# The Neutrino Platform and the neutrino programme at CERN Sandro Palestini

#### Colloquium Prague v19

2019

24-25 October 2019 J. Heyrovsky Institute of Physical Chemistry Europe/Prague timezone



# The CERN Neutrino Platform

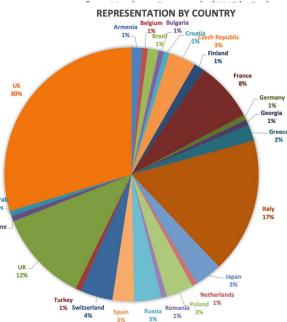
- We are following the mandate give by the EU strategy in 2013: "pave the way ...."
- Main goal : compact the European groups around the projects of the future short and long Neutrino baselines
- CERN as a facility for R&D on future technologies (HW and SW) and partner in several neutrino research programs
- A new phase of the Neutrino Platform has been recently approved
- including most notably a second campaign of test beam for ProtoDUNE detectors in2021.

#### CERN v Platform (today, 146 institutions)

Alikhanian National Science Laboratory (YerPhi), Armenia; Institute of Theoritical Physics and Modeling, Armenia, Armenia; Theoretical Nuclear Physics Research Group, Department of Physics and Astronomy, Ghent University, Belgium; Campinas University, Brazil, Brazil; Federal University of ABC, Brazil, Brazil; University of Sofia, Department of Physics, Bulgaria, Bulgaria; Ruder Boskovic Institute, Zagreb, Croatia, Croatia; Acad. of Sciences of the Czech Rep., Czech Republic; Charles University, Faculty of Mathematics and Physics, Czech Republic; Czech Technical University, Czech Republic; Institute of Experimnetal and Applied Physics, Czech Technical University in Prague, Czech Republic; Institute of Physics, Acad. of Sciences of the Czech Rep., Czech Republic; University of Jyvaskyla - Department of Physics, Finland, Finland; University of Oulu, Finland, Finland; CEA/IRFU,Centre d'etude de Saclay Gif-sur-Yvette - IRFU, France; Centre d'Etudes Nucléaires de Bordeaux-Gradignan, France; Centre National de la Recherche Scientifique - LAPP-Laboratoire d'Annecy-le-Vieux de Physique des Particules, France; IN2P3 - I ^L, France; Laboratoire de physique nucléaire et

Very intense R&D in various domains on new de hautes énergies Paris (LPNHE), France; OMEGA Ecole Polytechnique IN2P3 / CNRS. France: Universite Clause France; Universite de Paris VII - Laboratoire APC - Astroparticules et Cosmologie, France; Université 🗅 Diderot, France; Université Savoie Mont Blanc, France; High Energy Physics Institute -National Technical University of Athens, NTUA, Greece; The University of Athe Fisica E. Pancini, Università di Napoli Federico II, Italy; Dipartimento Active partnership of CERN on external Matematica e Fisica, Universita del Salento, Lecce, Italy Italy; INFN e Laboratori Nazionali di Frascati " experiments/facilities in the US, Japan Pavia, Italy, Italy; Laboratori Naziotechnologies Effort to bring to new communities the degli Studi e INFN Milano - Sezi 🗸 organization model of the LHC project and Universita e INFN, Napoli - Sezior Italy; High Energy Accelerator Res Poland, Poland; Institute of Nuclea Otwock, Poland, Poland; Warsaw U Russia, Russia; Lebedev Physical Ins Sciences - Institute for Nuclear Re Medioambientales y Tecnológicas (CIE 🗸 collaborations! Fisica Corpuscular (IFIC), Spain; Unive ean Organiz. for Nuclear Res. (CERN Zuerich - ETH Zurich Institute for Par Jept. de Phys. Nucl. et Corpuscul., Switzerland; Ul Hochenergiephysik, Switzerland; Univer-University (METU), Ankara, Turkey, Turke and College London, UK; Lancaster University, UK; Queen Mai

Rutherford Appleton Lab. - Rutherford Appleton Laboratory, UK; University of Birmingham, UK; University of Edinburgh, UK; University of Glasgow, UK; University of Liverpool, UK; University of London - University College Oxford - Particle Physics, UK; University of Sheffield, UK; University of Sussex, UK; University of Warwick, UK; New York University Abu Dhabi, United Arab Emirates; Argonne National Laboratory, US; Brookhaven Natior University, US; Drexel University, US; Duke University, US; Fermi National Accelerator Lab., US; Harvard University, US; Lawrence Berkeley National Lab., Berkeley, US; Los Alamos National Laboratory, US; Louisiana St State University, US; Princeton University, US; SLAC National Accelerator Laboratory, US; South Dakota Sch Methodist University, Dallas, US; State University of New York (Stonybrook), US; University of Boston, US; Ur University of California Los Angeles, US; University of California, Berkeley, US; University of California, Irvine Honolulu, US; University of Houston, US; University of Iowa, US; University of Minnesota, US; University Philadelphia, US; University of Pittsburgh, US; University of Rochester - Department of Physics and Astronomy, L Arlington, US; University of Texas at Austin, US; University of Wisconsin, Madison, US; Virginia Tech, US; William  yon I - Institut de Physique Nucleaire de Lyon, niversité Pierre et Marie Curie (UPMC) et Paris Justus-Liebig-Universität Gießen, Germany;
ia, Universita di Roma, Italy; Dipartimento di Jniversita di Bologna, Italy; Dipartimento di '; INAF - Osservatorio Astrofisico di Torino, 'ione di Lecce, Italy, Italy; INFN, Sezione di & INFN, Milano-Bicocca, Italy; Università NFN, Catania - Sezione di Catania, Italy; 'idova, Italy; University of Genova, INFN, 'ty of Science and Technology, Krakow, National Centre for Nuclear Research,



