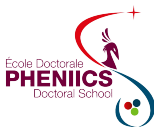


Searching for neutrinoless double beta decay with the SuperNEMO demonstrator

Cloé Girard-Carillo

Laboratoire de l'Accélérateur Linéaire

PHENIICS Fest 2019



What are we composed of?

Water

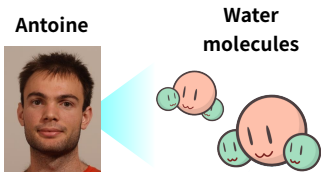


What are we composed of?

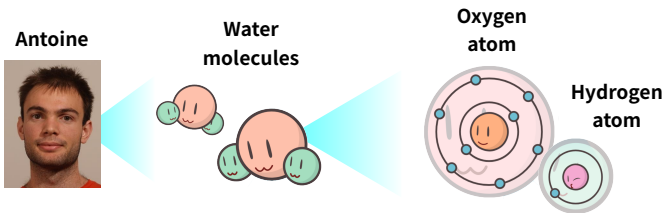
Antoine



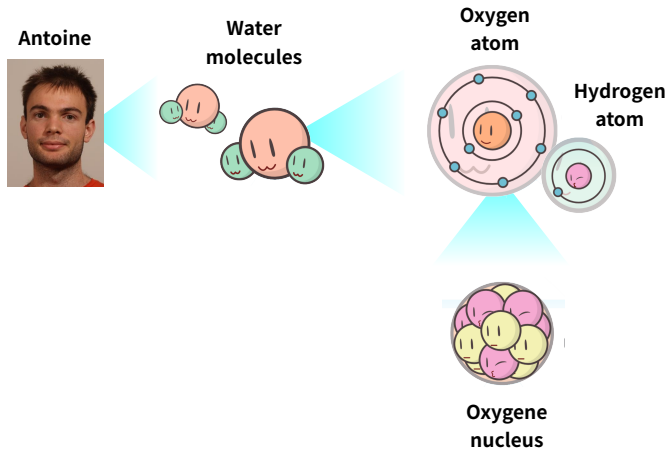
What are we composed of?



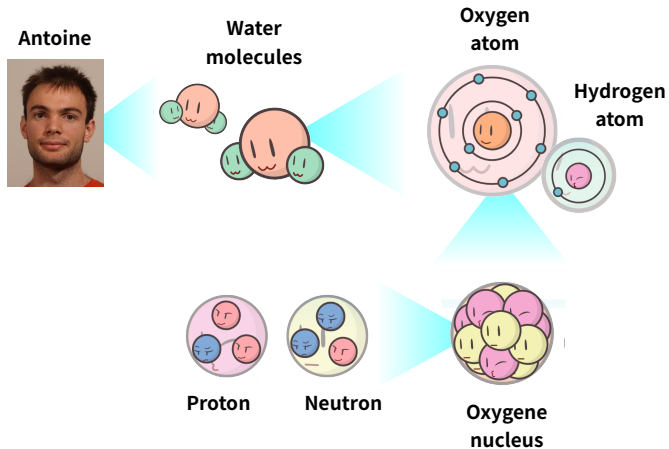
What are we composed of?



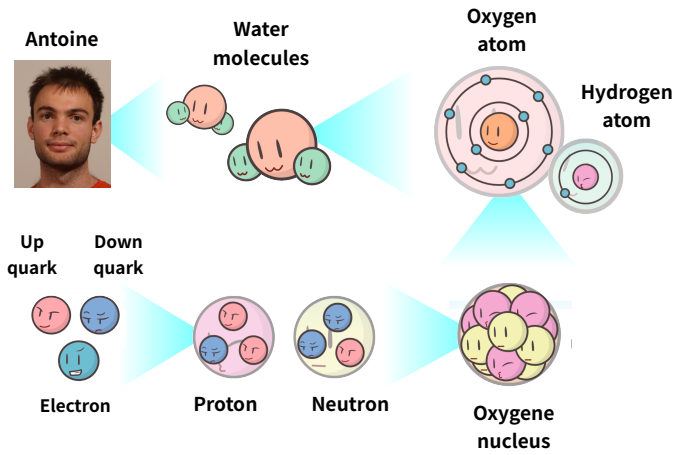
What are we composed of?



What are we composed of?

