

# *Scalar dark matter interacting through an extra $U(1)$ gauge interaction*

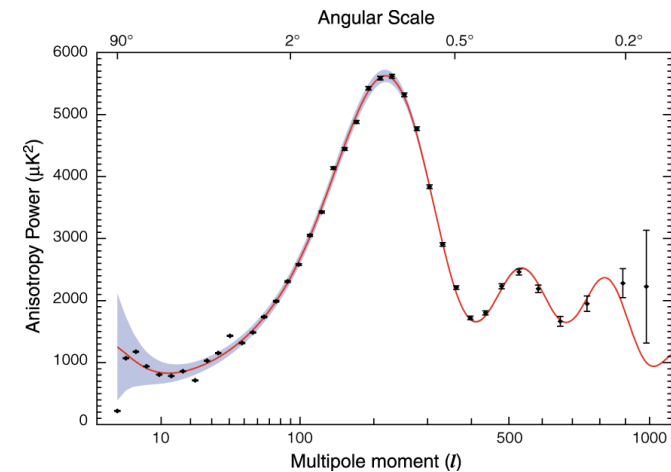
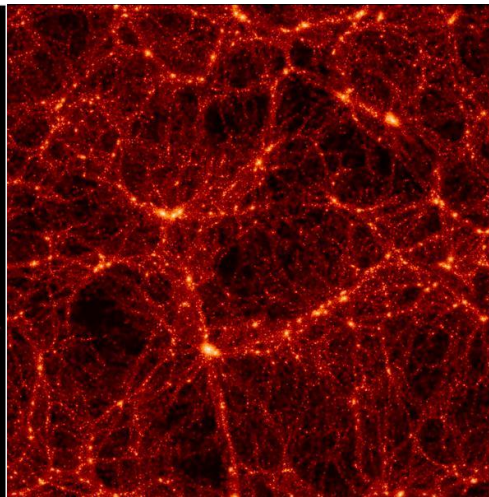
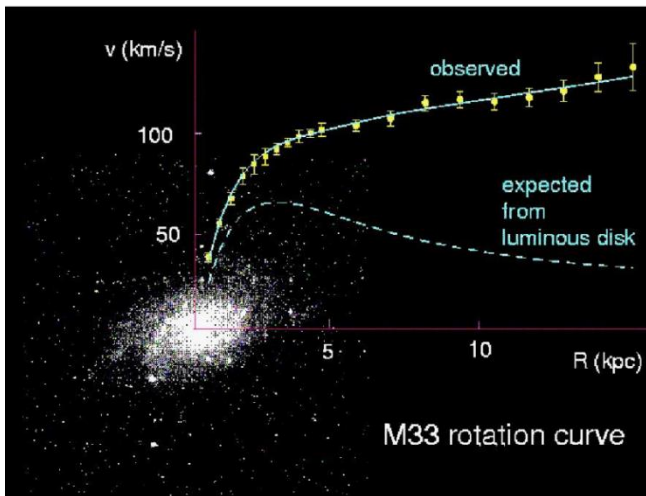
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(Hokkaido Univ.)

With: Nobuchika Okada (U. of Alabama)

Refs : 1908.09277

# § Introduction

Dark matters are everywhere!



- Its identification (Mass, ...) and properties are unknown

§ Low mass scale is interesting

# § § $g-2$ in muon

- Anomalous magnetic moment of muon [Brown et al (2001), Bennet et al (2006)]

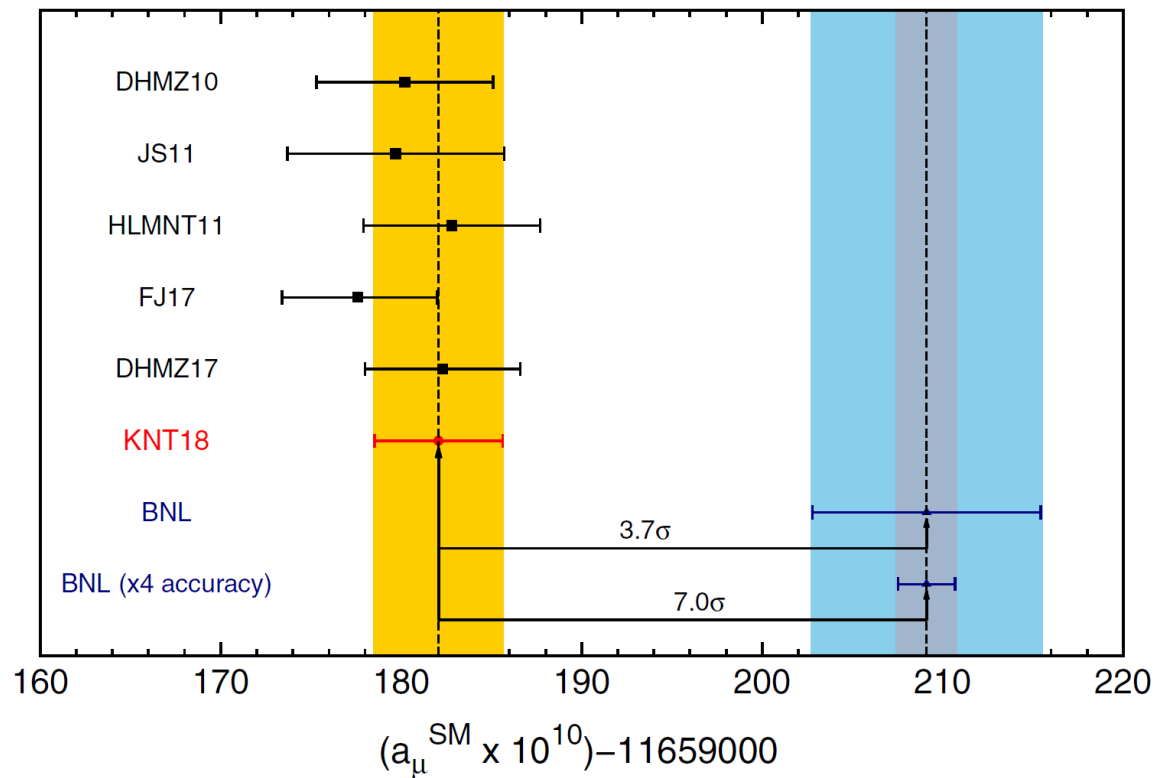
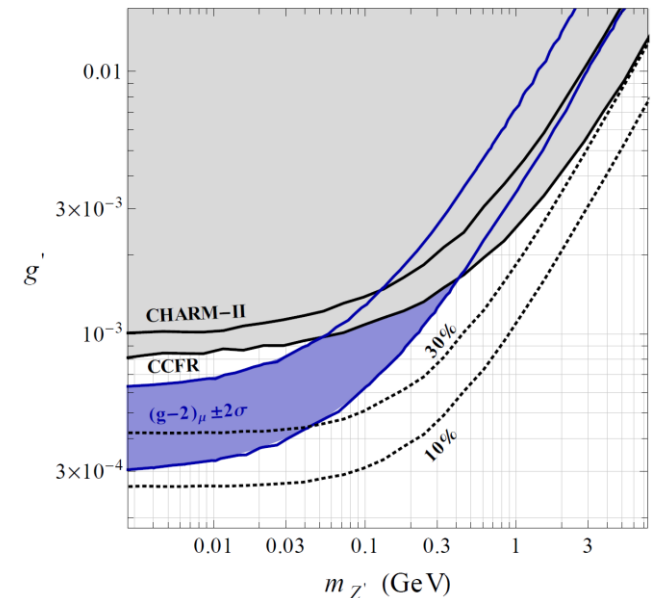
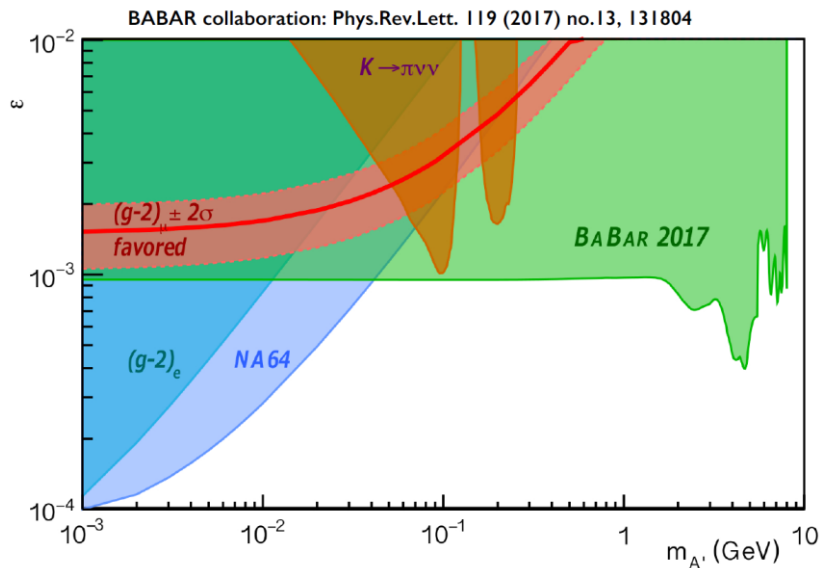


