# Electroweak and QCD effects in the W boson mass determination

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1st COFI Workshop on Precision Electroweak 16-19 July 2019, San Juan, Puerto Rico Electroweak precision measurements at hadron colliders

•  $M_W$ 

•  $\sin^2 \vartheta^{\ell}_{eff}$ 

# measured through C.C. and N.C. Drell-Yan

previous talk by S. Camarda for experimental status

 we have to remind that their most precise values come from EWPO @LEP1

talk by A. Freitas

# $M_W$ calculated in the Standard Model

$$M_W^2 = \frac{M_Z^2}{2} \left\{ 1 + \left[ 1 - \frac{4\pi\alpha}{\sqrt{2}G_\mu M_Z^2} \left( 1 + \Delta r \right) \right]^{1/2} \right\}$$
$$M_W^2 = 80.357 \pm 0.009 \pm 0.003 \text{GeV}$$

• one loop  $\mathcal{O}(\alpha)$  calculation

A. Sirlin, PRD22 (1980) 971

• two loop  $\mathcal{O}(\alpha \alpha_s)$ 

A. Djouadi, C. Verzegnassi, PLB195 (1987) 265

• three loop  $\mathcal{O}(\alpha \alpha_s^2)$ 

L. Avdeev et al., PLB336 (1994) 560;

K.G. Chetyrkin, J.H. Kühn, M. Steinhauser, PLB351 (1995) 331; PRL75 (1995) 3394

•  $\mathcal{O}(\alpha^2)$  for large top / Higgs mass

R. Barbieri et al., PLB288 (1992) 95; NPB409 (1993) 105

G. Degrassi, P. Gambino, A. Vicini, PLB383 (1996) 219

• exact  $\mathcal{O}(\alpha^2)$ A. Freitas et al., PLB495 (2000) 338; NPB632 (2002) 189 M. Awramik, M. Czakon, PLB568 (2003) 48; PRL89 (2002) 241801 A. Onishchenko, O. Veretin, PLB551 (2003) 111; M. Awramik et al., PRD68 (2003) 053004

G. Degrassi, P. Gambino, P.P. Giardino, JHEP 1505 (2015) 154; I. Dubovik et al., arXiv:1906.08815

$$\sin^2 \vartheta_{eff}^l = \frac{1}{4} \left( 1 - \operatorname{Re} \frac{g_v}{g_a} \right), \qquad \operatorname{Zl}\bar{l} \operatorname{vertex} \sim \bar{l} \gamma^{\mu} (g_v - g_a \gamma_5) l Z_{\mu}$$

- measured at Z peak:  $0.23153 \pm 0.00016$
- uncertainty in the Standard Model calculations:  $\sim 0.00005 \oplus 0.00004$ 
  - at one loop  $\mathcal{O}(\alpha)$

A. Sirlin, PRD22, (1980) 971, W.J. Marciano, A. Sirlin, PRD22 (1980) 2695

G. Degrassi, A. Sirlin, NPB352 (1991) 352, P. Gambino and A. Sirlin, PRD49 (1994) 1160

at higher orders:

 $\star \mathcal{O}(\alpha \alpha_s)$ 

A. Djouadi, C. Verzegnassi, PLB195 (1987) 265 B. Kiehl, NPB353 (1991) 567; B. Kniehl, A. Sirlin, NPB371 (1992) 141, PRD47 (1993) 883

A. Djouadi, P. Gambino, PRD49 (1994) 3499

 $\star \mathcal{O}(\alpha \alpha_s^2)$ 

L. Avdeev et al., PLB336 (1994) 560;

Chetyrkin, Kühn, Steinhauser, PLB351 (1995) 331; PRL75 (1995) 3394; NPB482 (1996) 213