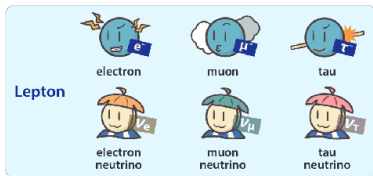
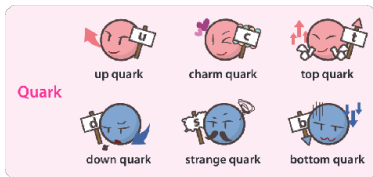


What are we composed of?



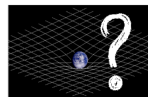
The Standard Model of particles

Fermions (matter)

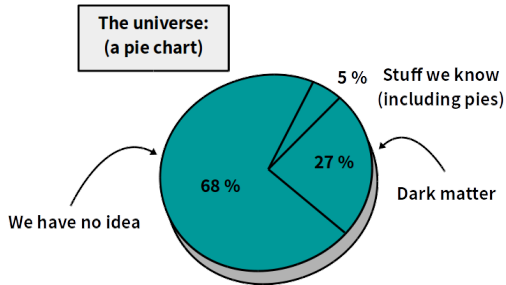
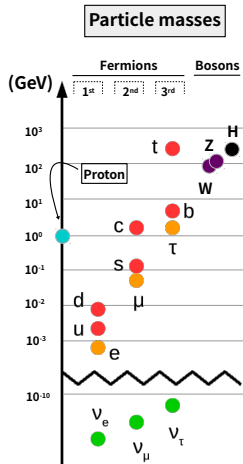


+ Antiparticles

Bosons



Where the Standard Model ends?



We need to go *beyond* the Standard Model



H.Becquerel (1896)



Discovery of radioactivity through β decay
Only electron observed
Non conservation of total energy

W.Pauli (1930)



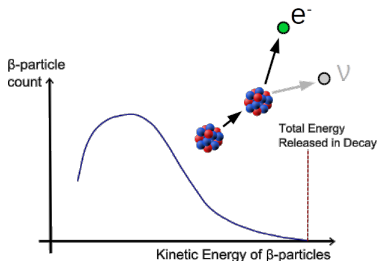
Solution to conserve total energy
"Neutrino"

- Neutral
- Spin 1/2
- Small or null mass
- Small interaction probability

E.Fermi (1934)



Effective theory:
Foundation stone of **weak interaction**



Neutrino oscillations: neutrino masses are not described by the SM

In the Standard Model neutrinos are massless

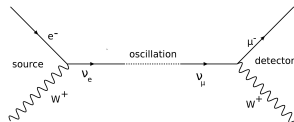
G.Pontecorvo (1957): Can neutrinos oscillate?

$$\nu_\alpha = U_{\text{PMNS}} \sum_{1,2,3} \nu_i$$

Oscillation probability (2 flavours):

$$\mathcal{P}_{e \rightarrow \mu}(t) = \sin^2 2\theta \sin^2 \frac{\Delta m^2}{2E} L$$

Possible only if neutrinos are massive particles



SuperKamiokande experiment (1998): **Observation of neutrino oscillations**

→ At least 2 massive neutrinos (considering 3 flavours)

Need to go *beyond* the Standard Model



Proof of massive neutrinos lead to extensions of SM

The mass generation mechanism depends on the neutrino nature

Dirac particles: neutrino & antineutrino are distinct particles

As other fermions: Higgs mechanism

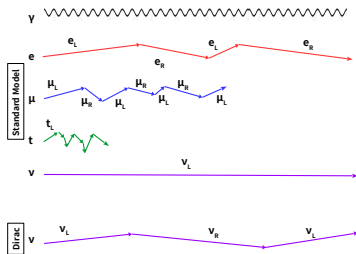
$$\mathcal{L}_\nu^{\text{Dirac}} = -\frac{v}{\sqrt{2}} \bar{\nu}_L Y^\nu \nu_R + \text{h.c.} \quad \rightarrow \quad \boxed{\text{Chirality}}$$

BUT weak interaction only *talk* to LH particles
AND only weak interaction for neutrinos

⇒ No LH neutrino described by the SM

⇒ Need to extend the SM

Sterile neutrino



Proof of massive neutrinos lead to extensions of SM

The mass generation mechanism depends on the neutrino nature

- $m_\nu \ll m_l$
 - Neutrinos are neutral
- } \Rightarrow origin of neutrino masses different from those of charged fermions ?