Fig from [Keshavarzi et al (2019)]

# § § $g-2$ in muon

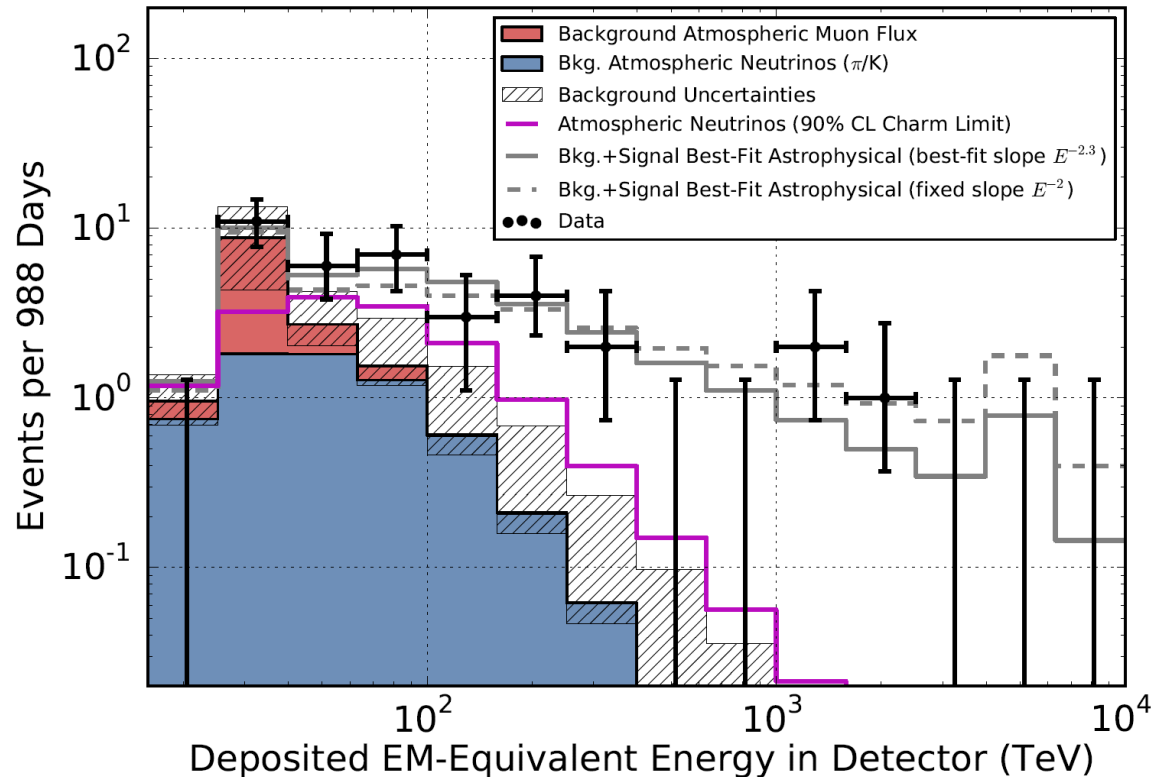
- Anomalous magnetic moment of muon [Brown et al (2001), Bennet et al (2006)]
- Dark photon interpretation



- $U(1)_{L\mu-L\tau}$  interpretation [Ma et al (2002), ...] is still viable [Altmannshofer et al (2014)]

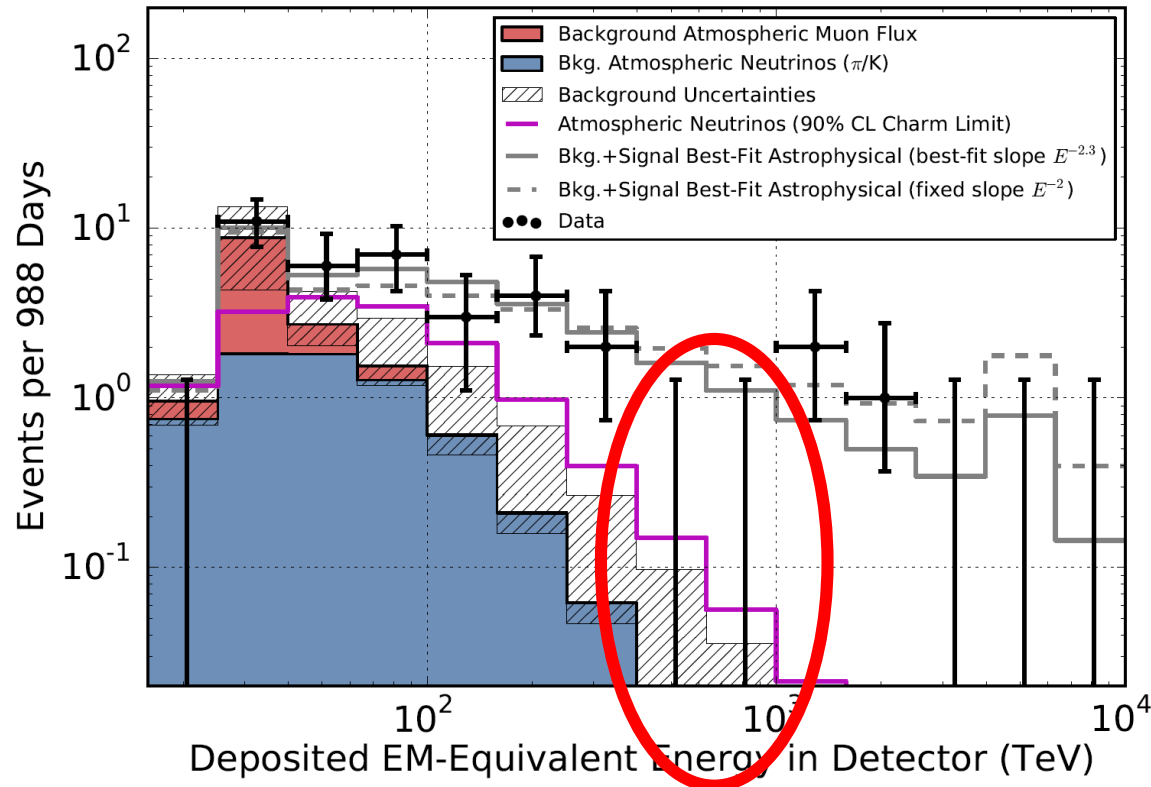
# § § A gap in IceCube?

