

*GroundBIRD*

# GB NOW



**Osamu Tajima (Kyoto University)  
on behalf of GB collaboration**

# The GB Collaboration



**RIKEN** Satoru Mima, Shugo Oguri, Chiko Otani (co-PI), Taketo Nagasaki



**Kyoto** Shunsuke Honda, Takuji Ikemitsu, Junta Komine, Junya Suzuki, Yoshinori Sueno, Osamu Tajima (co-PI)



**KEK** Masashi Hazumi, Hikaru Ishituka, Tomohisa Uchida, Mitsuhiro Yoshida, (O. Tajima as visiting staff)

**NAOJ** Makoto Nagai

**ISAS/JAXA** Yutaro Sekimoto



**Tohoku** Makoto Hattori, Fumiyasu Kanno, Hiroki Kutsuma, Tomoka Okada

**Tokyo** Kenji Kiuchi, Makoto Minowa, Nozomu Tomita

**Saitama** Ryo Koyano, Masato Naruse, Toru Taino

**Korea** Yongil Jo, Kyungmin Lee, Junhyeok Moon, Eunil Won

**IBS** Jihoon Choi

**TU Delft** Kenichi Karatstu

**IAC** Ricardo Génova-Santos, Rafael Rebolo, José Alberto Rubiño-Martín, Mike Peel

**Japan, Spain, Korea,  
and Netherlands**



# Led by young guys!



S. Honda  
(Kyoto)

J. Komine  
(Kyoto)

H. Kutsuma  
(Tohoku)

S. Oguri  
(RIKEN)

J. Suzuki  
(Kyoto)

T. Nagasaki  
(RIKEN)

# Science of CMB obs.

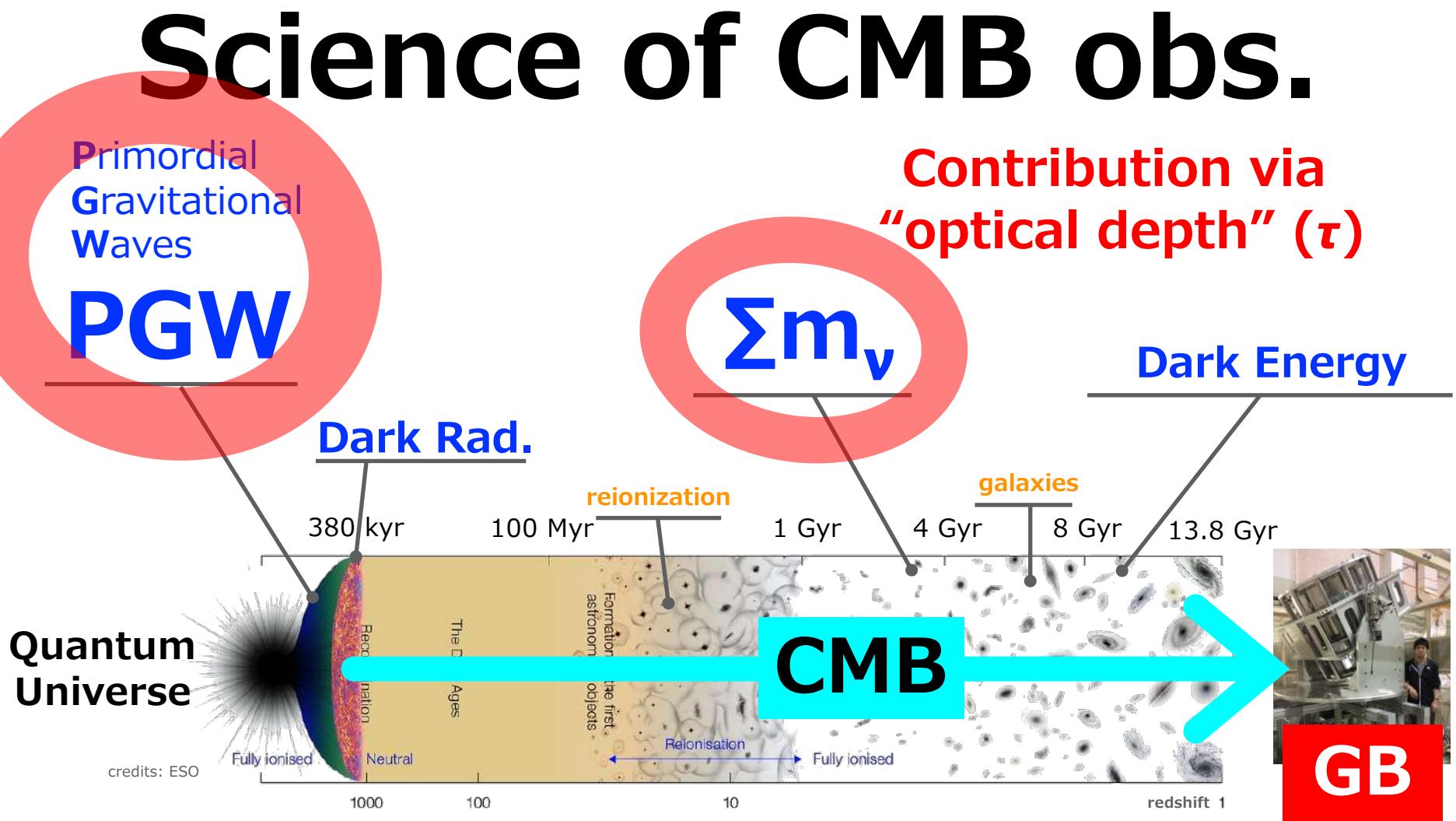
Primordial  
Gravitational  
Waves

**PGW**

Contribution via  
“optical depth” ( $\tau$ )

$\Sigma m_v$

Dark Energy



***GB aims CMB polarization patterns in large angular scale,  $O(1^\circ \sim 10^\circ)$***

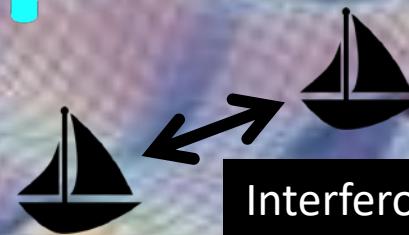
# B-modes from PGW



*Looking at asymmetric  
patterns on wave surface*

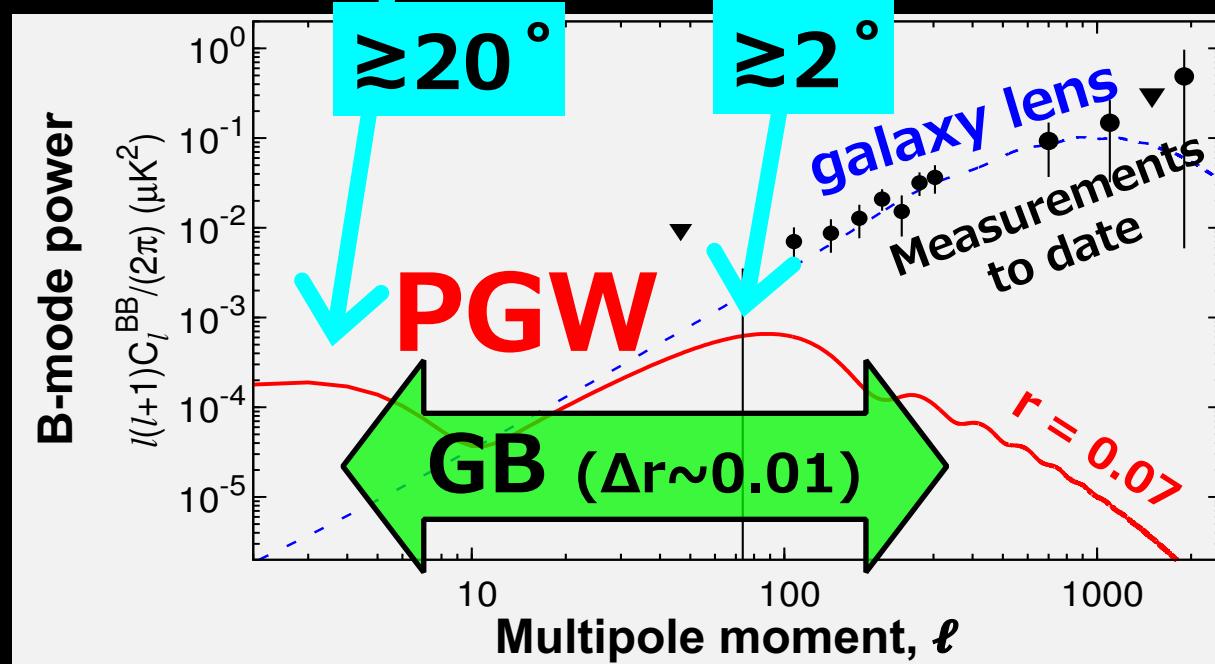
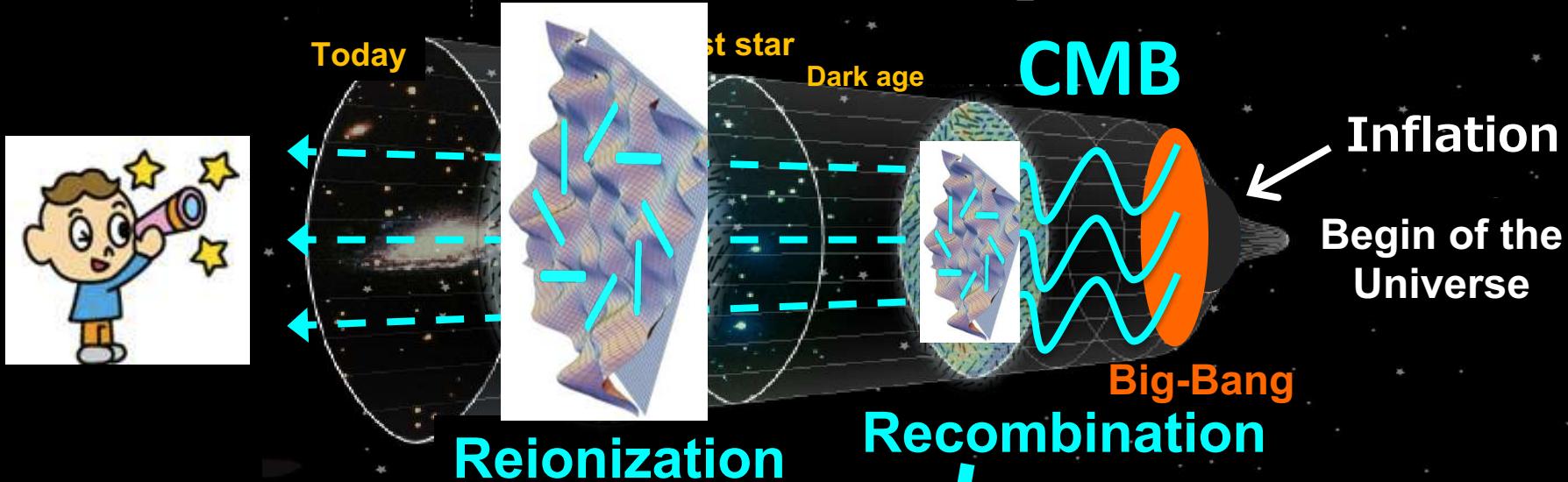
Large angular scale  $> 2^\circ$

