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Fusion of $^{12}\text{C} + ^{24}\text{Mg}$ at extreme sub-barrier energies

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The study of heavy-ion fusion reactions of light systems is essential for the understanding of the astrophysical reaction networks responsible for the energy production and elemental synthesis in stellar environments. At far sub-barrier energies, fusion is influenced by the hindrance phenomenon [1]. Fusion of light heavy-ions is characterised by a positive Q-value and establishing the presence of hindrance in such systems requires challenging measurements. In the two relevant cases $^{12}\text{C}+^{12}\text{C}$ and $^{16}\text{O}+^{16}\text{O}$ the available data give indication of the existence of hindrance, but the situation is far from being well established, especially for $^{12}\text{C}+^{12}\text{C}$ fusion where the several resonances observed [1], make very controversial any conclusion. The study of slightly heavier systems is then appealing, because their low-energy fusion trend may provide a reliable guidance for the extrapolation to the lighter cases of astrophysical interest.

We have measured [2] the fusion excitation function (and consequently the S-factor) for the system $^{12}\text{C}+^{24}\text{Mg}$, using the magnesium beam from the XTU Tandem accelerator of INFN-LNL and 50 micro-g/cm² ^{12}C targets enriched to 99.9% in mass 12, down to around $\sigma_{\text{fus}}=4.7$ micro-b. Fusion hindrance shows up because the S-factor displays a well-defined maximum vs energy below the barrier around $\sigma_{\text{fus}}=1$ mb. It is remarkable that the lowest cross sections are consistent with a simple one-dimensional barrier penetration calculation. The S-factor trend is well reproduced using an empirical formula in the spirit of the adiabatic model [3], as well as using the phenomenological hindrance model [4].

Measurements at energies slightly below the present ones are in schedule and will allow a deeper insight into the fusion dynamics far below the barrier. Far-reaching consequences may be envisaged for the lighter systems relevant for astrophysics.

[1] C.L.Jiang et al., Eur. Phys. J. A57, 235 (2021)

[2] G.Montagnoli et al., Phys. Rev. C101, 044608 (2020), and submitted to J.Phys. G

[3] T. Ichikawa, K. Hagino and A. Iwamoto, Phys. Rev. C75, 057603 (2007)

[4] C.L. Jiang et al., Phys. Rev. C75, 015803 (2007)

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