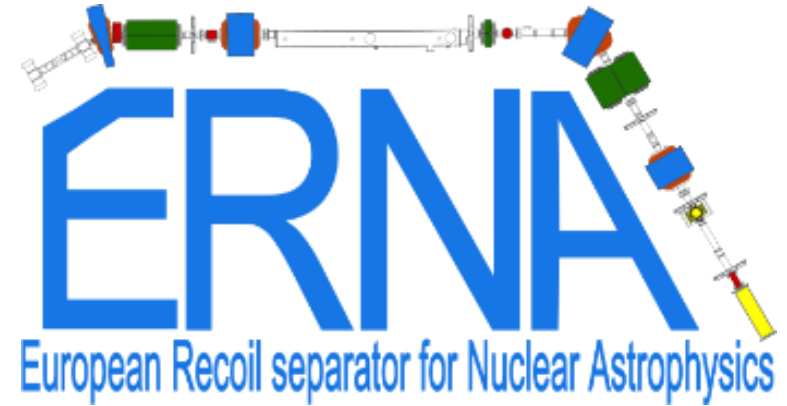


Measurement of the ${}^7\text{Be}(p,\gamma){}^8\text{B}$ reaction cross section with the recoil mass separator ERNA

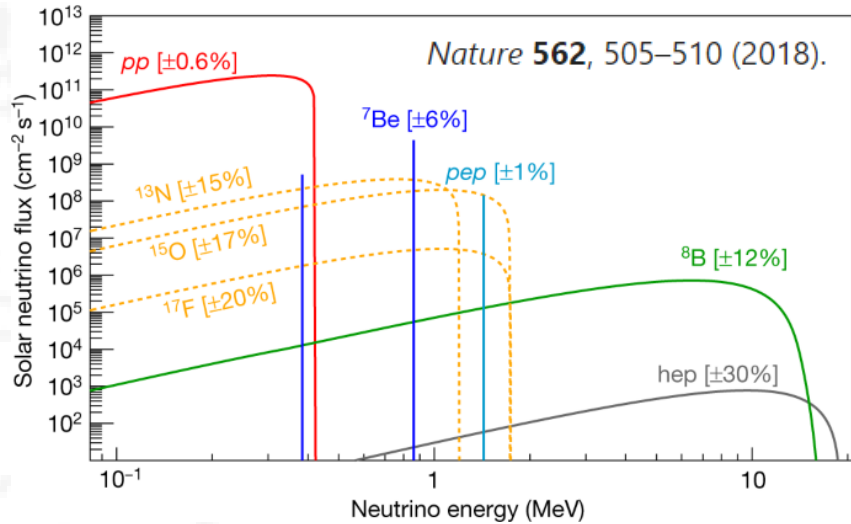
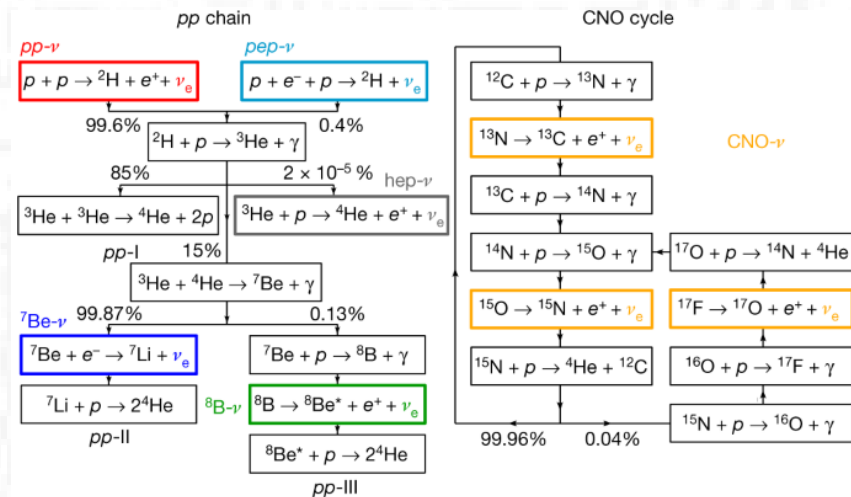


- **Raffaele Buompane**
- Università degli Studi della Campania “Luigi Vanvitelli” and INFN Naples

