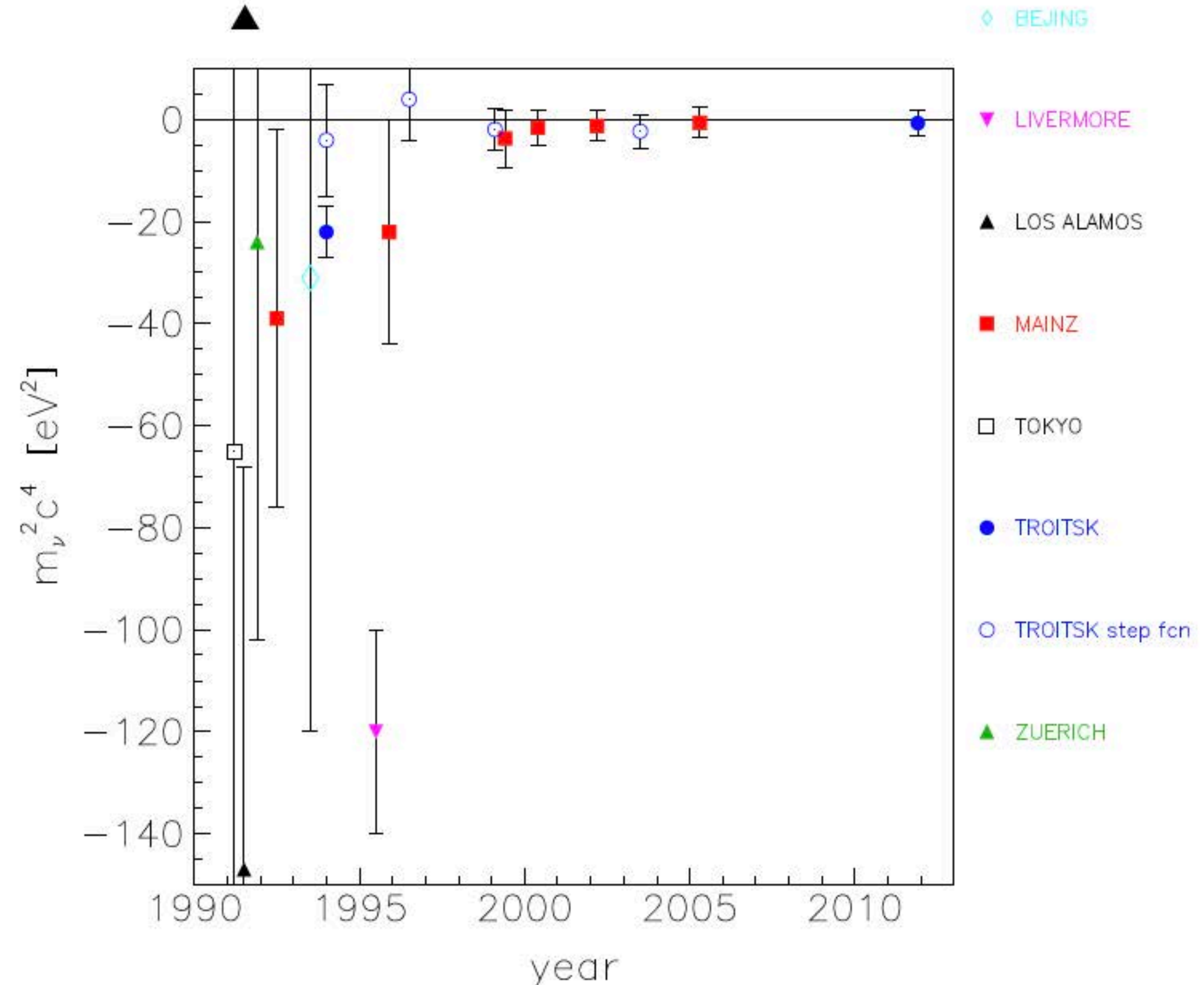
Neutrino Mass from Tritium Experiments

Tritium spectrometers have been workhorse of endpoint measurements for decades

Complicated molecular final states and incomplete de-excitation has yielded non-real mass values

More recent improved theory calculations can correct select previous results

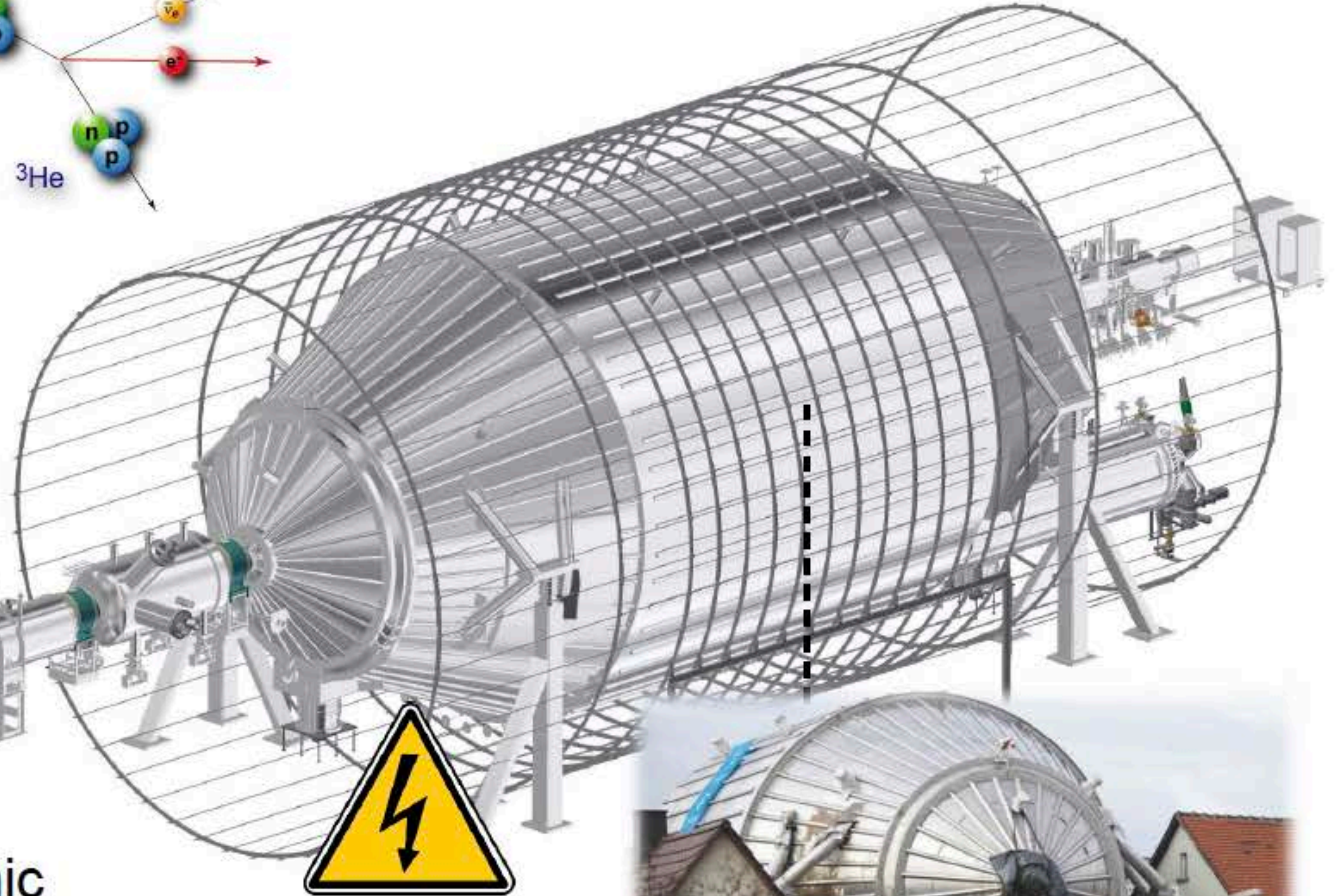
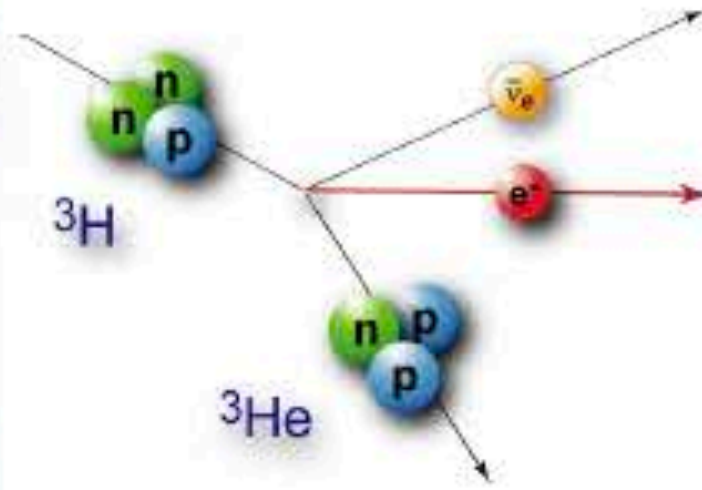
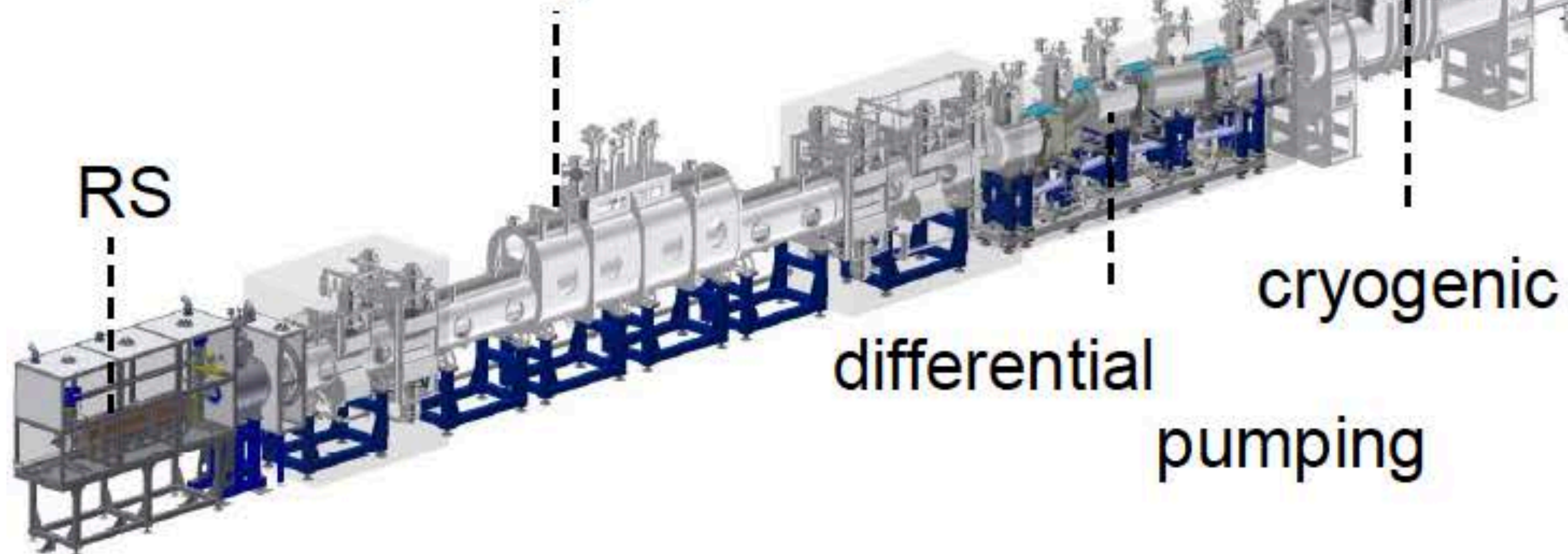
L. Bodine *et al.* PRC **91** 035505 (2015)



KATRIN Overview



Windowless Gaseous
Tritium Source cryostat



Main Spectrometer

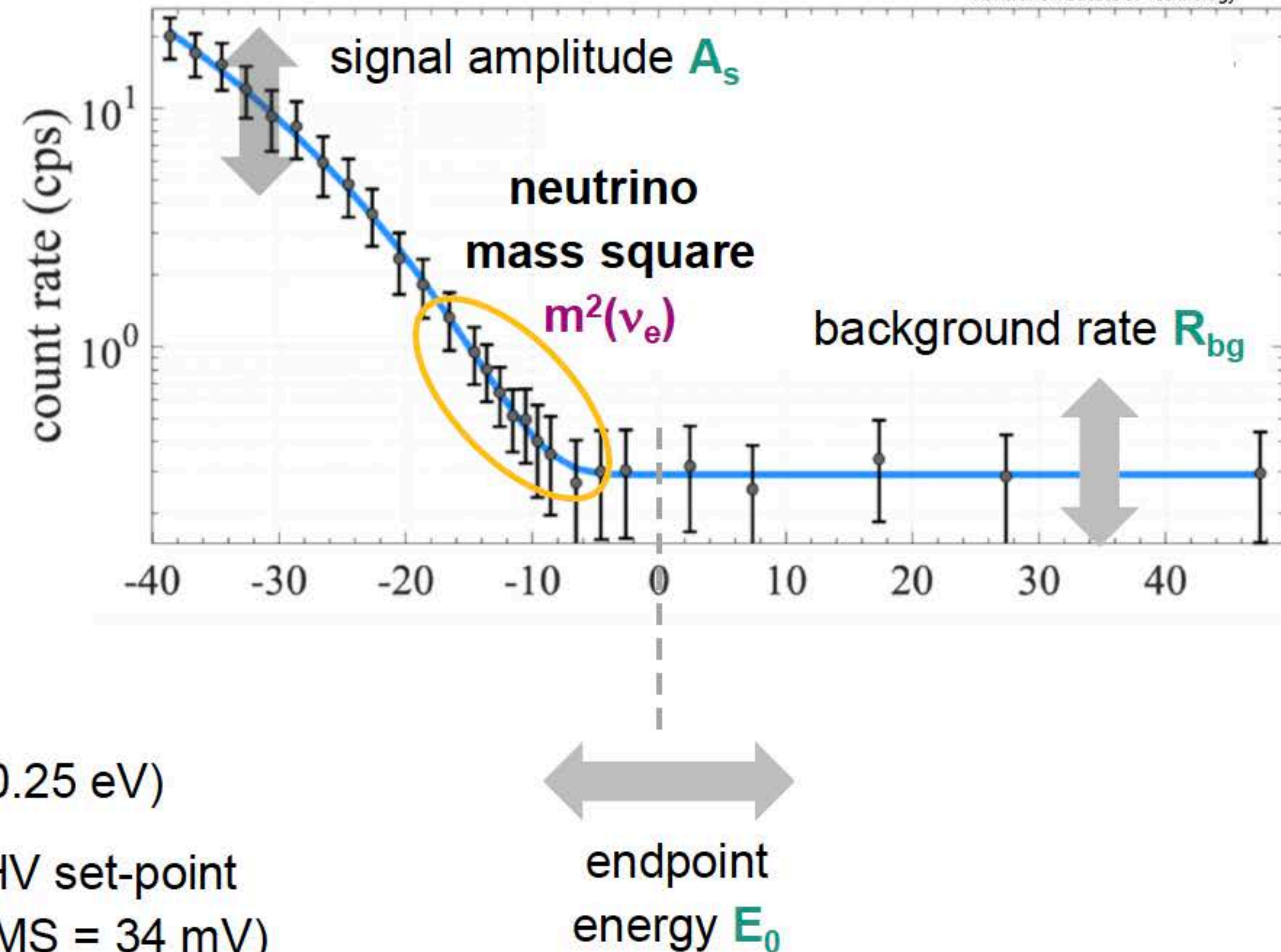


- fit of integrated experimental energy spectrum to theoretical model with **4 free parameters**

- leave parameters A_s and E_0 unconstrained
- 'shape-only' fit

- merged data set

- combine all 274 scans: excellent stability of all fitted β -decay endpoints E_0 ($\sigma = 0.25$ eV)
- ⇒ "stacking" of events at mean HV set-point (excellent reproducibility: RMS = 34 mV)



Drexlin, TAUP 2019

■ High-statistics β -spectrum

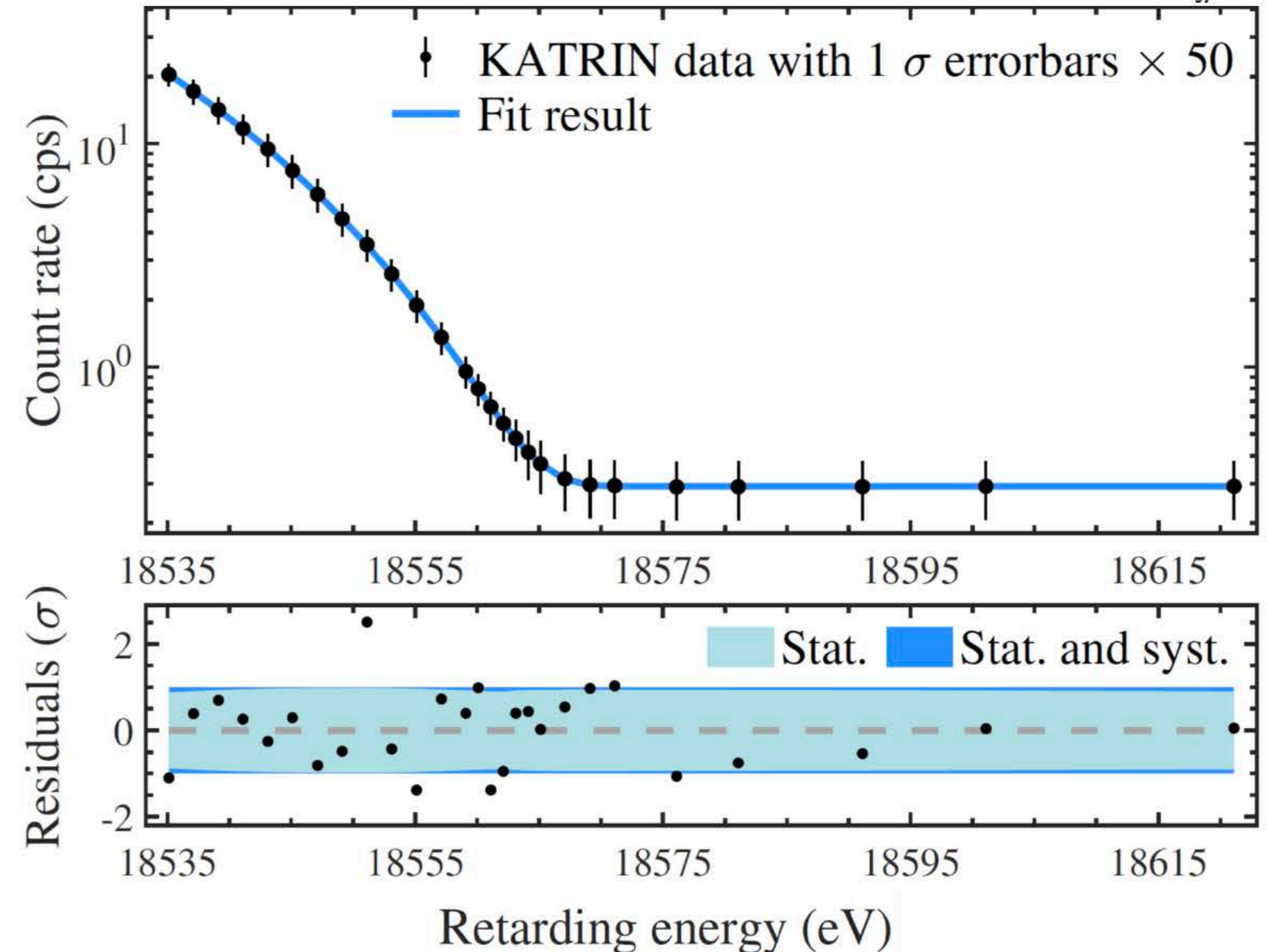
- 2 million events in
in 90-eV-wide interval
(522 h of scanning)
- excellent goodness-of-fit
 $\chi^2 = 21.4$ for 23 d.o.f.
(p-value = 0.56)

