



# *Gamma-ray calorimetry in Nuclear Physics experiments*



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**8<sup>th</sup> Beam Telescopes and Test Beams Workshop**  
**Tbilisi, 27<sup>th</sup> - 31<sup>st</sup> January 2020**

- **Introduction to calorimetry**
- **Gamma-rays detection**
  - **Gamma interactions**
  - **Detecting gamma-rays**
  - **Understanding a gamma-ray spectrum**
- **Gamma-ray emission in Nuclear Physics Experiments**
  - **Reactions and excited states**
  - **Some  $\gamma$ -ray calorimeters**
  - **Event reconstruction: addback**
- **Concluding remarks**

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**Calorimetry:** act of measure transferences of energy.

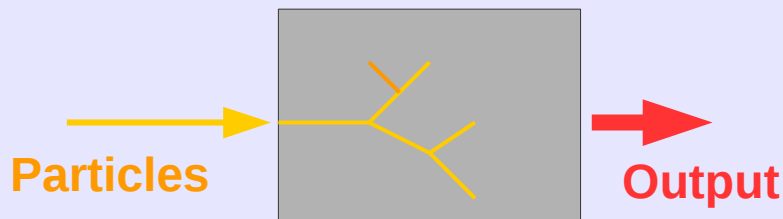
**Calorimeter:** device to measure the energy of particle through total absorption.

**Basic idea:** deposit the full energy in the medium (Shower formation).

Importance: know with precision the final states in particle collisions.

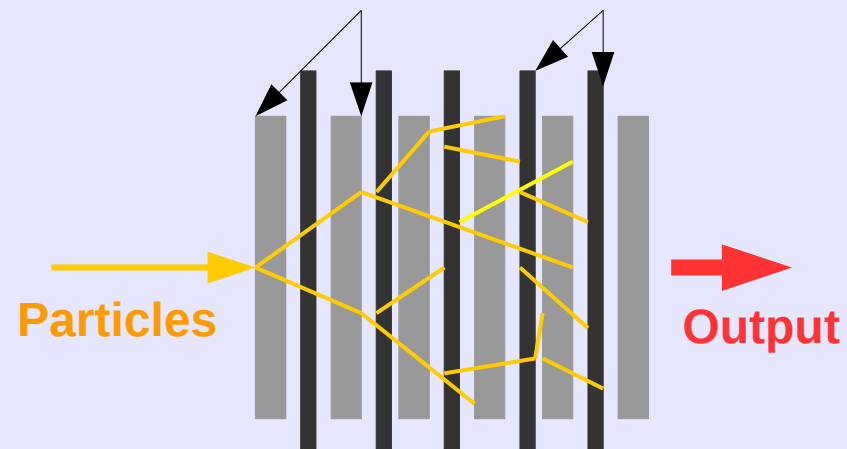
## Calorimeter configurations:

### Absorber + Detector



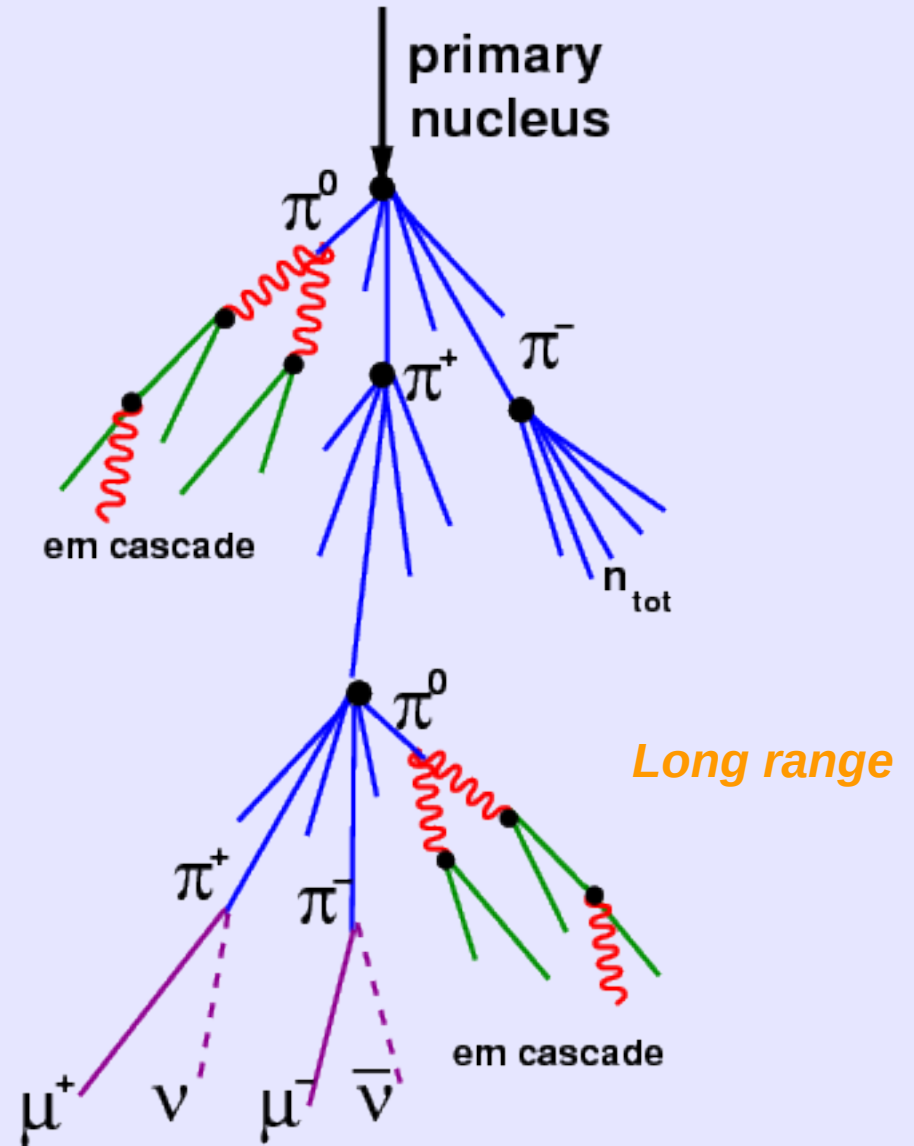
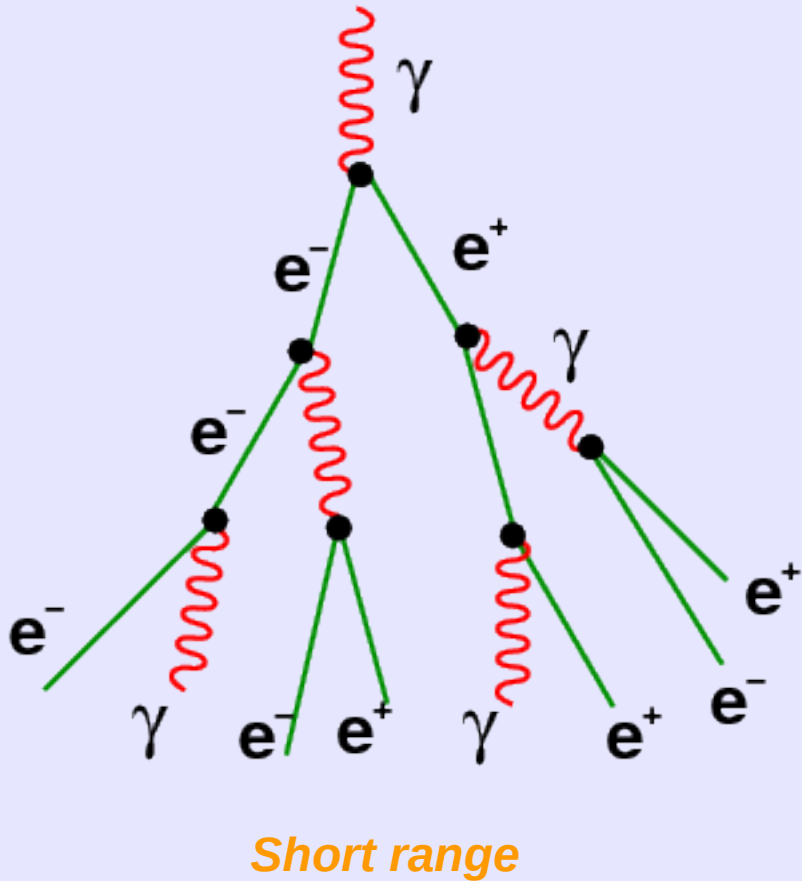
**Homogeneous calo**