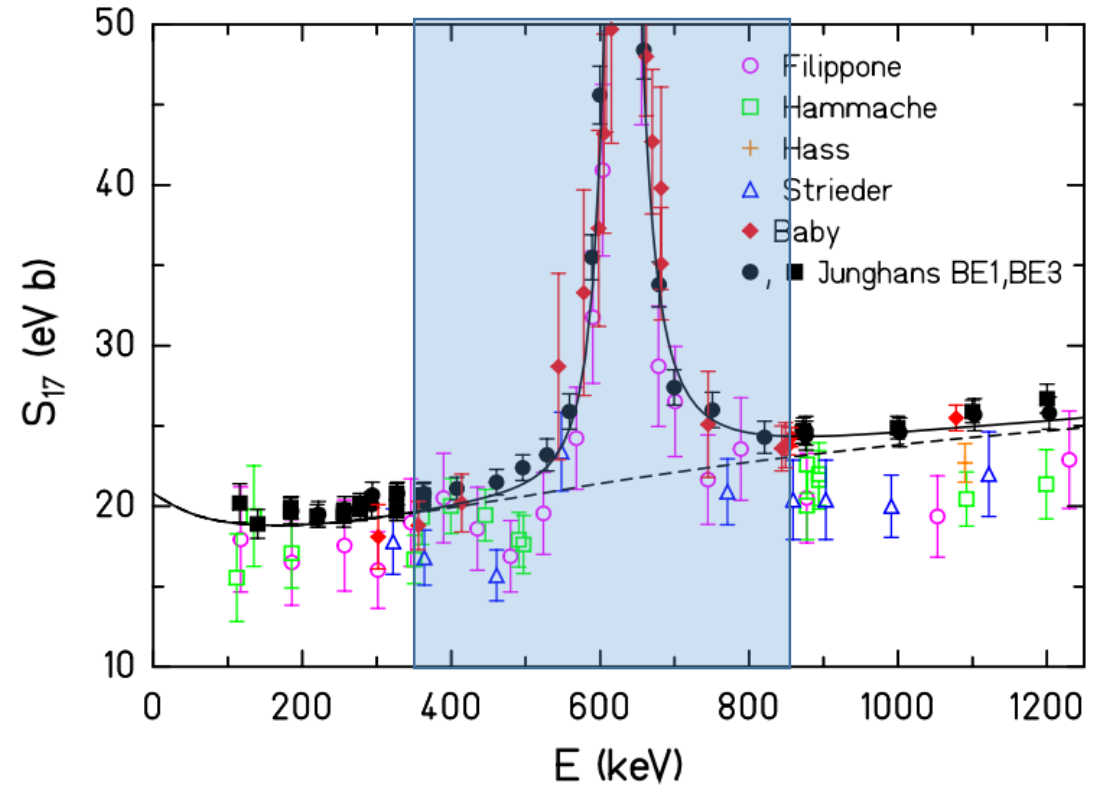
V : Università degli Studi della Campania *Dipartimento di Matematica e Fisica Luigi Vanvitelli*



${}^7\text{Be}(p,\gamma){}^8\text{B}$ - Solar Neutrino Flux



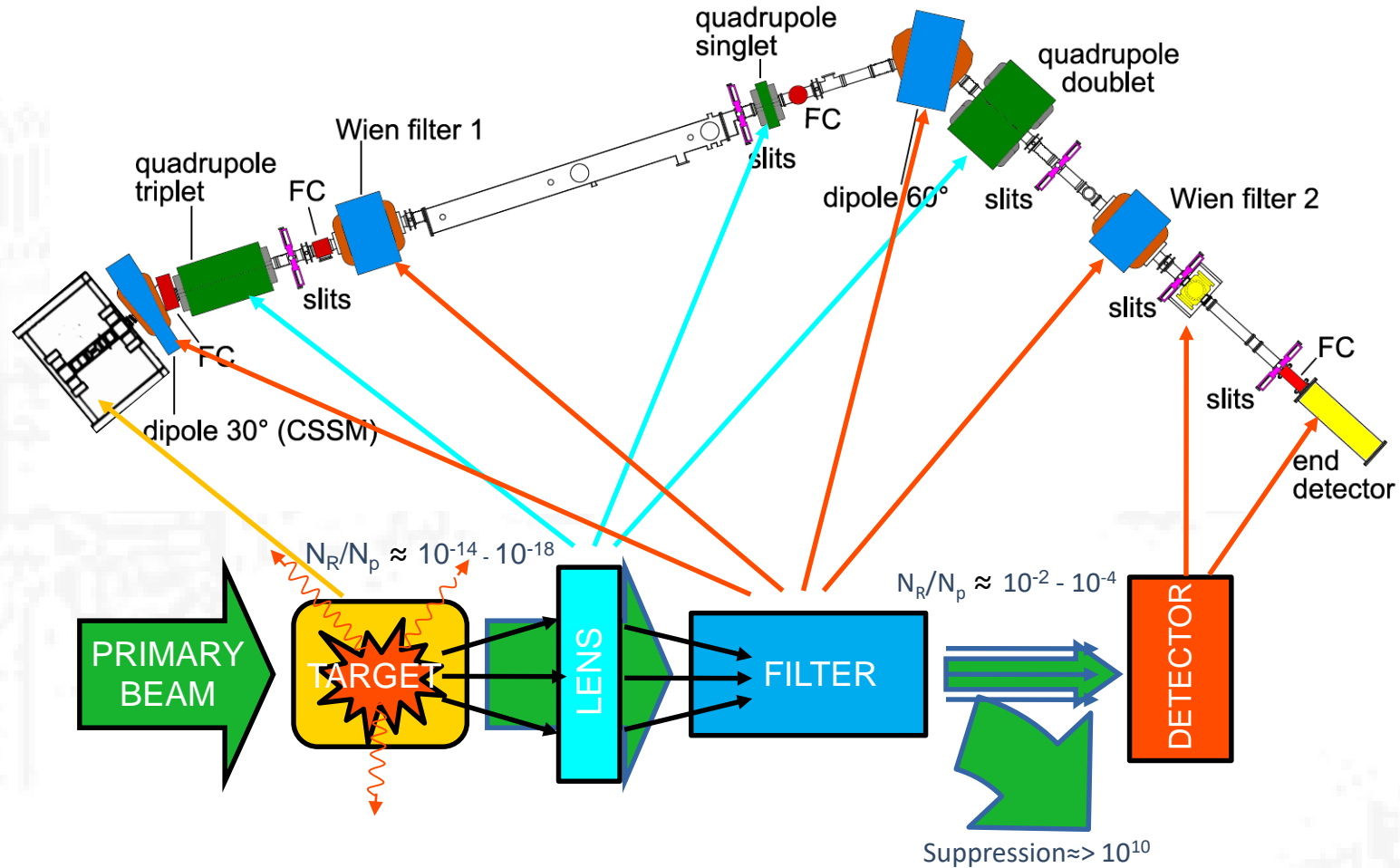
Large uncertainty on the ${}^8\text{B}$ component of predicted solar neutrino flux.