- High energy neutrino spectrum measured by IceCube [IceCube (2014)]



# § § A gap in IceCube?

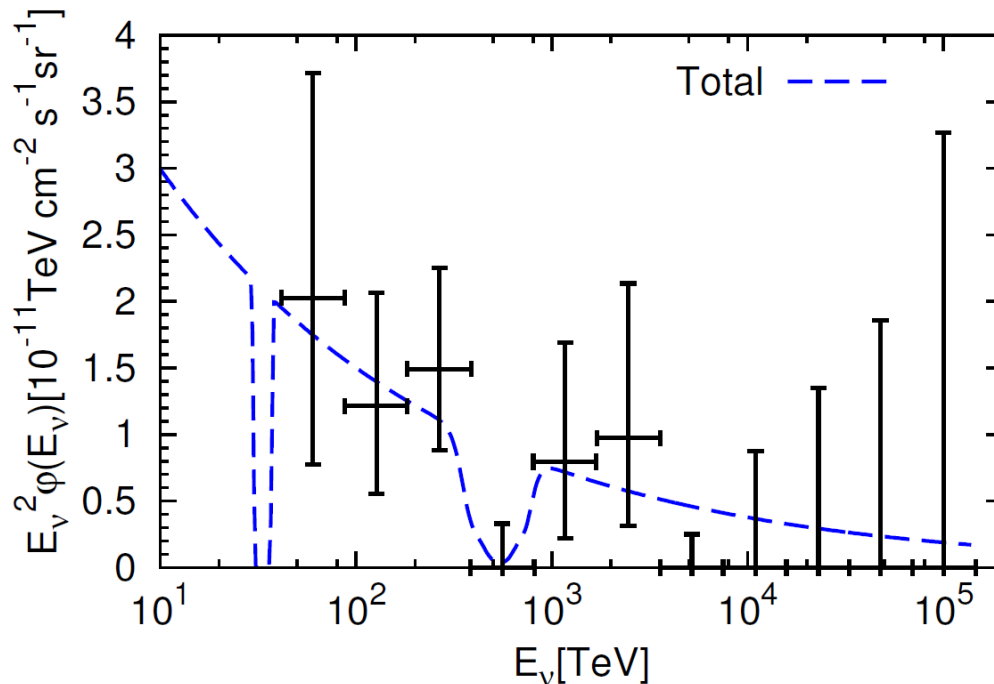
- High energy neutrino spectrum measured by IceCube [IceCube (2014)]



Gap!?

# § § A gap in IceCube?

- High energy neutrino spectrum measured by IceCube [IceCube (2014)]
- New physics interpretation
  - $Z'$  interpretation in  $U(1)_{L\mu-L\tau}$  model [Araki et al (2015)]



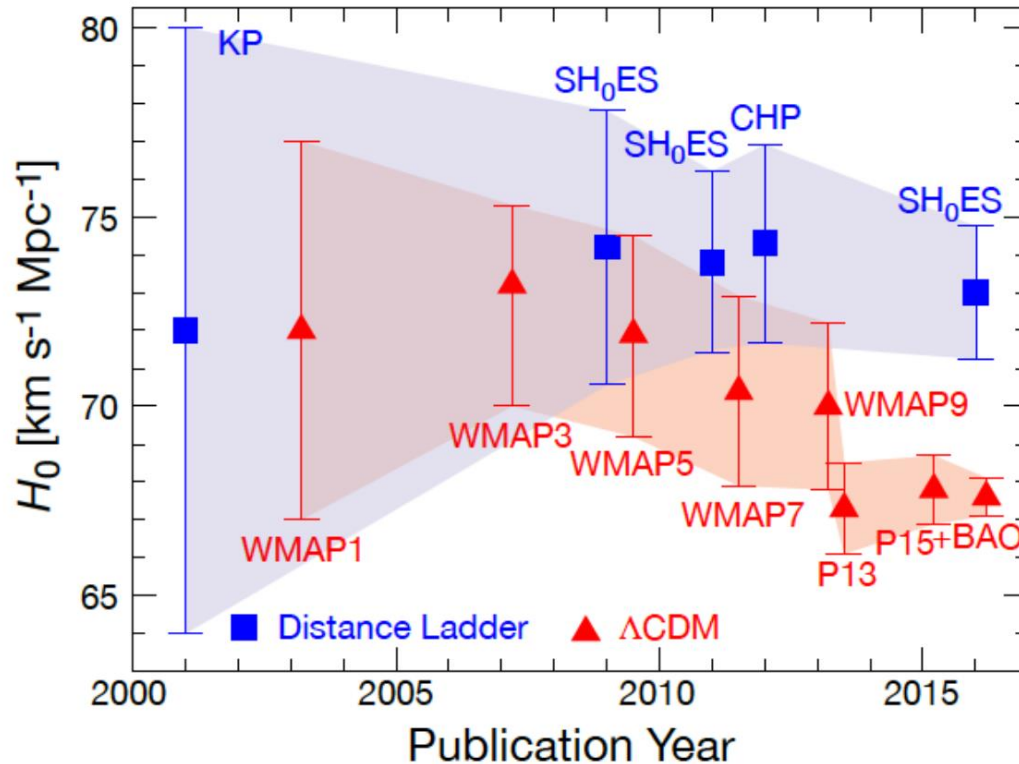
$$m_{Z'} = 1.9 \text{ MeV}$$
$$g' = 0.0005$$