### NP Projects – content of the talk

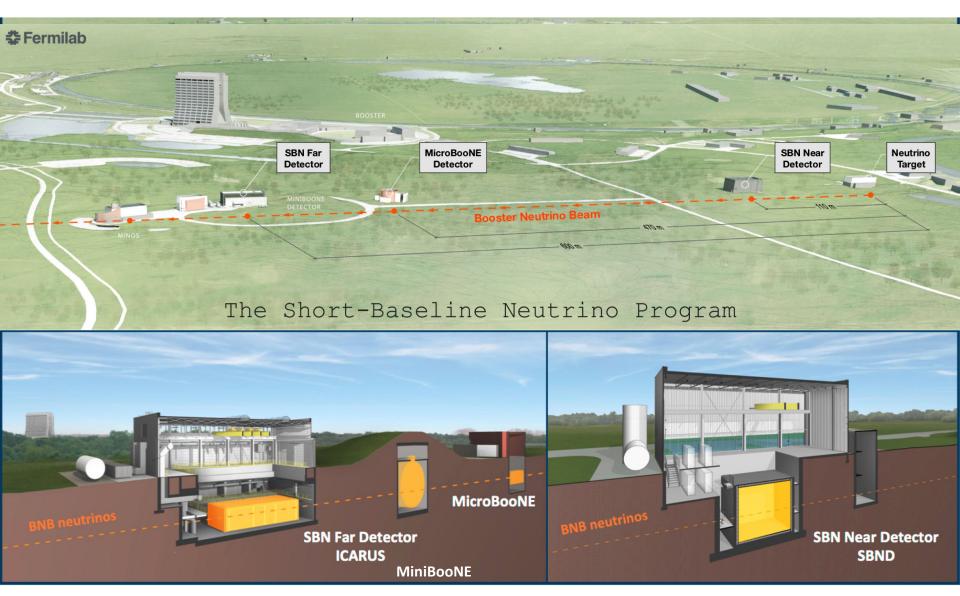
- NP01: ICARUS refurbishing and far detector in the SBN FNAL facility (almost ready for operation)
- *NP02: LAr double phase TPC demonstrator (ptotoDUNE)*
- NP03: PLAFOND generic detectors R&D
- *NP04: LAr single phase TPC demonstrator (protoDUNE)*
- NP05: BabyMind muon (near) detector for T2K (operational)
- NP06: ENUBET project
- NP07: ND280 T2K near detector upgrade

Participation in the construction and exploitation in the SBN and LBNF/DUNE in US, and T2K(ND) in Japan programmes.

Few words also about the projects FASER and XSEN.

Acknowledgement/thanks: material in this talk (often preliminary) made available by members of the initiatives listed above.

### NP01: SBN/ICARUS



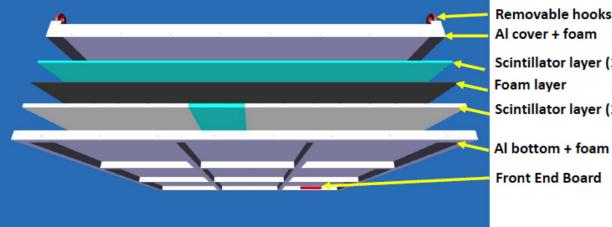
ICARUS: 476 tons, 600 m

MicroBooNE: 85 tons, 470 m

SBND: 112 tons, 110 m<sup>5</sup>

#### NEW LIGHT DETECTION SYSTEM (360 8' PMTs)





**Removable hooks** Al cover + foam Scintillator layer (10 mm) Foam layer Scintillator layer (15mm)

Cosmic Ray Tagger for position and time of charged tracks entering the TPC volume

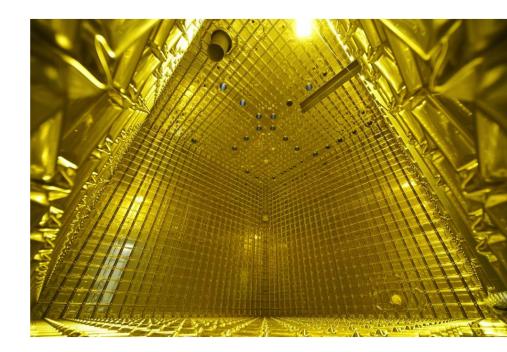
NEW

CABLIN



ICARUS detector preparation nearly complete Cool-down and filling planned for November 2019 – January 2020

## NP02, NP04: Membrane cryostat technology



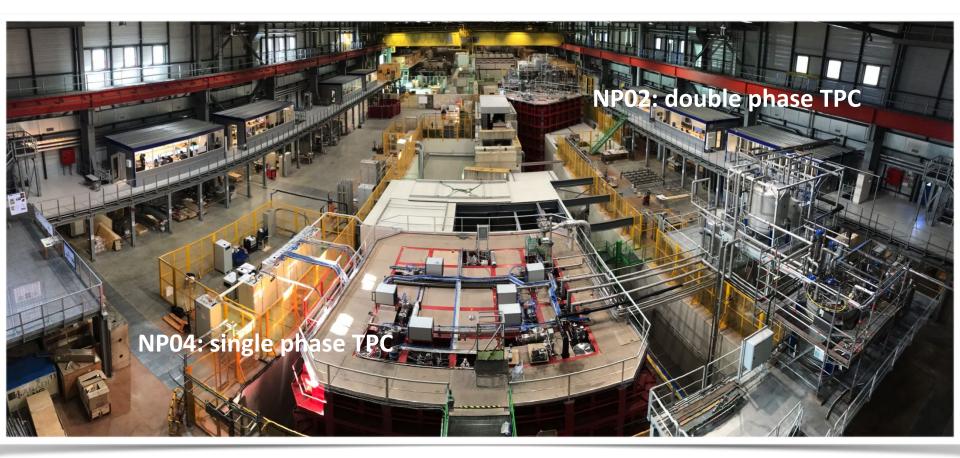
- Corrugated SS primary membrane in contact with LAr
- Plywood
- Insulation: reinforced polyurethane foam
- Secondary membrane for gas containment
- Insulation
- Plywood
- Outer vessel

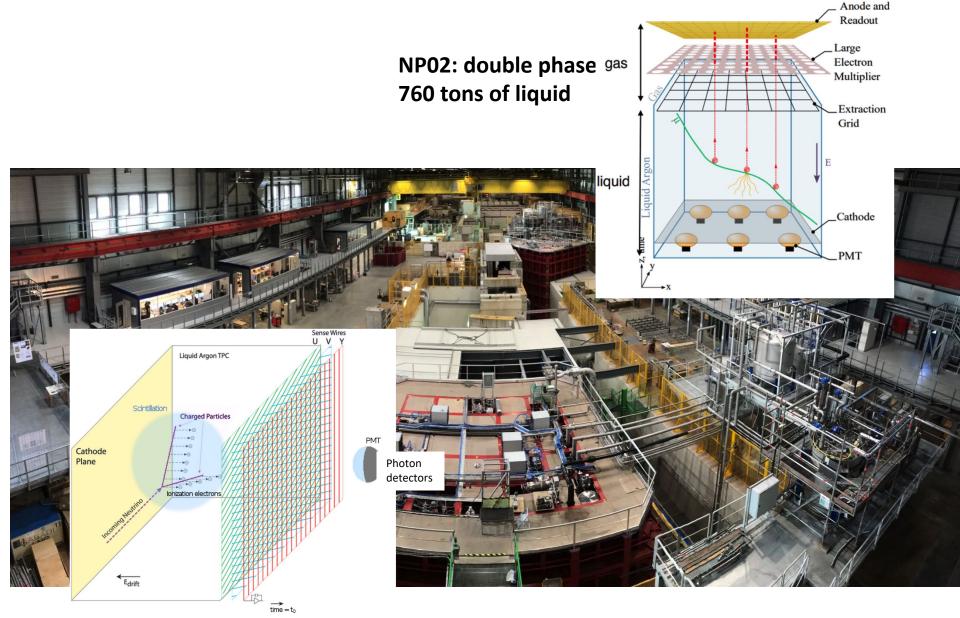
No vacuum, insulation purged with Ar

Concept developed for carrier ships. Developed for LAr in collaboration with industry (GTT)

Used for ProDUNEs, SBND, DUNE, DarkSide ...

