# Constraining NSI and Sterile Neutrino Physics with $\nu_{ au}$ Appearance in DUNE

#### Anish Ghoshal

Laboratorie Nazionale Frascati - INFN University Roma Tre

anishghoshal1@gmail.com

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Based on -

- (i) A. Ghoshal, A.Giarnetti and D.Meloni, arXiv:1906.06212
- (ii) A. Ghoshal, A.Giarnetti and D.Meloni, [draft in preparation 1911.xxxx]

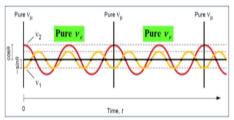
## Outline of talk:

- Neutrino Oscillation.
- Deep Underground Neutrino Experiment (DUNE).
- $\nu_{\tau}$  Appearance in DUNE.
- Standard Physics & Effect of Systematics.
- Sterile Neutrino in 3+1 scheme.
- Non-Standard Interaction (NSI).
- Neutrino Decay.
- Extracting Further New Physics with  $\nu_{\tau}$ .

#### Introduction:

# Neutrino Oscillation is now well-known phenomena.

# Flavor changes happen during the propagation of neutrinos!



#### Introduction:

#### Probability of Neutrino Oscillation:

$$i\frac{d}{dx}\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = H^\nu \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

$$H^{\nu} = H_{\text{vac}} + H_{\text{mat}}$$
 and  $H^{\bar{\nu}} = (H_{\text{vac}} - H_{\text{mat}})^*$ 

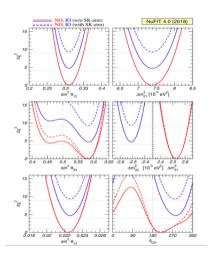
$$H_{vac} = U \cdot \mathrm{Diag}(m_1^2/2E_{\nu}, m_2^2/2E_{\nu}, m_3^2/2E_{\nu}) \cdot U^{\dagger}$$

$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{bmatrix} \begin{bmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{bmatrix} \begin{bmatrix} \cos\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = \delta_{\alpha\beta} - 4\Re \left[ \sum_{i>j}^{3} U_{\alpha i}^{*} U_{\beta i} U_{\alpha j} U_{\beta j}^{*} \sin^{2} \left( \frac{\Delta m_{ij}^{2} L}{4E} \right) \right] + 2\Im \left[ \sum_{i>j}^{3} U_{\alpha i}^{*} U_{\beta i} U_{\alpha j} U_{\beta j}^{*} \sin \left( \frac{\Delta m_{ij}^{2} L}{2E} \right) \right]$$

## Introduction:

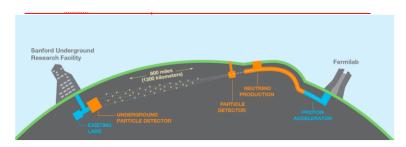
#### NuFit values of the oscillation parameters:



# Deep Underground Neutrino Experiment:

#### DUNE in a nutshell:

- Intense beam of  $\nu_{\mu}$  fired from FermiLab at a large detector 1300 KM away.
- Large (40 kt) Underground Liquid Argon detector at Sanford Underground Research Facility (SURF).



# Deep Underground Neutrino Experiment:

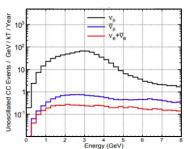
#### Standard Phenomenology in DUNE:

#### ν<sub>e</sub> appearance

$$P(\nu_{\mu} \to \nu_{e}) \approx \sin^{2} 2\theta_{13} \sin^{2} \theta_{23} \sin^{2} \left(\frac{\Delta m_{31}^{2} L}{4E}\right)$$

#### $v_{\mu}$ disappearance

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) \approx 1 - (\sin^2 2\theta_{23} \cos^4 \theta_{13} + \sin^2 2\theta_{13} \sin^2 \theta_{23}) \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E}\right)$$