Wave surface of PGW



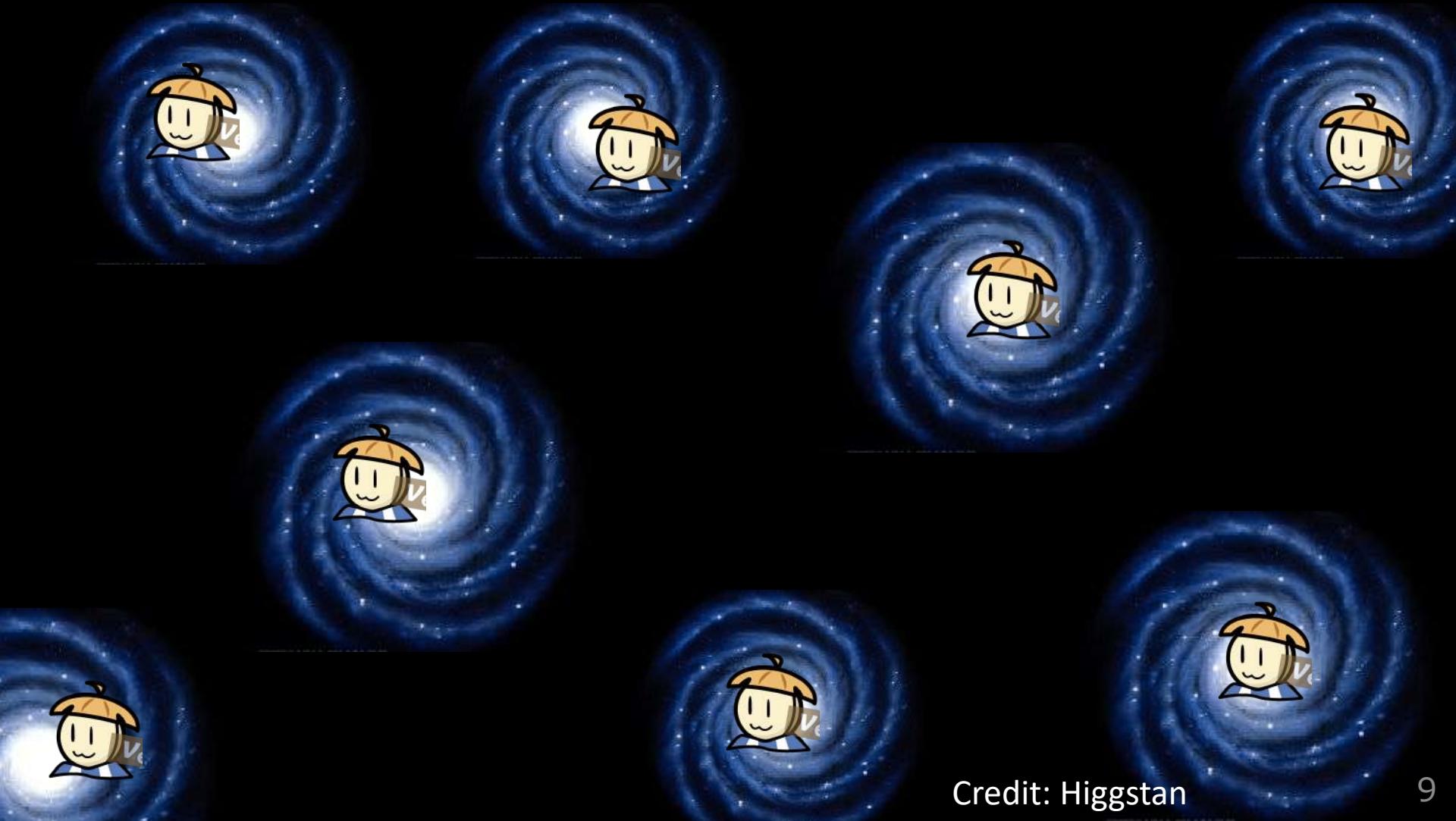
Interferometers, e.g. LIGO  
“look laser fluctuation”

# "Look" B-modes patterns



Courtesy of  
Y. Chinone

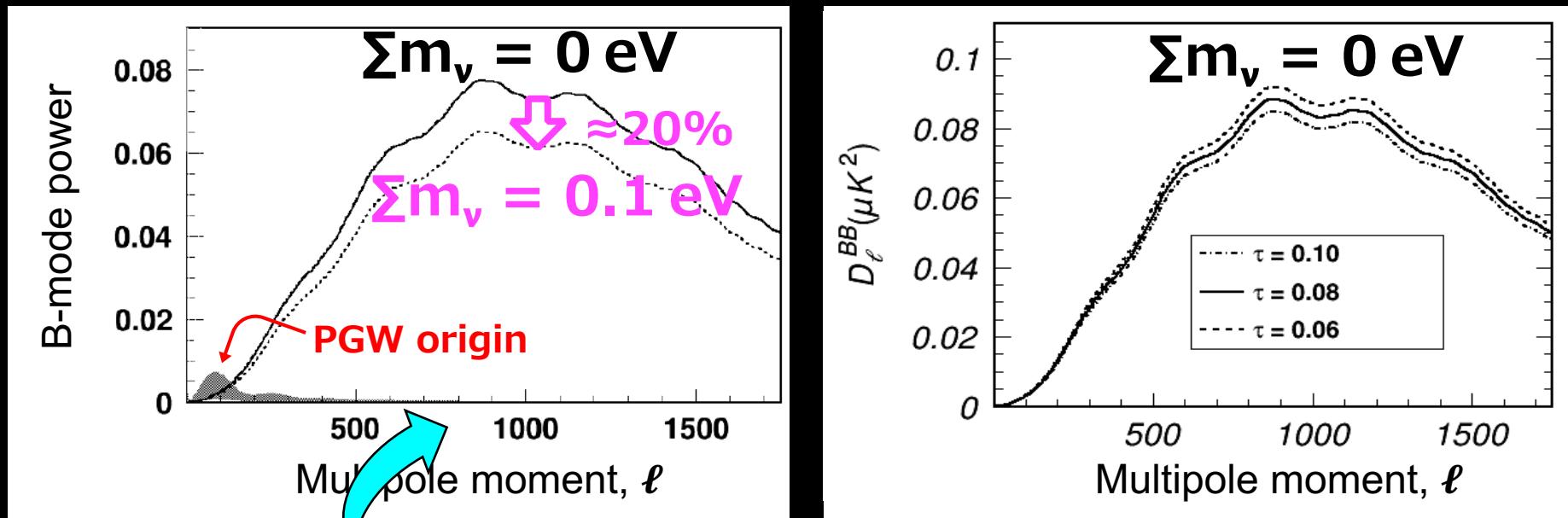
# CvB is unique massive particle NOT localized in galaxy haloes



