



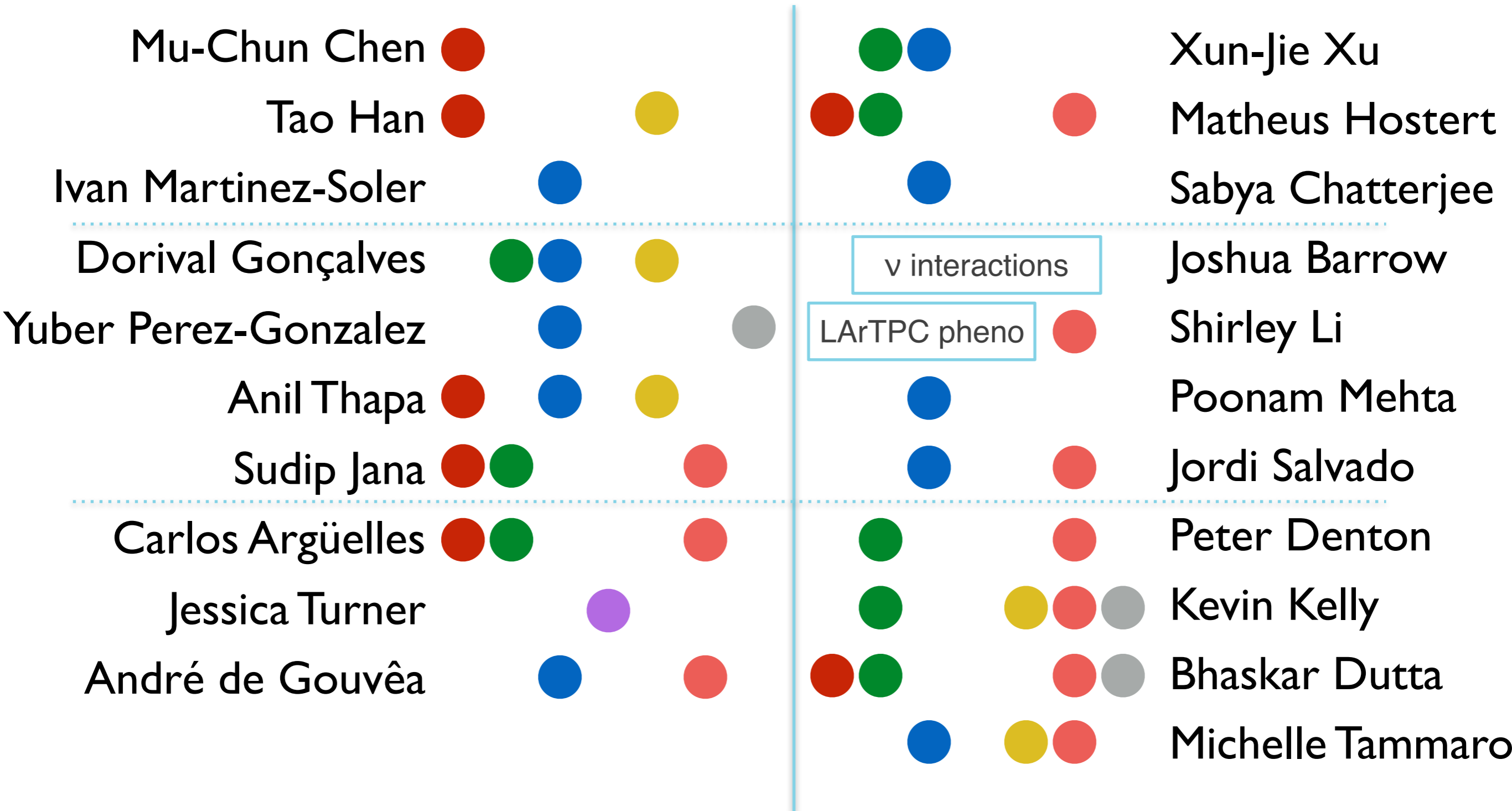
## NTN NSI Workshop Summary

*also known as "Bhupal's workshop"*

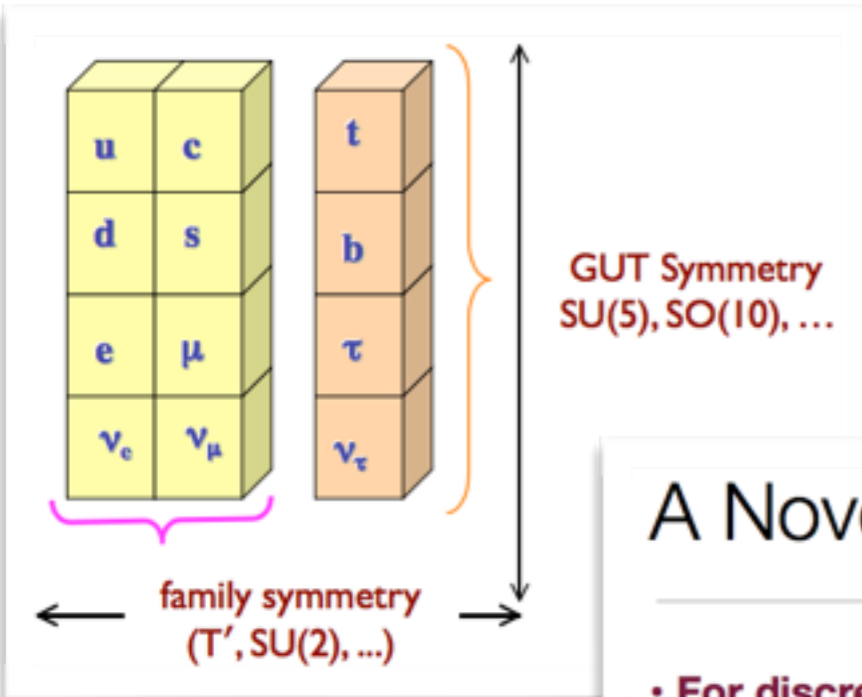
Pedro A. N. Machado

May 2019

**Mass and flavor**    **Traditional NSI**    **Colliders**    **Dark matter**  
**Light mediators**    **Leptogenesis**    **New signatures**

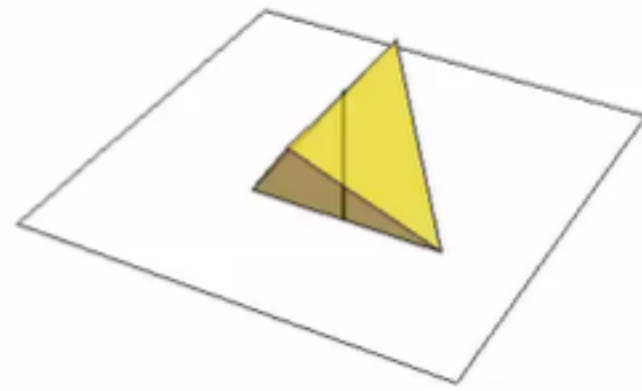
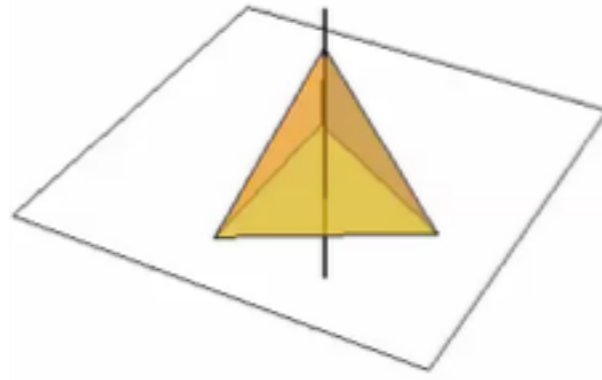


# Mass and flavor



T: (1234) → (2314)

S: (1234) → (4321)

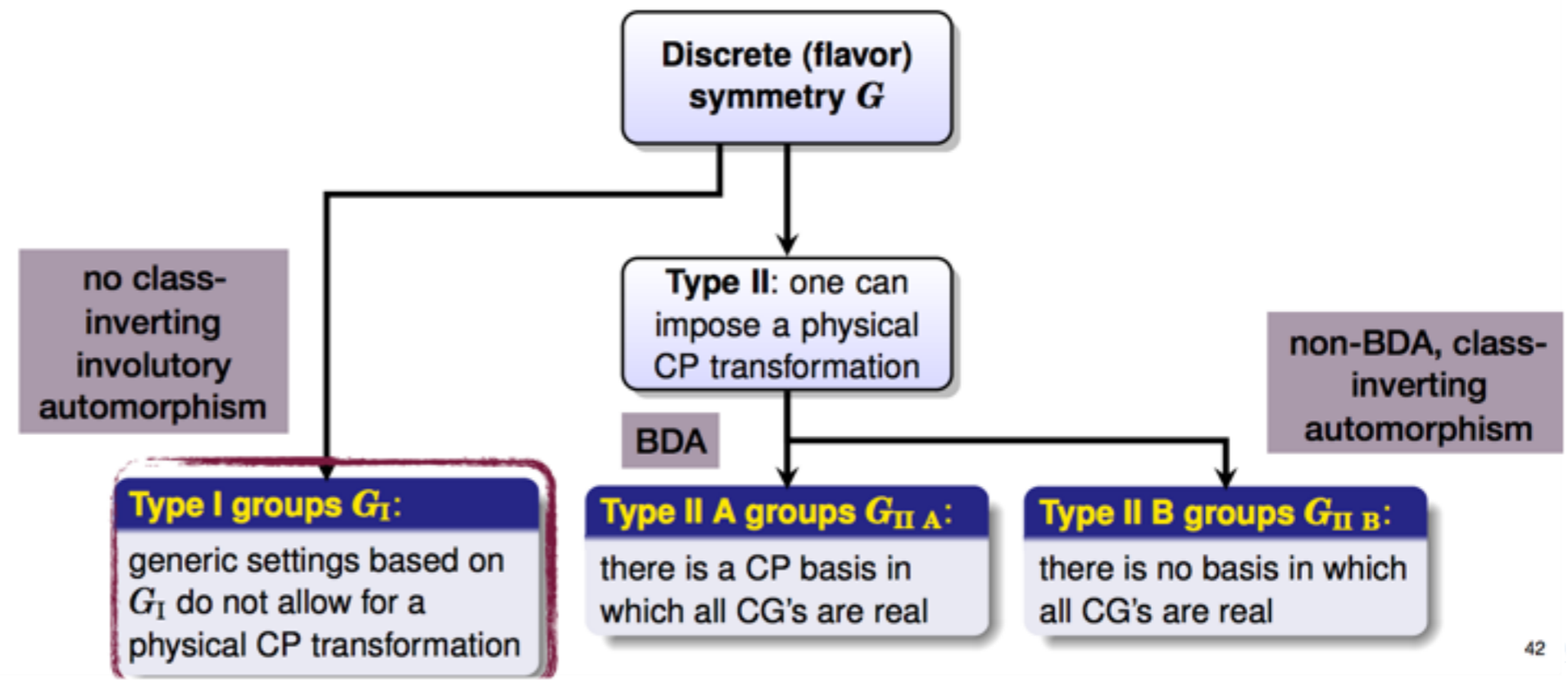


## A Novel Origin of CP Violation

M.-C.C, M. Fallbacher, K.T. Mahanthappa, M. Ratz, A. Trautner, NPB (2014)

- For discrete groups that do not have class-inverting, involutory automorphism, CP is generically broken by complex CG coefficients (**Type I Group**)
- Non-existence of such automorphism ⇔ Physical CP violation

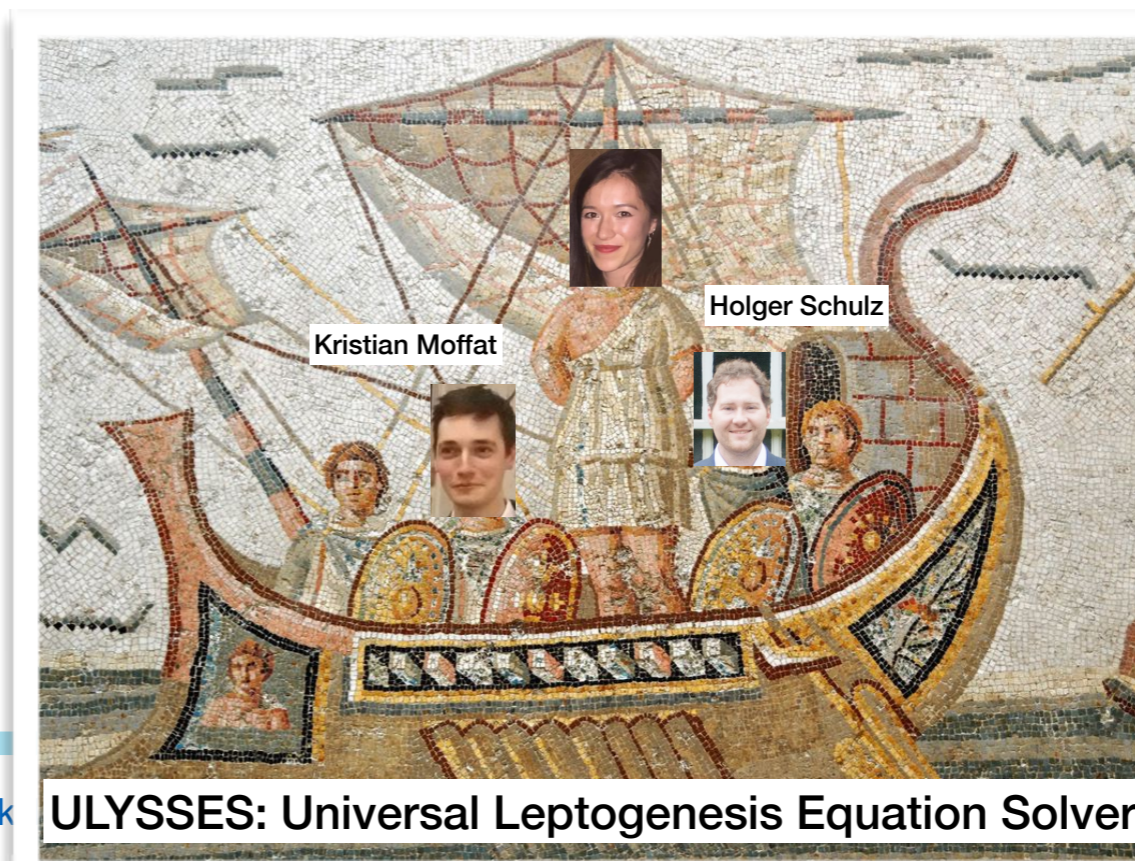
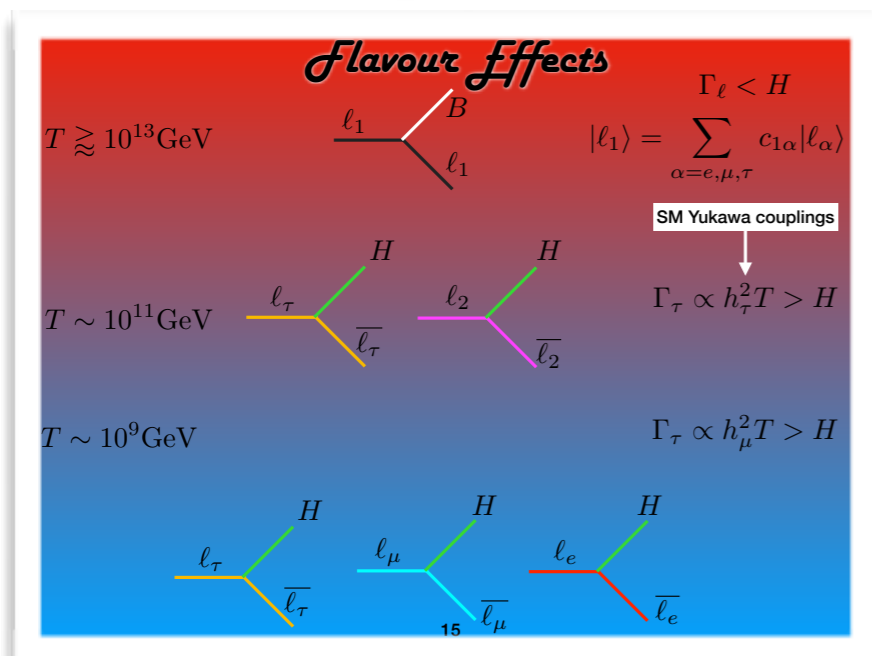
## CP Violation from Group Theory!



## Intermediate Scale Leptogenesis

## Leptogenesis via low energy $\mathcal{L}PV$

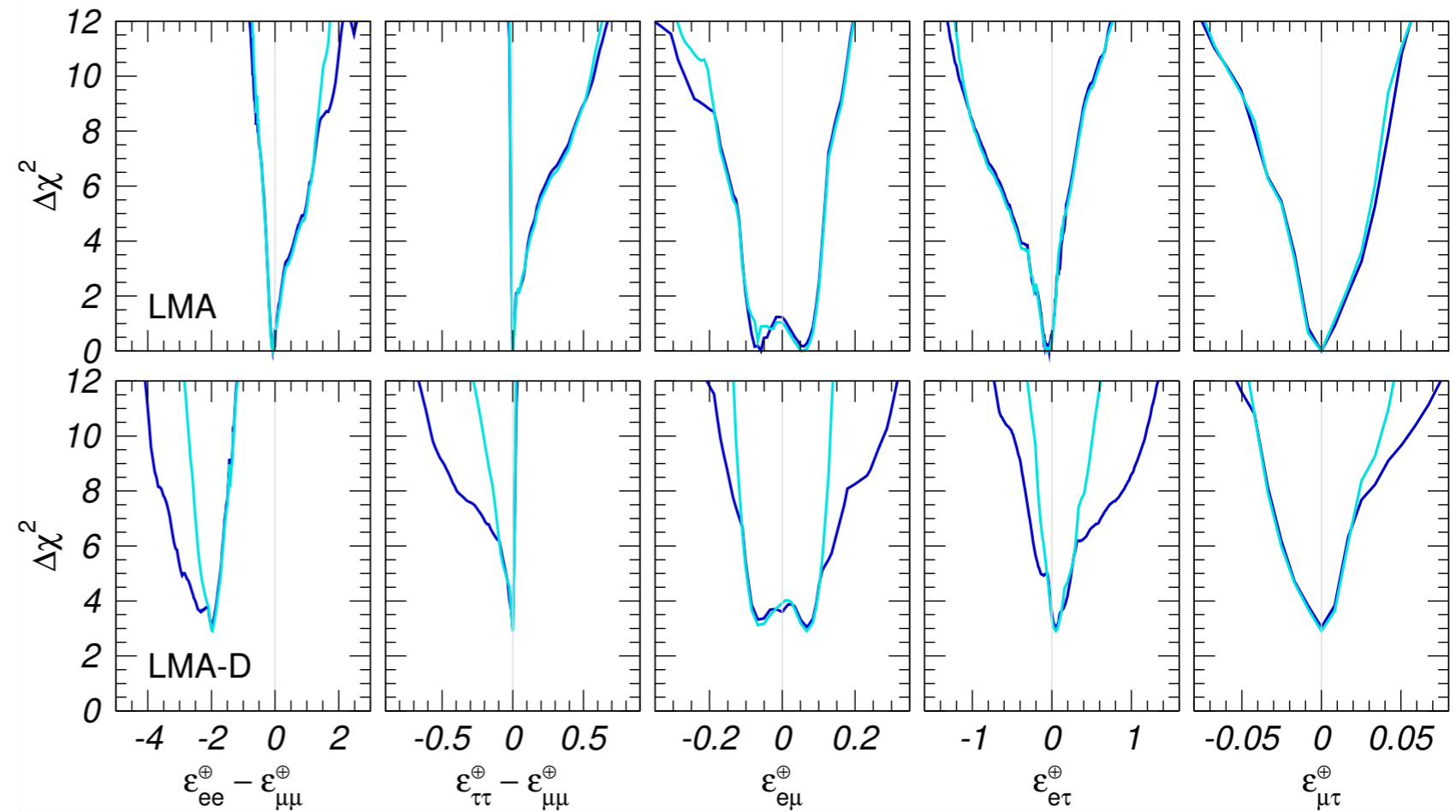
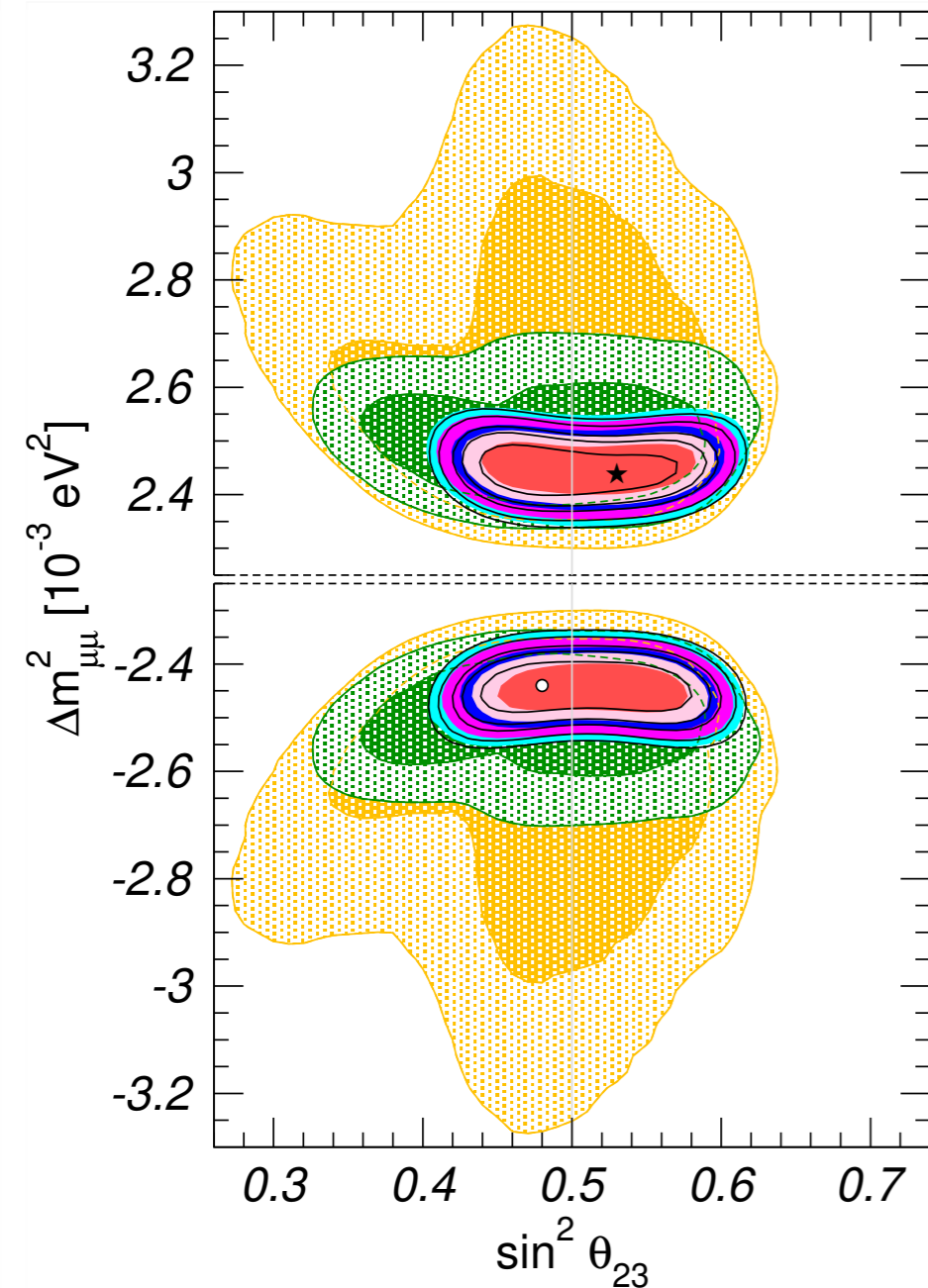
**Mini summary:** non-resonant thermal leptogenesis can explain BAU using only low scale phases at  $M_1 \sim 10^6$  GeV