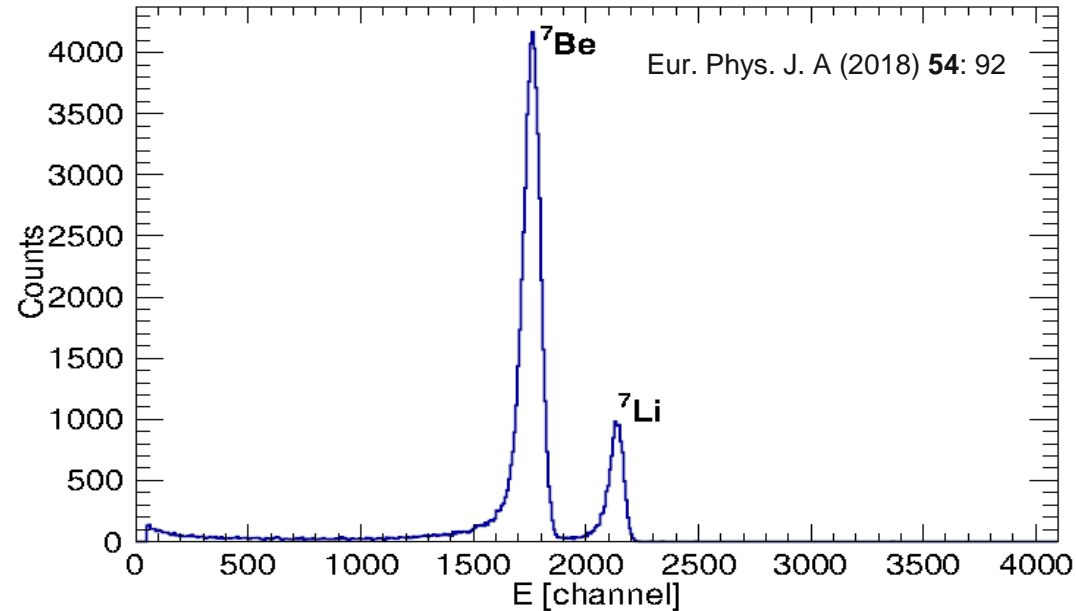
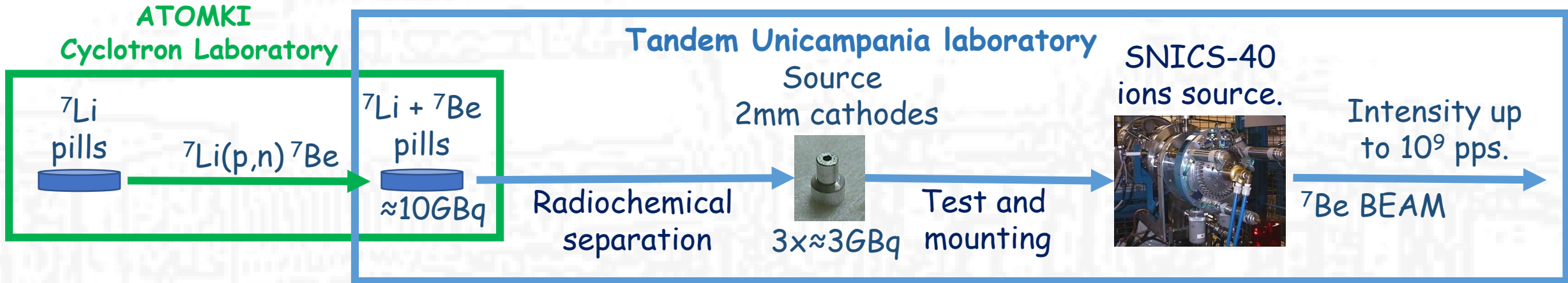
Adelberger et al. *Rev. Mod. Phys.* (2011)

Discrepancies between existing data sets limits the precision of the extrapolation to solar energy.

European Recoil for Nuclear Astrophysics - ERNA



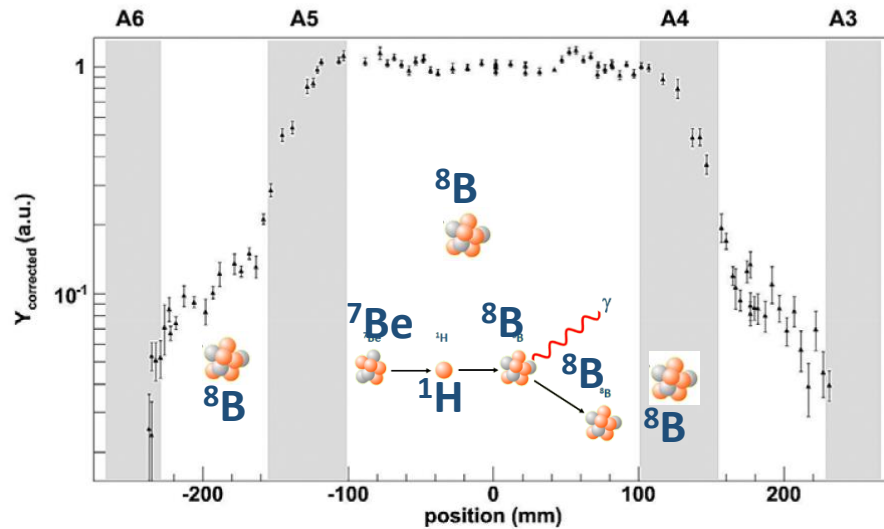
^7Be beam



The number of incident projectiles, including lithium contamination, is monitored on line through elastic scattering.