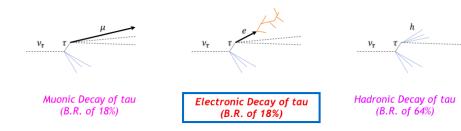
# Calculation of Events - Standard

$ u$ mode (150 kt $\cdot$ MW $\cdot$ year)			
$ u_e$ Signal NH (IH)	861 (495)		
$ar{ u}_e$ Signal NH (IH)	13 (26)		
Total Signal NH (IH)	874 (521)	_	
Beam $ u_e + ar{ u}_e$ CC Bkgd	159		
NC Bkgd	22		
$ u_{ au} + ar{ u}_{ au}$ CC Bkgd	42		
$ u_{\mu} + ar{ u}_{\mu} \; CC \; Bkgd$	3		
Total Bkgd	226	$ u$ mode (150 kt $\cdot$ MW $\cdot$ year)	
$\bar{\nu}$ mode (150 kt · MW · year)		$\overline{ u_{\mu}}$ Signal	10842
$ u_e$ Signal NH (IH)	61 (37)	$ar{ u}_{\mu}$ CC Bkgd	958
$ar{ u}_e$ Signal NH (IH)	167 (378)	NC Bkgd	88
Total Signal NH (IH)	228 (415)	$ u_{ au} + ar{ u}_{ au}$ CC Bkgd	63
Beam $ u_e + \bar{ u}_e$ CC Bkgd	89	$\bar{\nu}$ mode (150 kt $\cdot$ MW $\cdot$ year)	
NC Bkgd	12	$ar{ u}_{\mu}$ Signal	3754
$ u_{ au} + ar{ u}_{ au}$ CC Bkgd	23	$\nu_{\mu}$ CC Bkgd	2598
$ u_{\mu} + ar{ u}_{\mu} \; CC \; Bkgd$	2	NC Bkgd	50
Total Bkgd	126	$ u_{ au} + ar{ u}_{ au}$ CC Bkgd	39

#### Tau Detection

#### Detection

•  $\nu_{\mu} \rightarrow \nu_{\tau}$  channel has never been considered in the simulations of DUNE hitherto. In fact tau neutrinos are difficult to observe. Furthermore, the interactions of these neutrinos have a rather high energy threshold (3.4), which is why the number of events expected for this process is low.



Disclaimer - - we do not say this is the most suitable channel for detection.

# Signal and Backgrounds

We consider au o e as the detection channel. ICARUS proposal followed this strategy.

# $\begin{array}{c} \text{Signal} \\ v_{\mu} \text{->} \ v_{\tau} \ \text{oscillation} \end{array}$

# Backgrounds

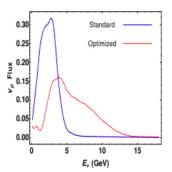
 $v_{\mu}$  ->  $v_{e}$  oscillation  $v_{e}$  ->  $v_{e}$  from beam

We consider various configurations to understand their impact on final sensitivities.

- 20% & 10% Signal Uncertainties in the  $\nu_{\tau}$  channel.
- $\bullet~100\%~\&~33\%$  of electrons being detected (detection efficiency).
- S/B values of 2.46 and 18.6.
- Standard and Tau-Optimized Fluxes.

#### Neutrino Flux

The standard flux consists of beam delivering  $1.47 \times 10^{21}$  protons on target (POT) per year with 80 GeV energy running with 1.07 MW beam power and having 1.5 m NuMi (Neutrino Main Injector) style target. The  $\tau-$  optimized flux is as per proposed by the DUNE collaboration consists of  $1.1 \times 10^{21}$  protons on target (POT) per year with 120 GeV energy running with 1.2 MW beam power and having 1m NuMi style target.



#### Rate Estimation

#### A comparison of $\nu_{ au}$ events:

u mode		$ar{ u}$ mode	
$ u_{ au} $ Signal	277	277 $\nu_{\tau}$ Signal	
$\bar{ u}_{ au}$ Signal	26	$\bar{ u}_{ au}$ Signal	85
Total Signal	303	Total Signal 15	
$\nu_e + \bar{\nu}_e$ CC Background (beam)	333 + 38	$3 + 38$ $\nu_e + \bar{\nu}_e$ CC Background (beam) 117	
CC Background (oscillation)	1753 + 12	$2 \mid \mid \nu_e$ CC Background (oscillation)   90 + 1	

Figure: Standard Flux

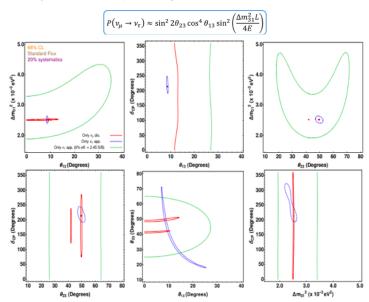
$\nu$ mode		$\bar{\nu}$ mode	
$ u_{ au} $ Signal	2673	$ u_{ au}$ Signal	98
$ar{ u}_{ au}$ Signal	34	$ar{ u}_{ au}$ Signal	983
Total Signal	2707	Total Signal	1081
$\nu_e + \bar{\nu}_e$ CC Background (beam)	688 + 63	$\nu_e + \bar{\nu}_e$ CC Background (beam)	176 + 177
CC Background (oscillation)	1958 + 11	$\nu_e$ CC Background (oscillation)	76 + 324

Figure: Optimized Flux

Experiment run-time of (3.5 + 3.5) years. Latest NuFit values of the Oscillation parameters used.