### Detectors Absorbers



**Sampling/segmented calo**

## Electromagnetic showers (electrons, positrons and gammas)



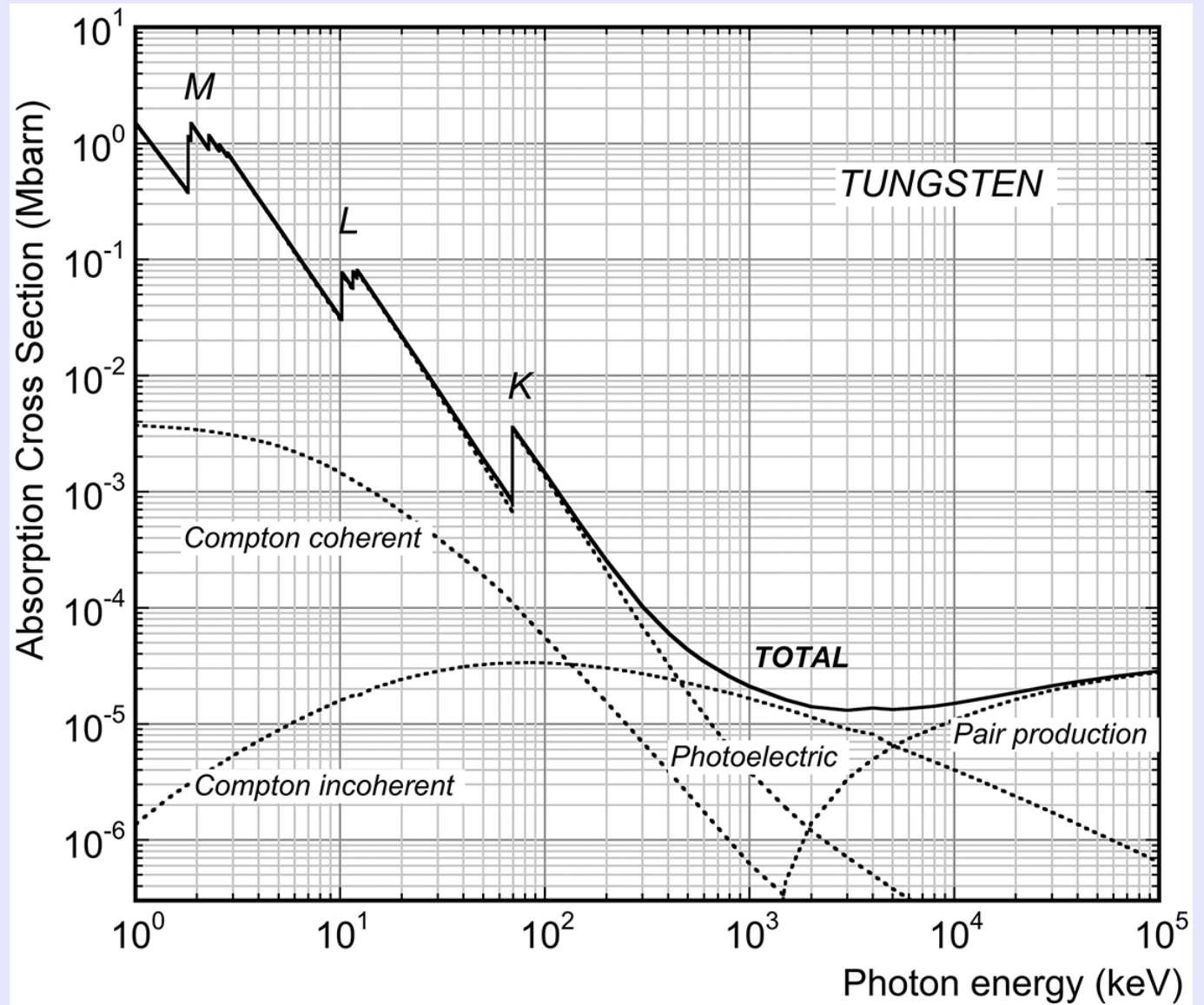
## Hadronic showers

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**$\gamma$  interactions  
cross-section  
diagram**

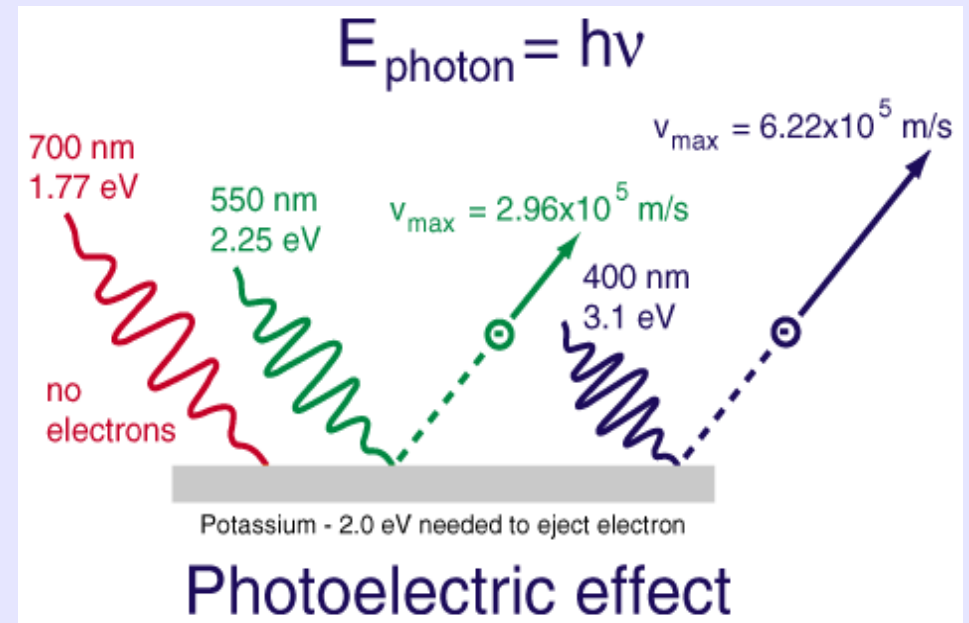
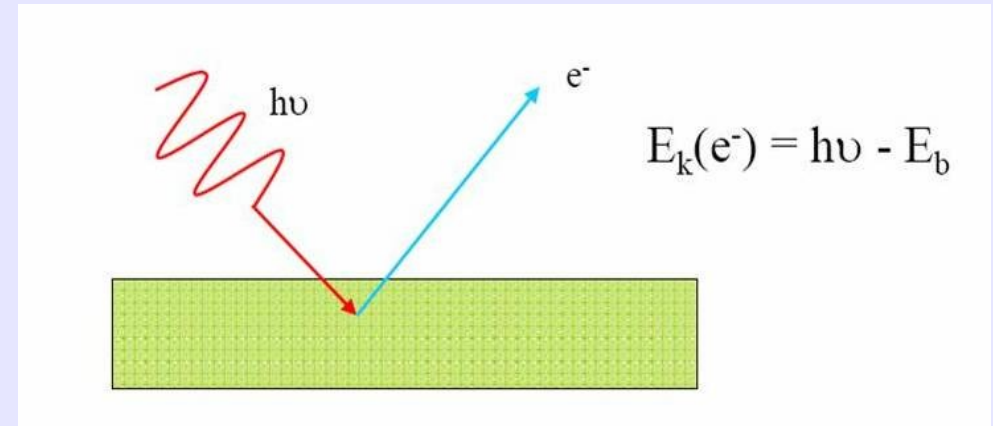


**Photons  
interact with  
matter by three  
different  
processes**



## Photoelectric Effect

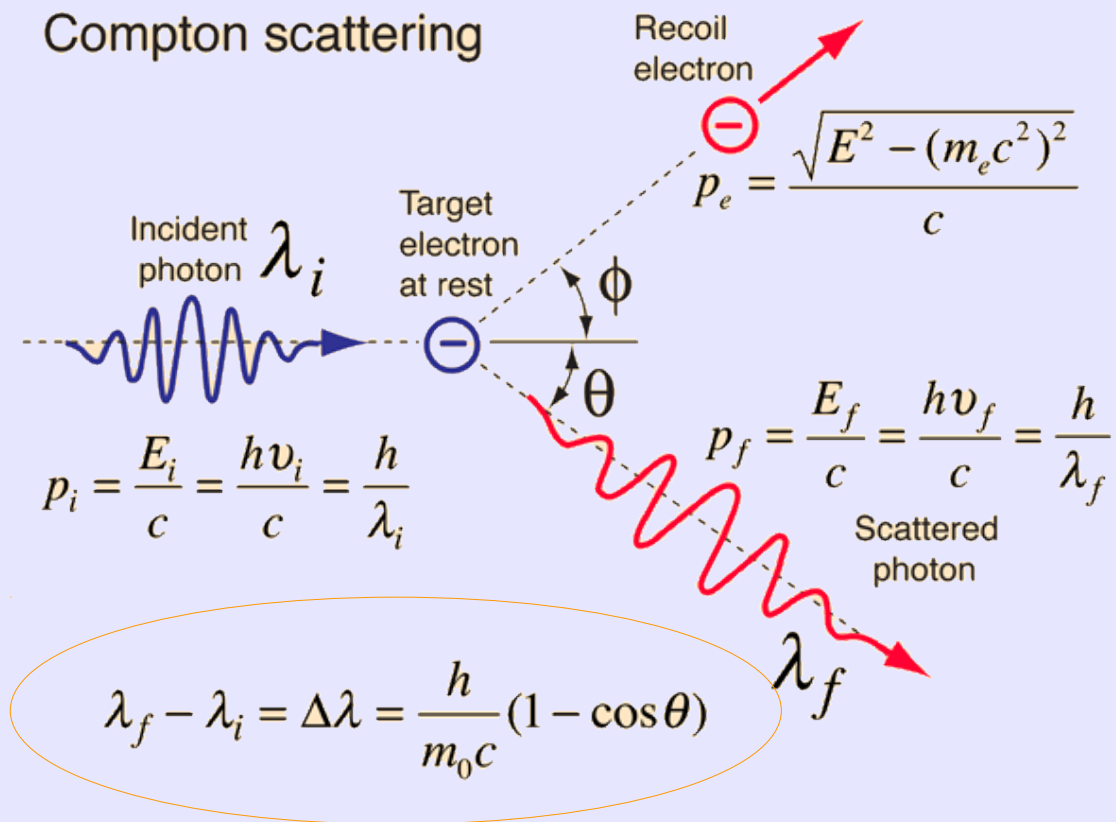
- The incident photon extracts an electron from the medium
- Dominates at low energies (<1 MeV)





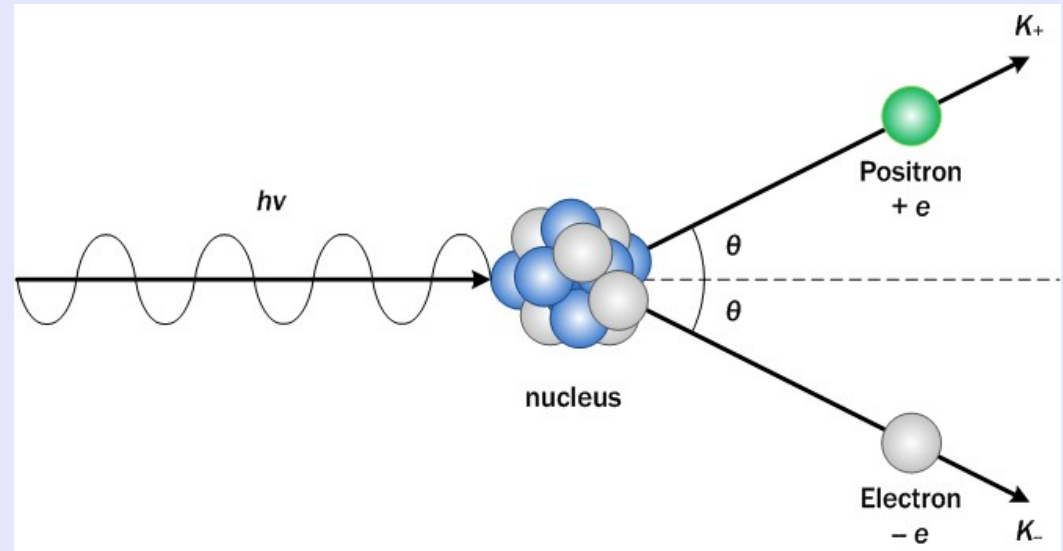
## Compton Scattering

- The incident photon “is scattered” and an electron is ejected at a given angle
- Becomes important at intermediate energies (~1 MeV)



## Pair Production

- **The incident photon interacts with the nuclear field and an electron-positron pair is created**
- **Is the main contribution at high energies (>1 MeV)**