 $\star \mathcal{O}(\alpha \alpha_s^3)$ 

**\*** exact  $\mathcal{O}(\alpha^2)$ 

Y. Schröder, M. Steinhauser, PLB622 (2005) 124;

K.G. Chetyrkin et al., hep=ph/0605201; R. Boughezal, M. Czakon, hep-ph/0606232

★  $\mathcal{O}(\alpha^2)$  for large Higgs / top mass

G. Degrassi, P. Gambino, A. Sirlin, PLB394 (1997) 188

M. Awramik, M. Czakon, A. Freitas, JHEP0611 (2006) 048

W. Hollik, U. Meier, S. Uccirati, NPB731 (2005) 213; I. Dubovik et al., arXiv:1906.08815

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# direct $M_W$ measurements at LEP2: $80.376 \pm 0.033$ MeV

ADLO and LEP EWWG arXiv:1302.3415

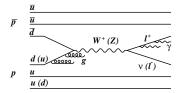
• cross section dependence on  $M_W$  at threshold  $\sim \beta = \sqrt{1 - \frac{4M_W^2}{s}}$ 

•  $M_W = 80.42 \pm 0.20 \pm 0.03 (E_{\text{LEP}}) \text{ GeV}$ 

- direct reconstruction of decay products in  $e^+e^- \rightarrow 4$  fermions
  - $q\bar{q}'q\bar{q}'$
  - $q\bar{q}'\ell\nu_\ell$

• 
$$M_W = 80.375 \pm 0.025 (\text{stat}) \pm 0.022 (\text{syst}) \text{ GeV}$$

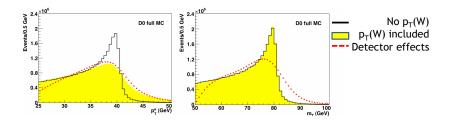
# direct $M_W$ measurement at hadron colliders



$$\begin{split} \sigma^{\text{theory}} &\equiv \sum_{a,b} \int_0^1 dx_1 dx_2 f_{a,H_1}(x_1,\mu_F^2,\mu_R^2) f_{b,H_2}(x_2,\mu_F^2,\mu_R^2) \times \\ &\times \int_{\Phi} d\hat{\sigma}_{a,b}(x_1,x_2,Q^2/\mu_F^2,Q^2/\mu_R^2) + \ \mathcal{O}\left(\frac{\Lambda_{QCD}^n}{Q^n}\right) \end{split}$$

- extremely important that the direct determinations of  $M_W$  and  $\sin^2 \vartheta^\ell_{eff}$  is be as much as possible independent of each other, also from a theoretical point of view
- $M_W$  and  $\sin^2 \vartheta_{eff}^{\ell}$  are tightly intertwined by with other electroweak parameters in the gauge sector (quite different w.r.t.  $m_{\text{top}}$  and  $m_H$ ) F. Piccinini (INFN Pavia) EW and QCD on  $M_W$  July 2019 6 / 32

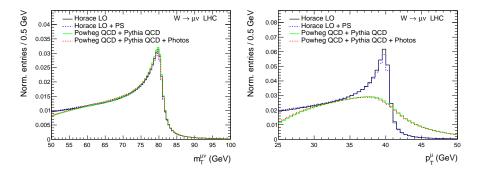
# Main relevant observables for $M_W$



### • $M_{\perp}$

- theoretically stable vs radiative corrections
- determination of neutrino momentum exp. challenging
- $p_{\perp}^{\ell}$ 
  - experimentally clean
  - theoretically very sensitive to ISR

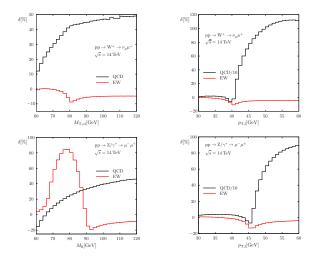
# sensitivity to radiative corrections



C.M. Carloni Calame et al., arXiv:1612.02841

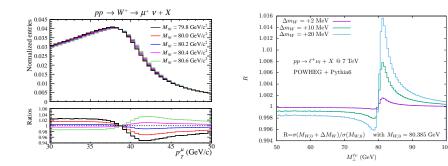
• crucial a good control of  $p_{\perp}^W$  (which affects directly  $p_{\perp}^\ell$ )