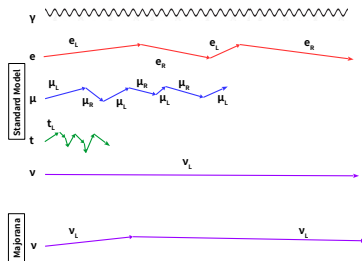
Majorana particles: The neutrino is its own antiparticle

Majorana mass term in the Lagrangian

$$\mathcal{L}_\nu^{\text{Majorana}} = \frac{1}{2} m_\nu \bar{\nu}_L^c \nu_L + \text{h.c.}$$

No extra particle

- Lepton Number Violation (LNV) with $\Delta L = 2$ (forbidden in the SM)
 \rightarrow could explain matter/antimatter asymmetry in the universe
- Seesaw mechanisms:
heavy ν_R mixes with ν_L and generates light Majorana masses for active neutrinos
 \rightarrow could explain smallness of neutrino masses



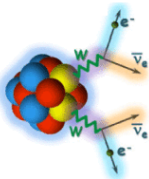
Probe: Neutrinoless double beta decay

Beyond the Standard Model

Probe neutrino nature with neutrinoless double beta decay

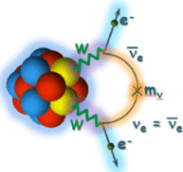
Double beta decay = 2 simultaneous neutron decays **inside the nucleus**

$2\nu\beta\beta$



- Allowed in SM
- Has been observed in several isotopes
- $T_{1/2}^{2\nu\beta\beta} \sim 10^{18} - 10^{21}$ years

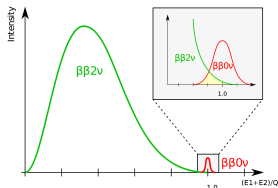
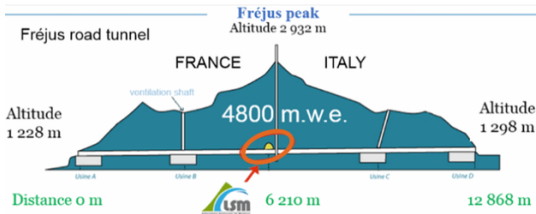
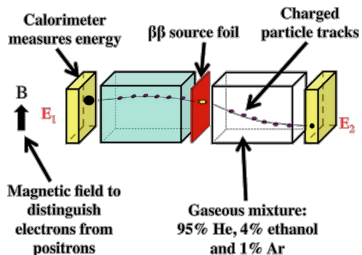
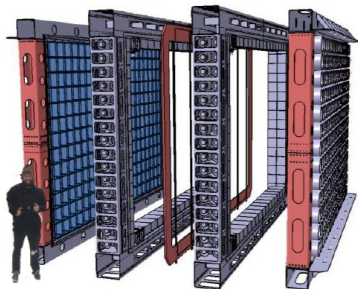
$0\nu\beta\beta$



- Forbidden in SM
- Possible only if neutrinos are Majorana particles
- $T_{1/2}^{0\nu\beta\beta} > 10^{24} - 10^{26}$ years

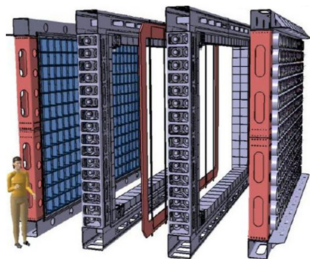
The SuperNEMO demonstrator installed @LSM

Tracko-calorimeter experiment + ultra-low background



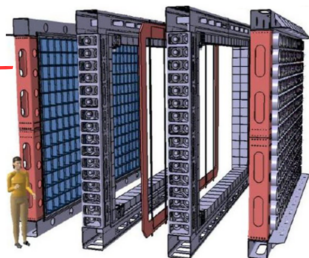
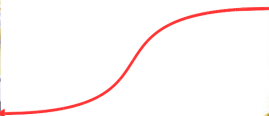
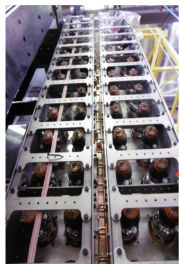
The SuperNEMO demonstrator

First commissioning datas were taken!



