



UNIVERSITÀ
DEGLI STUDI
DI PADOVA



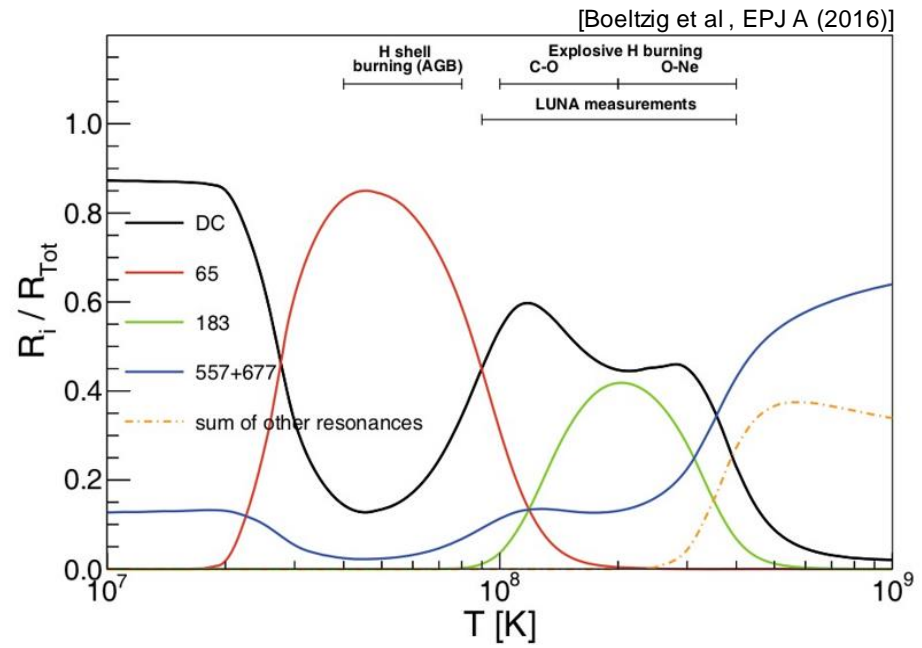
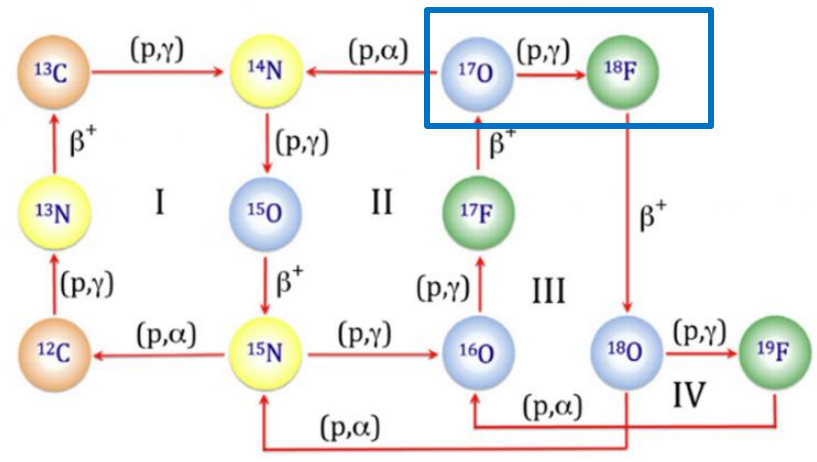
Towards a direct measurement of the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ 65 keV resonance strength at LUNA

Denise Piatti* for LUNA collaboration

* Università degli Studi di Padova and INFN Padova

Astrophysical Motivation

- $^{17}\text{O}(p,\gamma)^{18}\text{F}$ reaction ($Q = 5607 \text{ keV}$) takes part to CNO cycle, active during H-shell burning in AGB stars.
- Footprint of AGB nucleosynthesis and mixing processes is the oxygen isotopic ratio observed in presolar grains
- Model predictions still struggle to match observations
- For $30 < T < 100 \text{ MK}$ ($35 < E_G < 140 \text{ keV}$) the resonance $E_{\text{cm}}=65 \text{ keV}$ ($E_x = 5672 \text{ keV}$) dominates the reaction rate



Fractional contributions to the reaction rate of the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ as a function of the temperature

State of the Art

Only indirect measurements reported for the $E_{cm} = 65$ keV resonance :