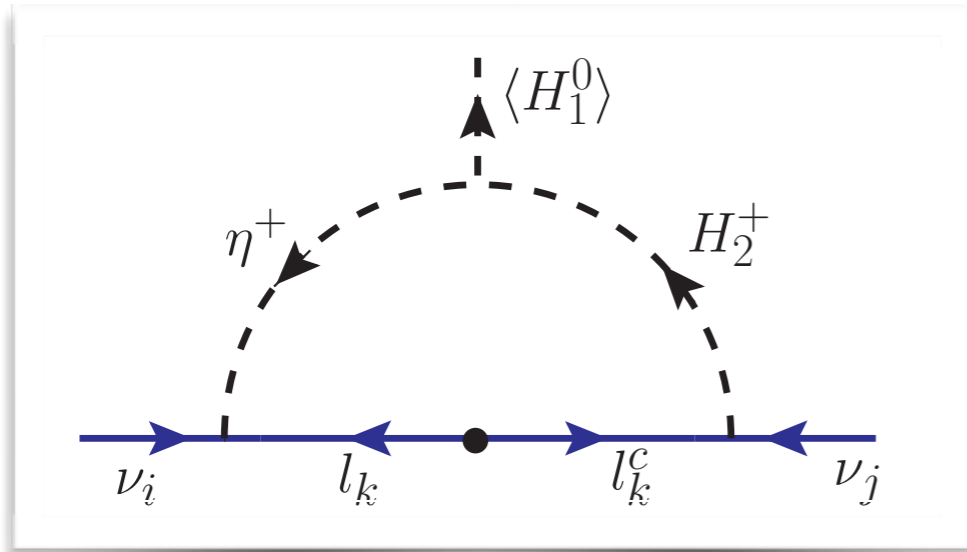
# Traditional NSI

Probing new physics

$$V_{mat} = \sqrt{2}G_F N_e(x) \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\tau\mu} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix}$$



# Mass and flavor Traditional NSI

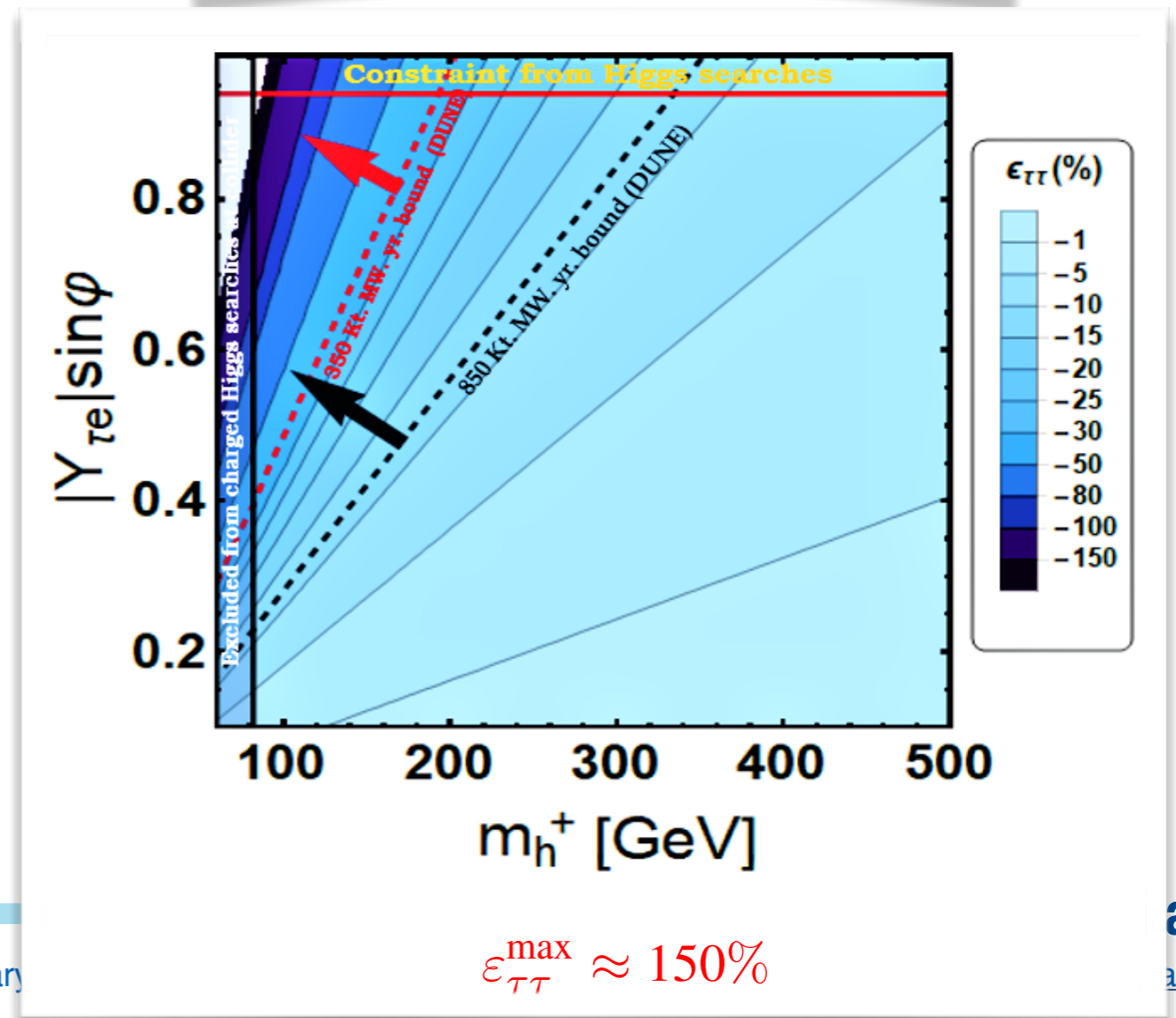
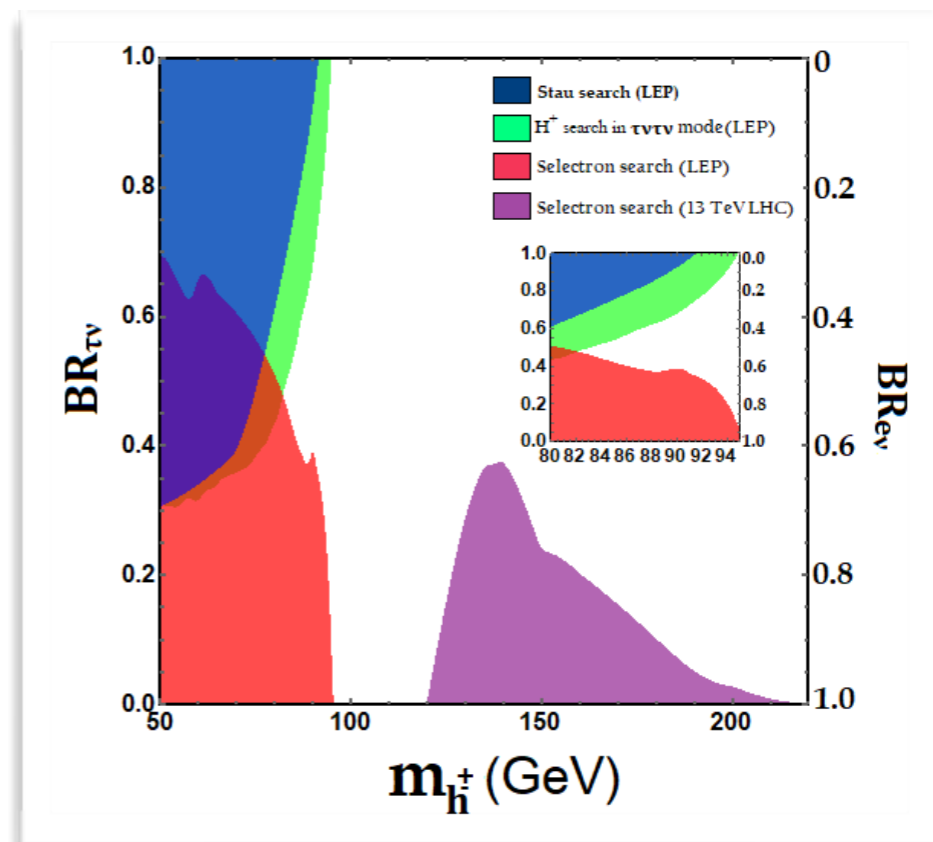
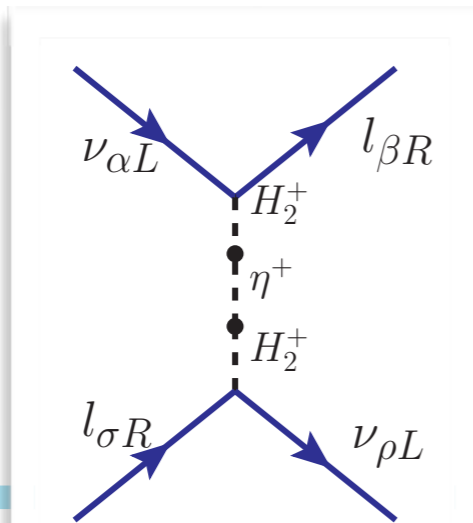


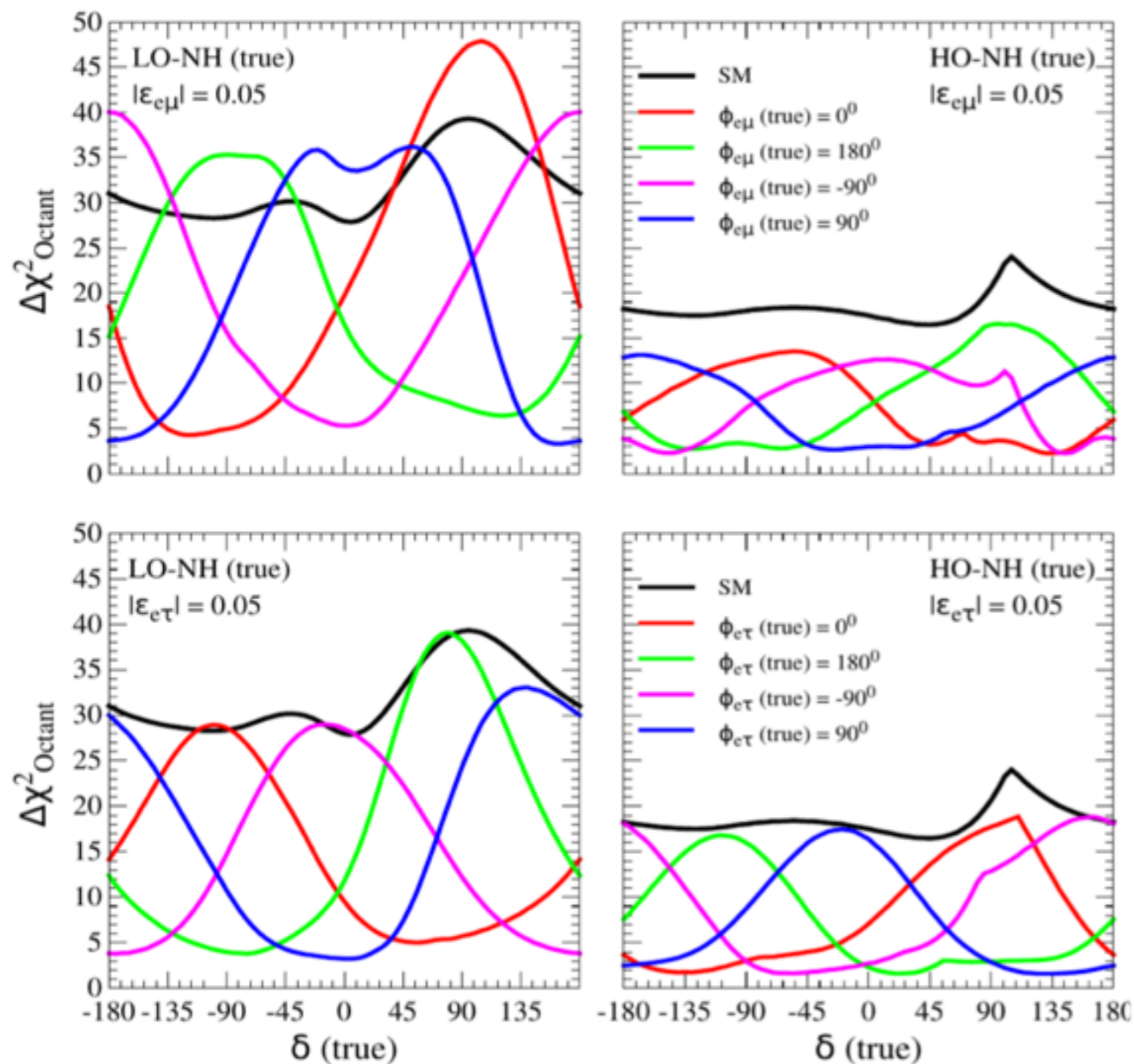
$$\mathcal{L}_Y = f^{ab} (\psi_{aL}^i C \psi_{bL}^j) \epsilon_{ij} \eta^+ + \bar{\psi}_{LY} \tilde{H}_1 e_R + \bar{\psi}_{LY} H_2 e_R + h.c.$$

$$V = \mu H_1^i H_2^j \eta^- + h.c. + \dots$$

$$y = \begin{pmatrix} y_{ee} & 0 & y_{e\tau} \\ 0 & y_{\mu\mu} & y_{\mu\tau} \\ 0 & 0 & y_{\tau\tau} \end{pmatrix}$$

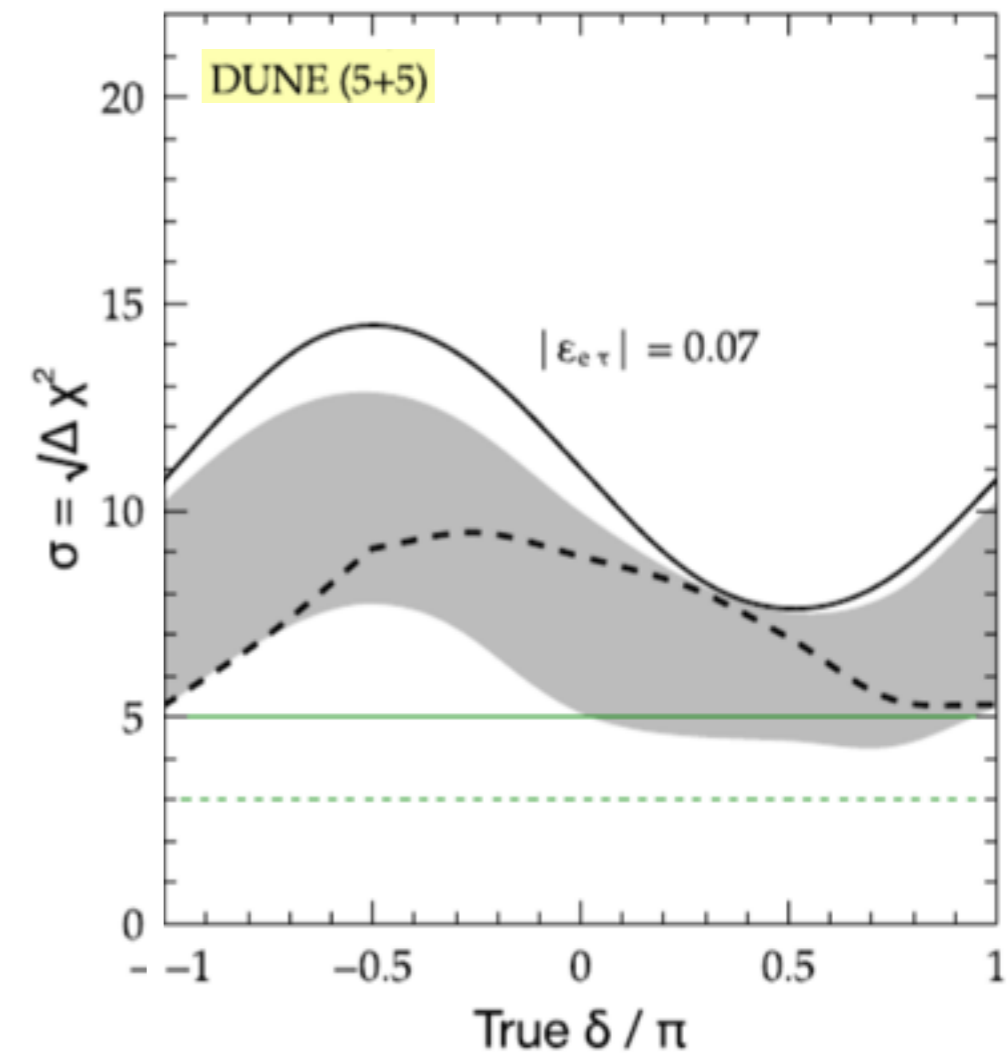
- The singly-charged scalars  $\eta^+$  and  $H_2^+$  induce NSI at **tree level**:





$$\Delta\chi^2_{\text{Octant}} = \chi^2_{\text{test Octant}} - \chi^2_{\text{true Octant}}$$

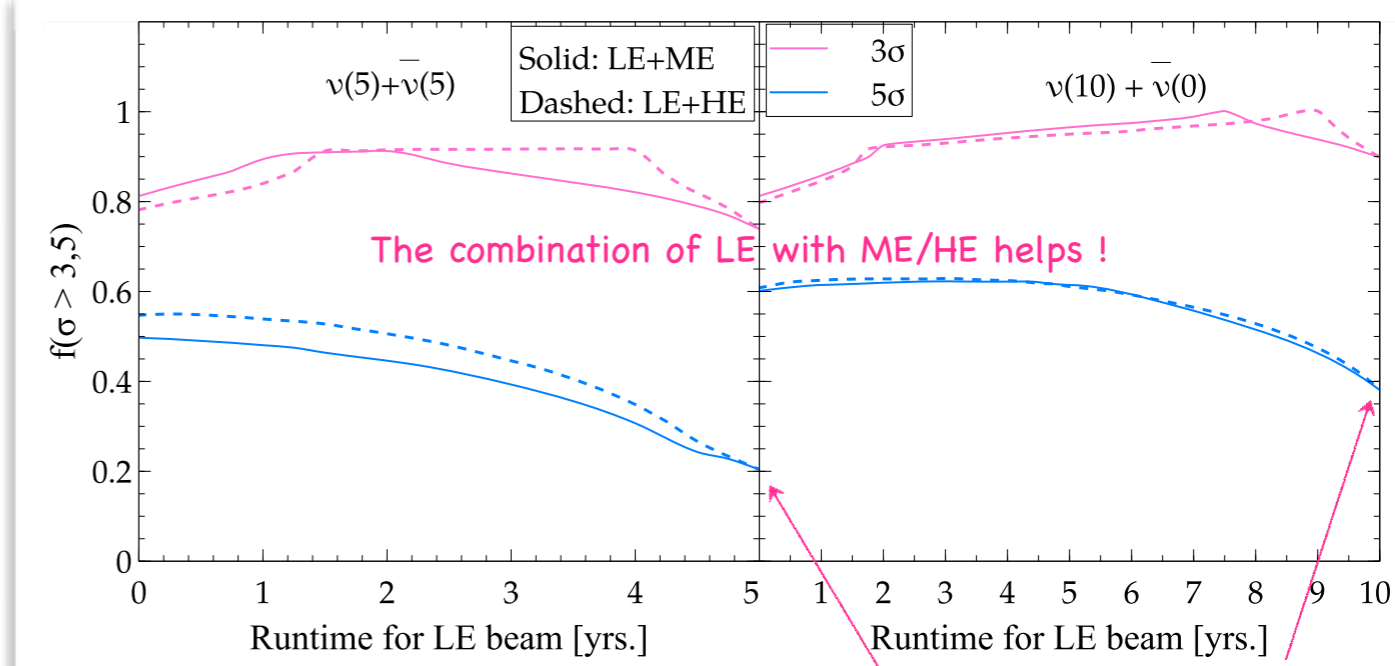
## Degeneracies



# Traditional NSI

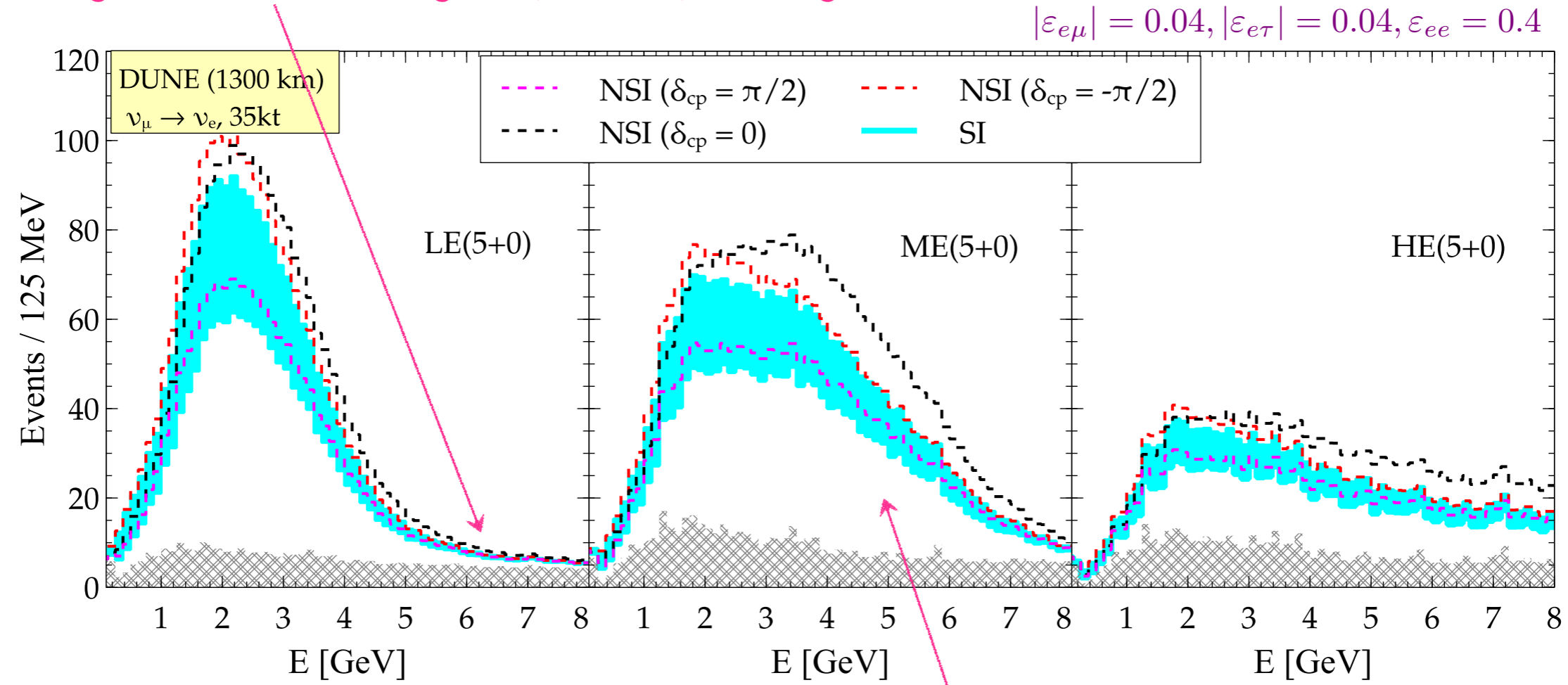
III. Can we separate new physics scenarios from the standard ?

