

Long-term monitoring of delivered luminosity & calibration stability in LHCb

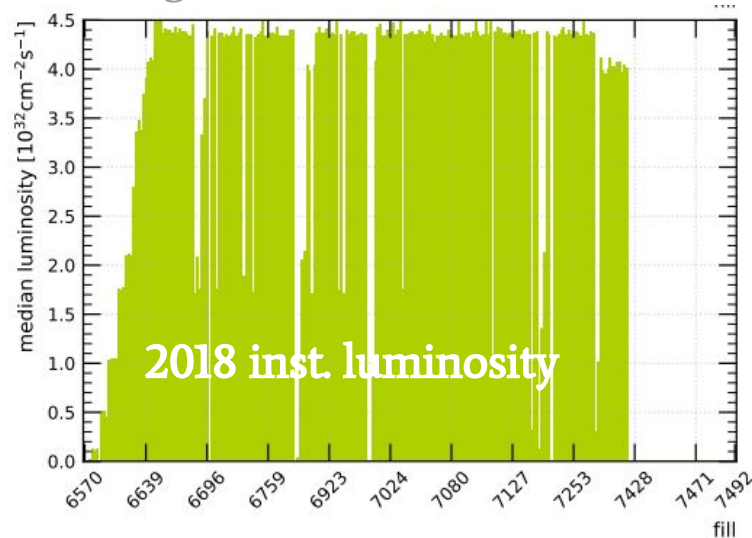
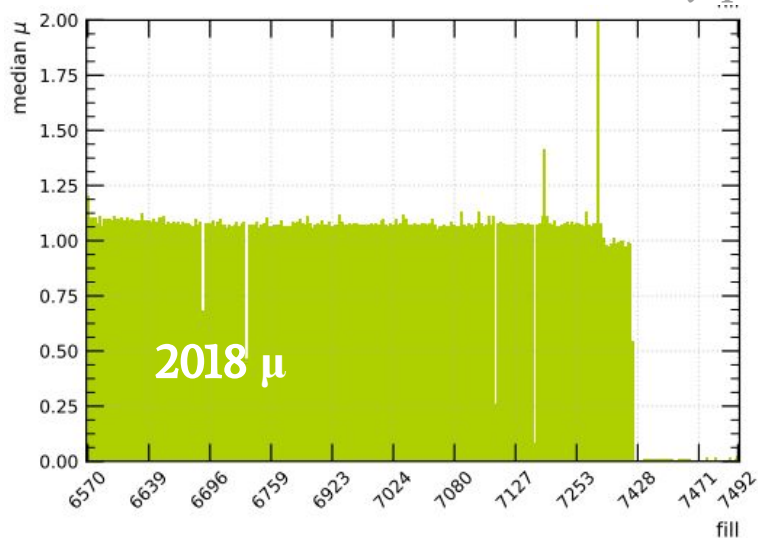


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LHC Lumi Days 2019, 4–5 June 2019




What's special in LHCb?

median μ and instantaneous luminosity per fill using online CALO-based measurement



Leveling at μ of 1.1 is a blessing for propagation of luminosity calibration!

Simply count the events

$$\int L dt = \frac{1}{\sigma_{eff}} \sum_i N_{int}^i$$


Absolute calibration

determined from
dedicated measurement
using BGI and van der
Meer methods

Relative luminosity

(interaction counting)
determined using randomly
triggered crossings and “-log0”
method

Relative luminosity

$$N_{\text{int}} = \frac{f_{\text{rev}}}{R_{\text{trig}}} \sum_j N_j^{\text{trig}} \mu_j \approx \frac{f_{\text{rev}}}{R_{\text{trig}}} n_{\text{bb}} N^{\text{trig}} \mu$$

$$\mu = \frac{\mu_{\text{bb}} - \mu_{\text{be}} - \mu_{\text{eb}}}{1 - I(0)}$$

$$\mu_{\text{bb}} = -\log P(0) = -\log \frac{N_0^{\text{trig}}}{N^{\text{trig}}}$$

“log0” method works really well at mu of 1.1 to get a value proportional to L

In Run 2, typically 1000 Hz of random “bb” events acquired, “nanofied” to O(100) bytes and stored in physics data. Aggregated in prompt offline processing and calibration applied in analysis jobs

f_{rev} = LHC revolution frequency

R_{trig} = random trigger rate (bb crossings)

j runs over all colliding BCIDs

Relative luminosity (2)

Method equivalent to Run 1
2014 JINST 9 P12005

$$N_{\text{int}} = \frac{f_{\text{rev}}}{R_{\text{trig}}} \sum_j N_j^{\text{trig}} \mu_j \approx \frac{f_{\text{rev}}}{R_{\text{trig}}} n_{bb} N^{\text{trig}} \mu$$

Bias is introduced due to spread of μ over bunches
- **correction is needed**

$$\mu = \frac{\mu_{bb} - \mu_{be} - \mu_{eb}}{1 - I(0)}$$

I is the detector response to one interaction.

$$\mu_{bb} = -\log P(0) = -\log \frac{N_0^{\text{trig}}}{N^{\text{trig}}}$$

$I(0)$ is the probability to get an “empty” event if one interaction took place. Depends on Z position of luminous region (LR) for VELO counters.
- **correction is needed**