$$\nu_{\text{CR}} + \nu_{\text{C}\nu\text{B}} \rightarrow Z'$$



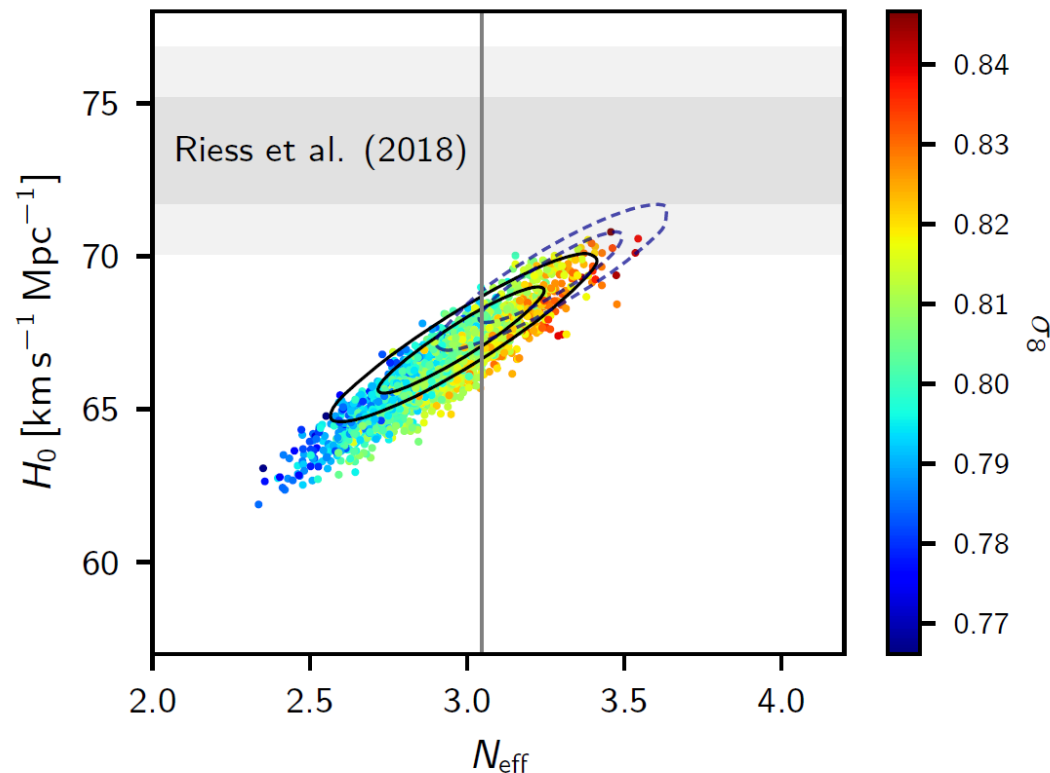
# § § Hubble tension

- Hubble parameter



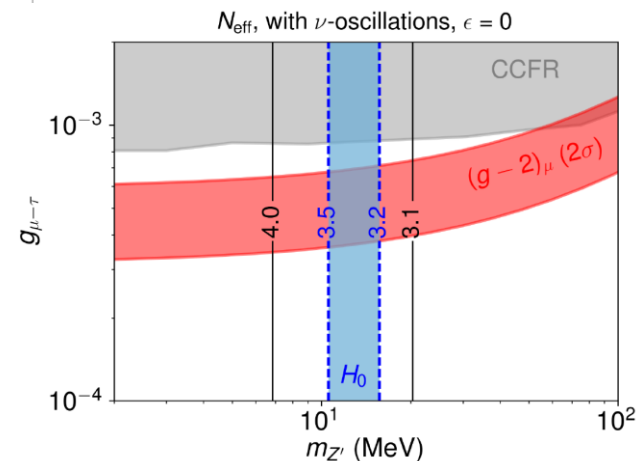
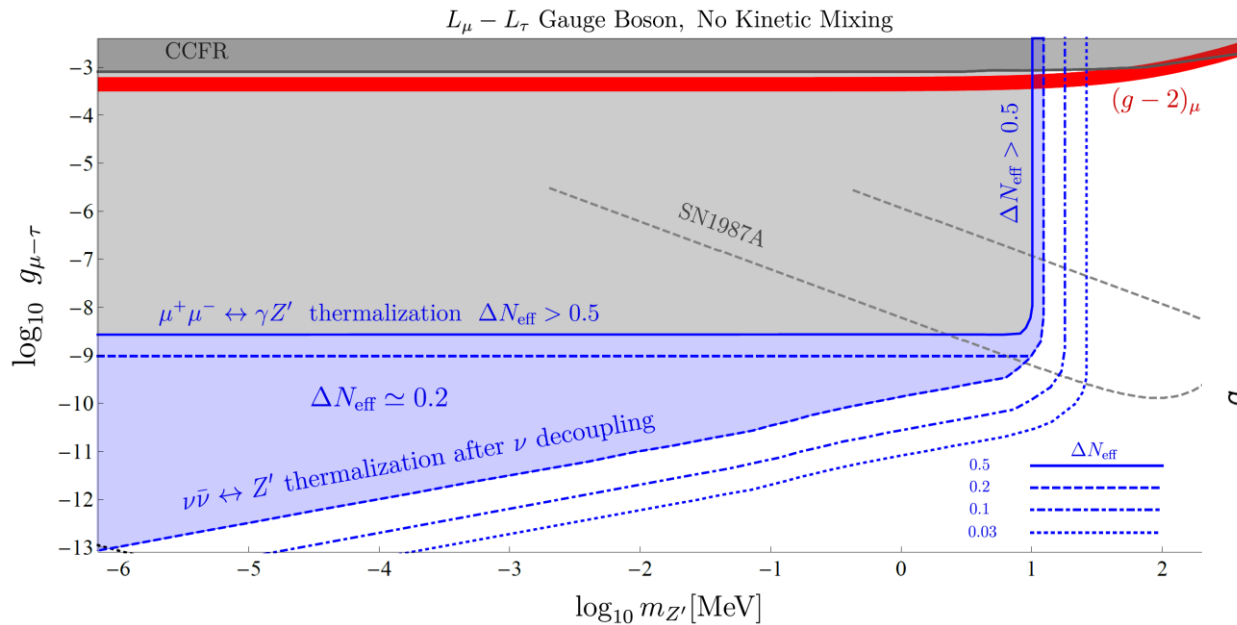
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- Hubble parameter
- New physics interpretation
  - $\Delta N_{\text{eff}}$  relaxes Hubble tension [D’Eramo et al (2018), Planck (2018), ...]



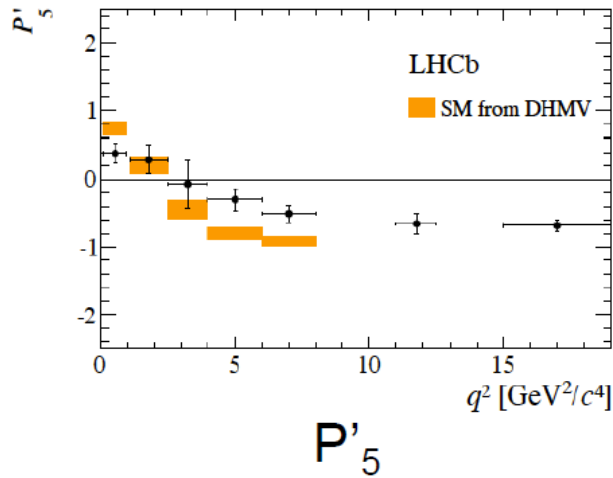
# § § Hubble tension

- Hubble parameter
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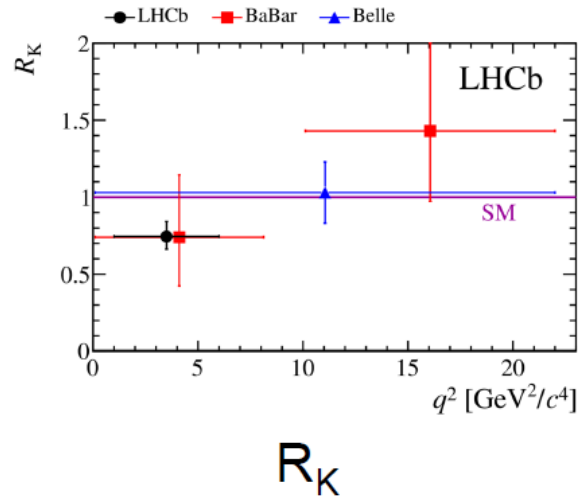