IV. Correlations and degeneracies of NSI parameters with high energy beam tunes...



LE only case - the separability between SI and NSI is less!

Falling flux kills the large asymmetry at large E

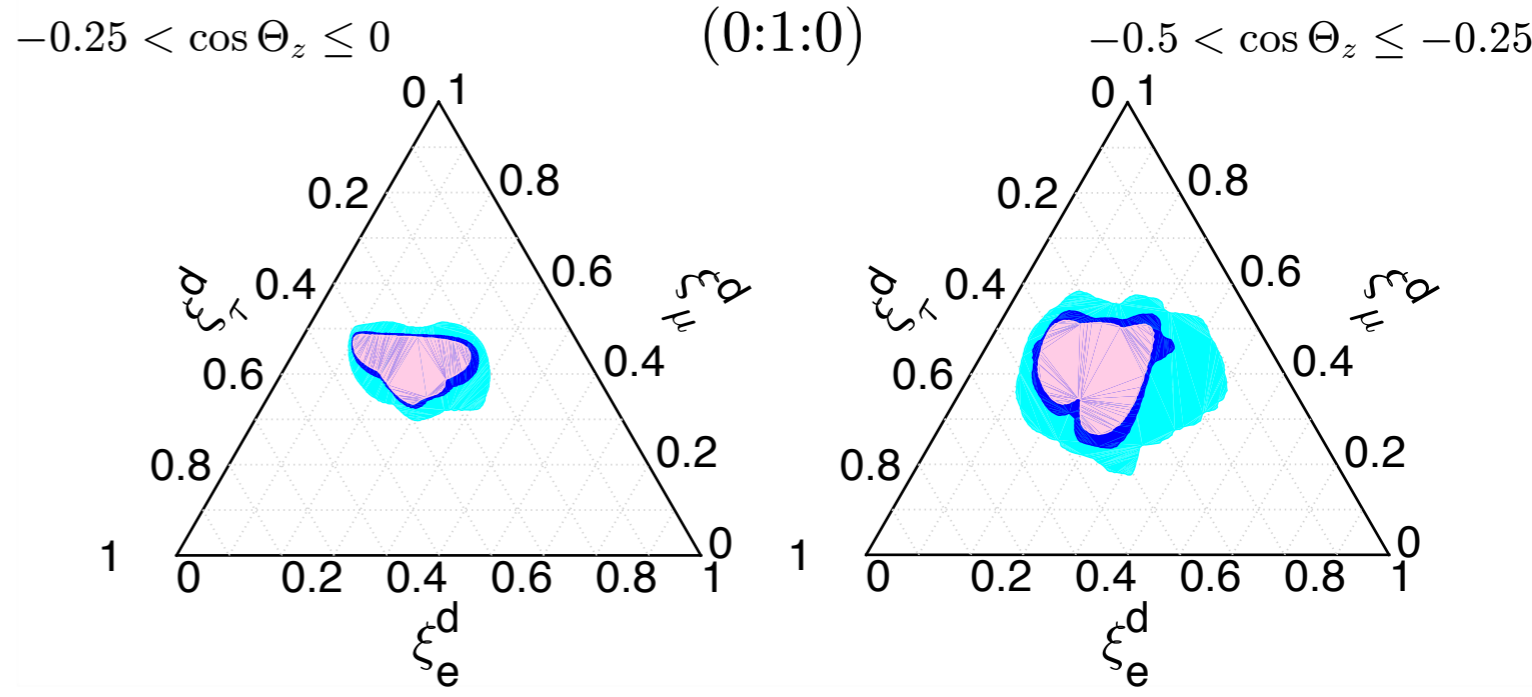
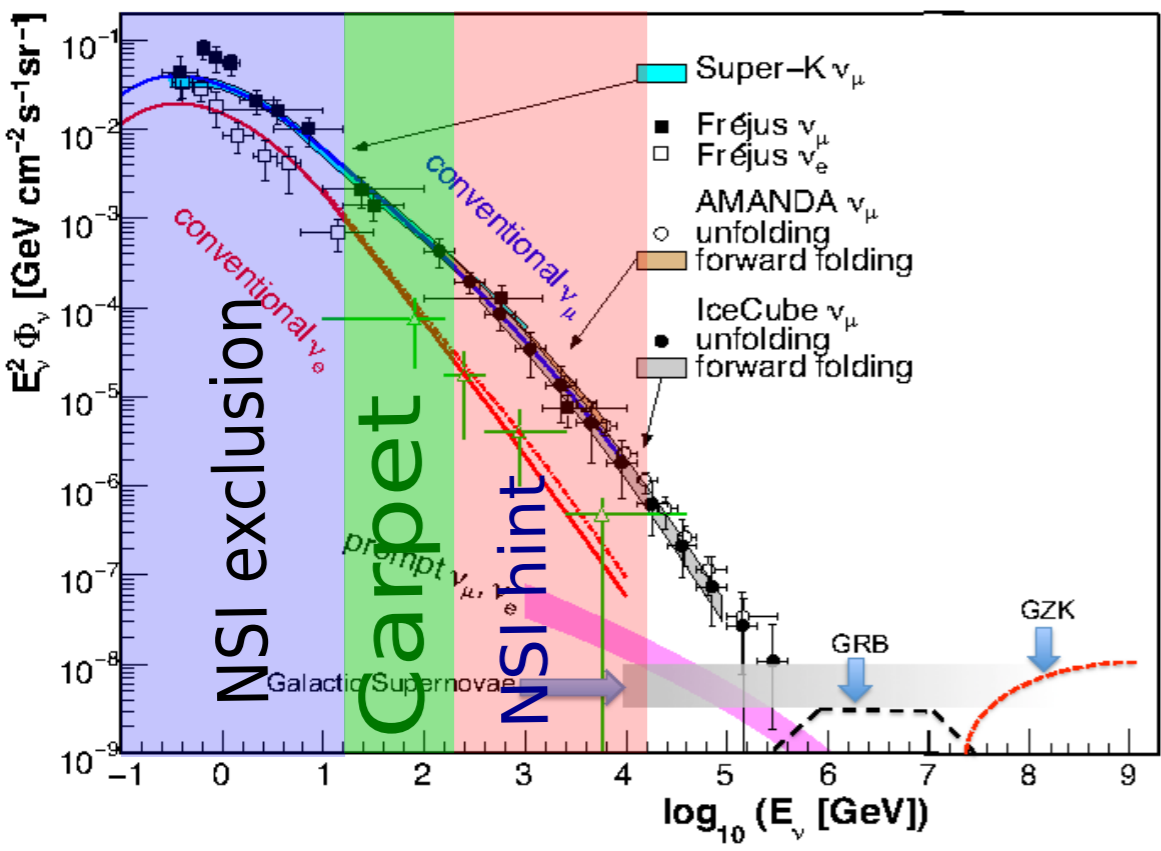
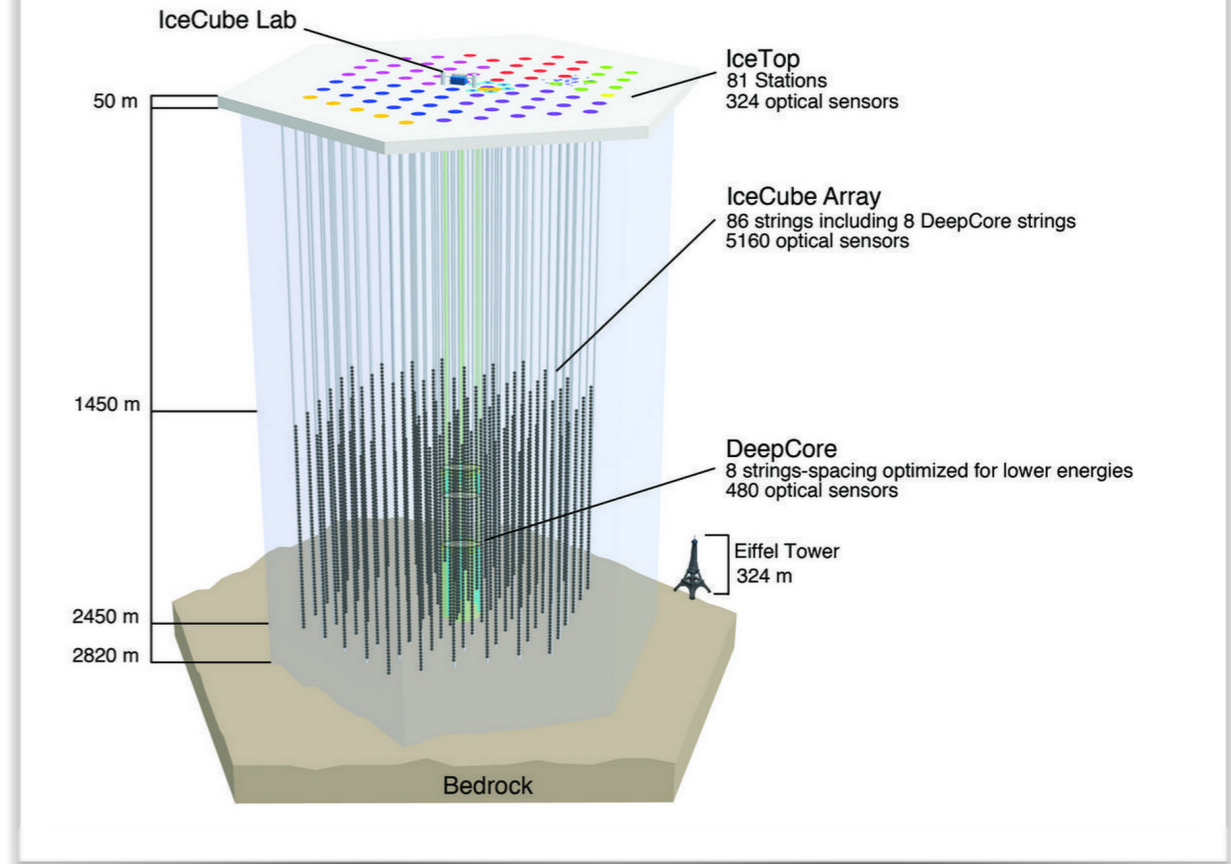
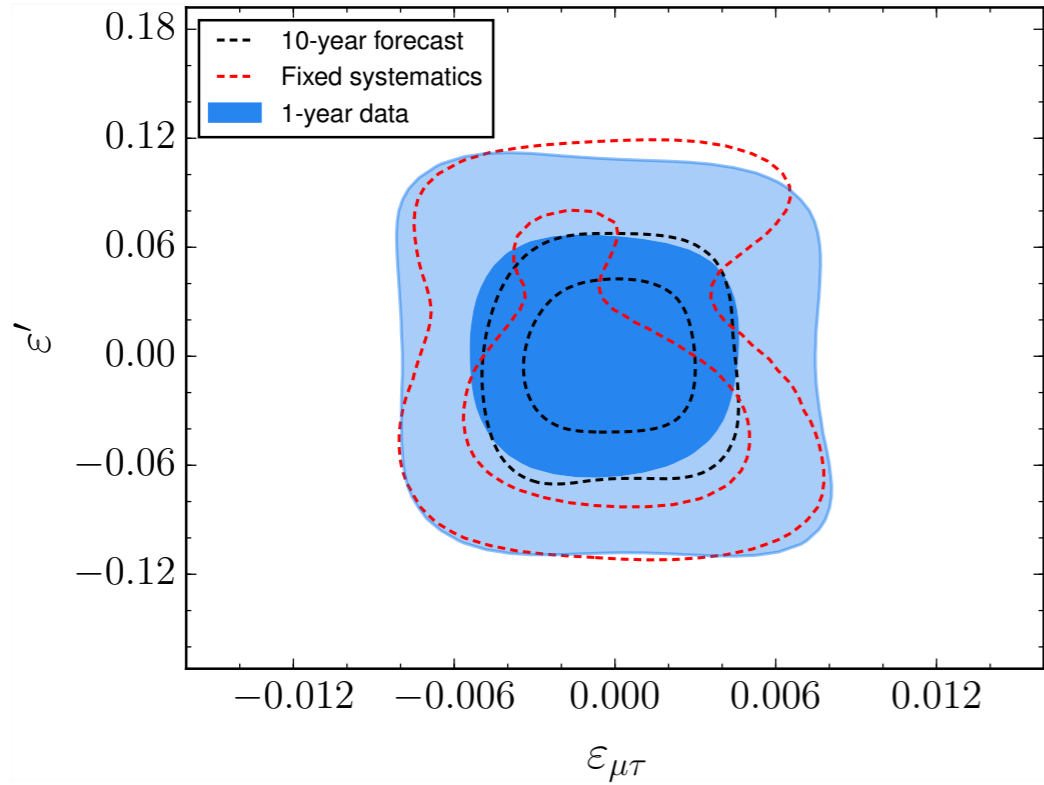


$$|\epsilon_{e\mu}| = 0.04, |\epsilon_{e\tau}| = 0.04, \epsilon_{ee} = 0.4$$

Better ability to separate black curve from cyan band



# Traditional NSI



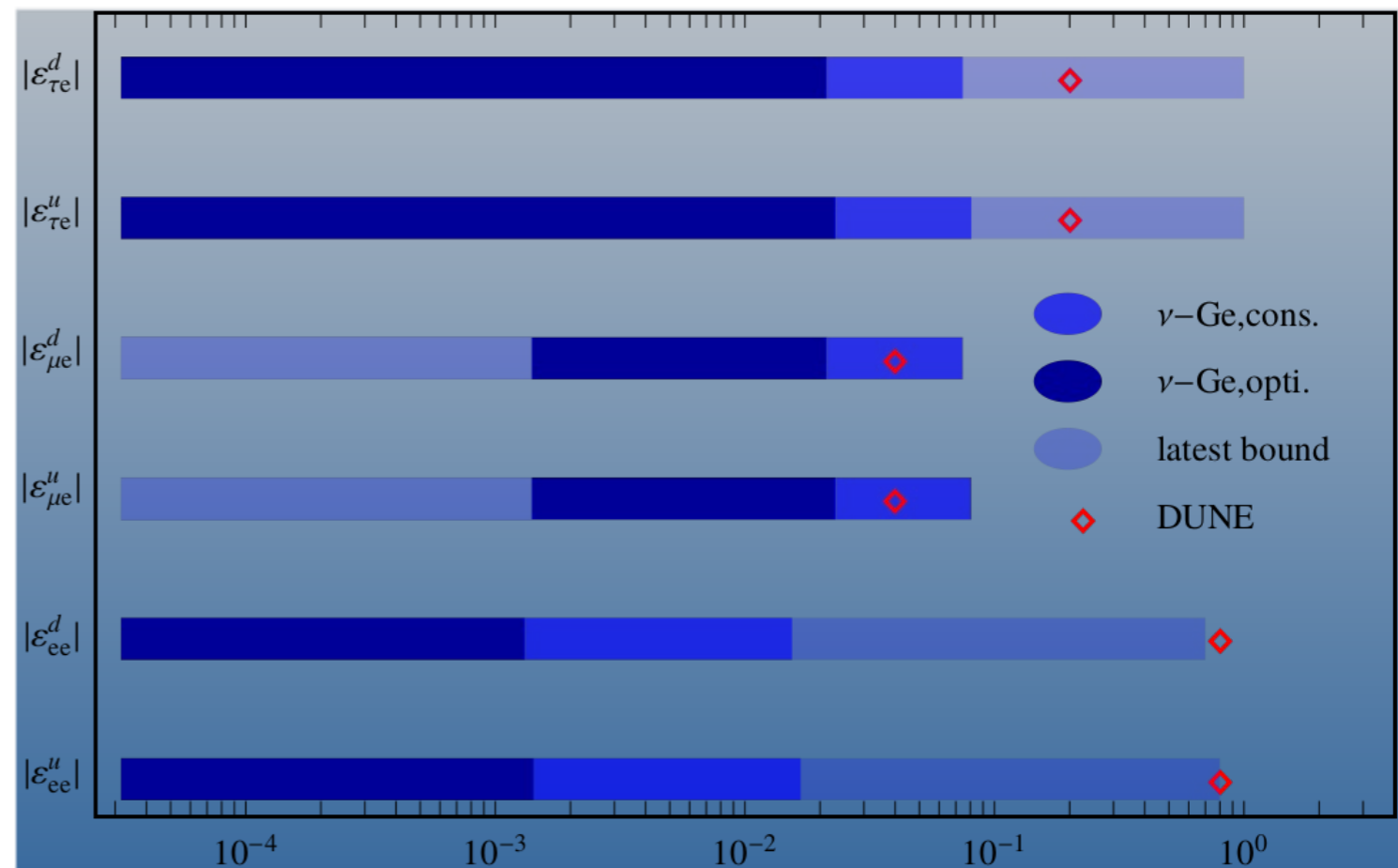
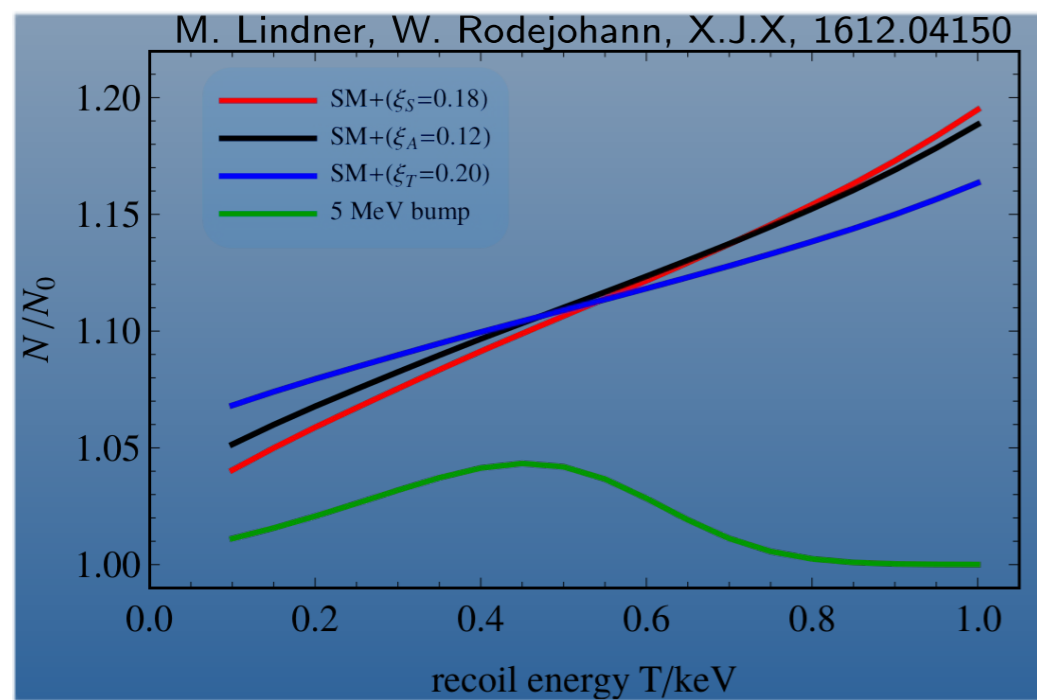
# Traditional NSI