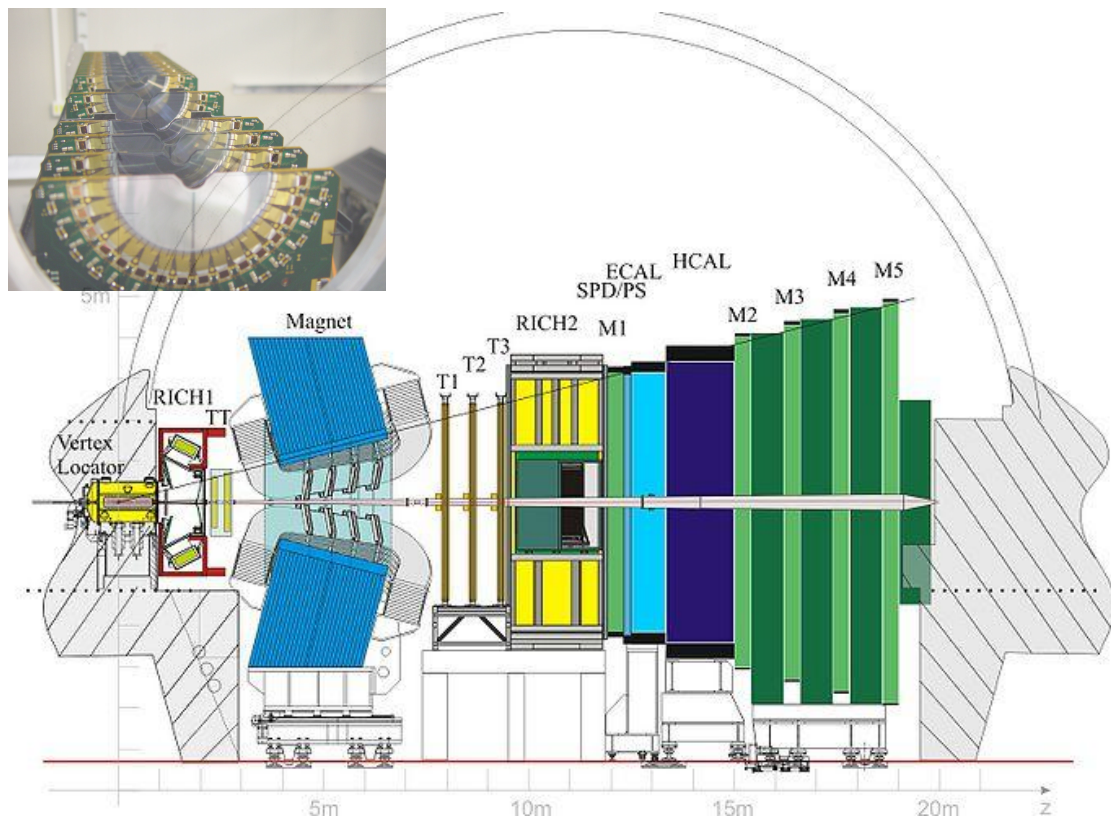
Number of “empty” events.

We can use various definitions of “empty” event (various lumi counters/observables).

- number of VELO tracks < 2
- number of PV < 1
- ...

f_{rev} = LHC revolution frequency
 R_{trig} = random trigger rate (bb crossings)
 j runs over all colliding BCIDs

Counting events



CALO (L0)

- SPD: # hits in 6k scintillators
- HCAL: max ET
- online: & of the two

MUON (L0)

- # muon candidates

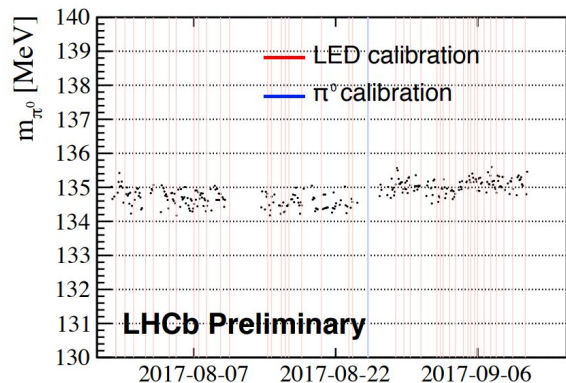
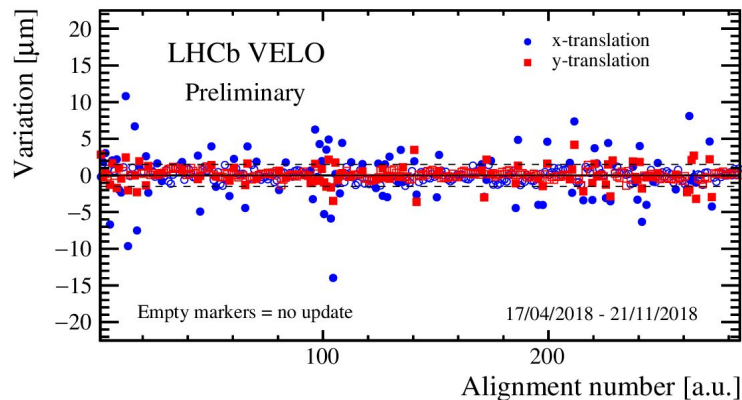
PU (L0)

- # hits

VELO (HLT)

- # tracks (baseline counter)
- # vertices
- variations of the above

Well calibrated!