■ bias-free analysis

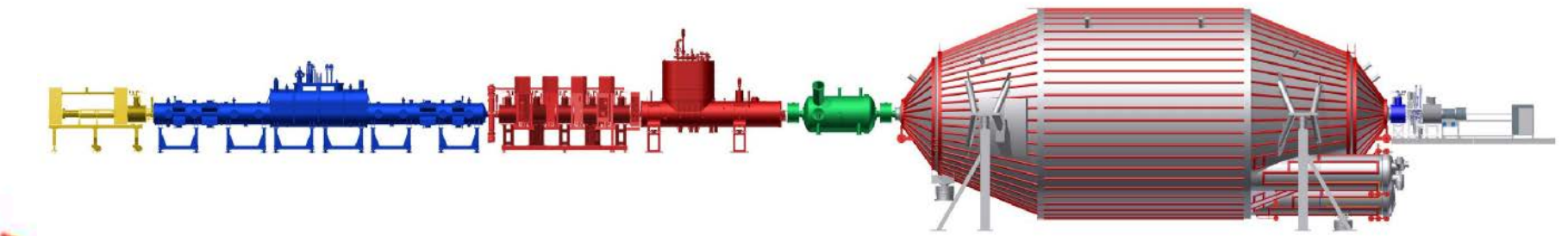
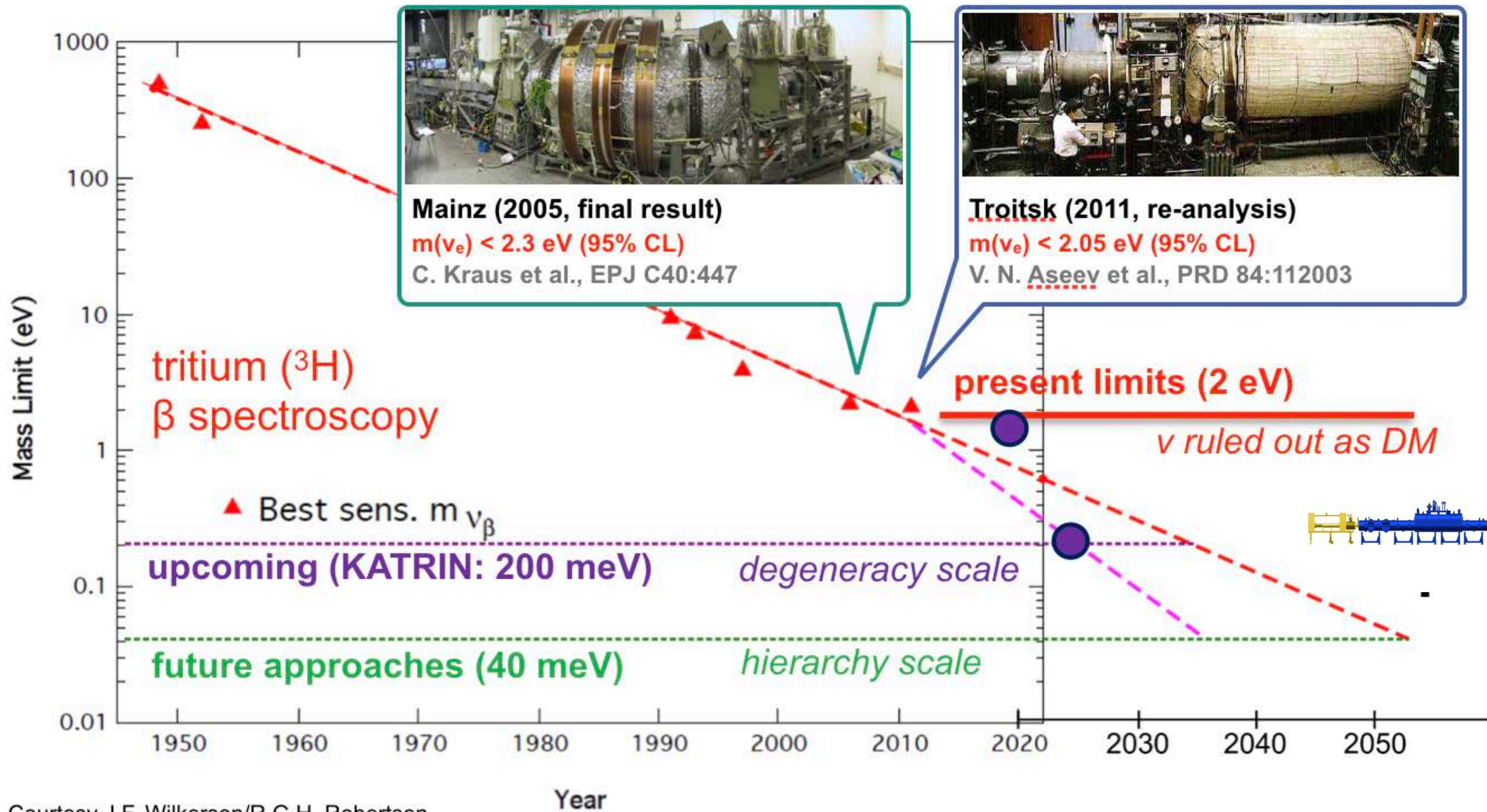
- blinding of FSD
- full analysis chain first on
MC data sets
- final step: unblinded FSD
for experimental data

■ ν -mass and E_0 : best fit results

$$m^2(\nu_e) = \left(-1.0^{+0.9}_{-1.1} \right) \text{eV}^2 \text{ (90\% CL)}$$



Direct Neutrino Mass Searches

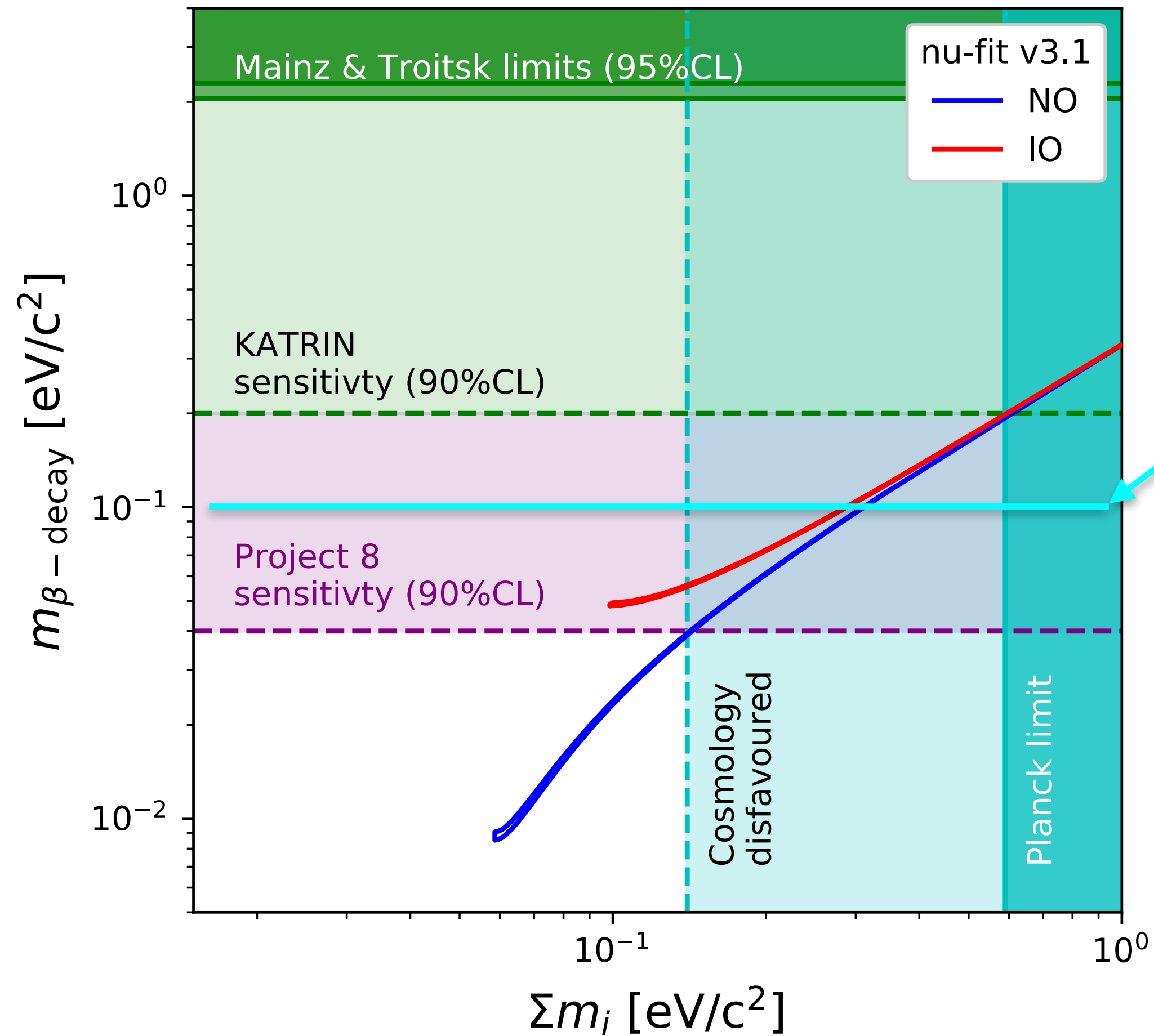


Schloesser

Courtesy J.F. Wilkerson/R.G.H. Robertson

Tritium experiments define the mass limit. Where will we be in 2030?

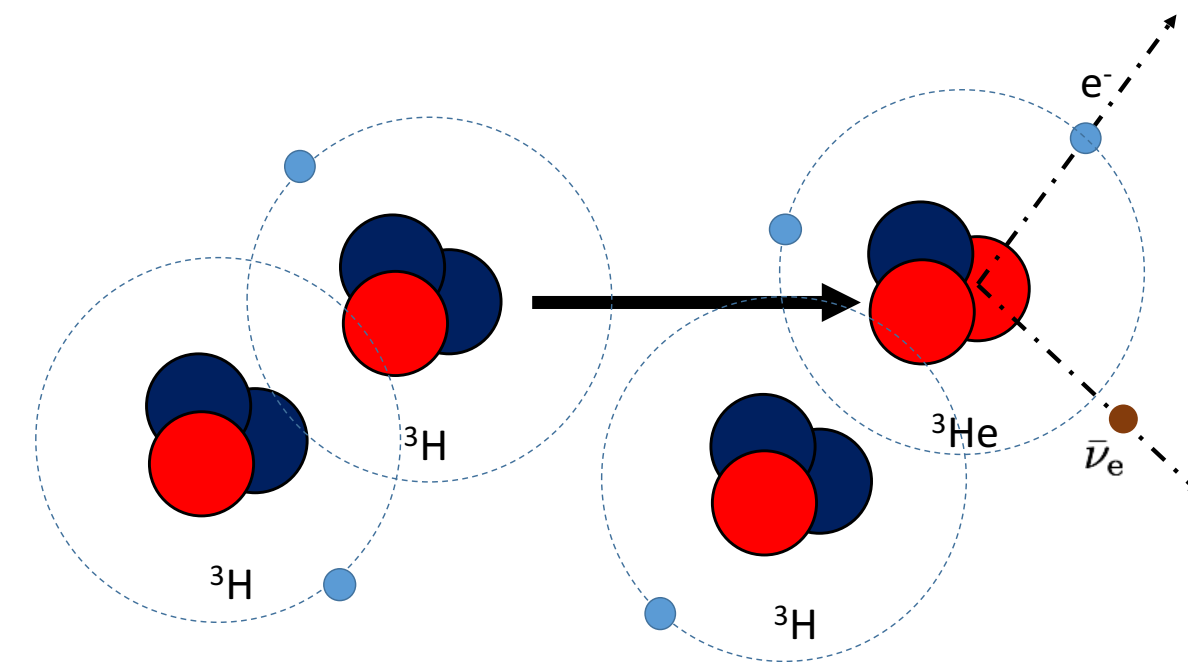
Challenges for Sensitivity Beyond KATRIN



S. Boeser with input from www.nu-fit.org

Source intrinsic limitation:

Irreducible excitation of ro-vibrational initial and final states of T_2 molecule.

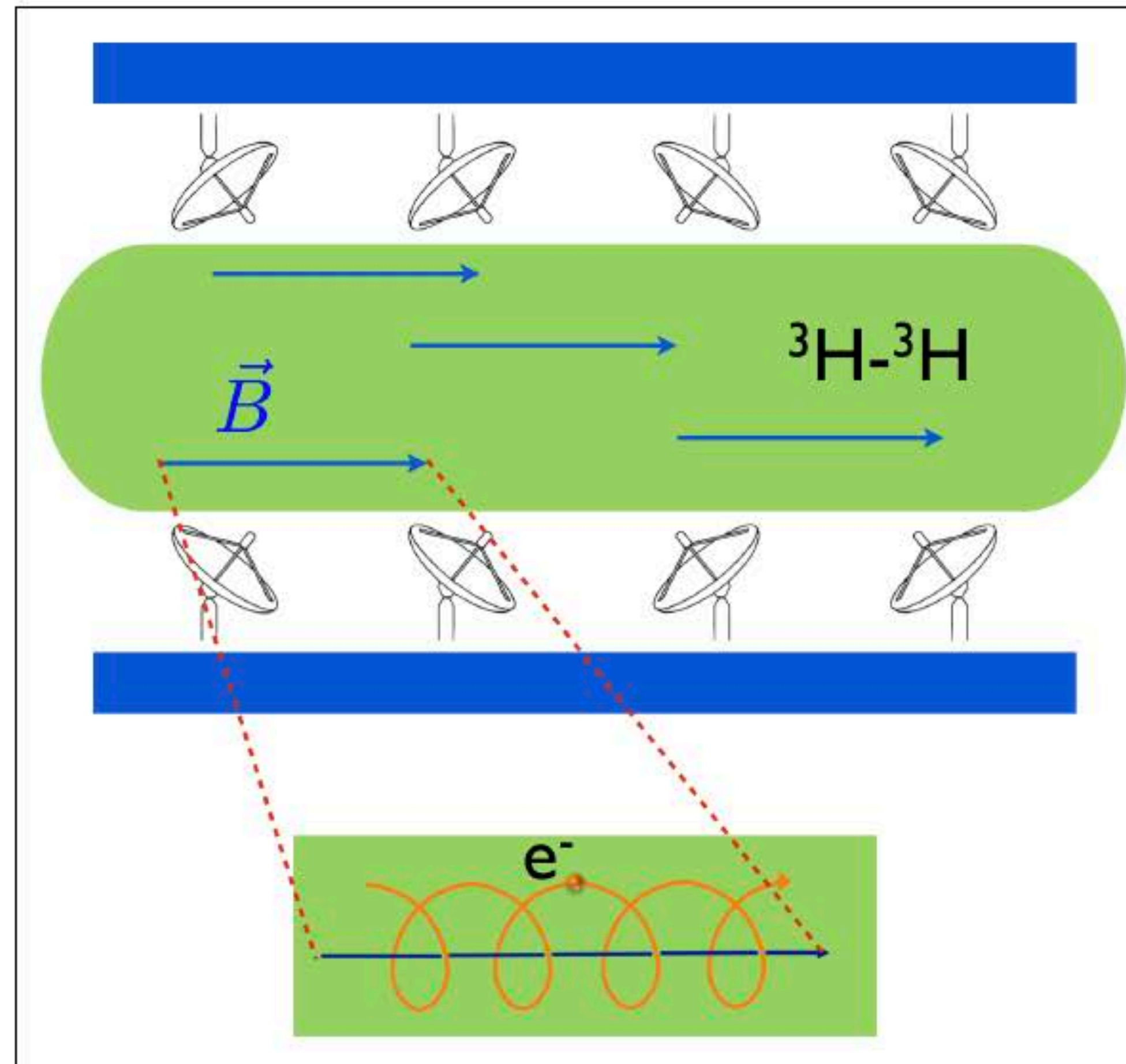


Technical challenges:

1. Statistical nature of e^- scattering in gas column
2. e^- can be scattered into angular acceptance cone of the MAC-E filter
3. Rydberg atoms as background source

