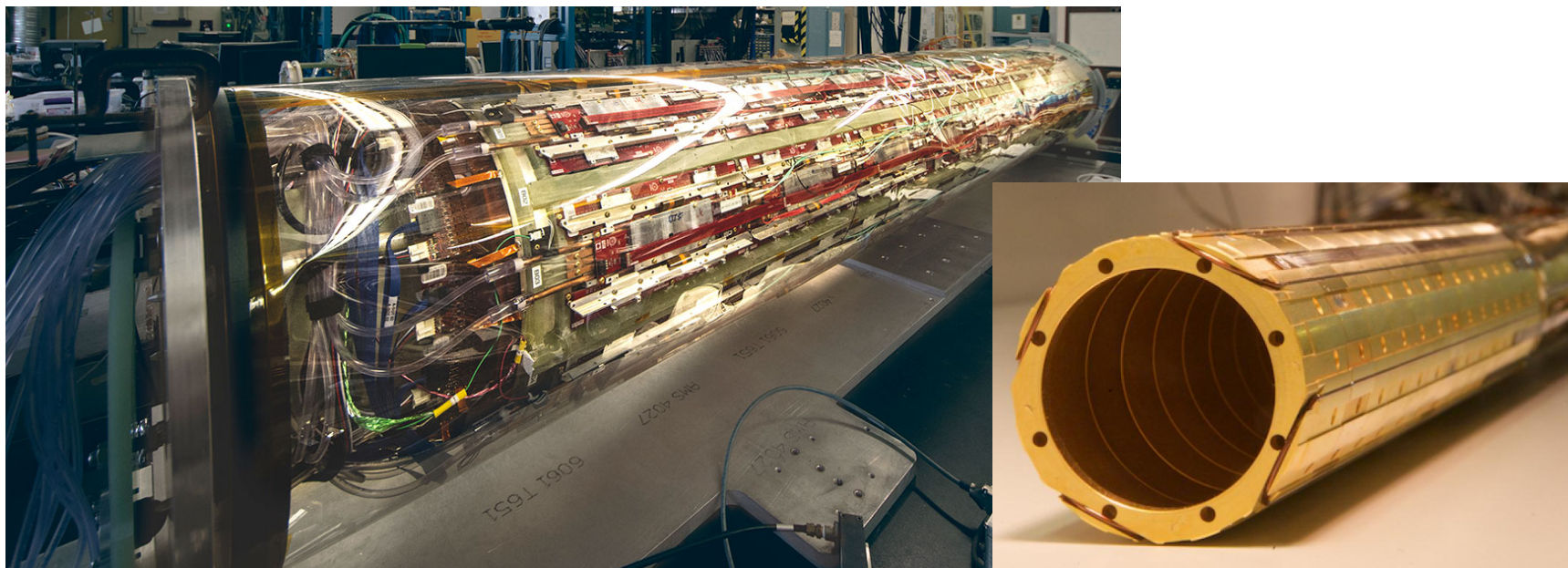


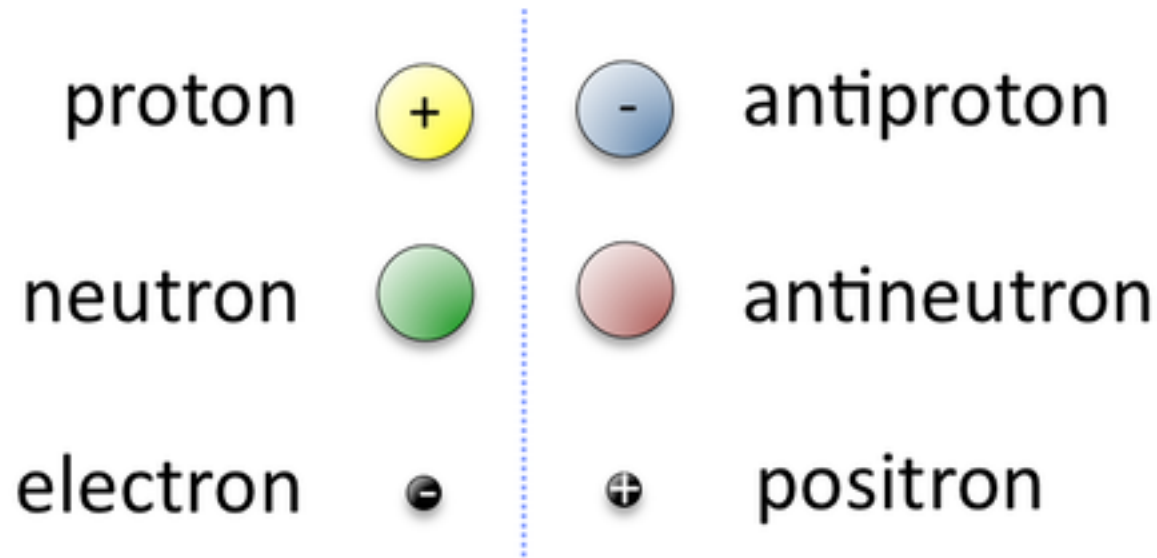
Antimatter at CERN's Antiproton Decelerator Facility (AD)



Daniel Maxwell
ALPHA Collaboration

What is antimatter ?

- For every particle there is a corresponding antiparticle with the same mass and the opposite charge.

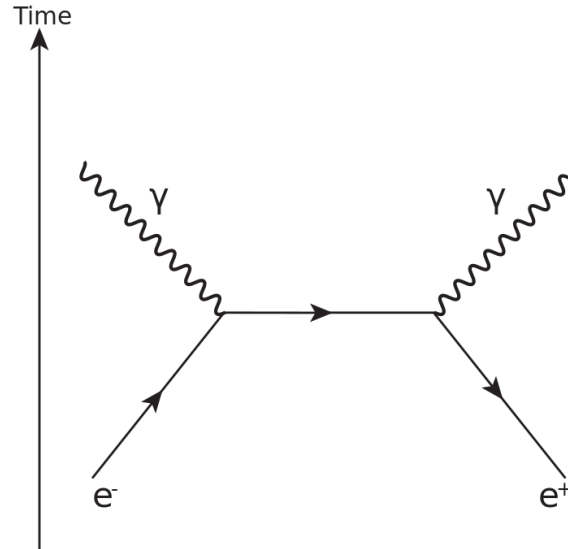


<https://web2.ph.utexas.edu/~coker2/index.files/particlesanti.htm>

Matter-antimatter annihilation

When a matter particle comes into contact with its corresponding antimatter particle they annihilate.

Electron-positron annihilation:

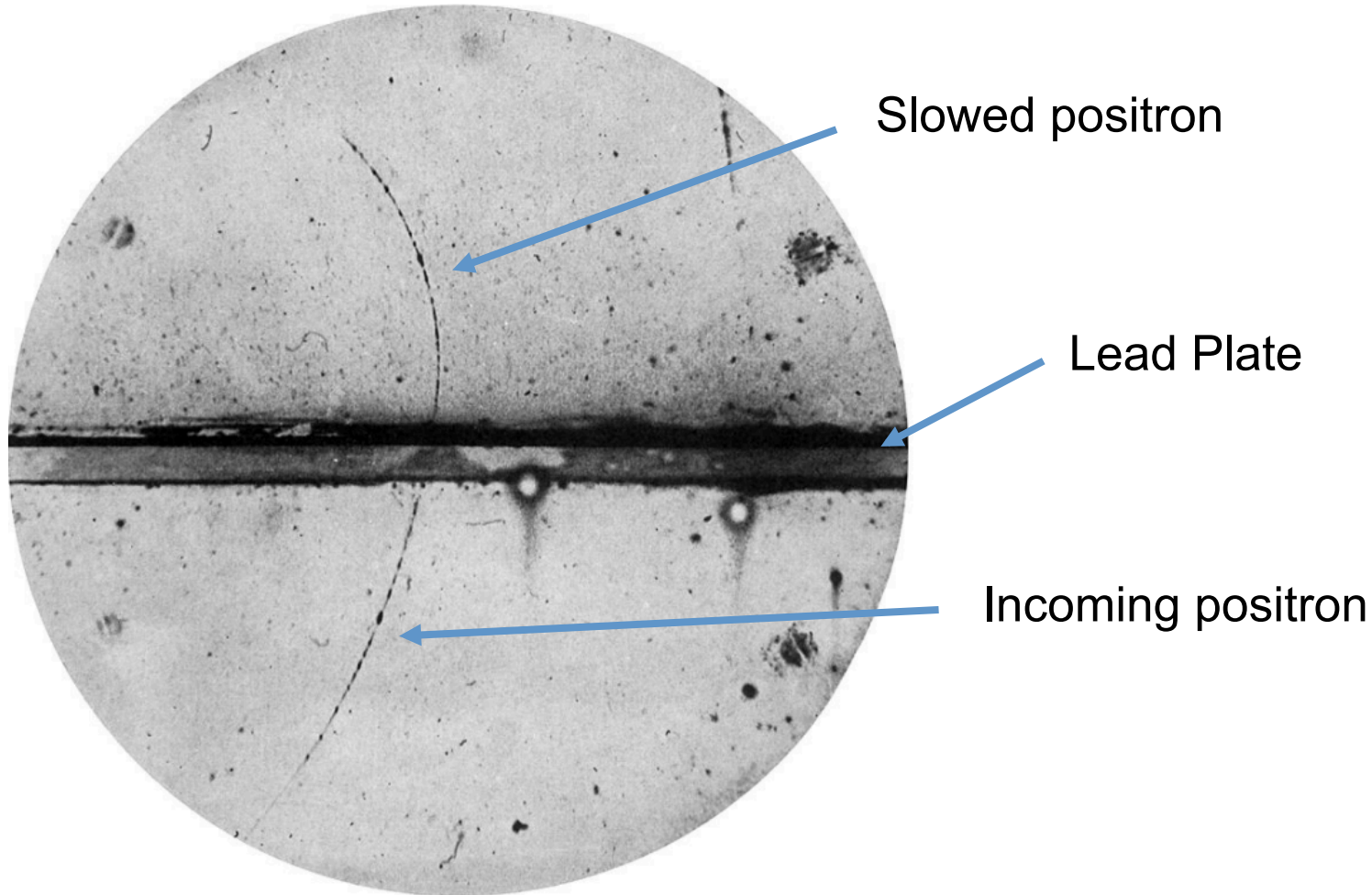


$$E = m c^2$$

Proton-antiproton annihilation is more complex and typically produces three charged pions.

The positron

- First predicted by Dirac in 1928 (although initially misunderstood).
- Observed in 1932 by Carl Anderson's measurements of cosmic rays.



Sources of antimatter

- Beta decay: $p \rightarrow n + e^+ + \nu_e$

Sources of antimatter

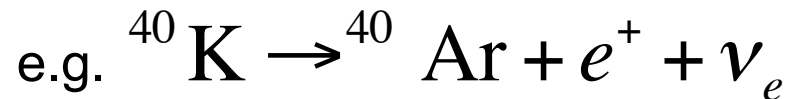
- Beta decay: $p \rightarrow n + e^+ + \nu_e$

e.g. $^{40}\text{K} \rightarrow ^{40}\text{Ar} + e^+ + \nu_e$



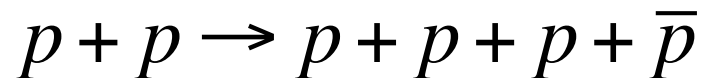
Sources of antimatter

- Beta decay: $p \rightarrow n + e^+ + \nu_e$



- Cosmic rays: high energy charged particles arrive at the earth from outer space (mostly protons). Positrons can be produced by the interaction of these particles with the earth's atmosphere.

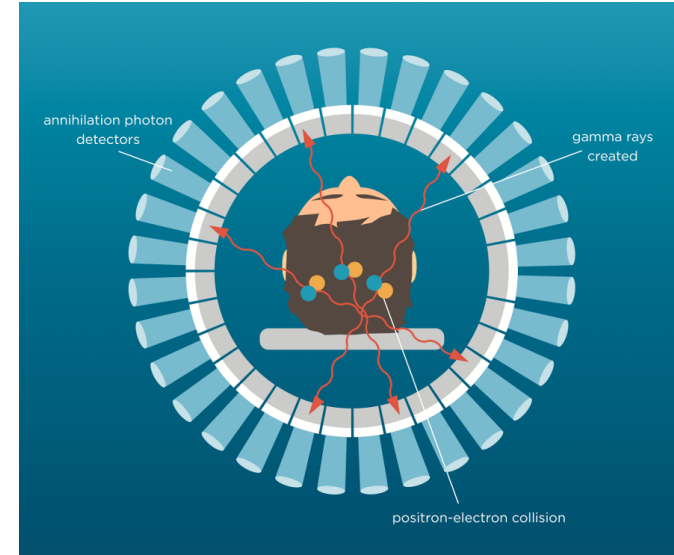
- Particle accelerators.



$$E = m c^2$$

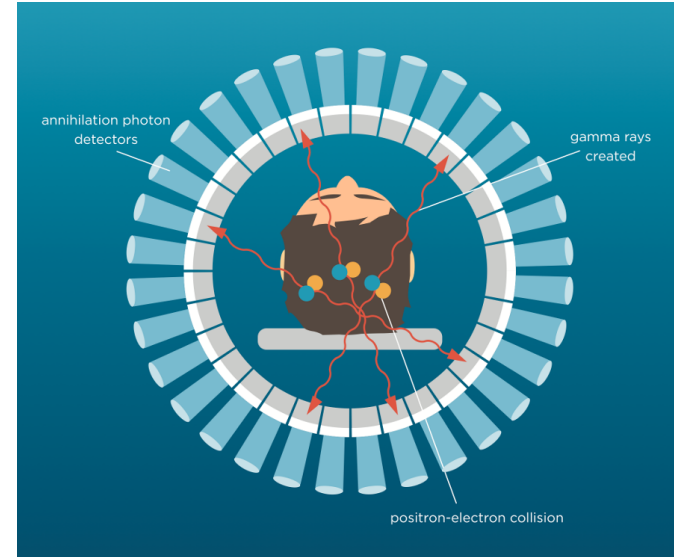
Applications of antimatter

- PET scanners.



