



Recoil separators in astrophysics

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Outline

- Introduction
- Radiative capture in inverse kinematics
- DRAGON@TRIUMF
 - Separator
 - Gas target
 - BGO array
 - Focal plane detectors
- Data analysis
- Summary





Radiative capture

- Defined as any reaction A(b,γ)C whereby a bound state of the final nucleus is reached via emission of a photon (can be direct, or via particle-unbound states)
- Ubiquitous in astrophysics, from BBN (e.g. D(p,γ)³He ... ³He(α,γ)⁷Be), to the ppchain, to the CNO cycles, to explosive nucleosynthesis
- Radiative capture reactions are relatively weak due to the fine structure constant, but span a large range at applicable astrophysical energies: (range of cross sections) → intense beams, low background,...

Neutron number



 Often in competition only with β+ decay (limiting our discussion to charged-particle induced), or p, α, n emission

25 26

27 28

- Traditionally have been studied using γ rays, via light particle induced reactions
- Some radiative capture reactions on stable nuclei remain unknown to sufficient precision for stellar evolution or astronomy
- Most radiative capture reactions on short-lived radioactive nuclei in stars remain <u>unmeasured</u>

Inverse Kinematics

- Necessary due to not being able to make short-lived radioactive targets
- Detect recoiling product nucleus: forward focused, 100% efficiency
- Becomes problem of separating rare reaction products from abundant beam

 \rightarrow zero-degree electromagnetic separator

• Additional advantages:

-Target either H_2 or He: gaseous at stp \rightarrow windowless (jet,extended), purified etc...

-Still detect gamma rays (tag)

-Particle ID on reaction products



- **Maximum** possible **recoil angle** when E_{γ} is maximized for $E_{\gamma} = Q + E_{c.m.}$
- AND emission perpendicular to beam axis ($\theta_{\overline{C}} \pi/2$)



Reactions per
incident ionStopping power
$$Y(\infty) = \frac{\lambda^2}{2} \frac{M+m}{m} \epsilon^{-1} \omega \gamma$$
Stellar reaction rate over isolated, narrow
resonance: $N_A < \sigma \nu >= 1.54 \times 10^{11} (\mu T_9)^{-3/2} \omega \gamma \exp(-11.605 \frac{E_R}{T_9})$





Detector of Recoils And Gammas Of Nuclear reactions

The DRAGON recoil separator



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- DRAGON designed for mass < 30
- Differentially pumped windowless hydrogen or helium gas target
- Suppression ~ $10^{-12} 10^{-15}$
- Acceptance around 20 mrad
- Existing magnets/dipoles field limits
 - MD1 ~ 5500 G (could go higher but at limit of NMR probe)
 - ED1 ~ 200 kV (rated to 300 kV)
- Lowest beam intensity ~ 10⁶ pps
- Highest beam intensty ~ few 10¹² pps
- Lowest mass reaction studied ⁴He(³He,γ)⁷Be
- Highest mass reaction studied ⁷⁶Se(α,γ)⁸⁰Kr
 - Required additional 'charge state booster'

DRAGON gas target

Windowless, differentially pumped, recirculating gas target (H₂ or He) 1-10 mbar (pumping constraints) $\rightarrow \leq 6 \times 10^{18}$ at/cm² H₂

LN₂ cooled zeolite cleaning trap





Large pressure drop to $\sim 3x10^{-6}$ Torr

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DRAGON gas target – DON'TS!

We learned the hard way...



- Pumping Xenon resulted in catastrophic failure of 6 turbo pumps!
- High atomic weight → noble gases generate large quantities of heat when striking the rotor
- Low specific thermal capacity → little heat transfer to stator or housing

→ High rotor temperatures!

DRAGON BGO array

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DRAGON BGO array

- **BGO** (Bi₄Ge₃O₁₂) array (30 detectors)
- High γ-ray detection efficiency (40 to 80%, depending on multiplicity & energy)
- Combined with TOF → low random coincidence rate!
- <u>Caveat:</u>

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- Rely on **simulation** for detection efficiency
- → dominates syst. error of the experiment!
- Limited γ-ray energy resolution (FWHM ~9%)
- Segmented BGO array along beam axis
 → Information about location of reaction
- BGO hit pattern → resonance energy (0.5%)



BGO "Hit pattern" Method



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Use BGO crystal z-coordinate to extract resonance position within extended gas targetbut needs sufficient statistics

