FERMION MODEL

Yu Hang Ng

Department of Physics and Astronomy Univeristy of Nebraska-Lincoln

Particle Physics on the Plains 2019, University of Kansas October 12, 2019

► Can the global and gauge symmetries of elementary particle physics that are broken at zero temperature be restored at high temperature?

Introduction

- ► Can the global and gauge symmetries of elementary particle physics that are broken at zero temperature be restored at high temperature?
 - ► Answered by Kirzhnits & Linde (1972) and Weinberg (1974).

Introduction

- ► Can the global and gauge symmetries of elementary particle physics that are broken at zero temperature be restored at high temperature?
 - ► Answered by Kirzhnits & Linde (1972) and Weinberg (1974).

FERMION MODEL

► Weinberg's (1974) paper also showed that not all broken symmetries at zero temperature are restored at high temperature. Particularly, he showed the existence of

Introduction

- ► Can the global and gauge symmetries of elementary particle physics that are broken at zero temperature be restored at high temperature?
 - ► Answered by Kirzhnits & Linde (1972) and Weinberg (1974).

- ► Weinberg's (1974) paper also showed that not all broken symmetries at zero temperature are restored at high temperature. Particularly, he showed the existence of
 - ► Symmetry Non-Restoration (SNR): Symmetry that is broken at lower temperature is never stored at higher temperature
 - ► Inverse Symmetry Breaking (ISB): Symmetry that is unbroken at lower temperature becomes broken at higher temperature

► Meade & Ramani (arXiv:1807.07578) showed that Electroweak (EW) symmetry will never be restored or only temporarily restored (TR) at higher temperature with a simple scalar extension of SM.

FERMION MODEL

► Is it possible to achive SNR or TR for EW symmetry by adding new fermions to SM?

SCALAR MODEL WITH $O(N_s)$ GLOBAL SYMMETRY

► Extend SM by adding SM-singlet scalar fields s_i with $O(N_s)$ global symmetry

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{2}\mu_s^2(s_is_i) - \frac{1}{4}\lambda_s(s_is_i)^2 - \frac{1}{2}\lambda_{hs}h^2(s_is_i)$$

FERMION MODEL

► The leading term in temperature of thermal mass correction of *h* is

$$\Pi_h = \frac{\partial^2 V_{th}}{\partial h^2} = T^2 \left(\frac{3}{16} g^2 + \frac{1}{16} g'^2 + \frac{1}{4} \lambda_t^2 + \frac{1}{2} \lambda + \frac{N_s}{12} \lambda_{hs} \right)$$

► SNR can be achieved by choosing λ_{hs} < 0 and large N_s .

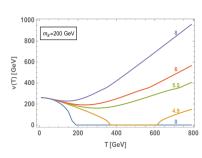
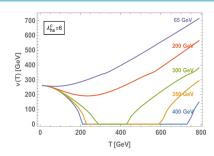


Figure: Temperature dependent vev for different values of $N_s |\lambda_{hs}|$.

► Temporary restoration phase becomes longer as $N_s|\lambda_{hs}|$ decreases.

Meade & Ramani, arXiv:1807.07578



FERMION MODEL

Figure: Temperature dependent vev for different values of m_s .

► Temporary restoration phase becomes longer as m_s increases.

A TOY MODEL WITH NEW FERMIONS

$$\mathcal{L} = \mathcal{L}_{\mathit{SM}} + \sum_{i=1}^{N_{
u}} \mathcal{L}_{D+M}^i \; , H = egin{bmatrix} G^+ \ (h+iG^0)/\sqrt{2} \end{bmatrix}$$

Each \mathcal{L}_{D+M}^{i} has the same form:

$$\mathcal{L}_{D+M} = -y_{\nu} \left(\overline{\nu_R} \widetilde{H}^{\dagger} L_L + \overline{L_L} \widetilde{H} \nu_R \right) + \frac{1}{2} m_R \left(\nu_R^T \mathcal{C}^{\dagger} \nu_R + \nu_R^{\dagger} \mathcal{C} \nu_R^* \right)$$

$$\supset -m_D(h) \left(\overline{\nu_R} \nu_L + \overline{\nu_L} \nu_R \right) + \frac{1}{2} m_R \left(\nu_R^T \mathcal{C}^{\dagger} \nu_R + \nu_R^{\dagger} \mathcal{C} \nu_R^* \right)$$

$$\supset \frac{1}{2} N^T \mathcal{C}^{\dagger} M N + h.c.$$

with

$$m_D(h) = \frac{y_\nu h}{\sqrt{2}}, \quad M = \begin{bmatrix} 0 & m_D \\ m_D & m_R \end{bmatrix}, \quad N = \begin{bmatrix} \nu_L \\ \nu_R^C \end{bmatrix}$$
 (1)

M can be diagonalized by a unitary transformation U:

$$U^{T}MU = \begin{bmatrix} m_1 & 0 \\ 0 & m_2 \end{bmatrix}, \ N = U \begin{bmatrix} \nu_1 \\ \nu_2 \end{bmatrix}$$