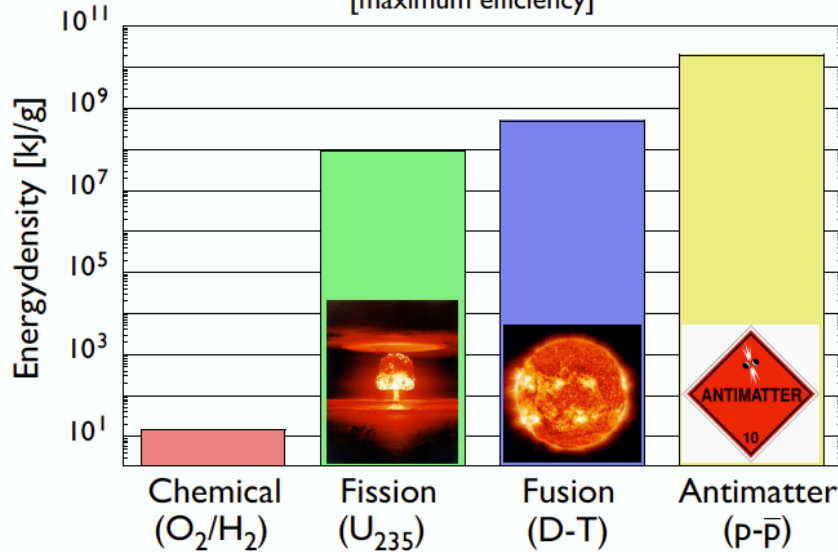
Applications of antimatter

- PET scanners.

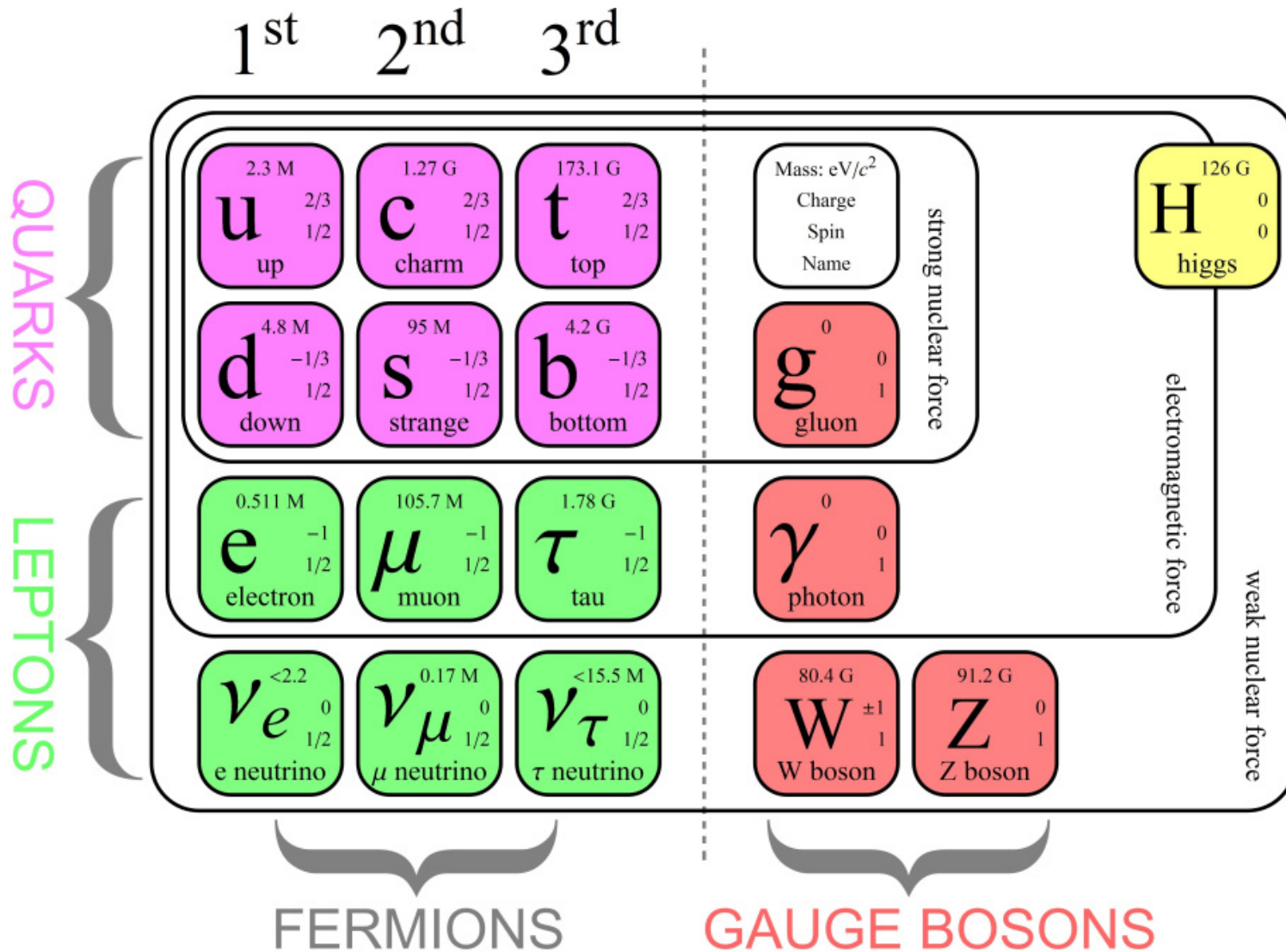


- A future fuel

Energy density comparison
[maximum efficiency]



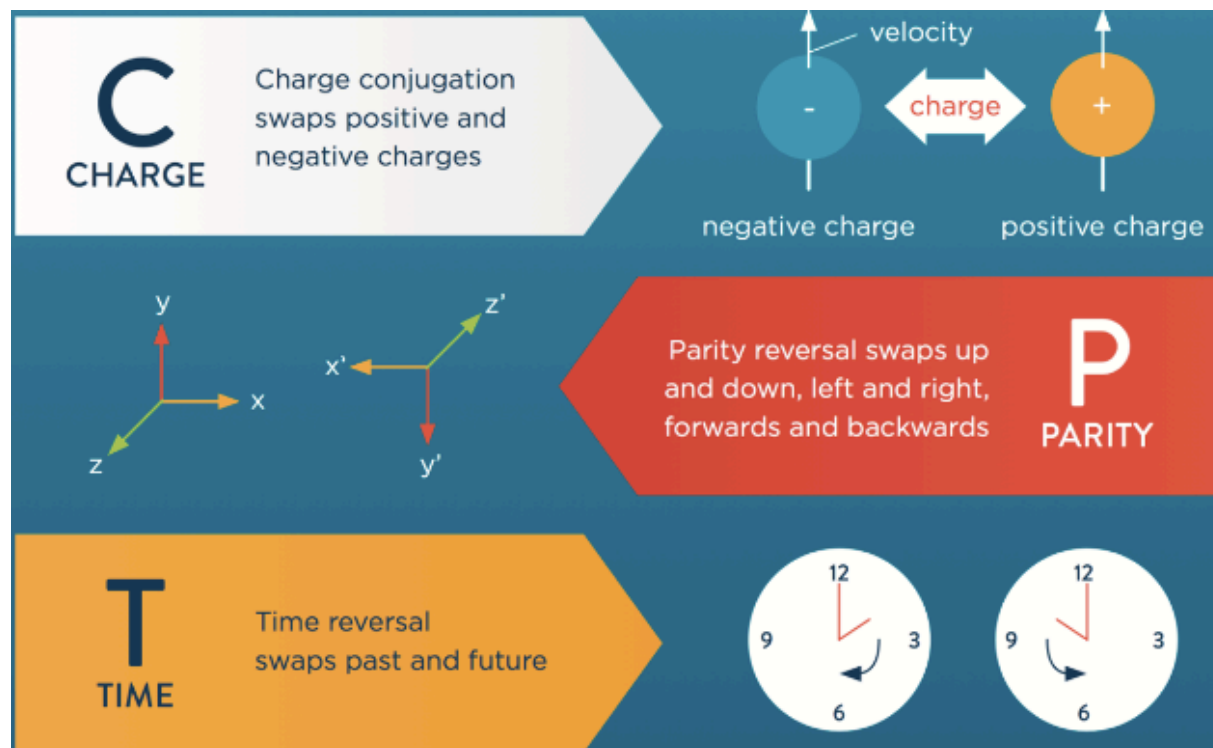
Standard Model



<https://sciencesprings.wordpress.com/2016/05/13/from-physicsworld-parallel-universe-search-focuses-on-neutrons/>

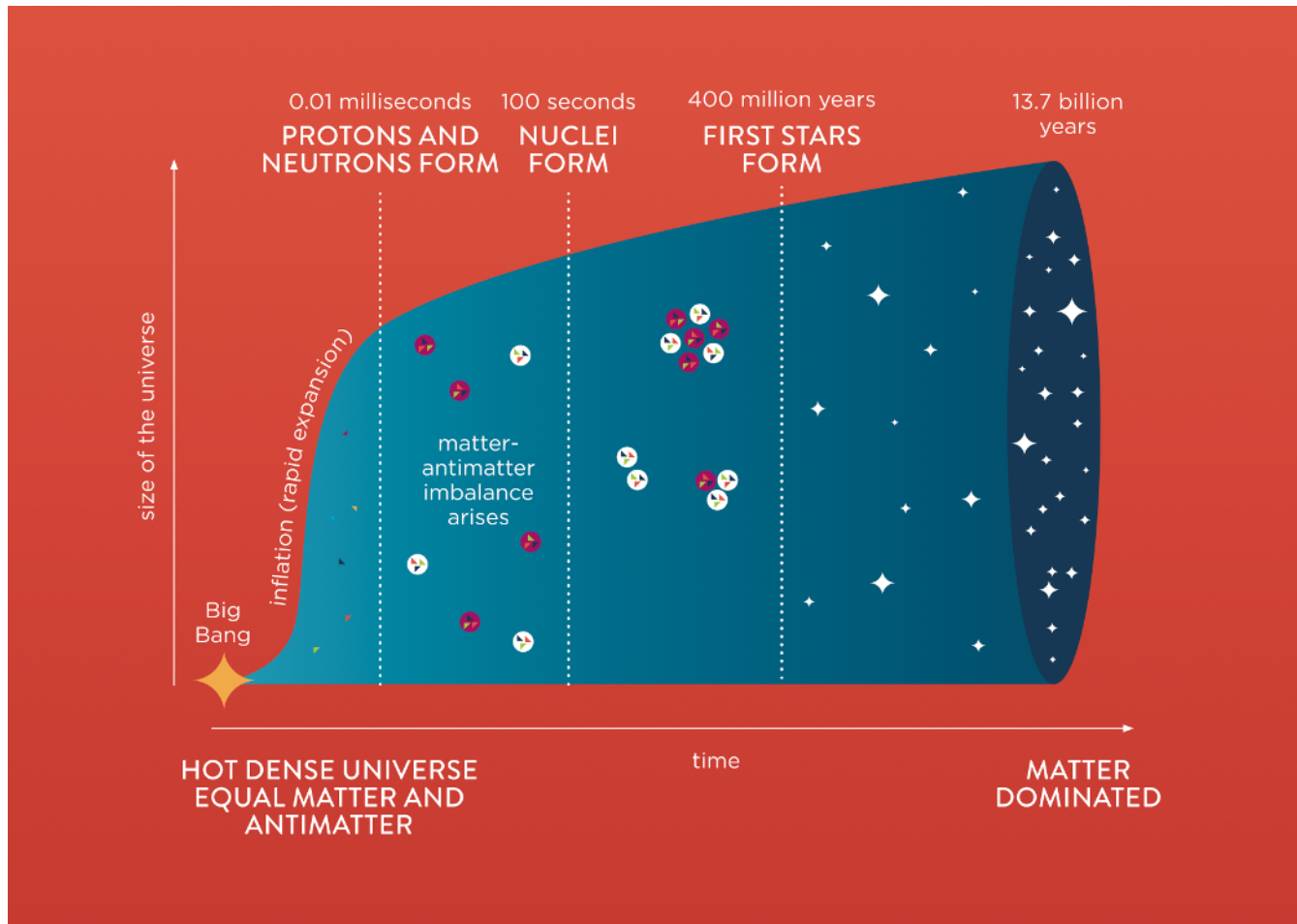
Symmetries

The physical laws described by the standard model have symmetries (only CPT (Charge-Parity-Time) is exact).



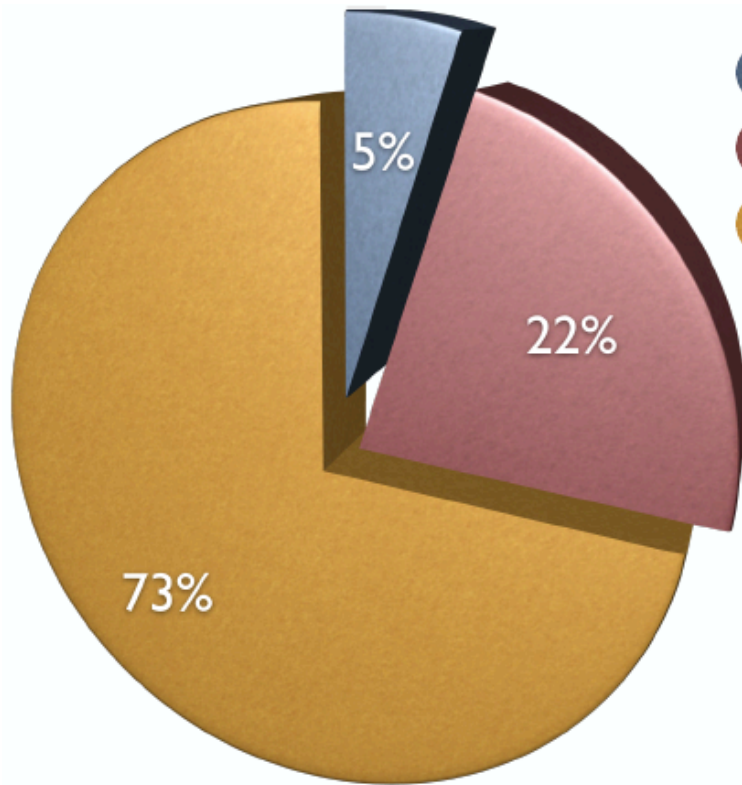
Antimatter experiments at CERN are testing these symmetries.

Why are we interested in antimatter?