- $\Gamma_{\alpha} = (130 \pm 5)$ eV from $^{14}\text{N}(\alpha, \alpha)^{14}\text{N}$ measurement [H-B. Mak et al. NPA (1980)]
- $\Gamma_{\gamma} = (0.44 \pm 0.02)$ eV from the measured $\omega\gamma(\alpha, \gamma)$ [I. Berka et al. NPA (1977)]
- $\Gamma_p = ?$ = indirect measurements and the direct measurement performed at LUNA [(p, α) channel – C. Bruno et al. PRL 117 (2016)] disagree by a factor of ~ 2

Updated estimate: $\omega\gamma = (1.6 \pm 0.3) \times 10^{-11}$ eV [M.Q. Buckner et al PRC 91, 015812 (2015)]

↓ Typical Direct Measurement ingredients:

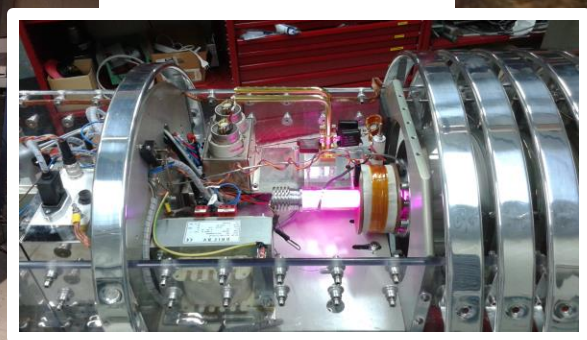
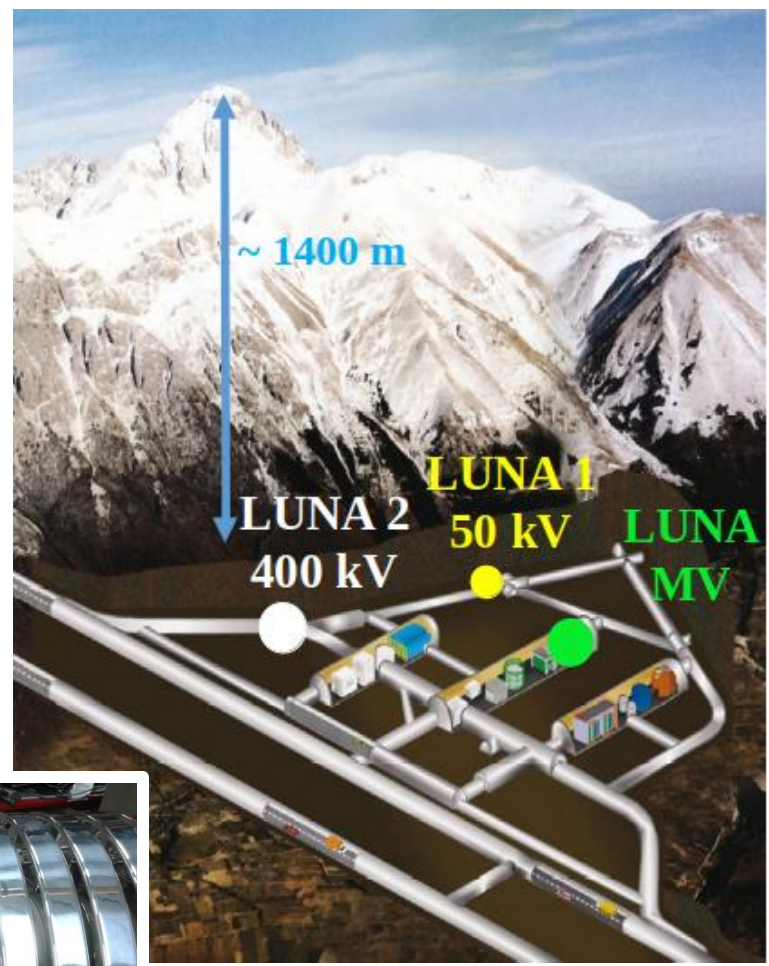
- 100% ^{17}O target
- Beam current $\sim 100 \mu\text{A}$

↓ 0.08 reactions/h

↓ **HIGH SENSITIVITY SETUP IS REQUIRED**

Laboratory for Underground Nuclear Astrophysics

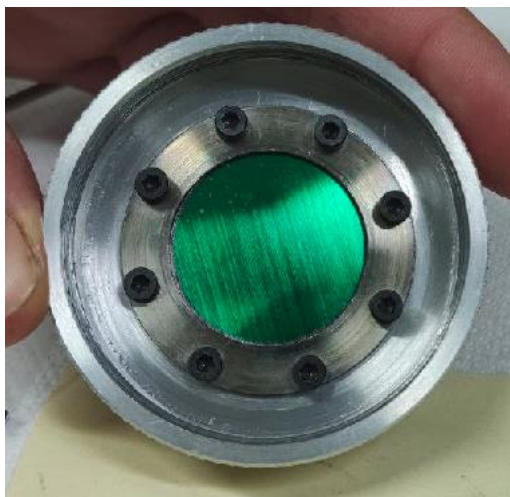
- Under 1400 m of rock = natural shielding
- Cosmic-ray background reduction
 - Muon = $\times 10^{-6}$
 - Neutron = $\times 10^{-3}$
 - In particle spectra = $\times 10^{-1}$
- High intensity and stability accelerator (up to mA and 0.5 mA for H⁺ and He⁺)