# § § B anomaly

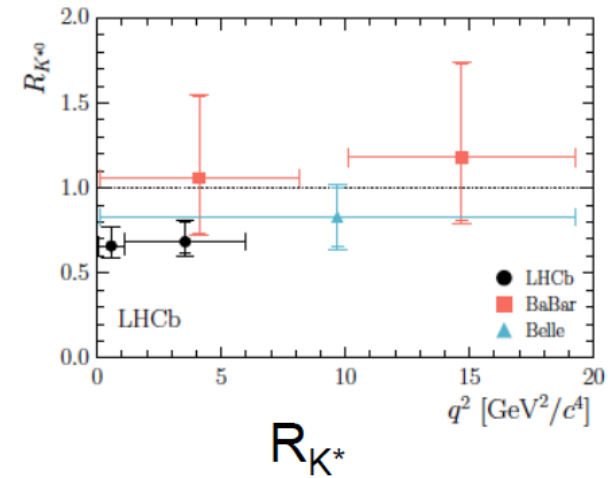
- $B \rightarrow K ll$  anomaly at LHCb [(2013), ...] and Belle [(2016)]



LHCb, JHEP 1602 (2016) 104



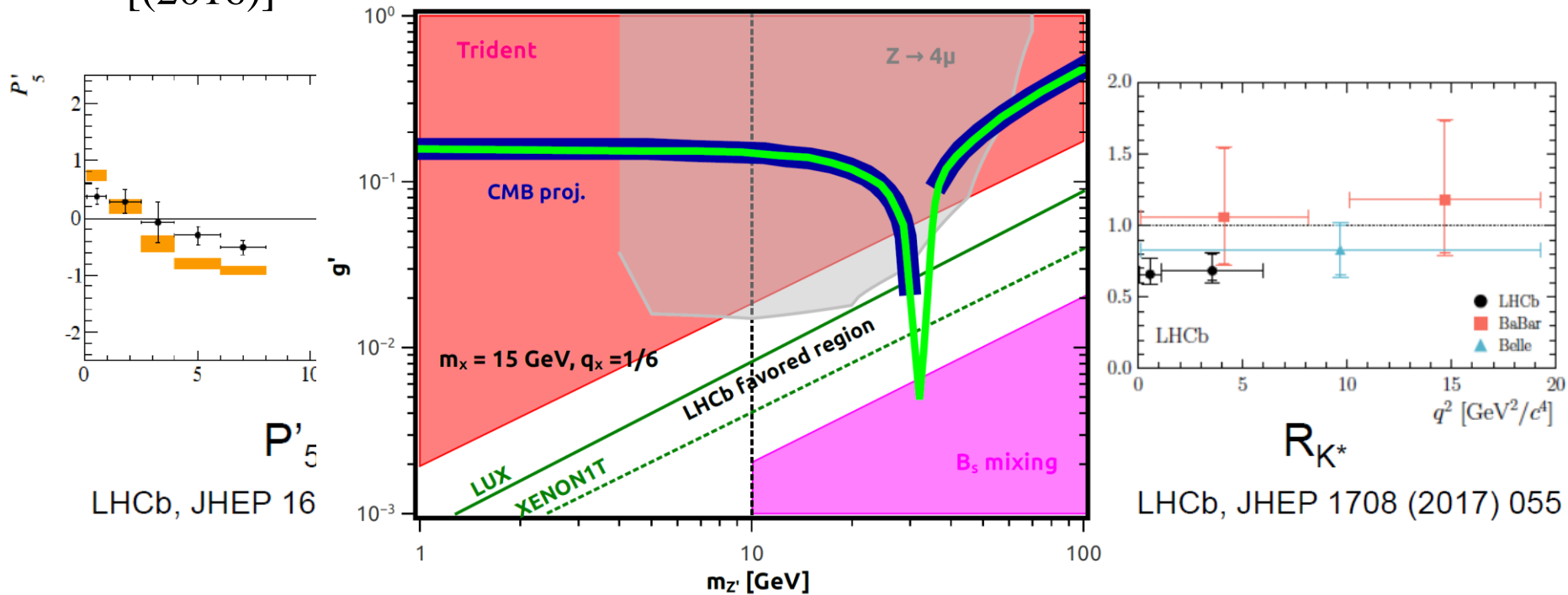
LHCb, PRL 113 (2014) 151601



LHCb, JHEP 1708 (2017) 055

# § § B anomaly

- $B \rightarrow K ll$  anomaly at LHCb [(2013), ...] and Belle [(2016)]



- Light  $U(1)_{L\mu-L\tau}$  gauge boson interpretation [Altmannshofer et al (2014, 2016)]

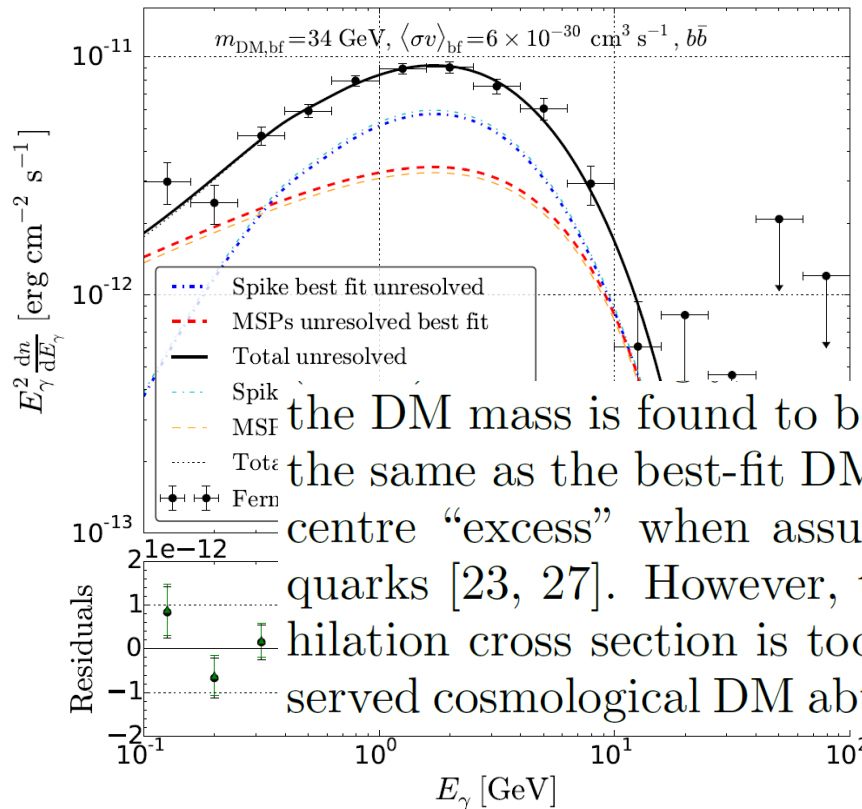
# § § Excess in globular cluster 47 Tuc

Understanding the  $\gamma$ -ray emission from the globular cluster 47 Tuc: evidence for dark matter?