Why is the universe composed almost entirely of normal matter ?

Why are we interested in antimatter?



- Known [normal matter]
 - Known unknowns [dark matter]
 - Unknown unknowns [dark energy]
- Antimatter ? : 0%



Antimatter at CERN



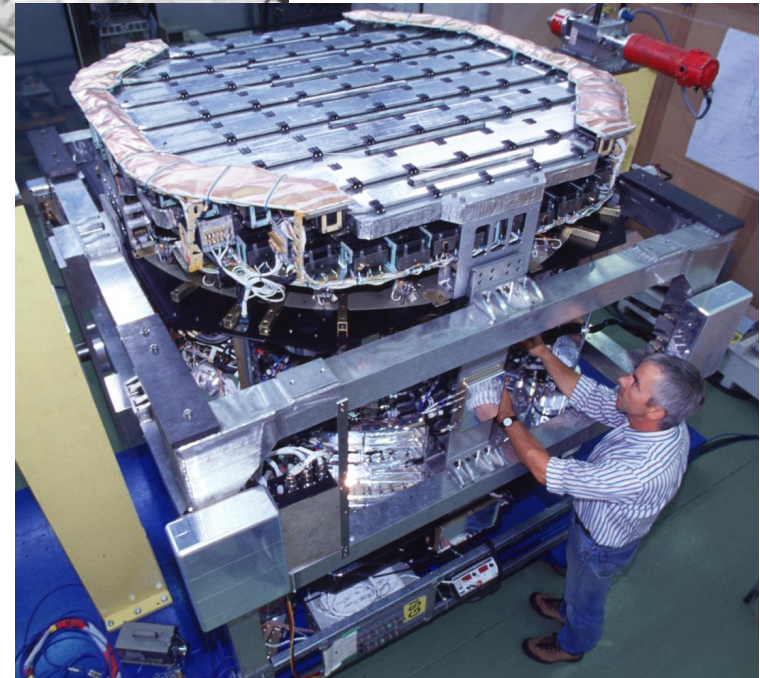
- High energy proton-proton collisions in LHCb. Looking at the relative amounts of matter and antimatter produced to test the symmetries of the standard model.
- Low energy experiments at the antiproton decelerator (AD) facility.



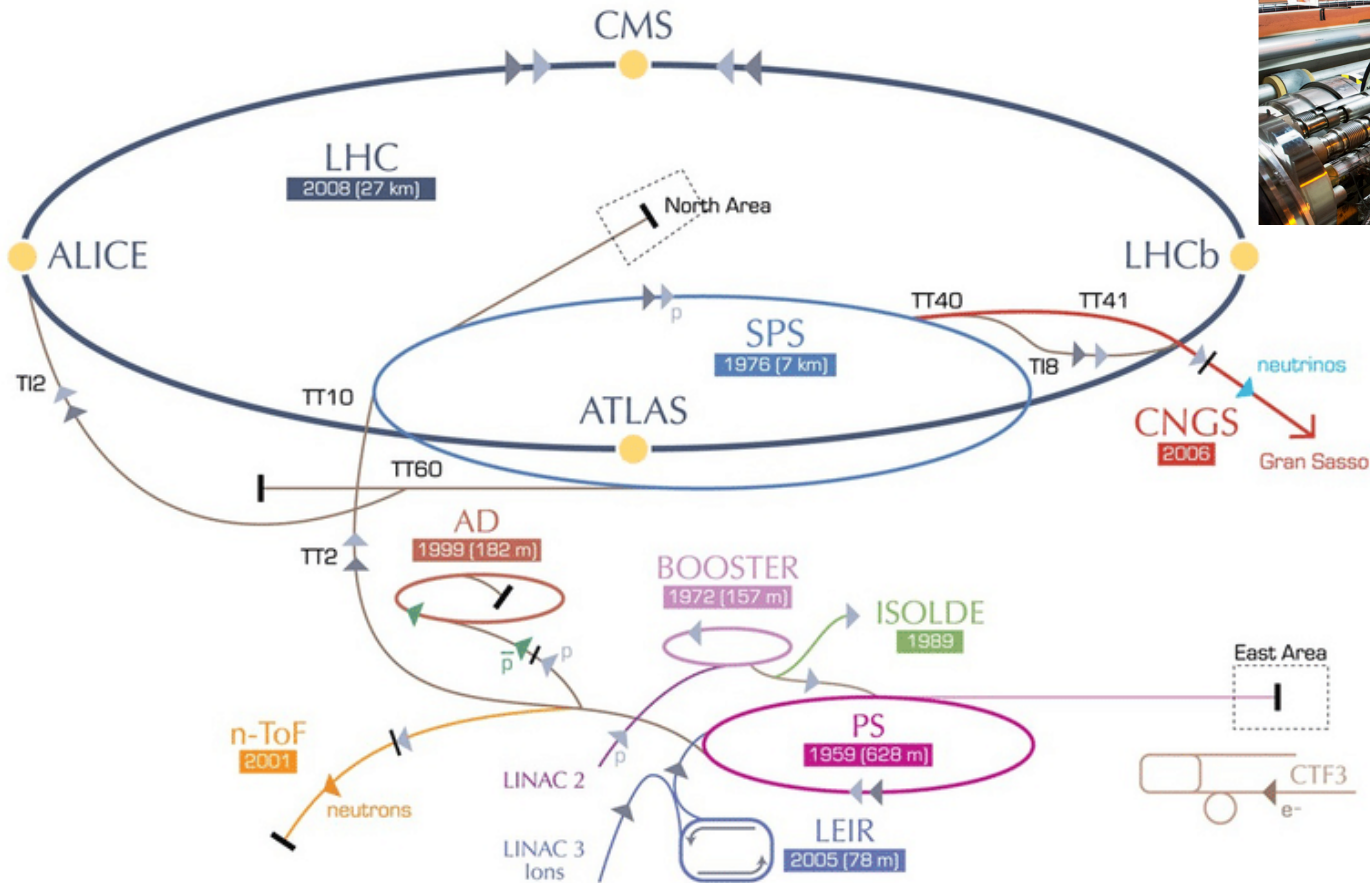
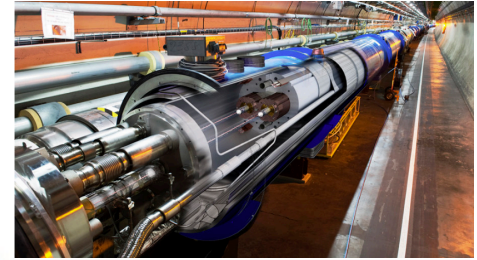
Antimatter in space



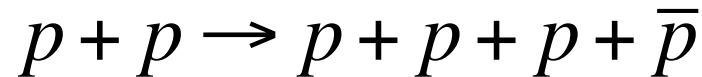
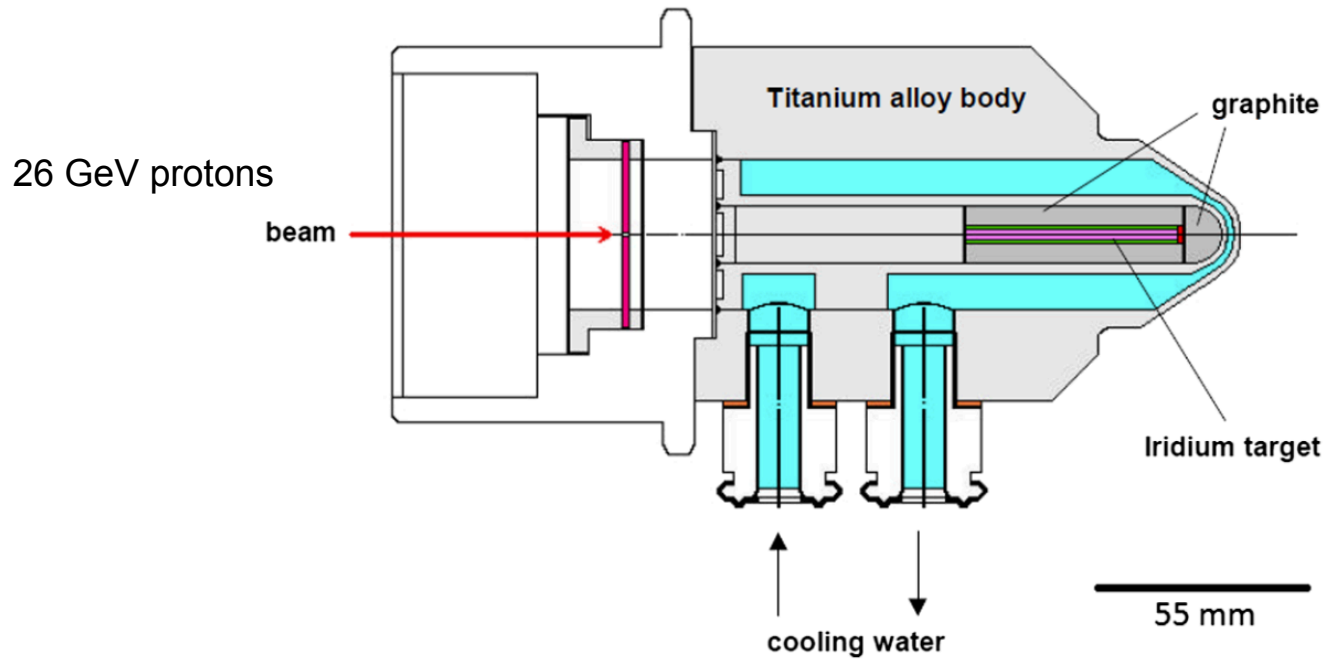
- The AMS detector has measured the flux of antiprotons to be 10,000 times smaller than the proton flux.
- Heavy nuclei could originate from regions of space with an abundance of antimatter e.g. anti-stars!



Antiprotons at CERN



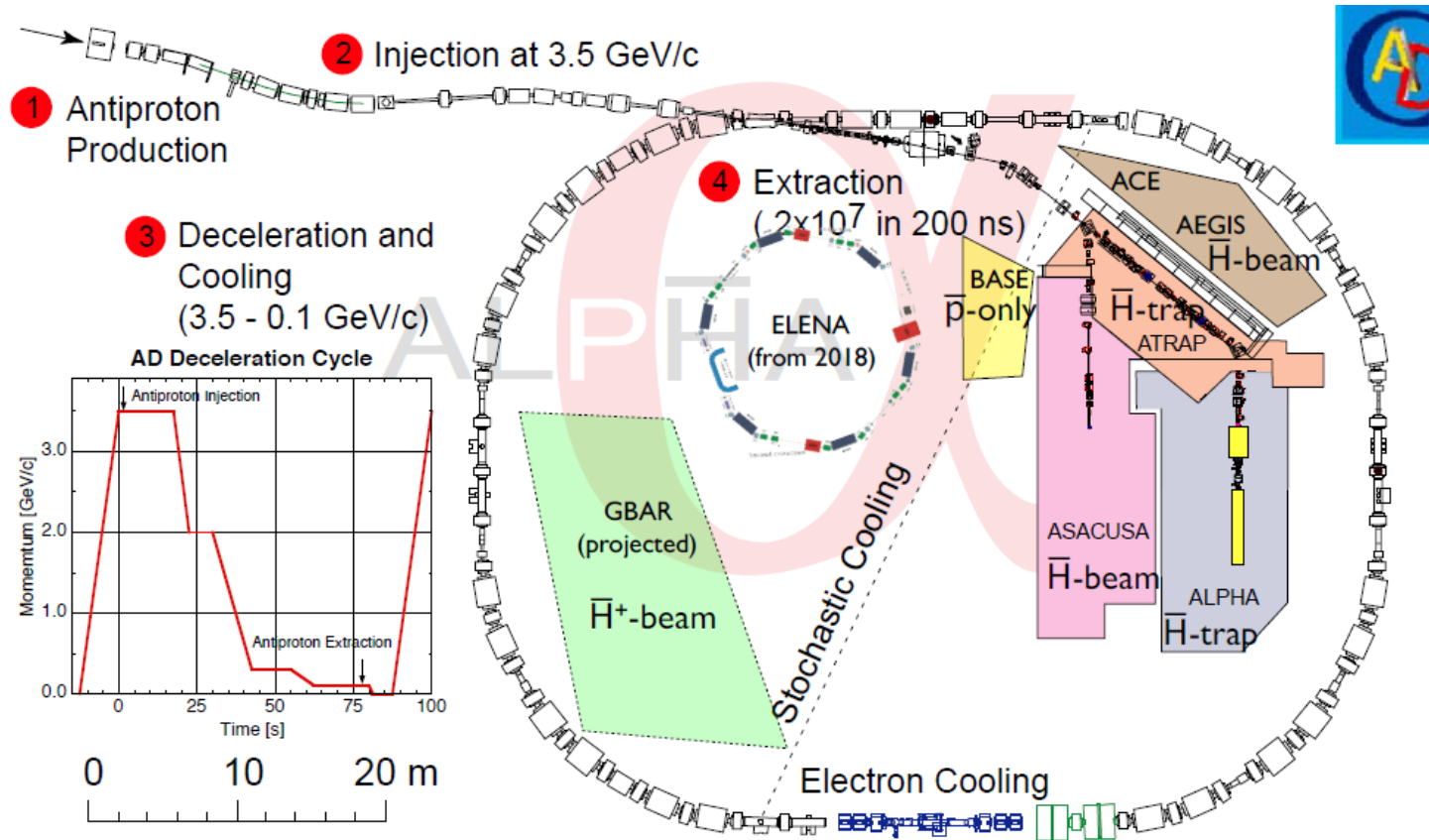
Making antiprotons



Around 5×10^7 antiprotons are produced every 120s at an energy of 3.5 GeV

The antiproton decelerator

The antiproton decelerator (AD) provides $\sim 2 \times 10^7$ antiprotons every 120 s at ~ 5 MeV.