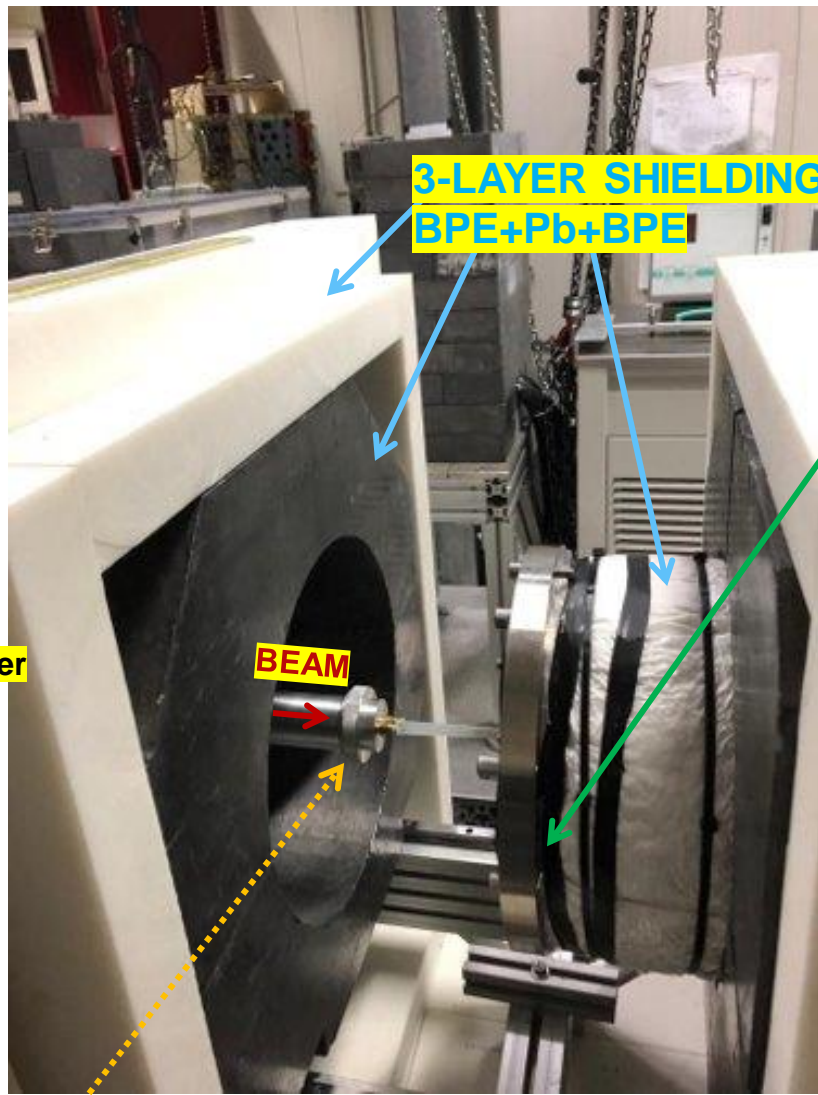
Setup



Ta₂O₅ target made by anodic oxidation with ¹⁷O (90%) enriched water

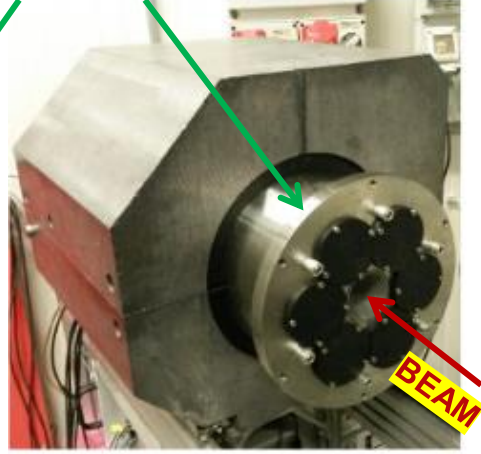


Al CHAMBER AND TARGET HOLDER



3-LAYER SHIELDING:
BPE+Pb+BPE

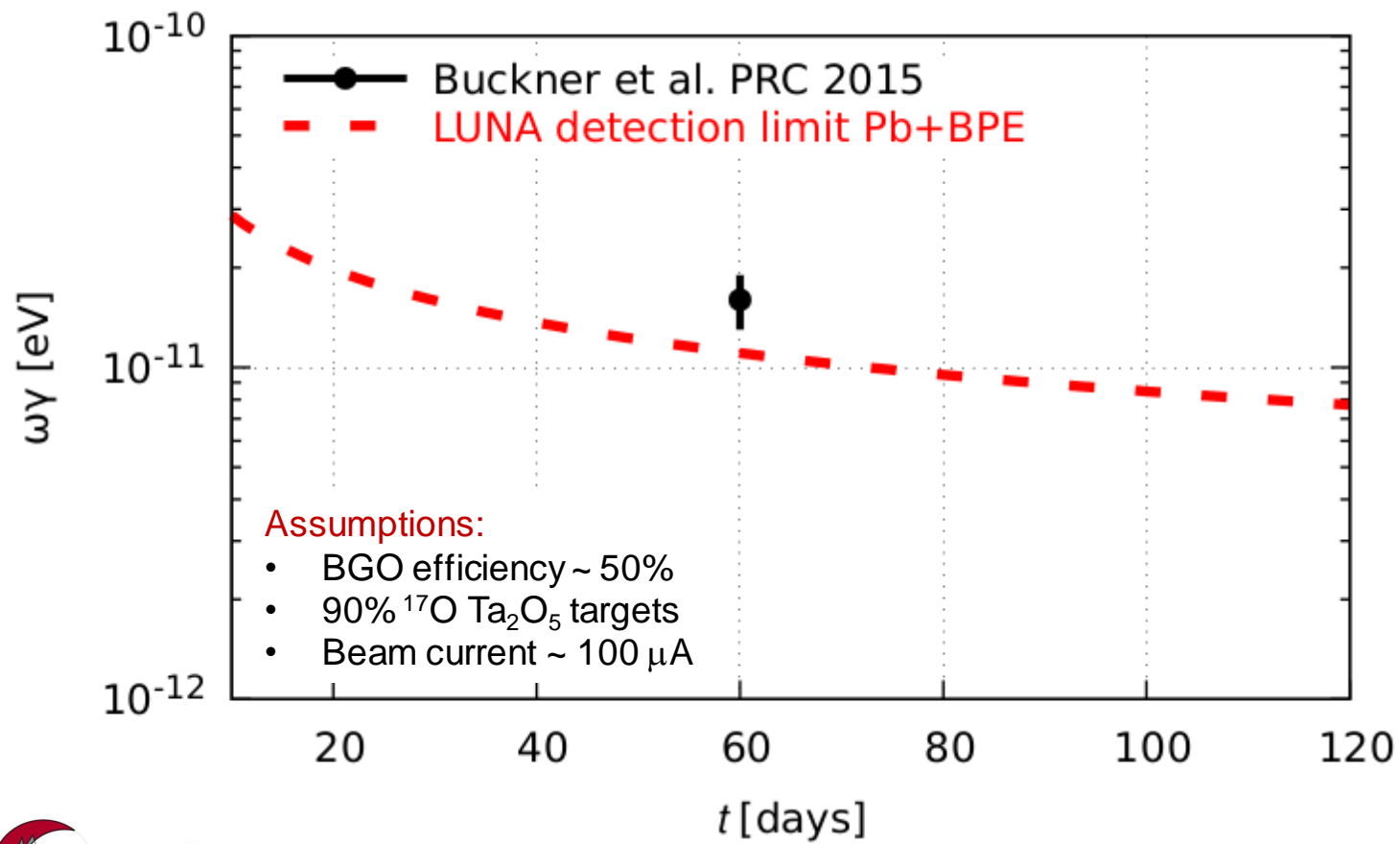
4II BGO DETECTOR
6 independent crystals
TAS mode possible



This part can move back and forth

Achieved Sensitivity

- 3 layer shielding: BPE + Pb + BPE -> reduction of the background by a factor ~5
- Al chamber and target holder -> ~20% increase in efficiency w.r.t. previous brass and stainless steel setup