**Energy threshold!!**

$$E = h\nu = 1.022 \text{ MeV}$$

## Two basic types of gamma-ray detectors:

### Scintillating detectors

- Scintillating light produced after ionization
- Transparent to that light
- Organic (plastic), inorganic (NaI, CsI, BGO), ceramic (GAGG)
- Read-out with light sensors (PMTs, APDs, SiPMs)

### Semiconductor detectors

- Electron-hole pairs are produced after ionization
- Very good resolution
- Silicon, Germanium, Cd-based

### KEY FACTOR:

**The output should be proportional to the energy deposited by the incident photons**

**Output pulses must be sorted by magnitude**

**Basic tools:  
Multichannel Analyzers (MCA)  
Analog-to-Digital Converters (ADC)**



**HISTOGRAMS**

***Important parameter: Energy Resolution***

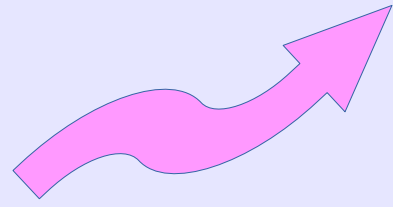
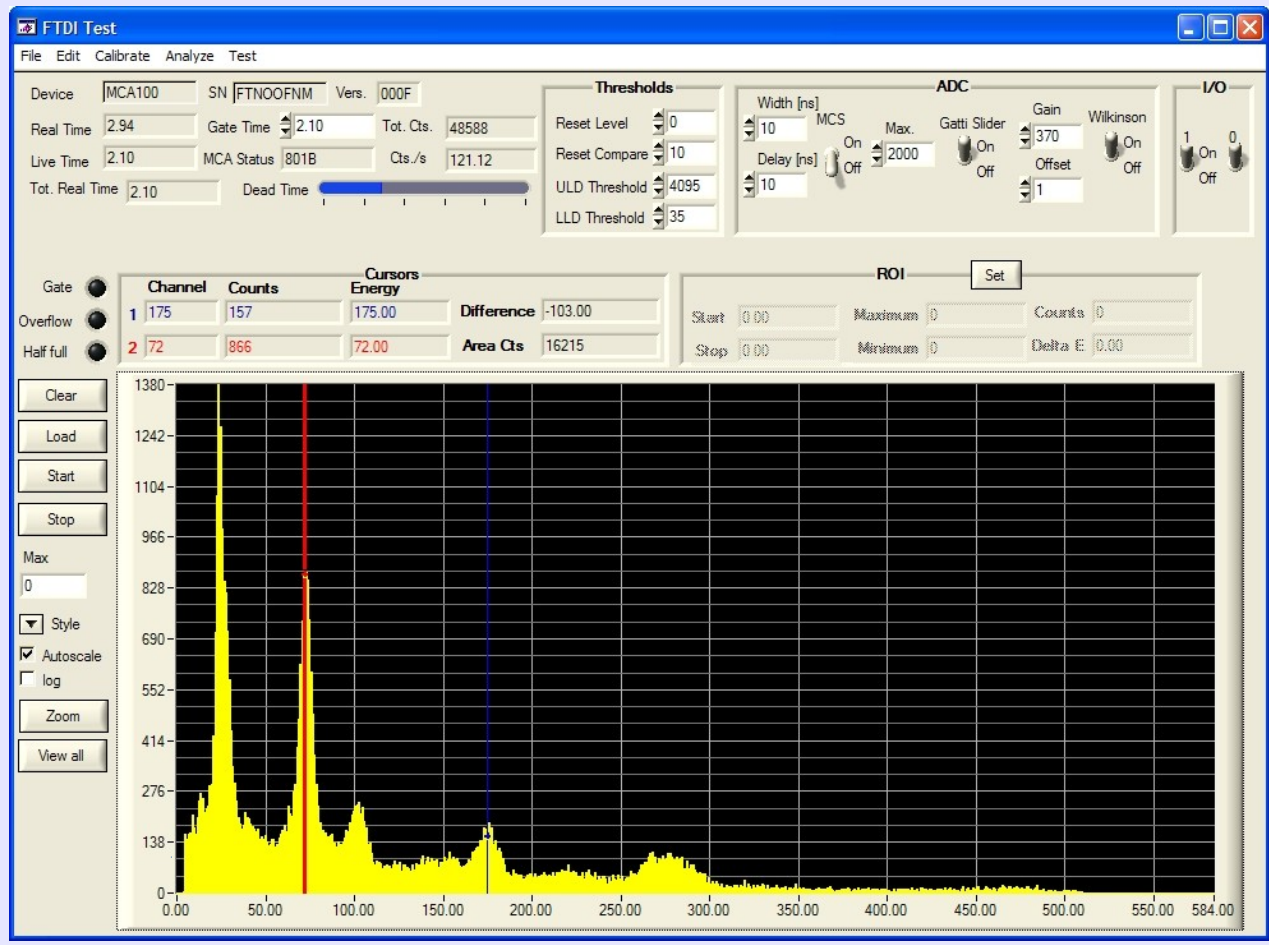
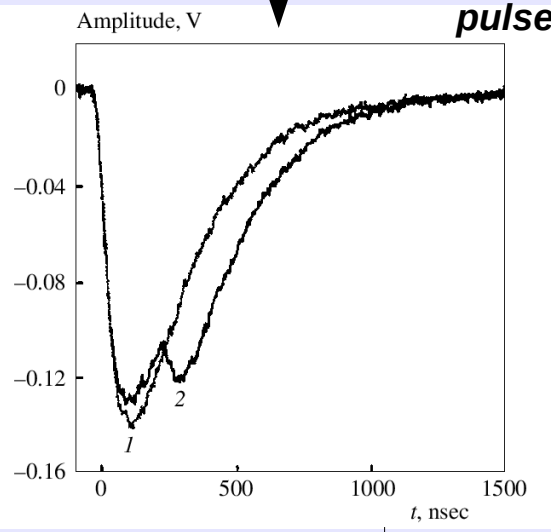
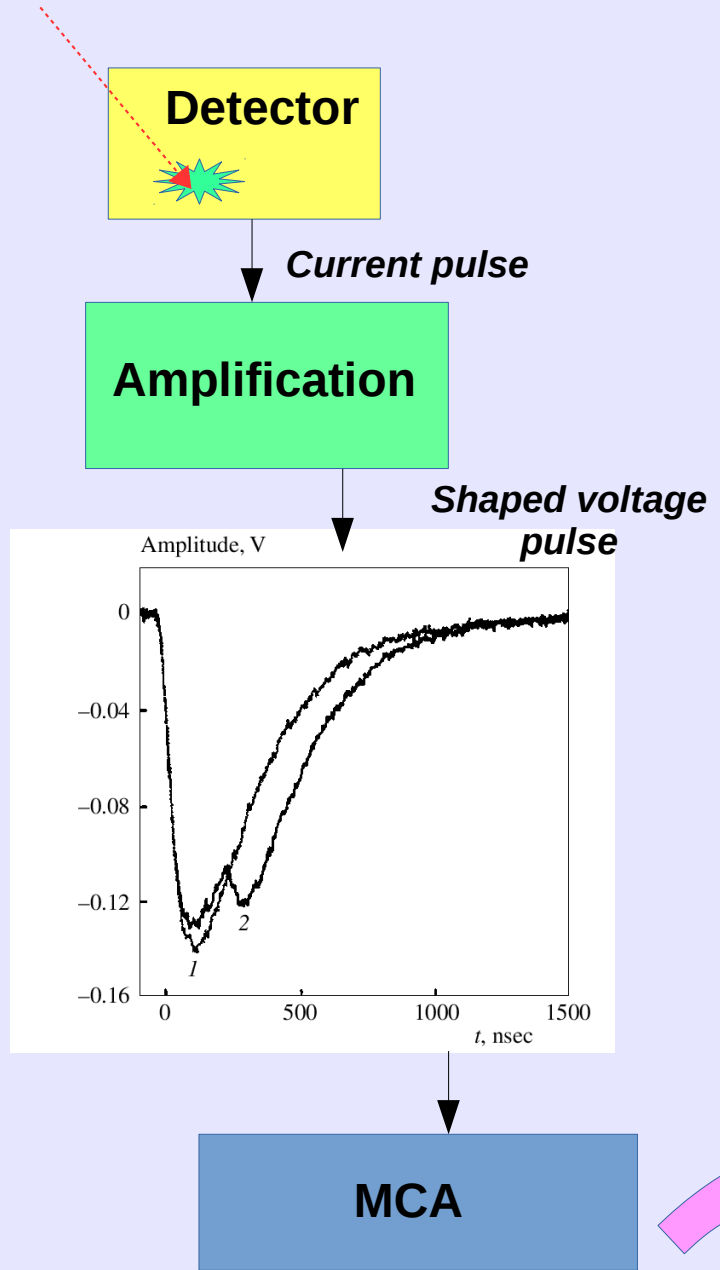
$$\sigma(E) = \text{FWHM} = 2.34\sigma$$



$$R(\%) = \sigma(E)/E = a/\sqrt{E} + b/E + c$$

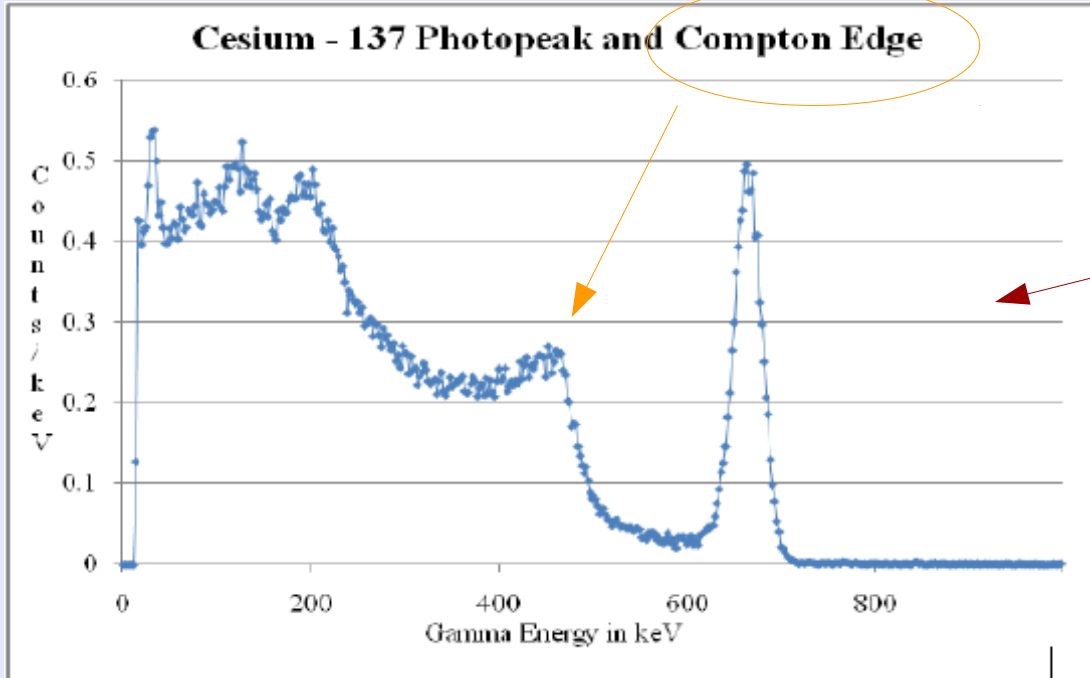
**Fitting the photopeak  
to a Gaussian curve**