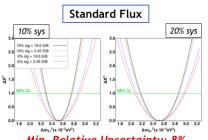
## Correlation Studies

Standard Physics does not improve using  $\nu_{ au}$  channel.

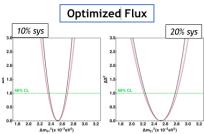


# Comparison with OPERA

OPERA has observed 10 events in the  $\nu_{\mu} \to \nu_{\tau}$  channel. Using these events the  $\Delta m_{31}^2$ parameter uncertainty is about 26%. Using the  $\tau$  events in DUNE this can be largely improved.



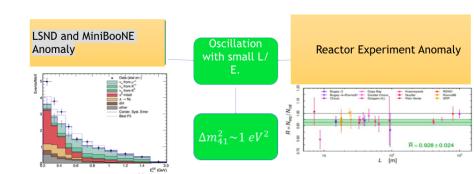
Min. Relative Uncertainty: 8%



Min. Relative Uncertainty: 4.5%

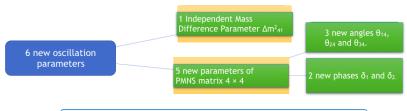
# Short Baseline Anomaly

#### Long Discussions Yesterday:



### Sterile Neutrinos in 3+1 Scheme

The simplest model that includes sterile neutrinos is the 3 + 1 model, in which only one sterile neutrino is added.



 $U_{PMNS} = R(\theta_{34}) R(\theta_{24}) R(\theta_{23}; \ \delta_2) R(\theta_{14}) R(\theta_{13}; \ \delta_3) R(\theta_{12}; \ \delta_1)$ 

#### **Parameters**

(4.2)

$$\begin{split} P(\nu_{\mu} \rightarrow \nu_{e}) = & 2|U_{14}|^{2}|U_{24}|^{2} + 4\Re[U_{23}^{*}U_{13}(U_{23}U_{13}^{*} + U_{24}U_{14}^{*})]\sin^{2}\left(\frac{\Delta m_{32}^{2}L}{4E}\right) + \\ & - 2\Im(U_{23}^{*}U_{13}U_{24}U_{14}^{*})\sin\left(\frac{\Delta m_{32}^{2}L}{2E}\right) \end{split}$$

 $U_{34}=\cos\theta_{14}\cos\theta_{24}\sin\theta_{34}.$ 



$$\begin{split} P(\nu_{\mu} \rightarrow \nu_{\tau}) = & \underbrace{(U_{34})^{3} |U_{24}|^{2} + 4\Re[U_{23}^{*}U_{33}(U_{23}U_{33}^{*} + U_{34}U_{34}^{*}) \sin^{2}\left(\frac{\Delta m_{32}^{2}L}{4E}\right)}_{+ 2\Im(U_{23}^{*}U_{34}U_{34}U_{34}^{*}) \sin\left(\frac{\Delta m_{32}^{2}L}{2E}\right) \end{split} \tag{4.3}$$

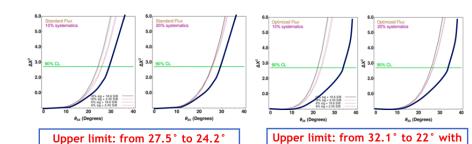
We expect the addition of  $v_{\mu}$  ->  $v_{\tau}$  appearance channel to improve  $\theta_{34}$  sensitivity.

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) = 1 - 2|U_{24}|^2(1 - |U_{24}|^2) - 4[|U_{23}|^2(1 - |U_{23}|^2 - |U_{24}|^2)]\sin^2\left(\frac{\Delta m_{32}^2 L}{4E}\right) + \frac{4}{4}(4.4)$$

# Effect of Systematics

Effect of Systematics, detection efficiencies, S/B values and two fluxes on the measurement of  $\theta_{34}$ .

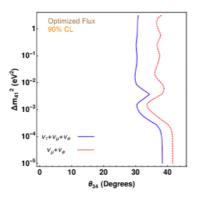
with the standard flux.