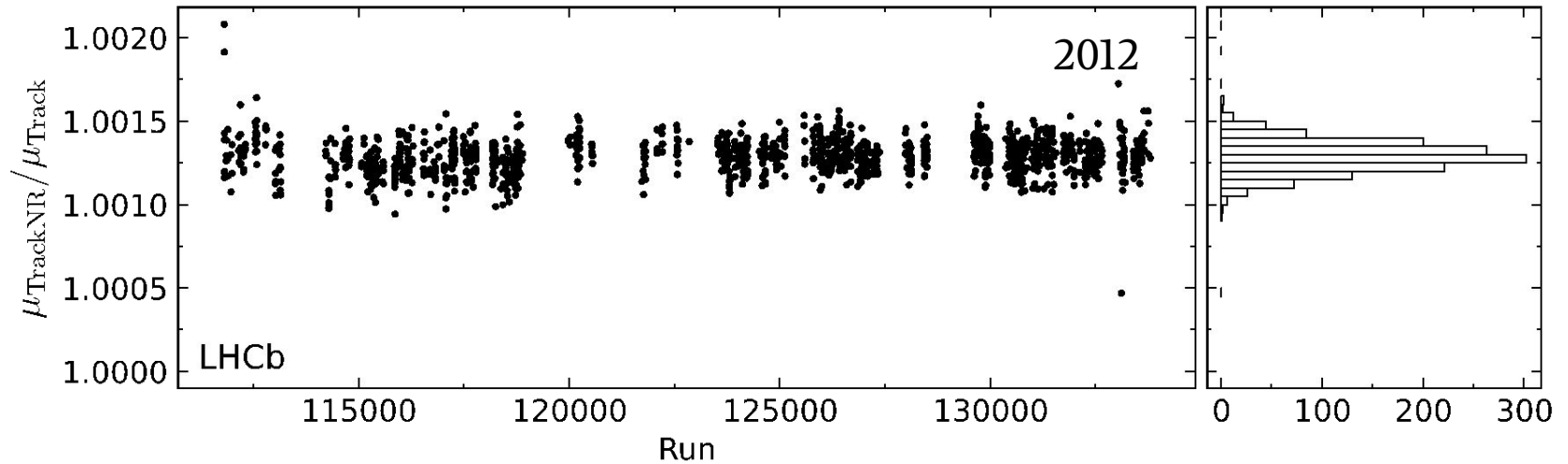
I like LHCb a lot, too!

Luminosity wouldn't be so easy without all calibrations.

VELO: Regular IV, CCE scans, HV adjustments; align every fill!

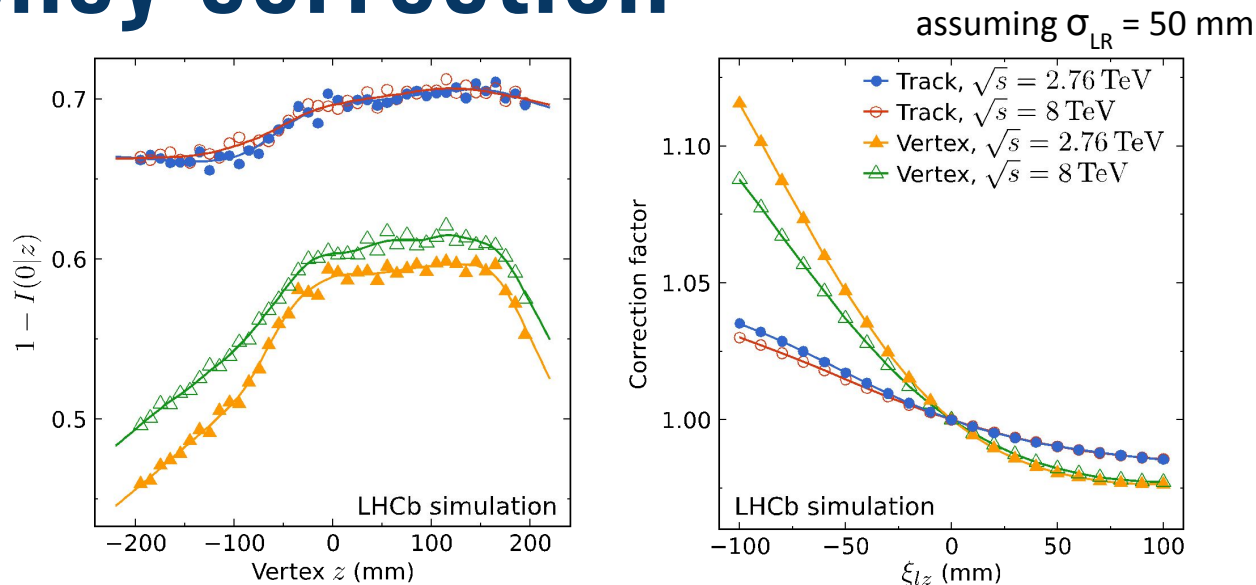
CALO: timing; caesium scan every TS for HCAL, π^0 -based for ECAL; update HV every fill based on a LED system