**Stochastic + noise + constant  
(calibration, materials,  
instruments...)**



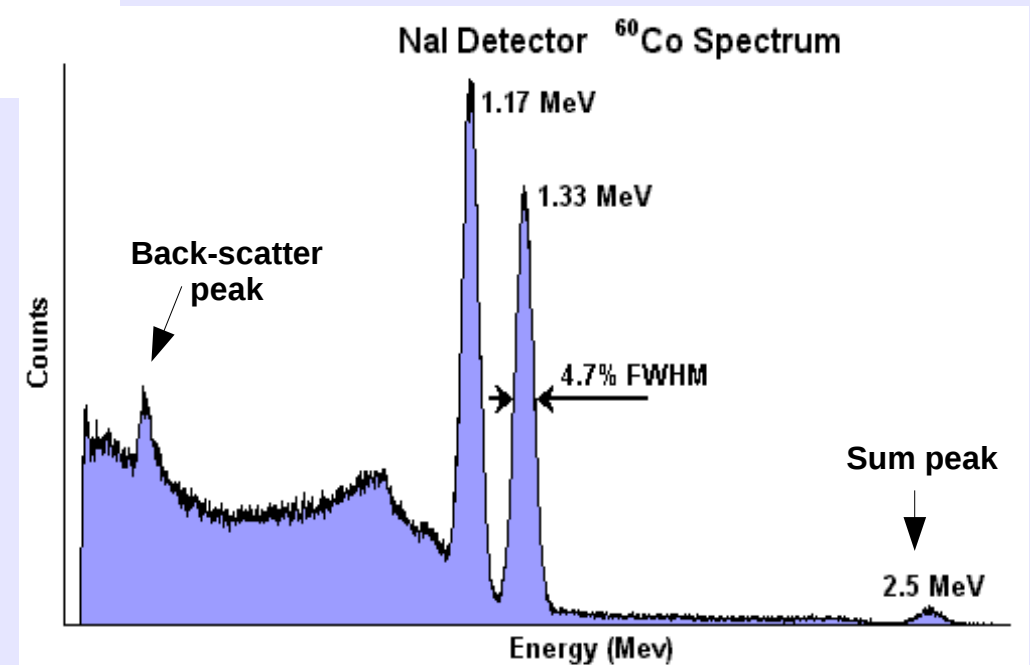
**Sorted pulses by amplitude:**  
**Histogram**  
**Calibration linear with  $E$**

# Understanding the spectrum

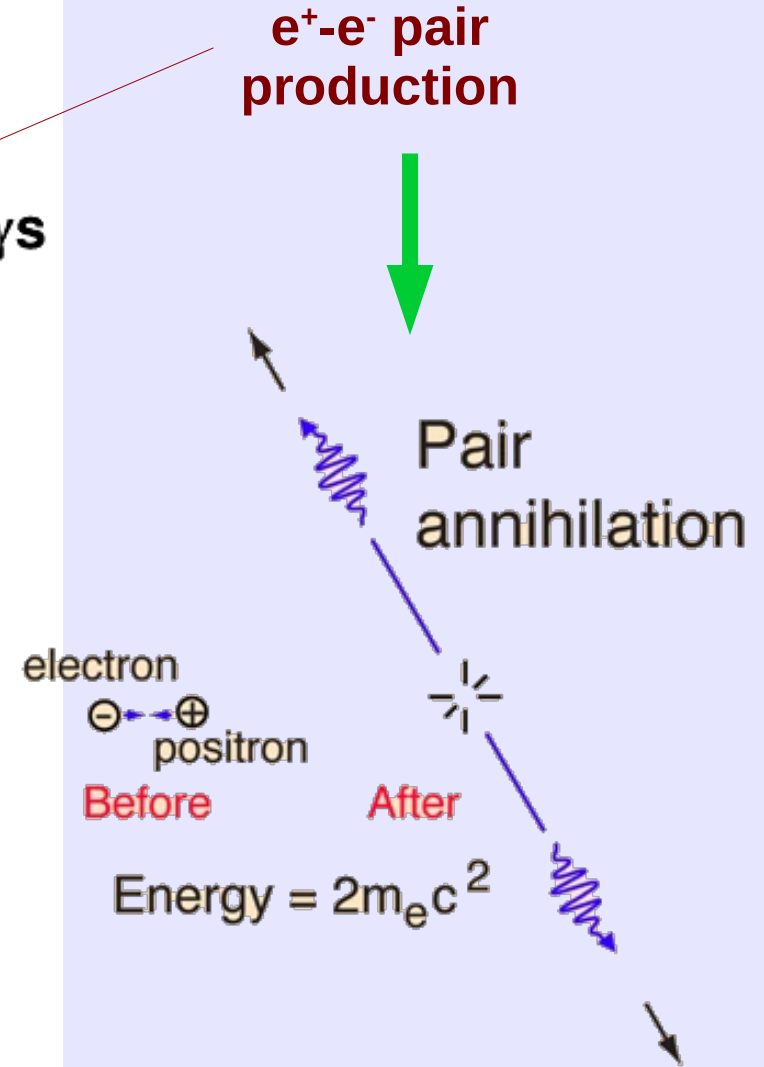
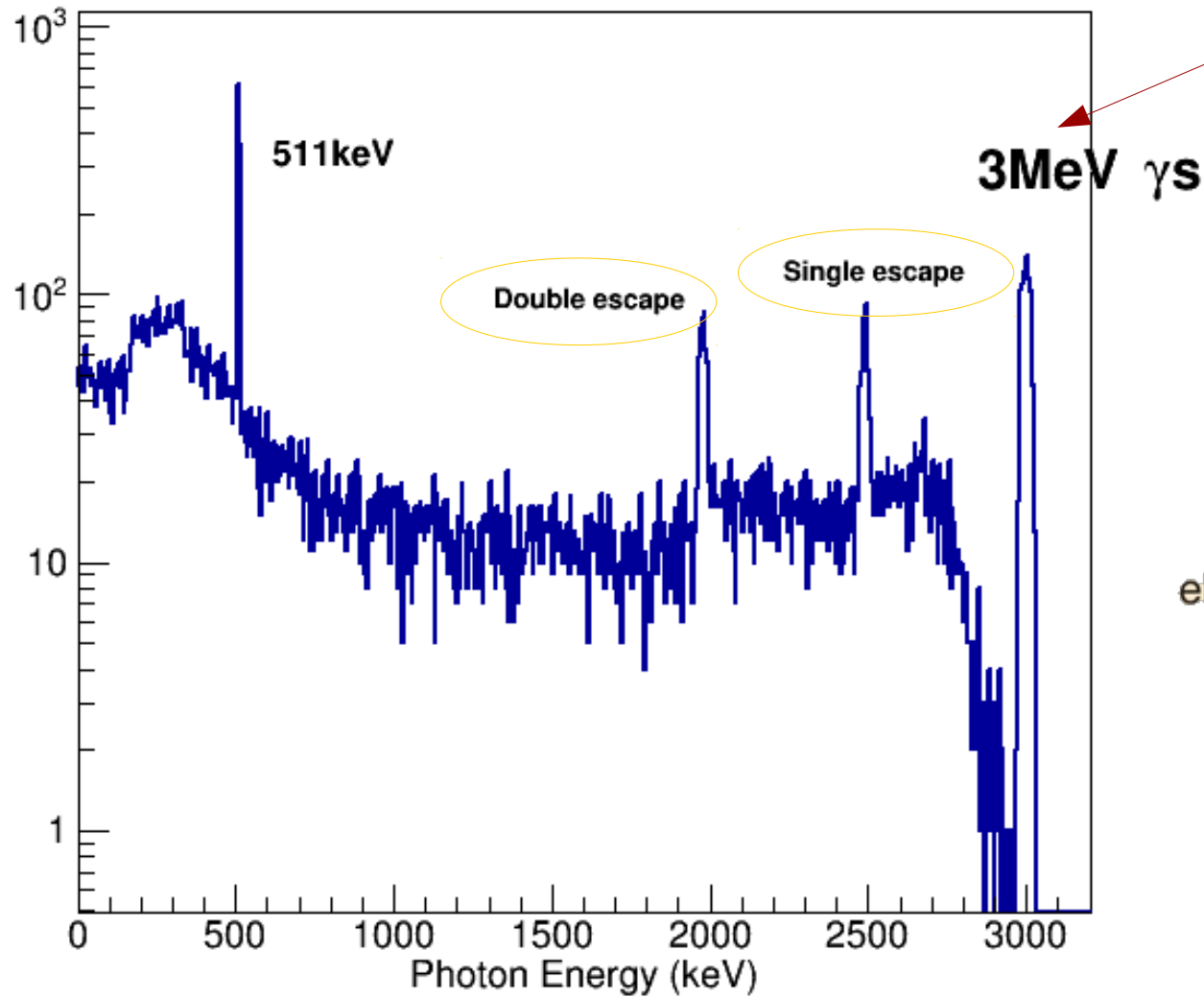


Mid-range energies:  
Compton domination

$$\Delta\lambda = \lambda_c (1 - \cos\theta)$$



# Understanding the spectrum



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**Nuclear reactions have intermediate excited states**