### EHN1 layout





NP04: single phase 760 tons of liquid

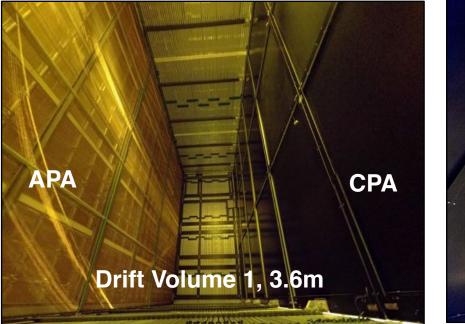
# NP04: single phase LAr TPC

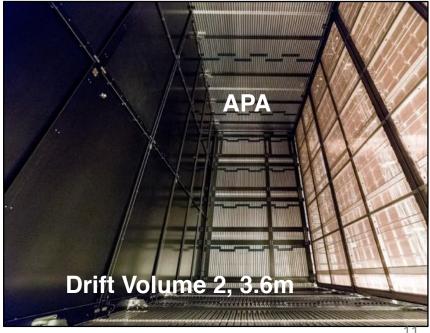
- Assembled and cold tested by May 2018
- From May to August 2018 cooling and filling
- Ready for beam in Fall 2018

6 x Anode Plane Assembly

Cathode Plane Assembly

Photon Detection





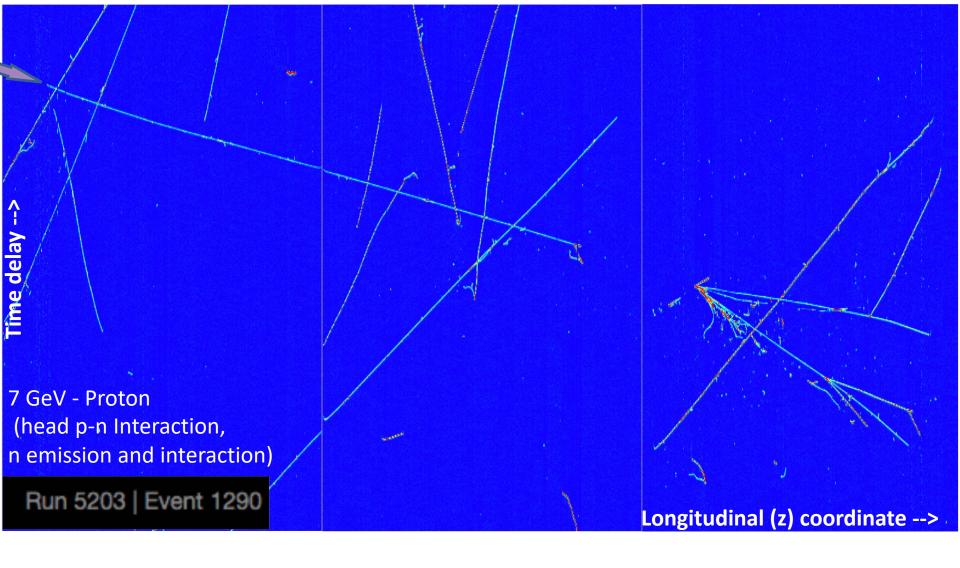
# Data – Collected beam events

Momentum	Total Triggers	Expected Pi trig.	Expected Proton trig.	Expected Electr. trig.	Expected Kaon trig.
0.3 GeV/c	269K	0	0	242K	0
0.5 GeV/c	340K	1.5K	1.5K	296K	0
1 GeV/c	1089K	382K	420K	262K	0
2 GeV/c	728K	333K	128K	173K	5K
3 GeV/c	568K	284K	107K	113K	15K
6 GeV/c	702K	394K	70K	197K	28K
7 GeV/c	477K	299K	51K	98K	24K
All momenta	4175K	1694K	779K	1384K	73K

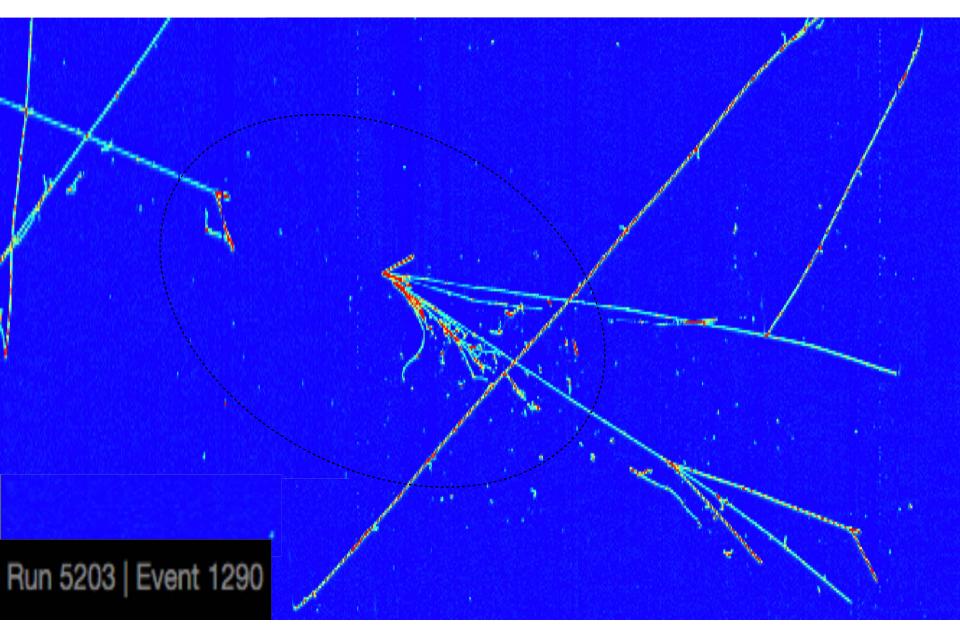
\* Numbers based on beam simulation only

