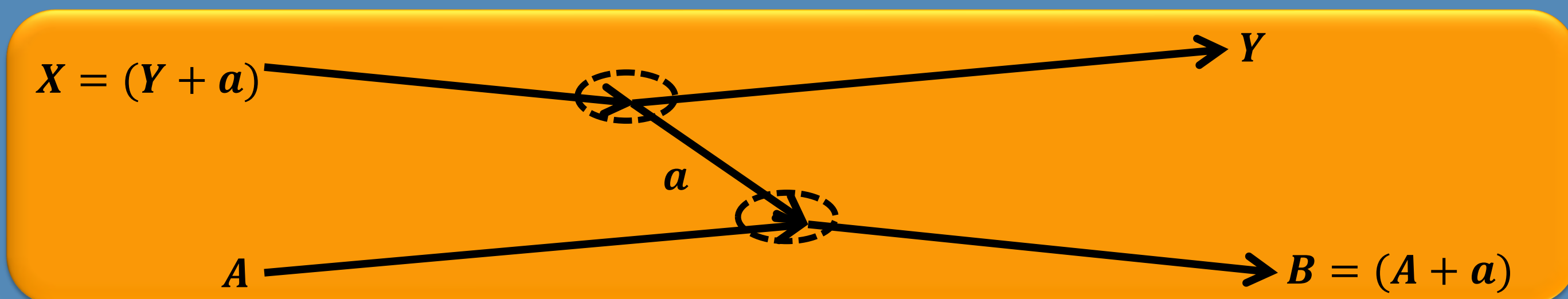


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ASYMPTOTIC NORMALIZATION COEFFICIENT (ANC) METHOD

The Asymptotic Normalization Coefficient (ANC) method is widely used to gain information about **DIRECT RADIATIVE CAPTURE** $A + a \rightarrow Y + \gamma$ from a suitable transfer reaction $A + X \rightarrow Y + B$, being $X = (Y + A)$ and $B = (A + a)$, under the condition that the second one is **PERIPHERAL** [1,2,3]



Under such conditions, in DWBA the cross section can be approximated in terms of the so-called single-particle ANC (that depends on the shape of the potential wells involved for the B and X nuclei) b^2 and the proper ANC c^2

$$\frac{d\sigma}{d\Omega} = \sum_{j_B, j_X} (C_{Aa, l_B, j_B}^B)^2 (C_{Ya, l_X, j_X}^X)^2 \frac{\sigma_{l_B, j_B, l_X, j_X}^{DWBA}}{b_{Aa, l_B, j_B}^2 b_{Ya, l_X, j_X}^2}$$

Here, c^2 must have a small dependence from the potential, while the ratio above is nearly independent from b^2

EXTENTION FOR MIRROR NUCLEI

Proton and neutron channel ANC of the g.s. for **MIRROR PAIR NUCLEI** are related [4,5]

$$(c_p)^2 = R (c_n)^2, \text{ with } R = \left| \frac{F_l(ik_p R_N)}{k_p R_N j_l(ik_p R_N)} \right|^2$$

- F_l : regular Coulomb wave function
- j_l : Bessel function
- R_N : radius of the nuclear interior
- $k_{p(n)} = \left(\frac{2\mu\epsilon}{\hbar^2} \right)^{1/2}$

Also, for **EXCITED STATES** [6]:

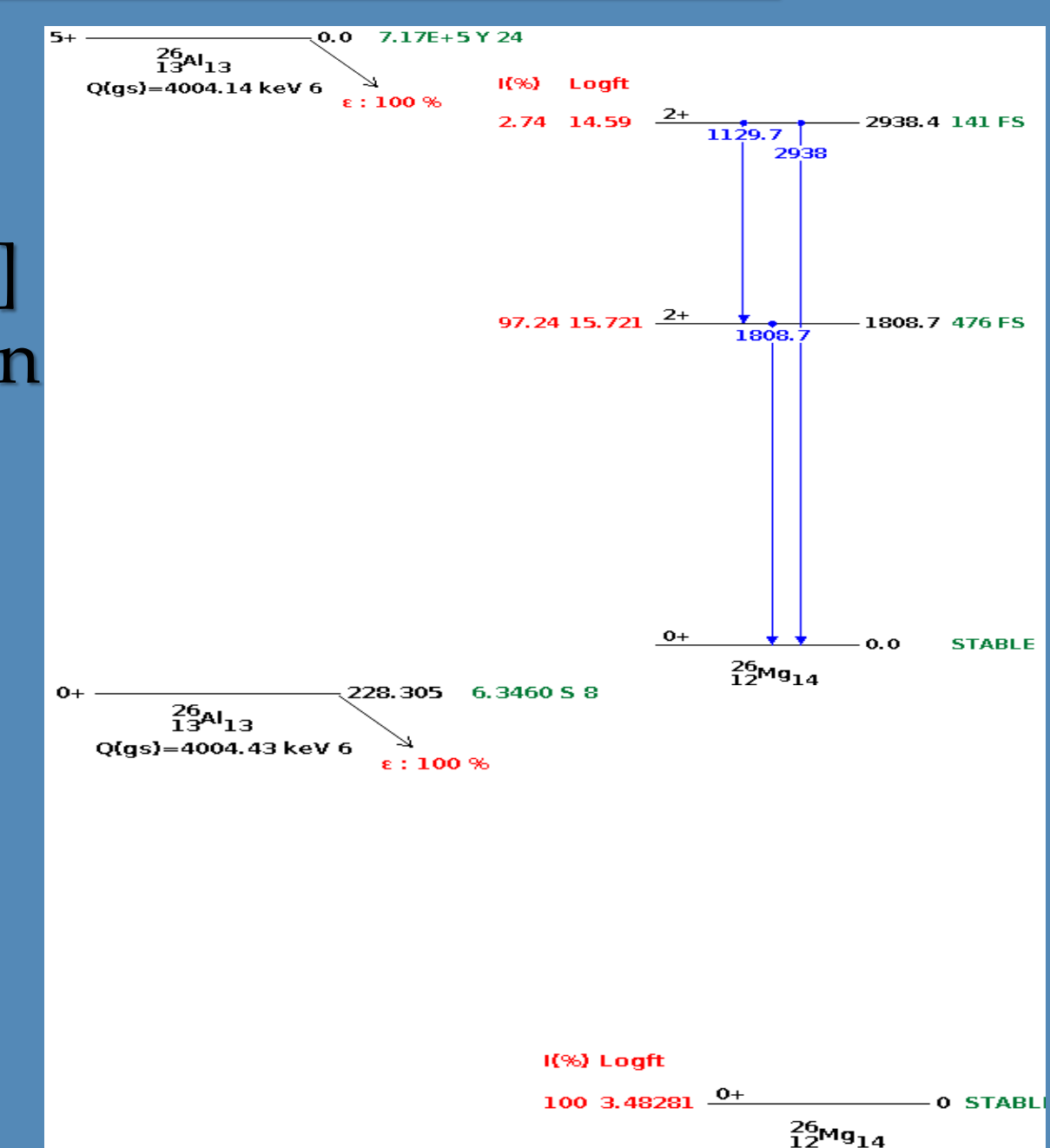
$$\frac{\Gamma_p}{|c_n|^2} = R_\Gamma \approx R_0^{res} = \frac{\kappa_p}{\mu} \left| \frac{F_l(\kappa_p R_N)}{\kappa_p R_N j_l(ik_p R_N)} \right|^2$$

ASTROPHYSICAL SCENARIO

^{26}Al can be produced in **Core-Collapse Supernovae, Wolf-Rajet Objects and AGB Stars** [7], **Novae** [8] and **X-ray bursts** [9] Its production can occur via the reaction chain $^{24}\text{Mg}(p, \gamma)^{25}\text{Al}(\beta^+)^{25}\text{Mg}(p, \gamma)^{26}\text{Al}$.