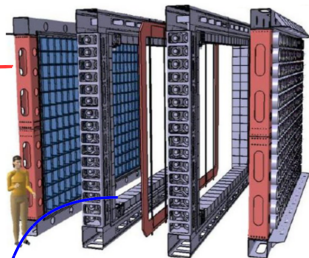
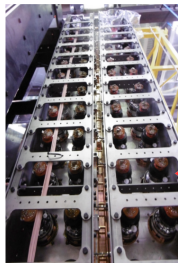
The SuperNEMO demonstrator

First commissioning data were taken!



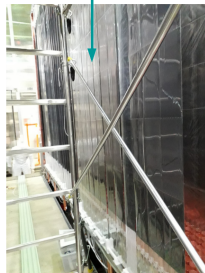
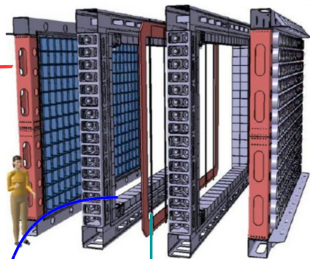
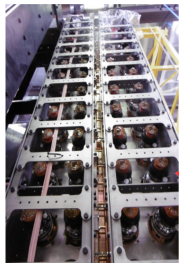
The SuperNEMO demonstrator

First commissioning datas were taken!



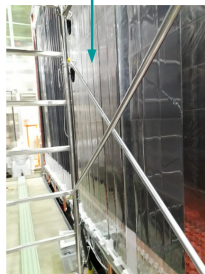
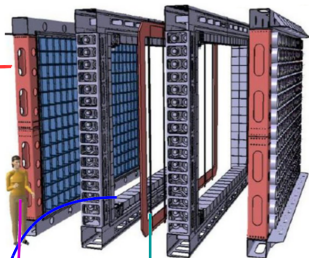
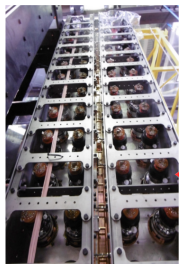
The SuperNEMO demonstrator

First commissioning data were taken!



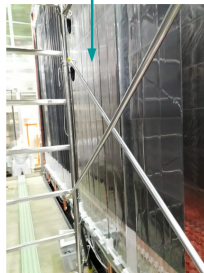
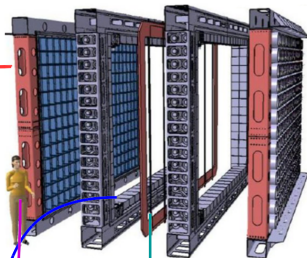
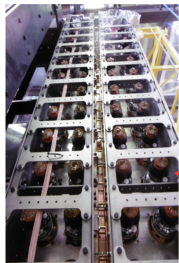
The SuperNEMO demonstrator

First commissioning datas were taken!



The SuperNEMO demonstrator

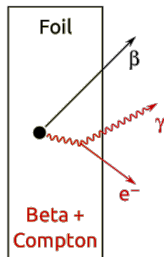
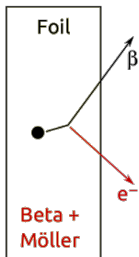
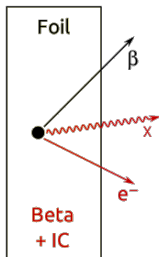
First commissioning datas were taken!



Noise background: events which mimic the signal

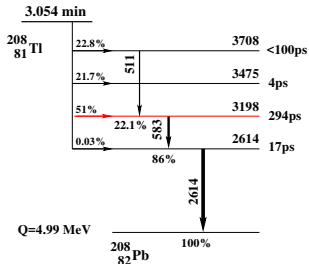
Main backgrounds:

- External γ from natural radioactivity Background events for $2\nu\beta\beta$
 - Origin: detector PMs, laboratory
 - $E_\gamma < 2.6$ MeV
- Internal contamination of Radon in the tracker Background events for $0\nu\beta\beta$
- Internal contamination in the source foils Background events for $0\nu\beta\beta$
 - ^{208}Tl from ^{232}Th decay chain
 - ^{214}Bi from ^{238}U decay chain



Analysis to reject ^{208}Tl background events

Use the electron of internal conversion of ^{208}Tl :