# NLO QCD and NLO EW



S. Dittmaier, A. Huss, C. Schwinn, arXiv:1403.3216

# sensitivity to $M_W$



Farry, Lupton, Pili, Vesterinen, arXiv:1902.04323

by A. Vicini

### • control of shapes below 1% scale for $\Delta M_W \sim 10-20$ MeV

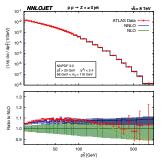
lepton pair (Z/W)  $p_{\perp}$ : two regimes

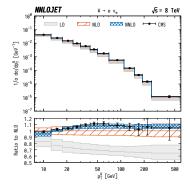
large p⊥ (≥ 20 GeV), where pert. th. is reliable
 state of the art is NNLO QCD

• small  $p_{\perp}$  ( $\lesssim 20$  GeV): ~90% of the cross section

- resummation of  $\log\left(\frac{M_V}{q_\perp}\right)$  is needed
- sensitivity to the non-perturbative model of the MC Evt Gen

### Large $p_{\perp}$ region





- A. Gehrmann-De Ridder et al., arXiv:1605.04295
- A. Huss,  $p_T^{\,Z}$  and  $p_T^{\,W}$  theory meeting, CERN 2018
- R. Boughezal et al., 1512.01291, 1602.05612, 1602.08140

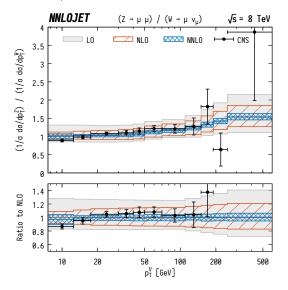
### yellow dash: ew corrections

A. Gehrmann-De Ridder et al., arXiv:1712.07543 A. Huss at  $p_T^Z$  and  $p_T^W$  theory meeting, CERN 2018 R. Boughezal et al., arXiv:1602.06965

A. Denner et al., arXiv:1103.0914

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### and on the ratio W/Z



A. Gehrmann-De Ridder et al., arXiv:1712.07543

# Small $p_{\perp}$ region: resummation techniques

- recent progress by different groups on resummation
  - $q_{\perp}$  resummation in impact parameter space
    - \* DYRES, DYTURBO S. Catani et al., arXiv:1507.06937; G. Ferrera, S. Camarda
    - ★ ReSolve T. Cridge and F. Coradeschi
    - ★ Resbos2, CSS formalism J. Isaacson
  - SCET based resummation
    - \* GENEVA, SCETIIb S. Alioli et al., arXiv:1211.7049, arXiv:1508.01475; F. Tackmann et al
    - CuTe
       T. Becher et al., arXiv:1109.6027, arXiv:1212.2621

T. Becher, Hager, arXiv:1904.08325

- resummation in direct space (RadISH) W. Bizon et al., arXiv:1705.09127; 1604.02191
- resummation throuh TMD factorisation (NangaParbat) V. Bertone and G. Bozzi
- recent progress in Monte Carlo generators
  - inclusion of NLO splitting kernels (DIRE)

S. Höche, F. Krauss, S. Prestel, 1705.00982

S. Höche, S. Prestel, arXiv:1705.00742

- DY at NNLOPS accuracy with different methods
  - ★ MiNLO with POWHEG
- Karlberg, Re, Zanderighi, arXiv:1407.2940
- ★ UNNLOPS with SHERPA
- Höche, Li, Prestel, arXiv:1405.3607
- ★ GENEVA Alioli, Bauer, Berggren, Guns, Tackmann, Walsh, '15-'16