UPGRADE to LaBr would allow method using timing (BGO timing too slow)





Focal plane detectors – particle detection and identification



¹⁸F(p,γ)¹⁹Ne - Event identification



FIG. 1. Energy loss vs. energy loss plot obtained from the first two anodes in the IC. The attenuated beam run is shown in black and two well-separated loci are clearly visible signifying the presence of both ¹⁸F and ¹⁸O. Circles (triangles), both red online, correspond to observed ¹⁹Ne (¹⁸F) events when the separator was tuned to recoils. Squares (blue online) correspond to ¹⁹F recoils during the separate ¹⁸O beam run.

Beam intensity ~ 10^6 pps – too low for reliable FC reading. Normalisation using elastic scattering.

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⁵⁸Ni(p,γ)⁵⁹Cu measurement



Fig. 2. Energy loss vs. total energy for (a) singles, (b) recoils coincident with γ -rays (black circles) and (c) recoils coincident with γ -rays with additional condition on separator and MCP time-of-flight (red squares). In panels (b) and (c) the shaded area denotes the singles histogram plotted for reference. Black contour denotes the range of true events.



Resonance strength = 0.303(65) eV

A. Simon et al., Ep. J.

Hybrid detector

Combine properties of IC (ΔE) and DSSSD (operation, position sensitivity & resolution) in hybrid detector







Data analysis

Data analysis

To extract a cross section or resonance strength, we need to know:

Corrected yield -

 charge state fraction – measured or formula based on previous DRAGON data

•gas target transmission efficiency – measured using Faraday Cups
•separator transmission efficiency – GEANT simulation

- •focal plane detector efficiences measured
- •DAQ deadtime measured

•BGO efficiency, if running coincidences – GEANT simulation based on data

Beam normalisation

•Intensity measured in Faraday cups before and after every run

•Relative changes in intensity from elastic scattering rate in target

Stopping power – SRIM code shows 20-50% deviation from experiment
measure energy loss across target - using MD1 with and without gas
know effective length of target – extensive studies using sources and known resonances





Courtesy of Anna Simon

My Sayer code agrees with TUDA CSD \rightarrow I'm going to use it for other ions.

24 charge state fraction [%] 22 12 10 20 22 18 24 26 16 charge state

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Solid red line - Gaussian fit to my Sayer code

Dashed red line – the fit shifted to match the experimental results (the same area under curve and sigma, just varied the centroid)

Useful tools: rossum

- Automated running
 - runs started and stopped automatically every hour
 - key parameters recorded and saved into data stream
 - between each run, all Faraday cup readings taken
 - warnings by text and email
 - allows remote running with stable beam no night shifts!



Summary

DRAGON designed to study nuclear reactions relevant for nuclear astrophysics in inverse kinematics

<u>"Strengths"</u>

- •Gas target (variable gas & pressure) → enables radioactive beam exp.
- •High beam suppression
- •Location (access to beams)
- •γ-coincidence measurements
- Variable end-detector system
- •TOF (local & separator)
- •RF Timing
- Beam Diagnostics
- Stopping power measured

"Weaknesses"

- •Limited rigidity
- \rightarrow Limitations for higher masses
- •Simulation required for detection efficiency
- •Limited γ-energy resolution w BGO array
- •Limited angular acceptance
- •Due to extended target, separator transmission can depend on
- resonance position in gas target