Project 8: Cyclotron Radiation Emission Spectroscopy of T2

1. Start with an enclosed volume
2. Fill with tritium gas
3. Add a magnetic field



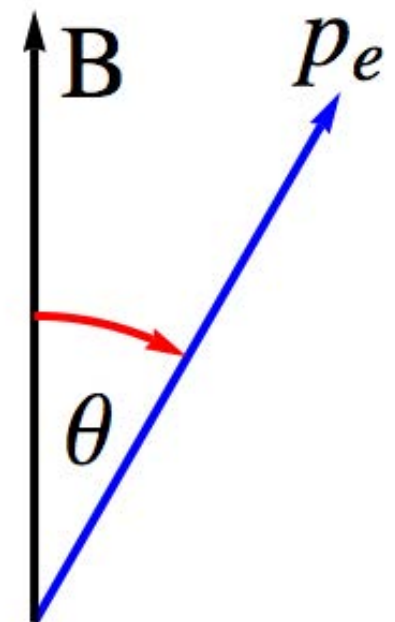
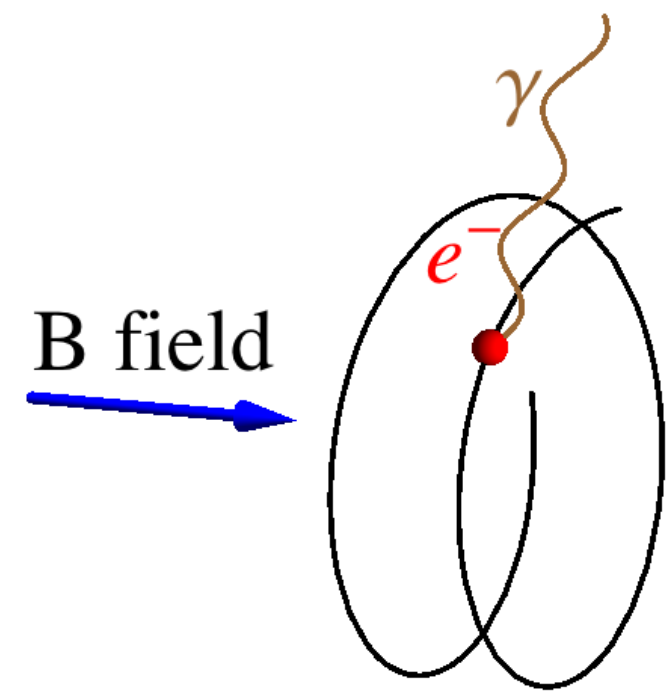
4. Decay electrons spiral around field lines
5. Add antennas to detect the cyclotron radiation

B. Monreal and J. Formaggio, Phys. Rev. D80 051301 (2009)

Project 8: Cyclotron Radiation Emission Spectroscopy of T2

Novel approach: J. Formaggio and B. Monreal, Phys. Rev. D 80:051301 (2009)

- Cyclotron radiation from single electrons
- Source transparent to microwave radiation
- No e⁻ transport from source to detector
- Highly precise frequency measurement



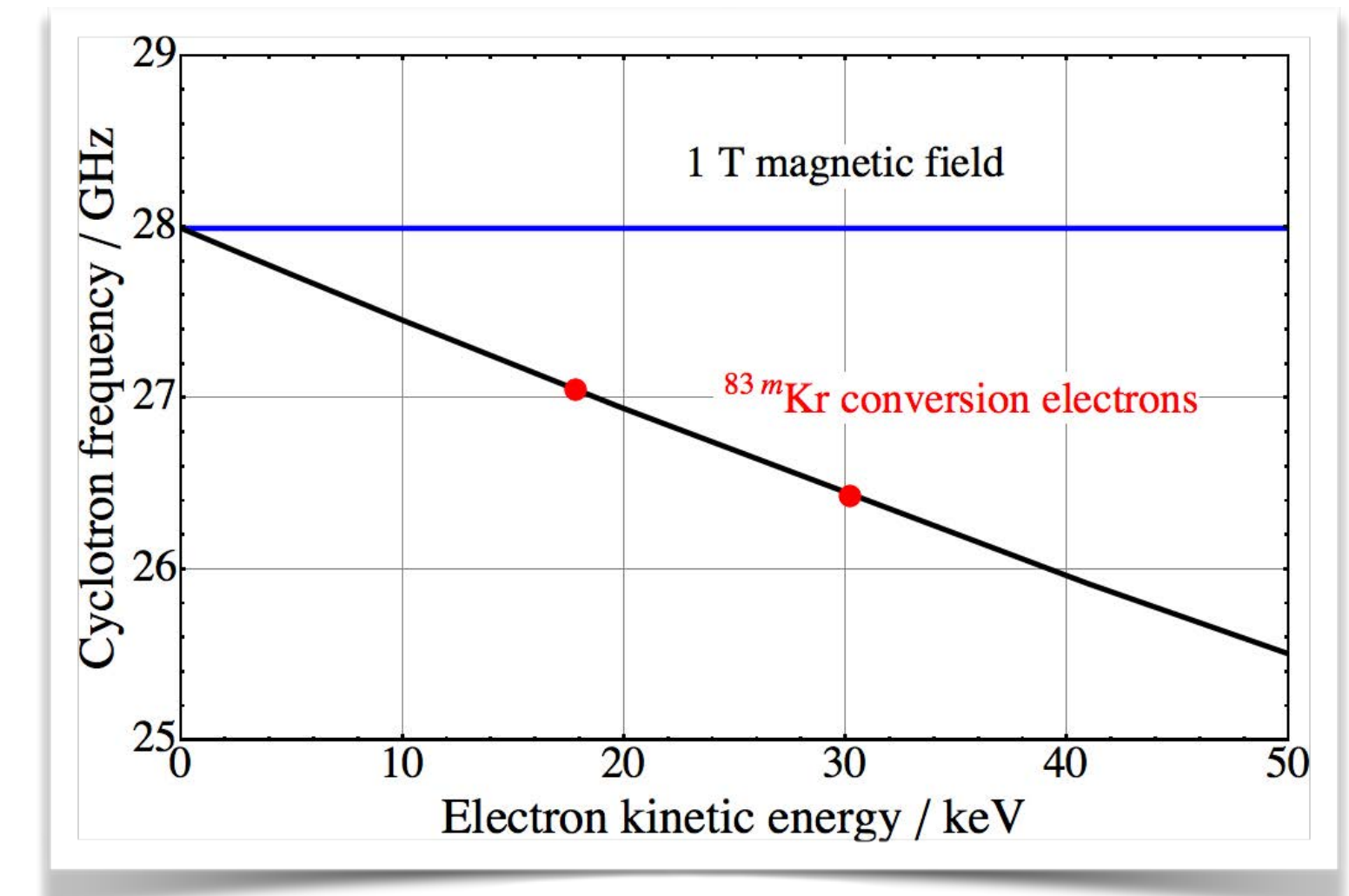
$$f_c = \frac{f_{c,0}}{\gamma} = \frac{1}{2\pi} \frac{eB}{m_e + E_{\text{kin}}/c^2}$$

$$P(E_{\text{kin}}, m, \theta) = \frac{1}{4\pi\epsilon_0} \frac{2}{3} \frac{e^4}{m^4 c^5} B^2 (E_{\text{kin}}^2 + 2 E_{\text{kin}} m c^2) \sin^2 \theta$$

$$P(17.8 \text{ keV}, 90^\circ, 1 \text{ T}) = 1 \text{ fW}$$

$$P(30.2 \text{ keV}, 90^\circ, 1 \text{ T}) = 1.7 \text{ fW}$$

Small but readily detectable with state of the art detectors



Project 8 in a waveguide with magnetic trap

Energy resolution vs. frequency resolution

$$\frac{\Delta E_{\text{kin}}}{E_{\text{kin}}} = \left(1 + \frac{m_e c^2}{E_{\text{kin}}}\right) \frac{\Delta \nu_c}{\nu_c}$$

≈ 28 for 18.6 keV electron

$$\Delta E_{\text{kin}} \approx 0.2 \text{ eV} \rightarrow \frac{\Delta \nu_c}{\nu_c} \approx 4 \times 10^{-7}$$

$$\nu_c \approx 27 \text{ GHz} \rightarrow \Delta \nu_c \approx 11 \text{ kHz}$$

Frequency resolution vs. observation time

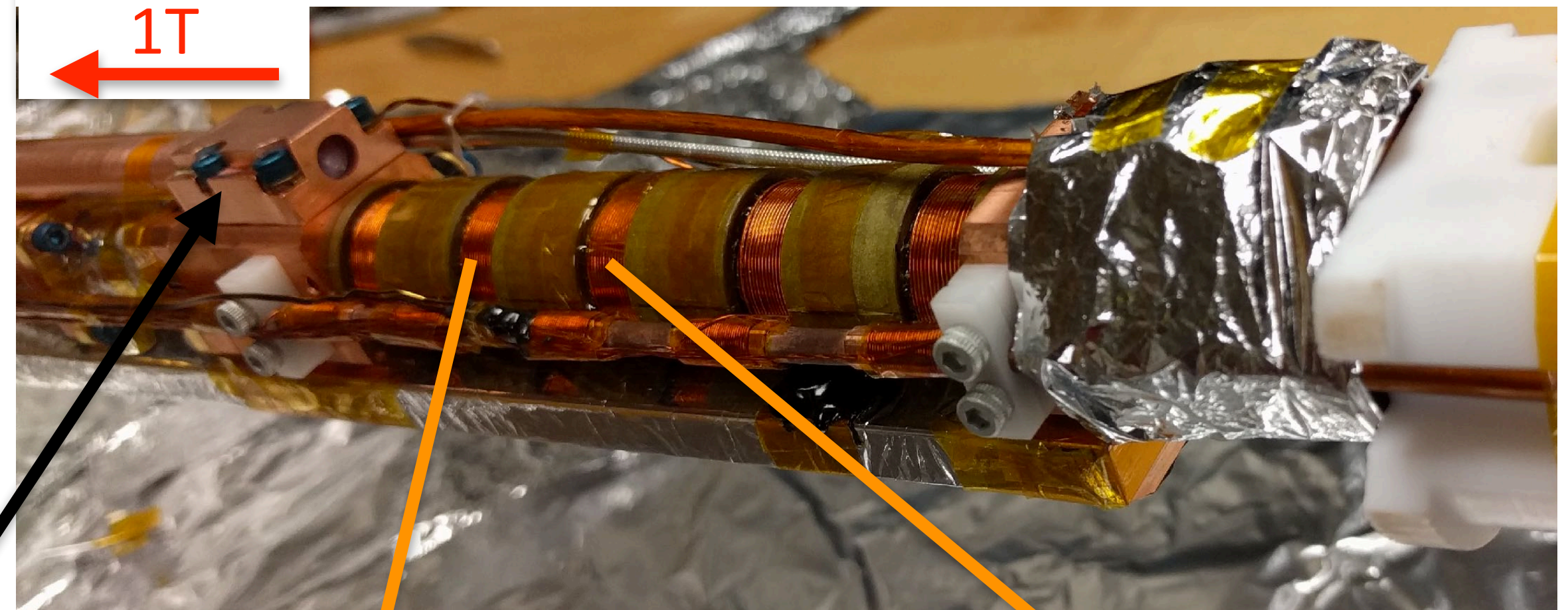
$$\Delta \nu_c \times t_{\text{obs}} \geq \frac{1}{2\pi} \rightarrow t_{\text{obs}} \geq 14 \mu\text{s}$$

→ **Need for a magnetic (no work) trap!**

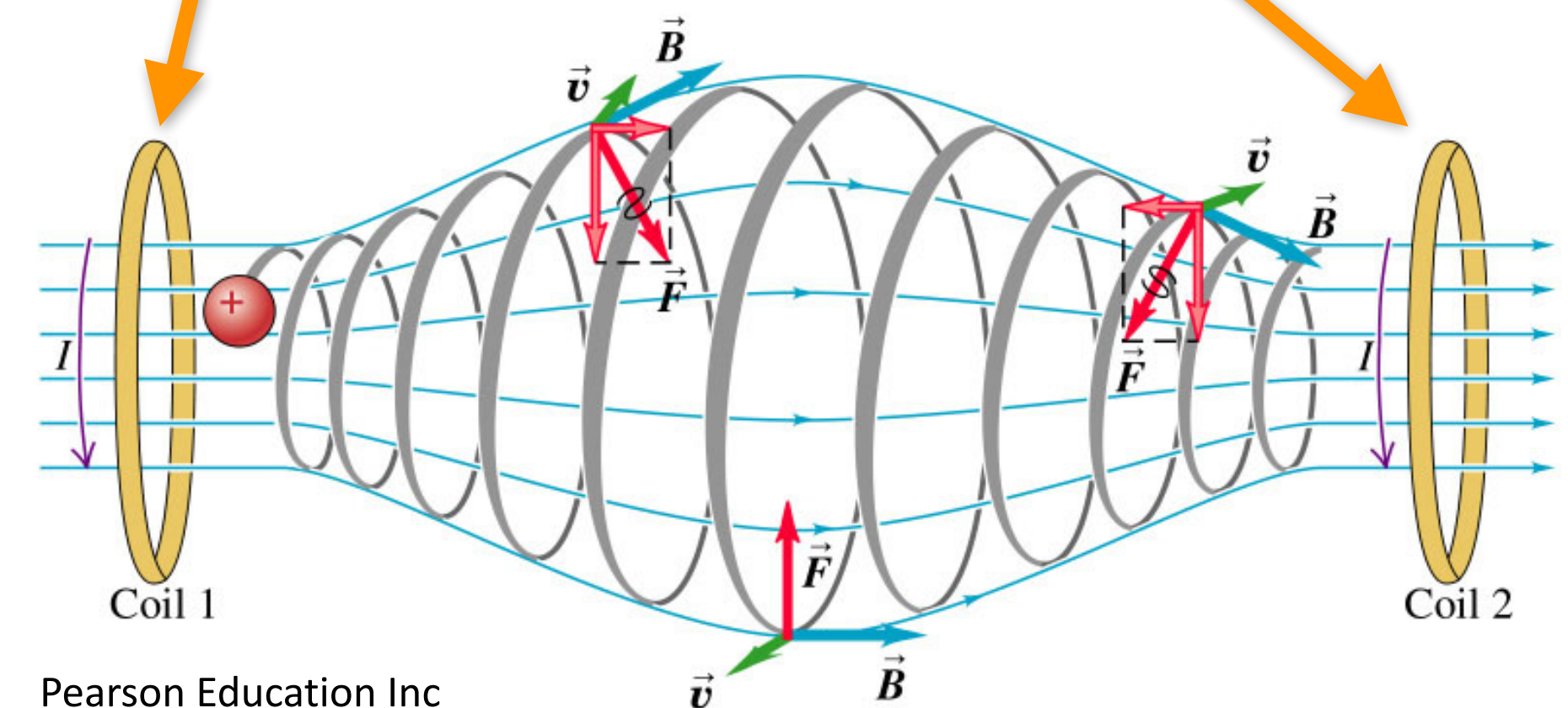
To cryogenic amplifier + heterodyne mixing stage + signal digitizer

^{83}mKr and T_2 supply line

Tritium compatible CRES cell

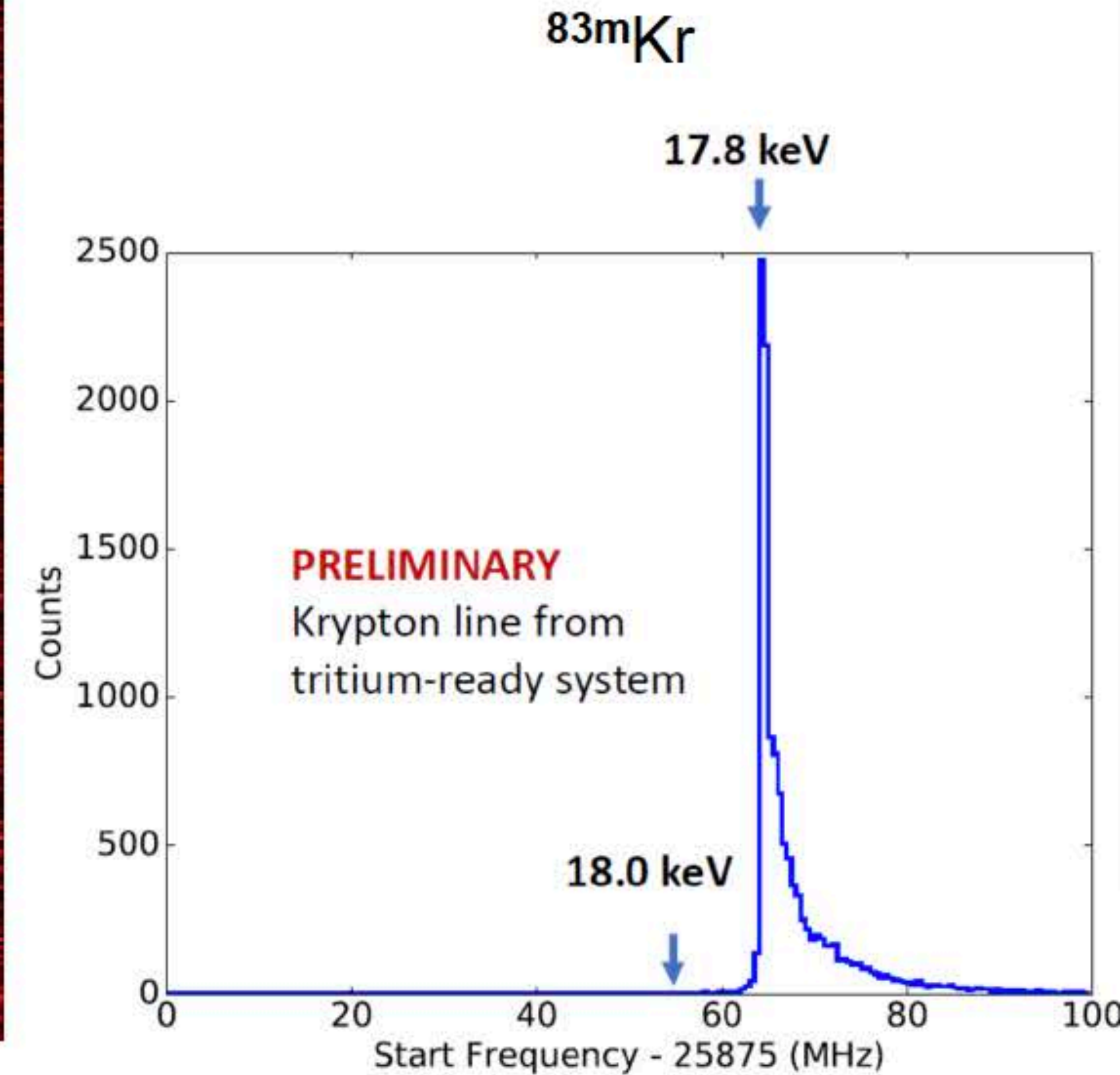
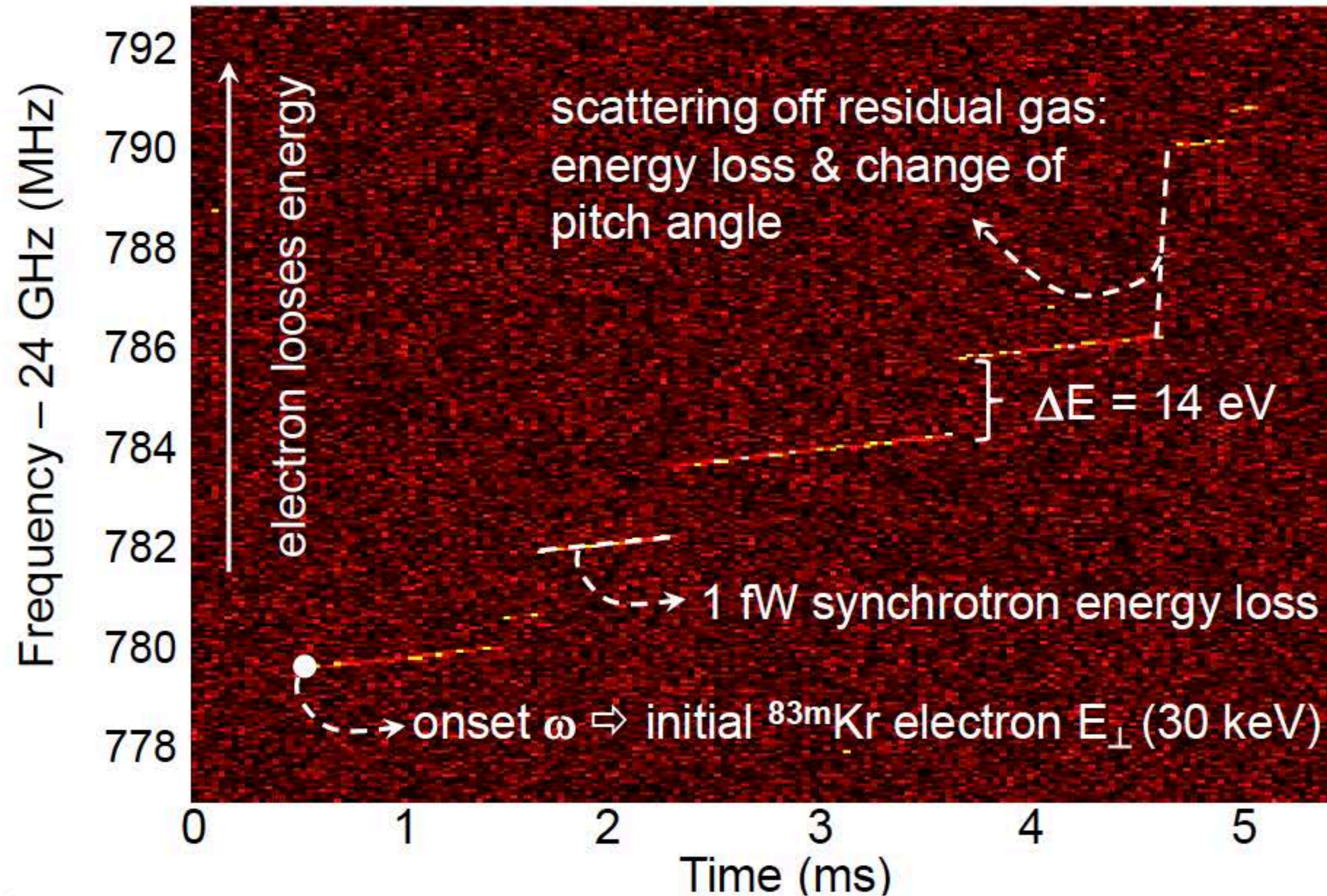