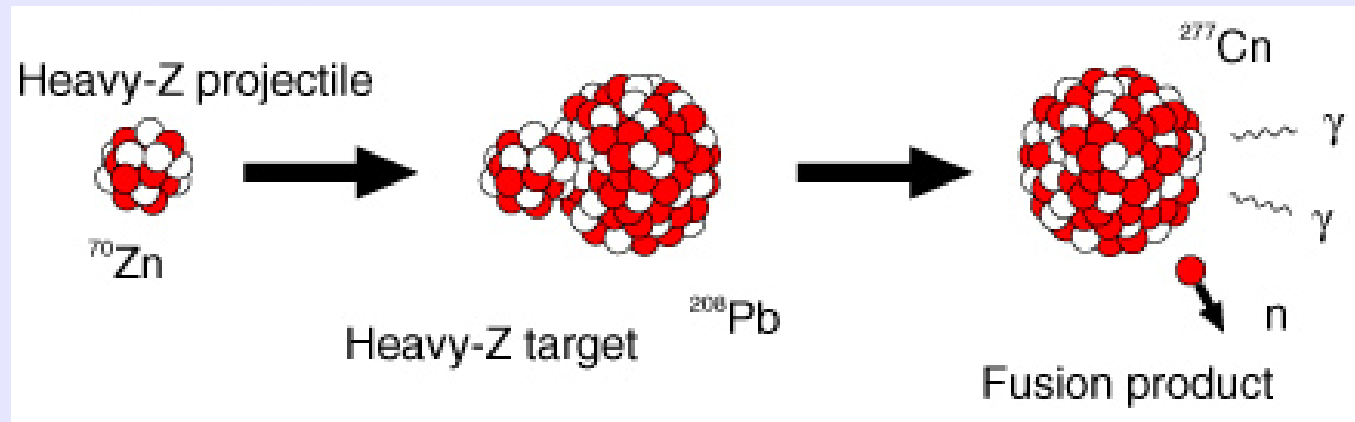
**Gamma emission**



**We can measure the energy, spin and parity of the excited states (EM transitions)**

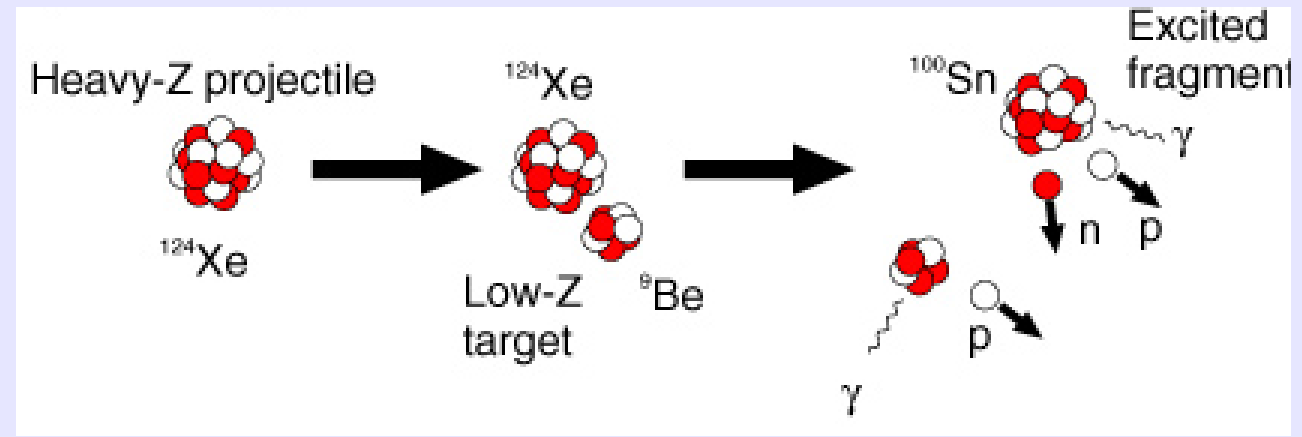
**We can understand:**

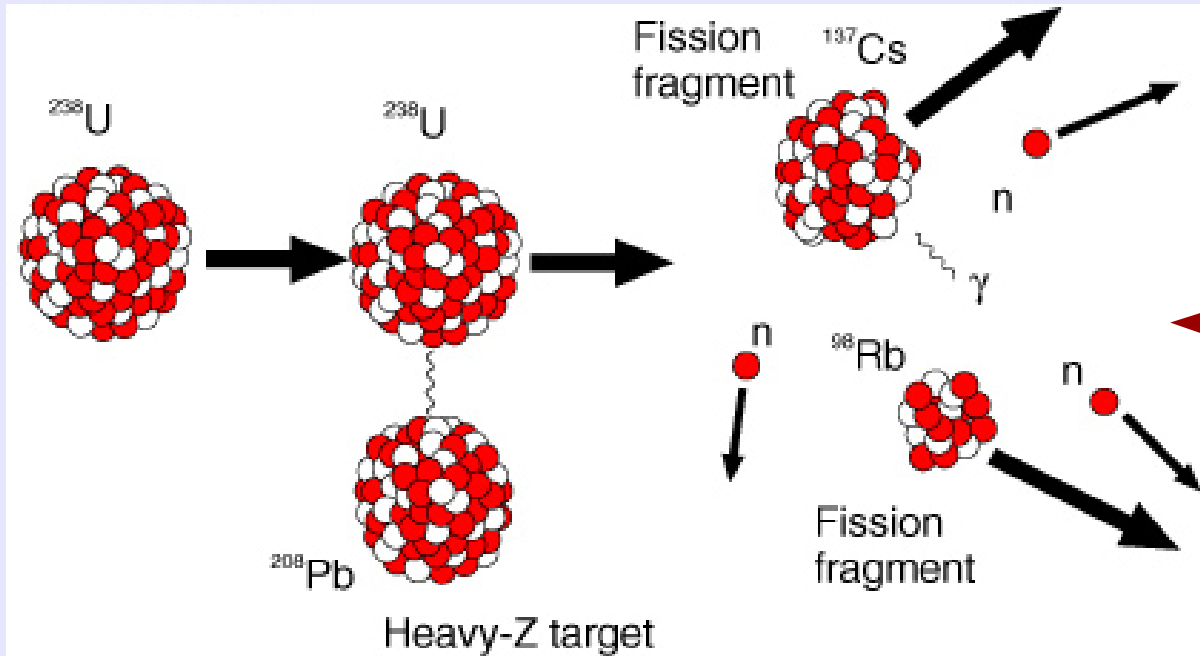
- **Collective excitations**
- **Phase transitions**
- **Nuclear structure and shell model**
- **EOS**