FERMION MODEL

0000

$$\mathcal{L}_{D+M} \supset \frac{1}{2} \sum_{k=1,2} m_k \nu_k^T C^{\dagger} \nu_k + h.c.$$

where physical masses are:

$$m_{1,2} = \frac{1}{2} \left[m_R \mp \sqrt{m_R^2 + 4m_D^2} \right]$$

FERMION MODEL

$$V_{th,\nu} = -N_{\nu} \frac{T^4}{2\pi^2} \sum_{k=1,2} J_F \left(\frac{m_k^2}{T^2}\right)$$

with thermal function

INTRODUCTION

$$J_F(y^2) = \int_0^\infty dx \, x^2 \log \left[1 + \exp\left(-\sqrt{x^2 + y^2}\right) \right]$$

The leading term in temperature of thermal mass correction of h is

$$\Pi_h = \frac{\partial^2 V_{th}}{\partial h^2} = T^2 \left(\frac{3}{16} g^2 + \frac{1}{16} g'^2 + \frac{1}{4} \lambda_t^2 + \frac{1}{4} \lambda + \frac{N_\nu}{24} y_\nu^2 \right)$$

- choose y_{ν} to be purely imaginary in order to have maximum negative mass correction from the new fermions.
- ► When $N|y_{\nu}|^2$ is large enough, the negative contribution outweight the usual positive contributions and SNR can be achieved at high temperature

$$\begin{split} &\delta m_h^2(h,T) \\ &= \sum_i \frac{\partial}{\partial h} \left[\frac{\partial V_{CW}^i}{\partial h} \left(m_i^2(h) + \delta m_i^2(h,T) \right) + \frac{\partial V_{th}^i}{\partial h} \left(m_i^2(h) + \delta m_i^2(h,T),T \right) \right] \end{split}$$

FERMION MODEL

0000

The effective potential is

INTRODUCTION

$$V_{\text{eff}}^{\text{pd}}(h,T) = V_0 + \sum_{i} \int_0^h dh' \left[\frac{\partial V_{CW}^i}{\partial h'} \left(m_i^2(h') + \delta m_i^2(h',T) \right) + \frac{\partial V_{th}^i}{\partial h'} \left(m_i^2(h') + \delta m_i^2(h',T),T \right) \right]$$

► The effective potential is calculated using *Partial Dressing* method, which resum the important higher-loop contributions (i.e. daisy and super-daisy diagrams) to thermal mass correctly.

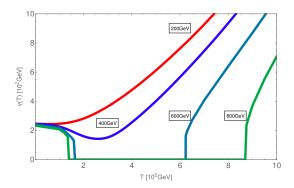
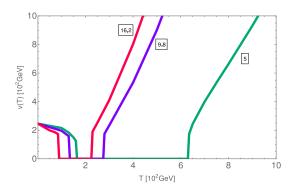


Figure: Temperature dependent vev for different values of m_R .

▶ Temporary restored period becomes longer as m_R increase. As $T < m_R$ and m_R is much heavier than SM particles, the thermal mass correction from the new fermions becomes less important.



FERMION MODEL

0

Figure: Temperature dependent vev for different values of $N_{\nu}|y_{\nu}|^2$.

► Temporary restored period becomes longer as $N_{\nu}|y_{\nu}|^2$ decrease as shown above.

SUMMARY AND OUTLOOK

SNR and TR can be achieved by adding new fermions to SM. The effect is stronger when $N_{\nu}|y_{\nu}|^2$ or m_R is larger.

- ► Find a realistic model with new fermions
- ► Investigate the cosmological implications (thermal evolution of universe, gravitational wave signal etc.) and collider signals.

DAISY AND SUPER-DAISY CONTRIBUTIONS TO THERMAL MASS

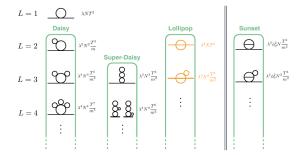


Figure 2. Complete set of 1- and 2- loop contributions to the scalar mass, as well as the most important higher loop contributions, in ϕ^4 theory. The scaling of each diagram in the high-temperature approximation is indicated, omitting symmetry- and loop-factors. Diagrams to the right of the vertical doubtle-lines only contribute away from the origin when $\langle \phi \rangle = \phi_0 > 0$. We do not show contributions which trivially descend from e.g. loop-corrected quartic couplings. Lollipop diagrams (in orange) are not automatically included in the resummed one-loop rotential.

EFFECTIVE POTENTIAL

► Tree-level potential:

$$V_0 = -\frac{1}{2}\mu^2 h^2 + \frac{1}{4}\lambda h^4$$

► Coleman-Weinberg potential (for *i*-th particle)

$$V_{CW}^{i} = (-1)^{F} g_{i} \frac{m_{i}^{4}}{64\pi^{2}} \left[\log \left(\frac{m_{i}^{2}}{\mu_{R}^{2}} \right) - c_{i} \right]$$

► One-loop thermal potential (for *i*-th particle)

$$V_{th}^{i} = (-1)^{F} g_{i} \frac{T^{4}}{2\pi^{2}} J_{B/F} \left(\frac{m_{i}^{2}}{T^{2}}\right)$$

where

$$J_B/F(y^2) = \int_0^\infty dx \, x^2 \log \left[1 \mp \exp\left(-\sqrt{x^2 + y^2}\right) \right]$$

► Effective potential

$$V_{ extit{eff}} = V_0 + \sum_i \left(V_{CW}^i + V_{ extit{th}}^i
ight)$$