#### **Over 4 million beam events!**

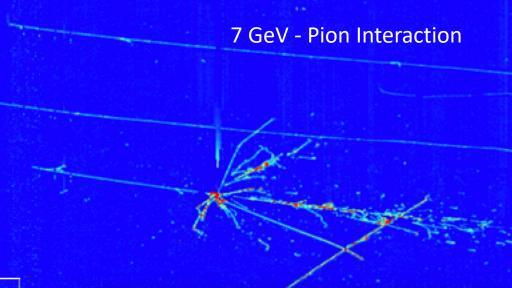
#### Over 20 million cosmics events! 12

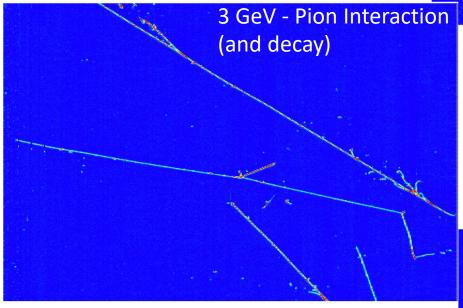


APA3



LAr TPC data of unprecedented quality





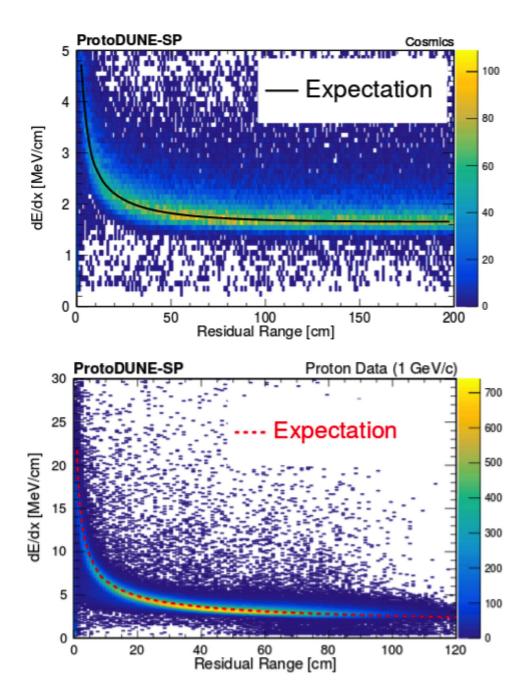


2 GeV - Electron shower

1 GeV - Pion Interaction (Absorption —> 2 p)

Detector Parameter	Minimal Requirement	Goal	ProtoDUNE Performance
Electric Drift Field	> 250 V/cm	500 V/cm	500 V/cm
Electron Lifetime	> 3 ms	10 ms	> 7 ms *
Electronics Noise	< 1000 enc	ALARA	450-750 enc s/n = 4050

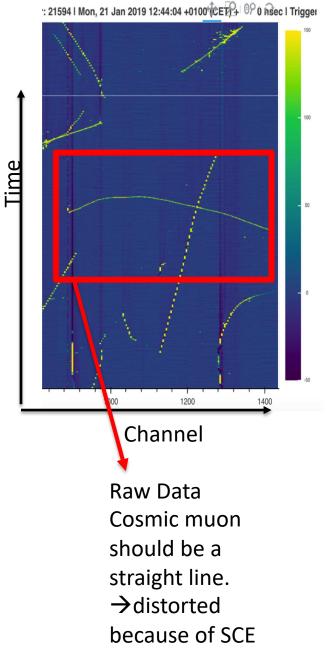
\* Saturation limit of the internal purity monitor, real value probably ~ 12-13 ms



dE/dx vs. residual track length for cosmic muons and beam protons, after detector calibration (mainly corrections of spacecharge effects)

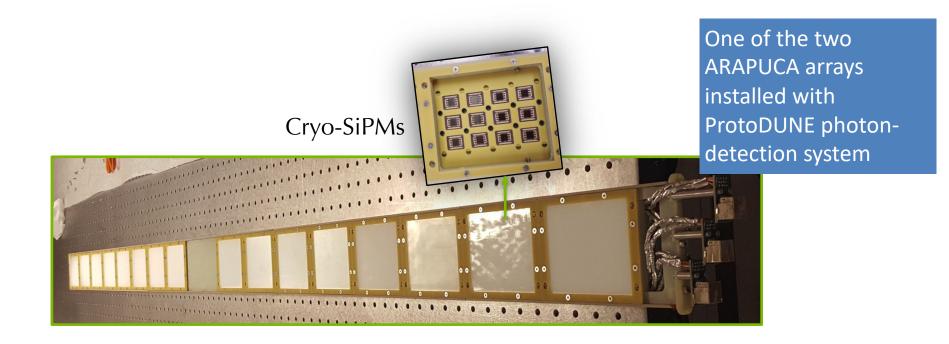
### Space Charge Effect (SCE) in large LAr TPC operated on surface

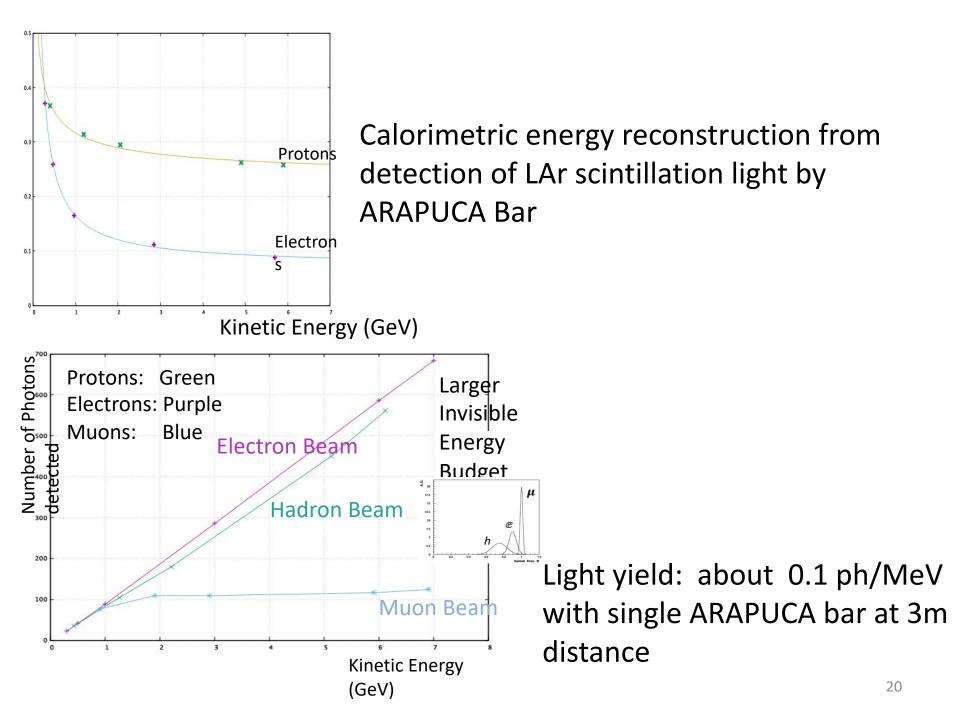
- Positive lons accumulate in the drift volume, affecting the electric field
- SCE depend on cosmic ray flux, ions mobility, E-field and drift distance in the LAr TPC
  - Drift field larger (lower) at cathode (anode)
  - x(t) distorted (partially compensated, larger effect near center of drift region)
  - Mismatch between V(x) induced by space charge and V(x) defined on field cage generates transverse field components near the detector side surfaces
  - This effect is cumulative and largest for electron drifting from the cathode: the apparent y, z coordinates are moved inside the detector (as much as 35 cm in ProtoDUNE/SP)
  - The drift velocity of is slower than the convective flow of LAr (due to thermal gradients and LAr recirculation) additional complexity
- Need a 3D correction map, use tracks with geometrical constraints:
  - Side walls crossing
  - Cathode crossing
  - Anode piercing
  - CRT tagged (cosmic ray tagger, beam tagger)
- SCE corrections are needed for track reconstruction and dE/dX calibration