S. Maury *et al.*, *Hyp. Int.* **109**, 43 (1997).

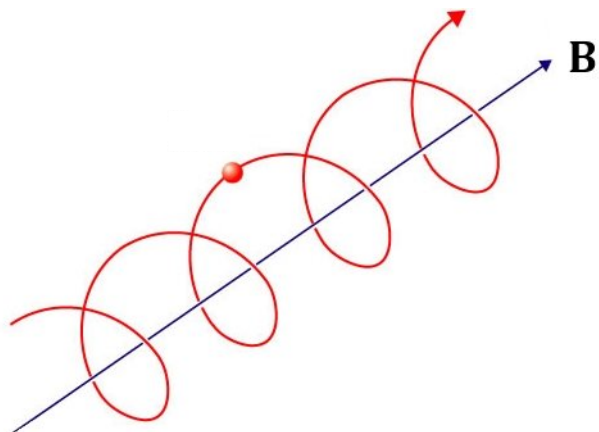
Trapping antiparticles

- Like charges repel.



$$F = q \vec{E}$$

- Moving charged particles experience a force in a magnetic field:

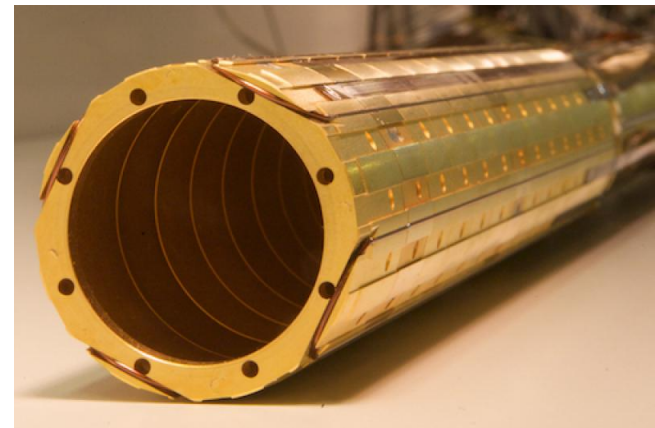
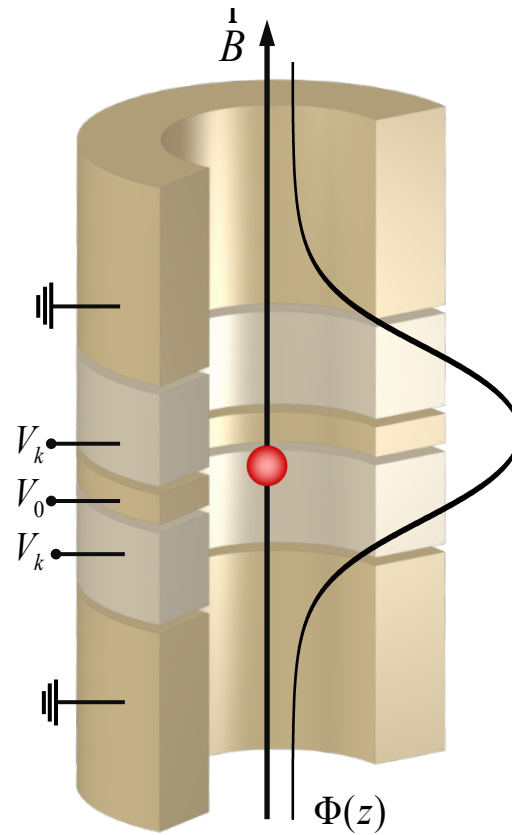
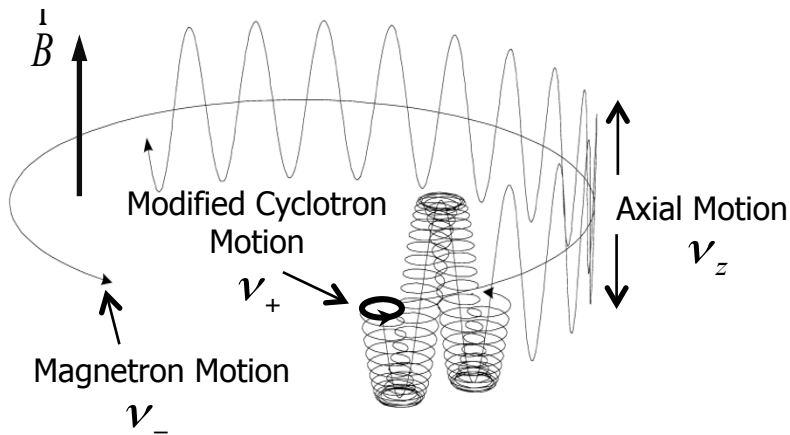


$$F = q \vec{V} \times \vec{B}$$

- Combine these two effects in a Penning trap.

Penning Trap

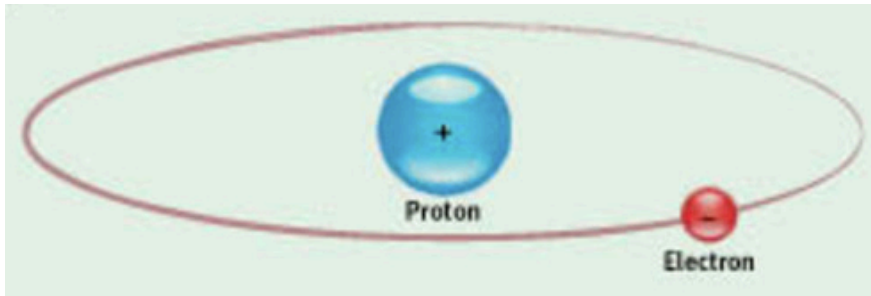
Traps are typically cooled down to $\sim 4\text{K}$ (-270 degrees C).



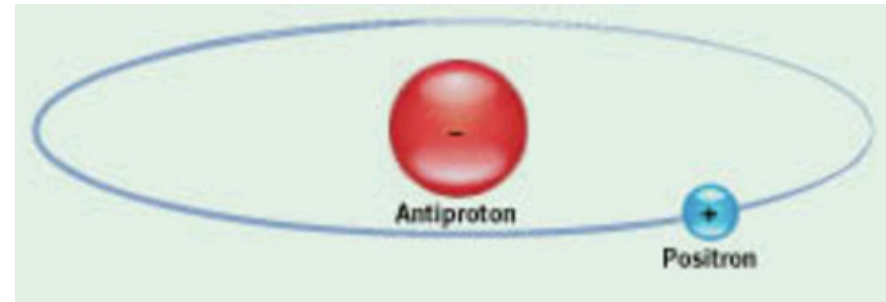
The ALPHA experiment

- ALPHA: **A**ntihydrogen **L**aser **P**hysics **A**pparatus.

hydrogen



antihydrogen



- Why antihydrogen?

- It's the simplest pure antimatter system.
- It is electrically neutral.
- Hydrogen is very well understood theoretically, and measurements of its properties have been performed to extraordinary levels of precision.

ALPHA-2 Apparatus

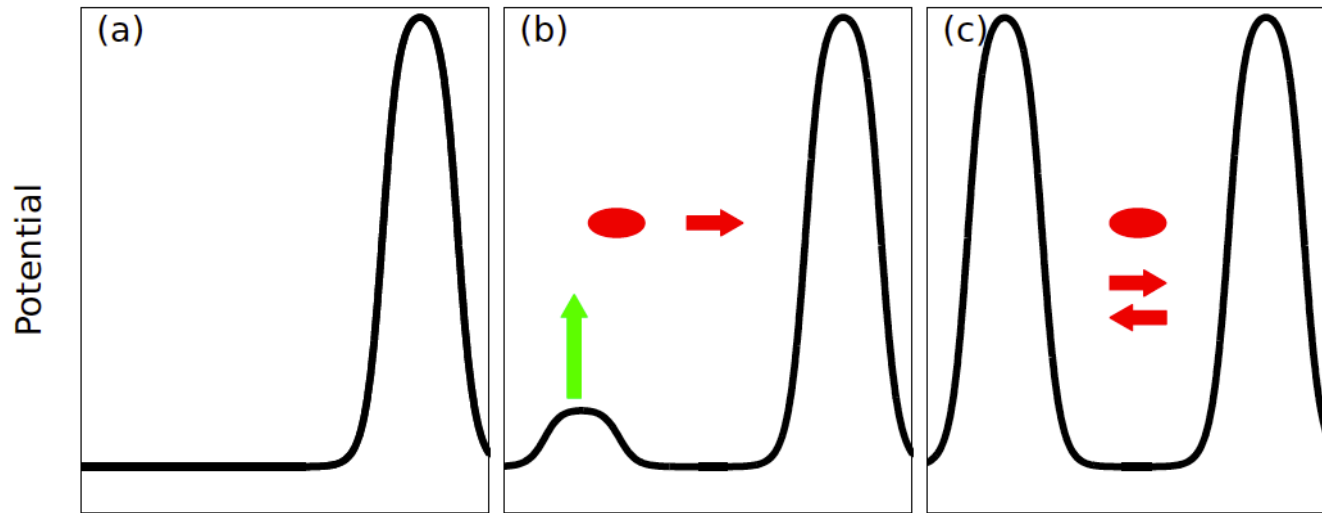


Antiproton “catching trap”

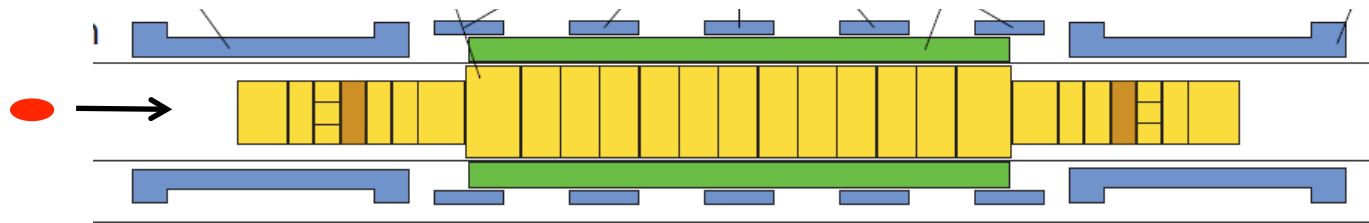
Antihydrogen “atom trap”

Catching antiparticles in the Penning traps

- Antiprotons from the AD are degraded from 5 MeV to 5 keV by a thin foil.



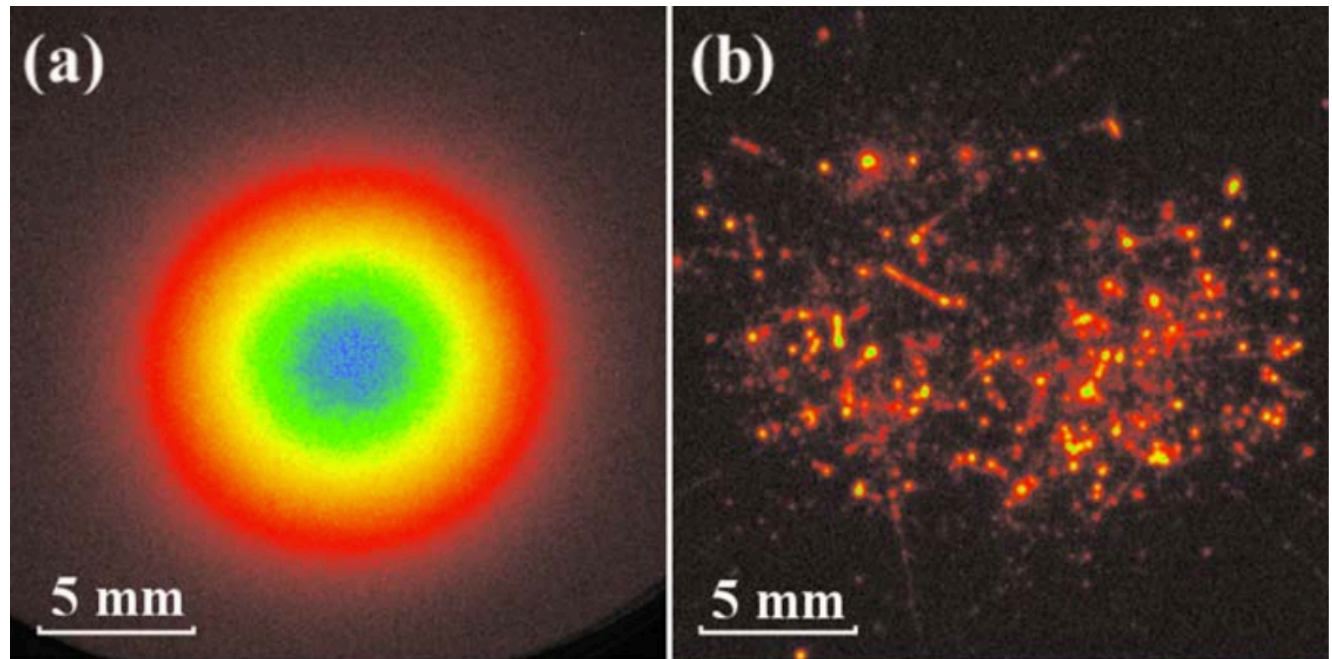
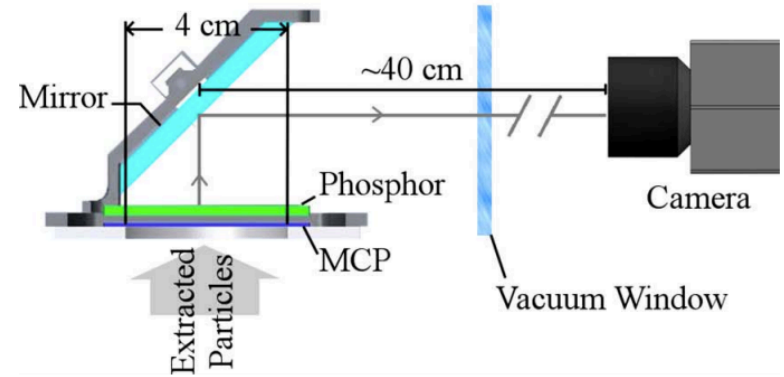
E. Butler PhD thesis.