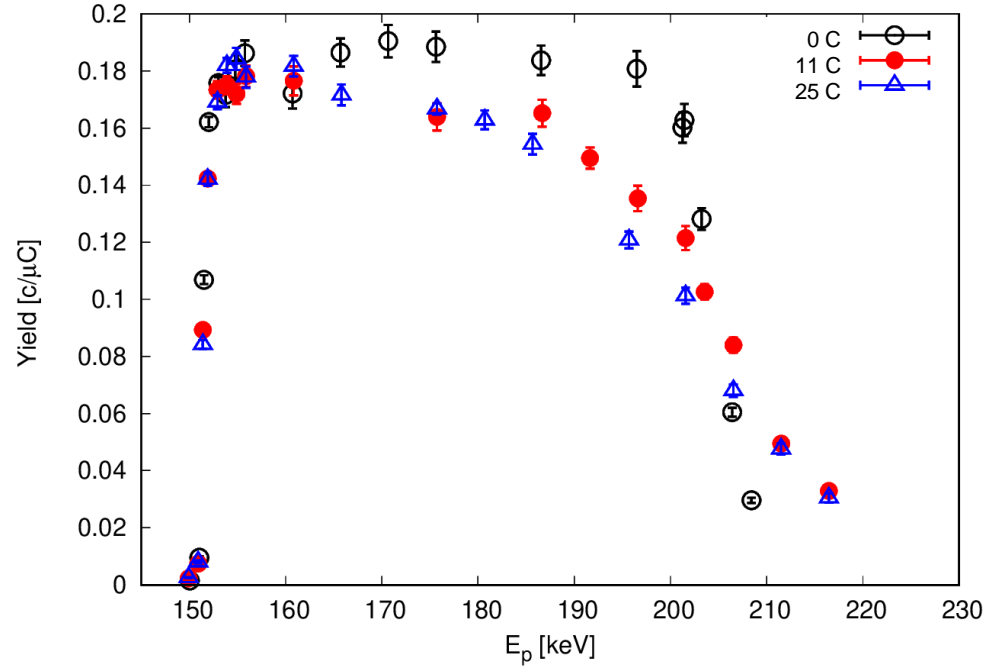
The Beam Induced Background nightmare

- BIB is due to contaminants in the oxide layer or in the backing that react with the beam
- The scariest contaminants are those that populates the ROI and with much higher cross section than the reaction of interest
- Ta is known for its H and D storage properties
- p+D reaction has a $Q = 5493.5$ keV (only ~ 100 keV lower than the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ reaction) and a cross section higher by many orders of magnitudes
- With BGO poor resolution the ROI for $^{17}\text{O}(p,\gamma)^{18}\text{F}$ reaction is 5200-6200 keV

No way to distinguish/resolve the p+D and the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ sum peak