Windowless gas target

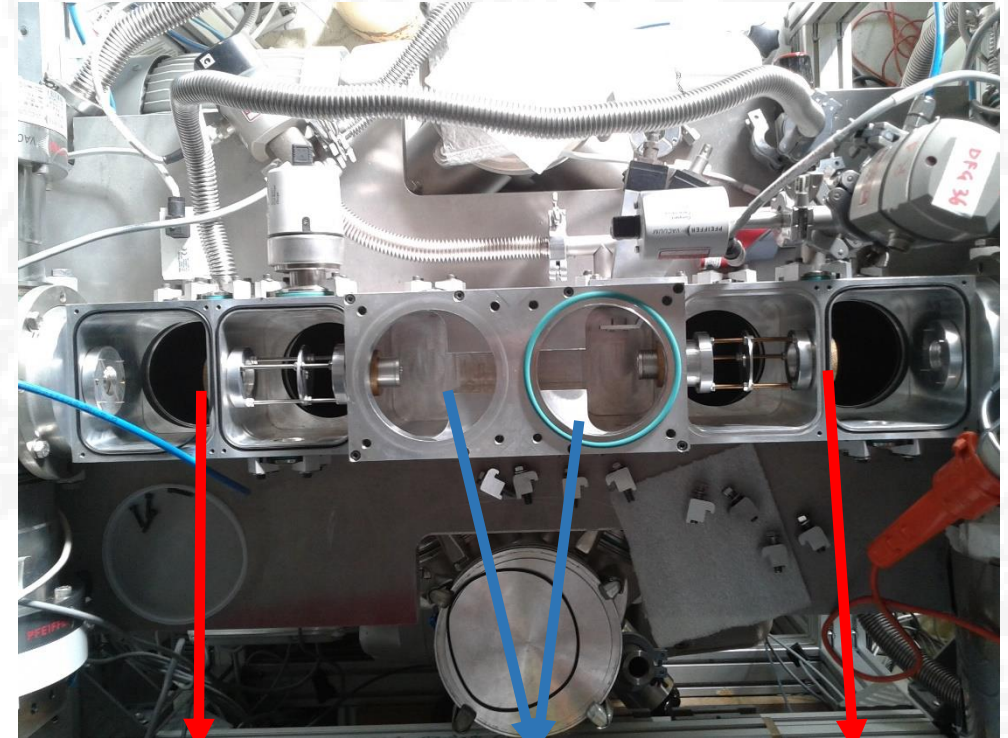
D. Schürmann et al., Eur. Phys. J. A (2013) 49: 80



Density profile of the gas target as seen in the yield of the 478 keV γ -ray line from the ${}^7\text{Li}(p, p'){}^7\text{Li}^*$.

Target density $n = 7.22 \pm 0.15 \cdot 10^{18}$ at/cm² at 4.9 mbar

The ${}^8\text{B}$ can be produced in different part along the gas target.