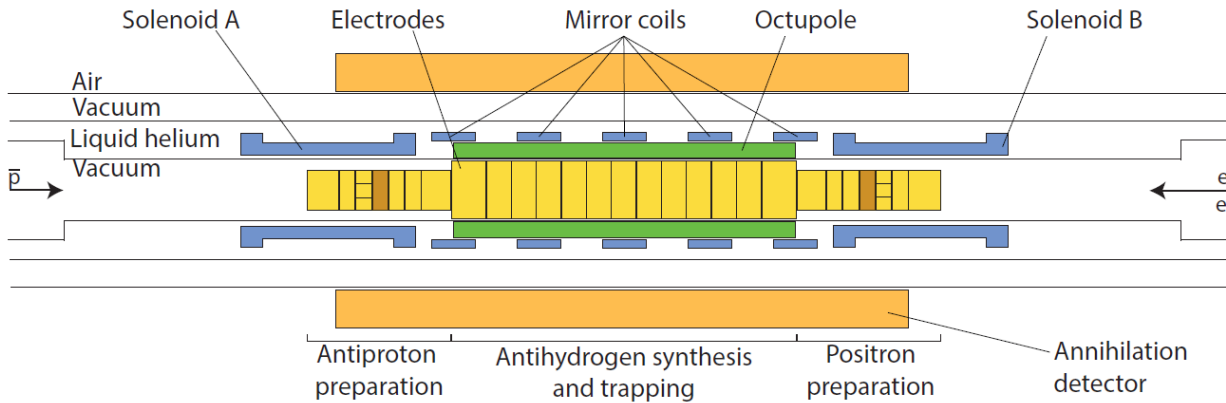
- Positrons from a radioactive sodium source are accumulated in a separate trap.

Imaging positrons and antiprotons

Light produced when particles hit a phosphor screen allows us to image the particles.

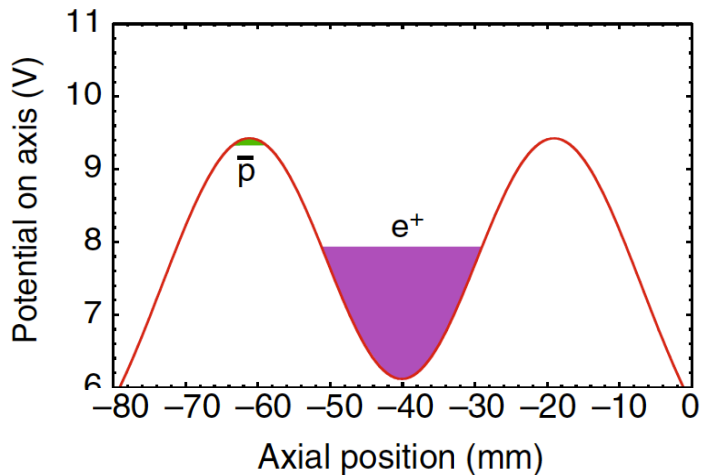


Antihydrogen synthesis

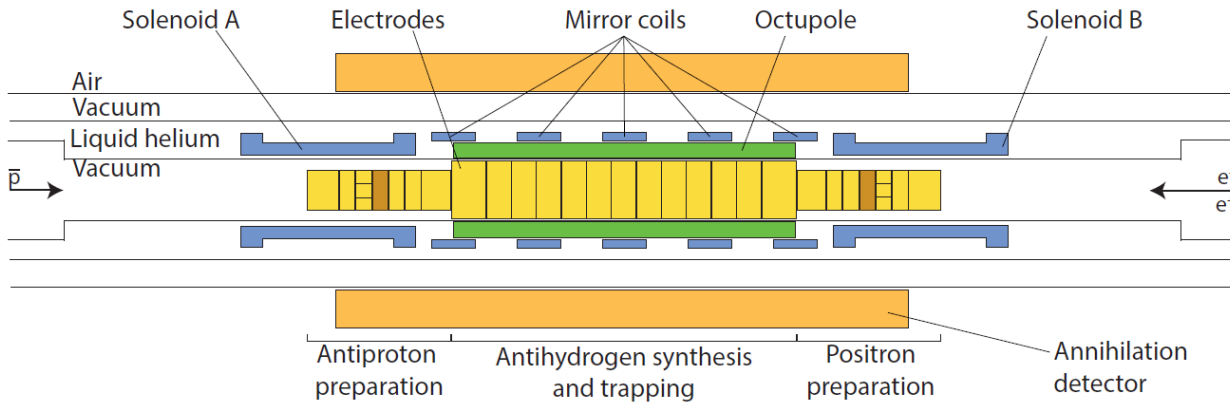


Trap positrons and antiprotons in adjacent potential wells of a Penning trap.

- Slowly merge the particles (in 1s) by lowering the barrier between them.
- We typically mix 3 million positrons (at $\sim 20\text{K}$) with 90,000 antiprotons (at $\sim 50\text{K}$) forming around **50,000 antihydrogen atoms**.

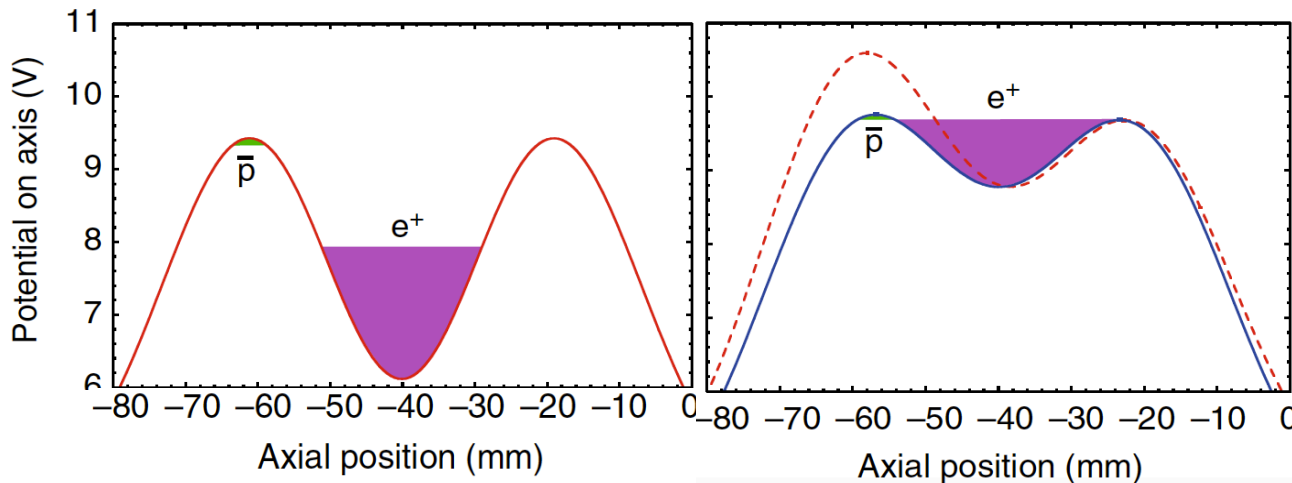


Antihydrogen synthesis

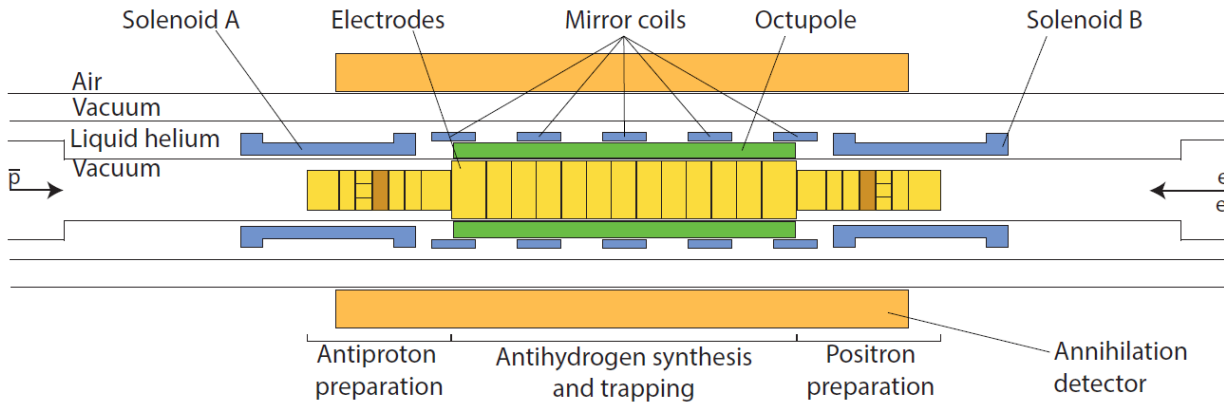


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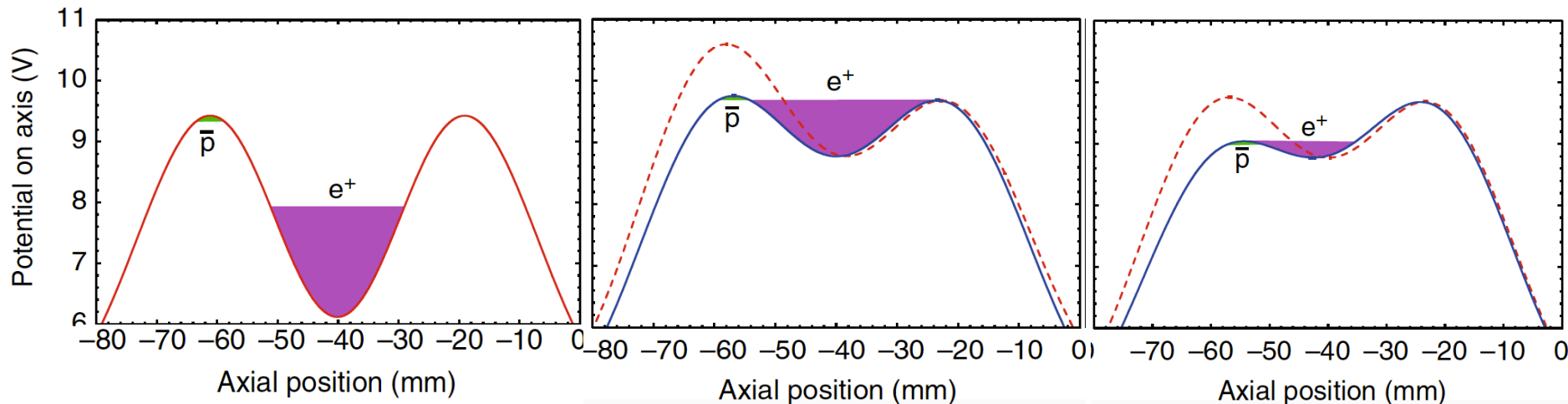


Antihydrogen synthesis



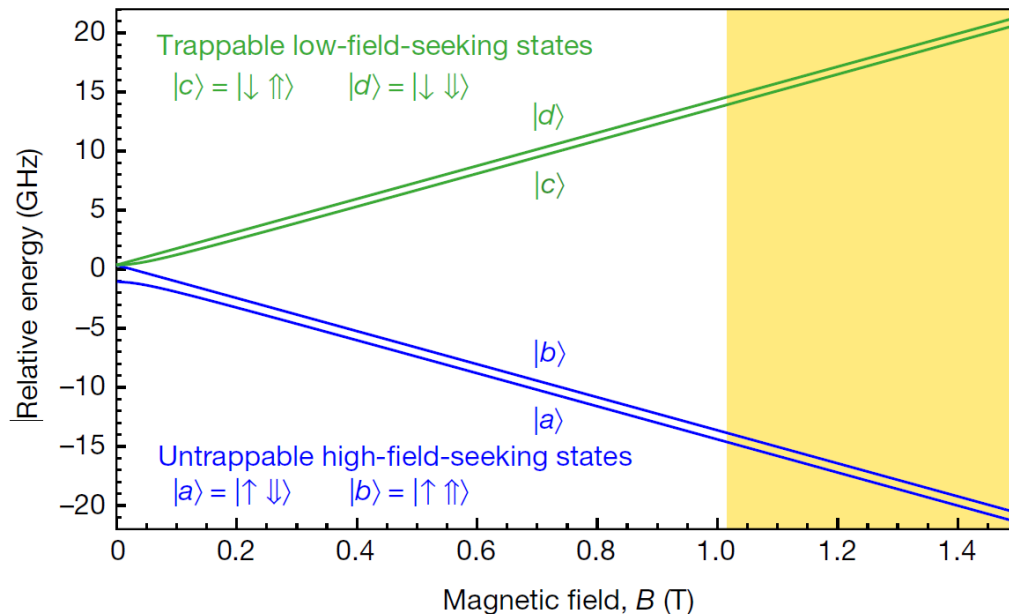
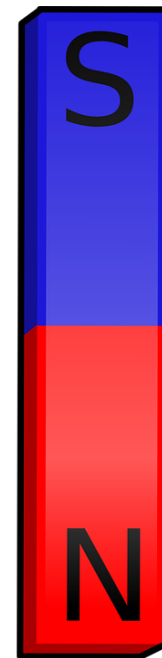
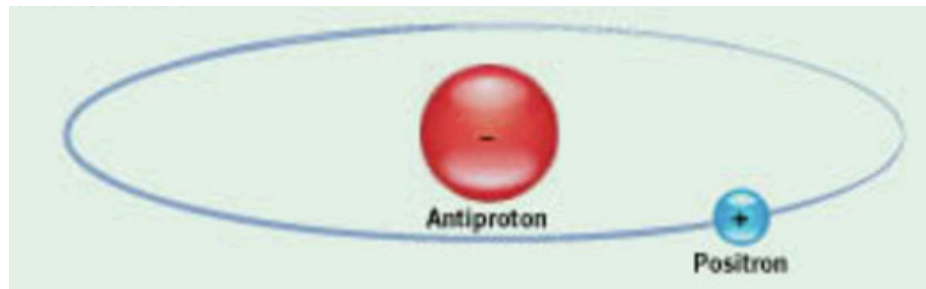
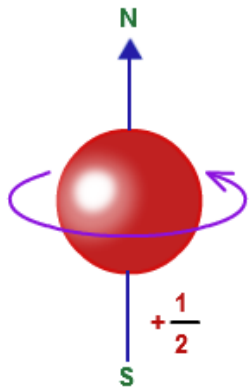
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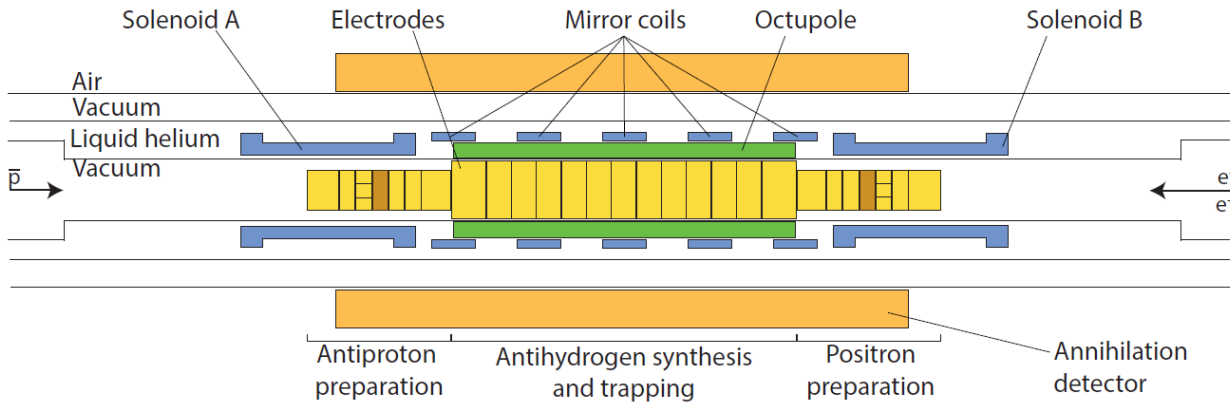