Double magnetic mirror trap



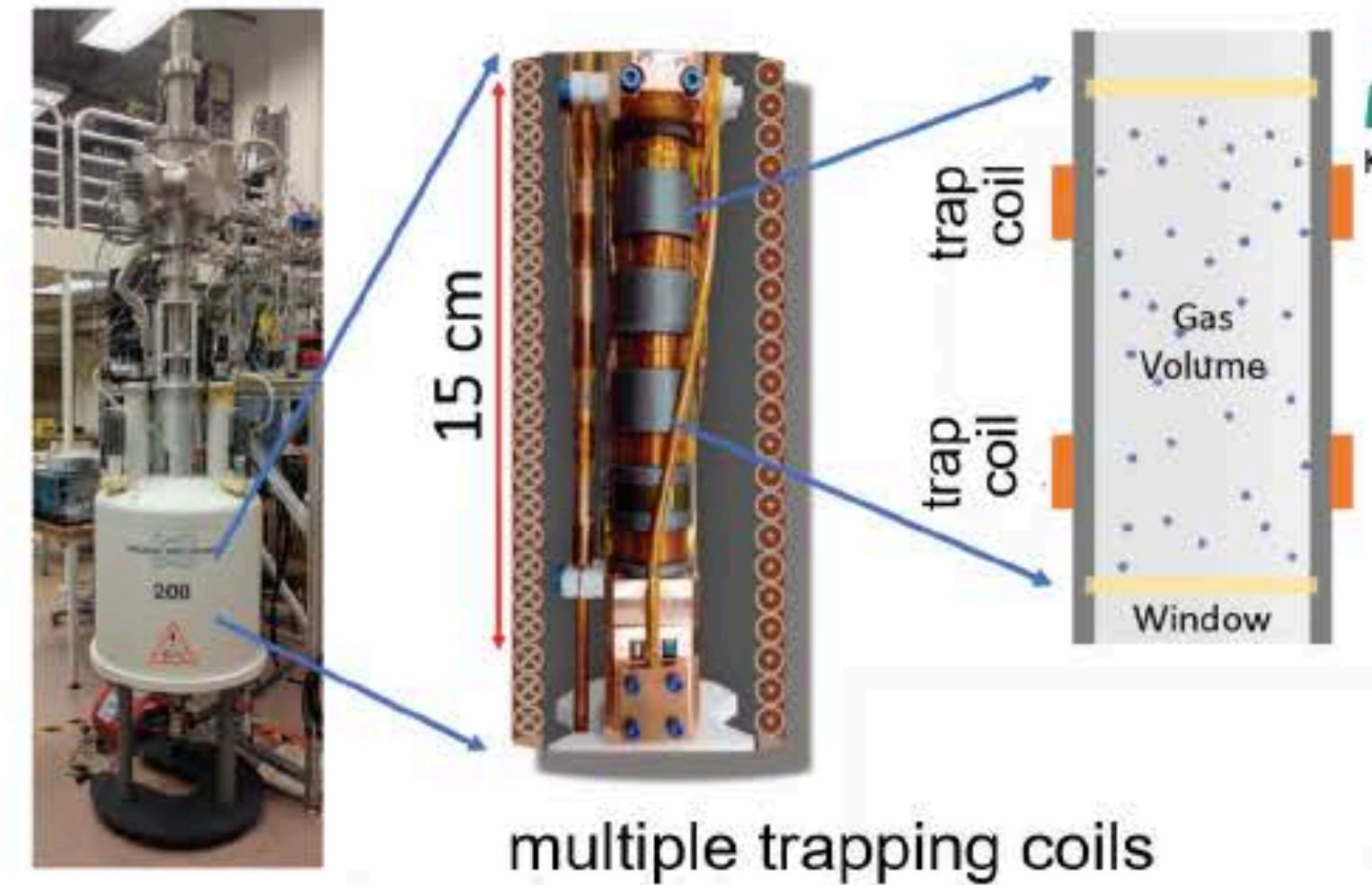
Project 8: Single Electron Detection

First Detection of Cyclotron Radiation from single keV electron

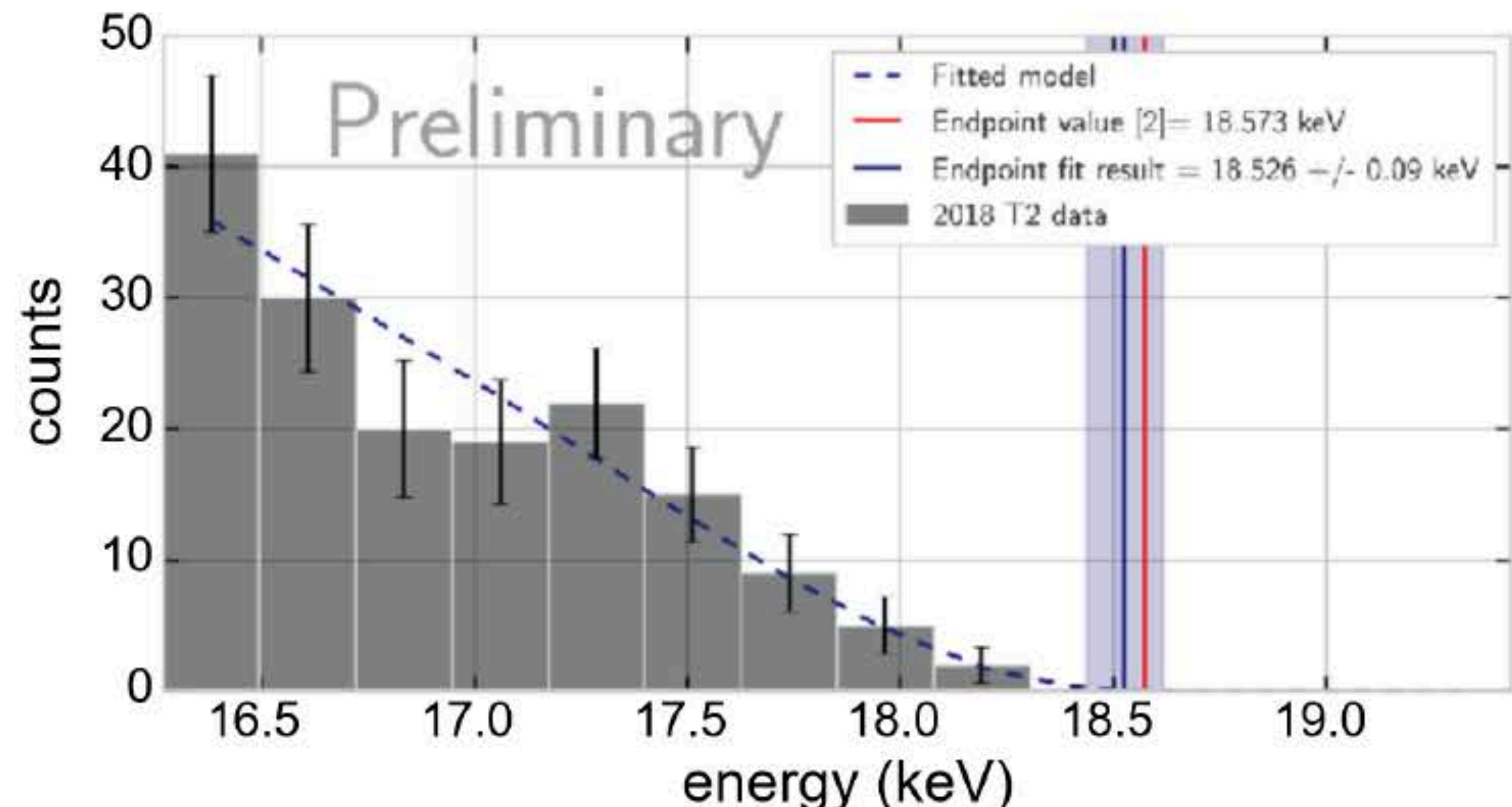


Project 8 - A Staged Approach

❶ Phase – I: 2010-2016 – proof-of-principle test measurements with ^{83m}Kr CRES observed for first time



❷ Phase – II: 2015-2019 - tritium CRES demonstrator
first tritium data 2018
several days of runs
fitted β -decay endpoint:
 $E_0 = (18.526 \pm 0.09) \text{ keV}$
new 2019 campaign to begin soon (100 d)