Credit: Higgstan

# $\Sigma m_\nu$ makes thinner lens

Correlation with  $\tau$  should be unfolded

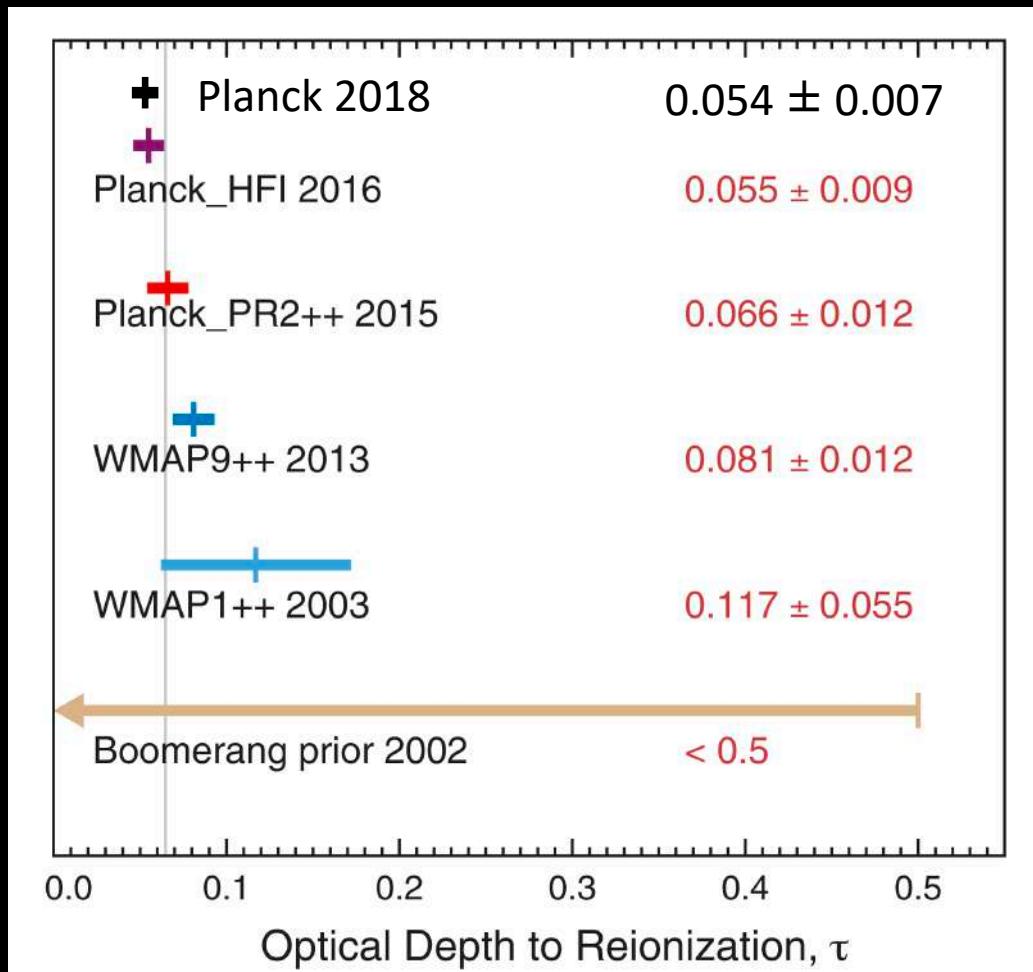


Moon scale  
( $0.5^\circ$ )

Planck 2018:  $\tau = 0.054 \pm 0.007$

$\Delta(\Sigma m_\nu) \sim 0.03 \text{ eV}$

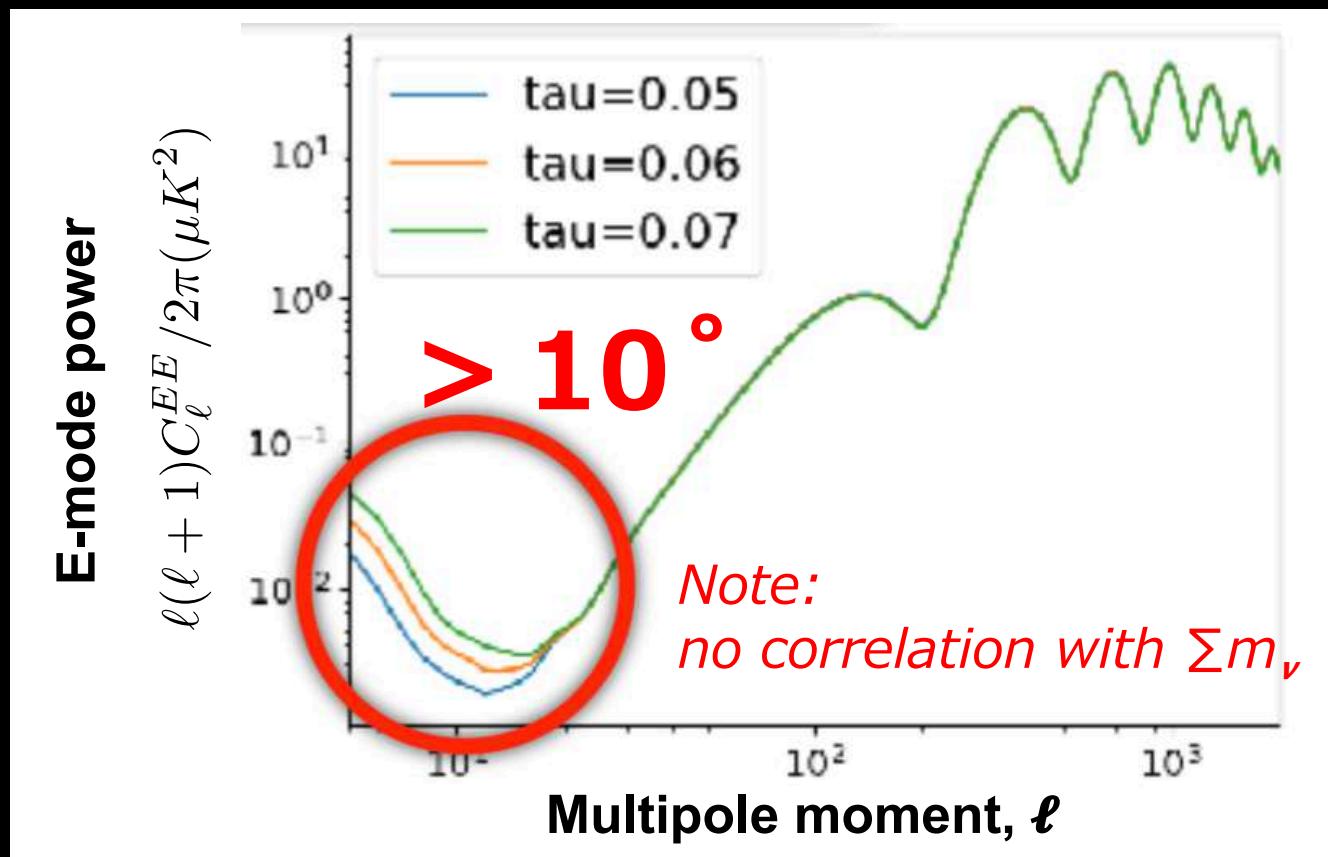
# History of measured $\tau$



collage plot of  
LAMBDA web

*Why don't you measure it ?*

# $\tau \leftrightarrow$ Reionization in E-modes (target $\Delta\tau \sim 0.01$ from single exp.)



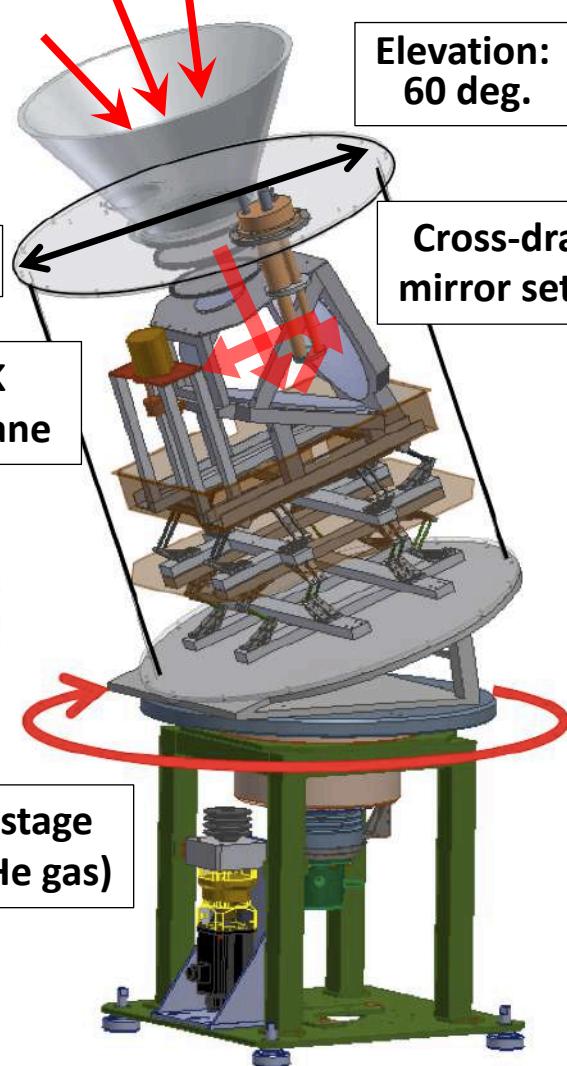
# Experimental concepts to measure $O(10^\circ)$ patterns

- **Wide scan with fast modulation**
  - High-speed Rotation-scan (HsRs)
- **Fast time response for HsRs**
  - MKID ( $\tau \lesssim 100 \mu\text{s}$ )
- **Be sensitive even though no-large focal plane**
  - Cryogenic optics (< 4K)
- **Be robust for foregrounds**
  - DUST-band (220 GHz) + CMB (150 GHz)
  - (future) alliance with QUIJOTE (10-40 GHz)