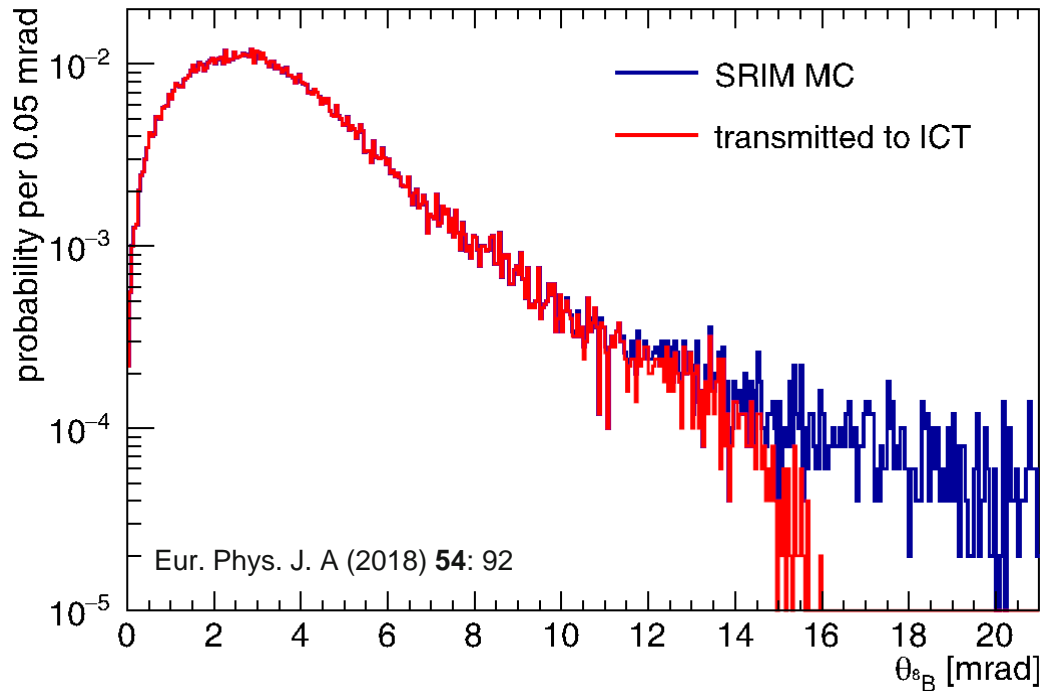
Ar

H₂

Ar

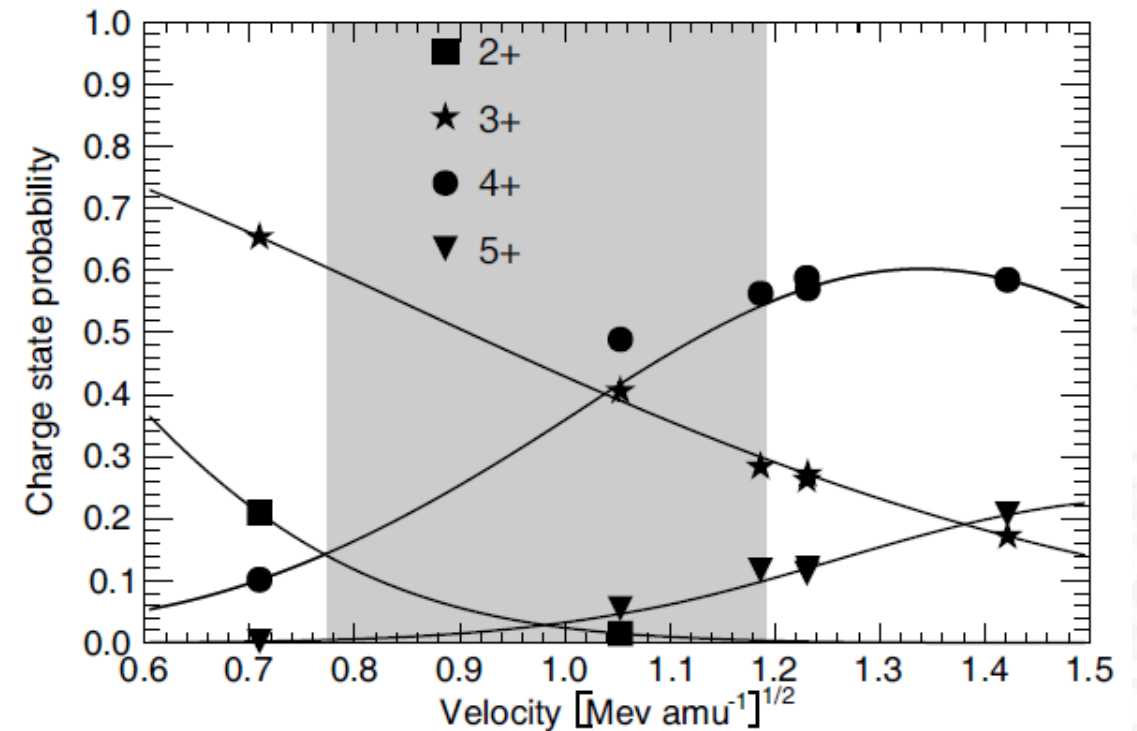
In all conditions full transmission to the end detector of the recoils in the selected charge state is mandatory.

Recoils transmission



Distribution of the ^8B recoils emerging from the target for $E_{\text{cm}} = 348$ keV (blue line).

The red line indicates the recoils reaching the end detector of ERNA using the **experimental acceptance curves**.



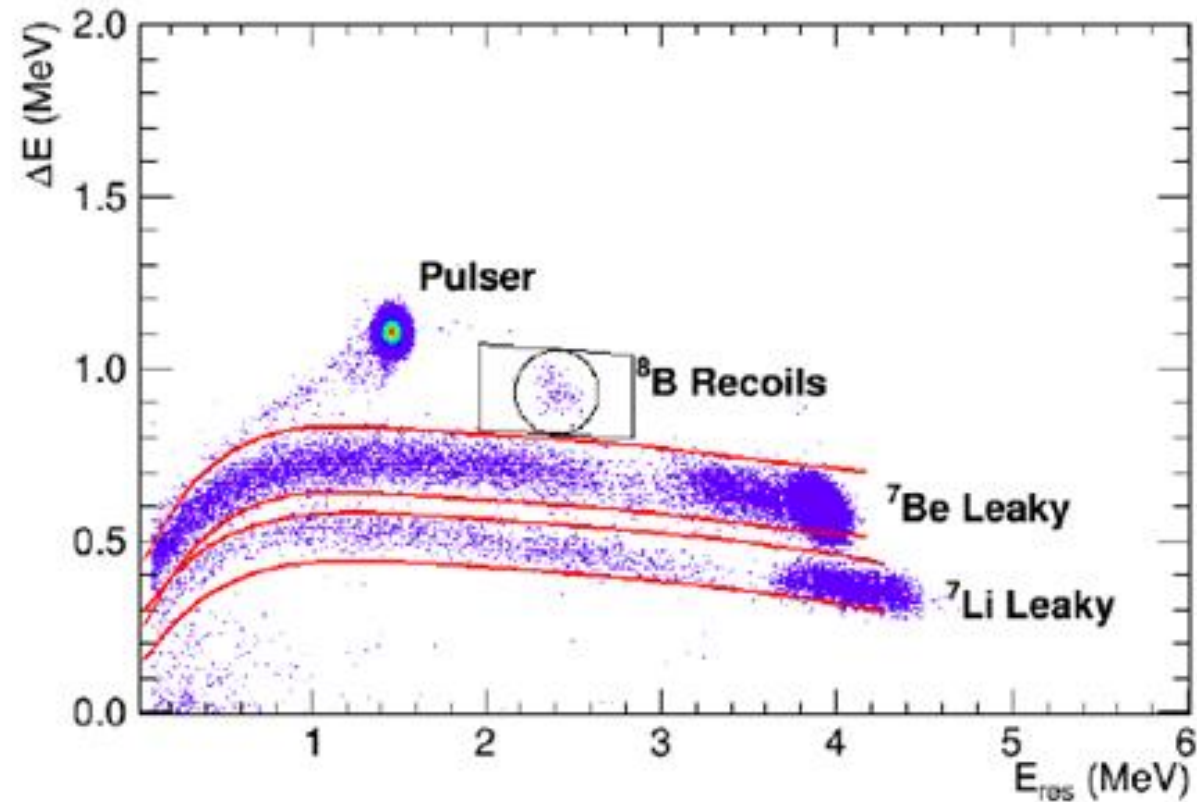
^8B ion charge state probability distribution as a function of the ion velocity.

