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## Mass measurement in the $N=40$ region with JYFLTRAP

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We report on a precise mass measurement of  $^{69}\text{Co}$  and  $^{70}\text{Co}$  performed at the IGISOL facility [1].

The nuclear structure above the  $N=40$  subshell closure remained for long unclear due to lack of information on neutron-rich nuclei in this region. This region is known for shape coexistence and intruder states have been observed in neutron-rich cobalt isotopes toward  $N=40$  [2]. The two-neutron separation energies and the empirical two-neutron shell-gap determined from the mass value give important insight on the evolution of the subshell closure. In nuclear astrophysics, the  $^{68}\text{Co}(n,\gamma)^{69}\text{Co}$  reaction have been highlighted in sensitivities studies to strongly influence the abundance pattern for the weak  $r$  process, which produces the lightest  $r$ -process elements [3]. The ratio of the photodisintegration to the neutron-capture rate depends exponentially on the reaction  $Q$ -value, stressing the need of precise mass value.

The double Penning trap JYFLTRAP [4] at the University of Jyväskylä has been successfully used to measure the masses of 12 nuclides in the  $N=40$  region. Among these, the masses of  $^{69}\text{Co}$  and  $^{70}\text{Co}$  were determined for the first time [1]. The isotopes of interest were produced via proton-induced fission at the IGISOL facility [5] during one week of experiment. The time-of-flight ion cyclotron resonance (TOF-ICR) technique [6] was used for the mass measurements. The  $\frac{1}{2}^-$  isomer in  $^{69}\text{Co}$  has been determined via ion beam manipulation technique, taking advantage of the difference of half-lives between the ground and the isomeric state in  $^{69}\text{Co}$ . In addition, the phase-imaging ion cyclotron resonance technique (PI-ICR) [7] was used for the first time online to determine the composition of the  $^{70}\text{Co}$  beam and the long living state of  $^{70}\text{Co}$  was precisely determined. Our experimental data were compared to large-scale shell model calculations and confirm the weakening of the  $N=40$  subshell closure below nickel. We also considerably reduced the mass-related uncertainty in the photodisintegration rate  $^{69}\text{Co}(\gamma,n)^{68}\text{Co}$  and in the ratio of the photodisintegration to the neutron-capture rate.

In this talk, the experimental method and the impact of our results on astrophysics and nuclear structure in the  $N=40$  region will be discussed.

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