*Compact size  
is preferred*

# Overview of GB

CMB ( $\text{FOV} \sim 21^\circ$ ,  $\vartheta \sim 0.6^\circ$ )



HsRs (120°/s in Az) mitigates effects of atmospheric fluctuation



This is real scan speed

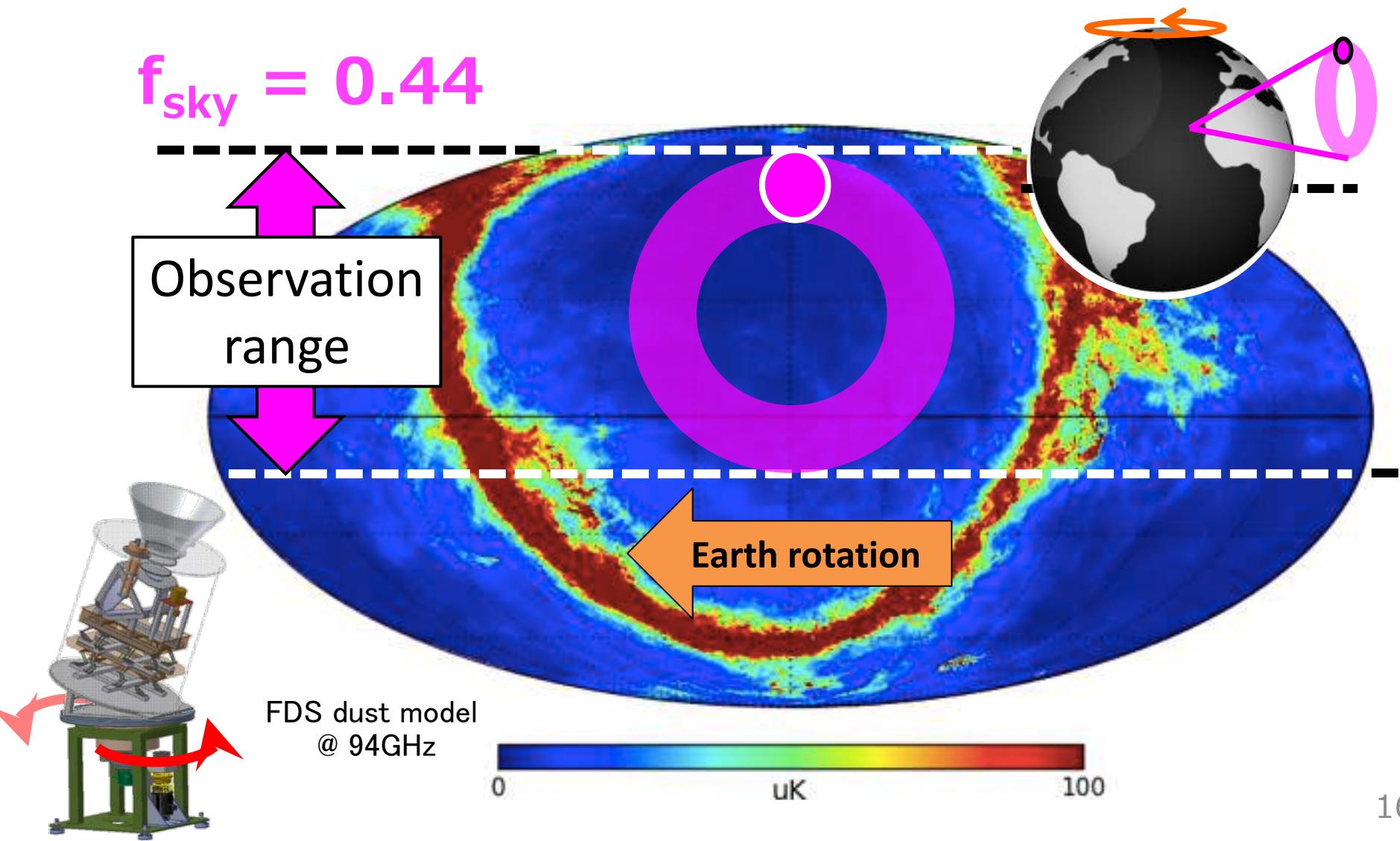
# To be deployed in the Canaries, soon Teide Observatory, 2,400 m alt.



Dry area above clouds  
 $+28.3^\circ$ ,  $-16.5^\circ$



# High-speed Rotation-scan provides large-sky coverage

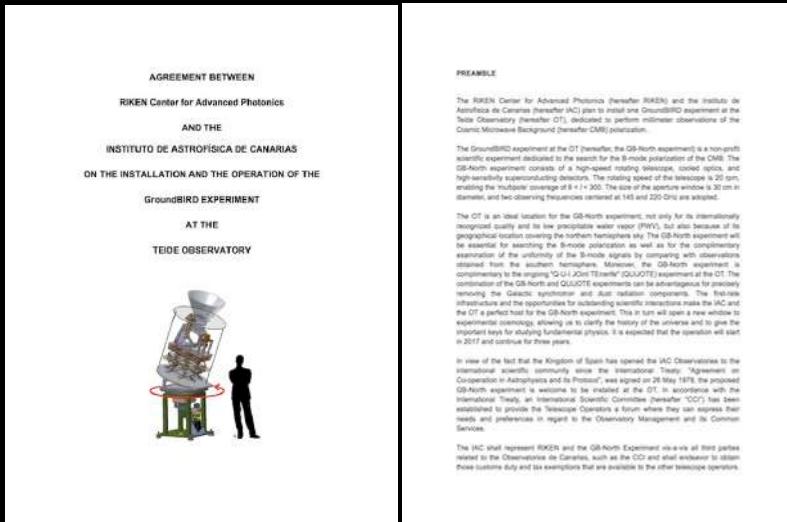


# Prep. observation site

MOU RIKEN btw IAC (May, 2015)



## Agreements (Aug. 2016)



**Ground shielded area for GB  
AC power & network lines are available**

# Dome for weather proofing end of Oct. 2018



# Miscellaneous constructions cable trays, etc. etc. ...

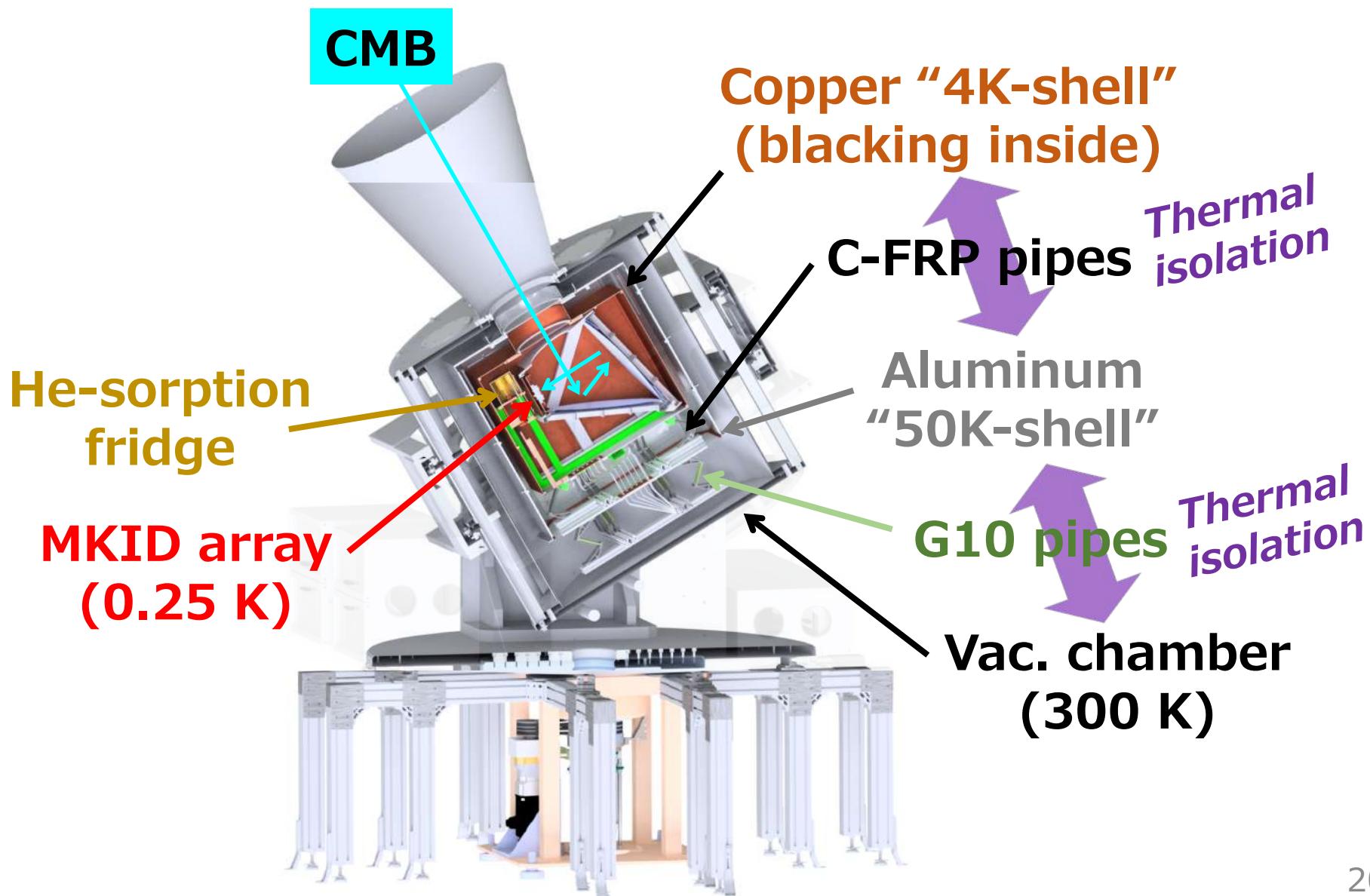
Shugo Oguri  
(RIKEN)