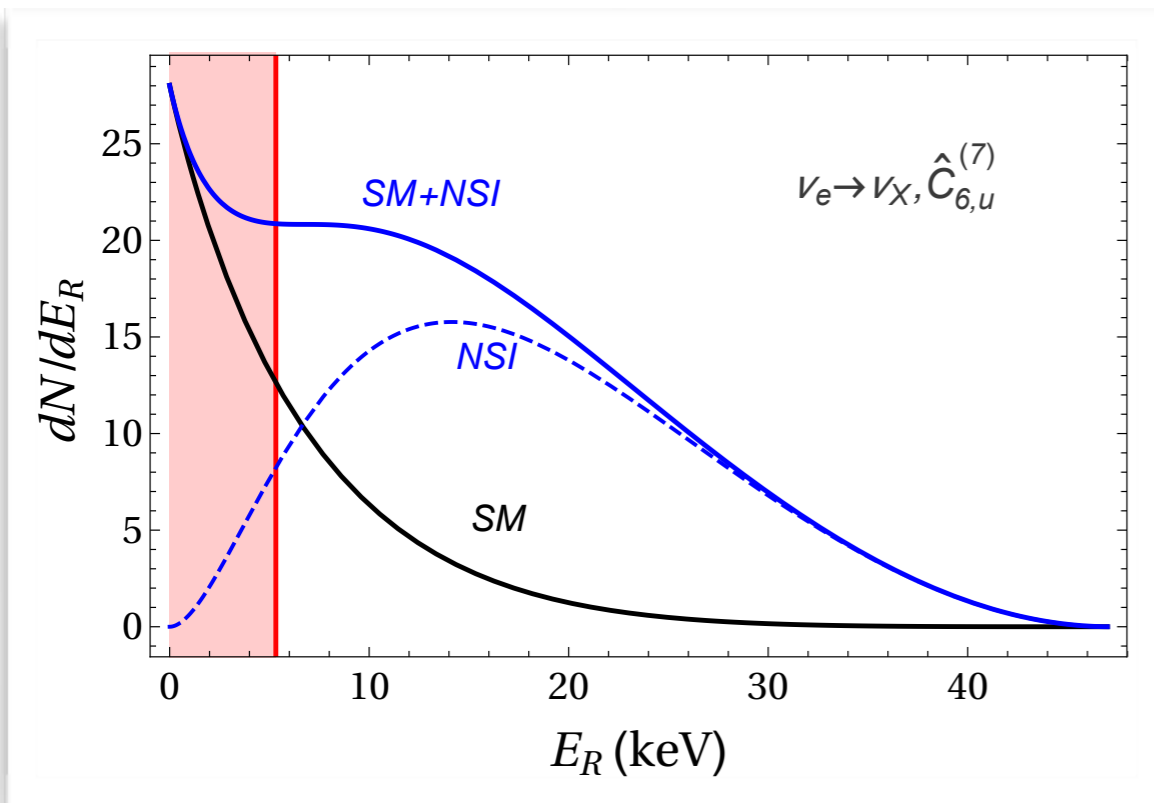
SPVAT (Scalar, Pseudo-S, Vector, Axial-V, Tensor)

$$\mathcal{L} \supset \frac{G_F}{\sqrt{2}} \sum_{a=S,P,V,A,T} \bar{\nu} \Gamma^a \nu [\bar{\psi} \Gamma^a (C_a + D_a i\gamma^5) \psi],$$

$$\Gamma^a = \{1, i\gamma^5, \gamma^\mu, \gamma^\mu \gamma^5, \sigma^{\mu\nu} \equiv \frac{i}{2} [\gamma^\mu, \gamma^\nu]\}.$$

Use coherent  $\nu$ -N scattering





## NSI complete operator basis

3 - Flavors basis ( $\mu \sim 2$  GeV):  $f = e, \mu, u, d, s$

Dimension five:

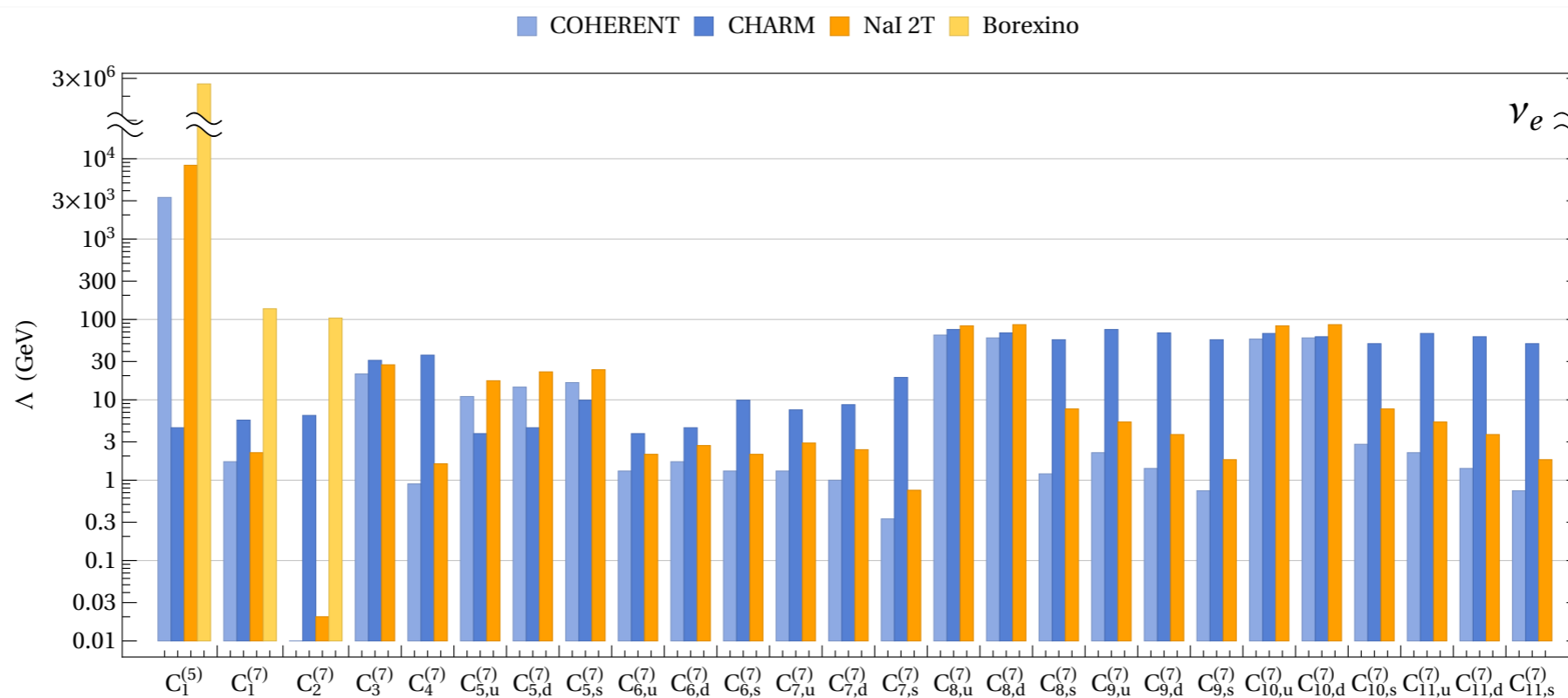
$$Q_1^{(5)} = \frac{e}{8\pi^2} (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) F_{\mu\nu},$$

Dimension six:

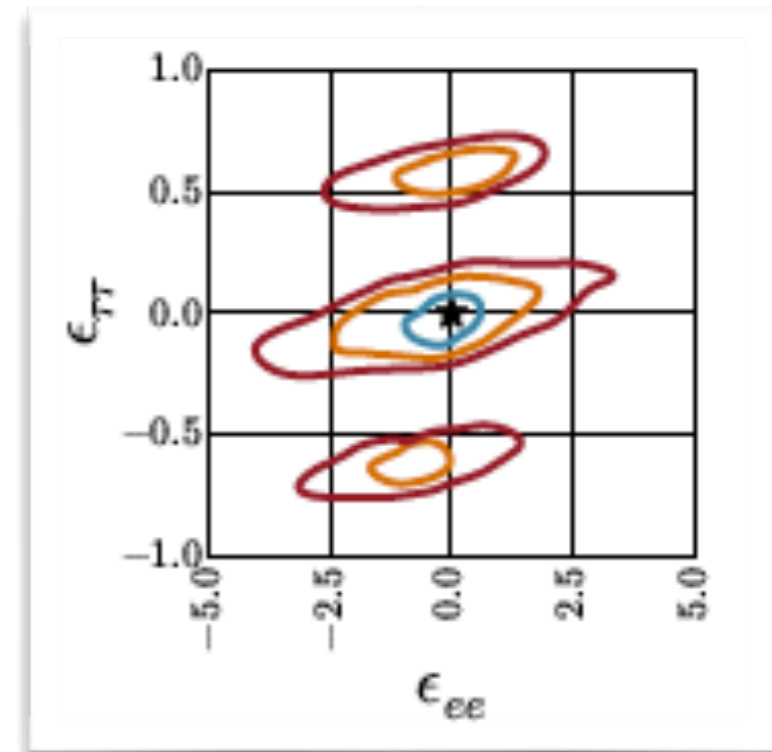
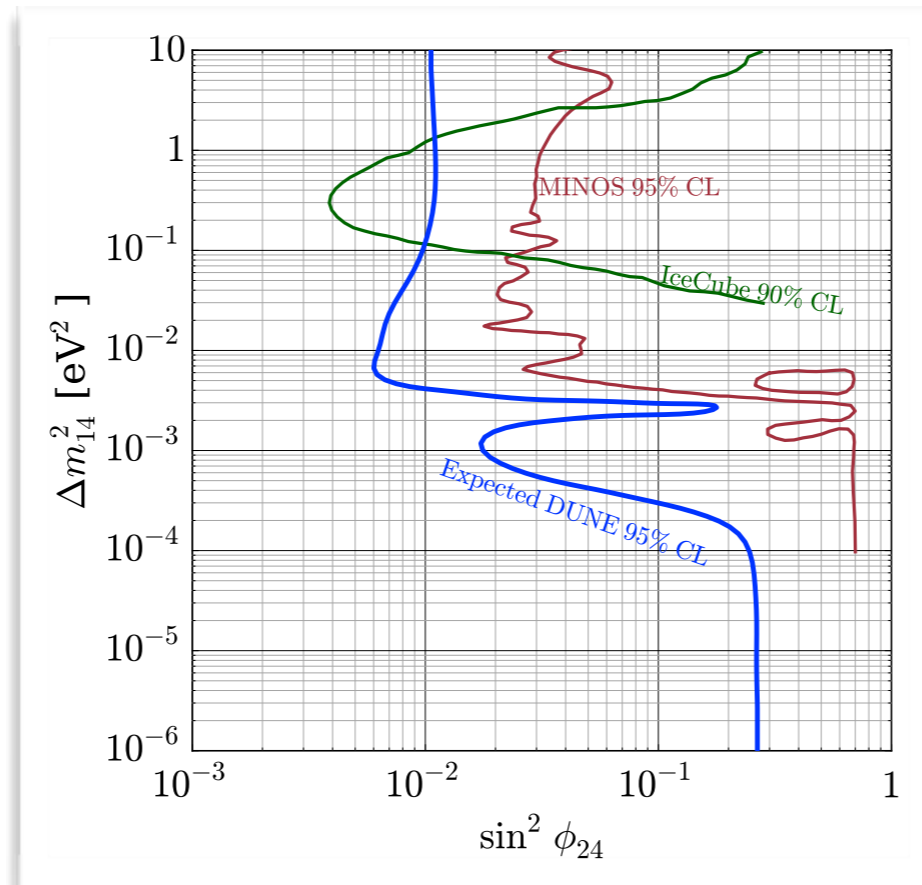
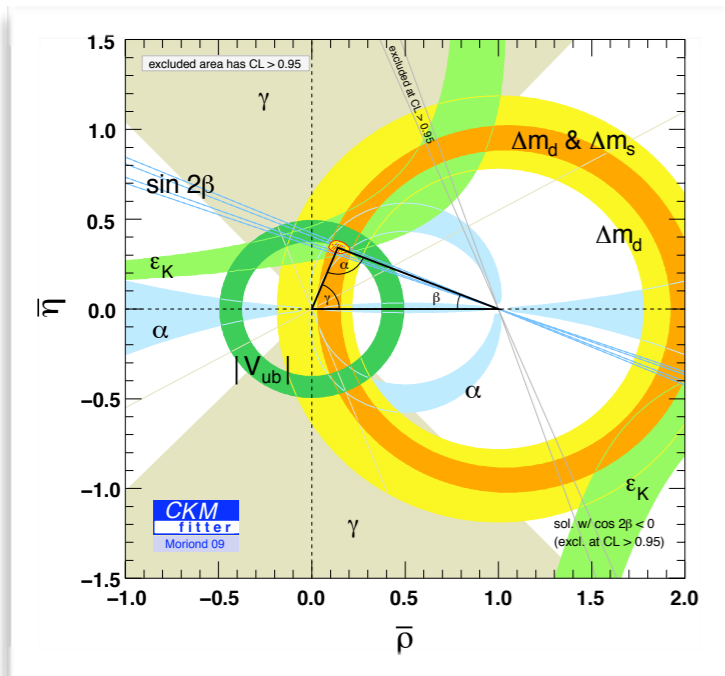
$$Q_{1,f}^{(6)} = (\bar{\nu}_\beta \gamma_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu f), \quad Q_{2,f}^{(6)} = (\bar{\nu}_\beta \gamma_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu \gamma_5 f),$$

Dimension seven:

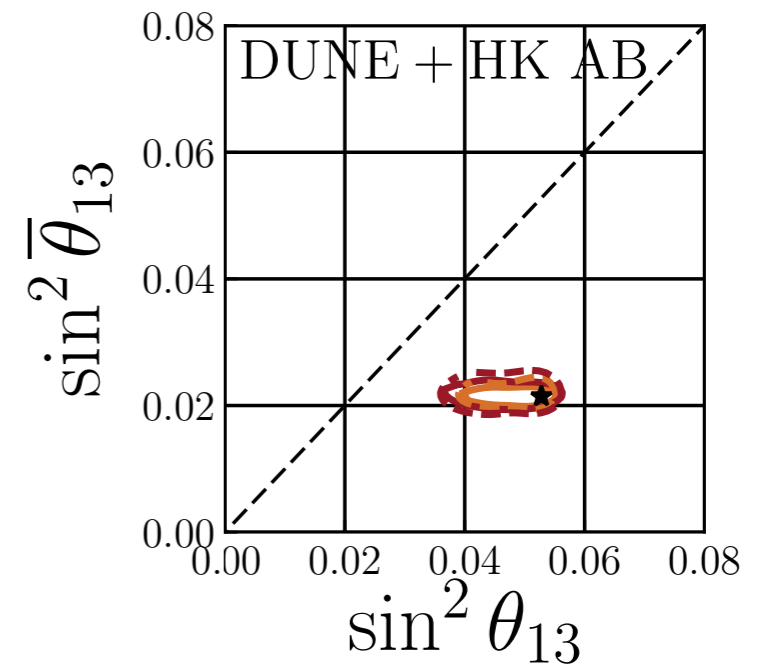
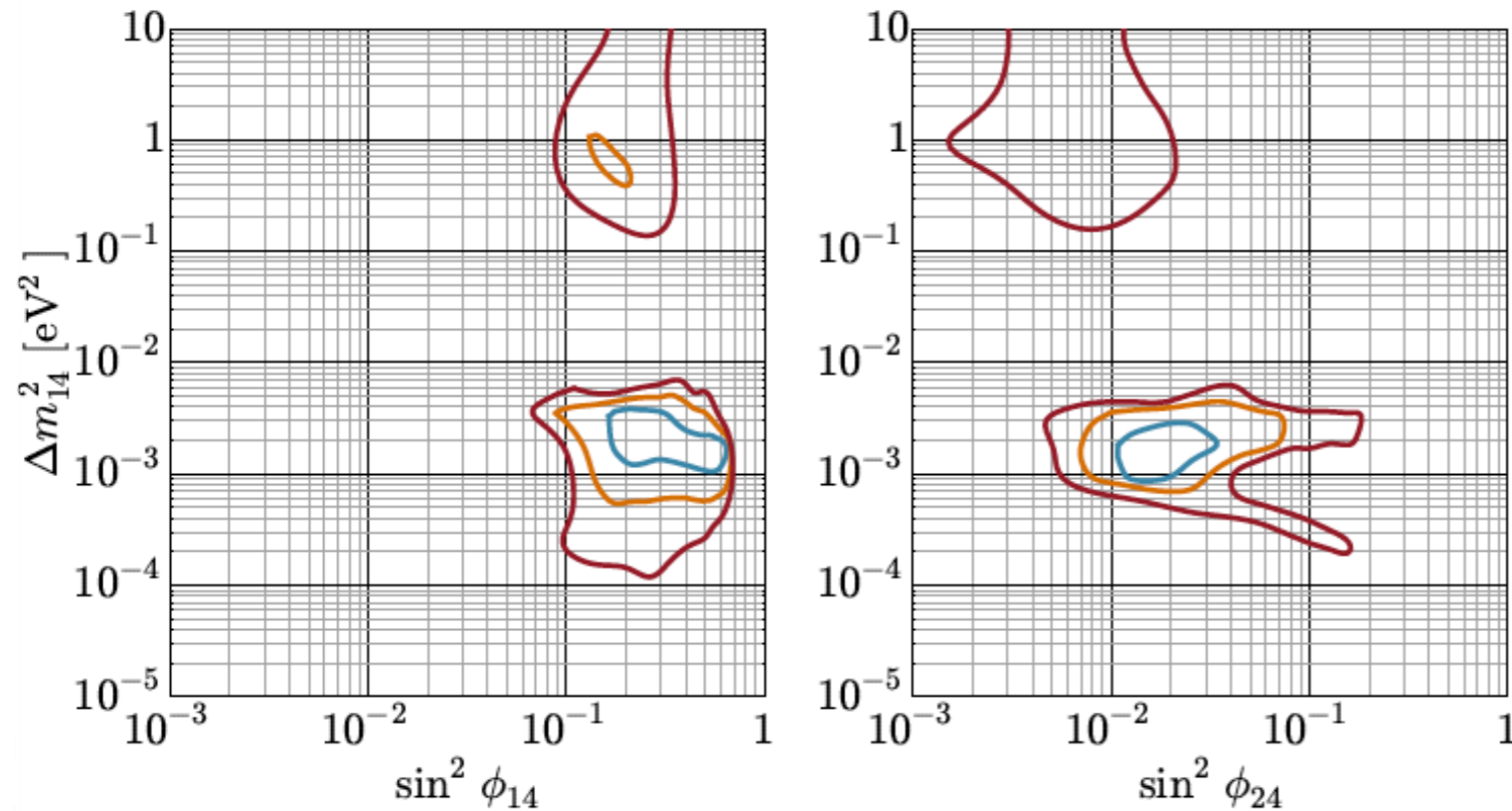
$$\begin{aligned} Q_1^{(7)} &= \frac{\alpha}{12\pi} (\bar{\nu}_\beta P_L \nu_\alpha) F^{\mu\nu} F_{\mu\nu}, & Q_2^{(7)} &= \frac{\alpha}{8\pi} (\bar{\nu}_\beta P_L \nu_\alpha) F^{\mu\nu} \tilde{F}_{\mu\nu}, \\ Q_3^{(7)} &= \frac{\alpha_s}{12\pi} (\bar{\nu}_\beta P_L \nu_\alpha) G^{a\mu\nu} G_{\mu\nu}^a, & Q_4^{(7)} &= \frac{\alpha_s}{8\pi} (\bar{\nu}_\beta P_L \nu_\alpha) G^{a\mu\nu} \tilde{G}_{\mu\nu}^a, \\ Q_{5,f}^{(7)} &= m_f (\bar{\nu}_\beta P_L \nu_\alpha) (\bar{f} f), & Q_{6,f}^{(7)} &= m_f (\bar{\nu}_\beta P_L \nu_\alpha) (\bar{f} i \gamma_5 f), \\ Q_{7,f}^{(7)} &= m_f (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) (\bar{f} \sigma_{\mu\nu} f), & Q_{8,f}^{(7)} &= (\bar{\nu}_\beta \overset{\leftrightarrow}{i} \partial_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu f), \\ Q_{9,f}^{(7)} &= (\bar{\nu}_\beta \overset{\leftrightarrow}{i} \partial_\mu P_L \nu_\alpha) (\bar{f} \gamma^\mu \gamma_5 f), & Q_{10,f}^{(7)} &= \partial_\mu (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) (\bar{f} \gamma_\nu f), \\ Q_{11,f}^{(7)} &= \partial_\mu (\bar{\nu}_\beta \sigma^{\mu\nu} P_L \nu_\alpha) (\bar{f} \gamma_\nu \gamma_5 f). \end{aligned}$$