Background



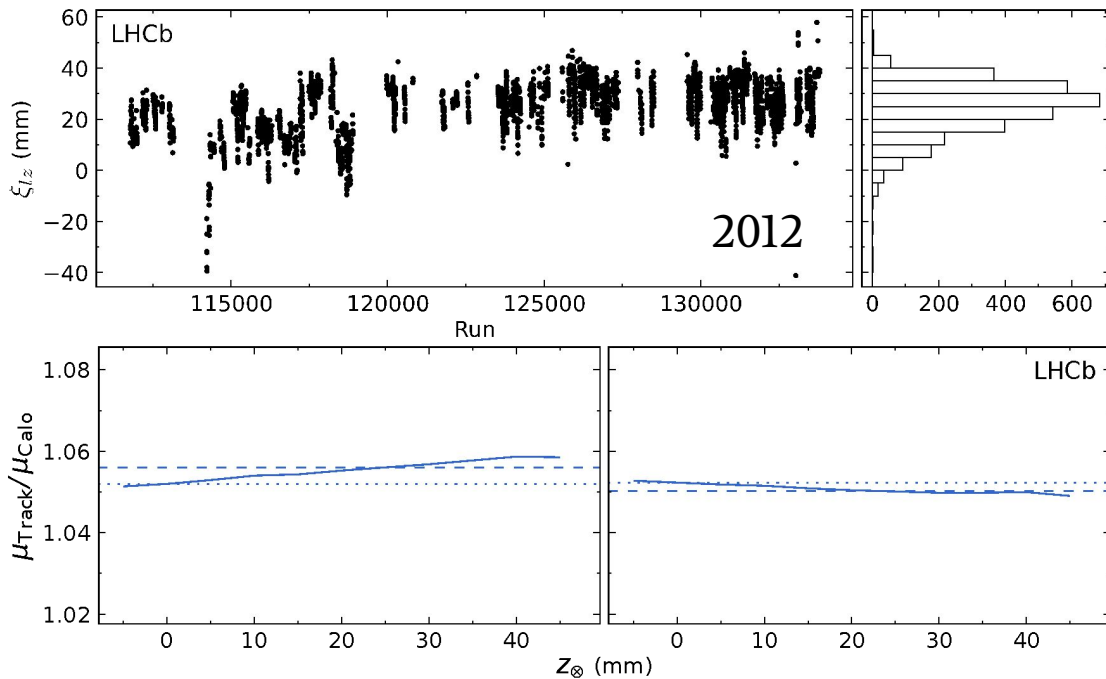
- Beam-gas background is well understood and subtracted (small)
- Beam-beam related background is estimated (e.g. main-satellite collisions)
 - take the difference between counting VELO tracks with and w/o fiducial volume cut around the luminous region as a systematic
 - the restricted observable is used for calibration

Efficiency correction



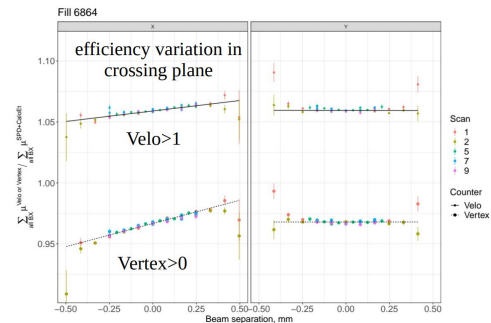
- Longitudinal luminous region (LR) movements not negligible compared to VELO length (worse in 2012, leveling in bad plane)
- The z -dependence of the efficiency is estimated from simulation
- Correction based on measured z position and length of LR

Efficiency correction (2)



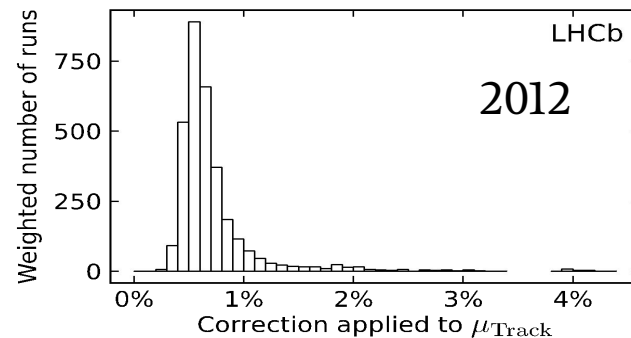
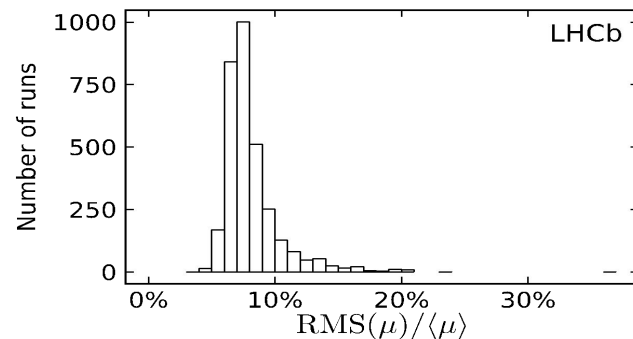
Cross check with an unaffected observable: online luminometer based on HCAL+SPD

Same procedure in VDM (bigger lever arm)



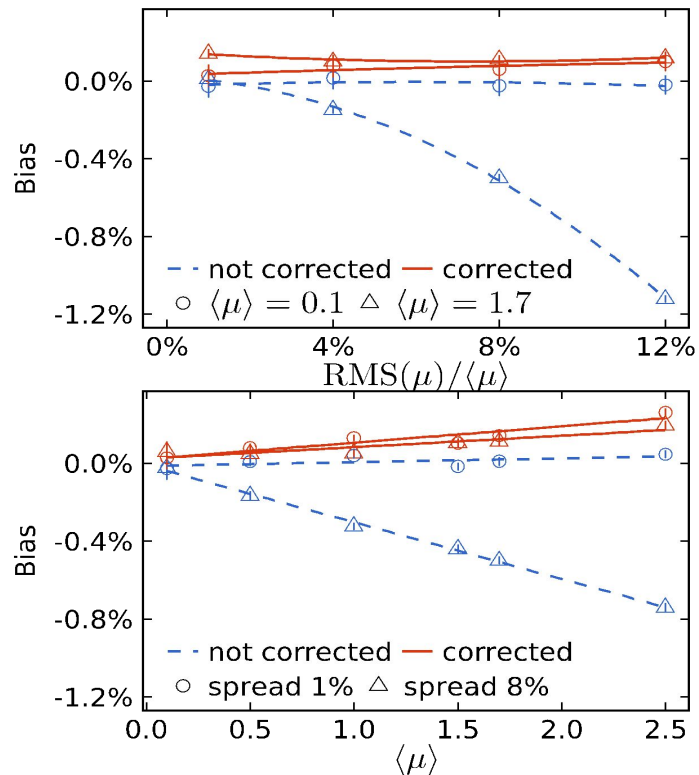
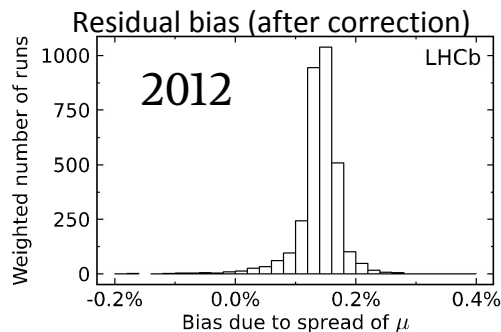
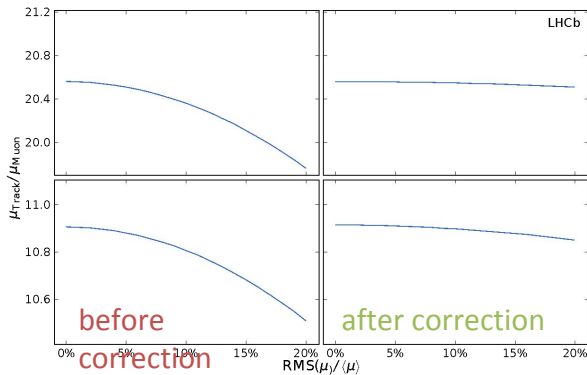
Spread of μ over bunches