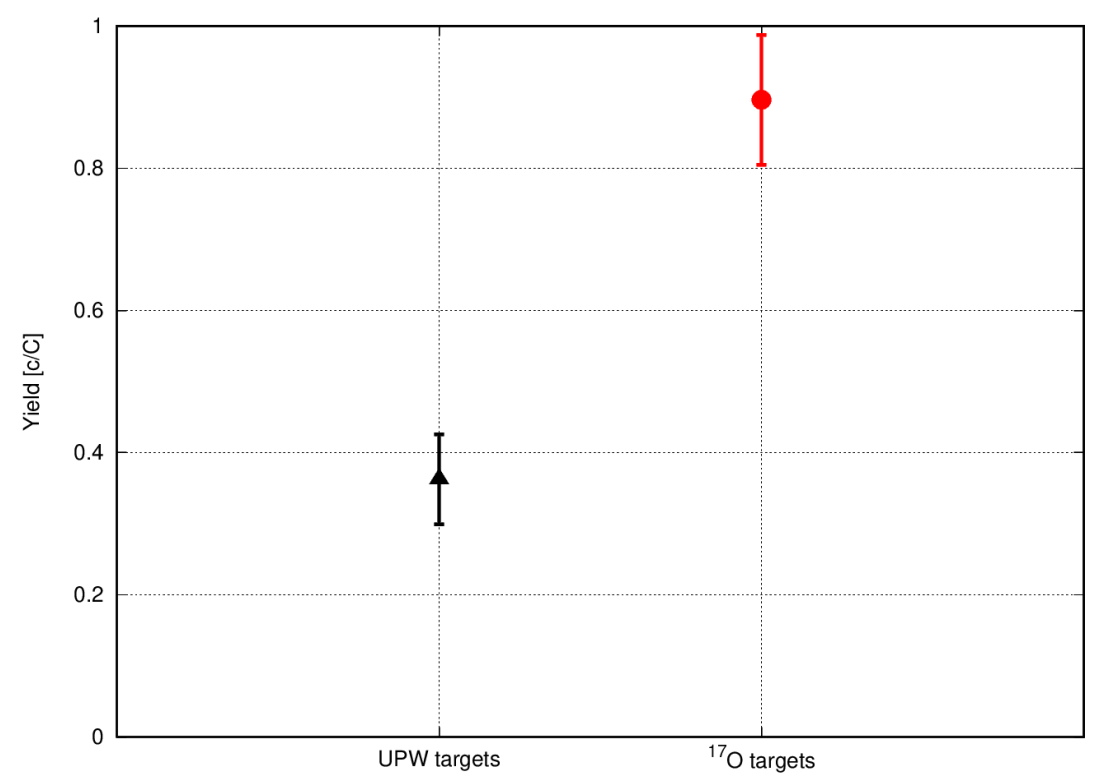
Data Acquisition

- BGO efficiency via simulations tuned on devoted measurements
- Target monitoring and online characterization via periodic scan of 151 keV resonance in $^{18}\text{O}(p,\gamma)^{19}\text{F}$ reaction + run on top of 193 keV resonance in $^{17}\text{O}(p,\gamma)^{18}\text{F}$
- Additional offline analysis for target characterization: SIMS and RBS
- 420 C on top of resonance with ^{17}O targets
- 300 C with UPW targets (NO ^{17}O) to monitor **Beam Induced Background**
- DC at $E_p=141, 184, 187, 190$ keV