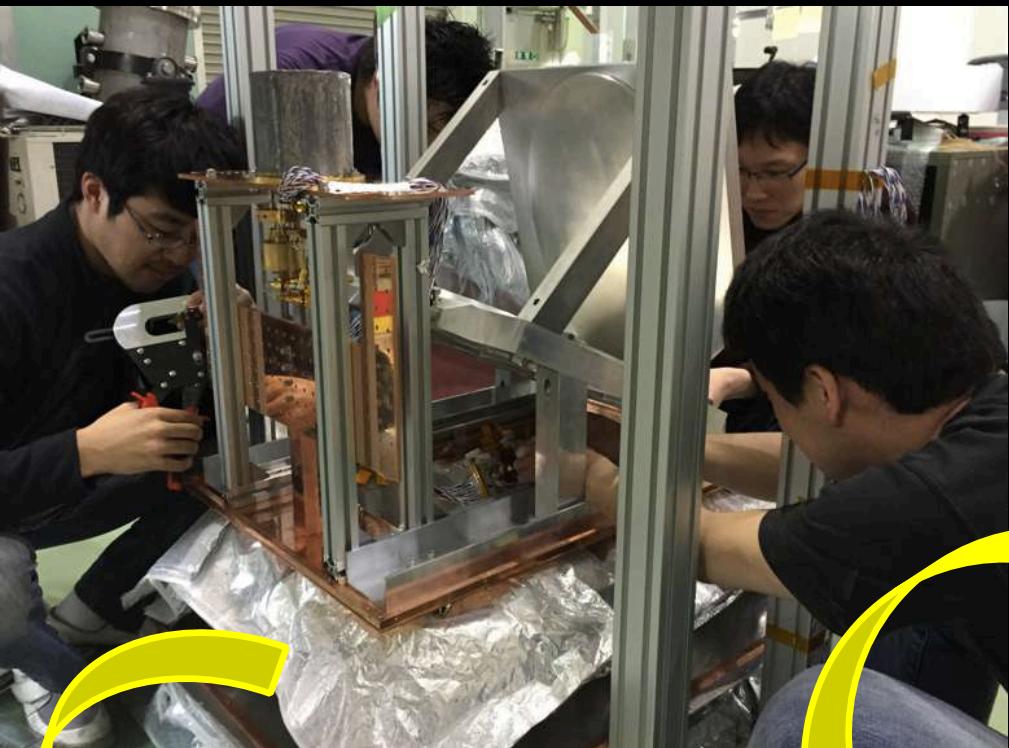


Kyungmin Lee  
(Korea U.)