**Nuclear Fusion**

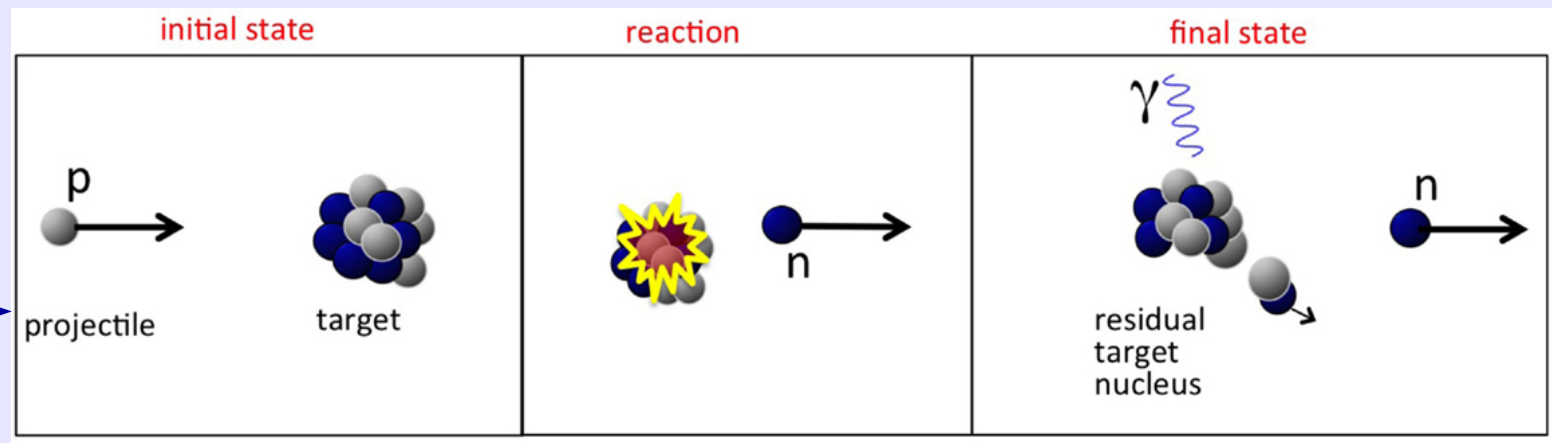
**Nuclear Fragmentation**

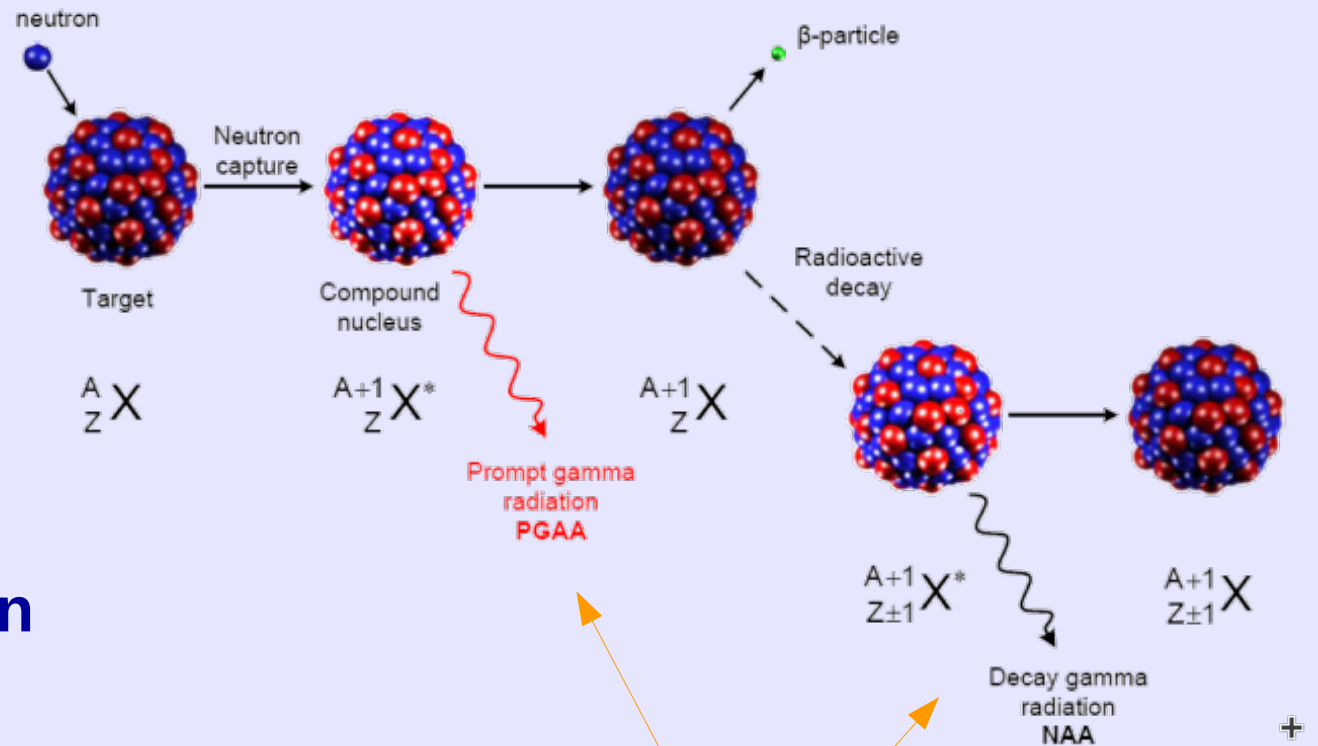




**Nuclear Fission**

**Light projectile fragmentation**



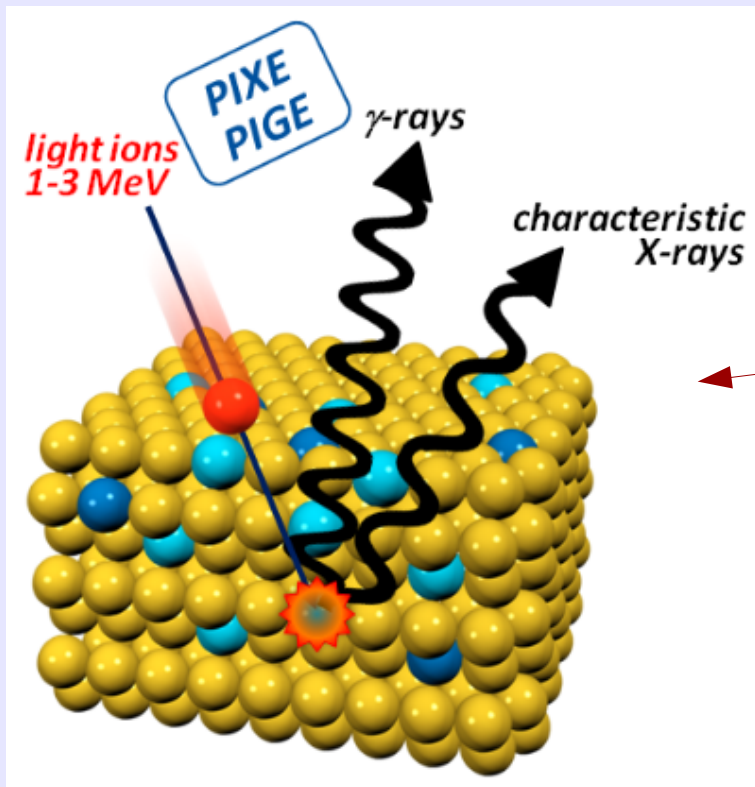


## Beta delayed neutron emission

### Essential to:

- Control reactions in Nuclear Power plants
- Stellar nucleosynthesis
- Nuclear Structure (n-rich nuclei)

$\gamma$  emission

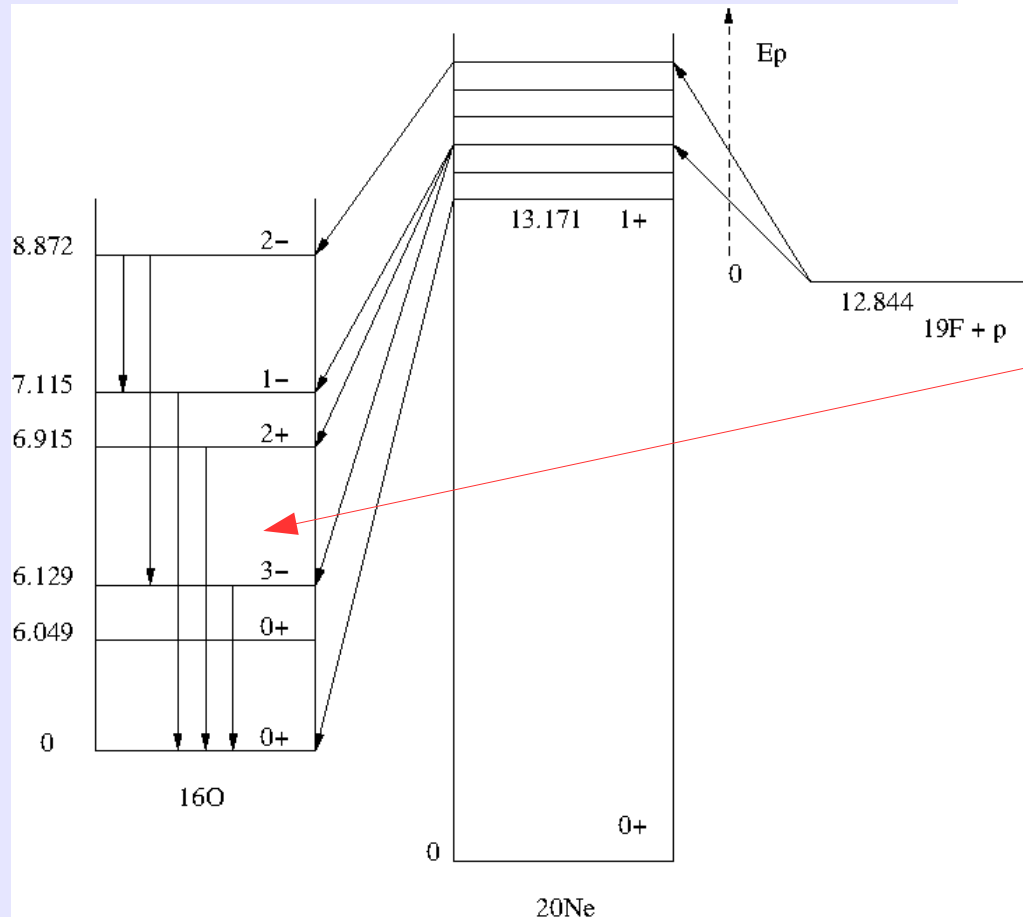
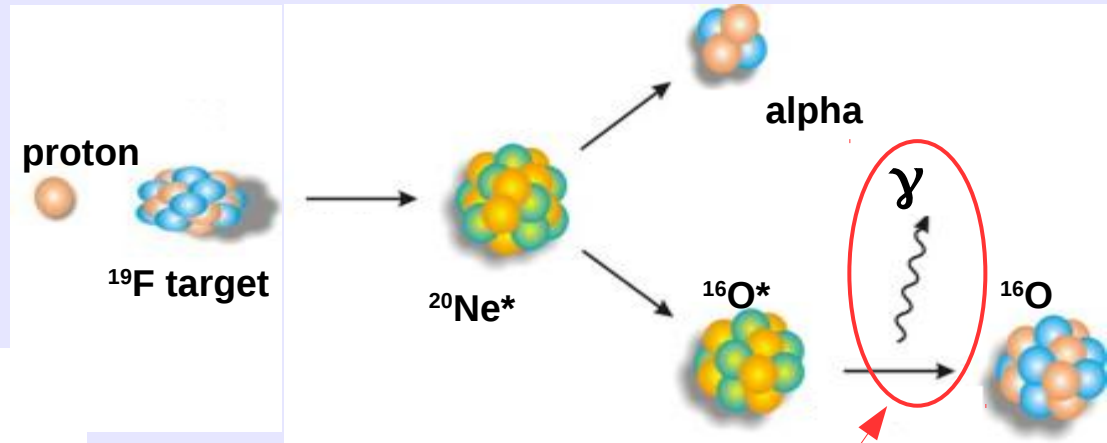


## Proton induced X-Ray and Gamma emission

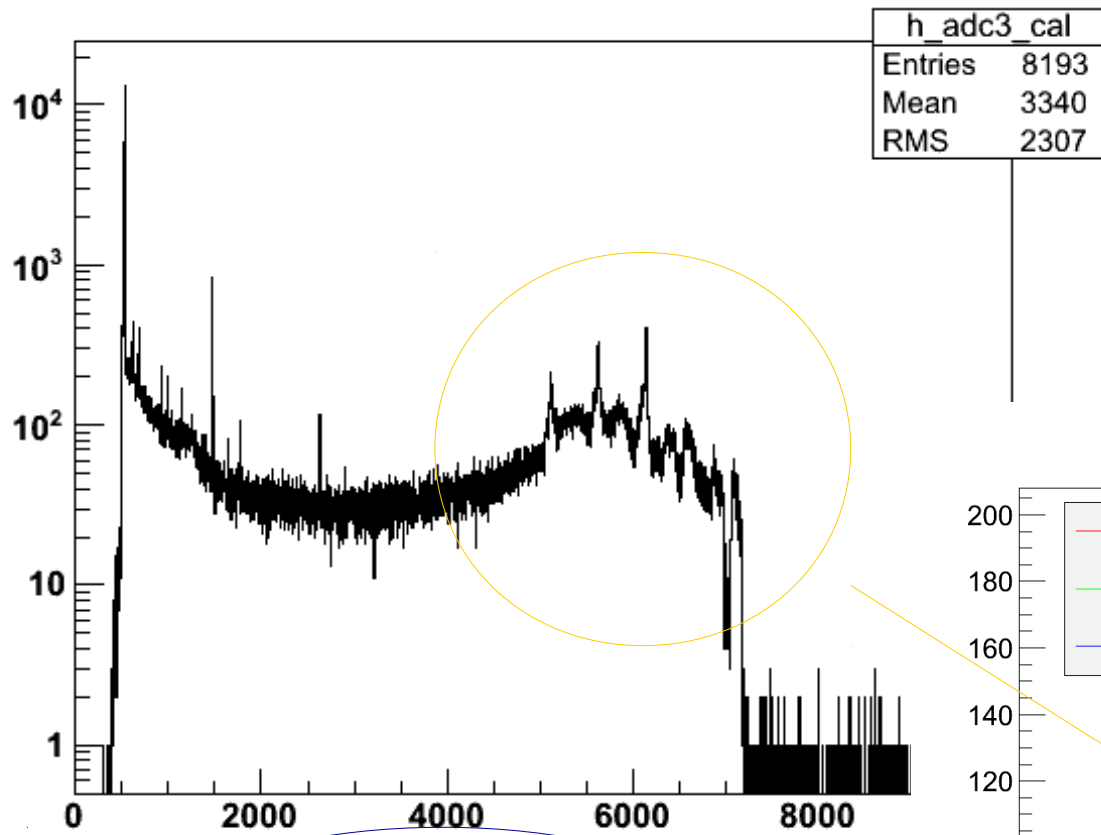
“Low scale” facilities  
(Tandem or Van de  
Graaff accelerators)