Anthony M. Brown,<sup>1,\*</sup> Thomas Lacroix,<sup>2</sup> Sheridan Lloyd,<sup>1</sup> Céline Boehm,<sup>3,4,5,6</sup> and Paula Chadwick<sup>1</sup>

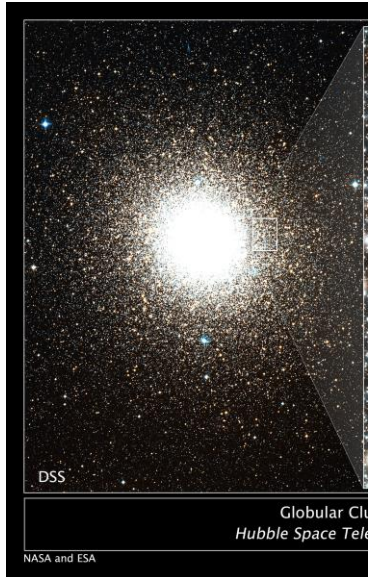
<sup>1</sup>Centre for Advanced Instrumentation, Department of Physics, University of Durham. South Road. Durham, DH1 3LE, UK

[ Brown et al (2018)]



the DM mass is found to be 34 GeV, which is essentially the same as the best-fit DM explanation for the Galactic centre “excess” when assuming DM annihilation into  $b$  quarks [23, 27]. However, the value of our best-fit annihilation cross section is too small to account for the observed cosmological DM abundance, but this might a hint

- Several GeV mass for annihilation into  $\tau\tau$ ??



# § Gauged U(1) scalar DM Model

# § Gauged U(1) scalar DM Model

- Particle content

	SU(3) <sub>c</sub>	SU(2) <sub>L</sub>	U(1) <sub>Y</sub>	U(1)
$Q^i$	<b>3</b>	<b>2</b>	1/6	$q_{Q^i}$
$u_R^i$	<b>3</b>	<b>1</b>	2/3	$q_{u^i}$
$d_R^i$	<b>3</b>	<b>1</b>	-1/3	$q_{d^i}$
$L^i$	<b>1</b>	<b>2</b>	-1/2	$q_{L^i}$
$e_R^i$	<b>1</b>	<b>1</b>	-1	$q_{e^i}$
$\Phi$	<b>1</b>	<b>2</b>	1/2	0
$N_R^i$	<b>1</b>	<b>1</b>	0	$q_{N^i}$
$\phi_1$	<b>1</b>	<b>1</b>	0	+1
$\phi_2$	<b>1</b>	<b>1</b>	0	+2

- $q_X$  to be anomaly free

- Dark matter with the fixed charge so that it interacts with the U(1) breaking Higgs field.

C.f. [Rodejohann and Yaguna (2015), Biswas et al (2016, 2018), Singirala et al (2016), Bandyopadhyay et al (2018)]



# § § Masses and interactions

- **Scalar potential**

$$\begin{aligned} V(\Phi, \phi_1, \phi_2) = & -M_\Phi^2 |\Phi|^2 + \frac{\lambda}{2} |\Phi|^4 + M_{\phi_1}^2 \phi_1 \phi_1^\dagger - M_{\phi_2}^2 \phi_2 \phi_2^\dagger \\ & + \frac{1}{2} \lambda_1 (\phi_1 \phi_1^\dagger)^2 + \frac{1}{2} \lambda_2 (\phi_2 \phi_2^\dagger)^2 + \lambda_3 \phi_1 \phi_1^\dagger (\phi_2 \phi_2^\dagger) \\ & + (\lambda_4 \phi_1 \phi_1^\dagger + \lambda_5 \phi_2 \phi_2^\dagger) |\Phi|^2 - A(\phi_1 \phi_1 \phi_2^\dagger + \phi_1^\dagger \phi_1^\dagger \phi_2) \end{aligned}$$

- **Mases of DM and Higgs bosons**

$$\begin{aligned} \mathcal{L}_{\text{mass}} = & -\frac{1}{2} (\varphi \ \varphi_2) \begin{pmatrix} -M_\Phi^2 + \frac{3}{2} \lambda v^2 + \frac{1}{2} \lambda_5 v_2^2 & \lambda_5 v v_2 \\ \lambda_5 v v_2 & -M_{\phi_2}^2 + \frac{3}{2} \lambda_2 v_2^2 + \frac{1}{2} \lambda_5 v^2 \end{pmatrix} \begin{pmatrix} \varphi \\ \varphi_2 \end{pmatrix} \\ & -\frac{1}{2} \left( M_{\phi_1}^2 + \frac{1}{2} \lambda_3 v_2^2 + \frac{1}{2} \lambda_4 v^2 - \sqrt{2} A v_2 \right) S^2 \\ & -\frac{1}{2} \left( M_{\phi_1}^2 + \frac{1}{2} \lambda_3 v_2^2 + \frac{1}{2} \lambda_4 v^2 + \sqrt{2} A v_2 \right) P^2 \end{aligned}$$

# § § Masses and interactions

- Interactions

- Gauge interactions

$$\mathcal{L}_{\text{int}} = g' Z'^{\mu} ((\partial_{\mu} S) P - S \partial_{\mu} P)$$

- Absence of DM-DM-Z': Inelastic

- Scalar interactions

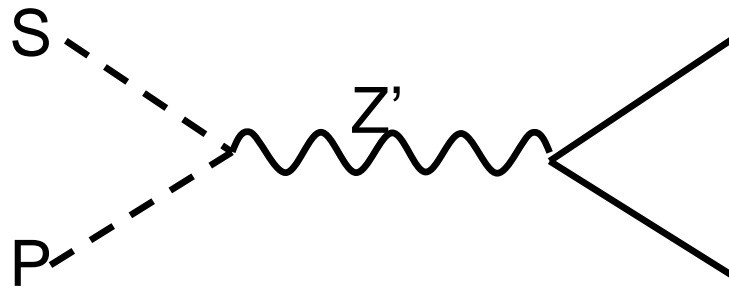
$$\mathcal{L}_{\text{int}} \supset \frac{1}{2} \left( \left( \lambda_4 v \cos \alpha - (\lambda_3 v_2 - \sqrt{2} A) \sin \alpha \right) h + \left( \lambda_4 v \sin \alpha + (\lambda_3 v_2 - \sqrt{2} A) \cos \alpha \right) H \right) S^2 \\ + \frac{1}{2} \left( \left( \lambda_4 v \cos \alpha - (\lambda_3 v_2 + \sqrt{2} A) \sin \alpha \right) h + \left( \lambda_4 v \sin \alpha + (\lambda_3 v_2 + \sqrt{2} A) \cos \alpha \right) H \right) P^2$$

- The direct DM search bound for Higgs bosons exchange processes is avoidable by taking those very small
      - Not used in freeze-out annihilation.