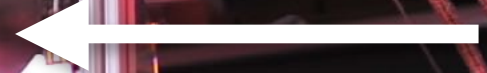


The newly assembled instrument

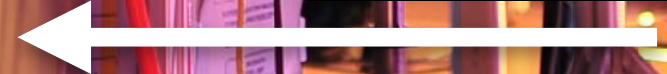
Kr/T₂ gas handling system attached



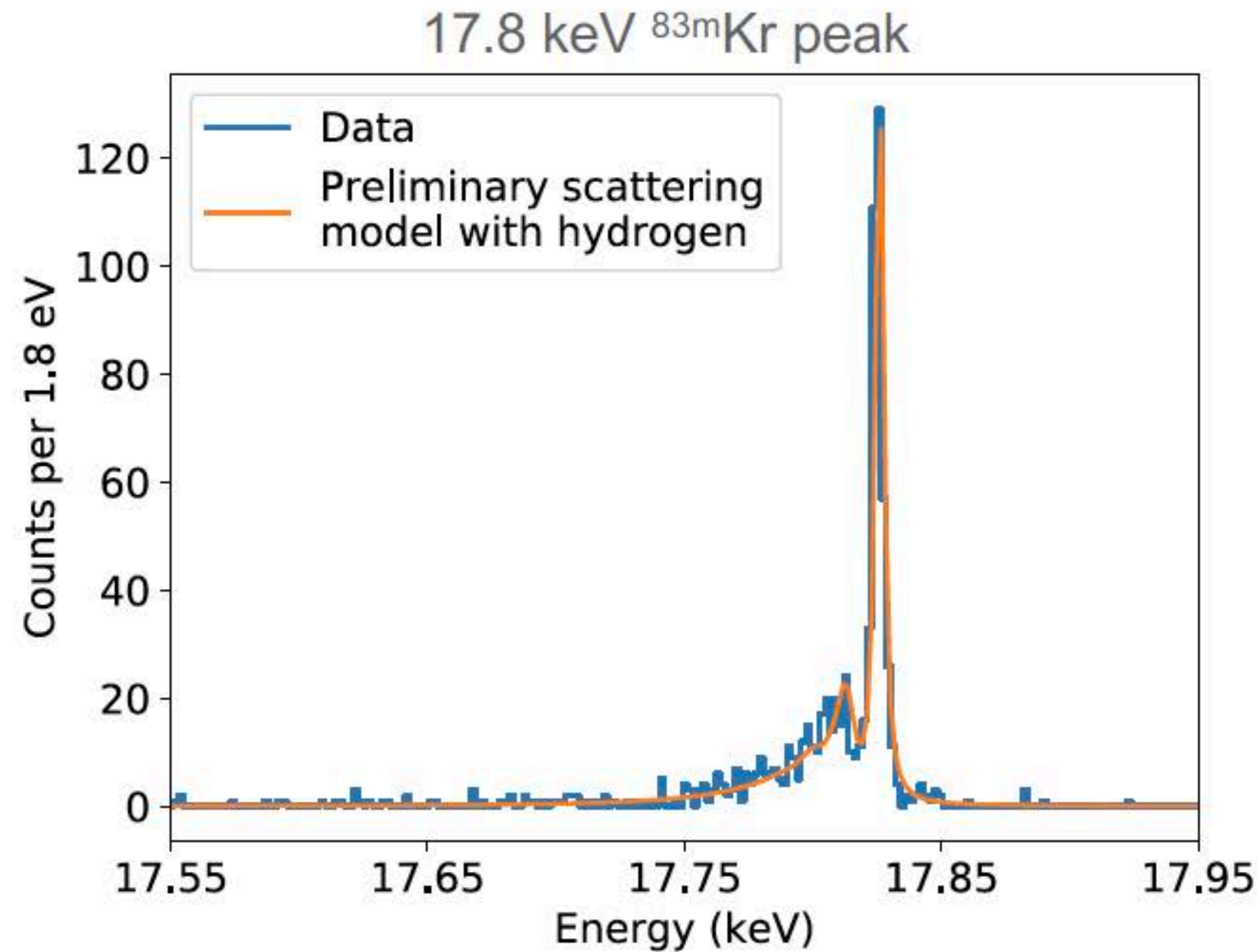
Insert cryostat



NMR magnet providing background magnetic field



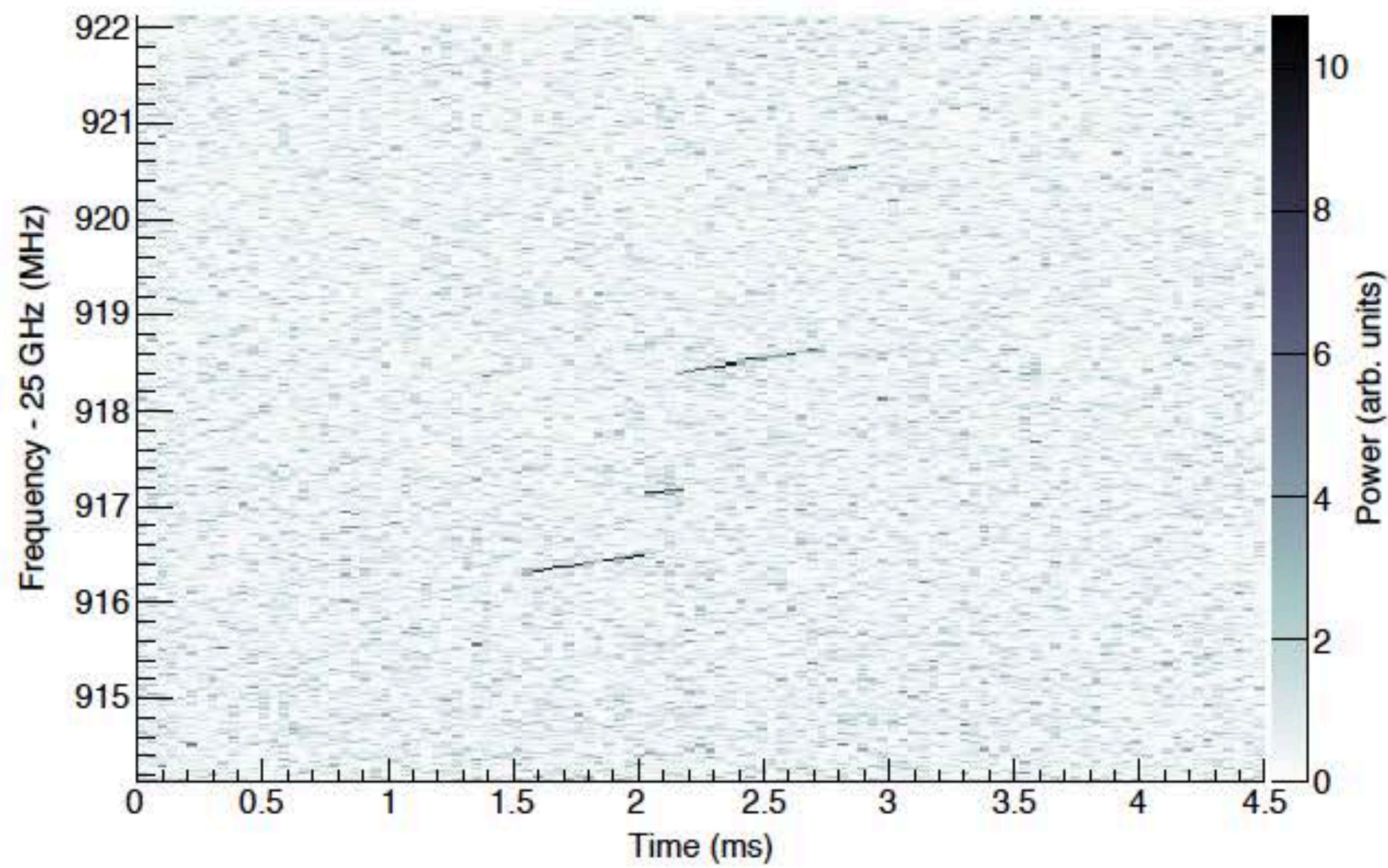
Project 8: A High Precision Measurement



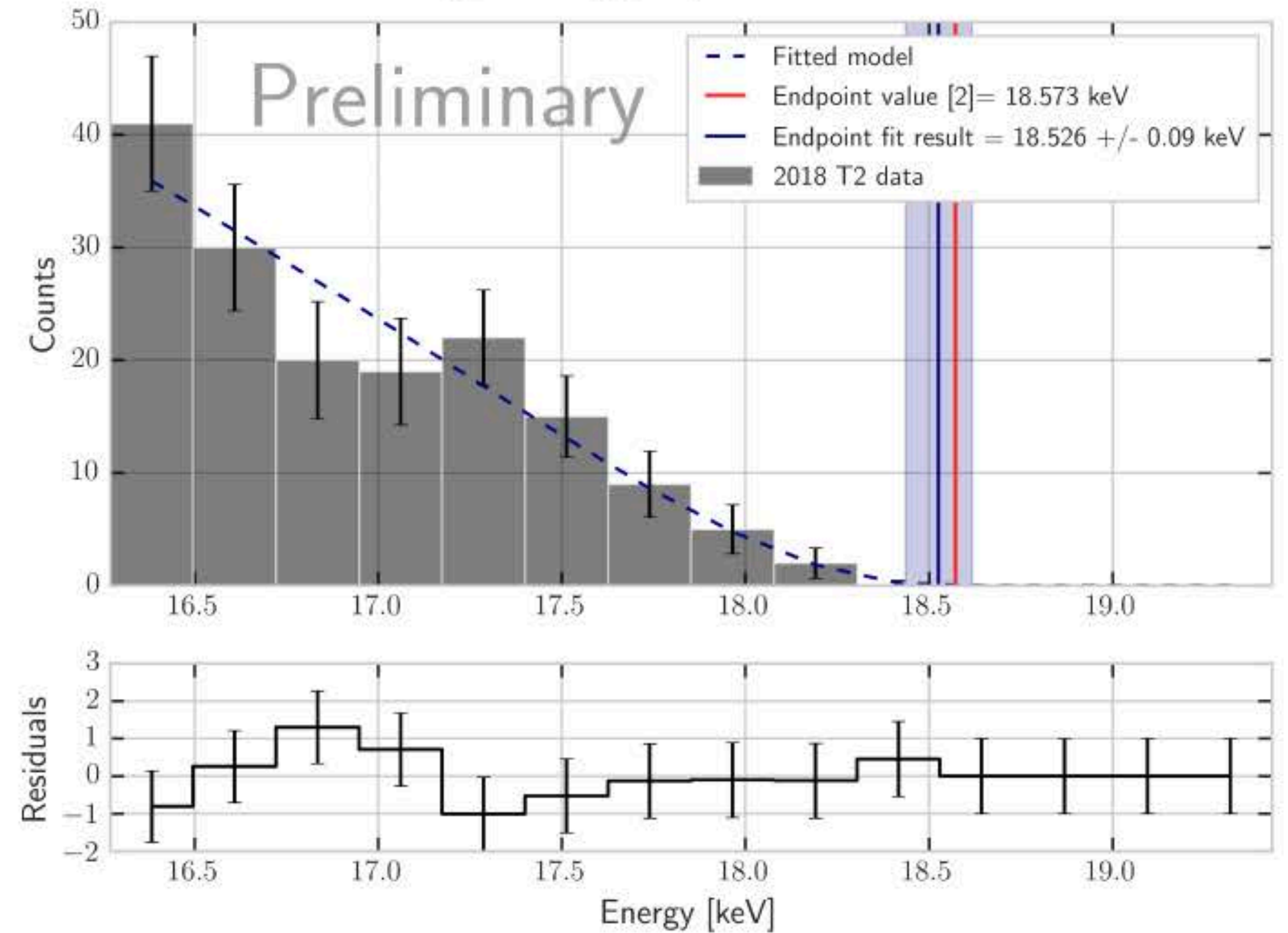
- Preliminary model includes intrinsic line width and Gaussian instrument resolution
- Line width: 2.8 ± 0.1 eV (FWHM)
- Instrumental width: **2.0 ± 0.5 eV** (FWHM)

Project 8 Phase II: First Observation of Tritium Events

First T₂ Electron Event



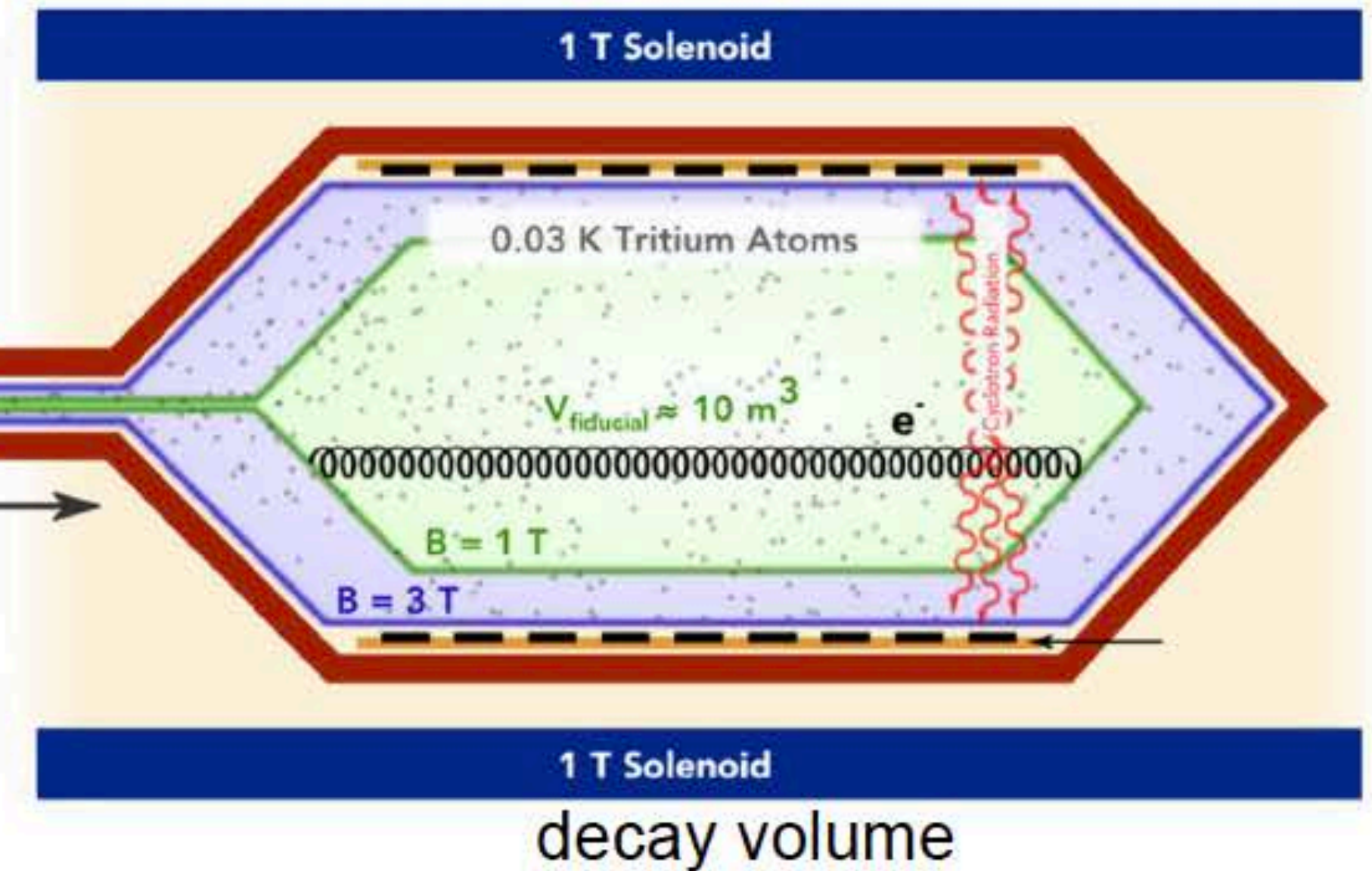
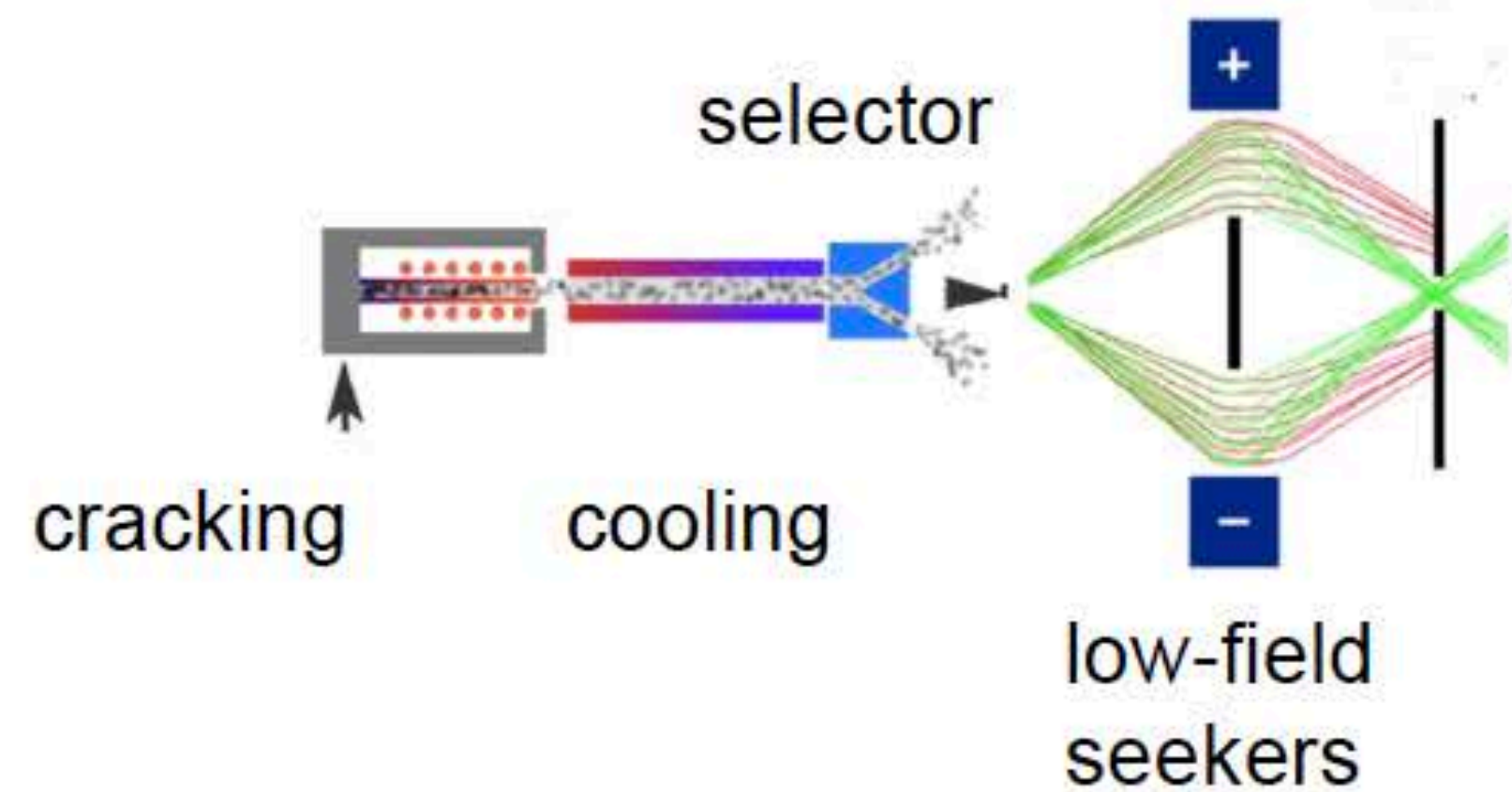
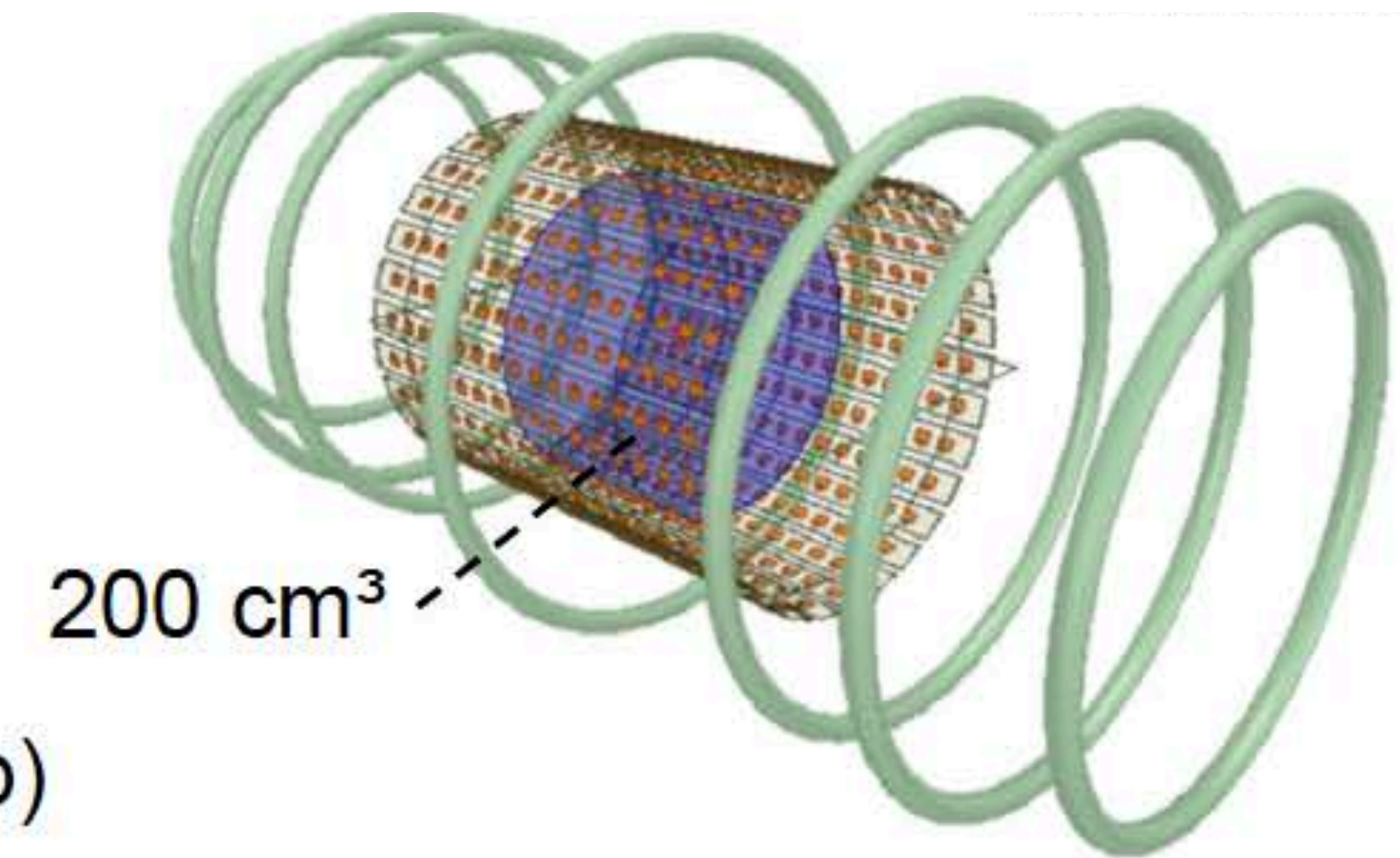
T₂ energy spectrum