R. Buompane et al., Eur. Phys. J. A (2018) 54: 92

Recoils detection

Physics Letters B 824 (2022) 136819

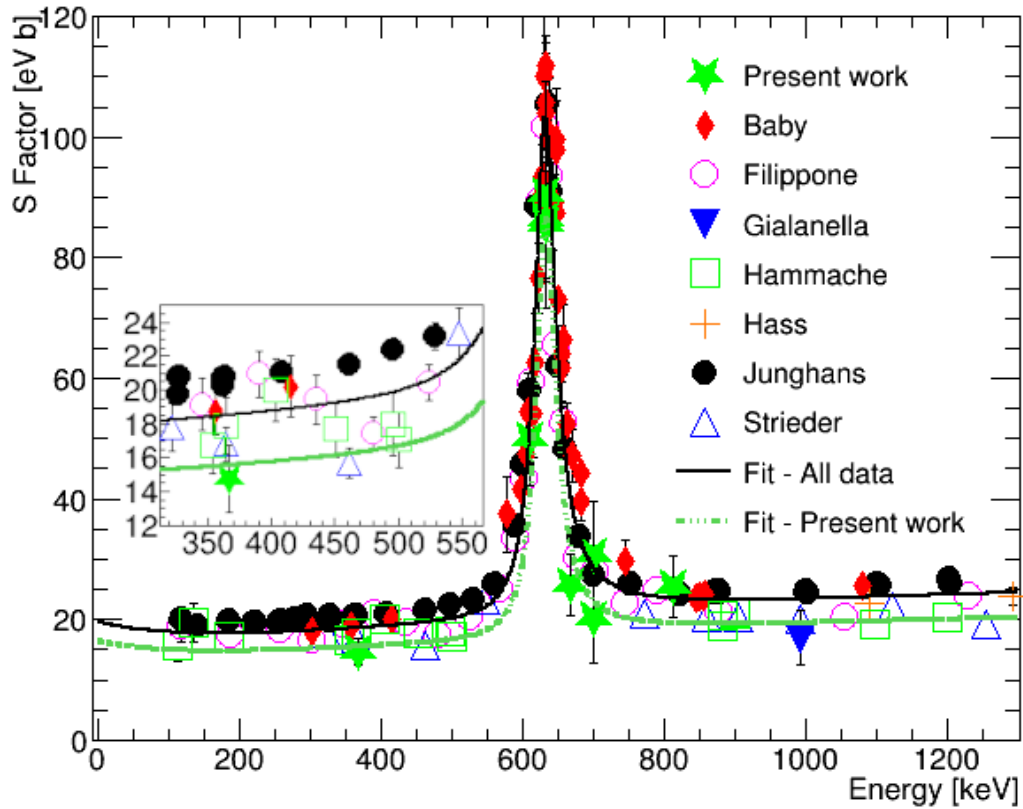
ΔE - E_{res} matrix of the end detector of ERNA for ${}^7\text{Be}(p;\gamma){}^8\text{B}$ at 632.4 keV. Both ${}^7\text{Li}$ and ${}^7\text{Be}$ "leaky" beam ions are visible with their long low energy tails.



The contour plots represent the predictions of a Monte Carlo simulation for the different ion species. The rectangular region around the elliptical area where the ${}^8\text{B}$ recoils are expected is used to estimate the background.