# Cryogenic optics





**4K-shell**

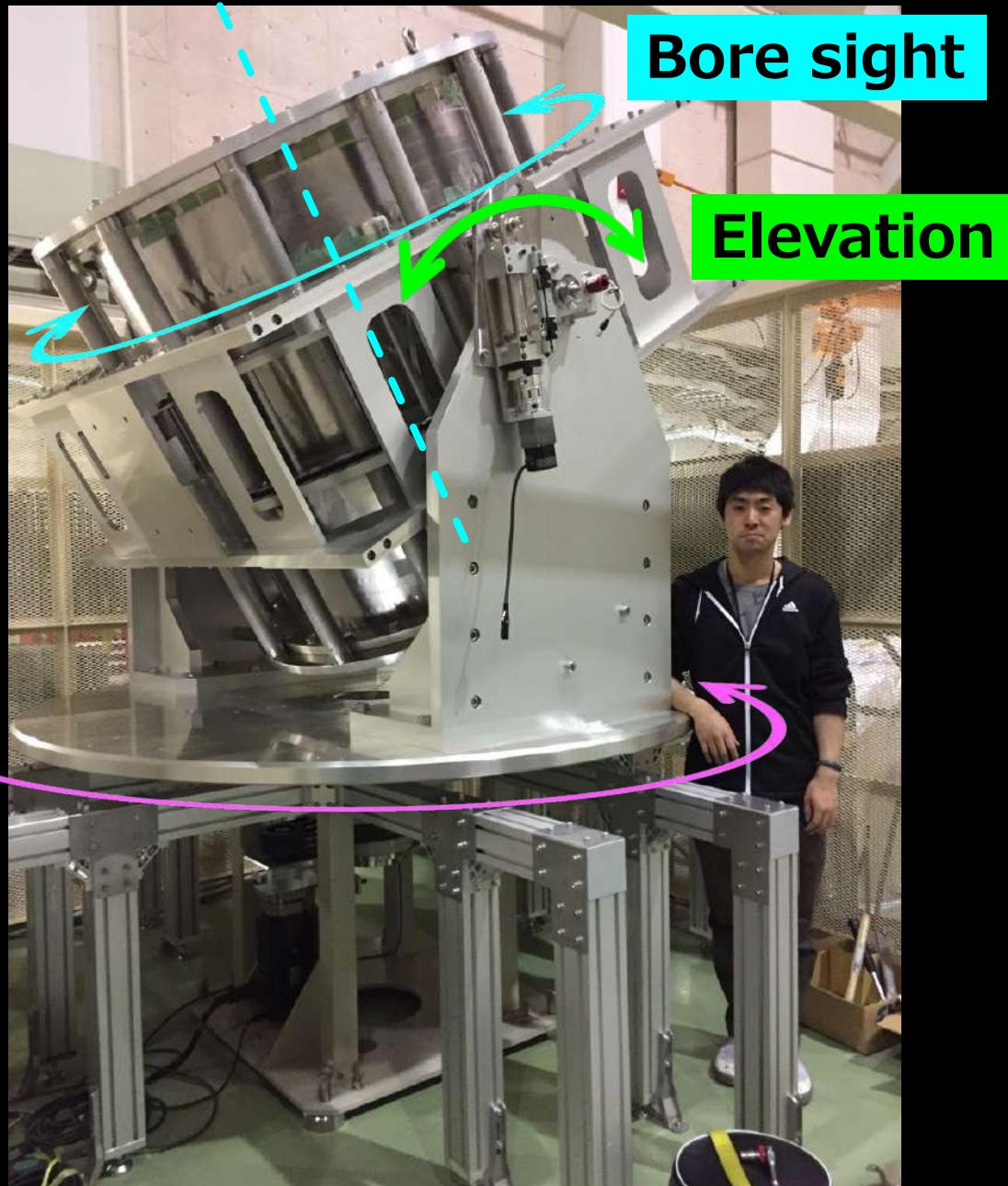


**50K-shell**



Receiver window is here

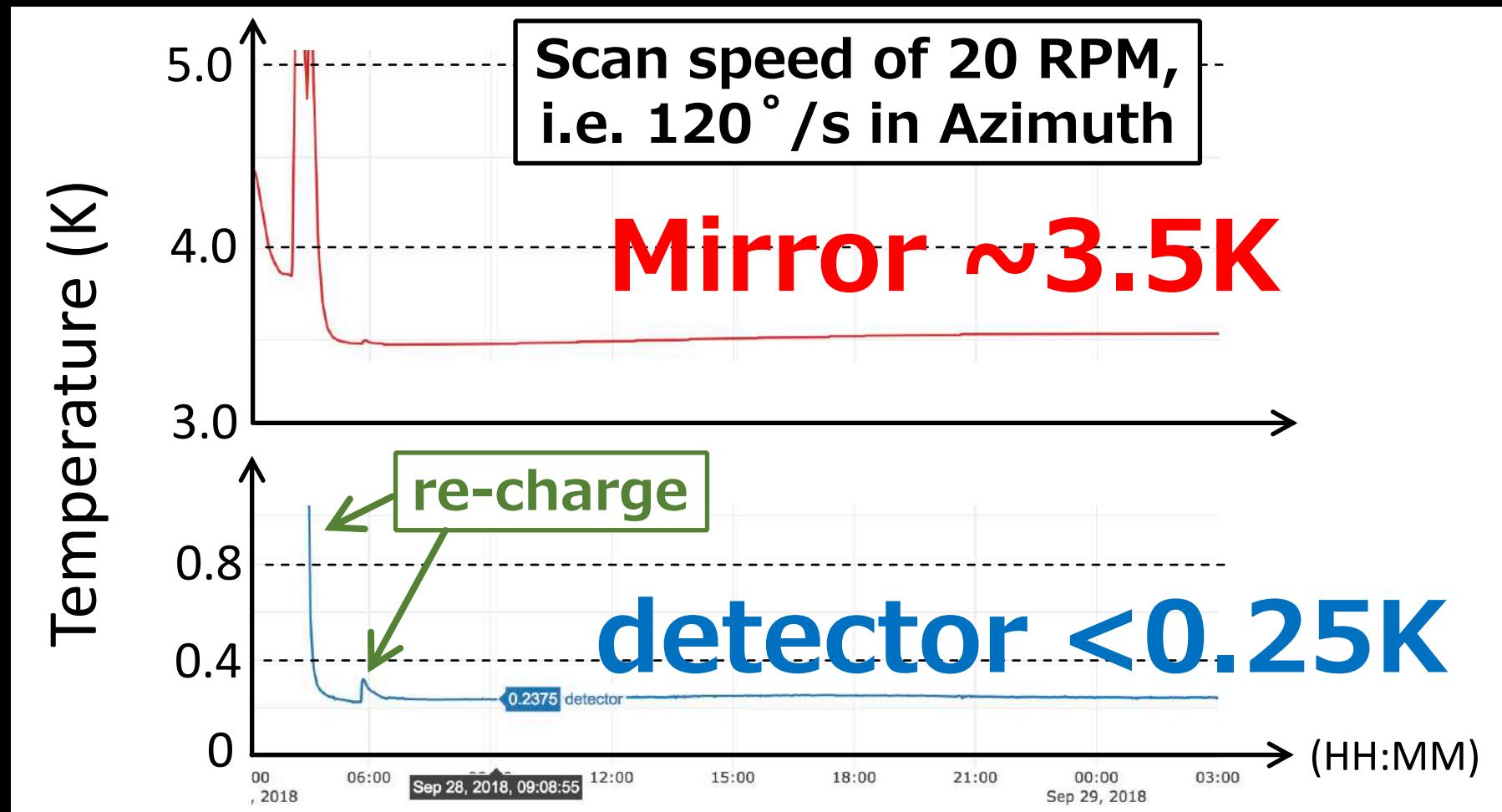
Vac. chamber



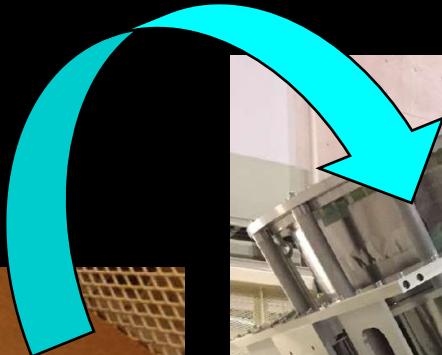
# High-speed Rotation-scan of $120^\circ/\text{s}$



# 24H $T$ trends with HsRs under Eccosorb in $T_{room}$

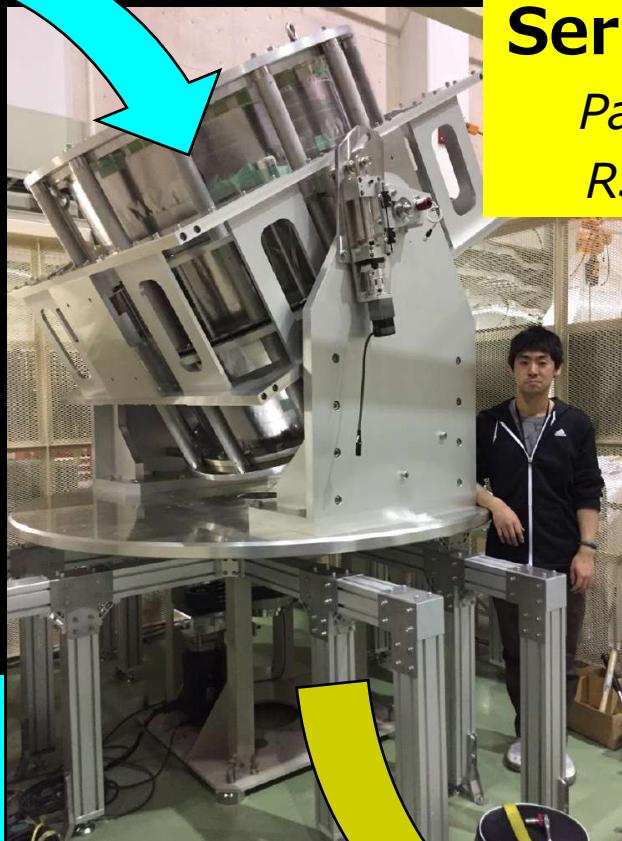


# “Cool” technologies for GB



**RT-MLI for  
IR-blocking**

*Patent: JP6029079  
RSI, 84, 114502 (2013).*



**Series of rotary joints**

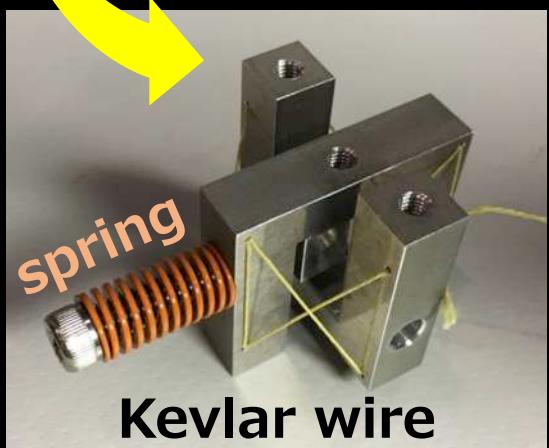
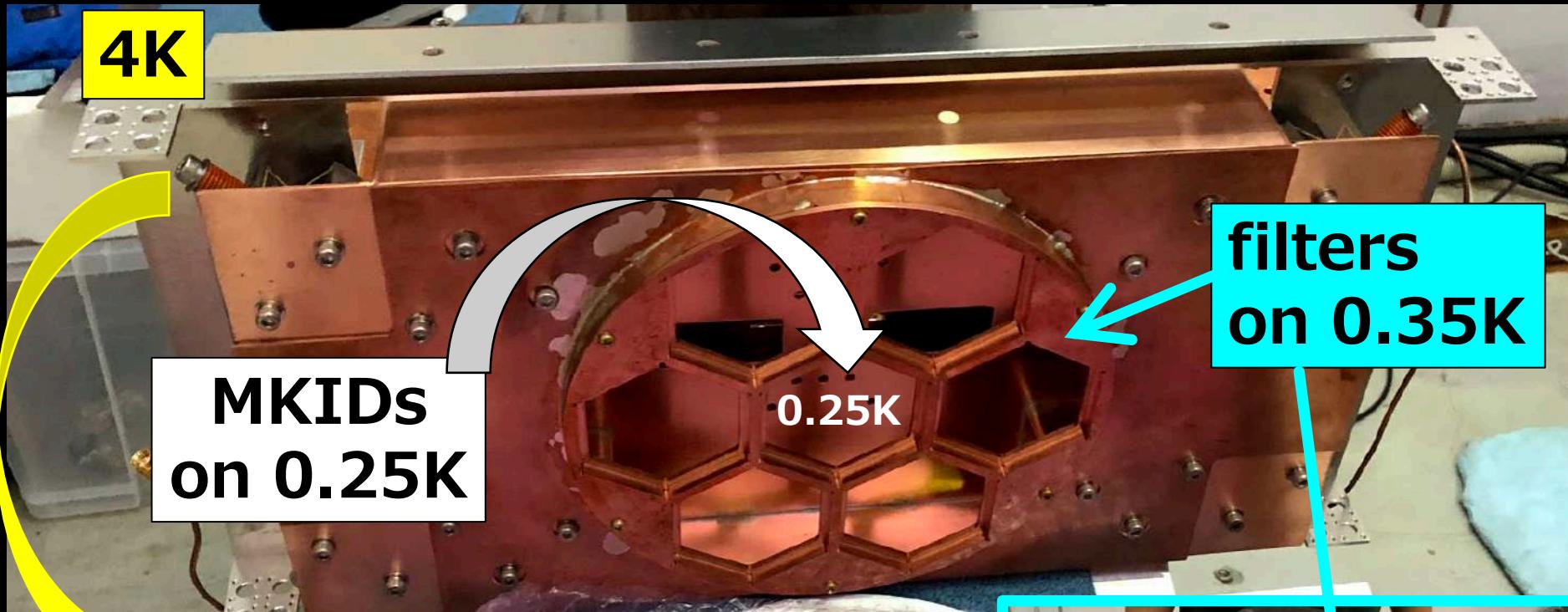
*Patent: US 9,316,418-B2  
RSI, 84, 055116 (2013).*

