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Stellar s -process neutron capture cross sections on ^ASe and ^ACe

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Neutron-induced reactions remain at the forefront of experimental investigations for the understanding of stellar nucleosynthesis and chemical evolution of the Galaxy in the region of medium- and heavy-mass nuclides [1]. We report on measurements of the cross section of neutron capture reactions $^{74,80,82}\text{Se}(n, \gamma)$ and $^{138,140,142}\text{Ce}(n, \gamma)$ relevant, respectively, to the weak and main s -processes. The ^ASe data complement our recent study of the $^{69,71}\text{Ga}$ stellar (n, γ) reactions in the weak s -process regime [2]. The proton rich isotope ^{74}Se is a p nuclide, shielded from the s and r processes by stable nuclei in the region. The disentanglement of the different heavy-nuclide synthesis modes (s -, r - and p -processes) requires reliable and precise stellar neutron-capture cross sections. Such is the case also for the Ce isotopes [3]. In particular, ^{140}Ce is found to be one of the most important nuclides in the network of s -process reactions, affecting the abundances of a large number of isotopes [4]. The experiments were performed by the activation technique using a high-intensity ($3\text{--}5 \times 10^{10}$ n/s) quasi-Maxwellian neutron beam that mimics conditions of stellar s -process nucleosynthesis. The neutron field was produced by a mA proton beam at $E_p = 1925$ keV (beam power of 2–3 kW) as part of our experiment campaign at the Phase I of Soreq Applied Research Accelerator Facility (SARAF) [5], bombarding the Liquid-Lithium Target (LiLiT) [6,7]. The cross sections were measured by counting the resulted nuclei activities via γ spectrometry with a high-purity germanium detector.

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