# **ARAPUCA Photon Detector design**

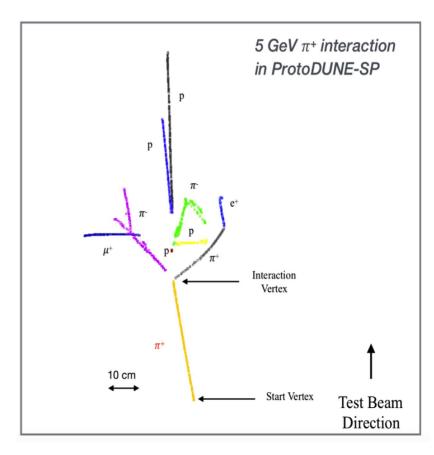
- 12 (or 6) cryo-SiPMs per Cell passively ganged
- Dichroic (short-pass) filter optical window: 9.8 x 7.8 cm<sup>2</sup>
- p-TP deposited on outer surface of Dichroic glass, TPB on inner surfaces deposited on VIKUITI Reflective Foil



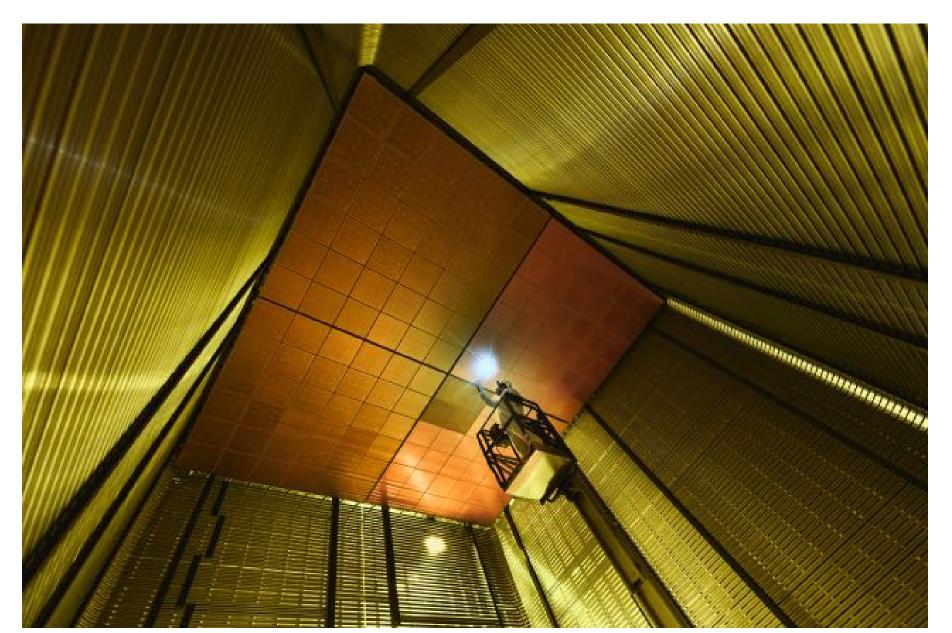


# Automated pattern recognition and reconstruction software

- Pandora progress to improve reconstruction specifically for ProtoDUNE.
- Adaptive Boost Decision Tree based Beam Particle ID:
  - Efficiencies: 72.3% for beam and 94.5% for cosmic muons

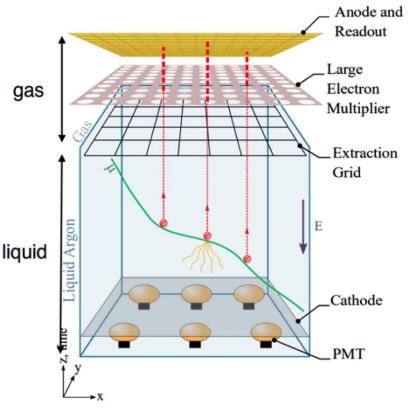


# NP02: double phase LAr TPC



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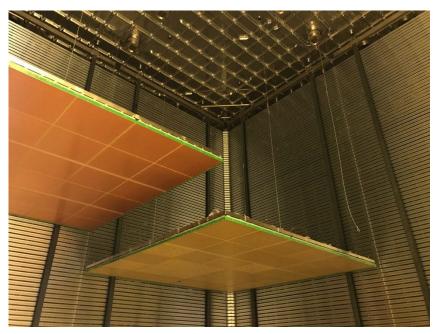
- ✓ ProtoDUNE/DP as a real first large scale demonstrator (6 m max. drift)
- ✓ More critical technology (LEM amplifiers stability, high HV system=600KV) which needs more R&D in the future to be ready for a final DUNE detector.
- ✓ Environmental conditions (liquid stability and purity, space charge effects, positive ions back drift into the liquid ...) are the critical points to first experimentally address and understand.



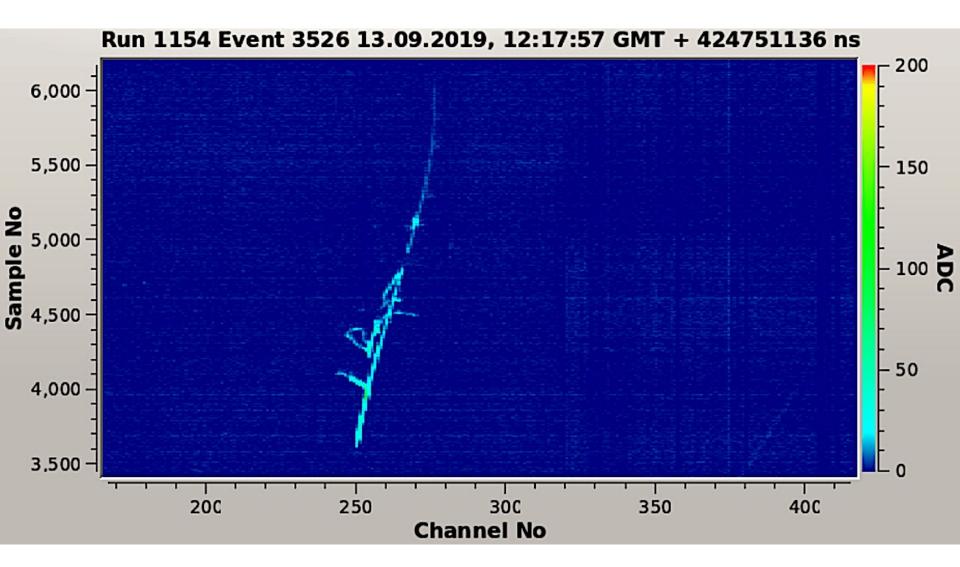
# NP02: double phase

- Final readout units assembled and cold tested by end of 2018
- Installation completed in spring 2019
- Purging, cooling and filling in summer