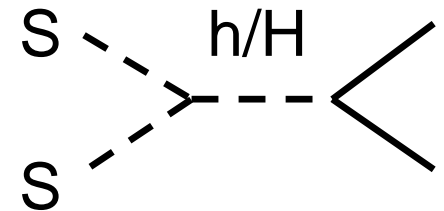
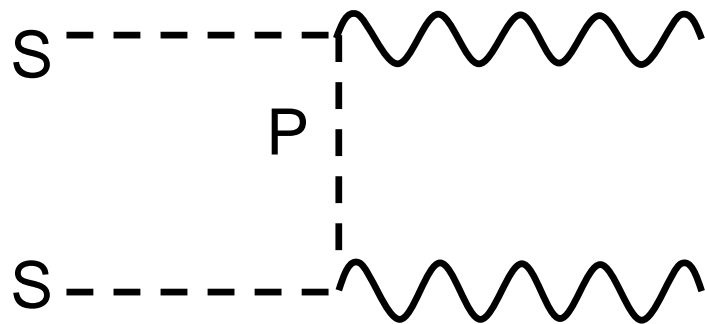
# § § Annihilation

- Annihilation modes

- Co-annihilation



- Into  $Z'$  pair

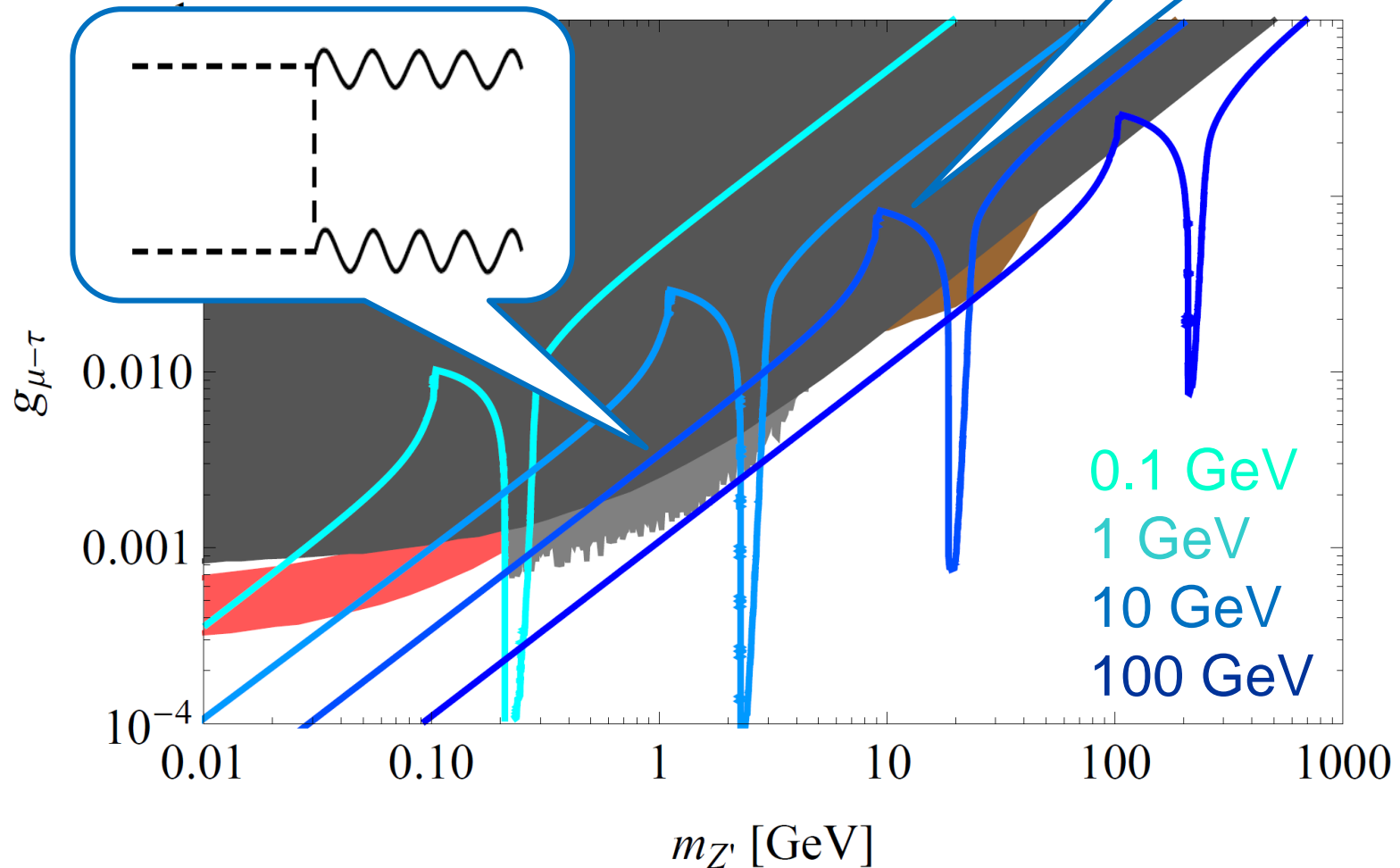


## § For specific U(1) models

- $U(1)_{B-L}$ 
  - At TeV scale
- $U(1)_{(B-L)3}$ 
  - At the weak scale
- $U(1)_{L\mu-L\tau}$ 
  - From MeV to the weak scale