Project 8 - The Future

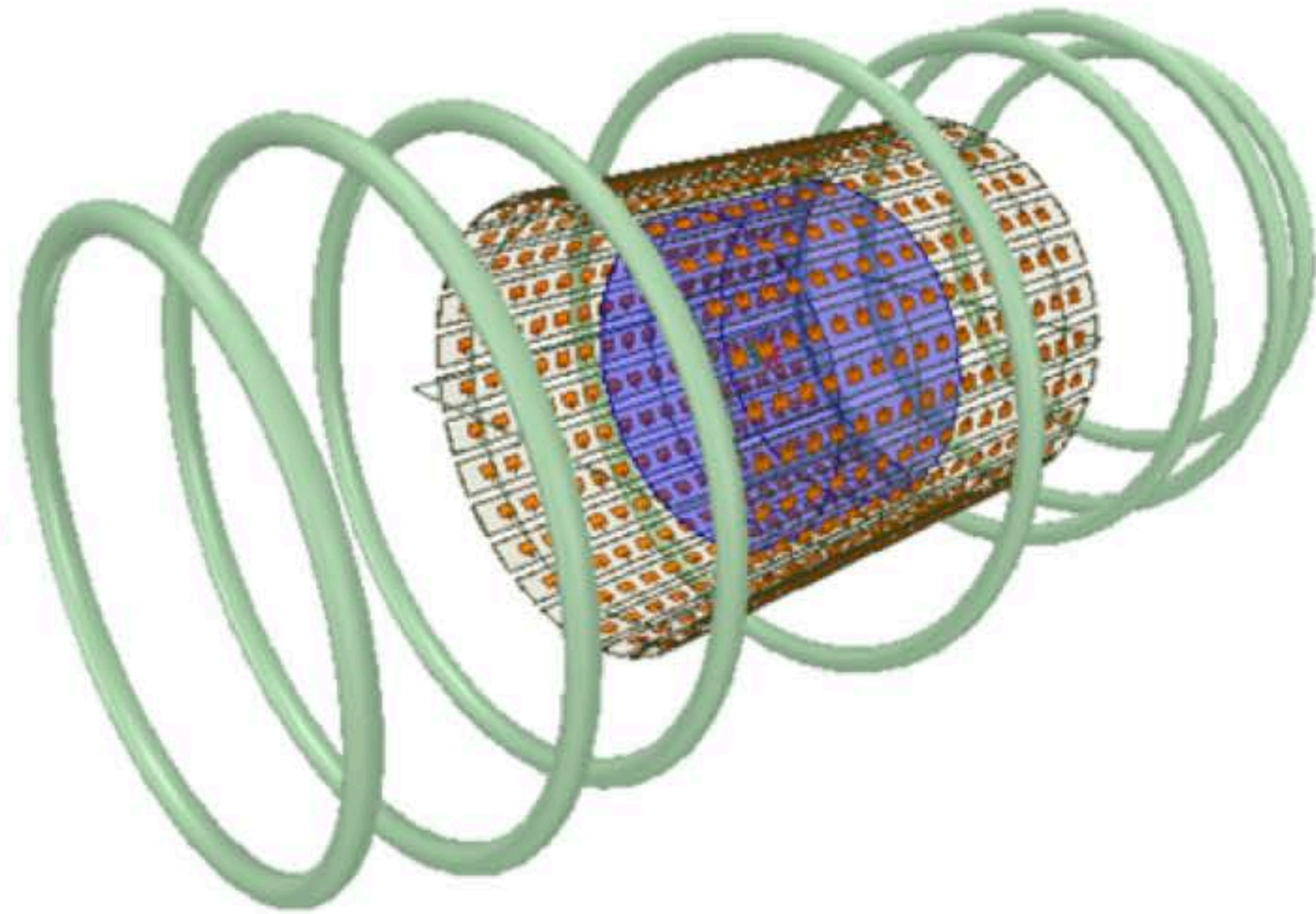
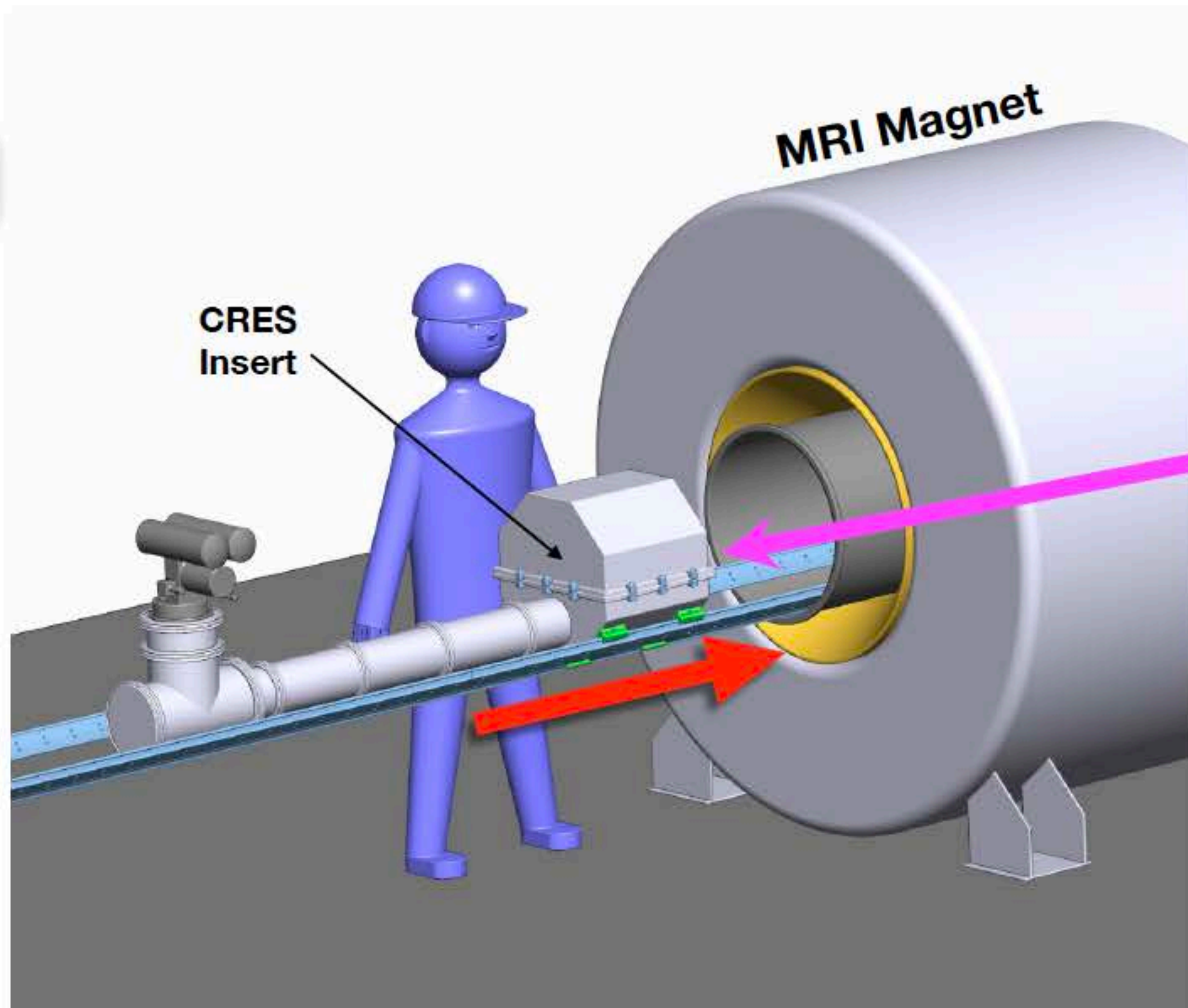
③ Phase – III: ... – a large volume demonstrator based on multi-antenna array in MRI tritium spectrum for $m(\nu_e) \sim 2$ eV

④ Phase – IV: ... – towards an atomic tritium source
R&D for an **atomic tritium source** (Ioffe trap)
goal: inverted mass hierarchy for $m(\nu_e)$



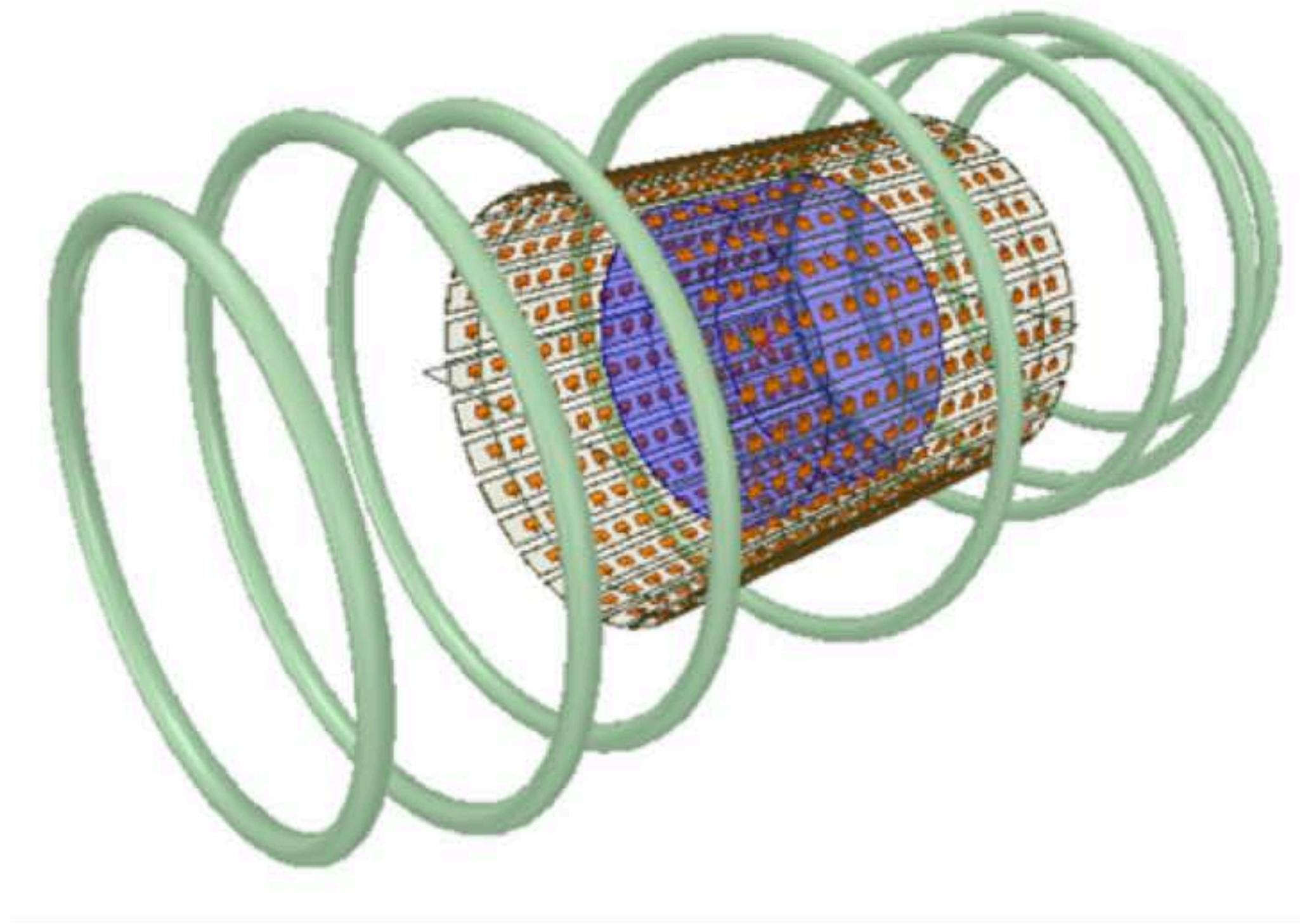
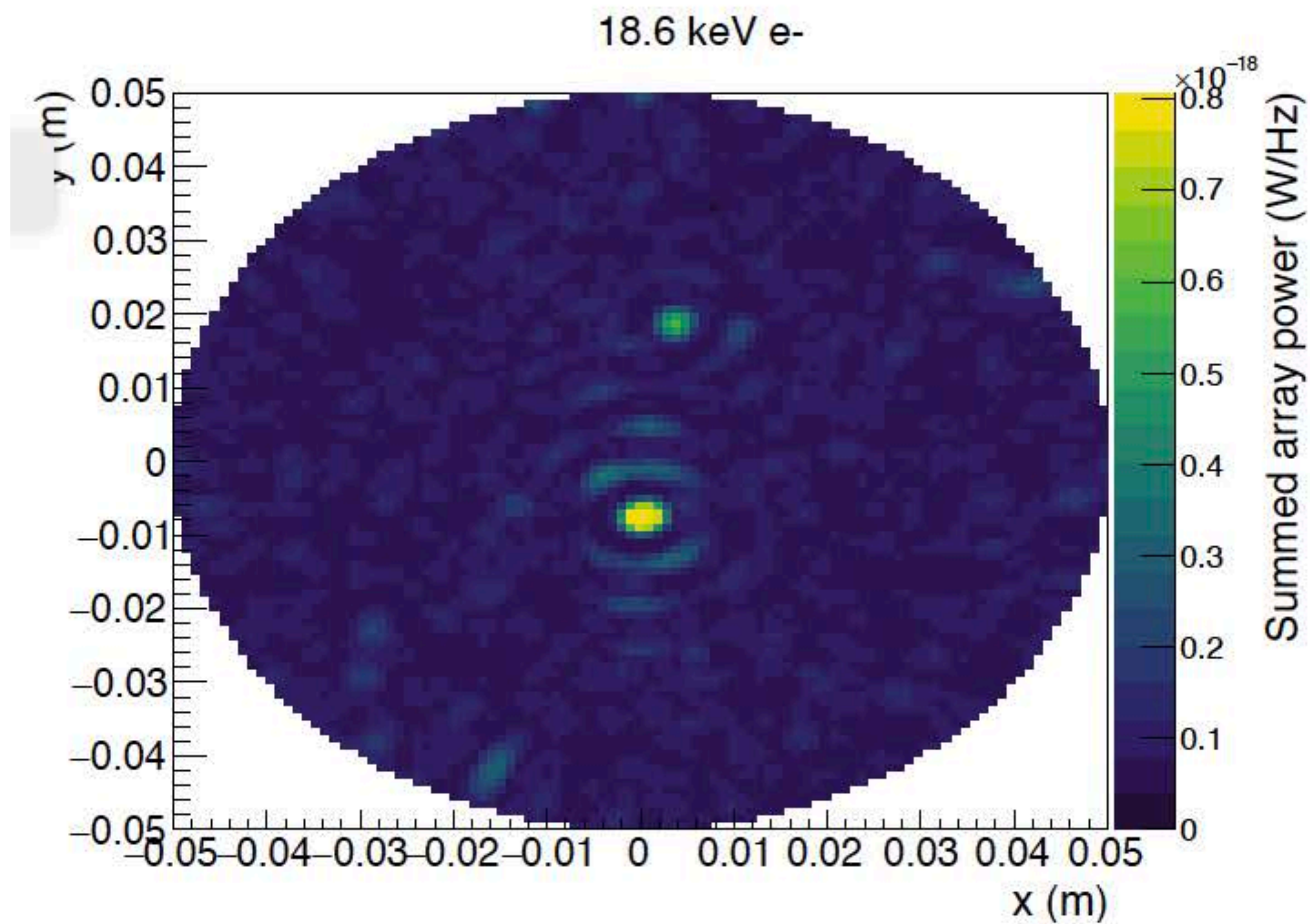
Project 8: Phase III

- > Inwards-looking antenna array watches electrons radiating cyclotron power in free space



Project 8: Phase III

> Single electrons resolvable in simulation

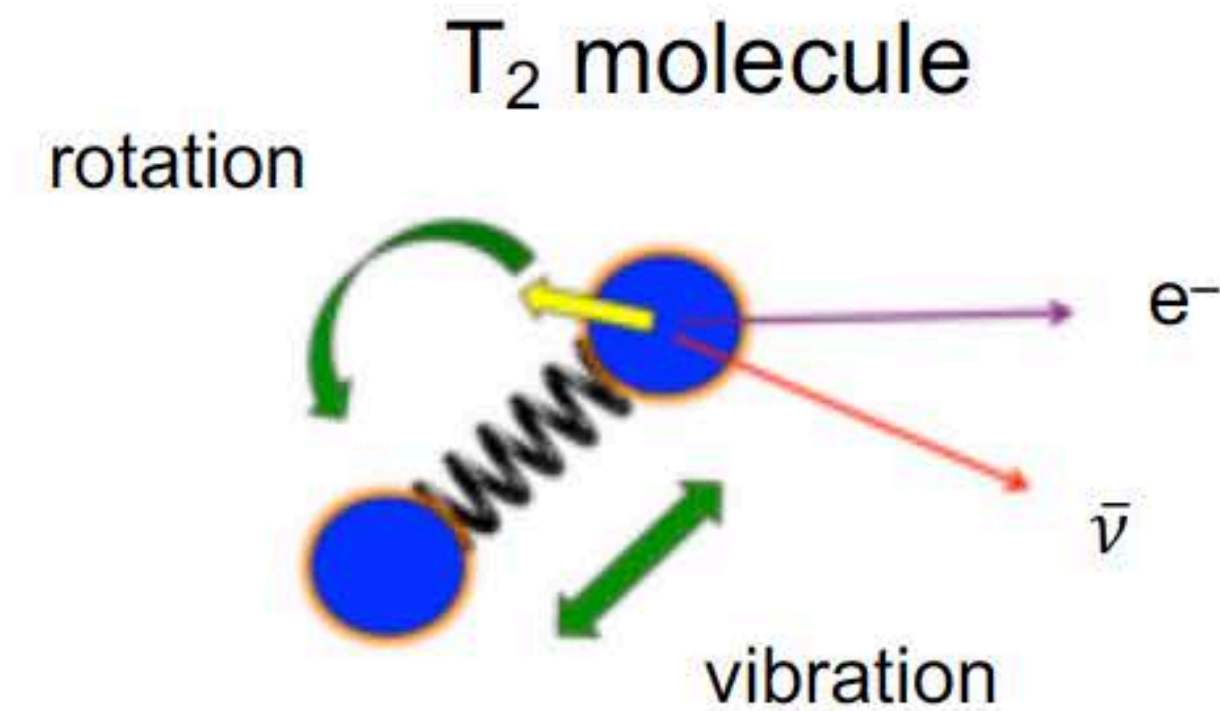
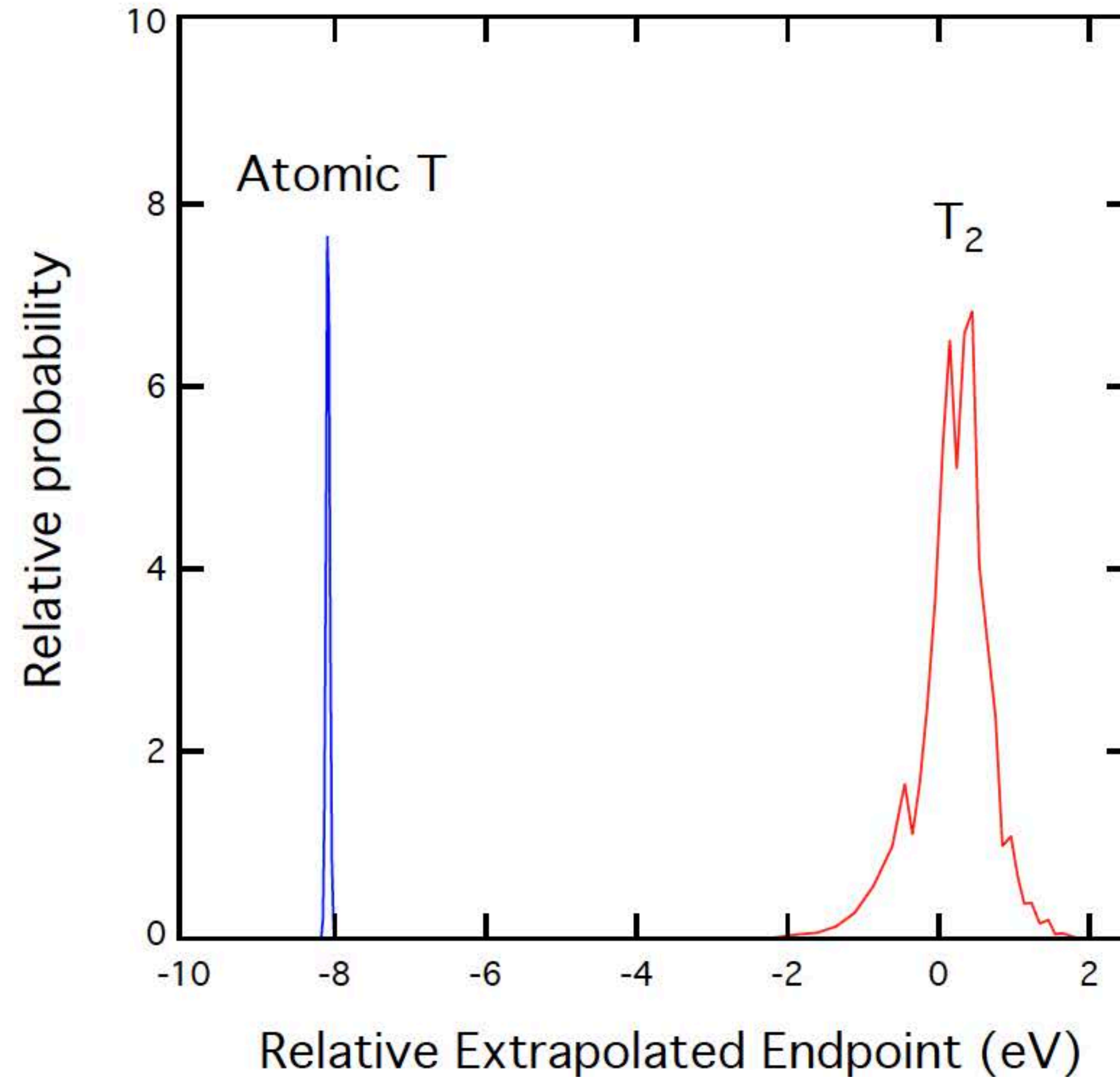


see also:

A. Ashtari Esfahani *et al.*
arXiv:1907.11124

Atomic Tritium

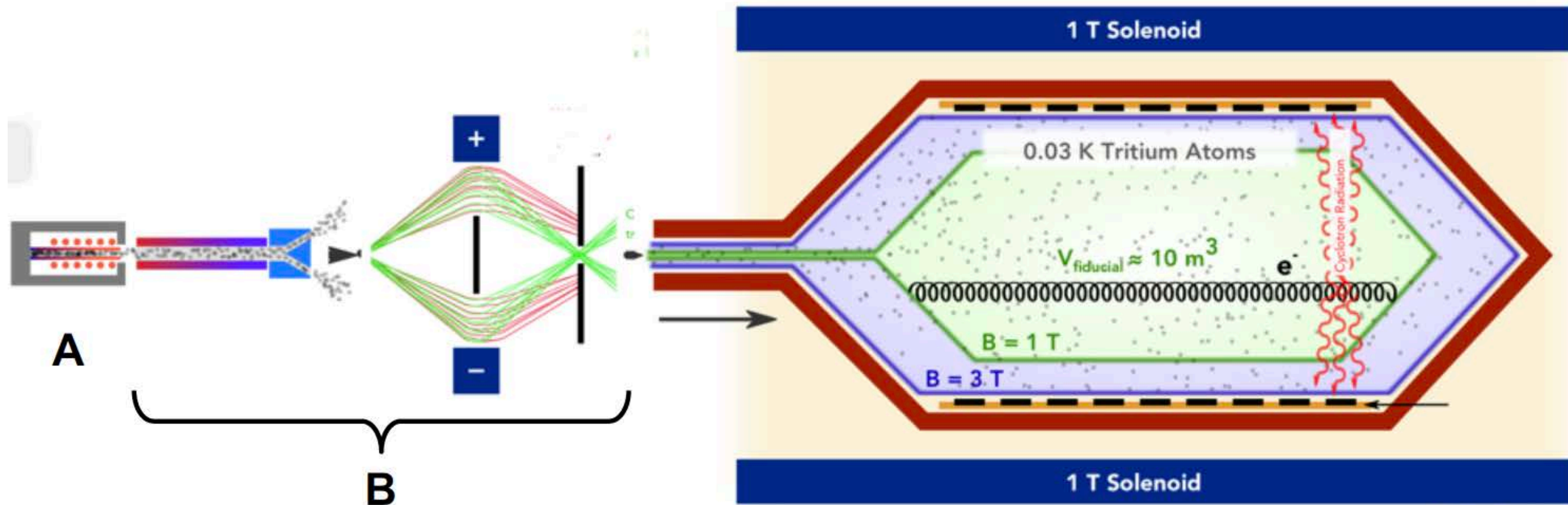
- > Sensitivity beyond inverted hierarchy requires atomic tritium



H. Robertson

Phase IV Concept

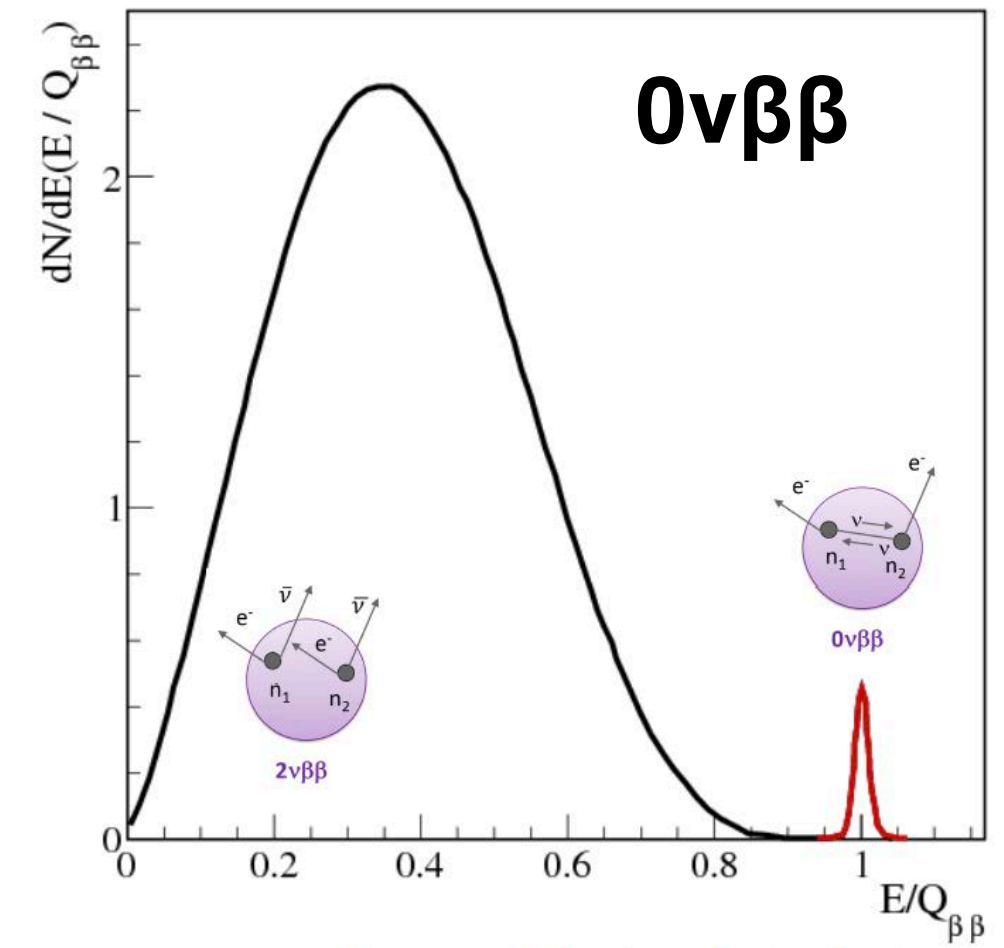
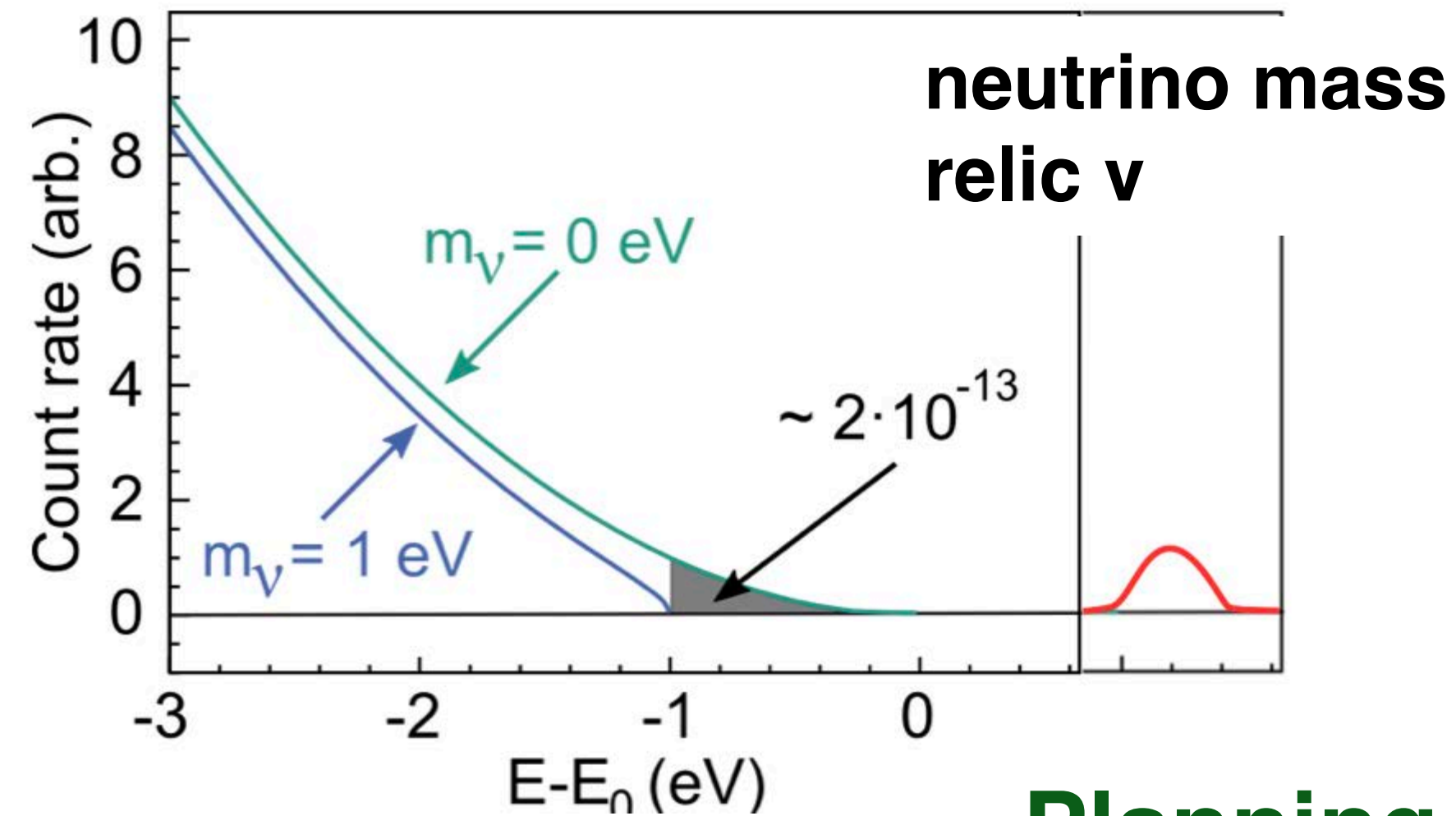
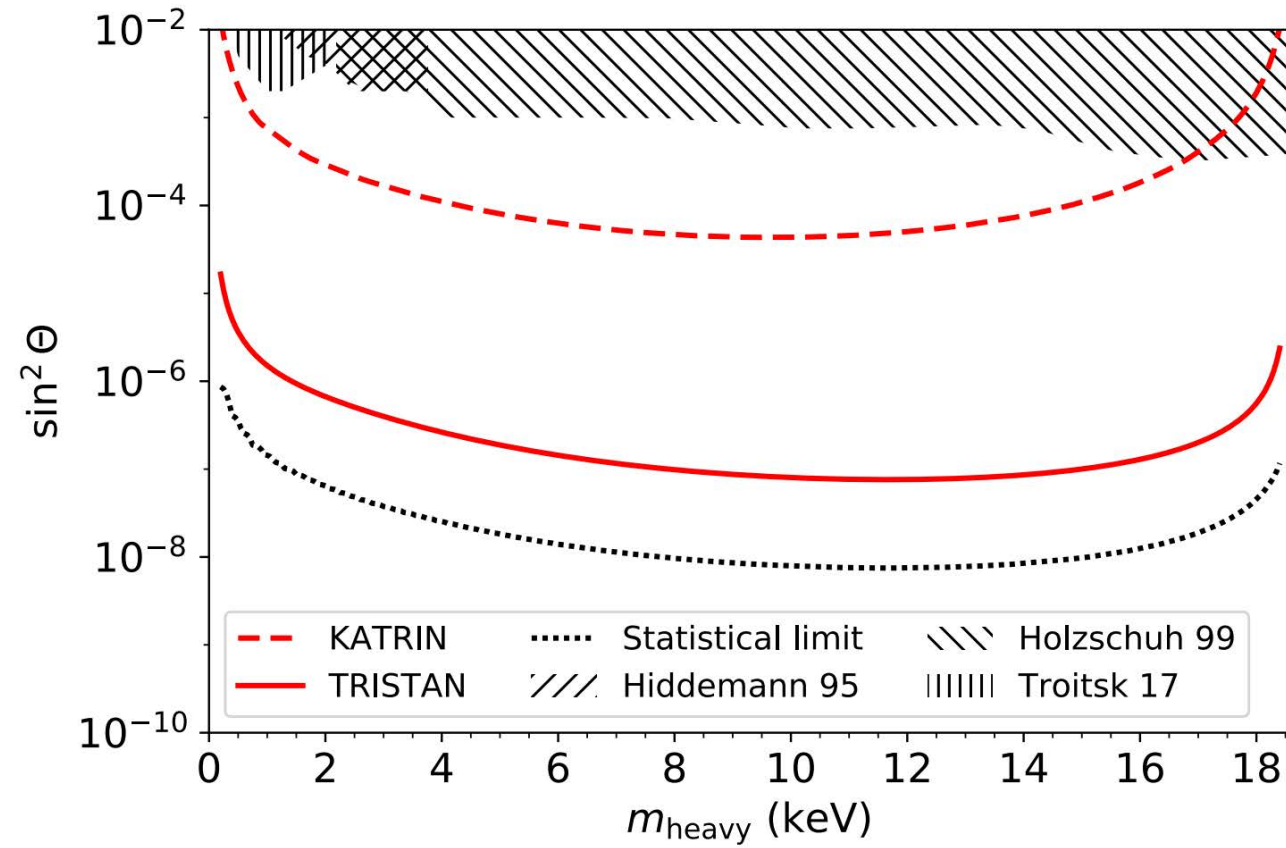
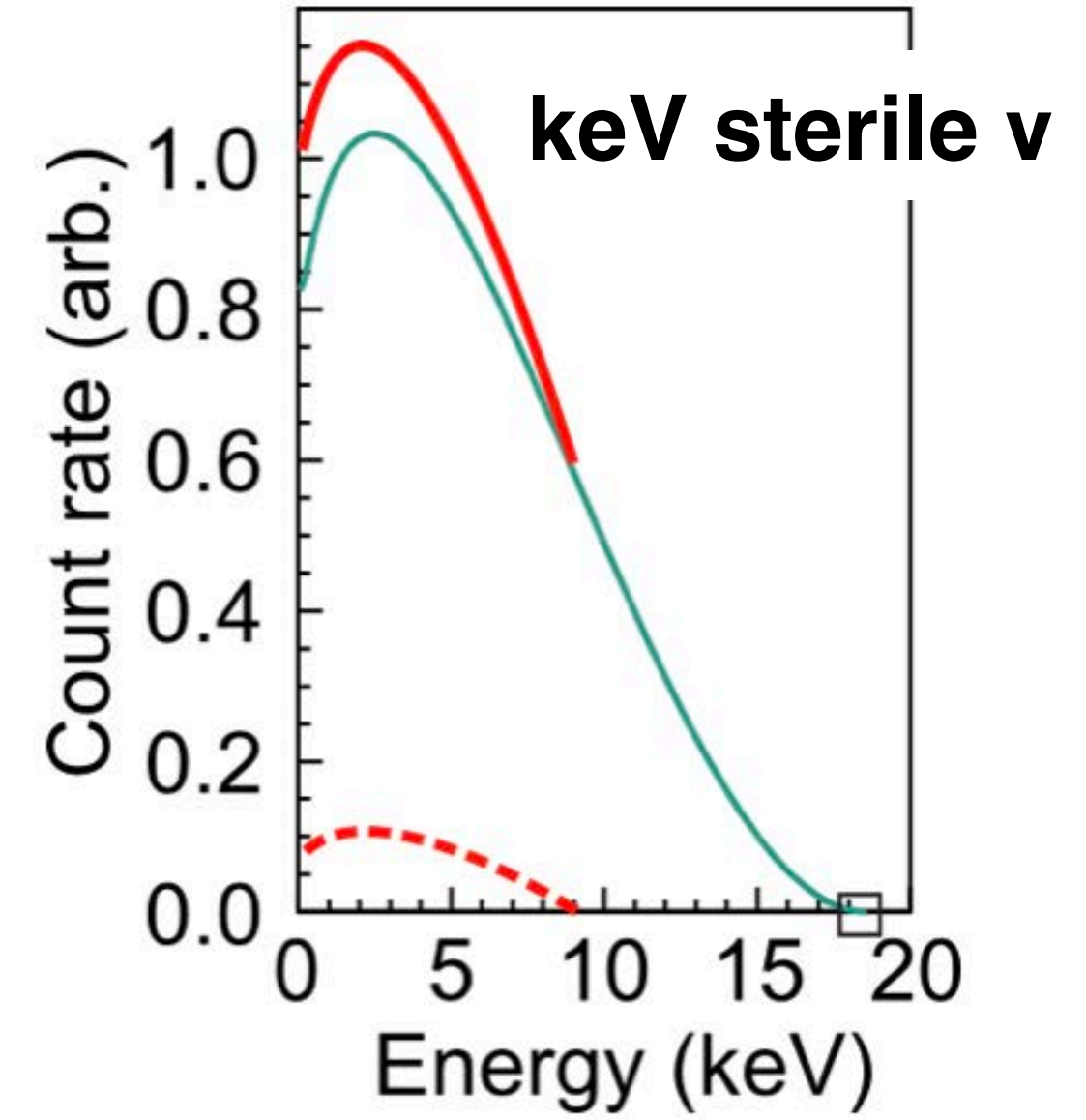
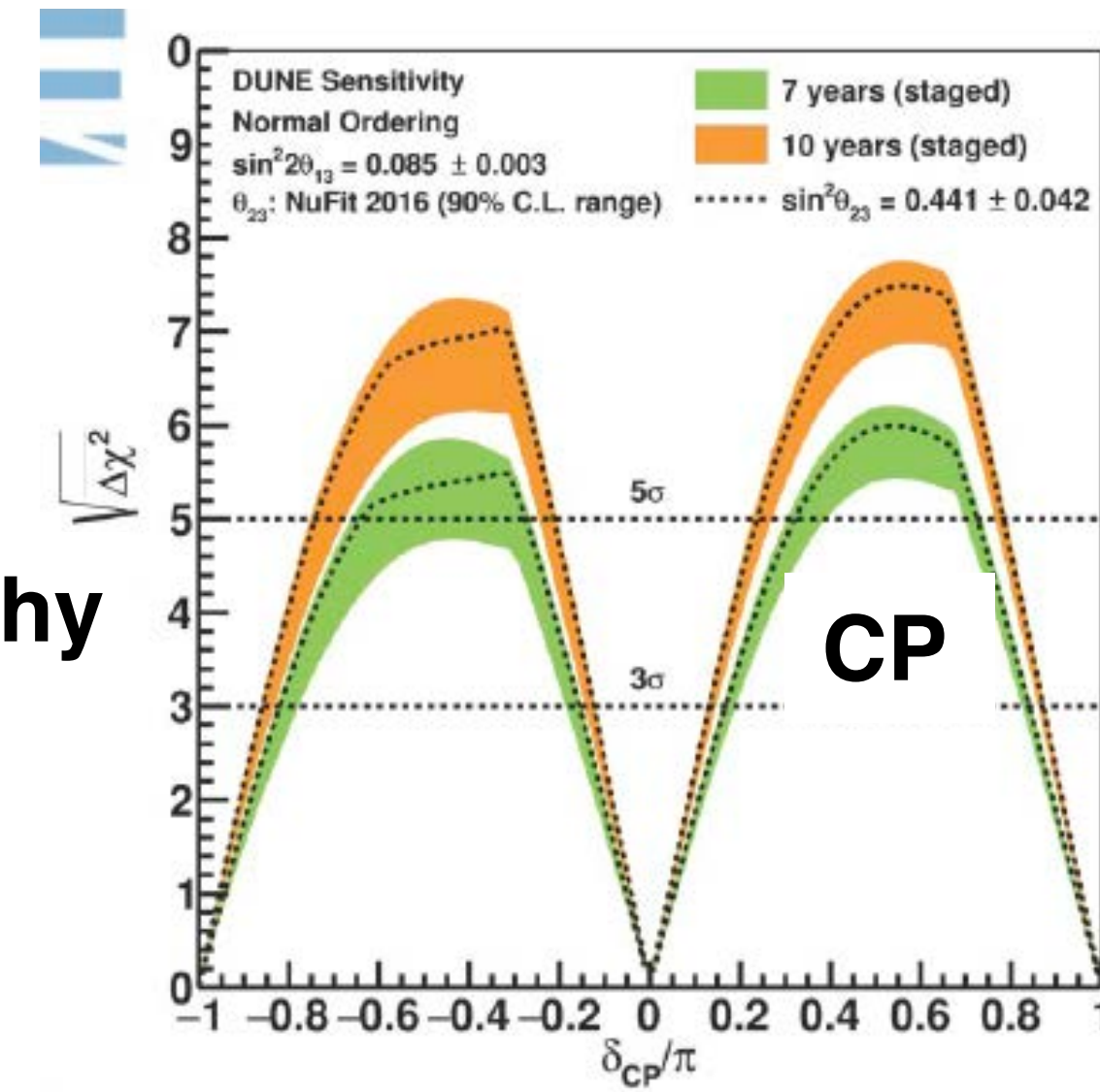
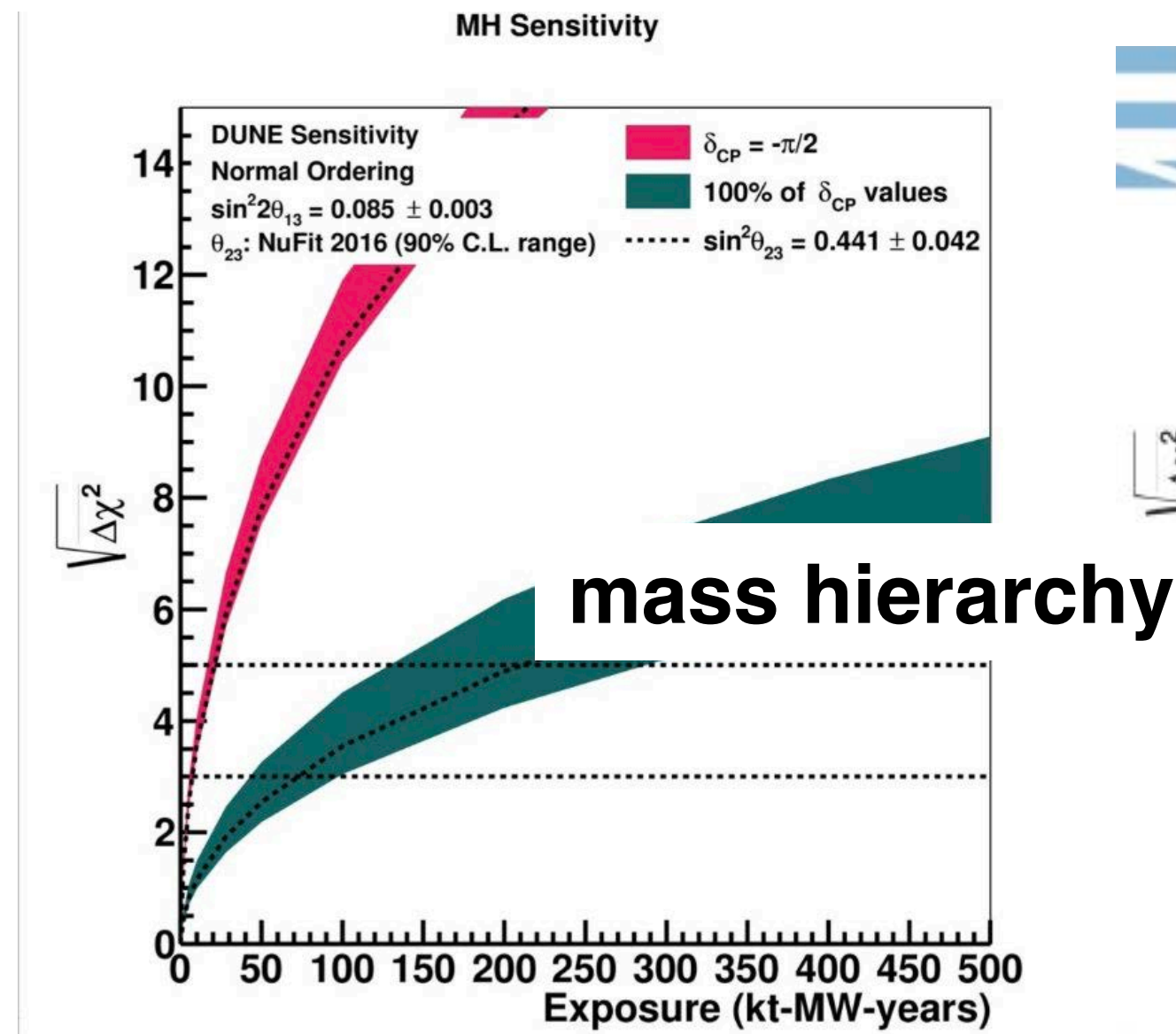
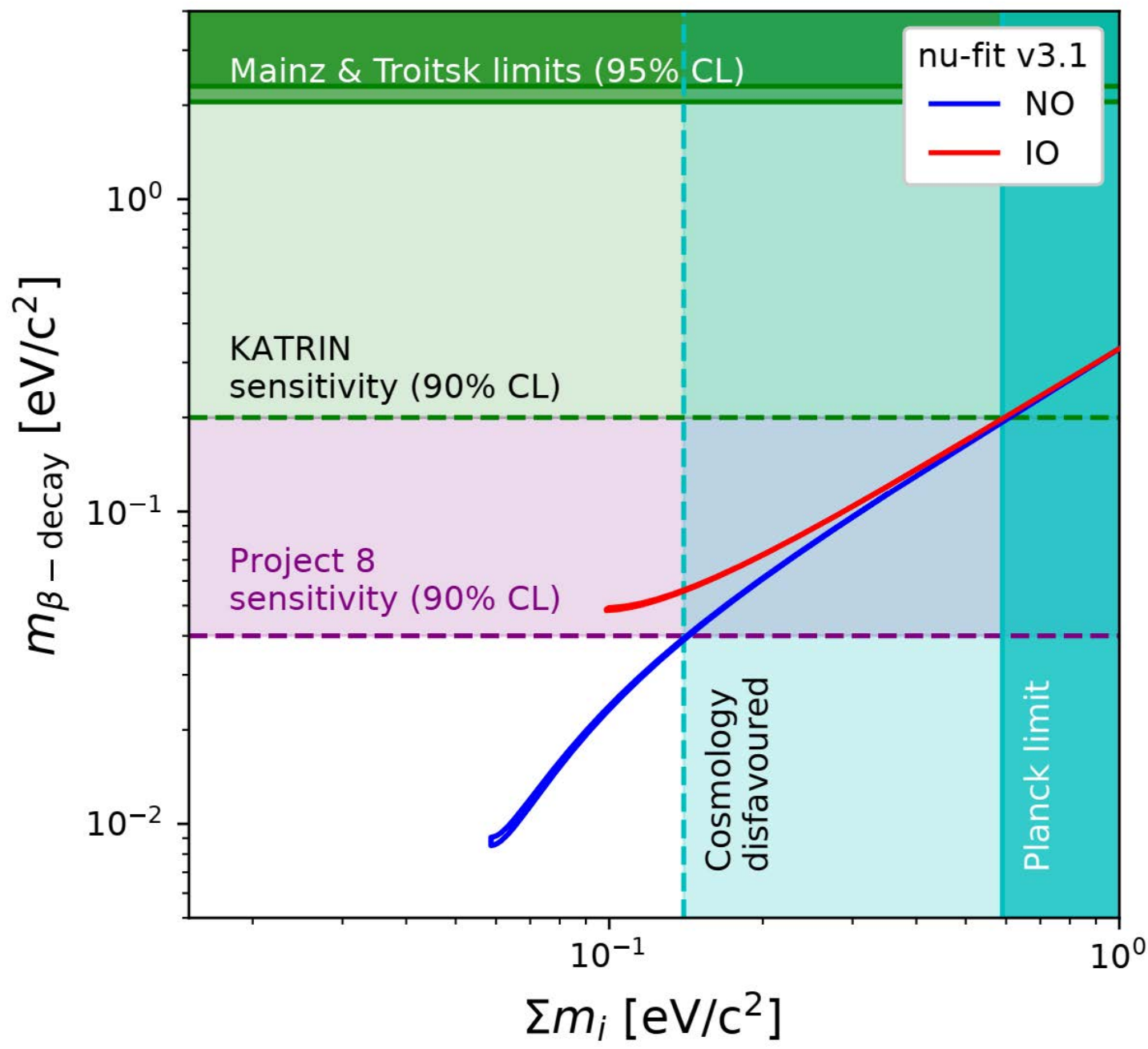
- > Use magnetic moment of atomic species to guide and trap
 - Unpaired electron of atomic T (or H, D) gives it magnetic moment