Successful DRAGON programme for over 15 years

Reaction	Motivation	Intensity (s ⁻¹)	Purity (beam:cont.)
²¹ Na(<i>p,γ</i>) ²² Mg	1.275 MeV line emission in ONe novae	5 x 10 ⁹	100%
¹² C(<i>α</i> ,γ) ¹⁶ O	Helium burning in red giants	3 x 10 ¹¹ to 1 x 10 ¹²	
^{26g} Al(<i>p,γ</i>)² ⁷ Si	Nova contribution to galactic ²⁶ Al	3 x 10 ⁹	30,000:1
¹² C(¹² C,γ) ²⁴ Mg	Nuclear cluster models	3 x 10 ¹¹	
⁴⁰ Ca(<i>α,γ</i>) ⁴⁴ Ti	Production of ⁴⁴ Ti in SNII	3 x 10 ¹¹	10,000:1 – 200:1
²³ Mg(<i>p,γ</i>) ²⁴ Al	1.275 MeV line emission in ONe novae	5 x 10 ⁷	1:20 – 1:1,000
¹⁷ Ο(<i>α,γ</i>) ²¹ Ne	Neutron poison in massive stars	1 x 10 ¹²	
¹⁸ F(<i>p,γ</i>) ¹⁹ Ne	511 keV line emission in ONe novae	2 x 10 ⁶	100:1
³³ S(<i>p</i> , <i>γ</i>) ³⁴ Cl	S isotopic ratios in nova grains	1 x 10 ¹⁰	
¹⁶ Ο(<i>α,γ</i>) ²⁰ Ne	Stellar helium burning	1 x 10 ¹²	
¹⁷ O(<i>p,γ</i>) ¹⁸ F	Explosive hydrogen burning in novae	1 x 10 ¹²	
³ He(<i>α,γ</i>) ⁷ Be	Solar neutrino spectrum	5 x 10 ¹¹	
⁵⁸ Ni(<i>p,γ</i>) ⁵⁹ Cu	High mass tests (p-process, XRB)	6 x 10 ⁹	
^{26m} Al(<i>ρ,γ</i>) ²⁷ Si	SNII contribution to galactic ²⁶ Al	2 x 10⁵	1:10,000
³⁸ K(<i>p,γ</i>) ³⁹ Ca	Ca/K/Ar production in novae	2 x 10 ⁷	1:1
¹⁹ Ne(<i>p,γ</i>) ²⁰ Na	¹⁹ F abundance in nova ejecta	2 x 10 ⁷	1:1 to 4:1
²² Ne(<i>ρ,γ</i>) ²³ Na	NeNa cycle; explosive H burning in classical novae	2 x 10 ¹²	
⁷ Be(α,γ) ¹¹ C	v-p process		1:200 to 1:1000

Thank you for your attention

DRAGON core collaboration

Barry Davids, Dave Hutcheon, Chris Ruiz (TRIUMF) Brian Fulton, Alison Laird (York) Gavin Lotay (Surrey) Uwe Greife (Colorado School of Mines) Ahmed Hussein (UNBC) Alan Chen (McMaster)

Thanks to Annika Lennarz and and Chris Ruiz for slides.



John D'Auria 1939-2017 TISOL → ISAC DRAGON

Radiative capture in inverse kinematics

- Radionuclei that can't be made into a target
 - accelerate them, make H or He target.
- Gaseous H₂ or He target
- Detect recoils in coincidence with γ rays
- Measure stopping powers directly (required for ωγ)
- Change target thickness easily
- Can make target 'extended'
 - (sensitivity to resonance position \Rightarrow resonance energy)
- Target does not degrade over time
- Target elastic scattering for normalization

Max recoil angle \leftarrow kinematics & decay scheme Need recoil separator with full acceptance *and* high beam suppression \leq



$$N_A \langle \sigma v \rangle = 1.54 \times 10^{11} (\mu T_9)^{-3/\omega} \gamma \exp\left(-11.605 \frac{E_R}{T_9}\right)$$

Thick target yield function in inverse kinematics

$$\underbrace{\mathcal{Y}(\underbrace{\underbrace{}})}_{\mathcal{Y}} = \frac{l^2}{2} \frac{M+m}{m} \frac{1}{\mathcal{C}} Wg$$



Opening angle and therefore detection efficiency depend on energy.

Separator beam suppression

- High intrinsic beam suppression: 10⁸ to 10¹³ (proton capture)
- Depends on **beam energy & emittance**
- >10¹⁴ raw suppression demonstrated for ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$
- Coincidence measurement with prompt γ-rays & PID cuts & TOF
 - \rightarrow suppression factor of ~10¹⁵ for

p-capture & ~5x10¹⁷ for α -capture

Beam suppression is **NOT** described by a single number, but determined by mass & charge difference, decay mode, energy, detectors, etc...



Figure from D. Hutcheon et al. NIMRB 266 (2008)

Event identification

- Energy / energy loss at focal plane
- MCP local time of flight
- RF timing
- BGO coincidences
 - Separator time of flight
 - BGO energy

Particle ID: dealing with contaminants, background ISOE beams may contain isobaric Atte

- ISOP beams may contain isobaric contaminants
- Tradeoff between ∆M/M of mass separator to transmission (beam intensity)