- Random triggers rate insufficient to measure μ for each bunch in a short period of time
- First, assume all bunches have the same μ
- Then, correct using long time periods (30m)
 - assume μ does not change during the period
 - measure relative μ values of bunches
 - calculate correction factor for each long period
 - estimate residual bias after such correction using a MC technique
- Run 3: more spread expected \Rightarrow to be revised
 - with distinct groups of bunches



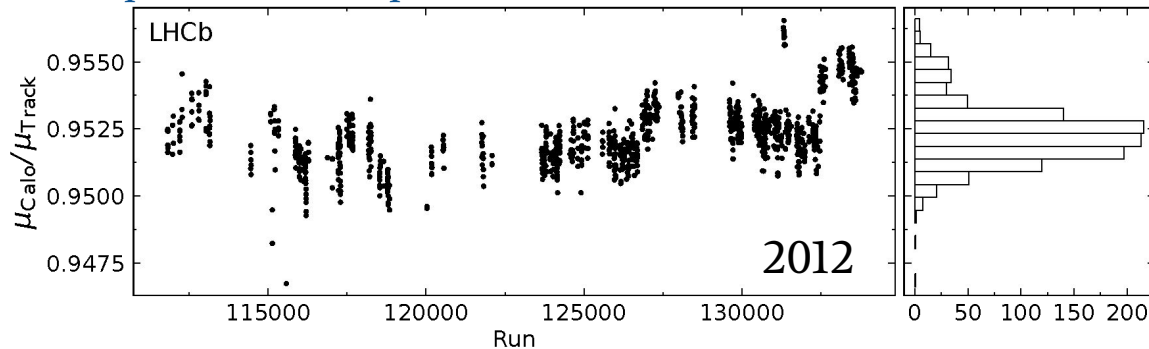
Spread of μ over bunches (2)

- Estimate residual bias using toy simulations
- Cross-check using a less affected counter
 - i.e. one with smaller efficiency (μ values)