Results

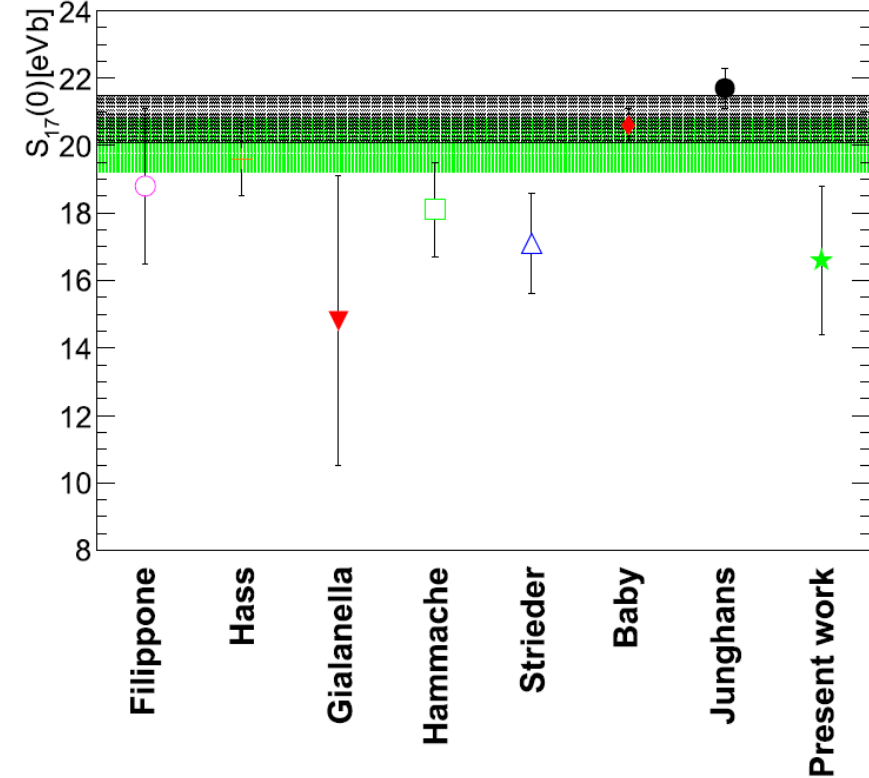
Physics Letters B 824 (2022) 136819



The solid green line represents the fit to the present work data alone $S_{17}(0) = 16.6 \pm 2.1$ eV b.

The solid black line (left) and the green shaded band (right) represents the result obtained including all data sets and considering the inflation of the systematic uncertainties to obtain a statistical compatibility $S_{17}(0) = 20.0 \pm 0.8$ eV b.

Physics Letters B 824 (2022) 136819

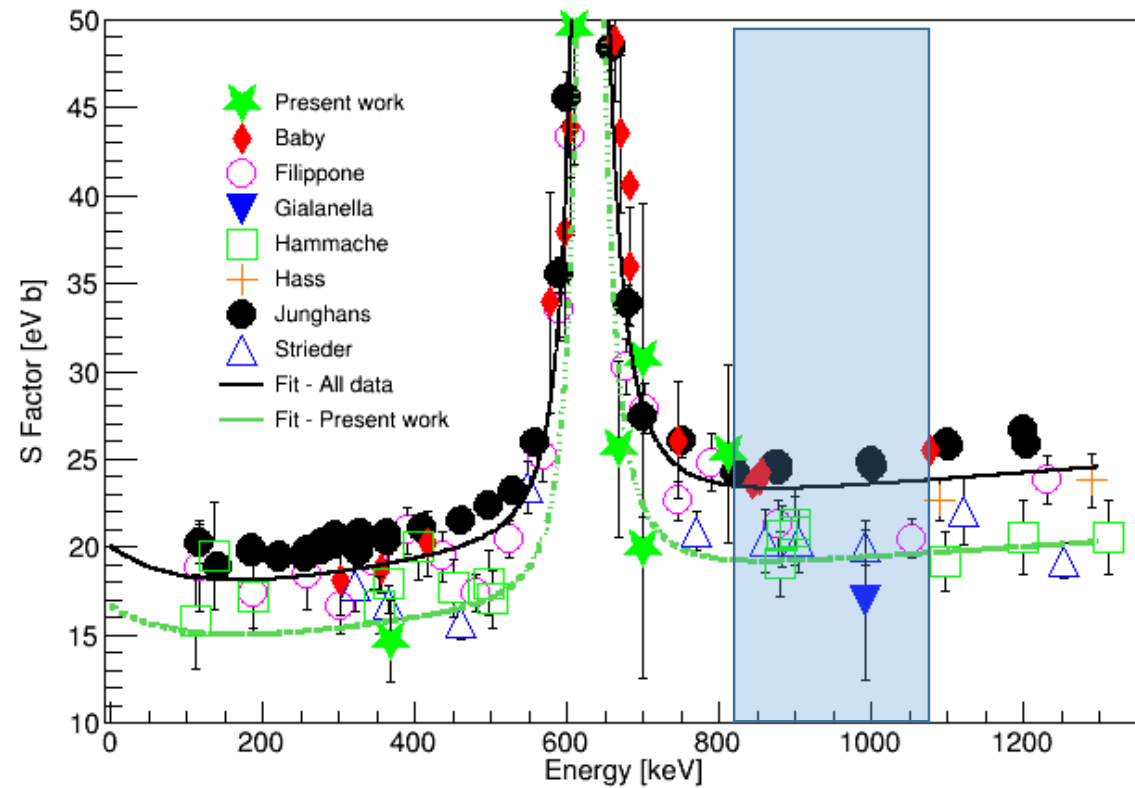


$S_{17}(0)$ values obtained fitting each data set, the error bars include systematic errors.

- Next -



A new solid stripper was installed at the tandem terminal.



The solid stripper increase the probability of higher charge state at terminal and than higher energy range will be explored.