Antihydrogen trapping

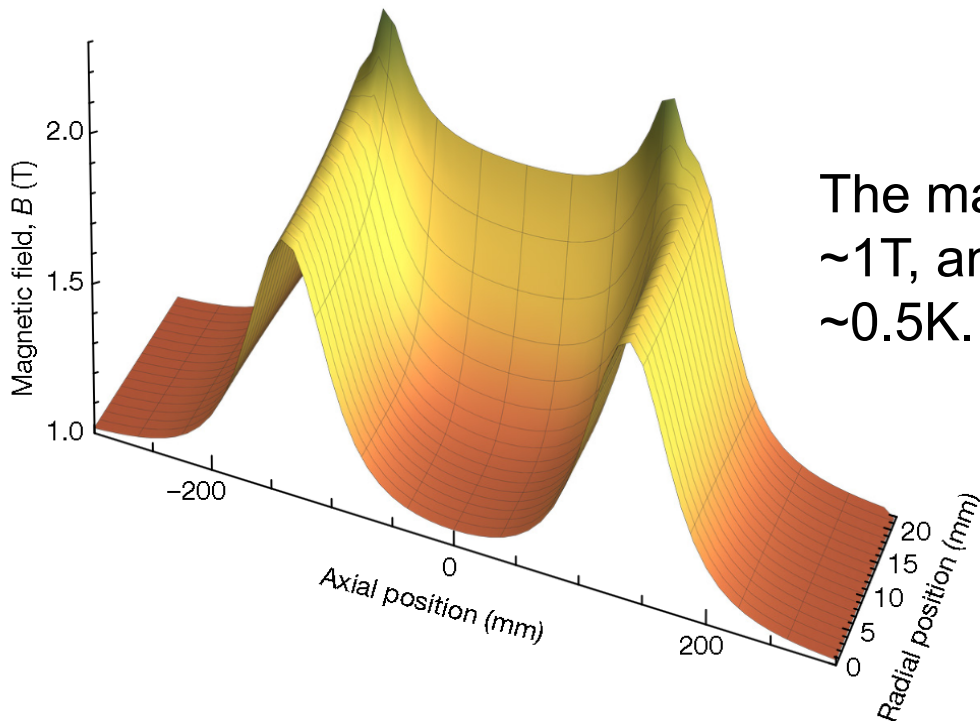
Antihydrogen is neutral – can't trap with a Penning trap. We make use of the fact that it is slightly magnetic (has a magnetic moment).



Antihydrogen trapping



Before mixing the particles, a magnetic minimum trap is energised consisting of an octupole and five mirror coils.

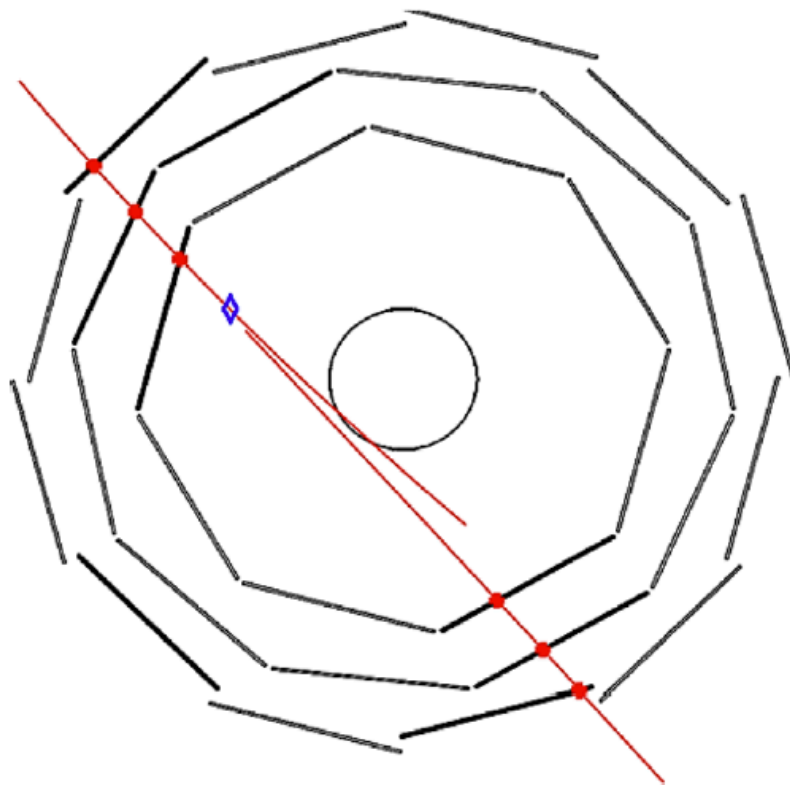


The magnetic field at the trap center is ~ 1 T, and the trap depth corresponds to ~ 0.5 K.

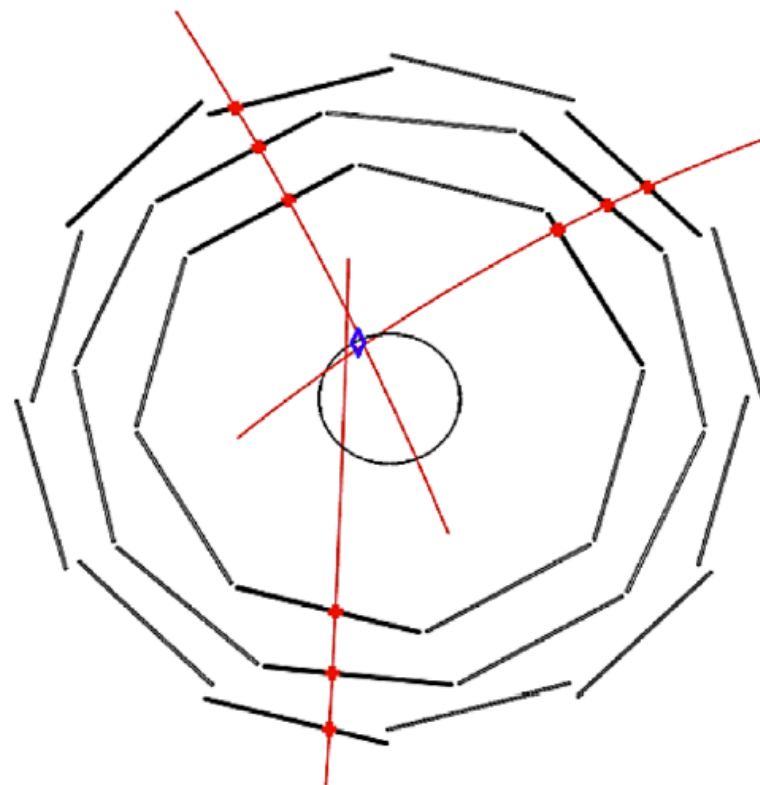
Antihydrogen detection

- We detect antihydrogen by ramping down the trap magnets to release the atoms – **we have to destroy the antihydrogen!**
- Image the annihilation products with a silicon vertex detector.
- Event topology allows us to distinguish antiproton annihilations from cosmic rays.

(a)

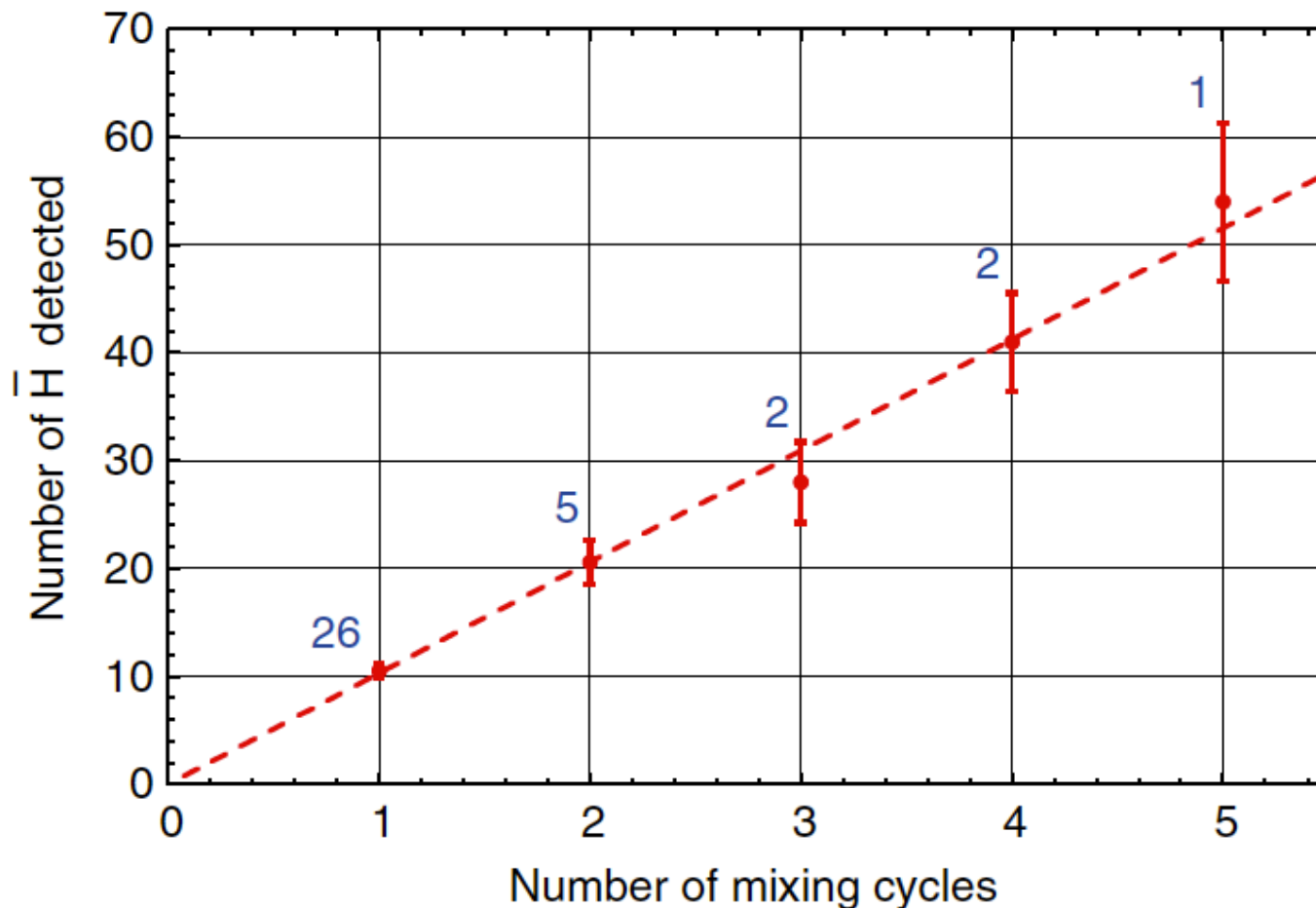


(b)



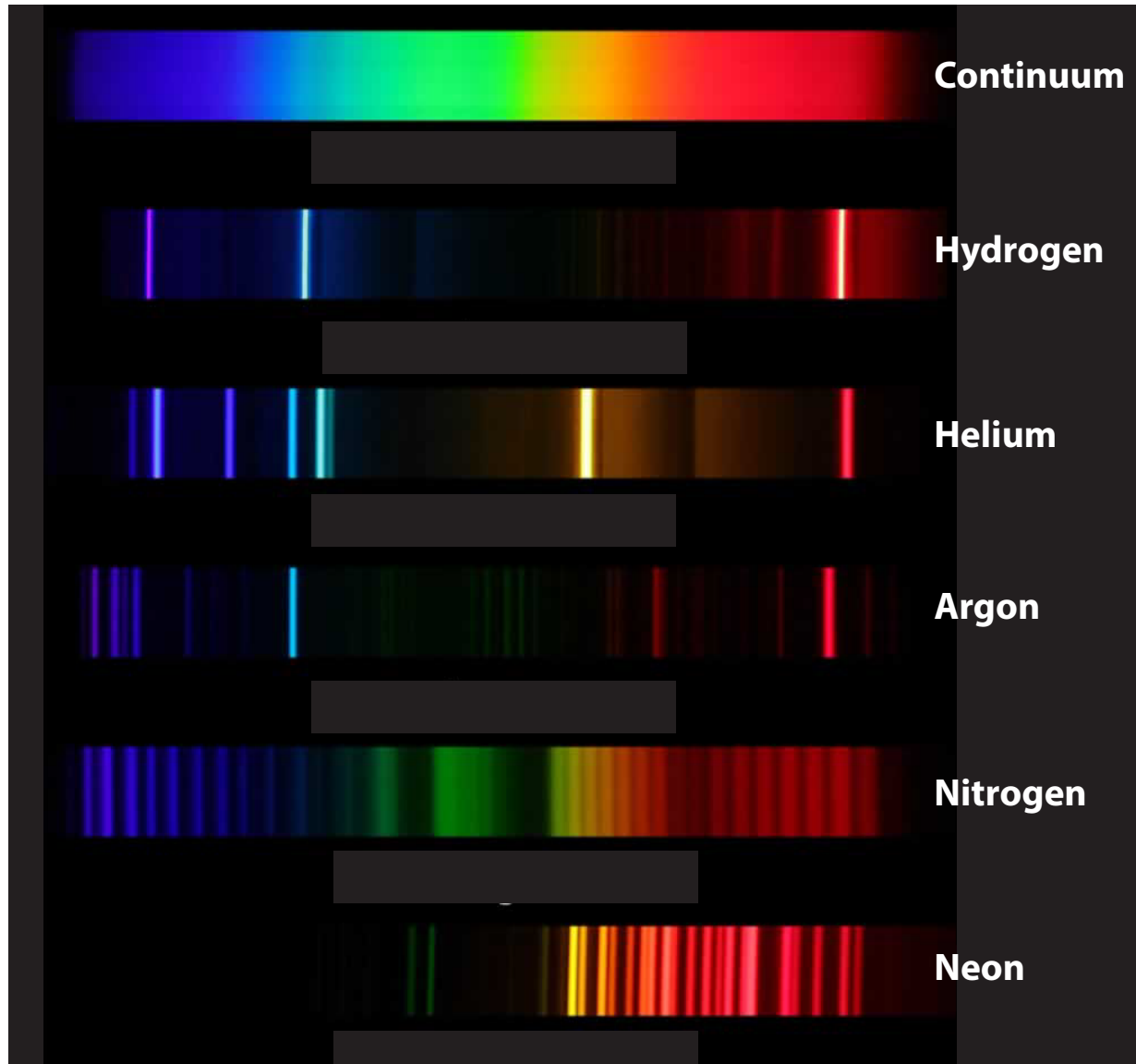
Antihydrogen accumulation

- Only trap **~20 atoms** per experimental cycle!
- However, we can accumulate trapped antihydrogen and have trapped **>1000 atoms** this way.