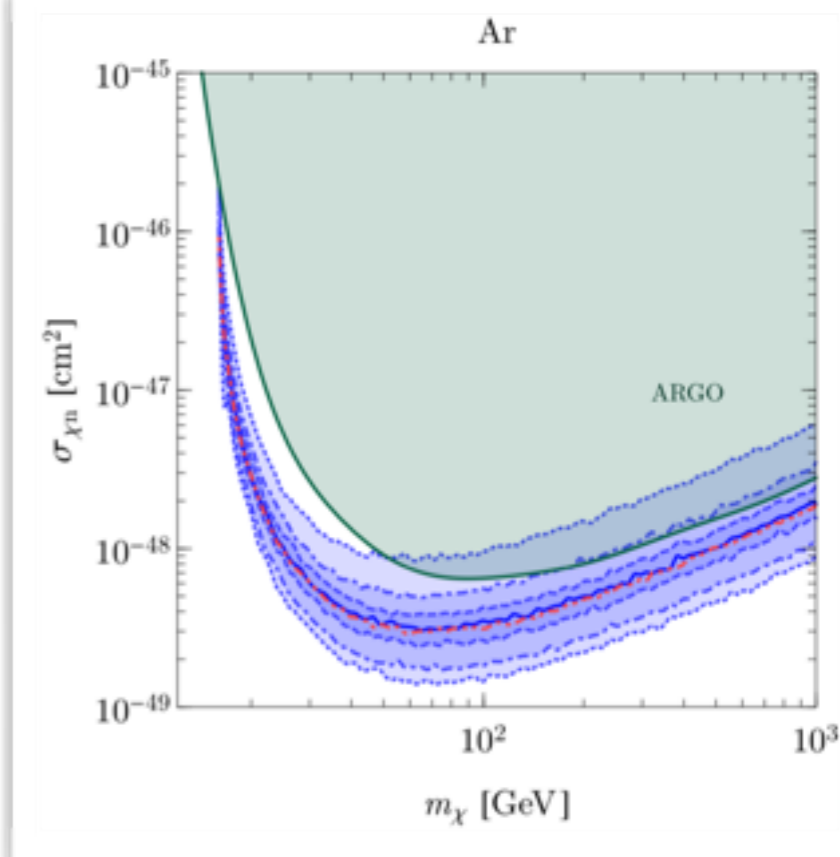
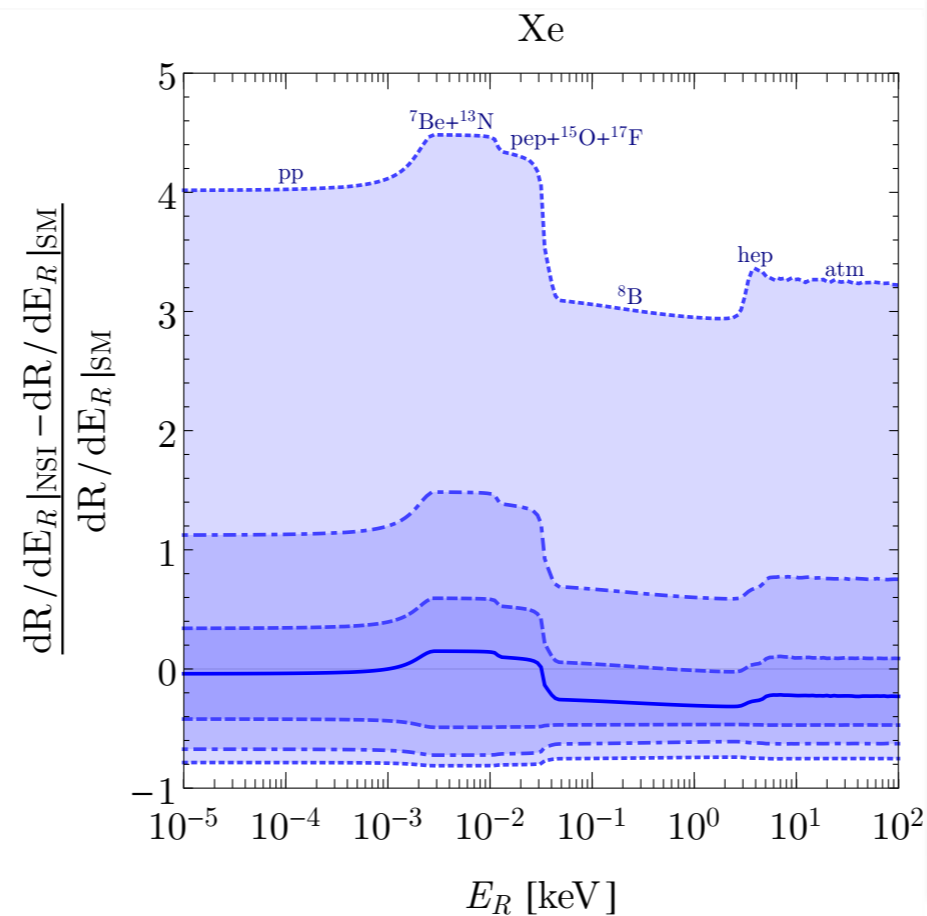
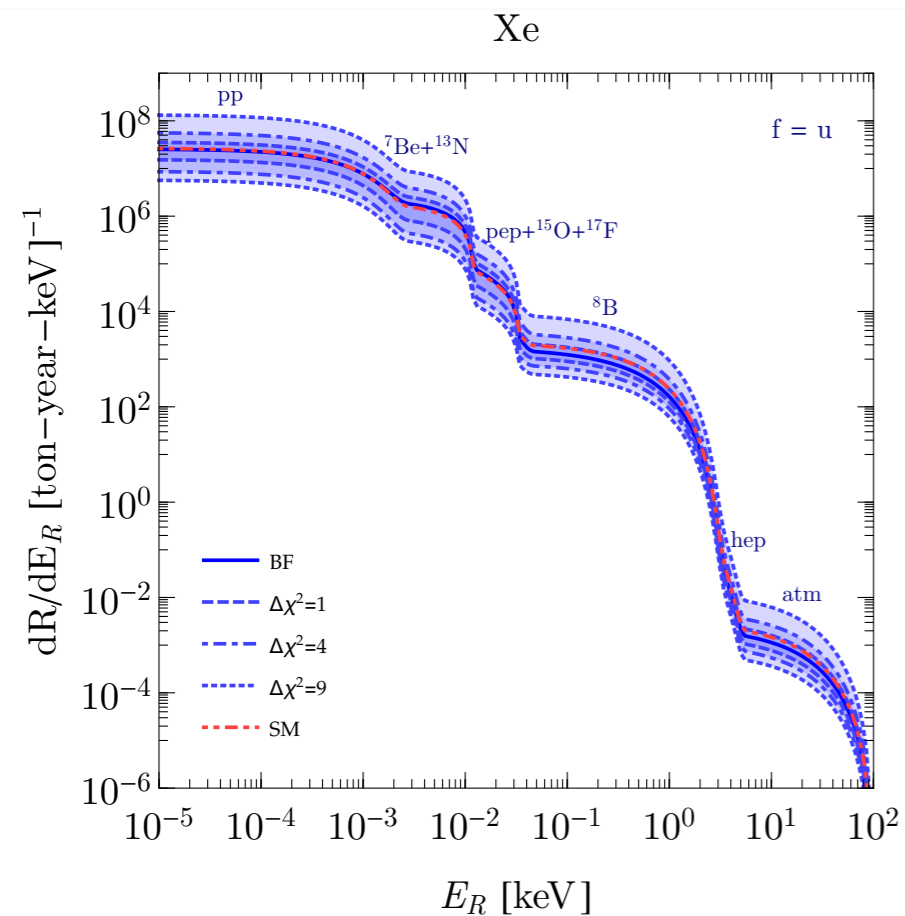
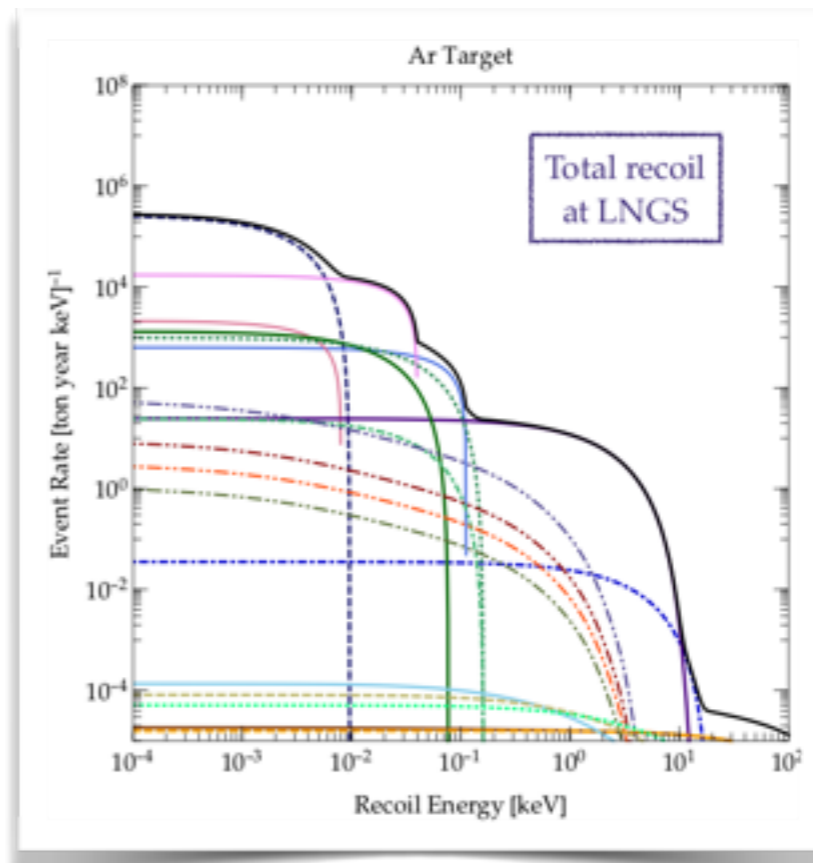
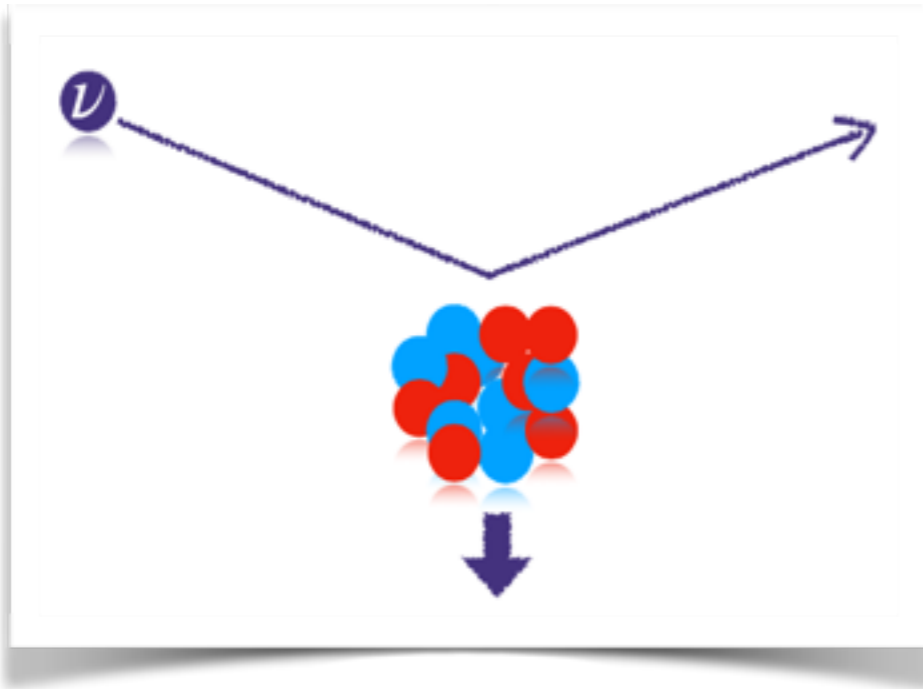
# Traditional NSI and more



## Telling Different Scenarios Apart:



# Dark matter Traditional NSI

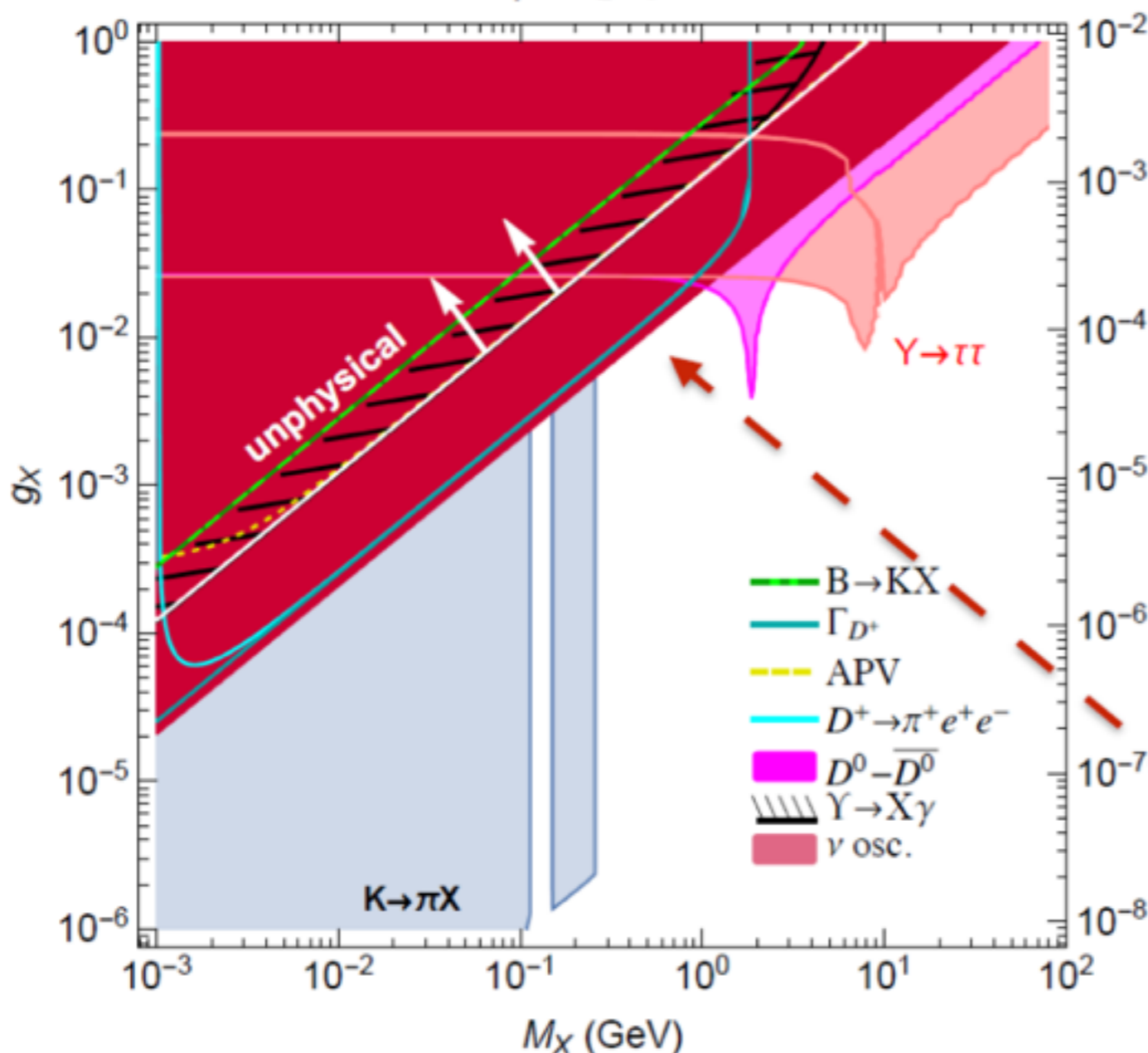


# B-L for the 3<sup>rd</sup> generation

## Neutrinos and low scale new Physics

Can there be a flavor mediators at low scale???

$$\tan\beta = v_2/v_1 = 10$$



U(1) B – L of the third family

Complete model, including scalar sector and CKM generation

Vast phenomenology:  
Z-X mixing ( $s_X$ )  
D oscillations

Atomic Parity Violation  
Upsilon, B, D, and K decays  
Higgs, top, Z, and W decays  
Neutrino oscillations

...

# Light mediators

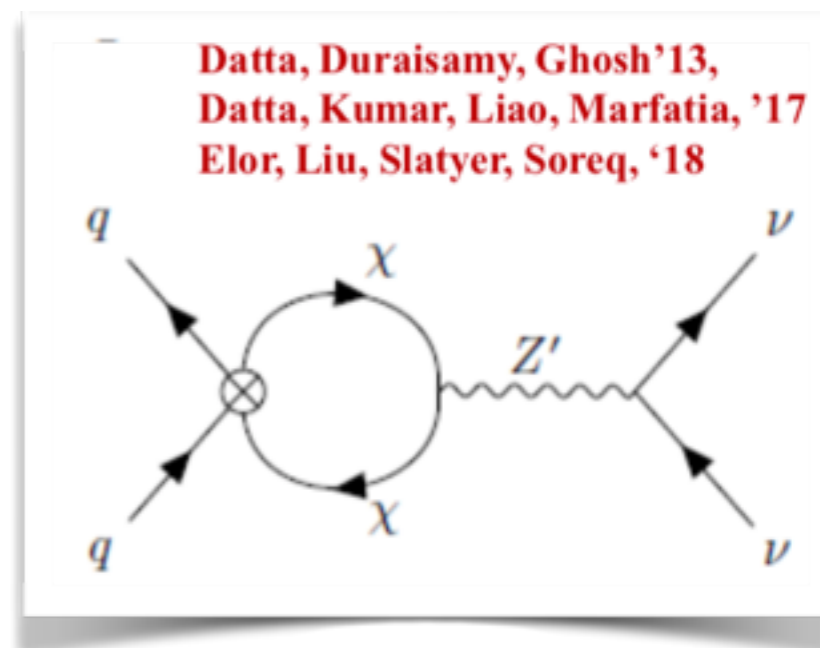
UV complete models

$$DM: SU(2)_L \times U(1)_Y \times U(1)_{T3R}$$

Anomaly free

field	$q_{T3R}$
$q_R^u$	-2
$q_R^d$	2
$\ell_R$	2
$\nu_R$	-2
$\eta_L$	1
$\eta_R$	-1
$\phi$	-2

New form factors



New DM phenomenology

*Migdal effect (Ionization and excitation of electron)*

Ibe, Nakano, Shoji, Sujuki, 2018  
Dolan, Kahlhoefer, McCabe, 2018

*Cosmic ray scattered*

Bringmann, Pospelov, 2018

Ema, Sala, Sato, 2018

Dent, Dutta, Newstead, Shoemaker, to appear



## 1. Flavor Puzzle is a much<sup>n</sup> harder problem

→ Scale and couplings wide open in theory space.

Will search for ANYTHING

new states of mass  $M$ , new couplings/mixings  $k, V_{ij} \dots$

Will search EVERY WHERE

low-energy & high-energy regimes.

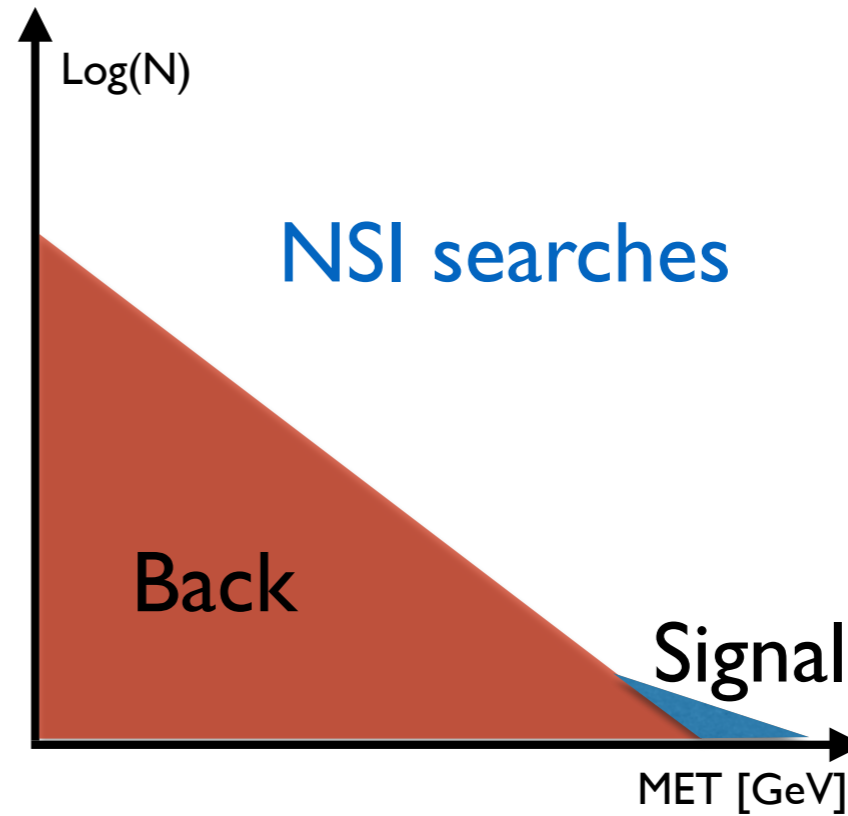
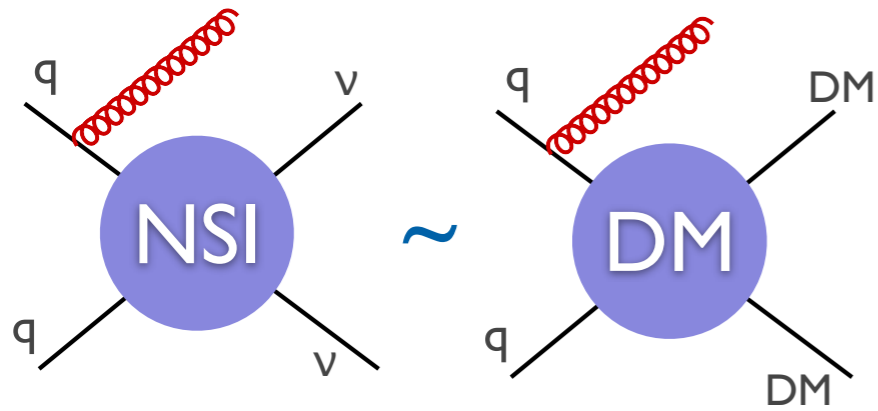
	$0\nu 2\beta$	$\mu$ - $e$ conversion $\mu \rightarrow e\gamma$ etc.	rare decays $\tau, K, D, B$	colliders $e^+e^-, pp$	features
Type-I	✓	✓	✓	✓	$N$
Type-II	✓	✓	✓	✓	$H^{\pm\pm}, W_R^\pm$
Type-III	✓	✓	?	✓	$T^\pm$
Zee-Babu	?	✓	?	✓	$k^{\pm\pm}, h^\pm$
Ma models	?	✓	?	✓	scalars, DM
RPV/leptoquarks	?	✓	✓	✓	$\ell_q$
extra-dim	?	?	✓	✓	KK states
Inverse/linear	?	?	✓	✓	
Pseudo Dirac	?	?	✓	✓	
NSI	?	?	?	✓	mediators
... ..					

It is of fundamental importance to test the Majorana nature of  $\nu$ 's.