### • first investigations of possible flavour dependence of non-perturb

partonic intrinsic  $k_{\perp}$ 

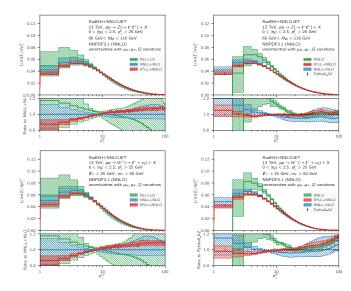
talk by A. Signori

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- different approaches, even if with the same nominal accuracy, can differ
  - subleading terms
  - different matching effects in the transition region
  - matching schemes (additive vs multiplicative)
  - non-perturbative corrections/MC tune
  - order of PDF evolution
  - thresholds and treatment of heavy quarks
- within the LHC EWWG, important benchmarking activity among different codes
  - this is the first time such an exercise is beeing performed, after the studies of s. Alioli, arXiv:1606.02330

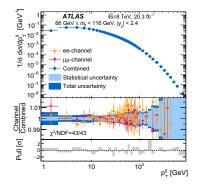
# latest results from RadISH+NNLOjet: N<sup>3</sup>LL+NNLO

W. Bizon et al., arXiv:1905.05171



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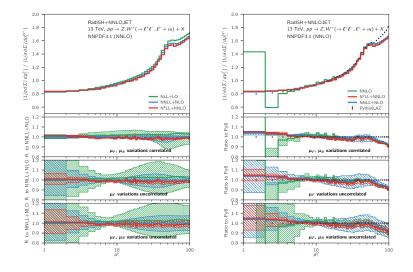


ATLAS coll., arXiv:1512.02192

$$p_{\perp}^{W} = (p_{\perp}^{Z})_{\text{measured}} \left(\frac{p_{\perp}^{W}}{p_{\perp}^{Z}}\right)_{\text{th}}$$

### • How to treat th uncertainties in numerator and denominator?

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W. Bizon et al., arXiv:1905.05171

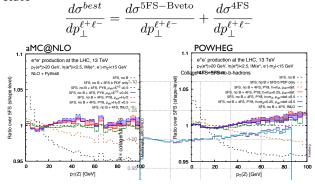
• stability of the best predictions vs (un)correlation of scales • remaining  $\mathcal{O}(\%)$  th. uncertainty

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# $b\bar{b}$ contribution to $p_{\perp}^{Z}$

E. Bagnaschi, F. Maltoni, A. Vicini and M. Zaro, arXiv:1803.04336

- two schemes
  - ▶ 5 flavour scheme (massless b, resummation of  $\log(p_{\perp}^b/m_b)$ )
  - ▶ 4 flavour scheme (finite  $m_b$  with exact kinematics of  $\ell^+ \ell^- b\bar{b}$ )
- improvement from a combination of the two schemes
  - $p_{\perp}^{Z}$  distribution split into 2 contributions: with and without B hadrons in final state



 $ightarrow \Delta M_W < 5\,\,{
m MeV}$  from  $p_\perp^\ell$  (with sensitivity to the fit window)

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Higher-order corrections (for  $M_W$  fit)

$$d\sigma = d\sigma_0 + d\sigma_{\alpha_s} + d\sigma_\alpha + d\sigma_{\alpha_s^2} + d\sigma_{\alpha\alpha_s} + d\sigma_{\alpha^2} + \dots$$

• multi-photon emission from the final state  $\rightarrow \delta M_W \simeq 10$  MeV for  $\mu \nu_\mu$  final state

Carloni Calame et al., PRD 69 (2004) 037301, JHEP 0710 (2007)

mixed QCD-EWK corrections

Dittmaier, Huss, Schwinn, NPB 885 (2014) 318, NPB 904 (2016) 216

NNLO EWK effects

C.M. Carloni Calame et al., Phys.Rev. D96 (2017) 093005

- EWK input scheme
- lepton pair emission

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# QCD-EWK interference

- ${\mbox{\circ}}$  the  ${\mathcal O}(\alpha\alpha_s)$  calculation involves as building blocks
  - ▶ NNLO virtual corrections at  $\mathcal{O}(\alpha \alpha_s)$  (not yet available)
    - ★ necessary two-loop master integrals (with m = 0 external particles and  $M_W = M_Z$ , or with one massive internal line)