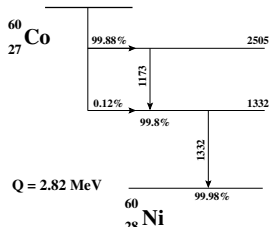
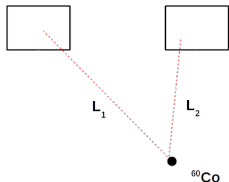
2e topologies in [2.7,3.2] MeV:
75% of ^{208}Tl decays are β decay +
internal conversion of the 2.615 MeV- γ

Delayed electron



\Rightarrow Possible rejection of ^{208}Tl background events

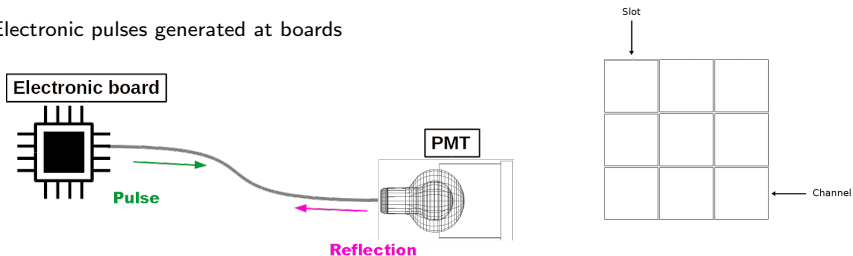
^{60}Co source to determine σ_t in optical modules:



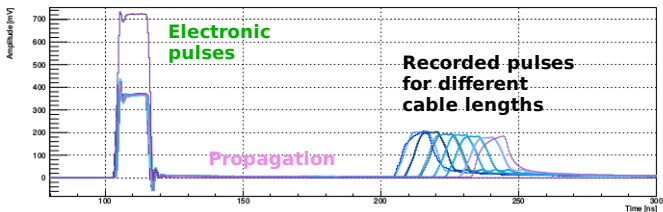
Calorimeter walls cabled: Commissioning data analysis

Reflectometry tests on signal cables for one board (slot)

Electronic pulses generated at boards



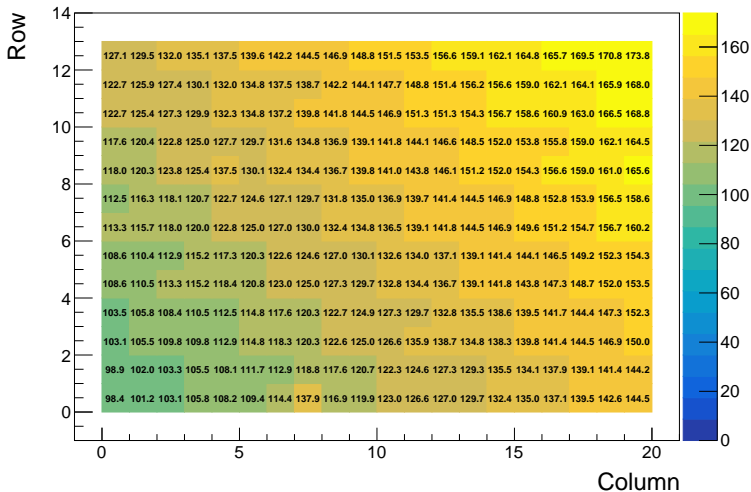
Multiple pulses for each PMT (13 channels - 20 slots)



Calorimeter walls cabled: Commissioning data analysis

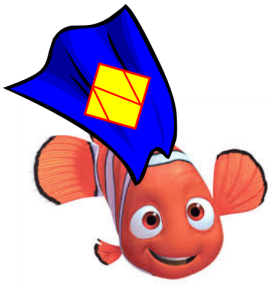
Reflectometry tests on signal cables for one board (slot)

Mean time difference between 2 peaks (CFD)

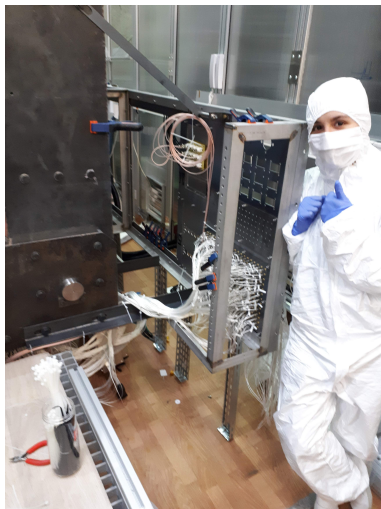


- Neutrinoless double beta decay is the best known process to probe LNV
- SuperNEMO is a tracko-calorimeter detector searching for $0\nu\beta\beta$
- In installation @Modane underground laboratory
- First commissioning data analysis in progress