**He-gas lines for PTC**

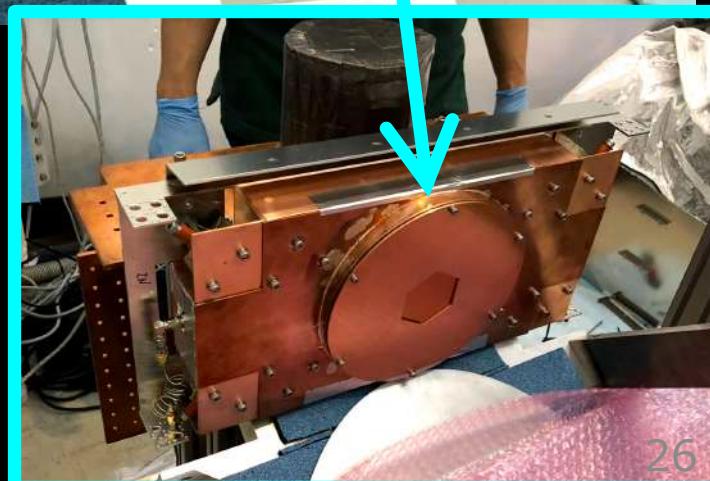


**Electric lines**

# Focal plane set in 0.35K box

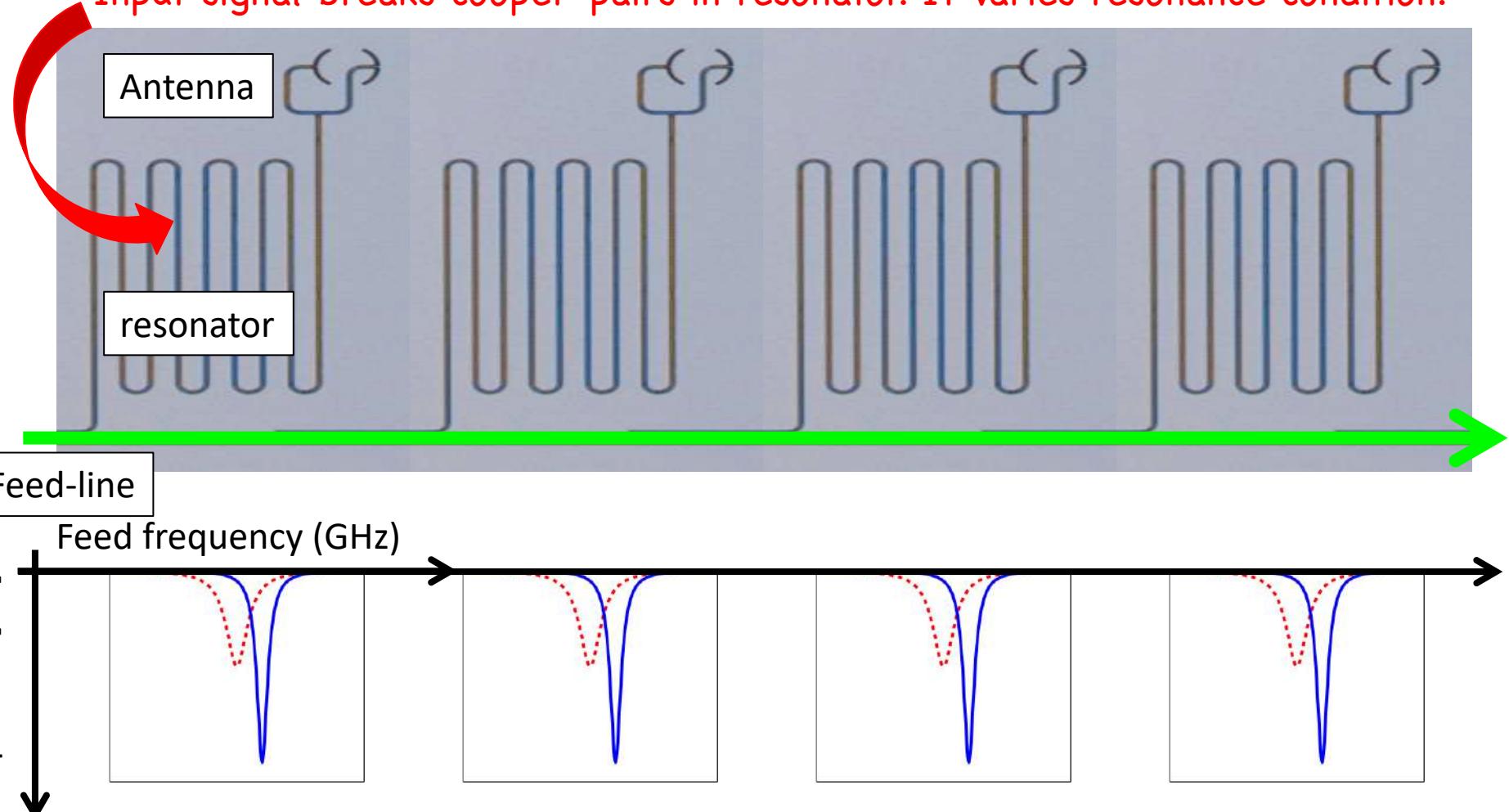


Blacking inside



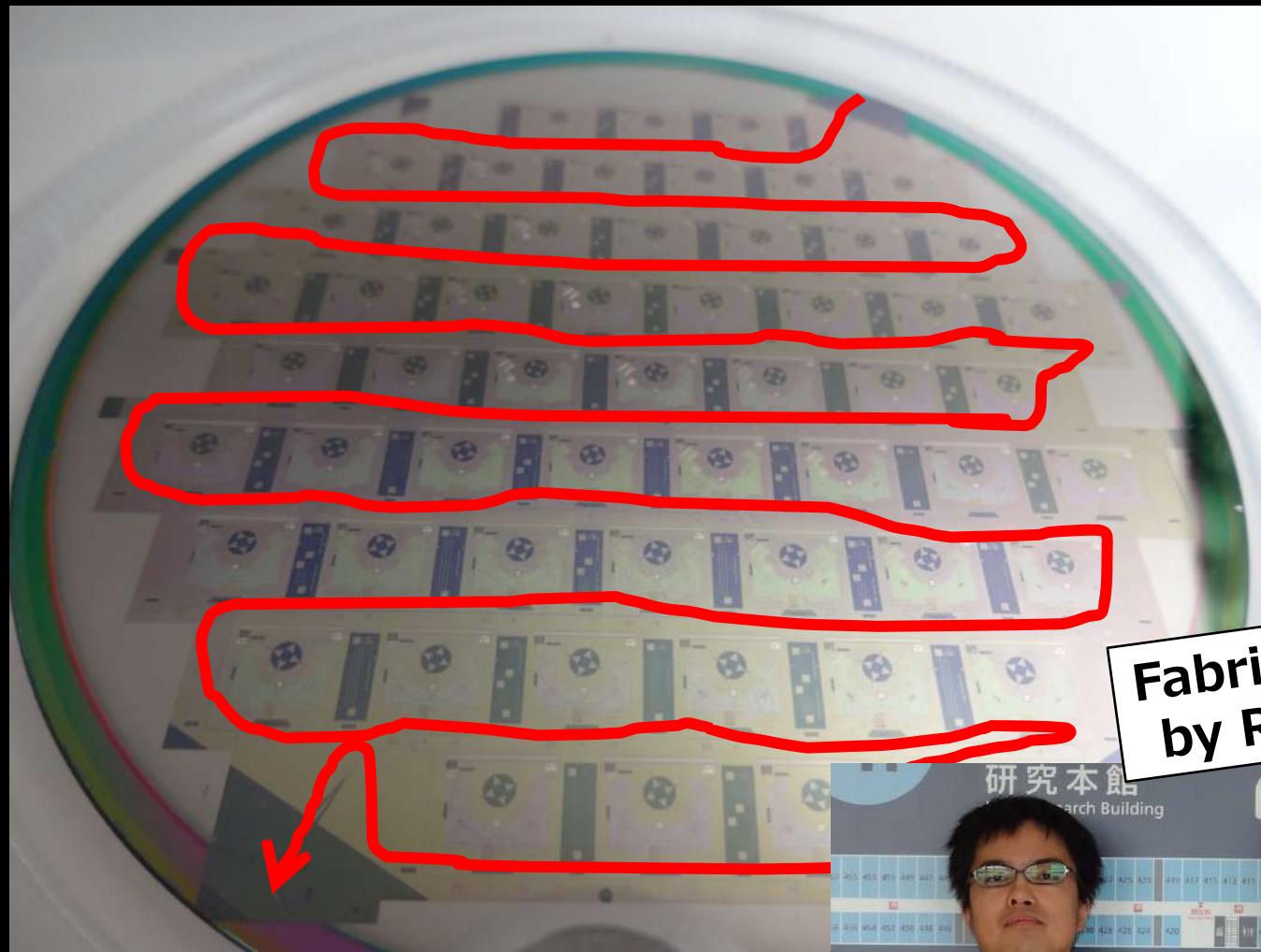
# Benefit of MKID “resonator = detector”