R. Bonciani et al., arXiv:1604.08581; A. von Manteuffel and R.M. Schabinger, arXiv:1701.06583

- NLO EW corrections to  $l\bar{l}^{(')}$ + jet
- $\blacktriangleright$  NLO QCD corrections to  $l\bar{l}^{(')}+\gamma$
- double real contributions  $l\bar{l}^{(')} + \gamma + jet$
- ▶ PDF's with NNLO accuracy at  $\mathcal{O}(\alpha \alpha_s)$  (not yet available)

very recent progress on NNLO mixed QCD-QED ISR corrections

talk by D. De Florian

### what is available:

- fixed order dominant  $\mathcal{O}(\alpha_s \alpha)$  corrections to DY in pole approximation Dittmaier. Huss. Schwinn, NPB 885 (2014) 318, NPB 904 (2016) 216
- $\blacktriangleright$  Monte Carlo estimates through NLO QCD  $\otimes$  NLO EW (with higher orders)

L. Barzè et al., JHEP 1204 (2012) 037, Eur. Phys. J. C73 (2013) 2474

# fixed order $\mathcal{O}(\alpha_s \alpha)$ in pole approximation

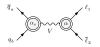
- two main classes of contributions:
  - factorizable
  - non-factorizable



(a) Factorizable initial-initial corrections



(c) Factorizable final-final corrections



(b) Factorizable initial-final corrections



(d) Non-factorizable corrections S. Dittmaier, A. Huss and C. Schwinn, arXiv:1601.02027

- a) not known but expected to be small recent progress covered by D. De Florian  $\begin{array}{c} (\mathcal{O}(\alpha) \text{ corrections in PA} \Longrightarrow M_{\perp} \text{ and } M(l^+l^-) \text{ insensitive to QED ISR} \\ \text{ in addition } M_{\perp} \text{ and } M(l^+l^-) \text{ mildly affected by NLO QCD corrections} \end{array}$
- b) this gives the bulk of the contribution
- c) no real contributions  $\implies$  no impact on shape of  $M_{\perp}$  and  $M(l^+l^-)$
- d) numerical impact below 0.1%

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POWHEG\_W\_ew\_BW, CC DY

OWHEG\_Z\_ew\_BMNNPV, NC DY

 $\mathcal{O}(\alpha_s \alpha)$  corrections through Monte Carlo

independent implementation

Barzè et al, JHEP 1204 (2012) 037

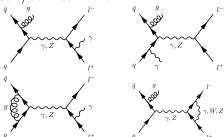
Bernaciak and Wackeroth, PRD 85 (2012) 093003

Barzè et al, EPJC 73 (2013) 6, 2474

Mück and Oymanns, JHEP 1705 (2017) 090

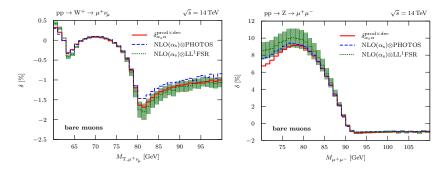
 correctly taken into account the NLO contribution with one additional radiation in the soft/collinear limit

The POWHEG-BOX includes NLO QCD & EW corrections interfaced



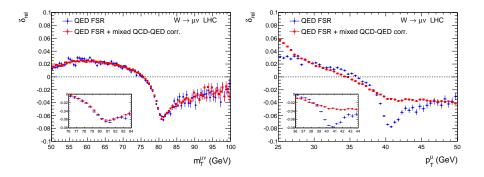
# $\mathcal{O}(\alpha \alpha_s)$ with other factorized approaches

- since the bulk of the  $\mathcal{O}(\alpha_s \alpha)$  corrections come from initial-final factorized contributions, it is interesting to compare the PA prediction for  $\mathcal{O}(\alpha \alpha_s)$  corrections with the factorized approximation NLO QCD  $\otimes$  FSR QED
- FSR QED treated with collinear structure functions or with PHOTOS