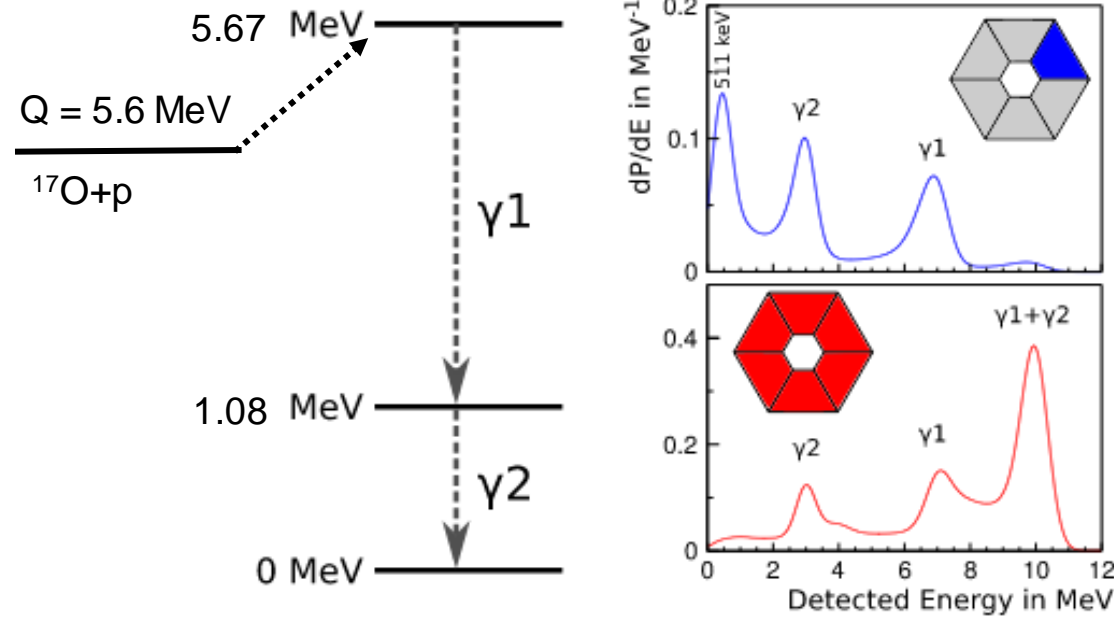


Analysis I

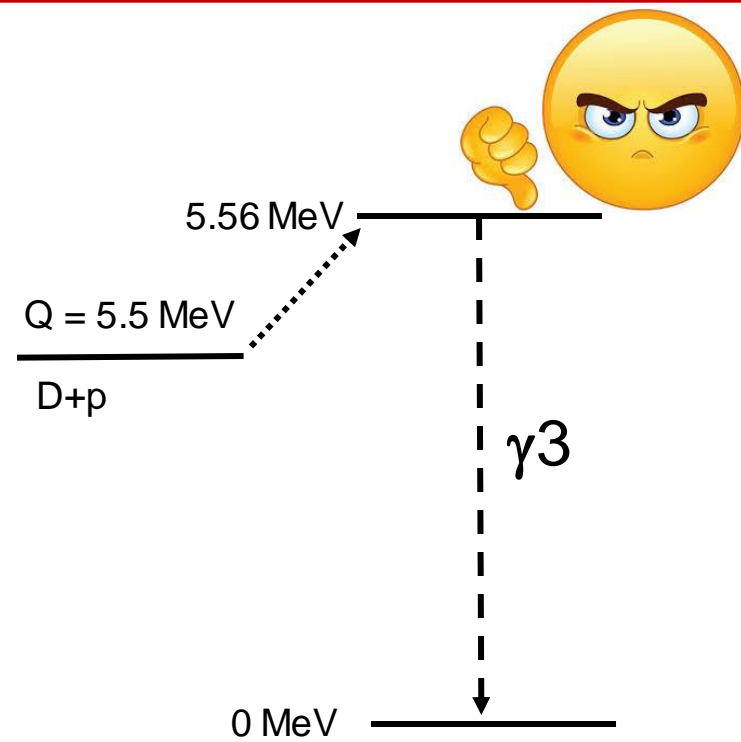
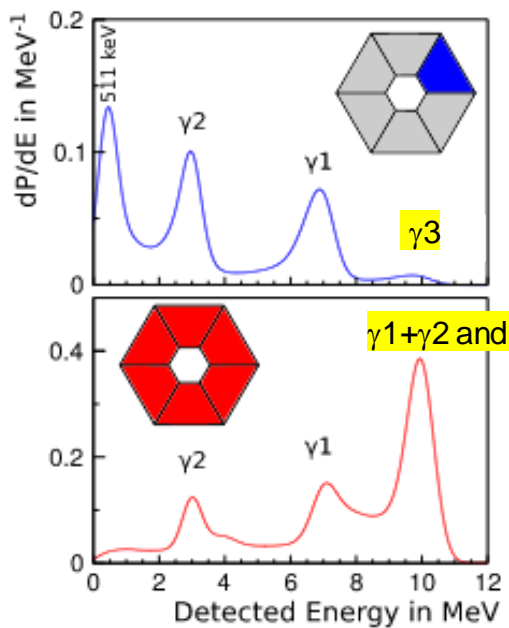
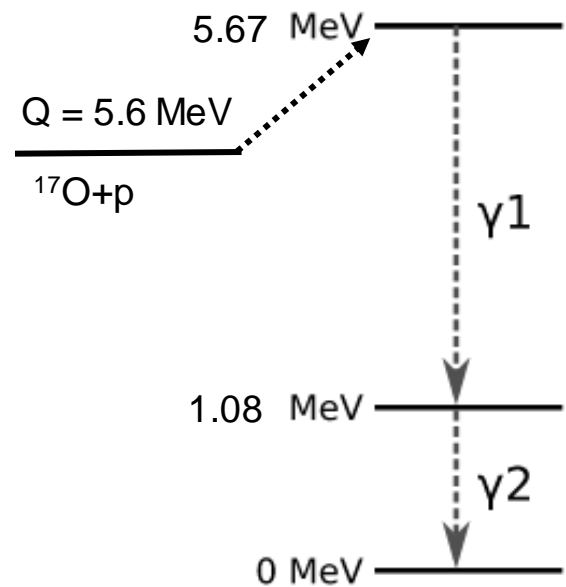
- Comparison between runs on the 65 keV resonance in UPW and ¹⁷O targets
- Looking for a significant excess in count rate
- Apply more refined statistical tools to get the $\omega\gamma$ -> Rolke Method