Pandas & Cartoon fishes



Thank you



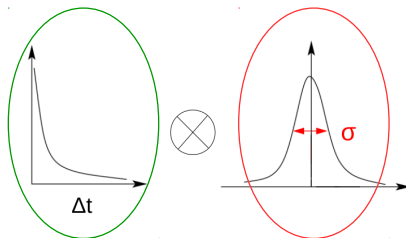
Back up

Exponentially modified gaussian

Convolution of an exponential and a gaussian function:

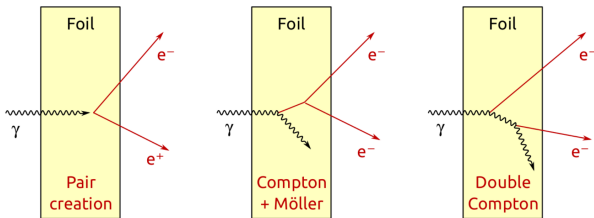
$$f(x, \mu, \sigma) = \frac{\lambda}{2} \exp\left(\frac{\lambda}{2}(2\mu + \lambda\sigma^2 - 2x)\right) \operatorname{erfc}\left(\frac{\mu + \lambda\sigma^2 - x}{\sqrt{2}\sigma}\right)$$

with $\mu = 0$, $\sigma = \sqrt{\sigma_t^2 + \sigma_{\left(\frac{L}{\beta c}\right)}^2 + \sigma_T^2}$ and $\tau = 294\text{ps}$

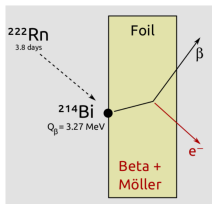


Background scheme for external gammas and Radon

External gamma background:

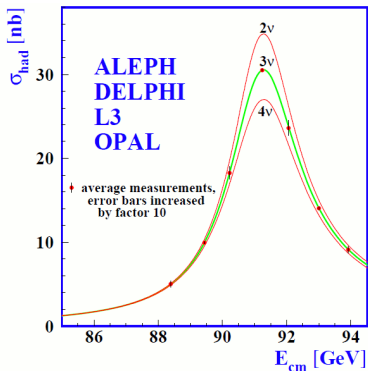


Internal Radon background:



Active light neutrino number with the total Z width

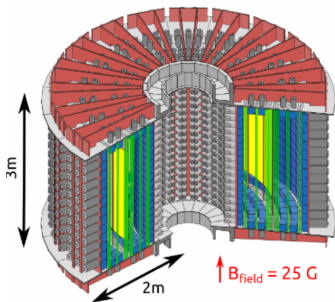
The most precise measurements of the number of light neutrino types come from studies of Z production in e^+e^- collisioners.



$N_\nu = 2,984 \pm 0,012$ light neutrinos

From NEMO3 to SuperNEMO

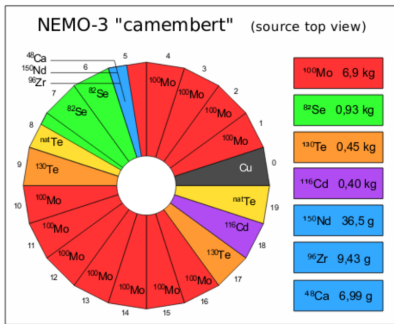
(Pour Laurent : je voudrais mettre ici un tableau comparatif NEMO3/SuperNEMO, est-ce que tu aurais a quelque part en stock ?)



sources
60 mg/cm² foils
10 kg of ββ isotopes

tracker
6180 Geiger cells
vertex resolution :
 $\sigma_{xy} - 3 \text{ mm}$ $\sigma_z - 10 \text{ mm}$

calorimeter
1940 optical modules :
polystyren scintillators
+ 3" and 5" PMTs
FWHM_E ~ 15% / $\sqrt{E_{\text{MeV}}}$
 $\sigma_t - 250 \text{ ps}$