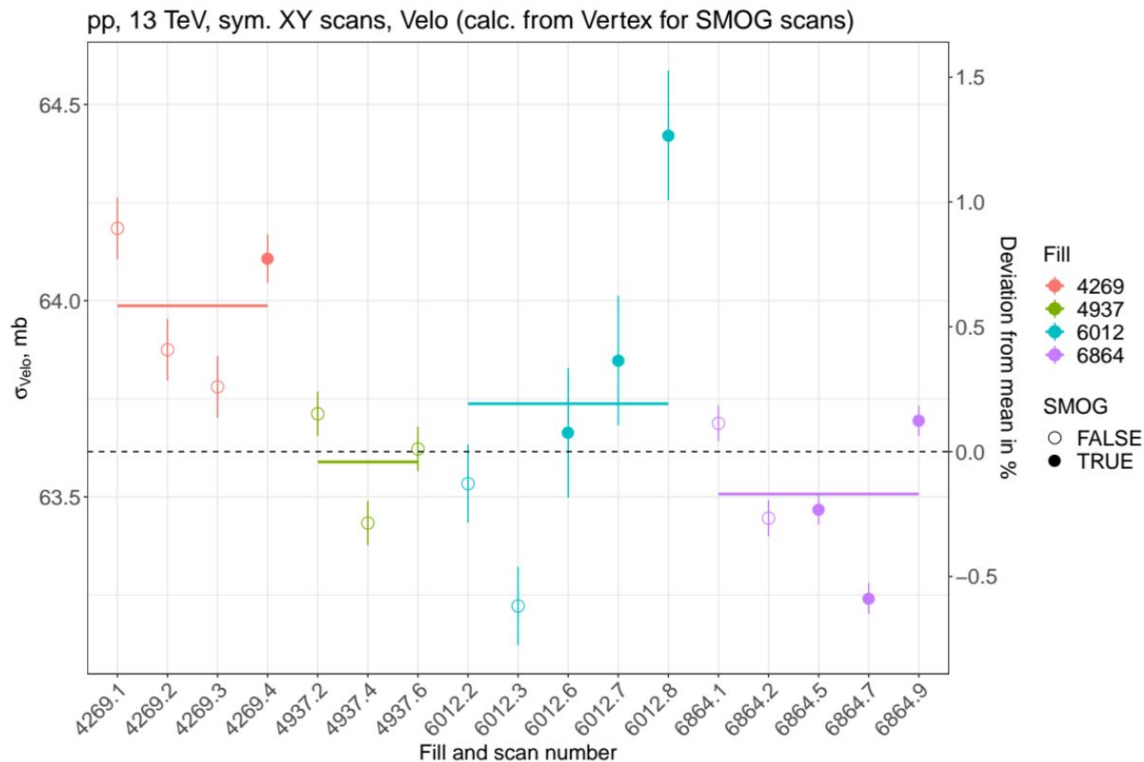


Stability of the effective process

- Compare Track counter with the CALO-based (online) counter
- Take the RMS of the ratio as an uncertainty on the Track counter
 - conservative, we “know” Tracks are inherently more stable
 - threshold changes in the calo correlate with steps in the ratio
- Analogous comparison with Vertex counter ($\#$ vertices ≥ 1)
 - Vertices composed of $>4(5)$ tracks \Rightarrow ratio sensitive to VELO efficiency
 - RMS compatible with expected statistical fluctuations



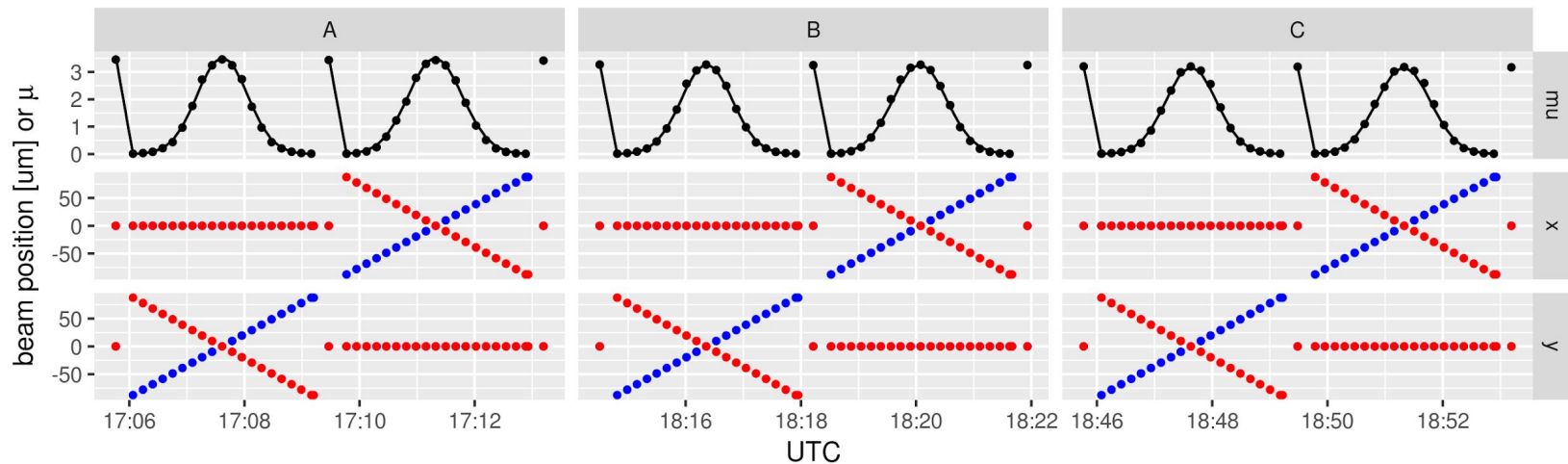
Ultimate stability check



The absolute calibration for each year of Run 2 yields stable values across different counters

Emittance scans?

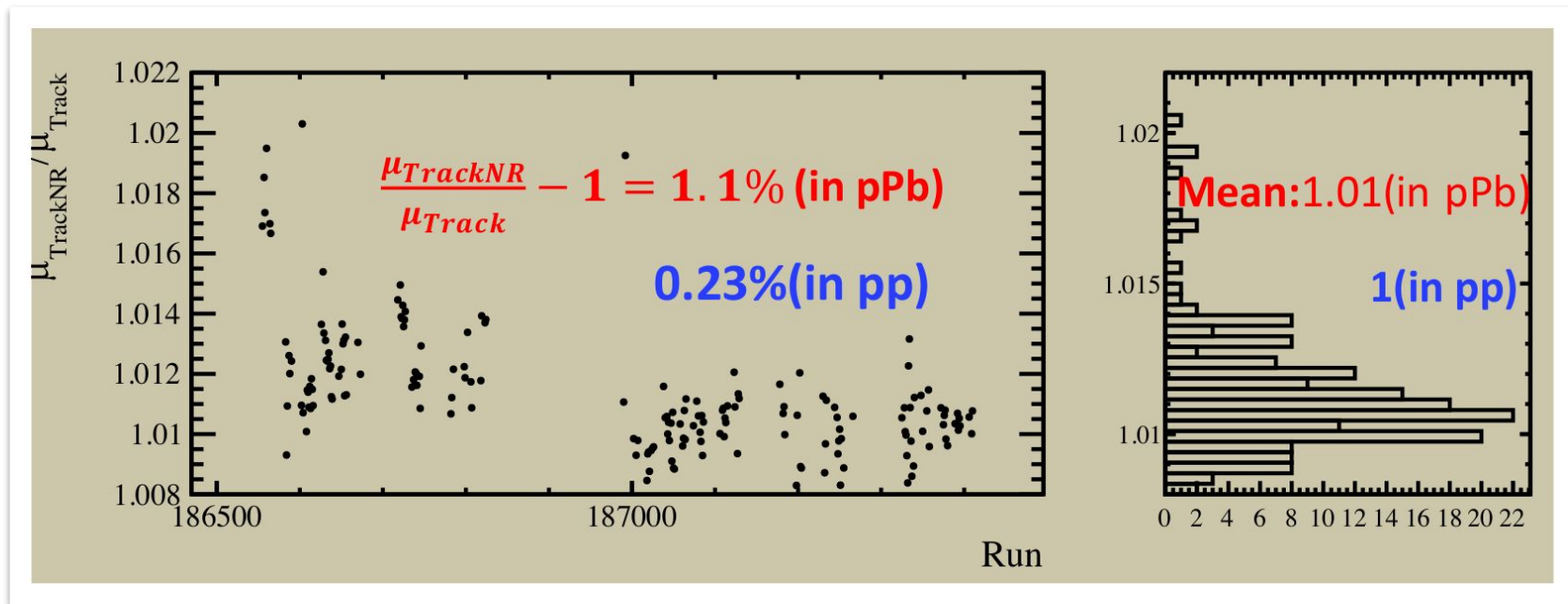
Online bunch-average analysis of VDM scans in fill 6913