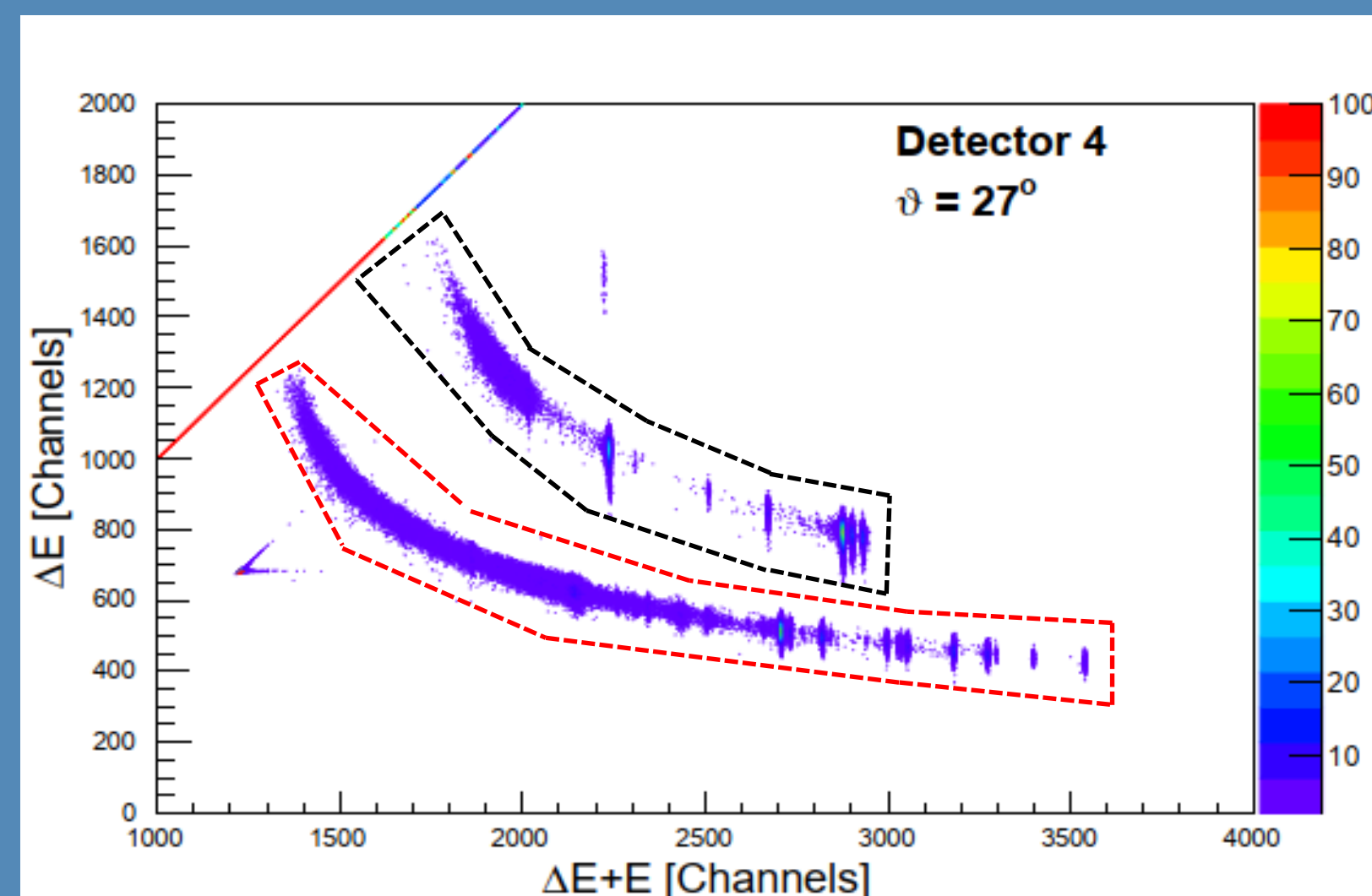
Production hampered by the isomer $^{26}\text{Al}^m$ ($T_{1/2} = 6.35$ sec).

The $^{25}\text{Al}(p, \gamma)^{26}\text{Si}(\beta^+)^{26}\text{Al}^m$, bypasses ^{26}Al production, and directly feed its isomer synthesis. ^{26}Si can be depleted by the $^{26}\text{Si}(p, \gamma)^{27}\text{P}$ reaction, interfering with both the ground state and isomeric state creation.