Conclusion

- A very intense ${}^7\text{Be}$ beam, up to 10^9 pps, is routinely produced and characterized at CIRCE laboratory @ Unacampania;
- The first significant total cross section measurement in inverse kinematic and using a recoil separator of the ${}^7\text{Be}(p,\gamma){}^8\text{B}$ reaction, from $E_{\text{cm}} = 366.9$ keV up to the $E_{\text{cm}} = 812.3$ keV, has been performed;
- A fit to our data alone yields $S_{17}(0) = 16.6 \pm 2.1$ eV b, this value are compatible only with a part of previous measurements, in particular those indicating a low value of the astrophysical S-factor at zero energy $S_{17}(0)$;
- A global fit including all data sets and considering the inflation of the systematic uncertainties provides an estimate $S_{17}(0) = 20.0 \pm 0.8$ eV b that we suggest here.
- New measurements at energy $E_{\text{cm}} > 1$ MeV are planned for the next year.

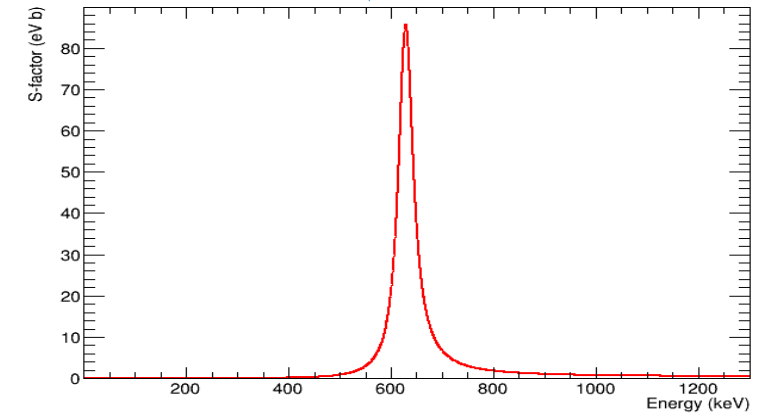
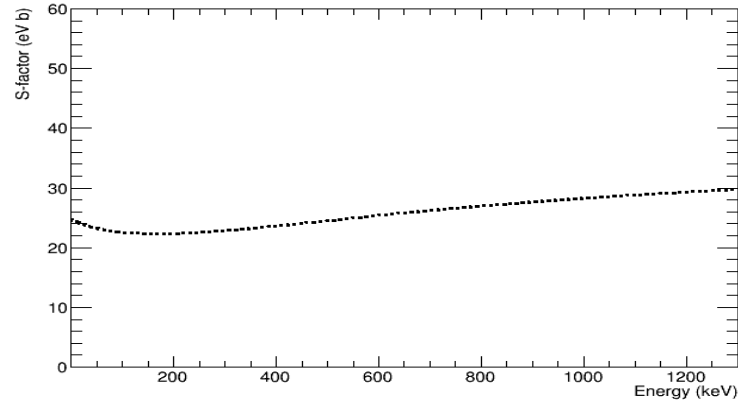
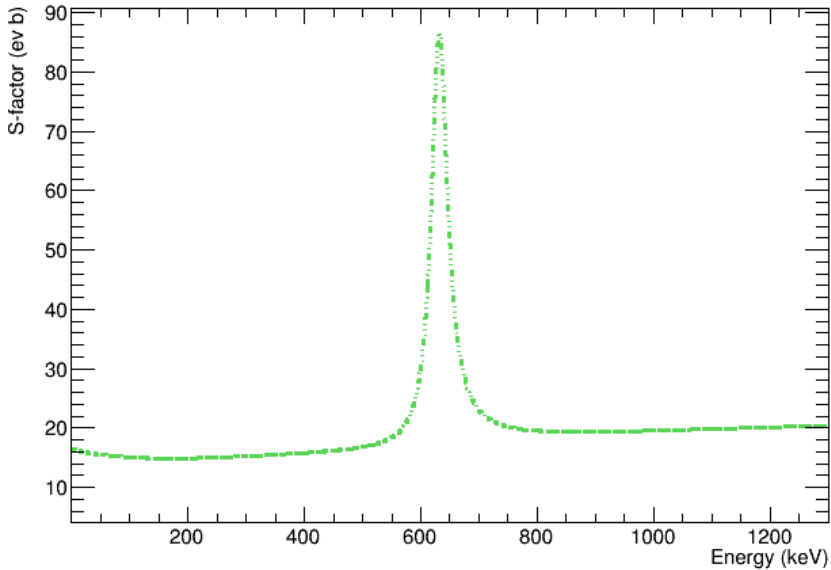


Thanks

NPA X, 5th-9th September 2022, CERN

Fit function

$$S_{\text{factor}}(E) = A \cdot S_{\text{factor}}(\text{MN})(E) + \pi\lambda^2 \cdot \frac{2J + 1}{(2J_p + 1) \cdot (2J_t + 1)} \cdot \frac{\Gamma_p(E) \cdot \Gamma_\gamma(E)}{(E - E_{\text{res}})^2 + (\Gamma_{\text{tot}}(E)/2)^2} \cdot E \cdot e^{2\pi\eta(E)}$$



P. Descouvemont, PRC 70, 065802 (2004)

The fit function is obtained summing the non resonant component S_{MN} from Descouvemont et al. 2004 and a resonant component.

The fit function included two parameters: A for the non resonant component and Γ_p .

The E_{res} and Γ_γ are fixed to the literature value from Tilley et al., 2004, NPA, 745(3), respectively of 632 keV and 25.2 meV.

Tandem Laboratory @ Unicampania