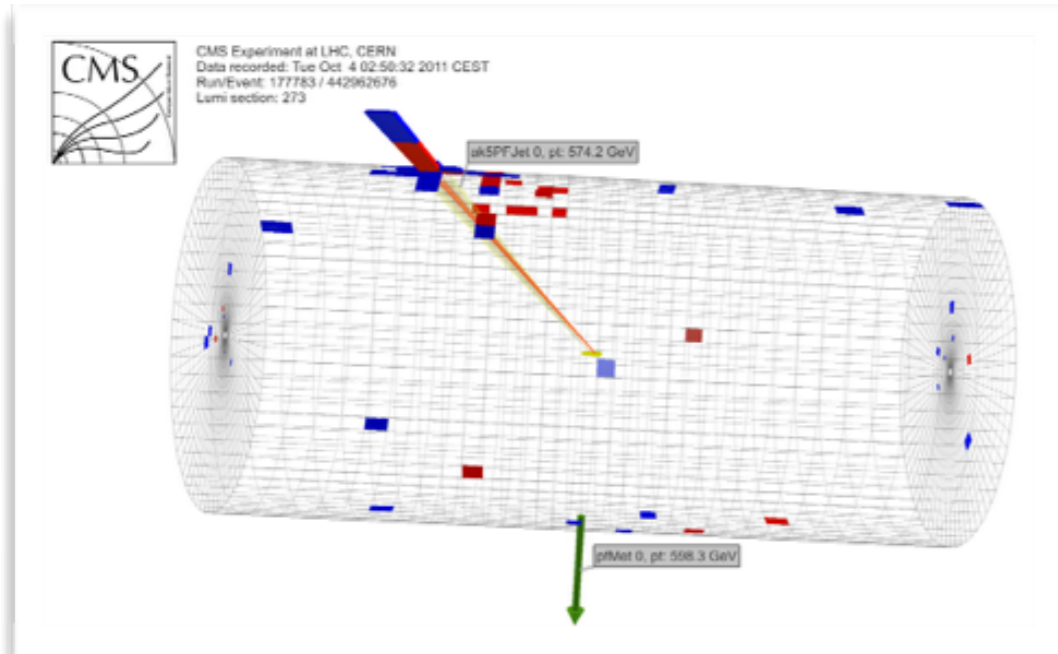
IF lucky, hadron colliders may serve as the discovery machine for Majorana nature of  $\nu$ 's.



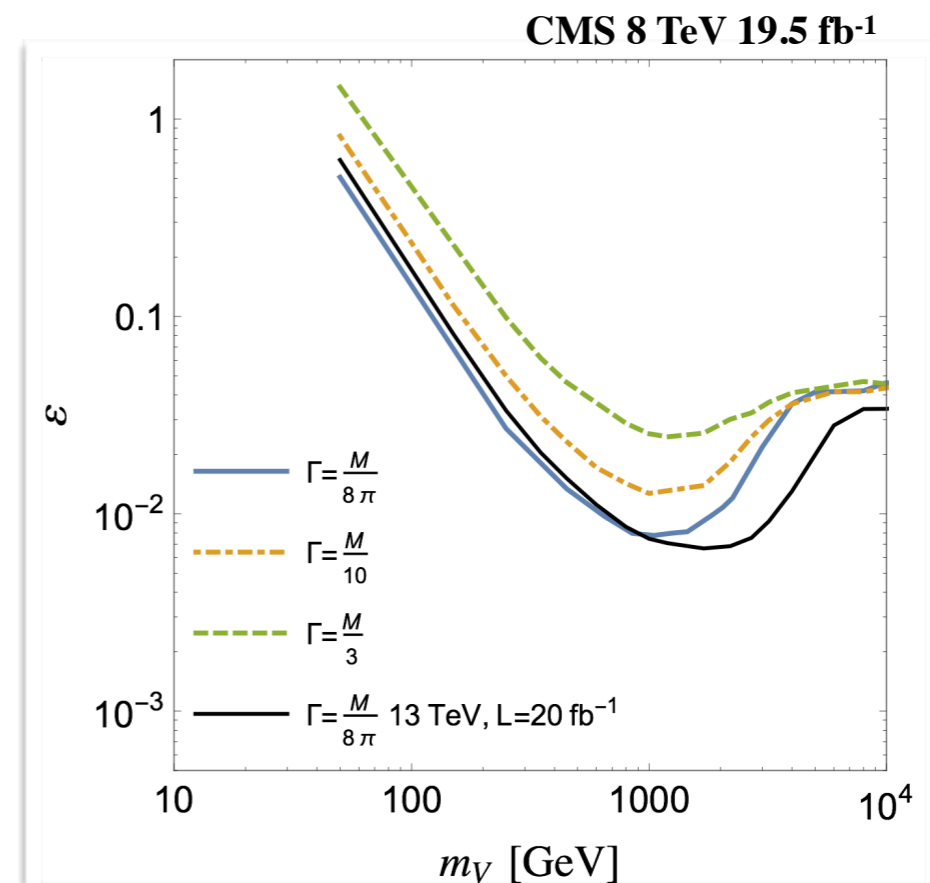
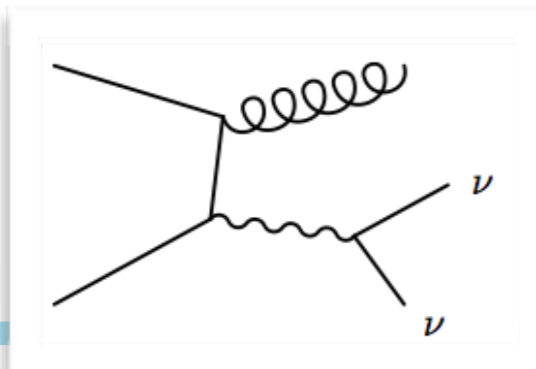
# Colliders



Systematic uncertainties are dominant



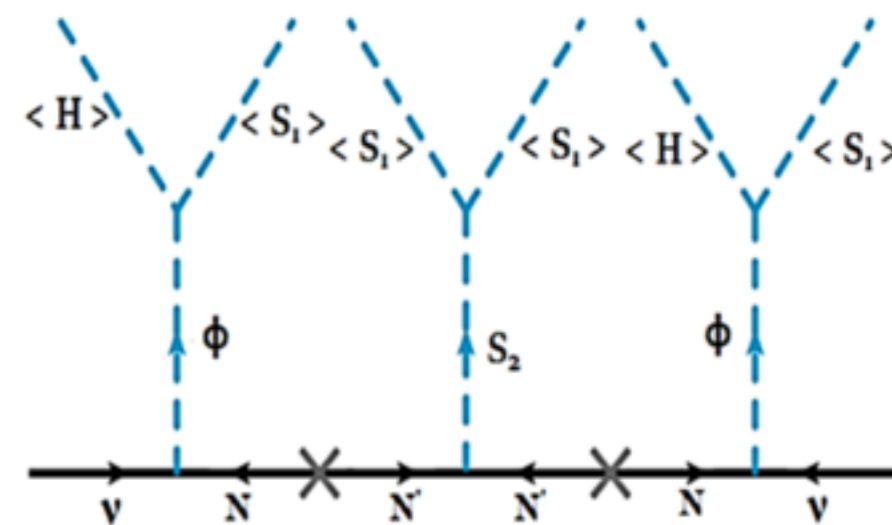
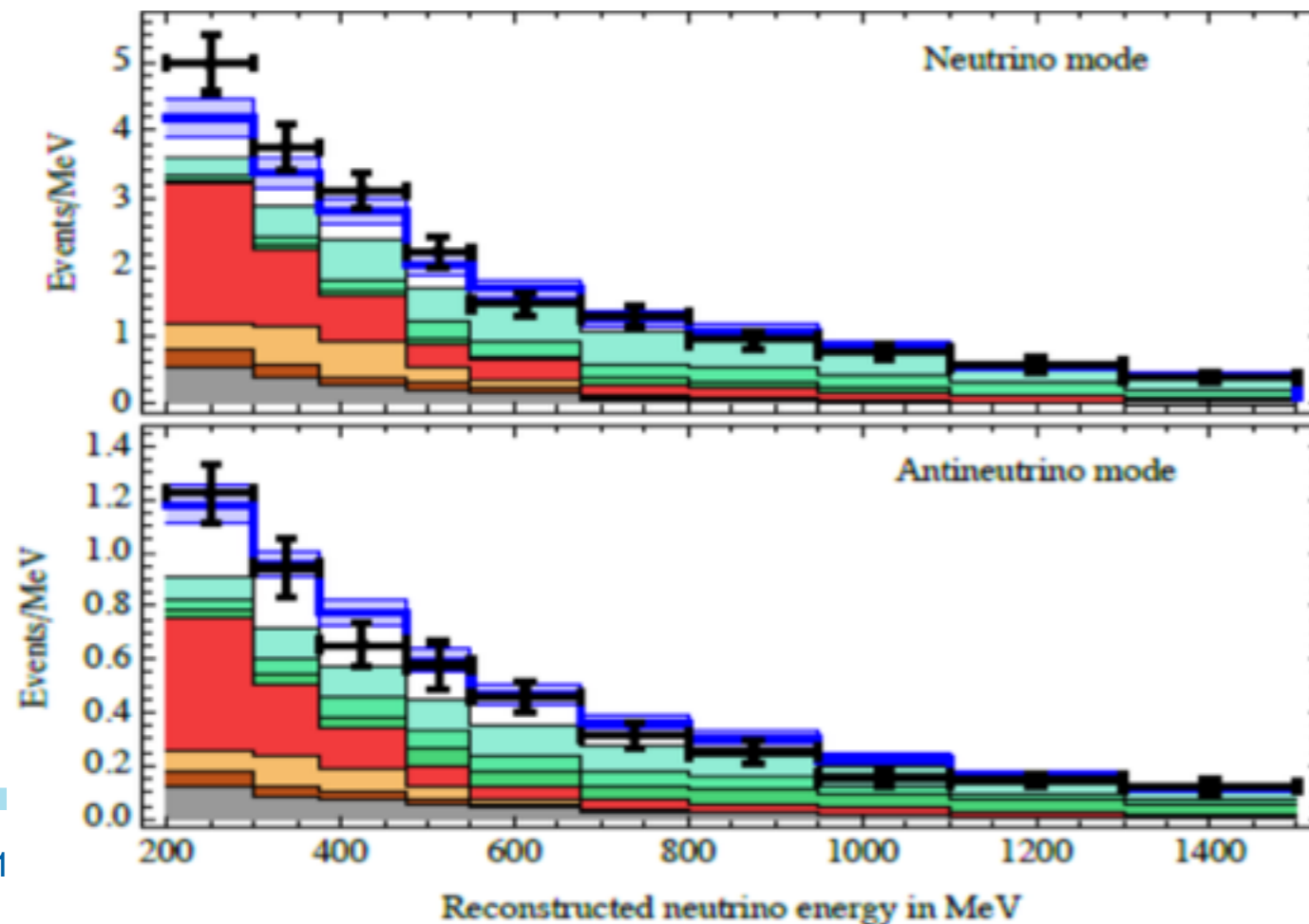
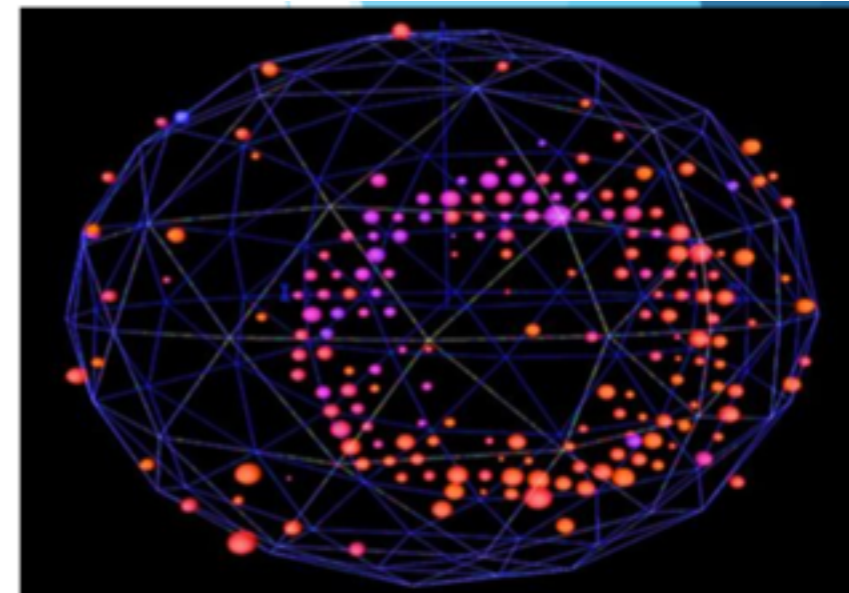
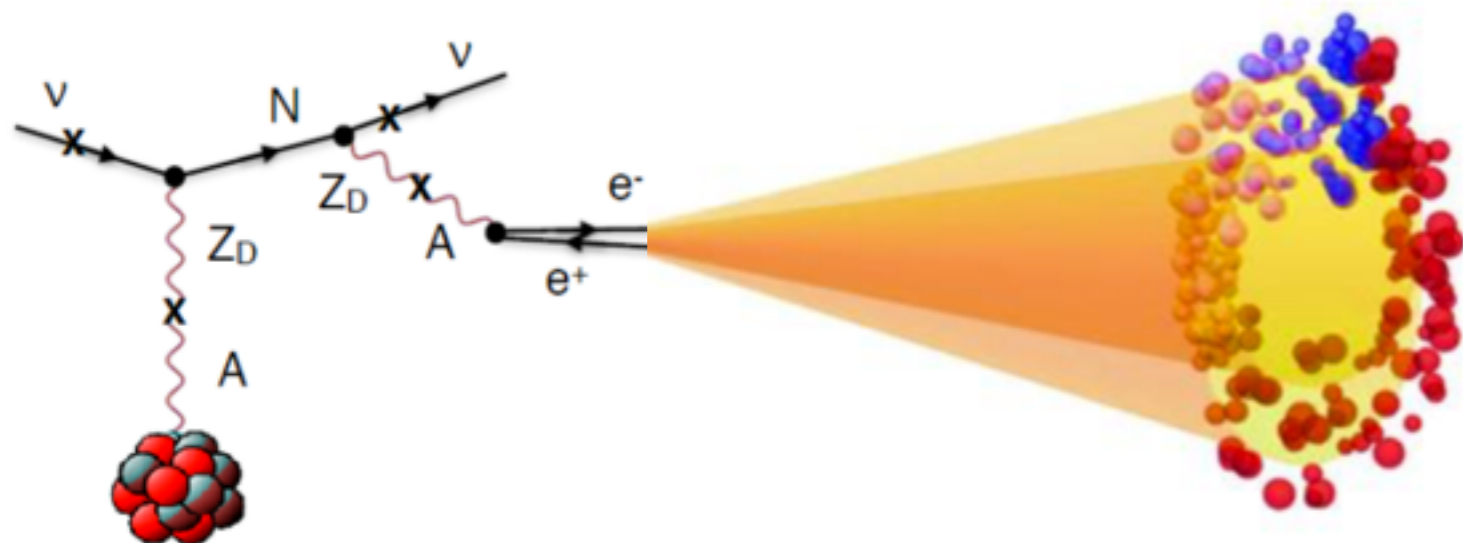
EFT vs simplified models?



# Light mediators

## ❖ Explanation of MiniBooNE's low energy excess

### A LIGHT DARK SECTOR – THE IDEA




# Light mediators





# Light mediators

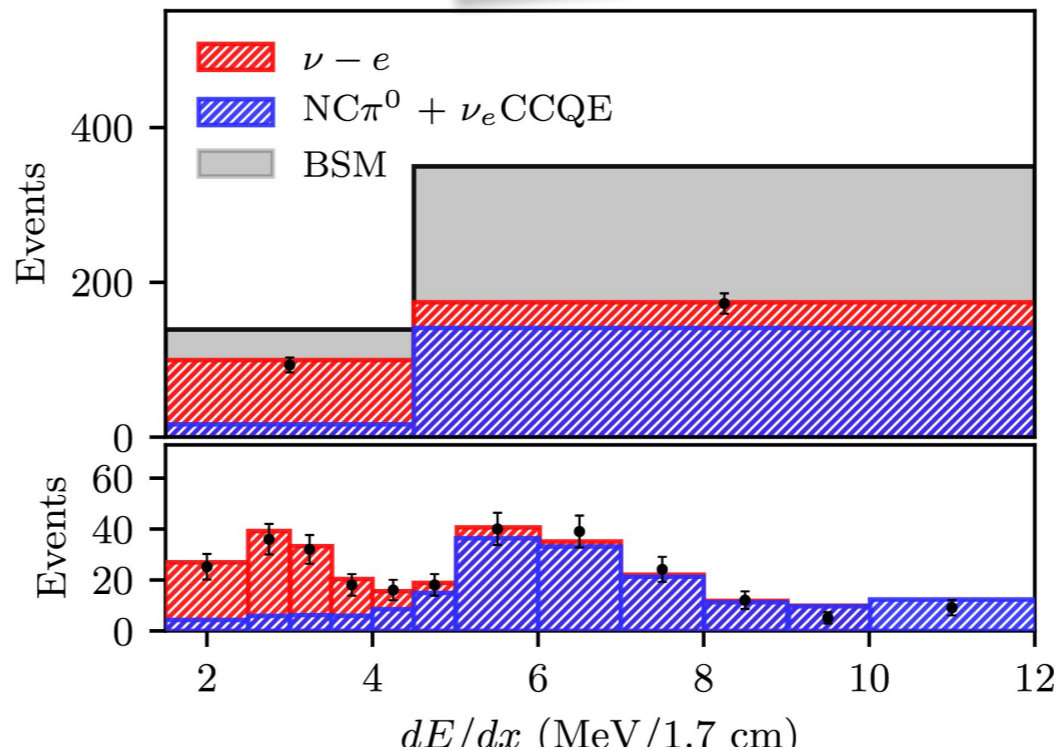
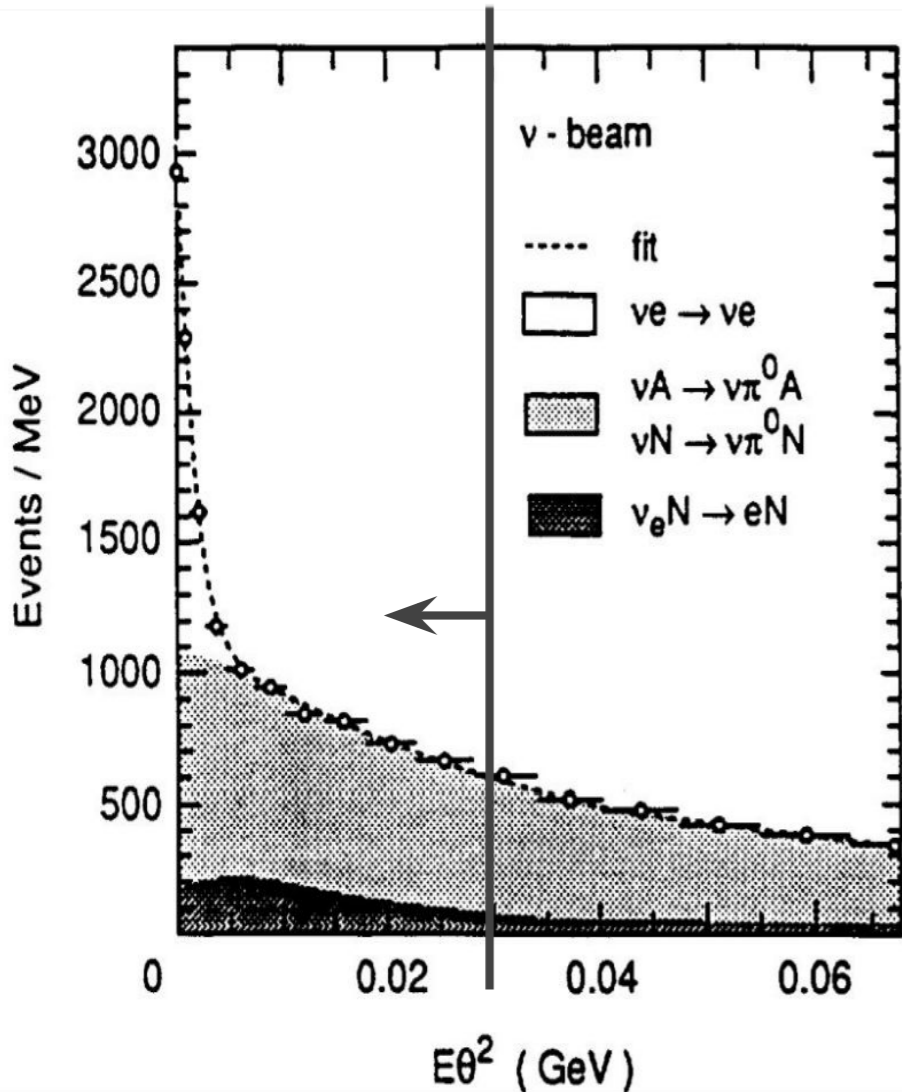
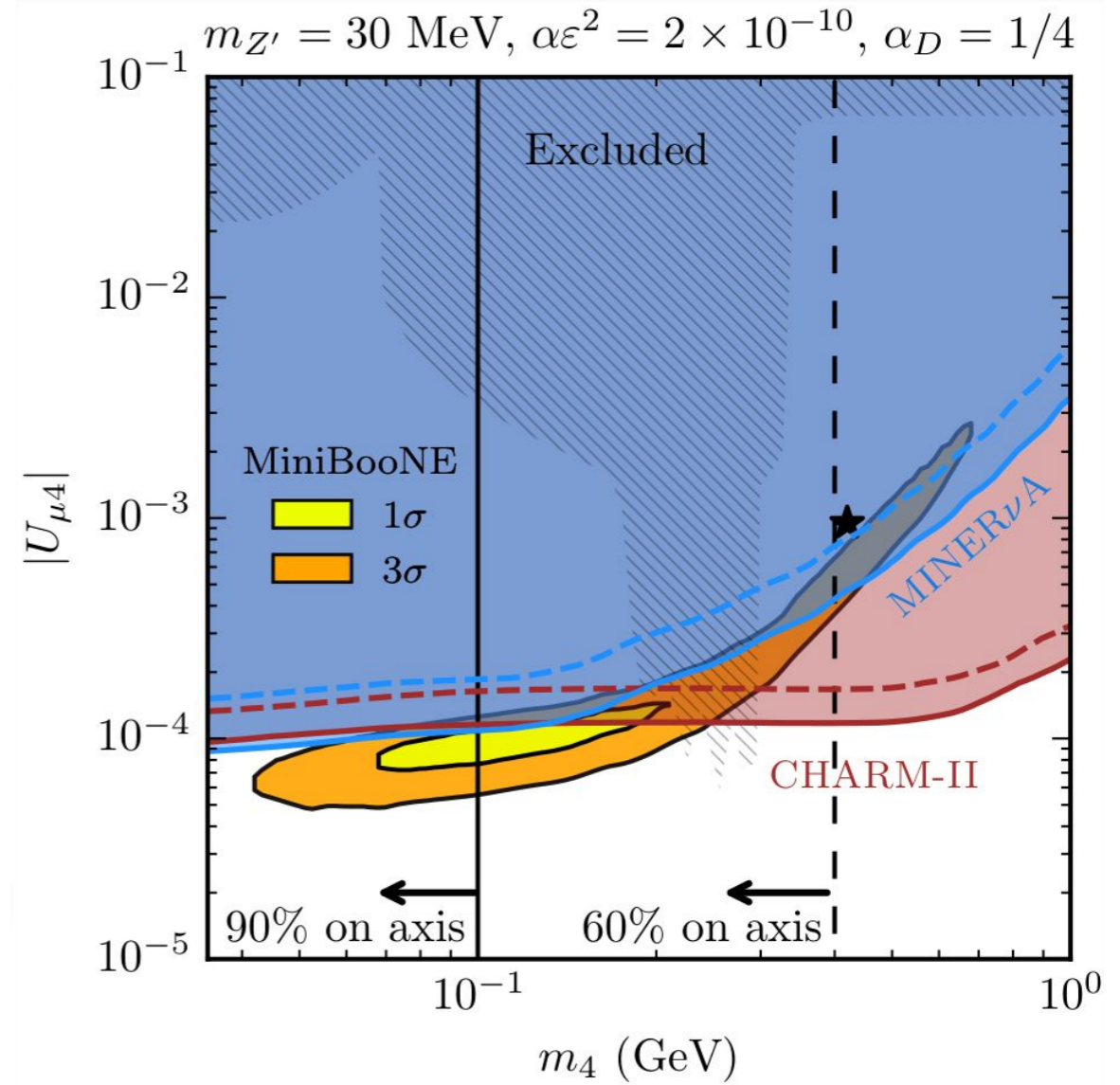
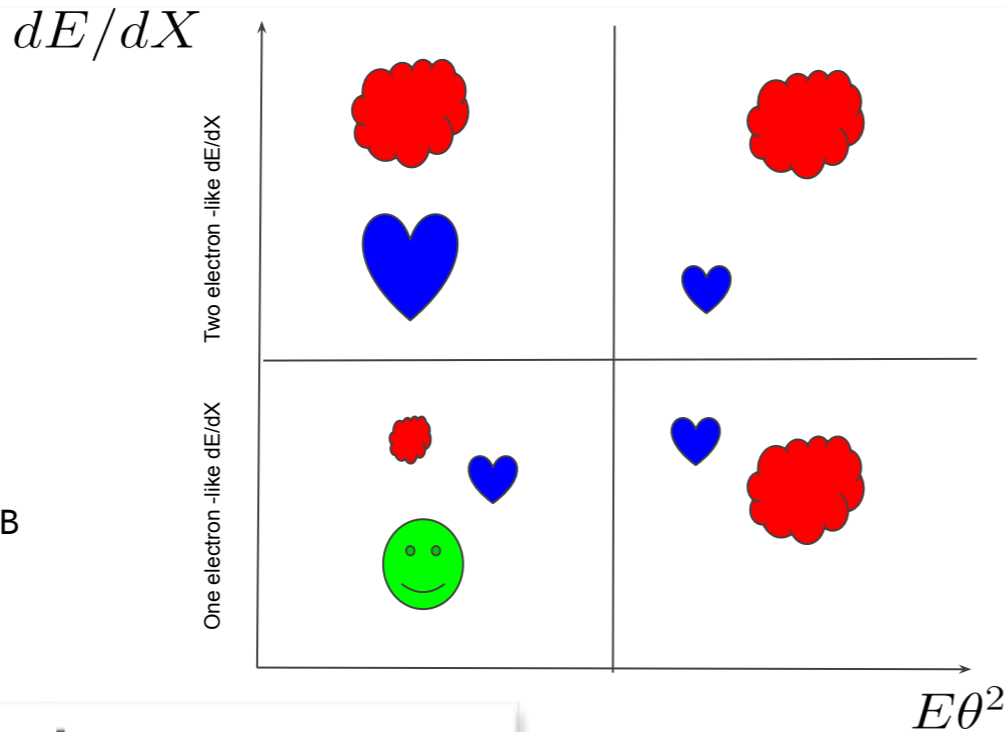