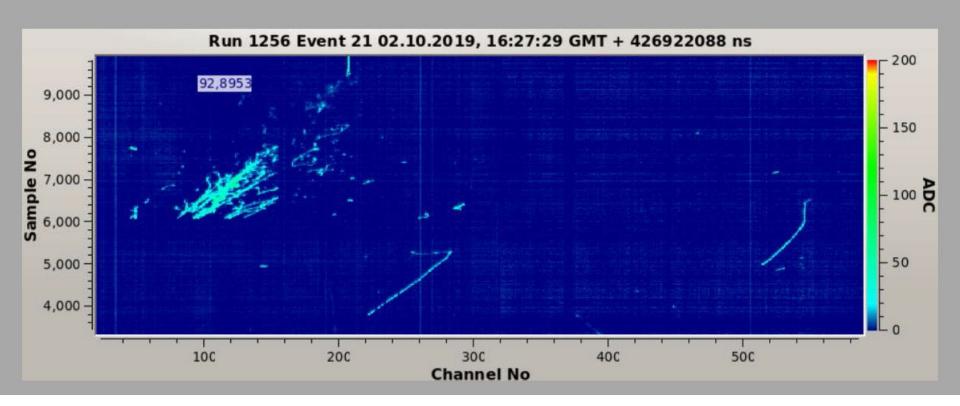








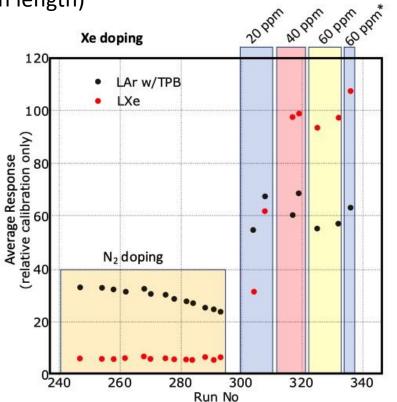
One of the first tracks observed in ProtoDUNE/DP [Fermilab News and CERN News, Oct. 9, 2019] Event with electromagnetic showers and two muon decays, LEMs at 3.1 kV at 1010 mbar, grid at 6100 V



A more recent event form ProtoDUNE/DP

## R&D activity on LAr detectors

- Other technologies for TPC sensors are considered for the 4<sup>th</sup> module of DUNE
- R&D also on **photon detection** for LAr scintillation light:
- Besides the wave-length-shifters used e.g. in ARAPUCA (p-terphenyl and TPB) and PMTs (TPB most frequently), the use of diffused Xenon in LAr is of particular interest (higher light collection, longer Rayleigh length)
- The stability of the Xe-Ar mixture is being tested in a small prototype – and being considered as the last test of the first run of ProtoDUNE/SP.
- Nitrogen pollution affects the light emission in LAr, and it is relevant to know how Xenon competes with it.
  - (very) **preliminary** results are encouraging:

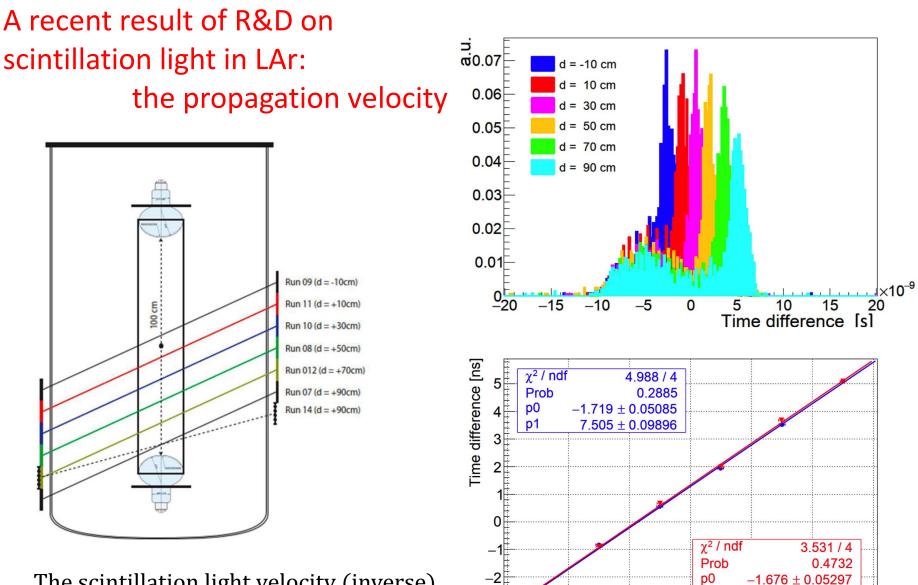


LAr w/TPB response drops to ~ 70 % with  $N_2$  doping to 5.2 ppm

All PMTs see an increase in the level of light

The amount of light seen is larger than the pure LAr response with the first Xe doping

\* HV=0



-3

-0.2

The scintillation light velocity (inverse)  $1/v_g = 7.50 \pm 0.07 stat \pm 0.05(syst) ns/m$ [Nucl.Instrum.Meth. A936 (2019) 178-179]

7.544 ± 0.1007

0.8

Distance [m]

p1

0.4

0.6

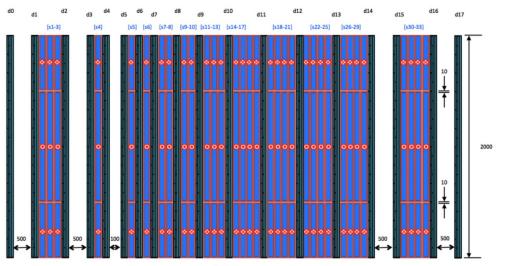
0.2

0

# NP05: BabyMind muon (near) detector for T2K

Muon spectrometer for the WAGASCI experiment (T2K beam line)

- Sandwich of 33 magnets and 18 scintillator modules
- Construction @ CERN and test beam on the PS line in summer 2017. Shipped soon after to Japan.
- Commissioning run in 2018 at J-PARC
- Physics run in 2019

