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New detection systems for an enhanced sensitivity in key stellar (n, γ) measurements

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Neutron capture cross-section measurements are fundamental in the study of astrophysical phenomena, such as the slow neutron capture (s-) process of nucleosynthesis operating in red-giant and massive stars. One of the best suited methods to measure neutron capture (n, γ) cross sections over the full stellar range of interest is the time-of-flight (TOF) technique.

TOF neutron capture measurements on key s-process branching isotopes are very challenging due to the limited mass (\sim mg) available and the high experimental background arising from the sample activity and the dominant neutron scattering cross section. As a consequence of these challenges, only five out of the 21 key s-process isotopes (^{63}Ni , ^{151}Sm , ^{171}Tm , ^{147}Pm y ^{204}Tl) have been measured by means of the time-of-flight technique to date. Overcoming the current experimental limitations requires the combination of facilities with high instantaneous flux, such as n_TOF, with detection systems with an enhanced detection sensitivity and high counting rate capabilities.

This contribution will review some of the latest developments in detection systems for (n, γ) measurements at n_TOF. The focus will be set on i-TED, an innovative detection system which exploits the Compton imaging technique to reduce the dominant neutron scattering background. The discussion will be illustrated with preliminary results of the first measurement of the s-process branching-point reaction $^{79}\text{Se}(n,\gamma)$.

Field of work

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