# Light mediators

 Background, e.g. NCπ<sup>0</sup>

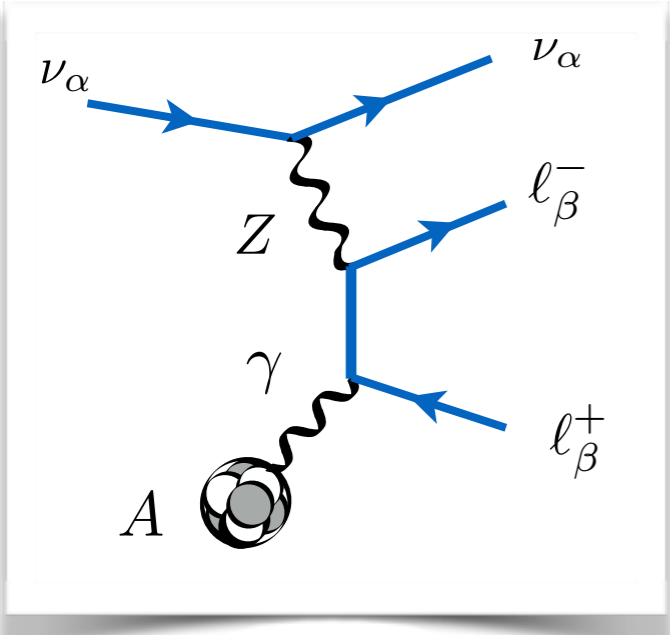
 Neutrino-electron scat.

 Recent BSM-MB explanations



# New signatures Light mediators

## Neutrino trident production



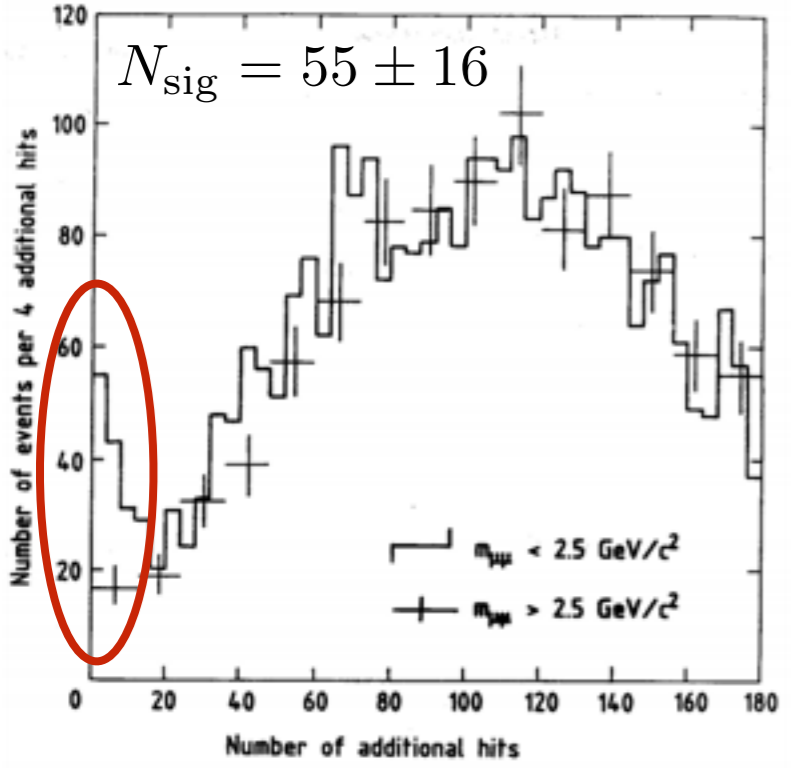
$e^+ e^-$        $\mu^\pm e^\mp$

Not yet observed or even searched for!

### CHARM-II

WANF beam (CERN)

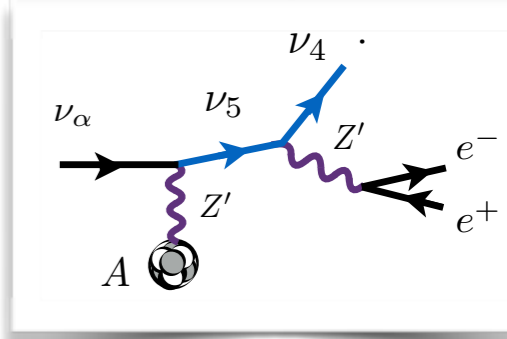
$\langle E_\nu \rangle = 25 \text{ GeV}$



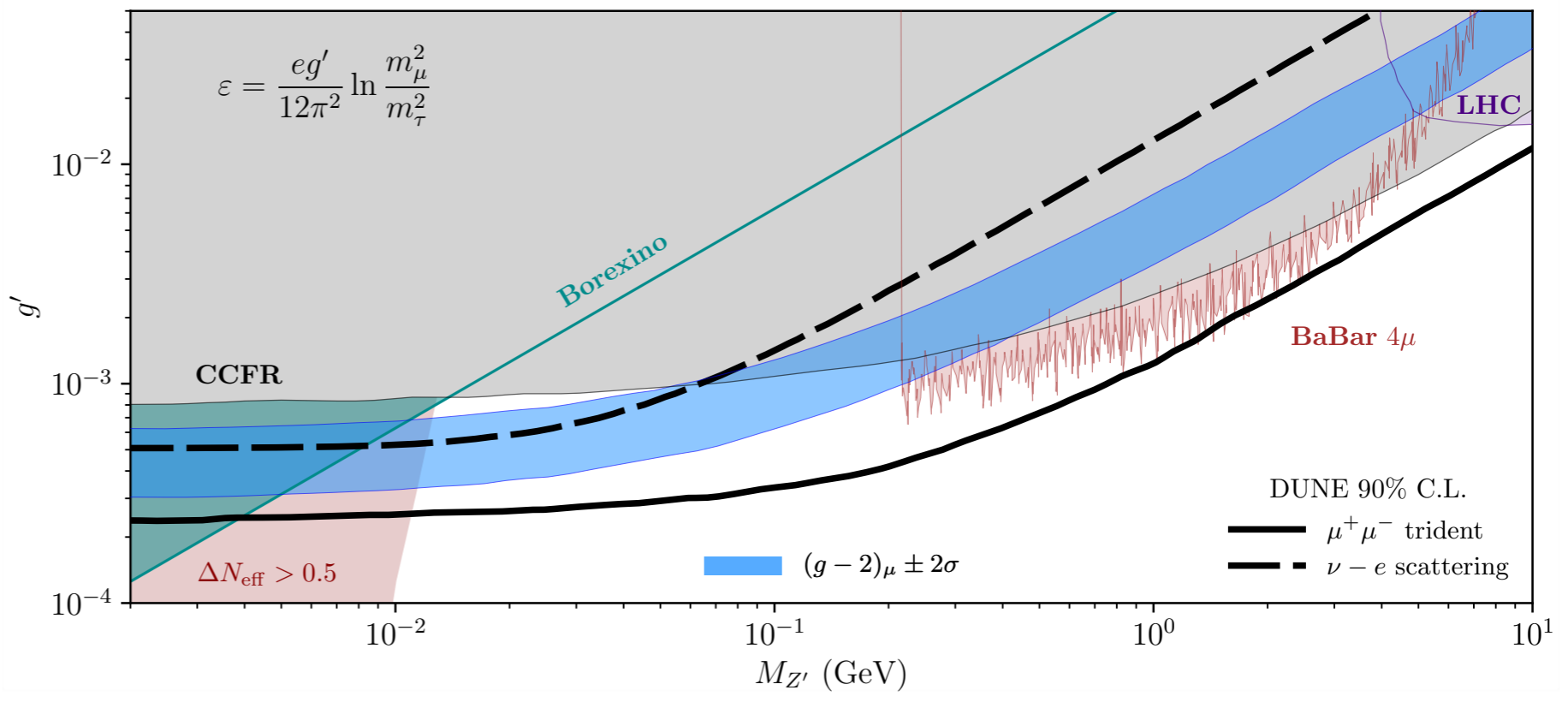
$$\frac{\sigma_{\text{CHARM-II}}}{\sigma_{\text{SM}}} = 1.58 \pm 0.57$$



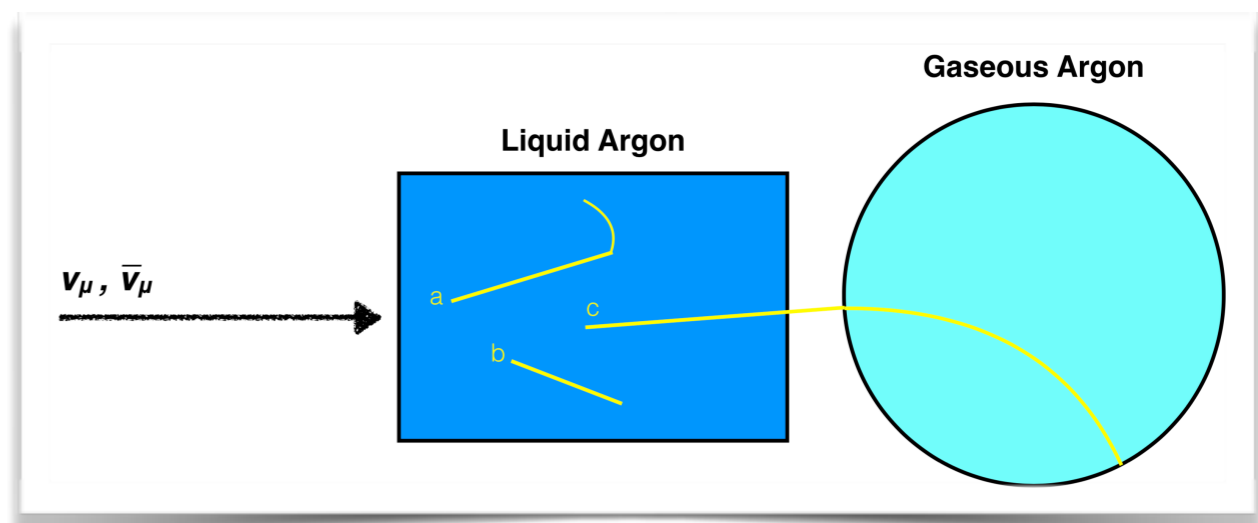
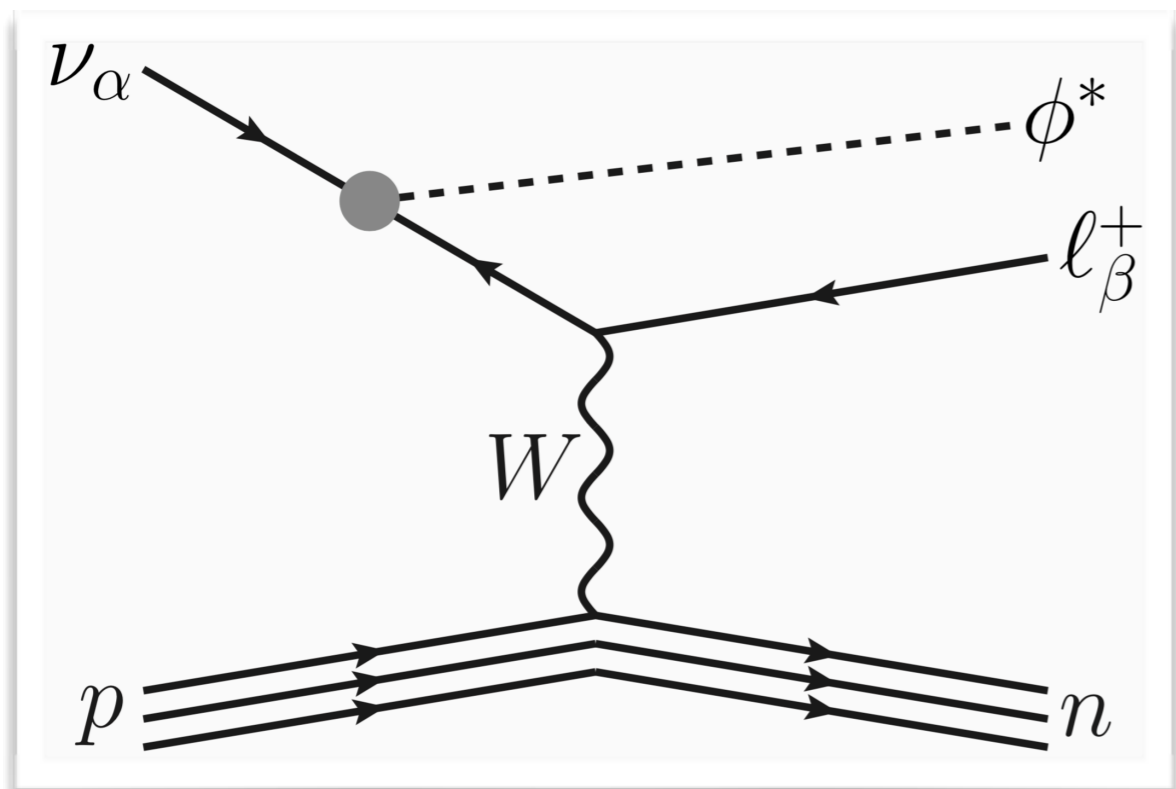
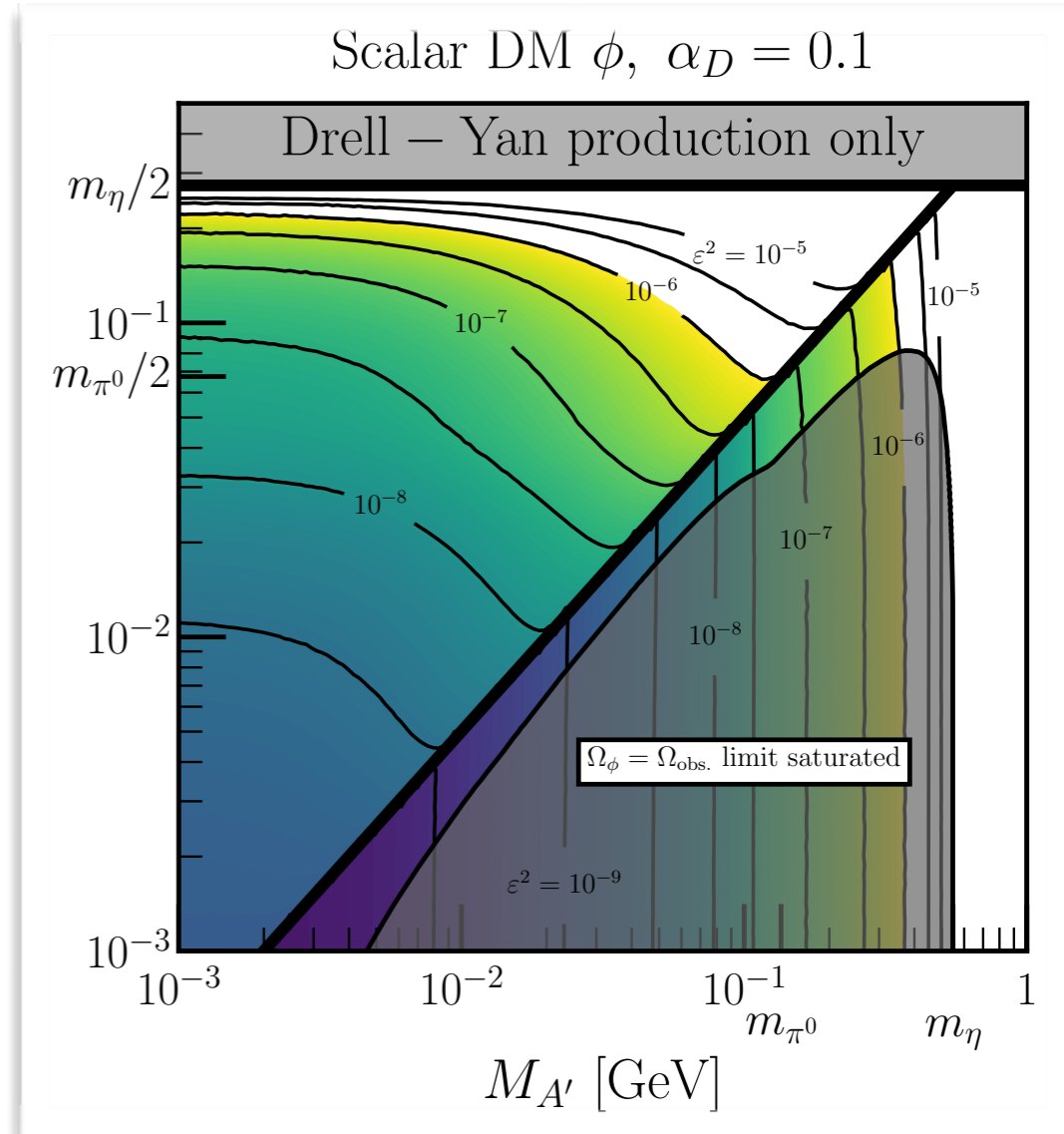
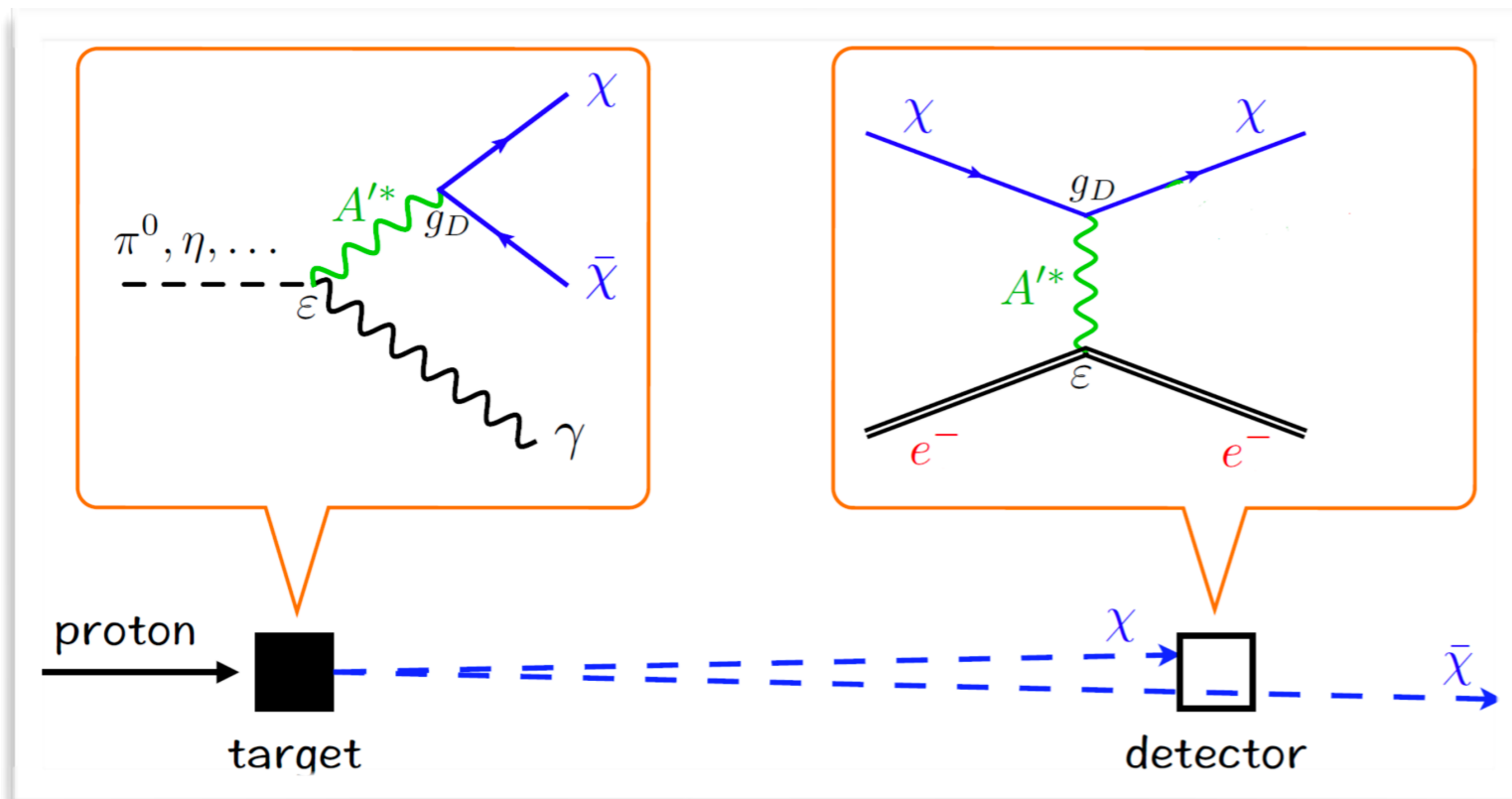
## Smiting New Physics