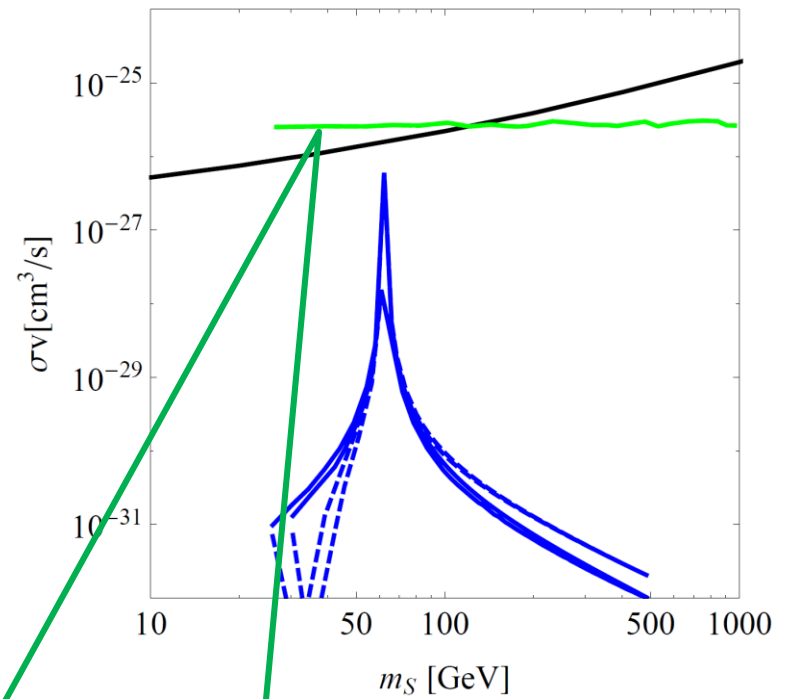
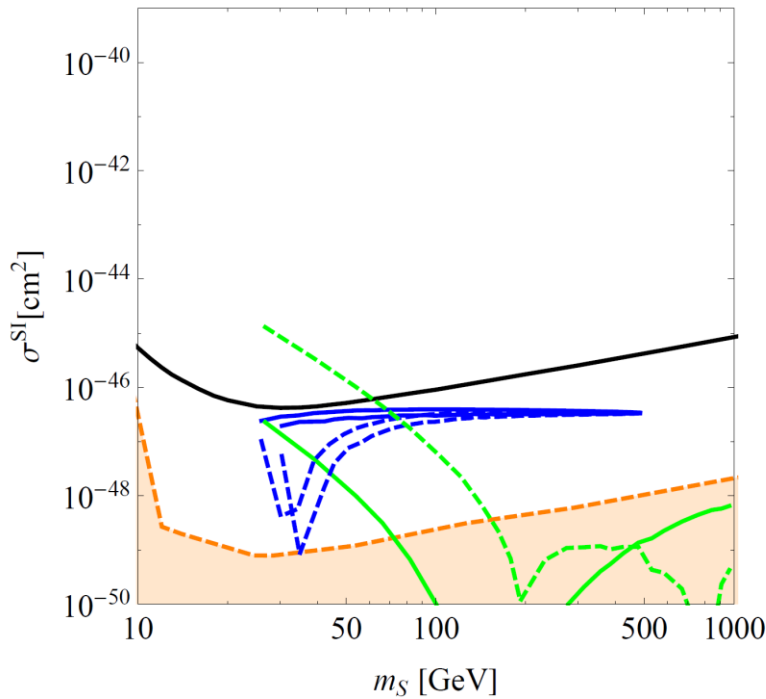
# § § $L_\mu - L_\tau$ model

- Thermal abundance and others

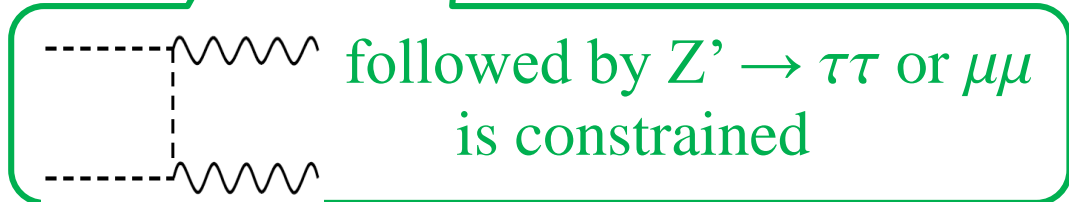


# § § $L_\mu - L_\tau$ model

- Constraints from and prospect for direct and indirect DM searches

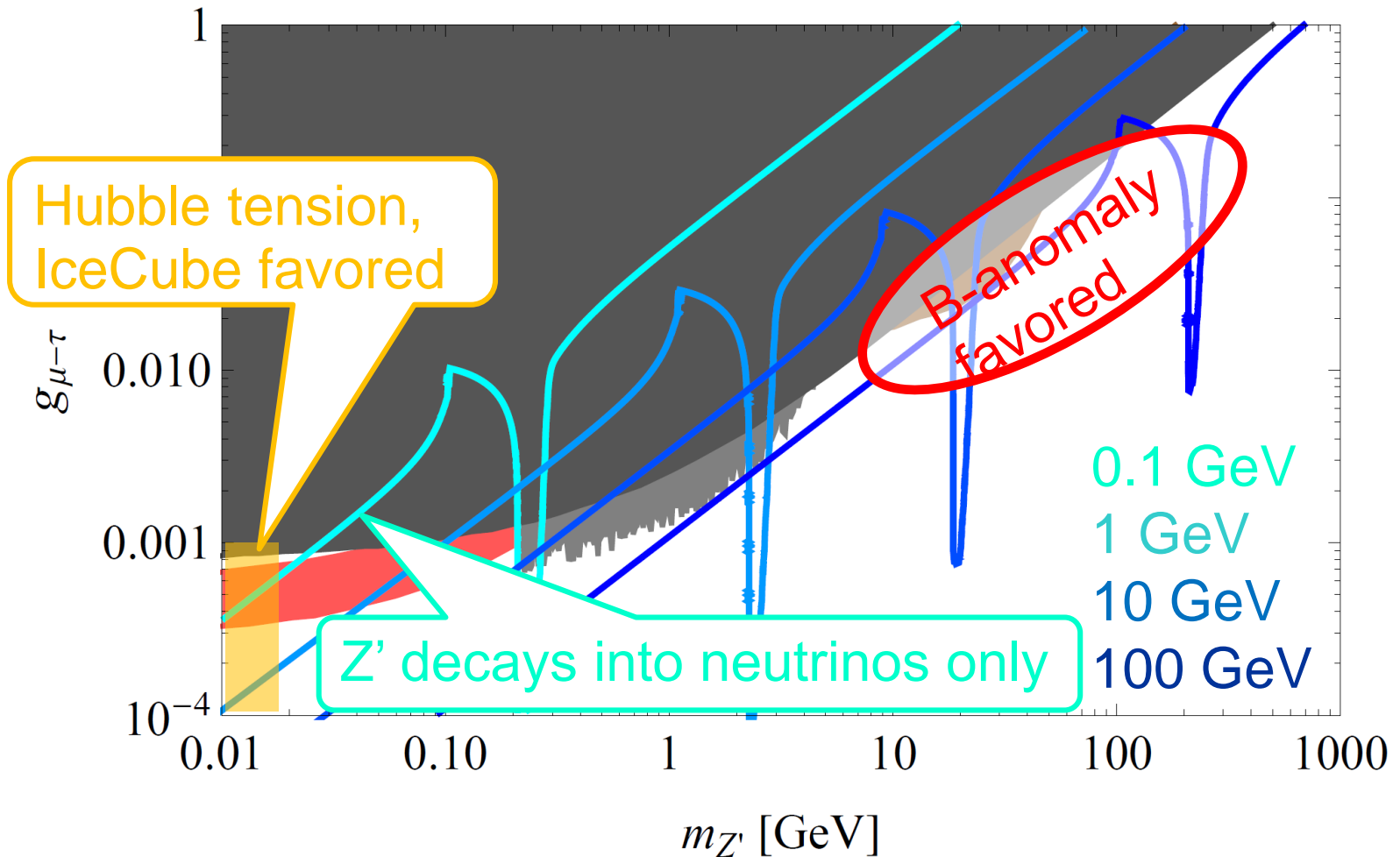


for  $g_{\mu-\tau} = 0.03$



# § § $L_\mu - L_\tau$ model

- Thermal abundance and others



# § Summary

- Light neutral gauge boson suggested?
  - g-2 of muon, IceCube, Hubble tension, B-anomaly...
- Gauged U(1) scalar DM model with DM-DM-U(1) Higgs coupling
- Gauged  $L_\mu - L_\tau$  offers the possibility of low mass dark matter and the mediator  $Z'$ 
  - 10 GeVish, B-anomaly, 47 Tuc excess
  - 10 MeVish, muon g-2, Hubble tension, IceCube