A. Lindman

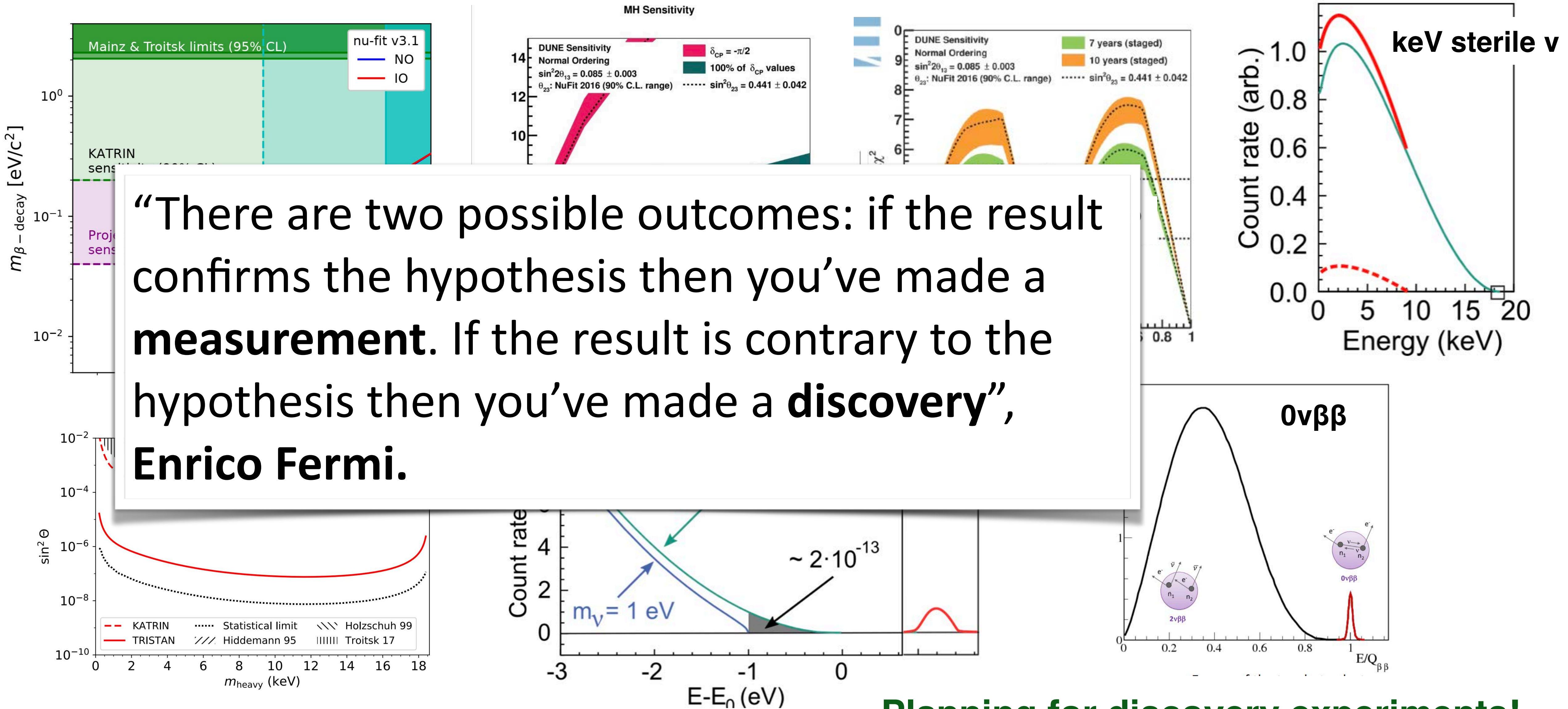
- A: Atomic tritium production**
- B: Transport and preparation**
- C: Trapping and measurement**

Towards the Future - Next Discoveries



Planning for discovery experiments!

Towards the Future - Next Discoveries



Planning for discovery experiments!

Summary

We have demonstrated Cyclotron Radiation Emission Spectroscopy as a novel technique with a promising future in a next-generation neutrino mass experiment

Phase I achieved few-eV resolution of $^{83\text{m}}\text{Kr}$ spectrum

- Approaching natural linewidth of $^{83\text{m}}\text{Kr}$ source

– **Phase II tritium** data taking ongoing

- Challenges of continuous spectrum measurement being met

R&D underway towards scaling CRES technique, atomic tritium for next generation endpoint measurement

Project 8 Collaboration



Case Western Reserve University

- Laura Gladstone, Benjamin Monreal, Yu-Hao Sun



Harvard-Smithsonian Center for Astrophysics

- Sheperd Doeleman, Jonathan Weintraub, André Young



Johannes Gutenberg-Universität Mainz

- Sebastian Böser, Christine Claessens, Martin Fertl, Michael Gödel, Alec Lindman



Karlsruher Institut für Technologie

- Thomas Thümmel



Lawrence Livermore National Laboratory

- Kareem Kazkaz, Lucie Tvrznikova



Massachusetts Institute of Technology

- Zachary Bogorad, Nicholas Buzinsky, Joseph Formaggio, Evan Zayas



Pacific Northwest National Laboratory

- Mathieu Guigue, Mark Jones, Benjamin LaRoque, Erin Morrison, Noah Oblath, Malachi Schram, Jonathan Tedeschi, Brent VanDevender, Mathew Thomas, Mauro Grandi



Pennsylvania State University

- Luiz de Viveiros, Timothy Wendler, Andrew Ziegler



University of Washington

- Ali Ashtari Esfahani, Raphael Cervantes, Peter Doe, Eric Machado, Elise Novitski, Walter Pettus, Hamish Robertson, Leslie Rosenberg, Gray Rybka



Yale University

- Karsten Heeger, James Nikkel, Luis Saldaña, Penny Slocum, Pranava Surukuchi, Arina Telles



This work was supported by the US DOE Office of Nuclear Physics, the US NSF, the PRISMA+ Cluster of Excellence at the University of Mainz, and internal investments at all collaborating institutions



P8 slides on behalf of collaboration