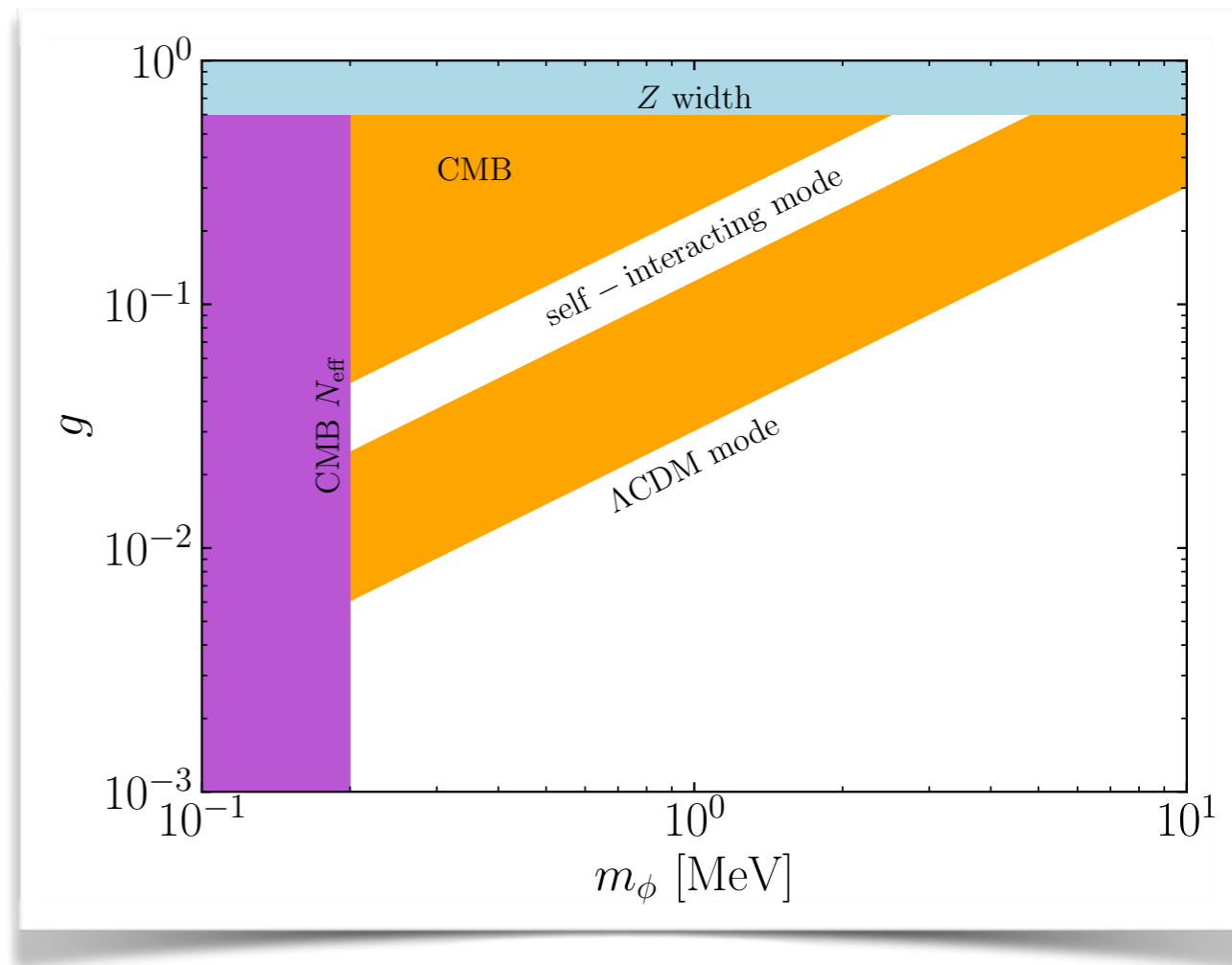
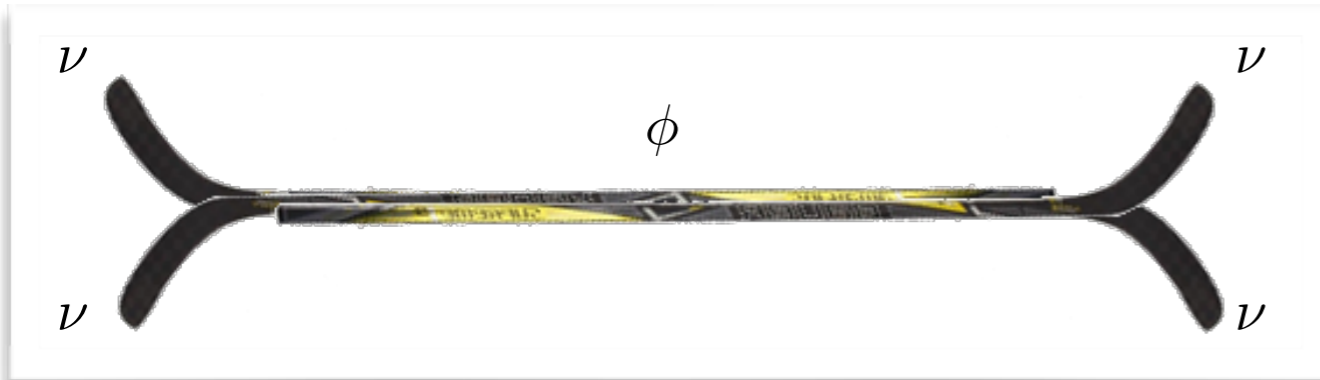
$L_\mu - L_\tau$ , DUNE ND, 75 tonnes, 5 y  $\nu$ -mode + 5 y  $\bar{\nu}$ -mode, 120 GeV  $p^+$ ,  $\sigma_{\text{norm}} = 5\%$



# New signatures Dark matter



# Light mediators and more



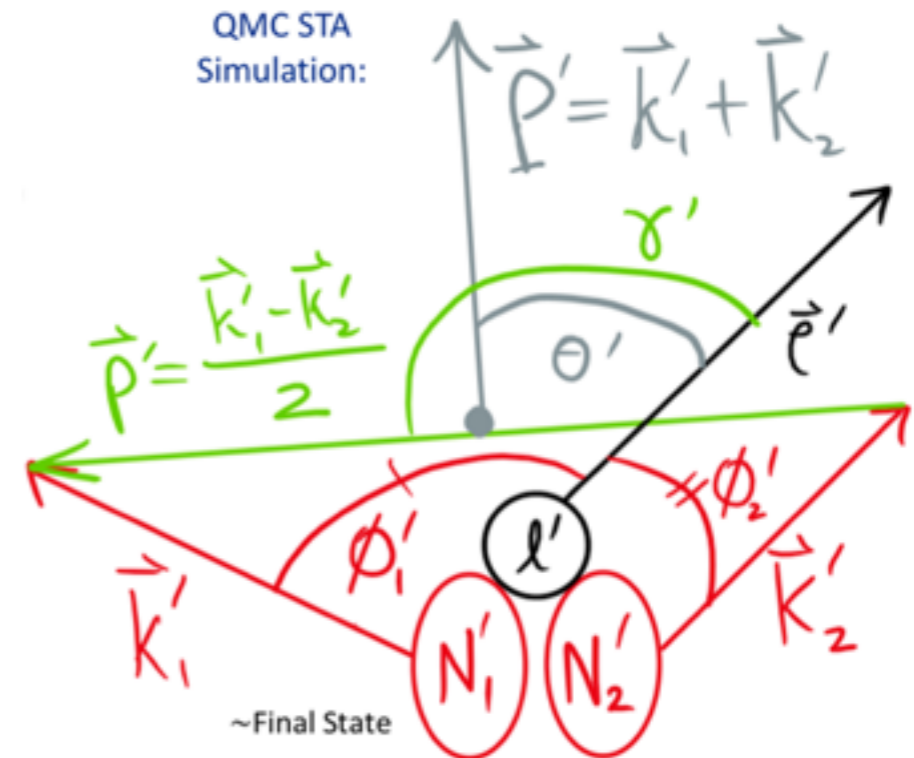
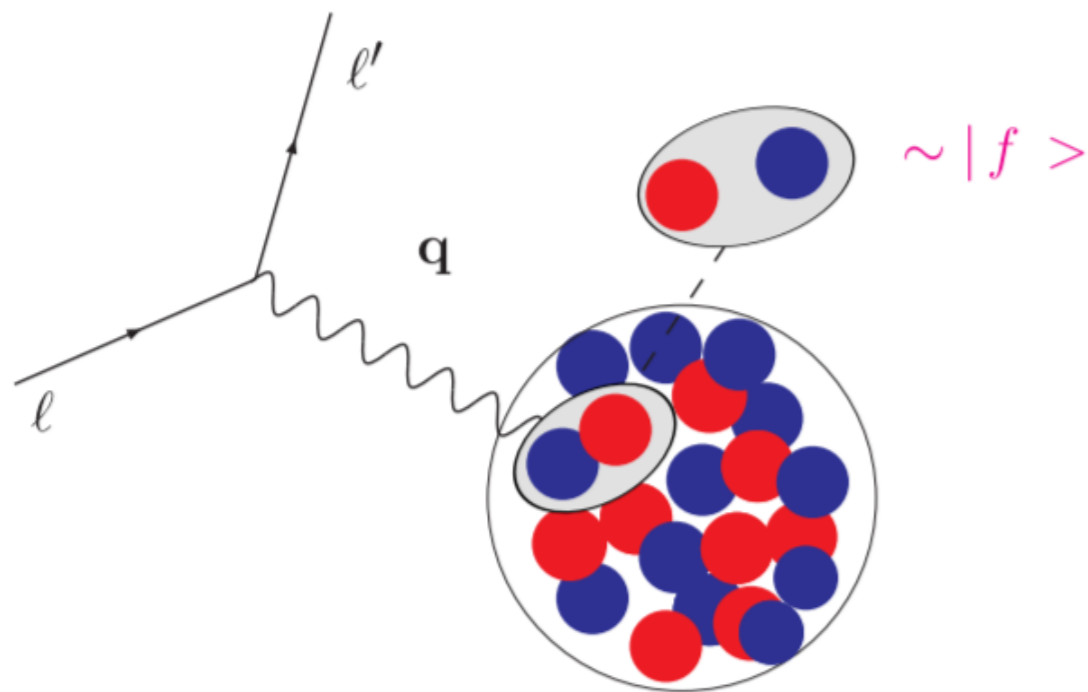
Early Universe Physics  
Inflation Models: Natural Inflation  
C $\nu$ B Scattering  
 $H_0$  Tension





# The Inclusion of Two-Body Physics: Nuclear Response

- The second order correction to a Hamiltonian describing a system of bound nucleons comes from two-body interaction terms in a high-order expansion:

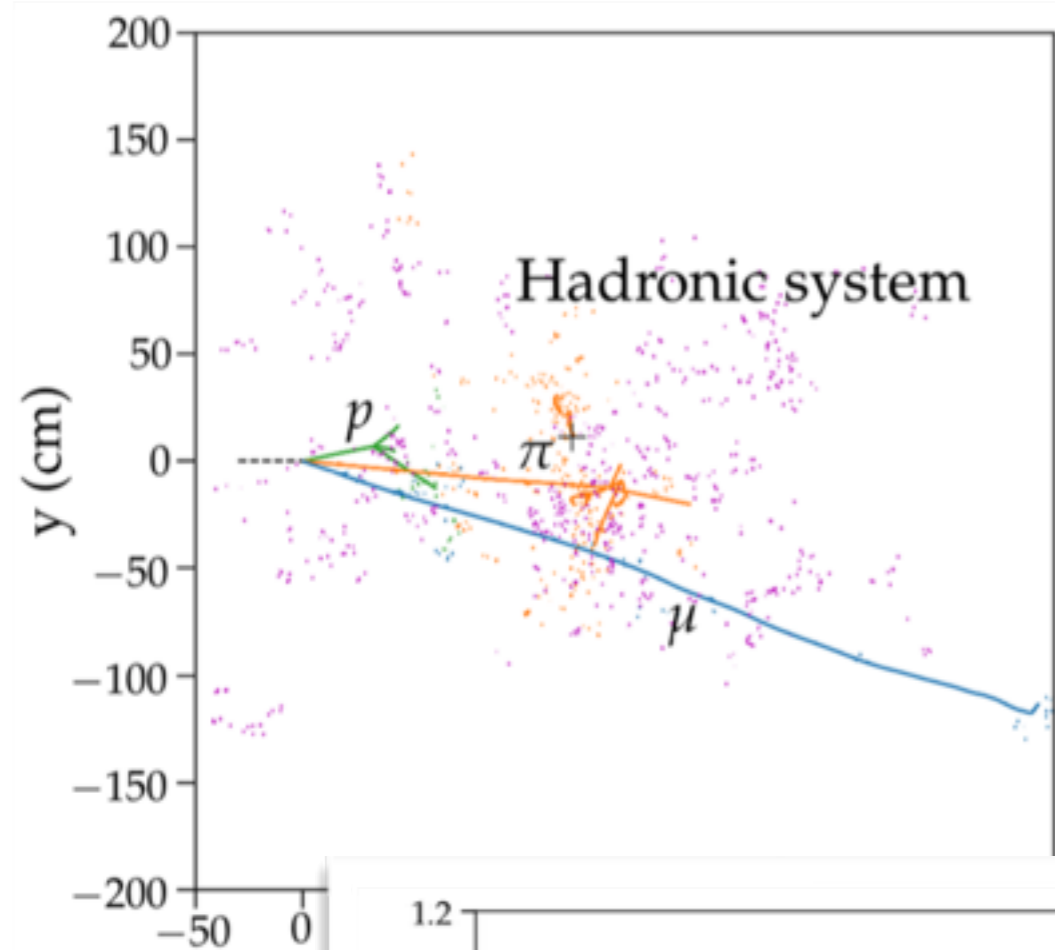


## Future Generator Validation

- Ample amounts of data are available for  $e^4_2\text{He}$  and  $e^{12}_6\text{C}$  scattering
- Once the generator is complete, tested, and validated on  $e$  data, we will proceed to  $\nu$  generation
- Publication will follow soon after

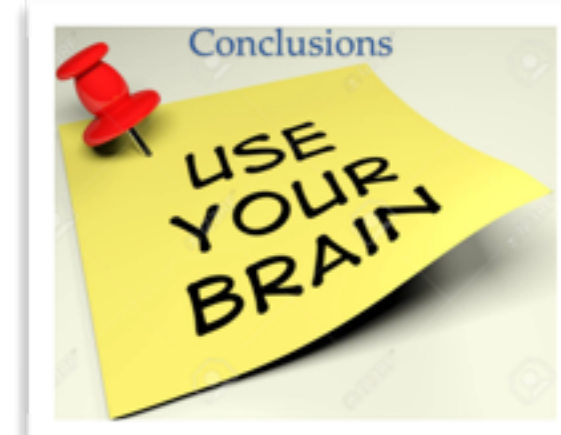
And much more

# What do particles do



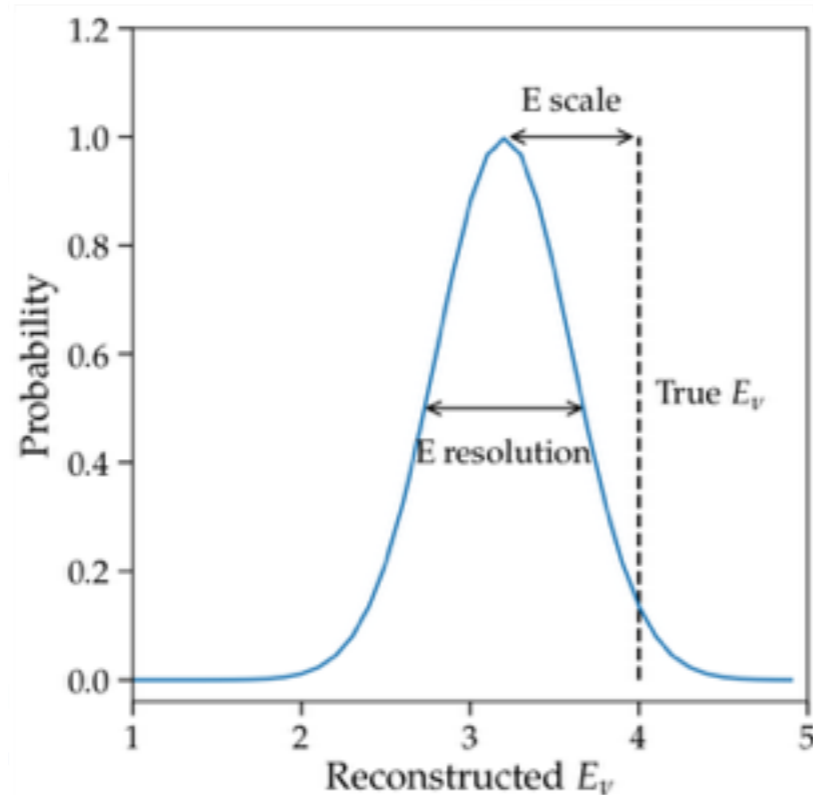
Particle type	Detection Threshold (KE)
$\mu^\pm$	30 MeV
$\pi^\pm$	100 MeV
$e^\pm/\gamma$	30 MeV
p	50 MeV
n	50 MeV
other	50 MeV

**DON'T TRUST ANYONE**



Mailing list:

Help with neutrino generators and propagation in liquid argon



**Shirley Li**

to me ▾

Carlos is great with names! His ideas:

[hitchhiker-guide@xx](mailto:hitchhiker-guide@xx)

[hitchhiker@xx](mailto:hitchhiker@xx)

[Pheno-Dont-panic@fnal](mailto:Pheno-Dont-panic@fnal)

**milab**

[@fnal.gov](http://@fnal.gov)

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

Mass and flavor models: Why do we want to push them to lower energies?

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

Mass and flavor models: Why do we want to push them to lower energies?

Leptogenesis: How to disprove it?

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

Mass and flavor models: Why do we want to push them to lower energies?

Leptogenesis: How to disprove it?

More general EFT-NSIs: How realistic are those in UV models?



# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

Mass and flavor models: Why do we want to push them to lower energies?

Leptogenesis: How to disprove it?

More general EFT-NSIs: How realistic are those in UV models?

Collider and neutrinos: How to clearly show their complementarity? Do we need more collaboration between these communities?

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

Mass and flavor models: Why do we want to push them to lower energies?

Leptogenesis: How to disprove it?

More general EFT-NSIs: How realistic are those in UV models?

Collider and neutrinos: How to clearly show their complementarity? Do we need more collaboration between these communities?

Generators: Can they be made more simple so they can be used as a tool for theorists? Will they keep up with new strategies/ideas/signatures???

# Few questions and comments

Traditional EFT NSI below weak scale: Are we happy with it?

UV complete models for NSI: proof of principle or should we focus on them?

Flavor models: Can we ever test the discrete symmetry based one?

Mass and flavor models: Why do we want to push them to lower energies?

Leptogenesis: How to disprove it?

More general EFT-NSIs: How realistic are those in UV models?

Collider and neutrinos: How to clearly show their complementarity? Do we need more collaboration between these communities?

Generators: Can they be made more simple so they can be used as a tool for theorists? Will they keep up with new strategies/ideas/signatures???

What do we need from EXPs to perform better analysis and propose new searches?

We plan on writing a report about this workshop.

**Background:** NSI community is changing. Several efforts on more coherent framework, new observables, UV complete models, complementarity with other experiments, greater precision, ...

**Goal:** Write a report to identify the recent progress we achieved and the new directions we are pursuing.

Each of you should write a 3 pages (maximum) summary related to your talk (including refs). We will put this all together, edit, and write intro and conclusions. You will get a more detailed email soon.

**Deadline: June 15th**

We plan on writing a report about this workshop.

**Background:** NSI community is changing. Several efforts on more coherent framework, new observables, UV complete models, complementarity with other experiments, greater precision, ...

**Goal:** Write a report to identify the recent progress we achieved and the new directions we are pursuing.

Each of you should write a 3 pages (maximum) summary related to your talk (including refs). We will put this all together, edit, and write intro and conclusions. You will get a more detailed email soon.

**Deadline: June 15th**

*What a great workshop!  
Thanks to all of you for coming here!*