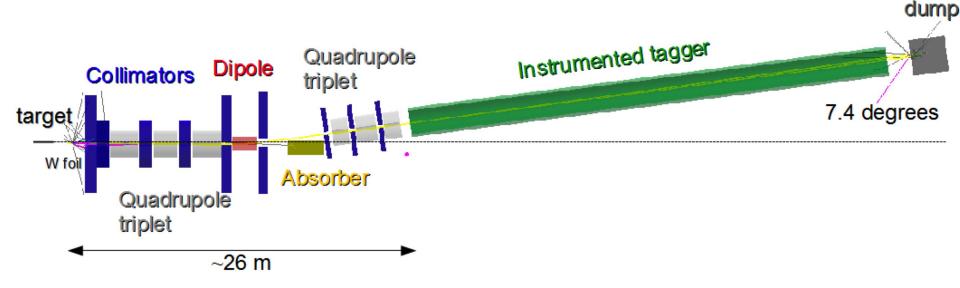






### NP06: ENUBET

#### instrumented decay tunnel for calibration of neutrino beam [CERN-SPSC-2018-034/SPSC-1-248]



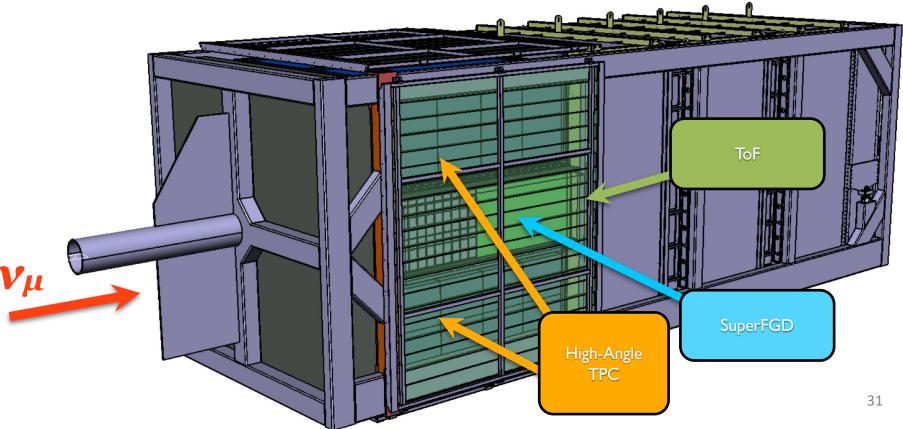
The wall of the decay tunnel is equipped with a calorimeter that can tag Ke3 decays  $(K^+ \rightarrow \pi^0 e^+ \nu_e)$ , and provide accurate measurements of  $\nu_e$  and  $K^+$  fluxes. CERN is providing support from accelerator experts. A demonstrator may come later.

Hadron

#### NP07: The upgrade of the T2K Near Detector (ND280)

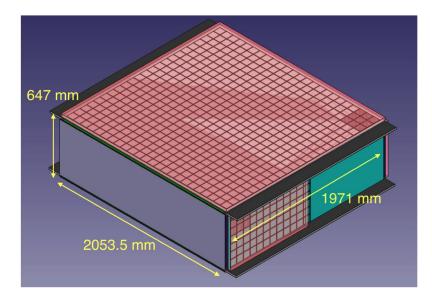
- Keep the electromagnetic calorimeter
- Horizontal active target detector: SuperFGD
- Two High-Angle TPCs
- Time-of-Flight detector around new tracker
- B-field of 0.2 T

	Current	Upgrade
Target Mass (tons)	~2	~4

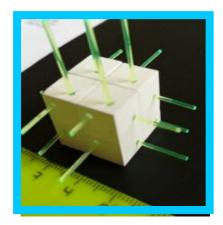


### SuperFGD for ND280

- CERN involved in the R&D of SuperFGD
- Design of the detector mechanics
- Scintillation light readout
- LED calibration system
- Building a small prototype (10k cubes)



#### 2018 JINST 13 P02006

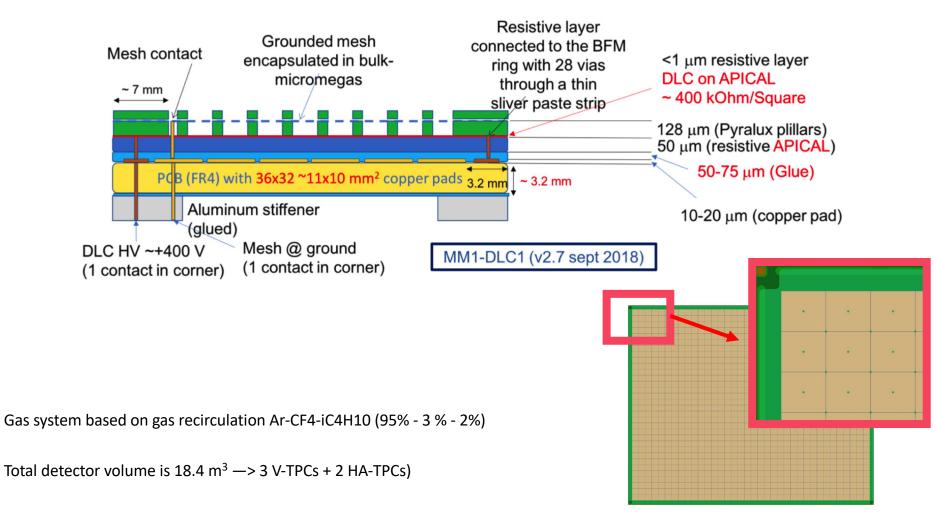




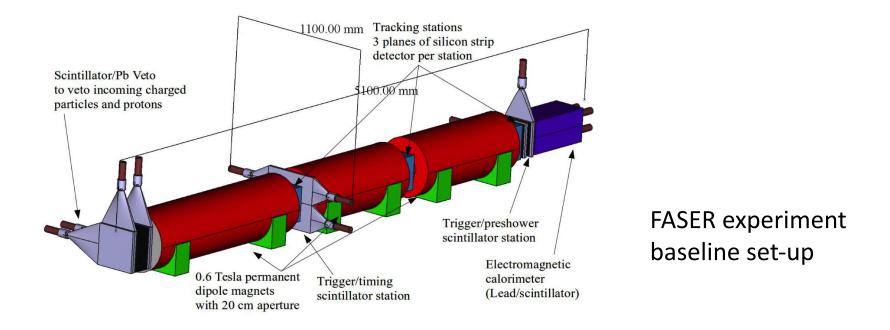
(Technology considered also for the Near Detector of DUNE)

### Resistive Micromegas for the (gas) TPC of NP280

• Resistive MicroMegas production will be done at CERN

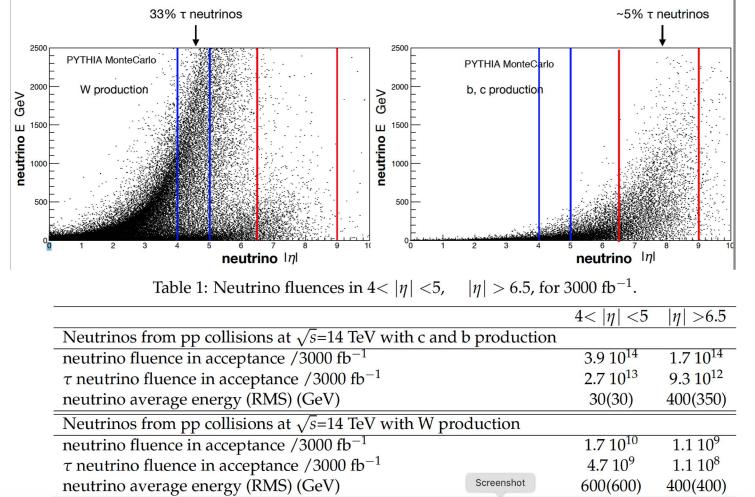


### Neutrinos from LHC: Faser- $\nu$ and XSEN



- On the LHC collision axis, at 480 m from the interaction point of ATLAS
- In an unused tunnel, shielded from LHC background
- Approved for a search of dark photon
- Preparing a proposal for an extension for the observation of neutrinos above 400 GeV, adding emulsion stacks and aiming at  $v_{\tau}$  interactions.