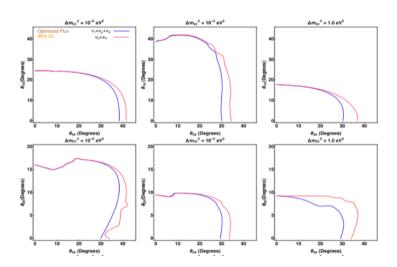
the optimised flux.

# **Correlation Studies**

We can see the maximum effect is on the improvement of  $\theta_{34}$  only.



# Correlation Studies



# New Physics: NSI

Diligent way to capture the effect of new physics, in terms of four-fermion interaction.

$$-\mathcal{L}_{ ext{NSI}}^{eff} = arepsilon_{lphaeta}^{fP} 2\sqrt{2}G_F(ar{
u}_lpha\gamma_
ho L
u_eta)(ar{f}\gamma^
ho Pf)$$

$$i\frac{d}{dt}\left(\begin{array}{c} \nu_e \\ \nu_\mu \\ \nu_\tau \end{array}\right) = \left[\begin{matrix} U_{PMNS} \left(\begin{array}{ccc} 0 & 0 & 0 \\ 0 & \Delta_{21} & 0 \\ 0 & 0 & \Delta_{31} \end{array}\right) U_{PMNS}^\dagger + A \left(\begin{array}{ccc} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\tau\tau}^* & \epsilon_{\tau\tau} \end{array}\right) \right] \left(\begin{array}{c} \nu_e \\ \nu_\mu \\ \nu_\tau \end{array}\right)$$

$$\begin{split} P(\nu_{\alpha} \to \nu_{\beta}; \varepsilon_{e\mu}, \varepsilon_{e\tau}, \varepsilon_{\mu\mu}, \varepsilon_{\mu\tau}, \varepsilon_{\tau\tau}) &= P(\nu_{\alpha} \to \nu_{\beta}; 2 \text{ flavor in vacuum}) \\ &+ P(\nu_{\alpha} \to \nu_{\beta}; \varepsilon_{e\mu}, \varepsilon_{e\tau}) \\ &+ P(\nu_{\alpha} \to \nu_{\beta}; \varepsilon_{\mu\mu}, \varepsilon_{\mu\tau}, \varepsilon_{\tau\tau}), \end{split}$$

# New Physics: NSI

#### Tau Appearance Probability with NSI:

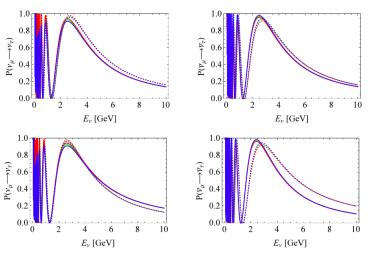


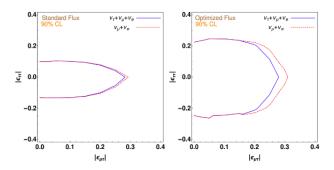
Figure: Solid/Dotted - - NH/IH. Green/Red/Blue - -  $\delta_{CP}=[0,\pi/2,-\pi/2]$ . Top/Bottom - - (No NSI;  $[\epsilon_{\mu\tau},\epsilon_{\tau\tau}]=(0.07,0.147)$ )



#### **NSI:**Correlation Studies

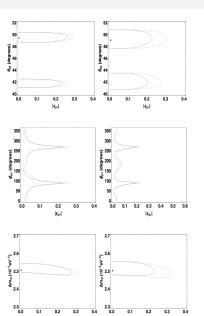
#### **NSI** probability

$$P_{\mu\tau} = P_{\mu\tau}^{SM} + \left(\frac{1}{2}\epsilon_{\tau\tau}\cos^2(2\theta_{23}) + 2\cos(2\theta_{23})\operatorname{Re}\{\epsilon_{\mu\tau}\}\right)(AL)\sin\left(\frac{\Delta m_{31}^2 L}{2E}\right) + \mathcal{O}(\epsilon^2)$$



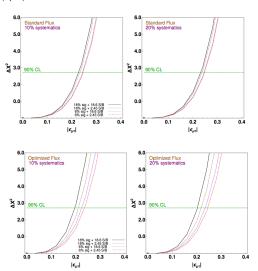
In the optimized flux, we do not get the advantage of increased tau-statistics as  $\nu_e$  &  $\nu_\mu$  channels are also increased.