Neutrinoless double beta decay sensitivity:

$$\tau_{1/2}^{0\nu\beta\beta} \propto \frac{a\epsilon}{m_{\text{mol}}} \sqrt{\frac{M \times t}{N_{\text{bkg}} \times \Delta E}}$$

a - abundance of isotope

ϵ - efficiency of the detector

$M \times t$ - exposure

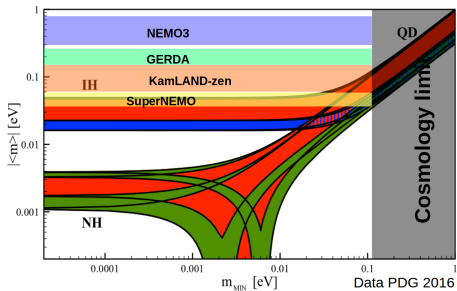
N_{bkg} - the background rate (counts.keV⁻¹.kg⁻¹.y⁻¹)

ΔE - energy resolution

To have a better sensitivity:

- High detection efficiency
- Long period data taking
- Big quantity of isotope
- Low background level

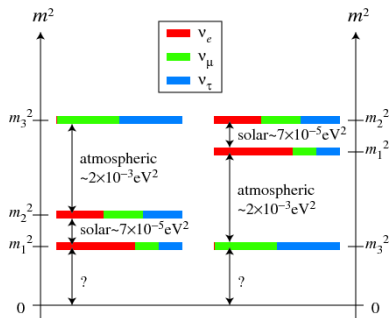
Double beta decay and mass hierarchy



Gas proportions chosen to have the wire chamber functioning at optimal performances

- 95% He: low atomic number \Rightarrow low energy losses for incoming particles (1 MeV electron: 50 keV energy loss)
- 4% ethanol: absorption of UV photons generated by Geiger plasma
- 1% Ar: ionisable \Rightarrow ease the Geiger plasma propagation

Neutrino mass ordering



The source foils

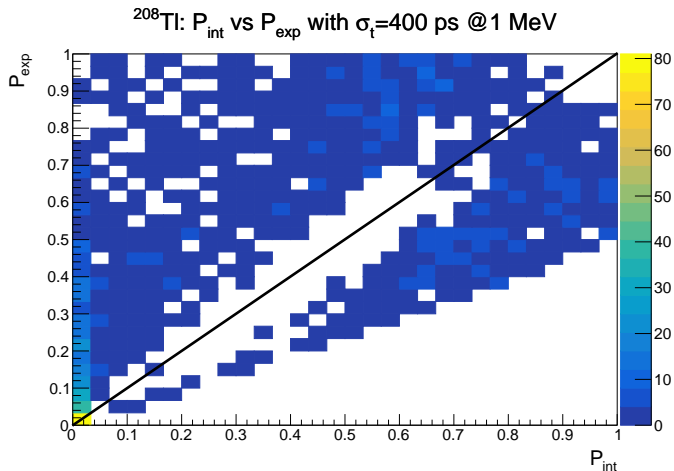
The choice of isotope

- High $Q_{\beta\beta}$ to reduce natural radioactivity background
 $Q_{\beta\beta} > 2.615$ MeV
- $(T_{1/2}^{0\nu})^{-1} = g_A^4 G^{0\nu} |M_{0\nu}|^2 \left| \frac{m_{\beta\beta}}{m_e} \right|^2$
High phase space factor $G^{0\nu}$ and matrix elements $M^{0\nu}$ to minimise $0\nu\beta\beta$ half life
- Isotopic abundance
- Highest $T_{1/2}^{2\nu}$ to reduce the $2\nu\beta\beta$ background contribution

...

Isotope	$Q_{\beta\beta}$ (MeV)	$G_{0\nu}$ (10^{-15} y^{-1})	$T_{1/2}^{2\nu}$ (y)	η (%)
⁴⁸ Ca	4.273	24.81	6.37×10^{19}	0.187
⁷⁶ Ge	2.039	2.363	1.926×10^{21}	7.8
⁸² Se	2.995	10.16	9.6×10^{19}	9.2
⁹⁶ Zr	3.350	20.58	2.35×10^{19}	2.8
¹⁰⁰ Mo	3.035	15.92	6.93×10^{18}	9.6
¹¹⁶ Cd	2.809	16.70	2.8×10^{19}	7.6
¹³⁰ Te	2.530	14.22	6.9×10^{20}	34.5
¹³⁶ Xe	2.458	14.58	2.165×10^{21}	8.9
¹⁵⁰ Nd	3.367	63.03	9.11×10^{18}	5.6