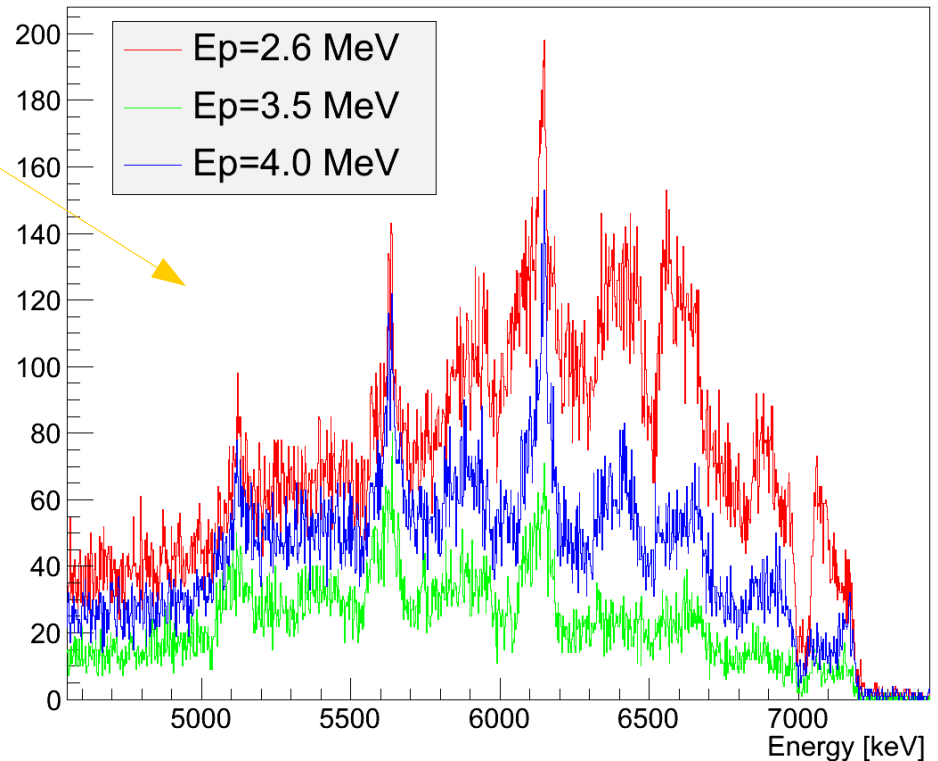
## PIGE Reaction



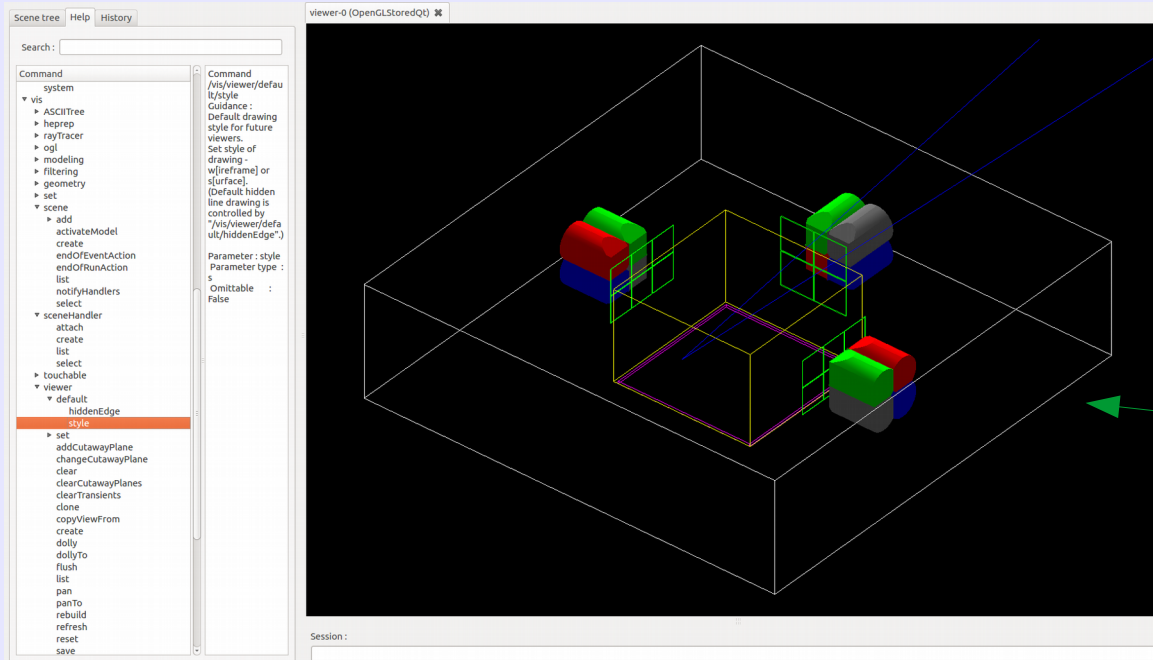
**Gamma emission**



**HE gamma rays!!!**  
**Easy to produce**



**PIGE Reactions:**  
**very powerful test beam**  
**candidates!**



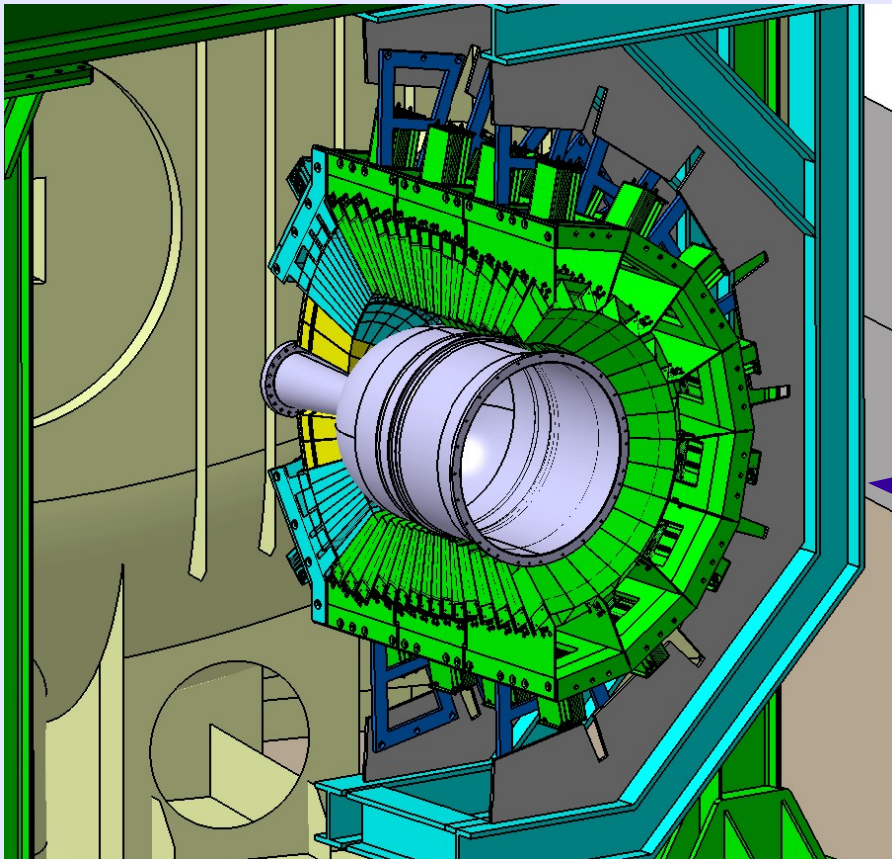
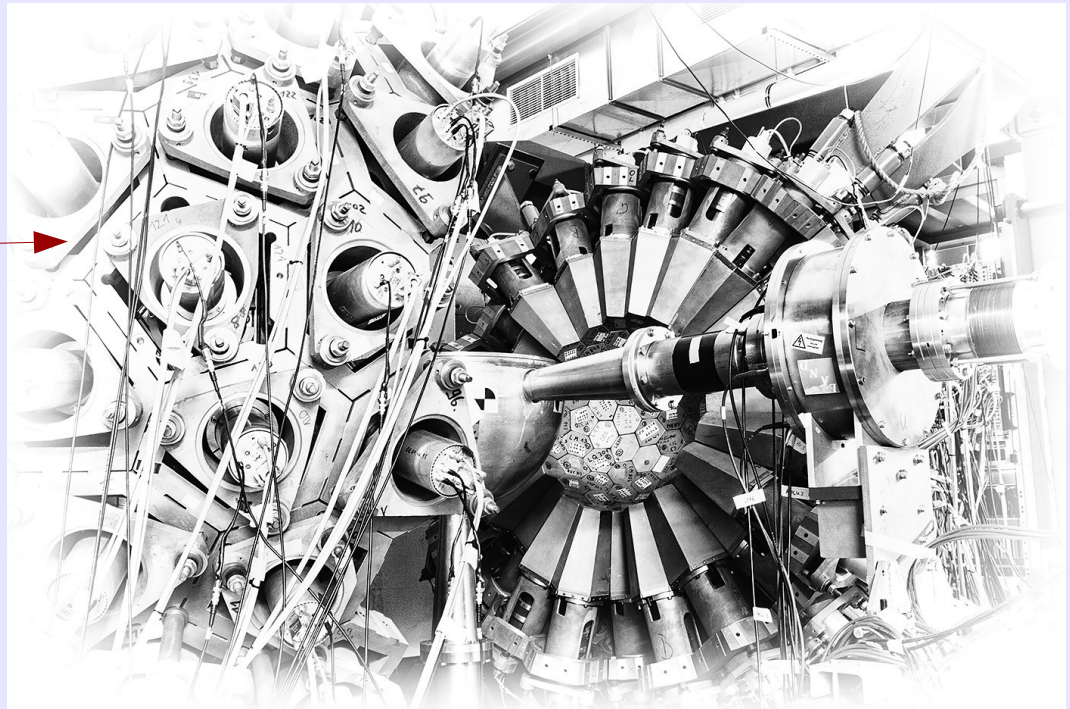
**ACTAR @ GANIL (France)**  
**3 Exogam clovers**  
**(4 HPGe detectors each)**