# **NSI** Correlation Studies



# NSI:Sensitivity on NSI parameter

Impact of Systematics, detection efficiencies, S/B values and two fluxes on the measurement of  $|\epsilon_{u\tau}|$ .



# Neutrino Decay

#### Neutrino Decay - - Introduction

$$\mathcal{L}_{\text{int}} = \frac{(g_s)_{ij}}{2} \bar{\nu}_i \nu_j S + i \frac{(g_p)_{ij}}{2} \bar{\nu}_i \gamma_5 \nu_j S$$

$$\nu_i \rightarrow \nu + S$$

$$\Gamma_i(L, E) = \exp(-\alpha_i \times L/E)$$

$$\alpha_i = rac{m_i}{ au_i}$$

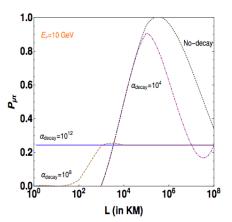
$$\begin{pmatrix} \nu_{\alpha} \\ \nu_{s} \end{pmatrix} = \begin{pmatrix} U_{PMNS} & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_{i} \\ \nu_{4} \end{pmatrix}$$

$$H=Uegin{bmatrix} 1 \ 2E egin{pmatrix} 0 & 0 & 0 \ 0 & \Delta m_{21}^2 & 0 \ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} -irac{lpha_3}{2E}egin{pmatrix} 0 & 0 & 0 \ 0 & 0 & 0 \ 0 & 0 & 1 \end{pmatrix} \end{bmatrix}U^\dagger + egin{pmatrix} A & 0 & 0 \ 0 & 0 & 0 \ 0 & 0 & 0 \end{pmatrix}$$

# Why Interesting for DUNE

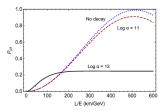
$$\begin{split} P_{\mu\mu}(E,L,\alpha_3) &= \left(\cos^2\theta_{23} + \sin^2\theta_{23} e^{-\frac{\alpha_3}{2E}L}\right)^2 \\ &- 4\cos^2\theta_{23}\sin^2\theta_{23} e^{-\frac{\alpha_3}{2E}L} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E}\right) \\ &+ 4\cos^2\theta_{23}\sin^2\theta_{23} e^{-\frac{\alpha_3}{2E}L} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E}\right) \\ \end{split}$$

Decay term causes vanishing of interference effects. Increase in Probability.



# Why Interesting for DUNE

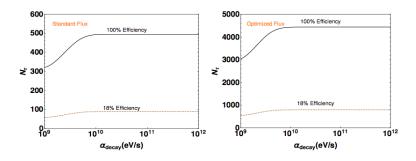
DUNE is very suited to explore this inrceased probability region due to its L/E.



# Some Preliminary Results

Expected number of  $\nu_{\tau}$  events:

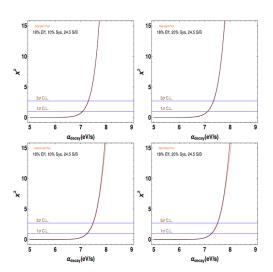
$$N_{ au} = \epsilon_{det} imes n_{p.o.t.} imes N_{Ar} imes \int dE_{\mu}(\phi) \sigma_{
u_{ au}}(E) P_{\mu au}(E,lpha_3)$$



Disclaimer - - preliminary calculations only. Final results may differ !

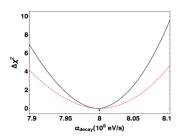
# Some Preliminary Results

#### Sensitivity of measurement of $\alpha$ :



# Some Preliminary Results

## Chi-squared Fit Analysis:



#### Conclusions

- In the case of standard physics, the addition of  $\nu_{\tau}$  appearance channel does not improve the sensitivities of any of the neutrino oscillation parameter set by the other two channels already being considered in DUNE.
- We studied the impact of various systematics,  $\nu_{\tau}$  detection efficiencies, experimental reaches (2 different S/B ratios) and the two fluxes on the sensitivities of the oscillation parameters. The performances of the tau optimized flux in the  $\nu_{e}$  appearance and  $\nu_{\mu}$  disappearance channels result in worsening the sensitivities overshadowing the advantage one may get from the increase in the  $\nu_{\tau}$  statistics. This is mainly due to the increased background events in both the  $\nu_{e}$  and  $\nu_{\mu}$  channels.
- In the new physics cases, NSI parameter sensitivities remains less unaffected after the addition of the new channel, except for the coupling  $|\epsilon_{\mu\tau}|$  for which improved limits (about 15% better) was found.
- For the sterile neutrino (3+1) case, the only parameter that shows an increase in sensitivity is the mixing angle  $\theta_{34}$  and we estimated the improvement to be about 20%.
- Neutrino Invisible Decay constant parameter can be constrained using the  $\nu_{ au}$  appearance channel due to a suitable L/E configuration that DUNE provides.