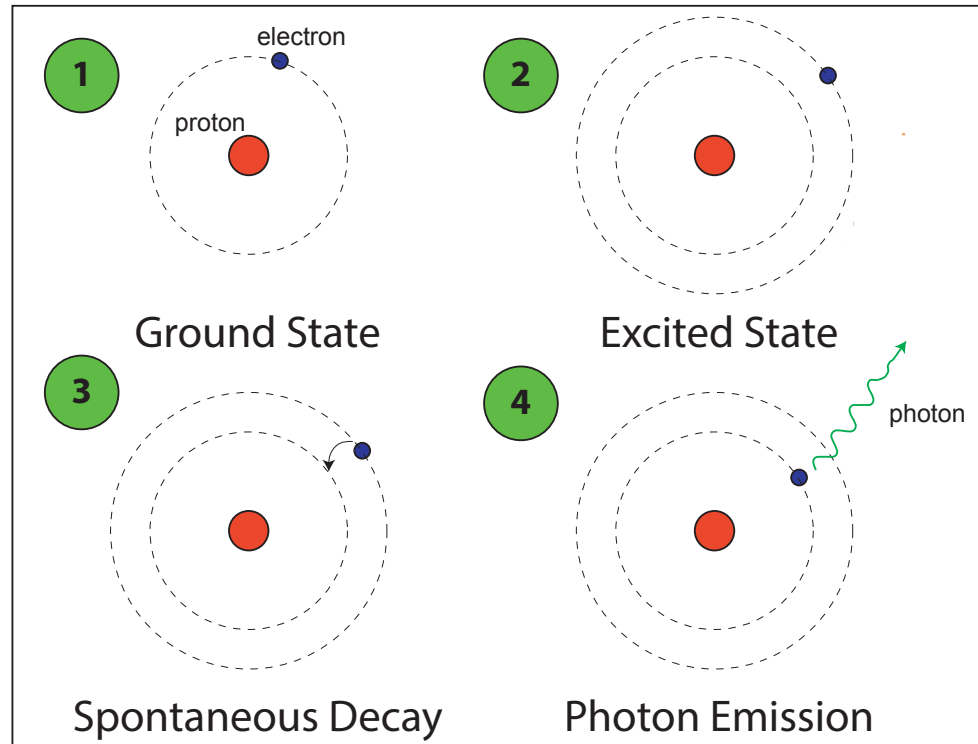
ALPHA Collaboration, Nat. Comms. **8**, 681 (2017).

Spectroscopy

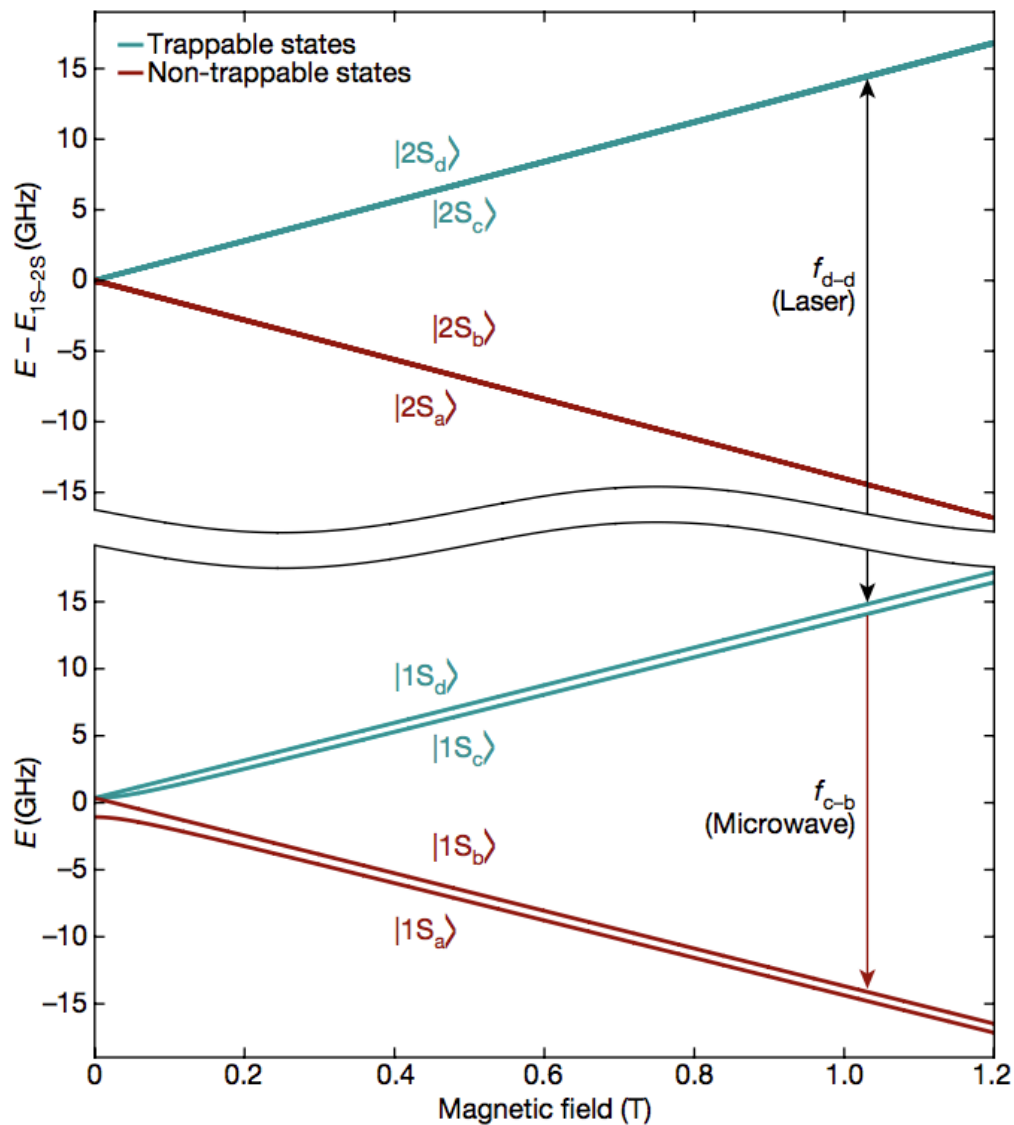


Spectroscopy

- The discrete wavelengths correspond to transitions of an atoms internal state.

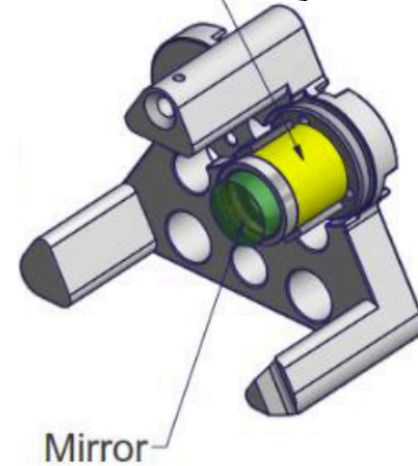
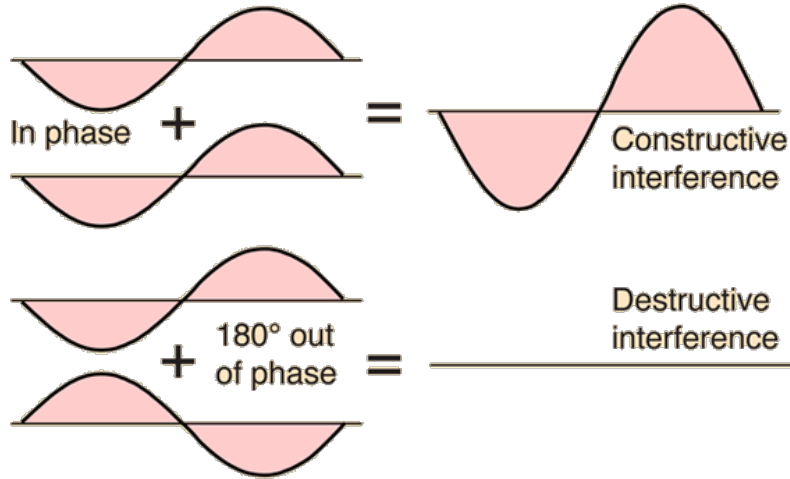
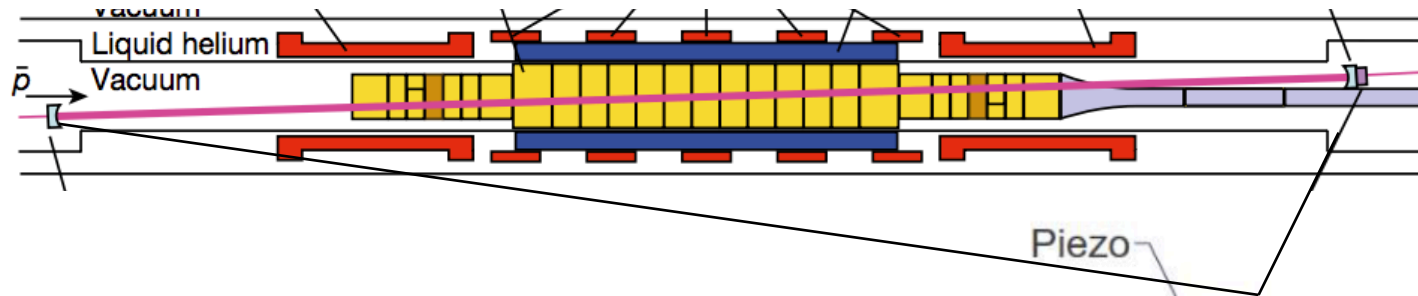


1S-2S transition



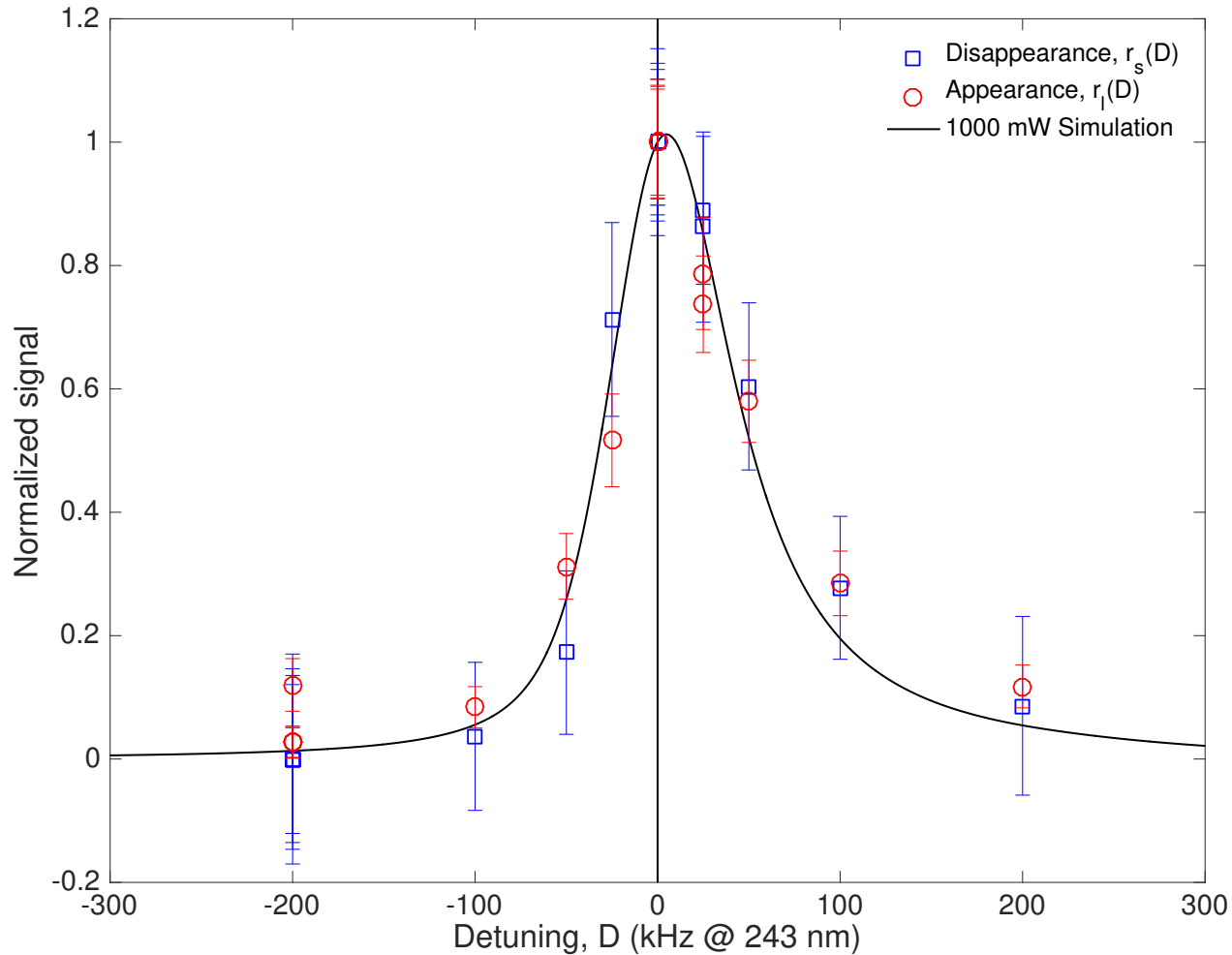
Measure the resonant frequency of the $1S_d$ - $2S_d$ transition, and compare with the expected value in hydrogen (in the same trap environment).

