



Contribution ID: 133

Type: Poster

## New results on the level structure of $^{26}\text{Si}$ and consequences for the $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$ reaction in Classical Novae environments

Tuesday 6 September 2022 18:50 (2 minutes)

The  $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$  reaction is of tremendous interest in nuclear astrophysics. The production of the  $\gamma$ -ray emitter  $^{26}\text{Al}$  ground state can be bypassed in classical novae via the production of  $^{26}\text{Si}$  which decays to an isomeric state of  $^{26}\text{Al}$ . In order to more precisely estimate the amount of  $^{26}\text{Al}$  that is of classical novae origin, it's crucial to determine the rate of the  $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$  reaction at nova-burning temperatures. The production of  $^{26}\text{Si}$  is dominated by resonant captures to several excited states above the proton threshold in  $^{26}\text{Si}$ . There has been considerable experimental effort in recent years to observe and identify these states [1], but the properties of the key resonances in  $^{26}\text{Si}$  remain unsettled.

The combination of GRETTINA [2] coupled with the Fragment Mass Analyzer (FMA) [3] at Argonne National Laboratory (ANL), provided a powerful opportunity to identify transitions in  $^{26}\text{Si}$ , owing to the large acceptance of the separator and the Doppler-reconstruction capabilities and high-energy efficiency of the GRETTINA array. The experiment, presented here, follows an earlier  $\gamma$ -ray spectroscopy study of the  $^{26}\text{Si}$  mirror nucleus,  $^{26}\text{Mg}$ , performed with Gammasphere at ANL where a  $l=1$  resonance was identified for the first time (fig.1) [4]. In the same study, the lifetime of the  $3+$ , 6125-keV state in  $^{26}\text{Mg}$  was measured via the Doppler shift attenuation method. The  $3+$ , 414-keV resonance in  $^{26}\text{Si}$  dominates the  $^{25}\text{Al}(p,\gamma)$  reaction over most of the novae peak temperature range, while the introduction of the new  $1-$  state increases the reaction rate by  $\sim 25\%$  at the highest novae temperatures.

In this talk, new results on  $^{26}\text{Si}$  from the GRETTINA+FMA study will be presented along with further information gained on the  $A=26$  system. Information on both the level structure of  $^{26}\text{Si}$  and the impact on the astrophysical  $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$  reaction will be discussed.

[1] K. Chipps, Phys. Rev. C 93, 035801 (2016).

[2] D. Weisshaar et al, Nucl. Instrum. Methods Phys. Res. A 847, 187 (2017).

[3] C.N. Davis et al., Nucl. Instrum. Methods Phys. Res. B 70, 358 (1992).

[4] L. Canete et al, Phys. Rev. C 104, L022802 (2021).

### Field of work

**Author:** Dr CANETE, Laetitia

**Presenter:** Dr CANETE, Laetitia

**Session Classification:** Poster session