Dittmaier, Huss, Schwinn, NPB 904 (2016) 216

# from Monte Carlo



C.M. Carloni Calame et al., arXiv:1612.02841

### comparison POWHEG-BOX-V2 vs NNLO in pole approx

C.M. Carloni Calame et al., Phys.Rev. D96 (2017) 093005

$$d\sigma_{\text{POWHEG}} = d\sigma_0 \left[ 1 + \delta_{\alpha_s} + \delta_\alpha + \sum_{m=1,n=1}^{\infty} \delta'_{\alpha_s^m \alpha^n} + \sum_{m=2}^{\infty} \delta'_{\alpha_s^m} + \sum_{n=2}^{\infty} \delta'_{\alpha^n} \right],$$

 $\Delta M_W^{\alpha_s \alpha}(\mu^+ \nu_\mu) = -16.0 \pm 3.0 \text{ MeV} \qquad \text{vs} \qquad \delta_{\text{NNLO}} = -14 \text{ MeV}$ 

Dittmaier, Huss, Schwinn, NPB 885 (2014) 318, NPB 904 (2016) 216

 $\bullet$  summary of residual effects present in (QCD  $\oplus$  EW)\_{\rm NLOPS} but missing in QCD\_{\rm NLOPS} \otimes QEDPS

		$\Delta M_W ({ m MeV})$		
	QED FSR model	$M_T$	$p_T^\ell$	
Tevatron	Pythia Photos	$^{+5~\pm~2}_{-2~\pm~1}$	$^{+17~\pm~5}_{-8~\pm~5}$	
LHC	Pythia Photos	$^{+6.2\pm0.8}_{-0.6\pm0.8}$	$^{+29}_{-2} \pm 4$	

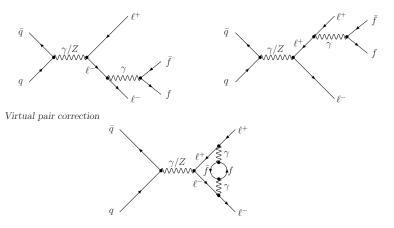
 differences in shifts induced by PYTHIA QEDPS and PHOTOS disappear when used on top of QCD⊕EW NLO

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# Lepton pair corrections: virtual and real contributions

- emission of a photon converting to a lepton pair  $\sim \mathcal{O}(\alpha^2 L^2) \sim$  two-photon contribution

Real pair emission



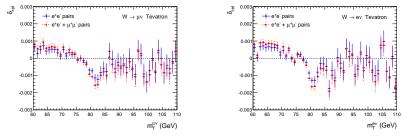
S. Antropov, A. Arbuzov, R. Sadykov, Z. Was, arXiv:1706.0557

### Lepton pair corrections: implementation in HORACE v3.1

C.M. Carloni Calame et al., arXiv:1612.02841

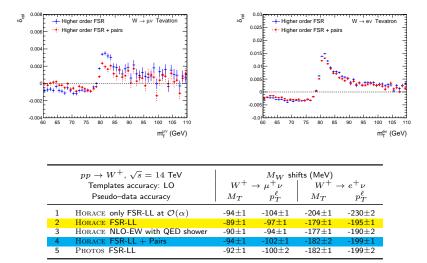
$$\alpha \Longrightarrow \alpha(s) = \begin{cases} \alpha/\left(1 - \frac{\alpha}{3\pi}\ln\frac{s}{m_e^2}\right) & \text{electrons only} \\ \alpha/\left(1 - \frac{\alpha}{3\pi}\ln\frac{s}{m_e^2} - \theta(s - m_\mu^2)\frac{\alpha}{3\pi}\ln\frac{s}{m_\mu^2}\right) & \text{electrons + muons} \end{cases}$$

 $\bullet$  running of  $\alpha$  included in the Sudakov form factor



• Normalization: multiphoton radiation

### Normalization: one-photon radiation from HORACE



 $\Delta M_W(\mu^+
u) \sim 5 \pm 1$  MeV (from  $M_\perp$ ) and  $\sim 3 \pm 2$  MeV (from  $p_\perp^\ell$ )