Applying cut $P_{int} > P_{exp}$ to reject ^{208}Tl background events

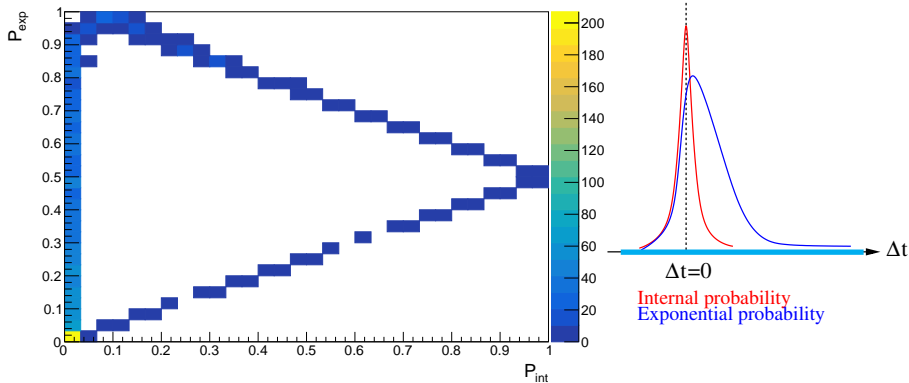


$P_{int} > P_{exp}$ allow to reject 70% of ^{208}Tl events ([2.7, 3.2] MeV and 2e topologies)

Internal and exponential probabilities

Specific case where $\sigma_t = 0$

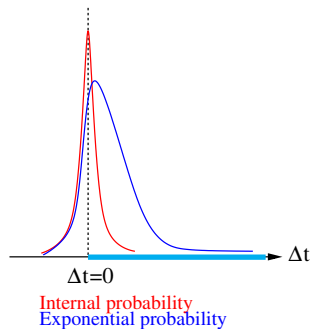
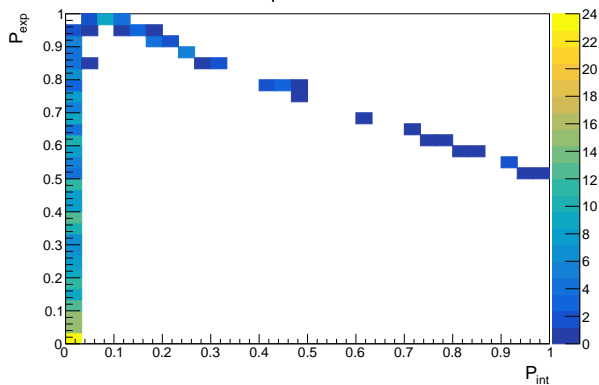
^{208}Tl : P_{int} vs P_{exp} with $\sigma_t=0$ ps @ 1 MeV



Internal and exponential probabilities

$$\Delta t > 0$$

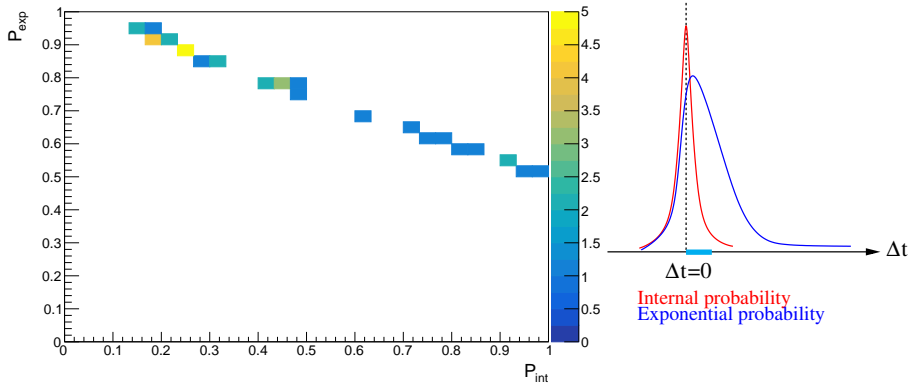
^{208}Tl : P_{int} vs P_{exp} with $\sigma_t=0$ ps @ 1 MeV



Internal and exponential probabilities

$$\Delta t \in [0, x_{max}^{exp}]$$

^{208}Tl : P_{int} vs P_{exp} with $\sigma_t=0$ ps @ 1 MeV



Internal and exponential probabilities

$$\Delta t < 0$$

^{208}Tl : P_{int} vs P_{exp} with $\sigma_t=0$ ps @ 1 MeV