The Gate Analysis

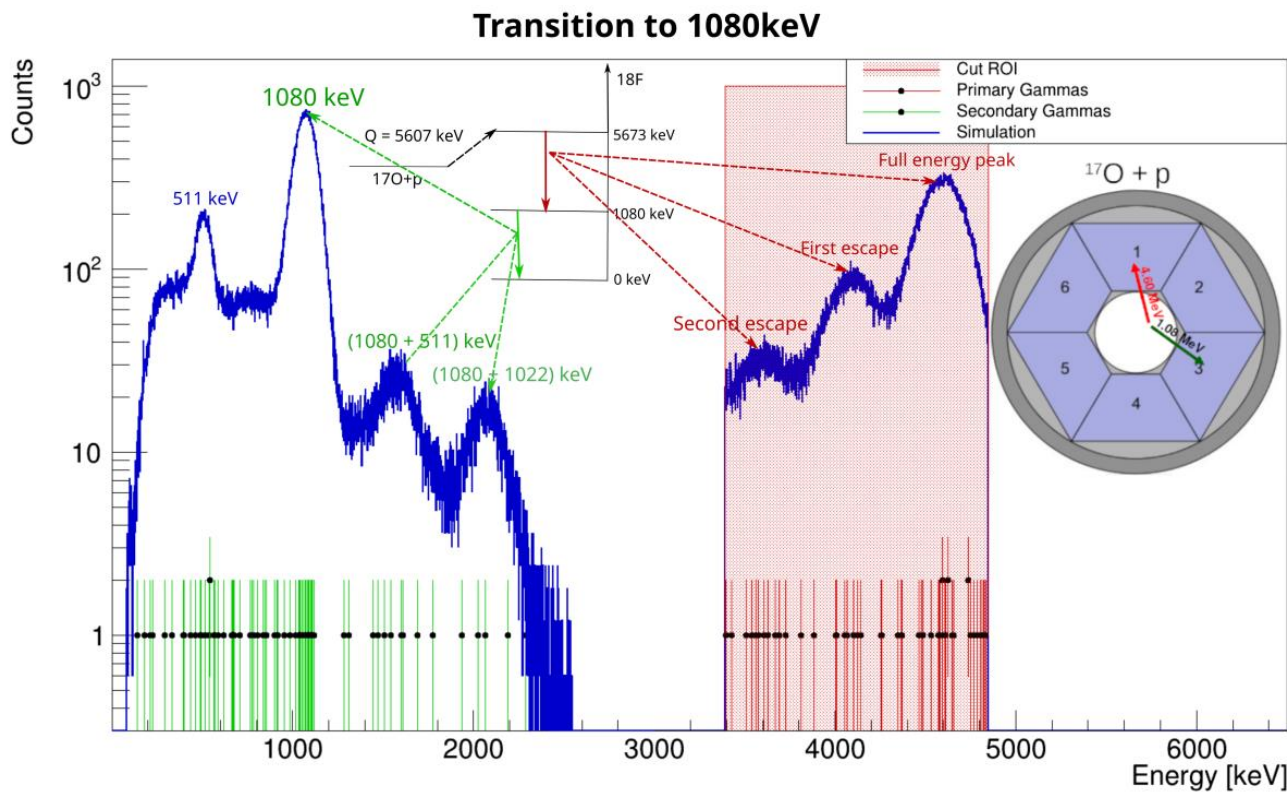


- Assume a cascade with 2 γ -rays emitted in coincidence
- The single BGO crystal will detect only one of the 2 γ -rays
- Since we are saving for each event the energy deposited and the timestamp we can produce the addback spectrum summing all the γ -rays in coincidence
- In the addback you observe the sum peak



BUT the sum peak might be also populated by a single γ -ray with $E_\gamma \sim \gamma_1 + \gamma_2$ produced by a contaminant reacting with the beam

- As well as you can construct the sum peak you can also de-construct it
- Gating in the ROI and looking at which γ -rays contributed to the sum peak
- You consider only coincident γ -rays corresponding to the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ cascade



Preliminary Results

- Target offline analysis is still ongoing
- As well as both 65 keV resonance data analysis
- They however reported preliminary results in good agreement suggesting a higher ω_γ than reported in literature

Preliminary
Results
Removed

LUNA COLLABORATION

A. Compagnucci*, R. Gesue', M. Junker | INFN LNGS

M. Axness-Ferrando, L. Barbieri, C. Broggin, A. Caciolli, P. Marigo, R. Menegazzo,
D. Piatti, J. Skowronski | Università di Padova and INFN Padova, Italy

A. Formicola, C. Gustavino | INFN Roma 1, Italy

D. Bemmerer, A. Boeltzig, E. Masha | HZDR Dresden, Germany

L. Csedreki, Z. Elekes, Zs. Fülöp, Gy. Gyürky, T. Szücs | MTA-ATOMKI Debrecen,
Hungary

M. Lugaro | Konkoly Observatory and ELTE University Budapest, Hungary

O. Straniero | INAF Osservatorio Astronomico di Collurania, Teramo, Italy

P. Corvisiero, P. Prati, S. Zavatarelli | Università di Genova and INFN Genova, Italy

R. Depalo, A. Guglielmetti, G. De Gregorio | Università di Milano and INFN Milano,
Italy

C. Ananna, A. Best, A. Di Leva, G. Imbriani, D. Rapagnani | Università di Napoli and
INFN Napoli, Italy

F. Cavanna, P. Colombetti, G. Gervino | Università di Torino and INFN Torino, Italy

M. Aliotta, C. Bruno, T. Davinson, R. Sidhu | University of Edinburgh, United
Kingdom

F. Barile, G. Ciani, V. Paticchio, R. Perrino, L. Schiavulli | Università di Bari and INFN
Bari, Italy

R. Perrino | INFN Lecce, Italy

* and GSSI, Italy

Thank You For Your Attention!