# NNLO uncertainty: input parameter scheme

- pert. EW calculations require a coherent set of input param. in the gauge sector, e.g.
  - $\alpha(0)$ ,  $M_W$  and  $M_Z$
  - $G_{\mu}$ ,  $M_W$  and  $M_Z$  to be preferred in the CC DY
  - we can define

$$\begin{aligned} \alpha_{\mu}^{tree} &\equiv \frac{\sqrt{2}}{\pi} G_{\mu} M_{W}^{2} \sin^{2} \vartheta \\ \alpha_{\mu}^{1l} &\equiv \frac{\sqrt{2}}{\pi} G_{\mu} M_{W}^{2} \sin^{2} \vartheta \left(1 - \Delta r\right) \end{aligned}$$

 $\blacktriangleright$  three possible different expression for the cross section, starting to differ at  $\mathcal{O}(\alpha^2)$ 

$$\begin{aligned} \alpha_0 &: \qquad \sigma = \alpha_0^2 \sigma_0 + \alpha_0^3 (\sigma_{SV} + \sigma_H) \,, \\ G_\mu I &: \qquad \sigma = (\alpha_\mu^{tree})^2 \sigma_0 + (\alpha_\mu^{tree})^2 \alpha_0 (\sigma_{SV} + \sigma_H) - 2\Delta r (\alpha_\mu^{tree})^2 \sigma_0 \,, \\ G_\mu II &: \qquad \sigma = (\alpha_\mu^{1l})^2 \sigma_0 + (\alpha_\mu^{1l})^2 \alpha_0 (\sigma_{SV} + \sigma_H) \end{aligned}$$

• potentially effects on  $M_W$  because of the different sharing among different photon multiplicities

	$p \bar{p}  ightarrow W^+$ , $\sqrt{s} = 1.96  { m TeV}$ Templates accuracy: LO			$M_W$ shifts (MeV) $W^+ \rightarrow \mu^+ \nu$	
	Pseudodata accuracy	Input scheme	$M_T$	$p_T^\ell$	
1	HORACE NLO-EW	$\alpha_0$	-101±1	-117±2	
2		$G_{\mu} - I$	-112±1	$-130 \pm 1$	
3		$G_{\mu}^{\mu} - II$	$-101 \pm 1$	$-117 \pm 1$	
4	HORACE NLO-EW+QED-PS	$\alpha_0$	-70±1	-81±1	
5		$G_{\mu} - I$	-72±2	-83±1	
6		$\begin{array}{c} G_{\mu} - I \\ G_{\mu} - II \end{array}$	-72±1	-82±2	

 differences present at NLO, after matching with higher orders, become much smaller

 $\Delta M_W \sim 2 \text{ MeV} \pm 1 - 2 \text{ MeV}$ 

# Summary

- aiming at a precision  $\delta M_W \leq 10$  MeV, the details of simulating radiation in MC's become relevant
- QCD: **impressive recent progress** in resummation matched to full fixed order results calculation
  - benchmarking activity started within LHCEEWG at Cern
- $\bullet\,$  mixed QCD $\times\,$  EW: comparison with fixed order in pole approximation nicely compatible, at the MeV scale
- the pragmatic recipe QCD NLOPS QEDLL (with PHOTOS) agrees at the MeV level with the factorized prescription QCD NLOPS EWNLOPS
  - ▶ the above prescription inherits an uncertainty of  $\sim$  5 MeV if QED FSR is simulated with PYTHIA  $(M_{\perp})$  and of  $\sim$  29 MeV  $(p_{\perp}^{\ell})$
- the differences between PYTHIA and PHOTOS disappear if used on top of EW NLO precision
- $\bullet\,$  leptonic pair corrections at the level of  $5~{\rm MeV}$
- $O(\alpha^2)$  uncertainties by exploring different input param schemes at the level of 1-2 MeV (with the available statistics)