#### **Future Directions**

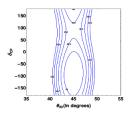
- Study involving shared run-time between Standard and Optimized fluxes so as to maximize the tau channel capabilities.
- Combining electronic and hadronic channels of tau decay so as to maximize the tau detection efficiency and consequently increase in tau-statistics.
- $\nu_{\mu} 
  ightarrow \nu_{ au}$  maybe suited to study Large Extra Dimension scenarios.
- $\nu_{\mu} \rightarrow \nu_{\tau}$  maybe suited to study dark matter scenarios especially searches in dark sectorinvolving  $L_{\mu} L_{\tau}$  symmetries and its corresponding mediators.
- Probe of Non-Unitary and Lorentz Violating Operators using  $\nu_{\tau}$  –appearance channel.
- · Other suggestions are welcome.

Stay tuned. Work in Progress !!

#### **Essential References**

- arxiv 1811.05487
- arxiv 1512.06148
- arxiv 1606.09550
- arxiv 0407333
- ICARUS: http://cds.cern.ch/record/574836/files/spsc-p-323.pdf
- arxiv 0110393
- arxiv 0402175
- arxiv 0705.0107
- arxiv 1209.2710
- arxiv 010317
- arxiv 1603.08696
- arxiv 1805.01747
- arxiv 1811.00095
- arxiv 1904.07265
- http://home.fnal.gov/ljf26/DUNEFluxes

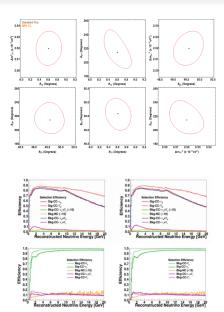
# Thank You



	Standard Flux		Optimized Flux		
	$\nu$ mode	$\bar{\nu}$ mode	$\nu$ mode	$\bar{\nu}$ mode	
$\nu_{\mu}$ CC	30175	3225	85523	4933	
$\bar{\nu}_{\mu}$ CC	1025	9879	1256	26221	
$\nu_e$ CC	371	136	856	258	
$\bar{\nu}_e$ CC	44	109	84	215	

	signal	b	ackgrou	$_{ m nds}$	
		intrinsic $\nu_e$	mis $\nu_{\mu}$	mis $\nu_{\tau}$	NC
	$\nu_{\mu} \rightarrow \nu_{e} \oplus \bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$				
neutrino mode	$1188 \oplus 11.5$	288.2	3.1	19.9	26
	$\nu_{\mu} \rightarrow \nu_{\mu} \oplus \bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu}$				
	$7601 \oplus 519.2$			28.2	75.3
	$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e} \oplus \nu_{\mu} \rightarrow \nu_{e}$				
antineutrino mode	$209 \oplus 64$	171.8	2.9	13.4	15.2
	$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu} \oplus \nu_{\mu} \rightarrow \nu_{\mu}$				
	$2591 \oplus 1489$			16.5	44.1

$\nu_e$ appearance channel			
Signal	$\nu_e$ and $\bar{\nu}_e$ CC events from $\nu_\mu$ oscillations 2%		
	Beam $\nu_e$ and $\bar{\nu}_e$ CC events	5%  sys	
Backgrounds	Misidentified $\nu_{\mu}$ and $\bar{\nu}_{\mu}$ CC events	5% sys	
Dackgrounds	Misidentified $\nu_{\tau}$ and $\bar{\nu}_{\tau}$ CC events	20%  sys	
	Misidentified NC events	10%  sys	
$\nu_{\mu}$ disappearance channel			
Signal	$\nu_{\mu}$ and $\bar{\nu}_{\mu}$ CC events	5% sys	
Backgrounds	Misidentified $\nu_{\tau}$ and $\bar{\nu}_{\tau}$ CC events	20%  sys	
	Misidentified NC events	10% sys	



Limits
(-0.2, 0.45)
< 0.1
< 0.3
(-0.02, 0.175)
< 0.03