- LHCb took part in a 2018 BSRT calibration fill. **Nominal optics \Rightarrow high μ !**
- Rudimentary online analysis gives cross-section surprisingly close to VDM
- Likely OK in Run 3 to exceed target μ as long as total lumi is $<$ design

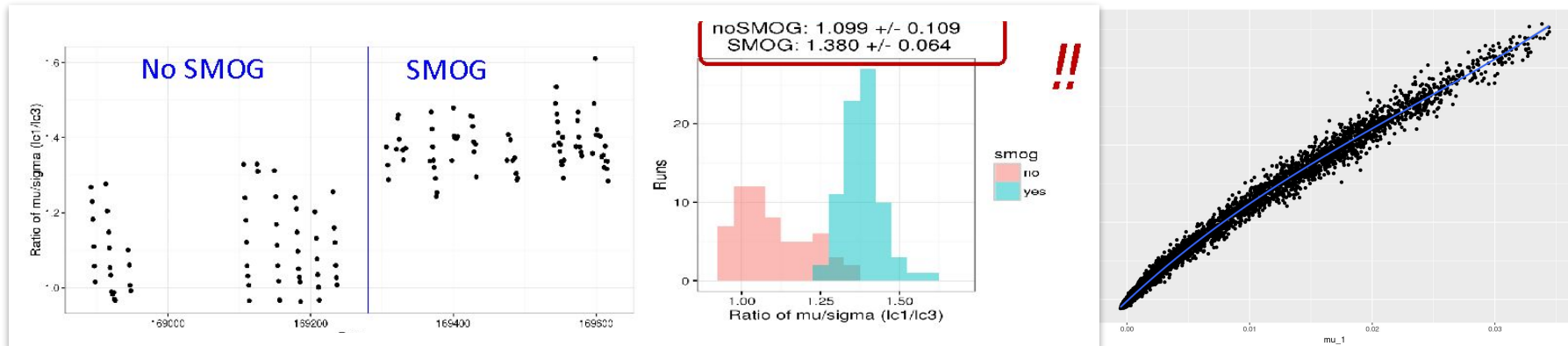
Proton-lead luminosity

- generally straightforward
- low mu \Rightarrow relative backgrounds higher \Rightarrow larger uncertainty



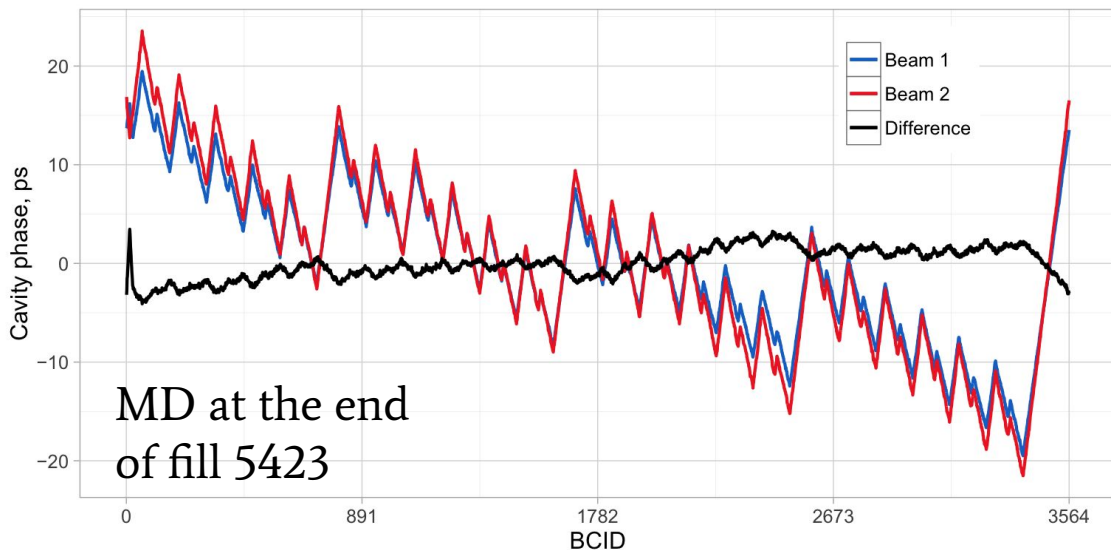
Lead-lead luminosity

- Large time-variation (non-linearity) of luminosity counters
- ~12% uncertainty on relative measurement
- seen in both 2015 and 2018
 - 2018: record high-rate random raw data to study effect
 - now known to be due to sensitivity to EM ($\mu \gg 1$)



Fun with RF detuning

Phase modulated RF voltage to minimize klystron power



$$\mu_{LR,z} = \frac{z_1 + z_2}{2} - \cos \alpha \sin \alpha \frac{x_1 - x_2}{2} \frac{\sigma_z^2 - \sigma_x^2}{\sigma_z^2 \sin^2 \alpha + \sigma_x^2 \cos^2 \alpha}$$

$$\Delta\mu_{LR,z} = \mu_{LR,z}(\Delta t_1, \Delta t_2) - \mu_{LR,z}(0, 0) = -\frac{c\Delta t_1 - c\Delta t_2}{2} \left(1 - \frac{\sigma_{LR,z}^2 \sin^2 \alpha}{\sigma_{LR,x}^2} \right)$$

not very visible?