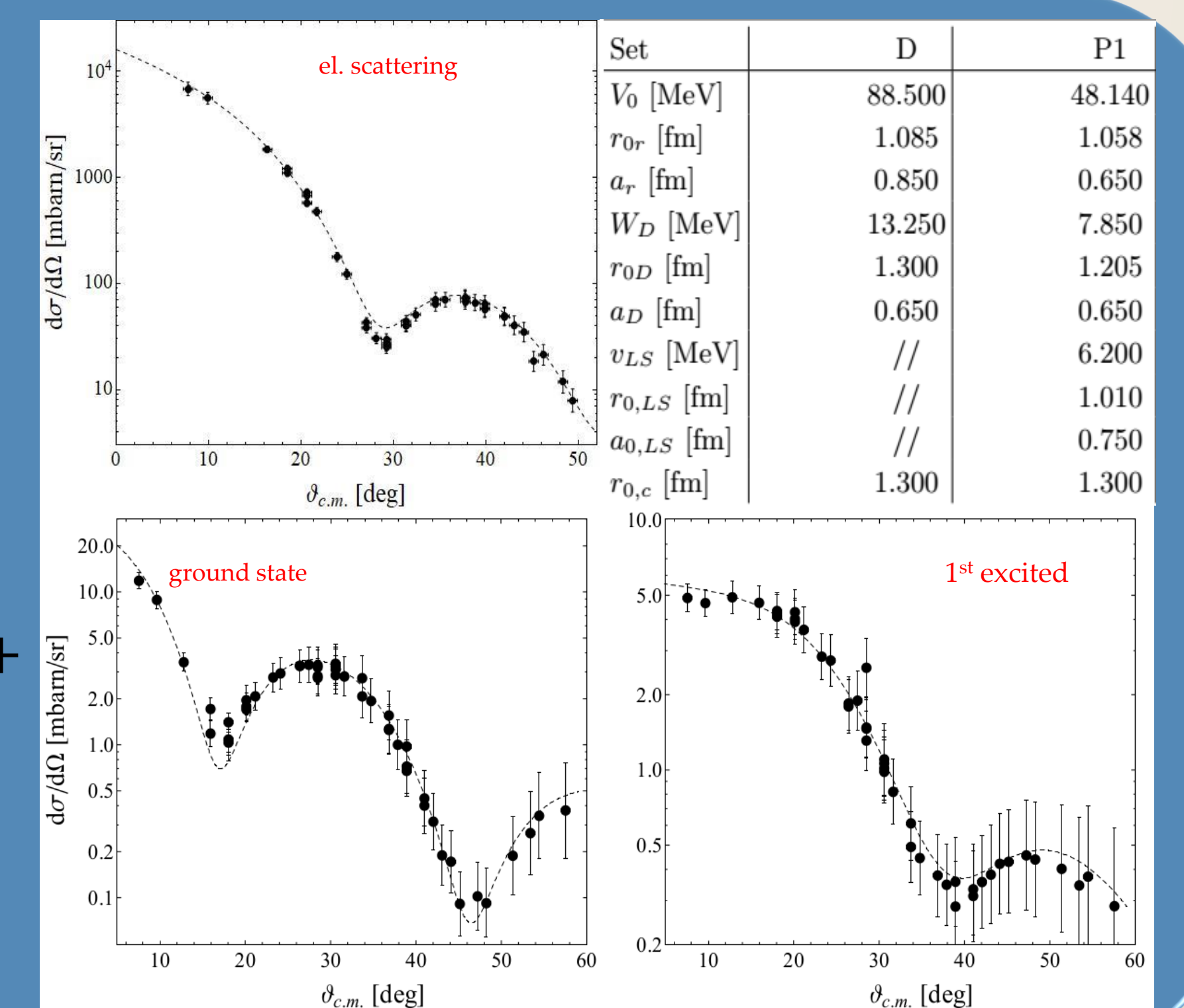
EXPERIMENTAL SET-UP AND OMP EXTRACTION

The ANC method has been applied on the $^{26}\text{Mg}(d, p)^{27}\text{Mg}$, that than has been used to retrieve information on the $^{26}\text{Si}(p, \gamma)^{27}\text{P}$. Protons and deuterons coming from the reaction have been detected and identified by means of the ΔE - E technique [10]. Such particles are then used to extract the Optical Models Parameters (OMP) for the entrance ($^{26}\text{Mg} + d$) and exit ($^{26}\text{Mg} + n$) channel respectively



In **Distorted Wave Born Approximation (DWBA, using FRESCO)** formalism [10]:

$$U = V_c(r_c) - V_0 f(x_0) - \left[W f(x_w) - 4W_D \frac{d}{dr} f(x_D) \right] + \frac{\hbar^2}{m_\pi c} V_{LS}(L\sigma) \frac{1}{r} \frac{d}{dr} f(x_{LS})$$