In a similar location, but somewhat off-axis, the **XSEN proposal** aims at neutrinos of higher energy, with a few tons of nuclear emulsions. [http://arxiv.org/abs/1903.06564]

# Conclusion

- Very intense 4 years of activities over many fronts, in line with the EU strategy recommendation (pave the way....)
- Community compact behind the project (more than 100 institutions collaborating)
- Important CERN effort as host lab (the buildings, the infrastructure, cryostats, cryogenics, DAQ, DCS, detectors components and assembly). A model of host lab interplay with the scientific collaborations for the future facilities.
- Results are very positive, with a large visibility in the international community (even outside Neutrino physics)

# **Technology** implications

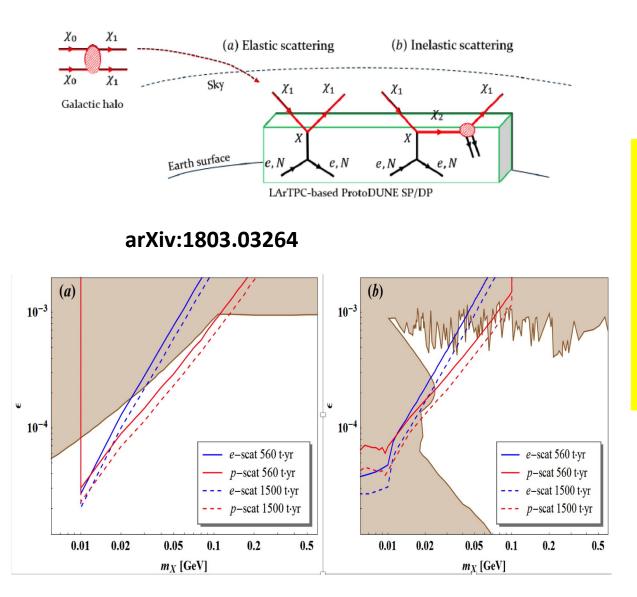
- We have now a strong and solid proposal for DUNE (basis of DUNE TDRs).
- Next phase is to apply the lessons learned and test a real DUNE module zero, before starting mass production (project submitted and approved by CERN RB).
- We have already one of our detectors taking regular data in T2K (Baby Mind). We are now opening to other activities within T2K (new T2K near detector).
- The developed LAr technology, in particular for all what concerns cryostats, cryogenics and cold electronics is opening new possibilities and interests in other fields of application ..... like dark matter search (example DARKSIDE-20k).
- In parallel we have developed many R&D applications, which will impact the future of these type of detector technologies.

# Physics research possibilities with protoDUNE

We start exploring several ideas on physics which can be done with the existing protoDUNEs (even with cosmics).

- Hadronic Argon cross section measurements (inclusive and semi-exclusive)
- Searches for boosted dark matter in cosmic rays
- Searches for fractional/mini/millicharged particles in cosmic rays
- Searches for monopoles in cosmic rays

# **Boosted Dark Matter**

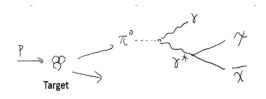


Dark matter with multiple components in order to be compatible with present DM limits

Can be searched for in the Near detector (beam Induced) and in ProtoDUNE (cosmic ray induced)

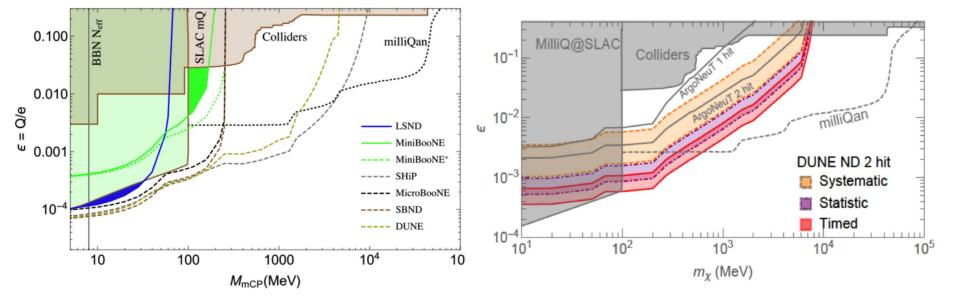
# Millicharge Searches

Recently there is a lot of revived interest to look for millicharged particles, ie particles with a charge much smaller than the electron. Signatures are elastic electron scattering and particles crossing the detector with small de/dx ionization losses



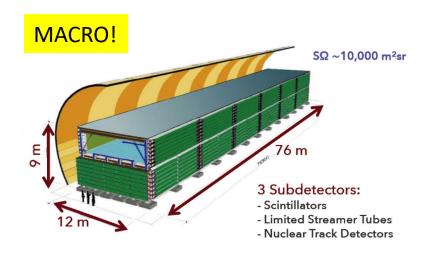
arXiv:1806.03310

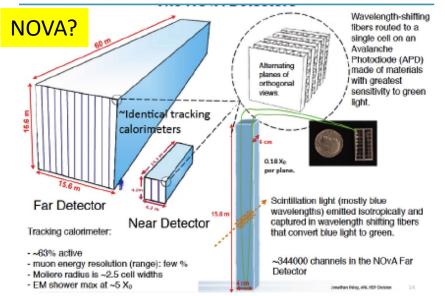
arXiv:1902.03246

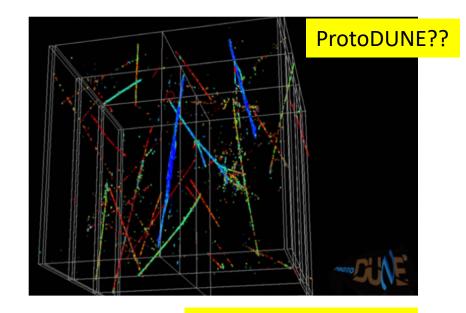


Sensitivity for DUNE ND and ProtoDUNEs under study

# **Monopole Searches**







Ionization saturation? Backgrounds? Triggers?

-Searching for highly Ionizing particles -Study of sensitivity with ProtoDUNE and the DUNE far detector launched