MOPPC015

Proceedings of IPAC2012, New Orleans, Louisiana, USA

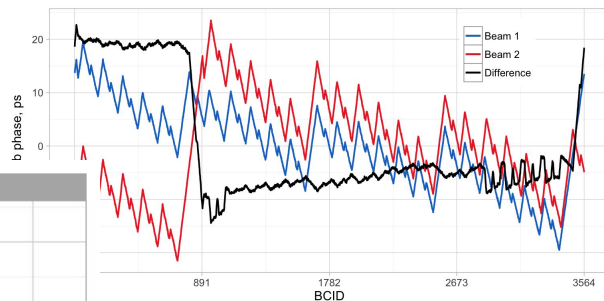
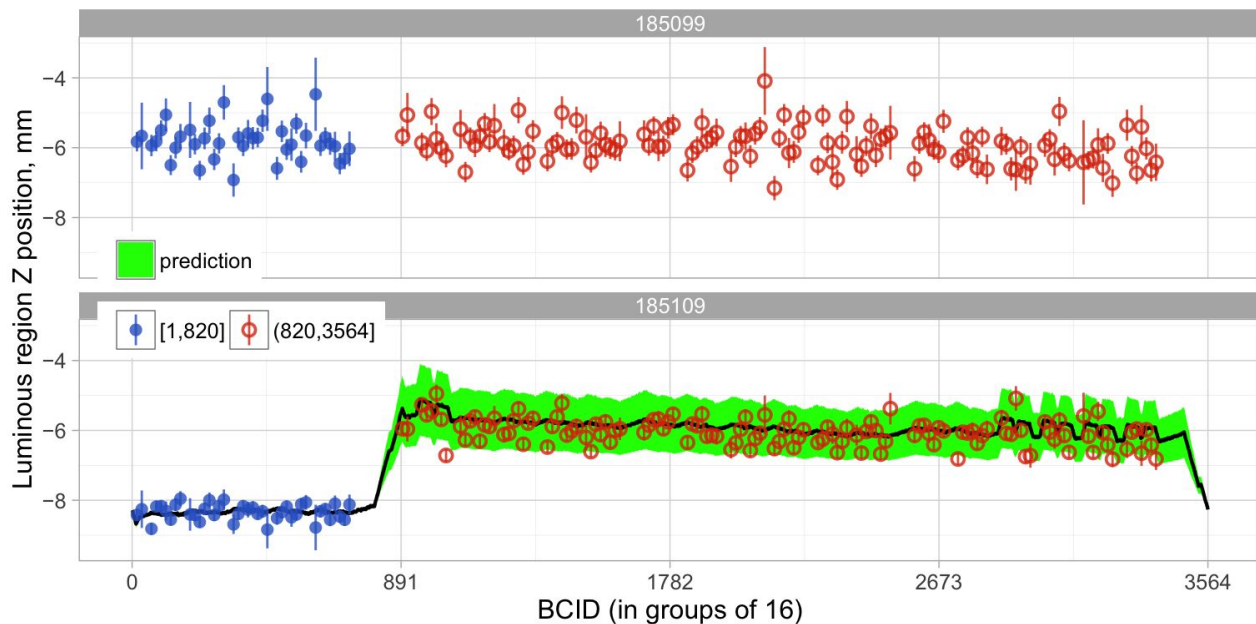
PROPOSAL FOR AN RF ROADMAP TOWARDS ULTIMATE INTENSITY
IN THE LHC

P. Baudrenghien, T. Mastoridis, CERN, Geneva, Switzerland

<https://accelconf.web.cern.ch/accelconf/IPAC2012/papers/moppc015.pdf>

Fun with RF detuning (2)

LHCb beam 2 shifted by 894 BCIDs!



Little impact on lumi measurement, but fun!

Outlook

- Work ongoing to finish Run 2 relative calibrations
 - promising Vertex/Tracks stability at the permil level \Rightarrow stable VELO efficiency
 - “out-of-the-box” stability of Tracks vs. CALO at $O(1\%)$
 - we know we can do better: $\sim 0.2\%$ in Run 1
 - tedious process of understanding every significant change in ratios
- Run 3
 - at **mu of 5.5**, log0 still works (keep statistical power by $\sim 10x$ rate increase)
 - **brand new detectors** \Rightarrow reassess choice of luminosity counters
 - larger **bunch-to-bunch mu differences** expected \Rightarrow more data per-bunch needed
 - ideas under consideration for new, **dedicated luminometer hardware**

Novelties in Run 2

- Readout supervisor firmware update
 - Per-BCID trigger masks for the various trigger sources: NoBias, Sequencer, L0
 - Could easily do physics + BGI at the same time (p-He + BGI, p-Pb + ghost charge)
 - Could take data efficiently for Pb VDMs, e.g. 5-8 of O(100) bunches triggered at 11kHz
- Continuous beam-shape monitoring
 - Important input for the LHC colleagues
 - Main limitation: resolution unfolding at 3m
 - **Run 3: improved VELO but even smaller beta**

	EE	B1	B2	BB	BL	BT
Physics	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
No bias	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
B1gas	<input type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
B2gas	<input type="checkbox"/>	<input type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
Sequencer	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input checked="" type="checkbox"/>	<input type="checkbox"/>