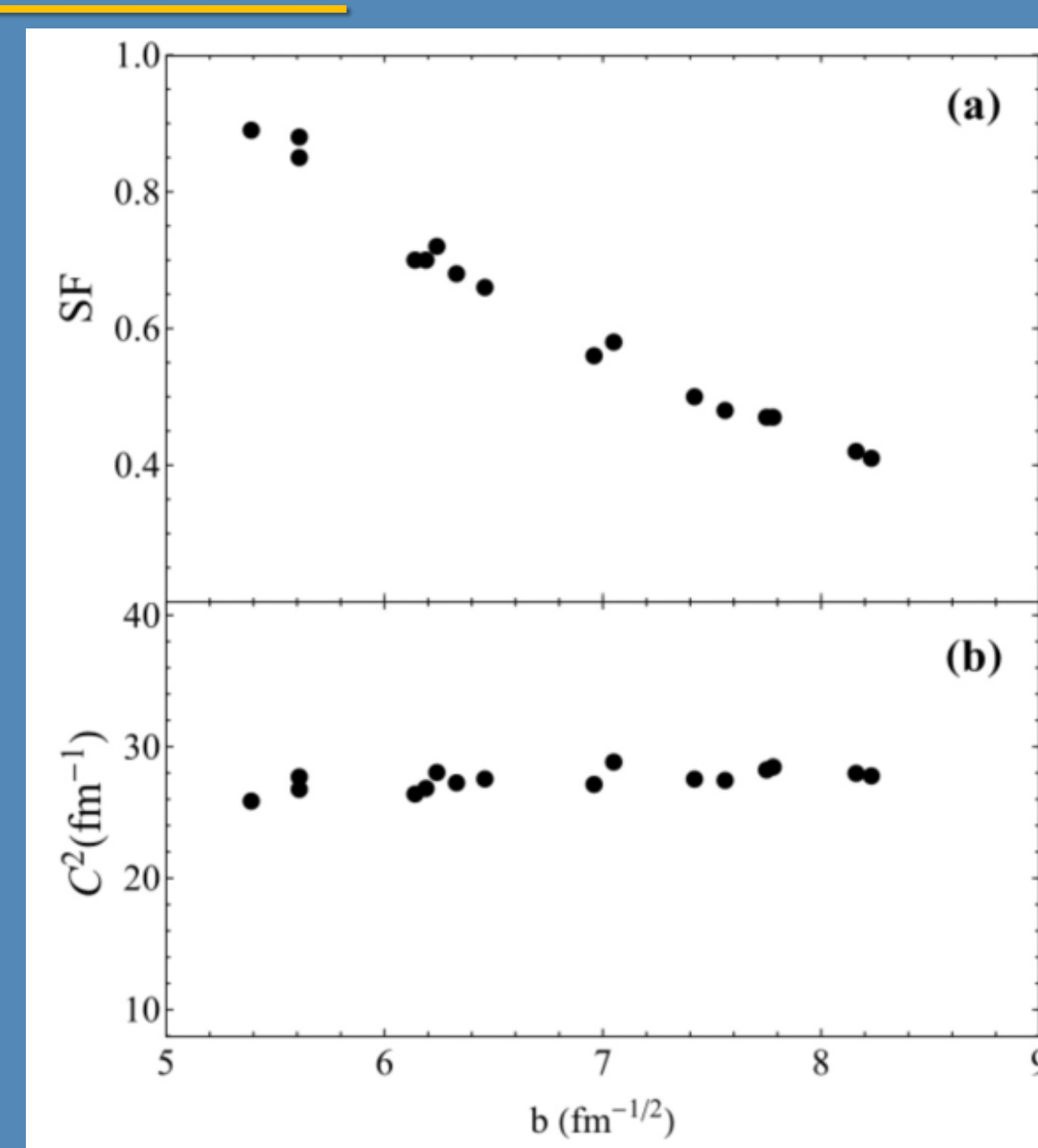


PERIPHERALITY

To check the peripherality one can change r_0 and a parameters to see the behaviour of ANC and S-factor with respect to the SPANC [10]:

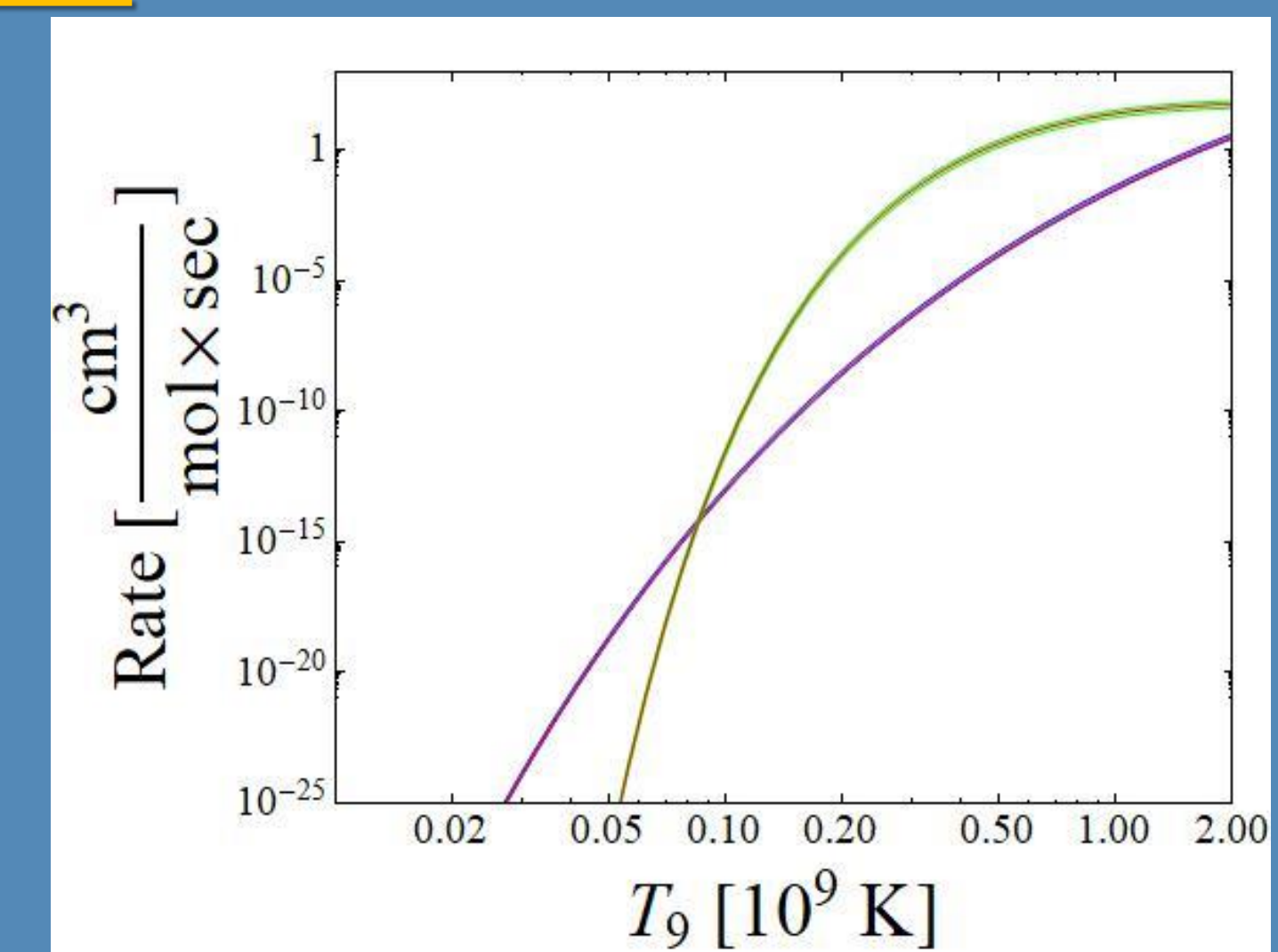
For $1.1 \leq r_0 \leq 1.4$ fm and $0.5 \leq a \leq 0.7$ fm the experimental spectroscopic factor shows a variation of $\approx 55\%$, while the ANC shows a variation of 13%.

THE REACTION CAN THEREFORE BE CONSIDERED PERIPHERAL!



RESULTS

Using the parametrization above the ANC for the two states has been calculated in $C_{gr}^2 = 28.26 \pm 4.24$ fm⁻¹ and $C_{1st}^2 = 1.51 \pm 0.27$ fm⁻¹, (in agreement with [11]). Finally, using the mirror nuclei procedure, the results have been used to extract the ANC for the radiative capture in the ground state ($C_{gr}^2 = 1420 \pm 255$ fm⁻¹), and $\Gamma_p = (6.76 \pm 1.35) \times 10^{-9}$ MeV for the first excited state. Such value have been used to calculate the reaction rate (blue for the non-resonant and green for the resonant contributions), using the ratio Γ_γ/Γ_p reported in [12].



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