**8 $\pi$  @ TRIUMF (Canada)**  
**40 HPGe detectors**





**CrystalBall @ GSI (Germany)**  
**162 NaI(Tl) crystals + PMTs**



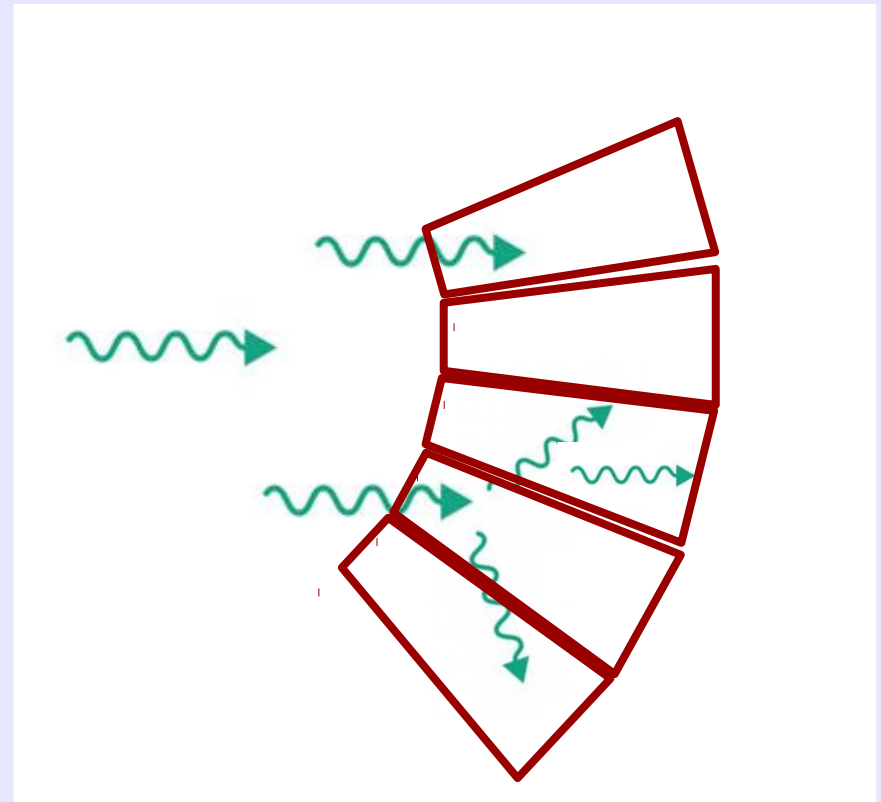
**CALIFA @ FAIR (Germany)**  
**2432 CsI crystals + APDs**  
**96 LaBr3/LaCl3 + PMTs**

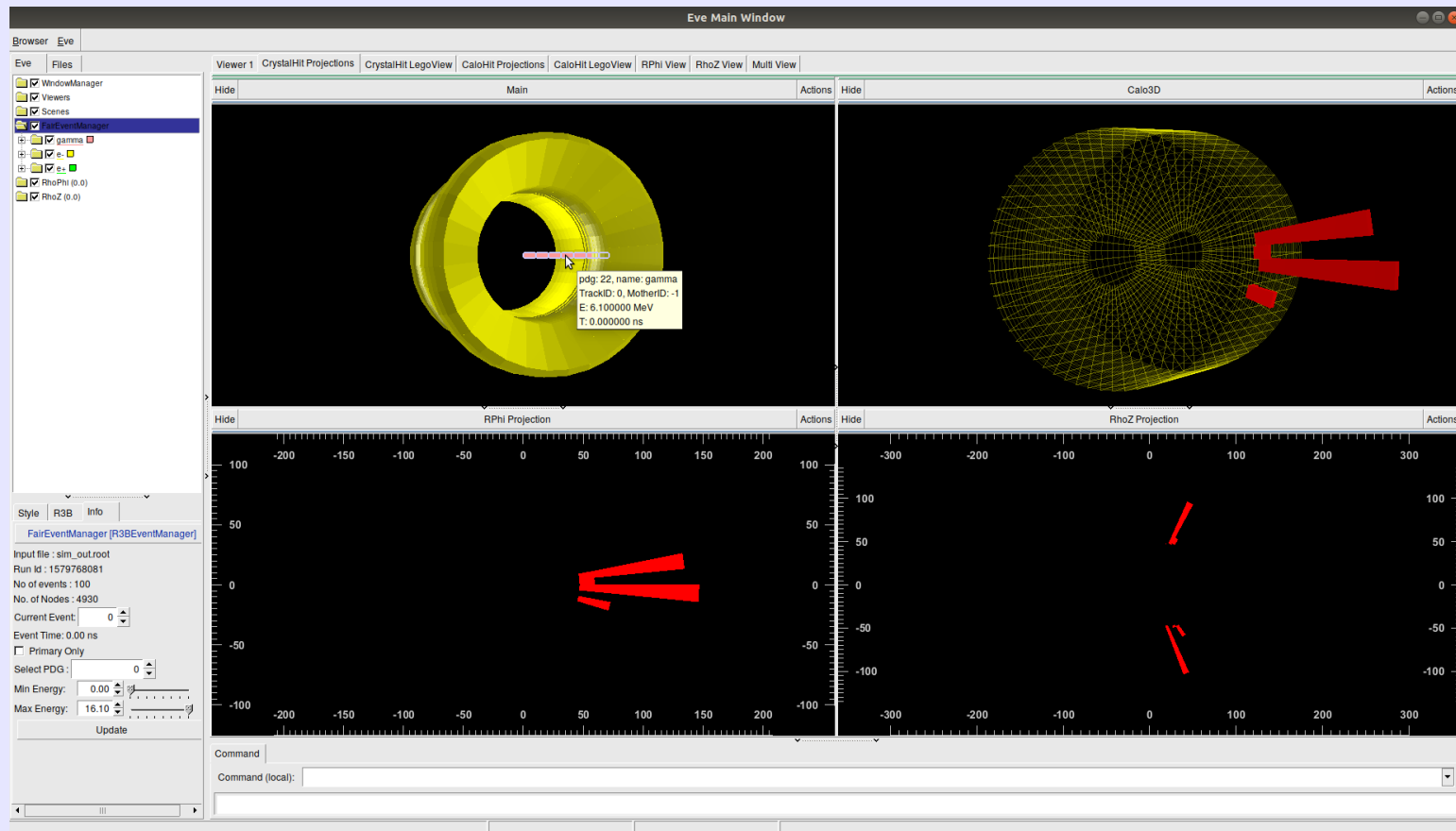
- **All are segmented arrays!**
- **Wide range of gamma ray energies!**
- **Photons can fired more than one detection unit**



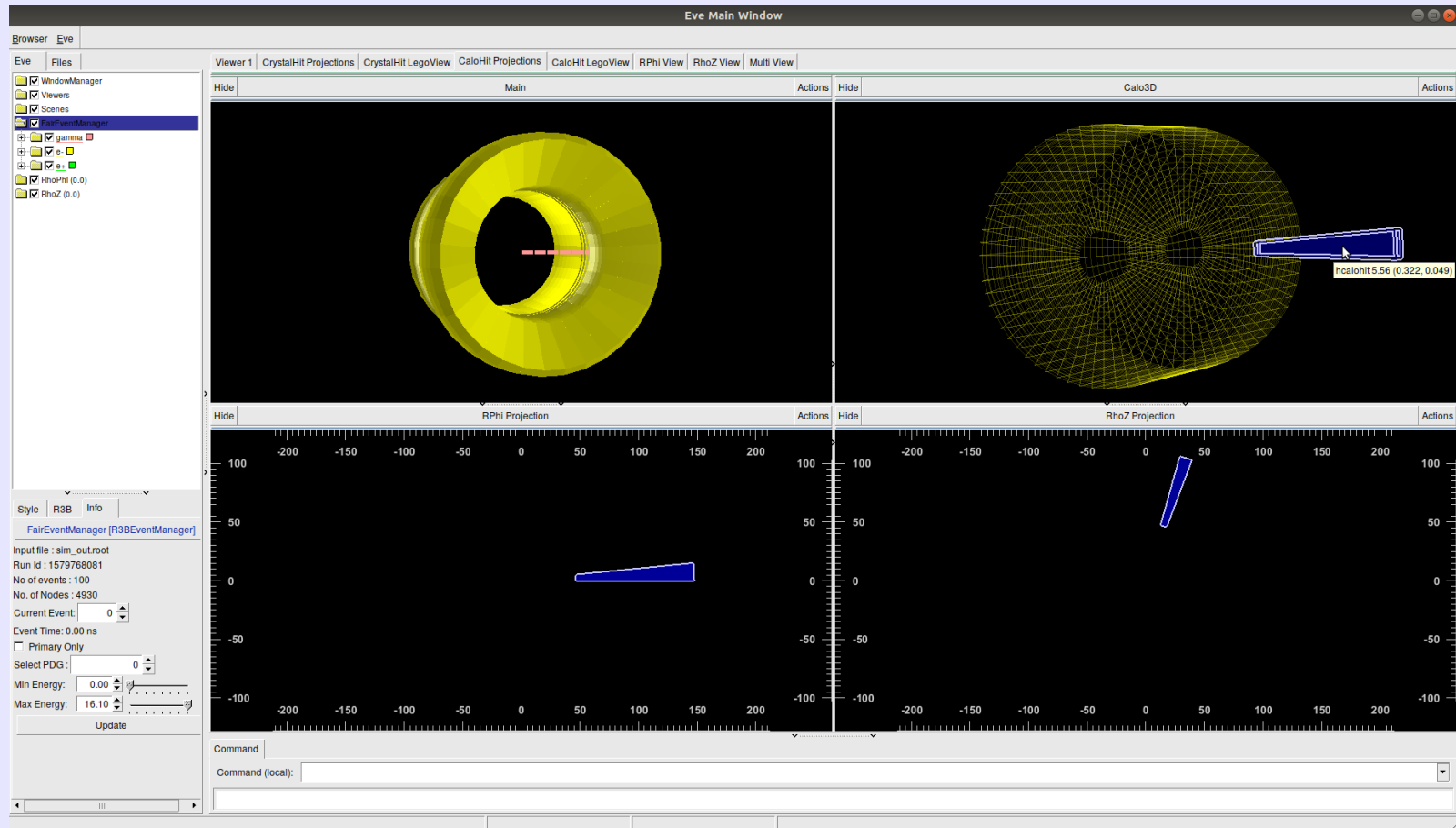
**Addback procedures needed!!**

- **Sum over neighbors**
- **Cluster finding methods**
- **Deep/machine learning techniques like ANN**





**Single 6.1 MeV  $\gamma$ -ray  $\longrightarrow$  At least 6 crystals fired!!**



**After Addback procedure**  $\longrightarrow$  **All released energy at one location!!**

- **Nuclear Physics reactions produce gamma-rays**
- **Gamma-ray calorimetry is required**
- **A good understanding of gamma interactions in our detectors is needed**
- **Experimental setups involve different kind of detector systems and gamma detectors arrays**
- **Dedicated analysis techniques are mandatory: calibrations, addback...**



***Thank you***  
***for your attention!***