1S-2S experiment setup



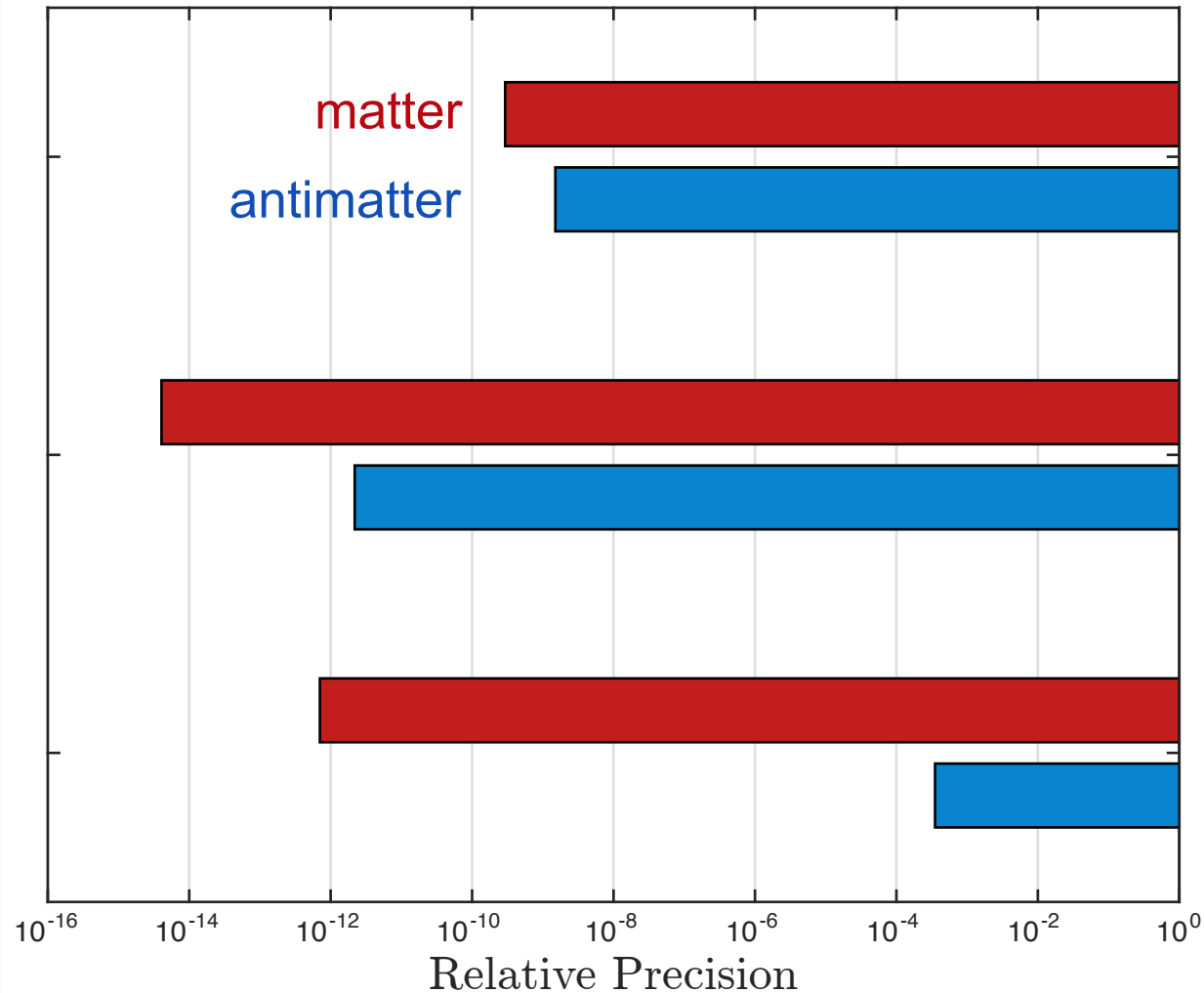
The distance between the mirrors which are 1 m apart is controlled to within 50 nm to achieve constructive interference of the laser beam.

1S-2S lineshape



Measured resonance frequency is consistent with the expected resonance frequency in hydrogen, and therefore consistent with CPT invariance, to a precision of 2×10^{-12} .

Matter-antimatter comparisons



G. Schneider *et al.*, *Science* **358**, 1081 (2017)

(anti)proton g-factor

C. Smorra *et al.*, *Nature* **550**, 371 (2017)

A. Matveev *et al.*, *Phys. Rev. Lett.* **110**, 230801 (2013)

(anti)hydrogen 1S-2S

M. Ahmadi *et al.*, *Nature* **557**, 71 (2018)

N. F. Ramsey, *Rev. Mod. Phys.* **62**, 541 (1990)

(anti)hydrogen GS HFS

M. Ahmadi *et al.*, *Nature* **548**, 66 (2017)

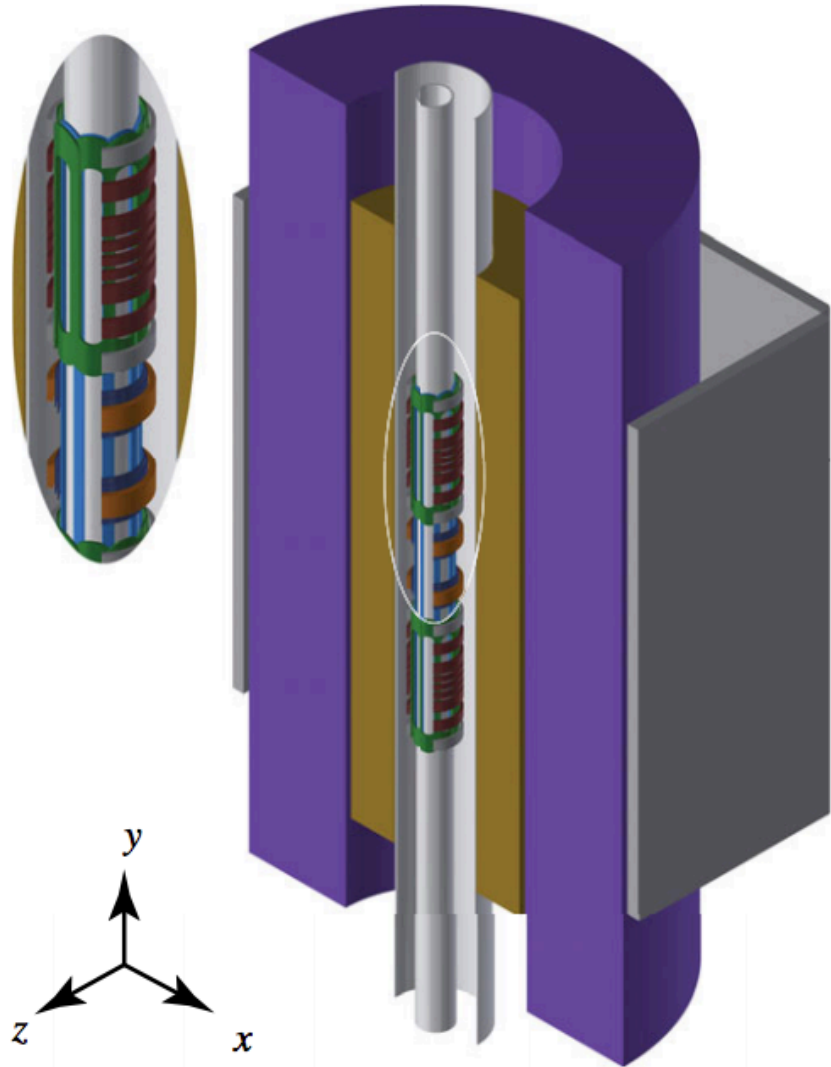
Measuring gravity with antimatter: ALPHA-g



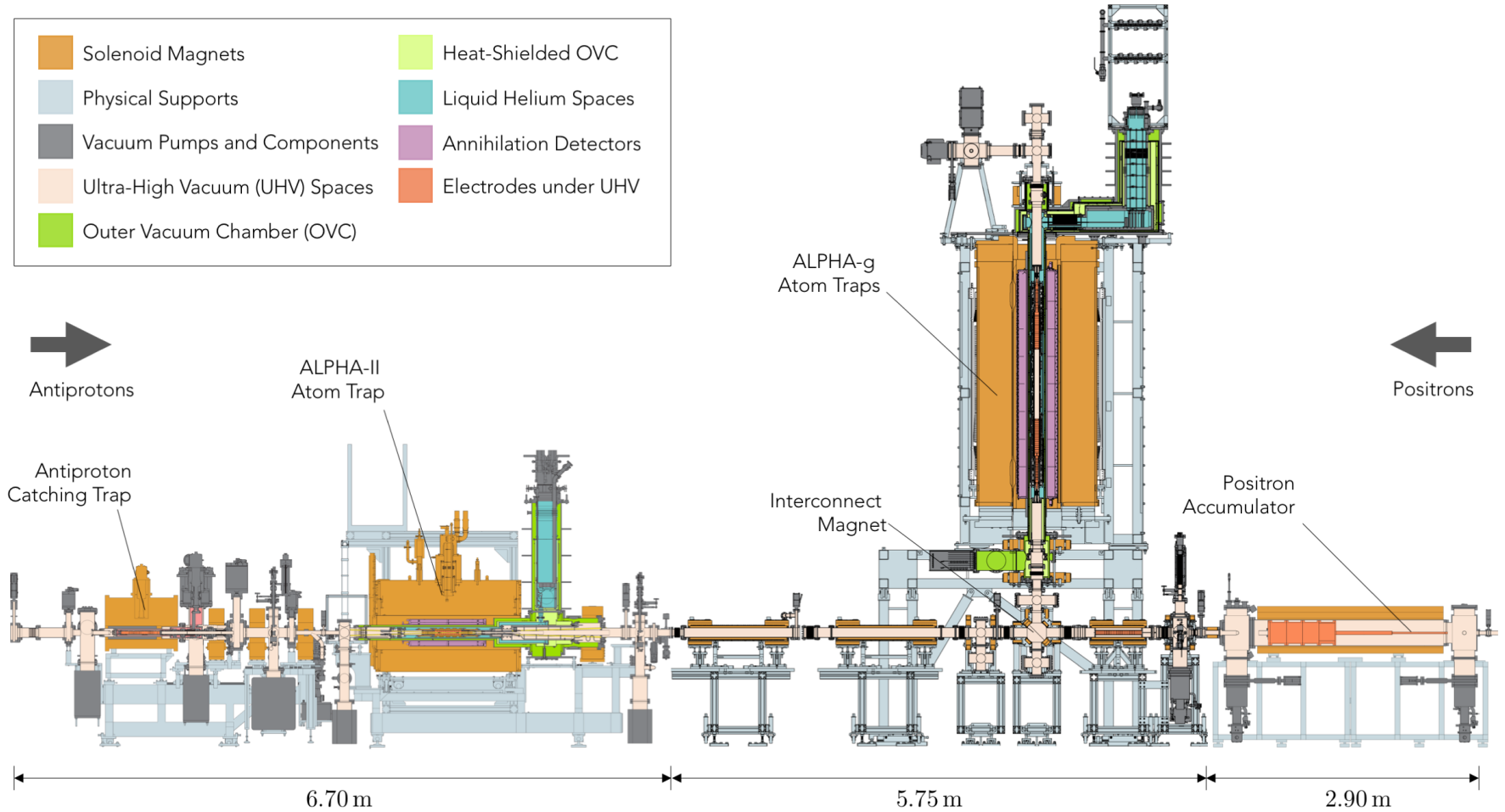
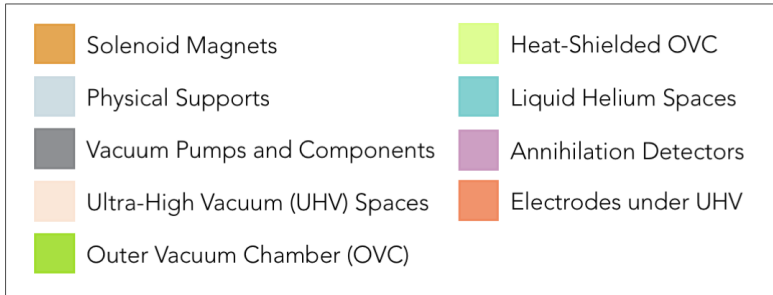
Which way does antimatter fall under the influence of gravity?

Never been measured directly!

Create antihydrogen in a vertical trap and release.



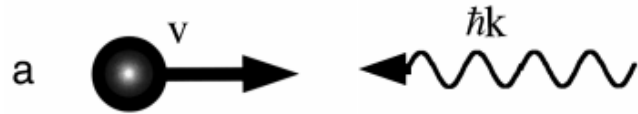
ALPHA-g



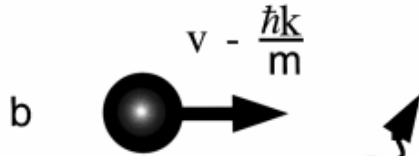
ALPHA-g



Laser cooling to make very cold antimatter



Photons carry momentum



This momentum is transferred to the atom which absorbs the photon



Make use of the Doppler effect to have a net friction force on the atoms

In 2018 we demonstrated this for the first time with antihydrogen – will ultimately lead to more precise measurements of gravity and with spectroscopy.

Current state of antimatter research



- Violations of CP symmetry have been measured, but these are not large enough to explain the matter-antimatter asymmetry of the universe.
- Electromagnetic properties of antimatter are now starting to be measured at levels of precision comparable to those measured in hydrogen.
- Whether antimatter falls up or down is not yet known!
- We still do not know why there is no significant amount of antimatter observable in the universe!

Thank you for listening!