*Input signal breaks cooper-pairs in resonator. It varies resonance condition.*



**Natural frequency domain MUX**

# Single line for readout



\*Now in UTokyo

K. Kiuchi\*

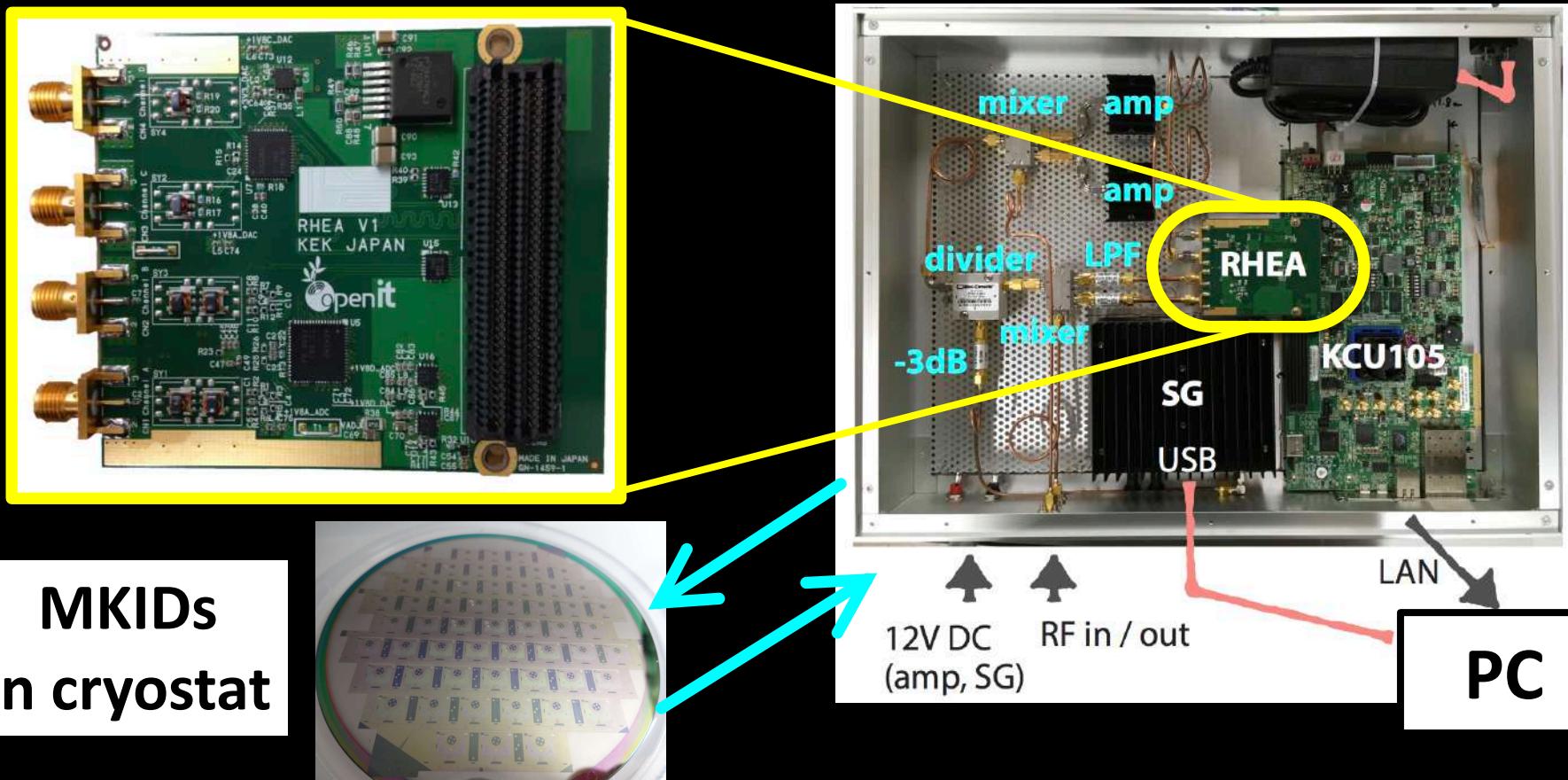
S. Mima

# MKID readout electronics

## DA/AD board “RHEA” :

- 120-MUX in 250 MHz band width
- 1 kSpS high-speed sampling w/o deadtime

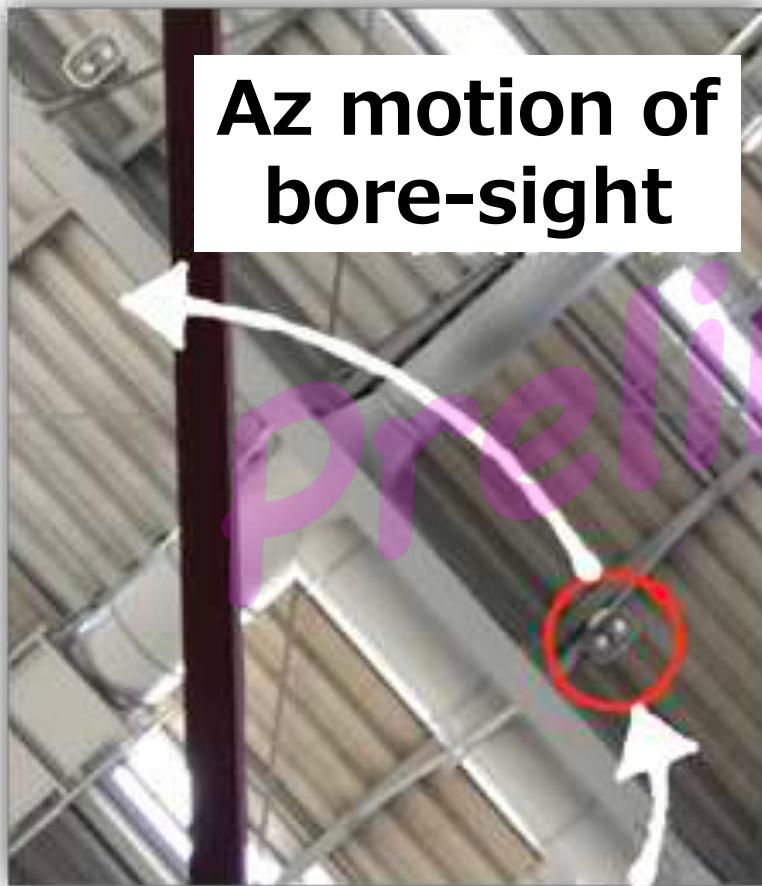
H. Ishitsuka et al, J. Low Temp.  
Phys., 184, Issue 1 (2016)



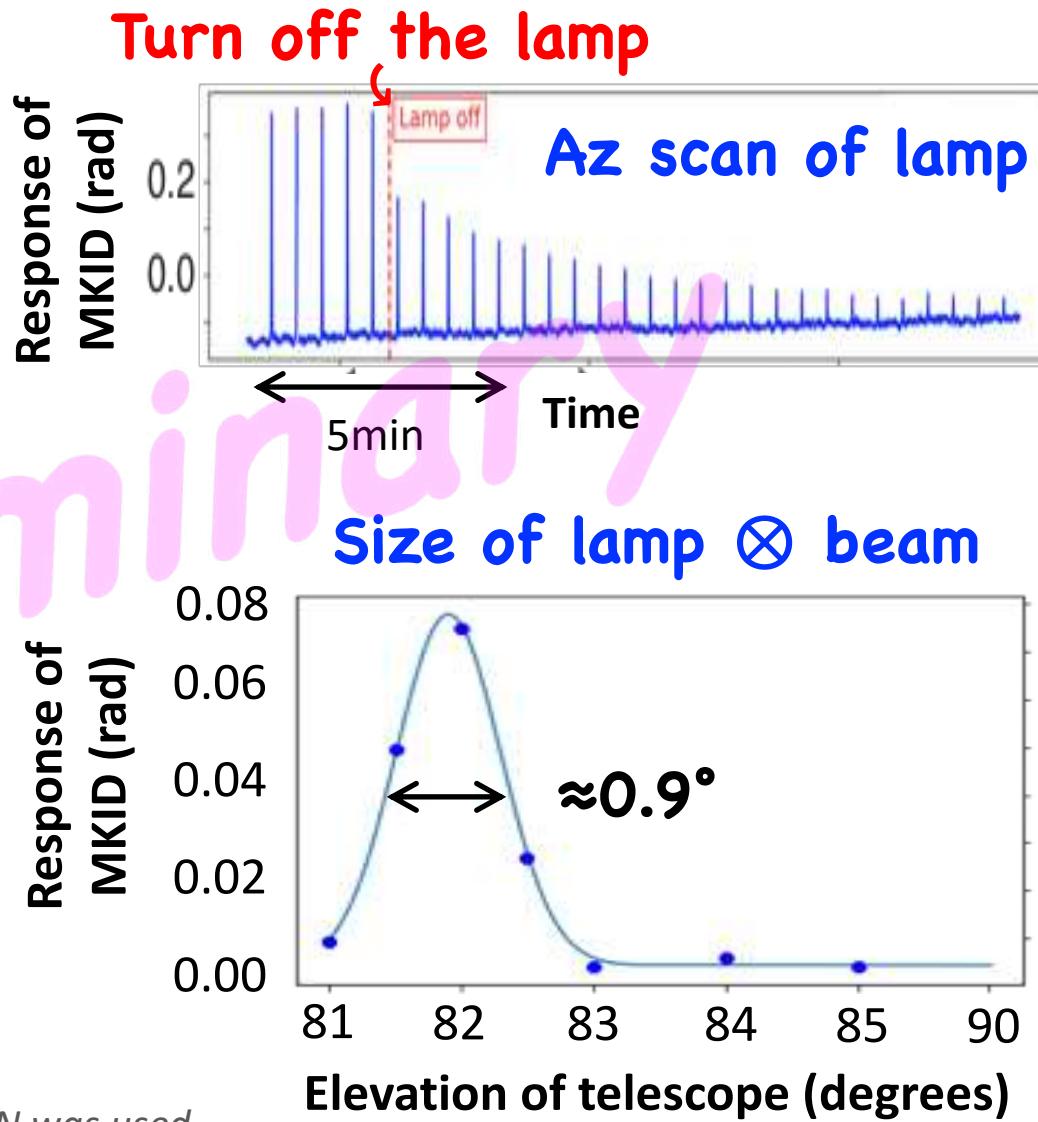
なんちゃってファーストライト

# “First light” in high-bay

Benefit of MKID:  
No saturation in  $T_{\text{room}}$  loads

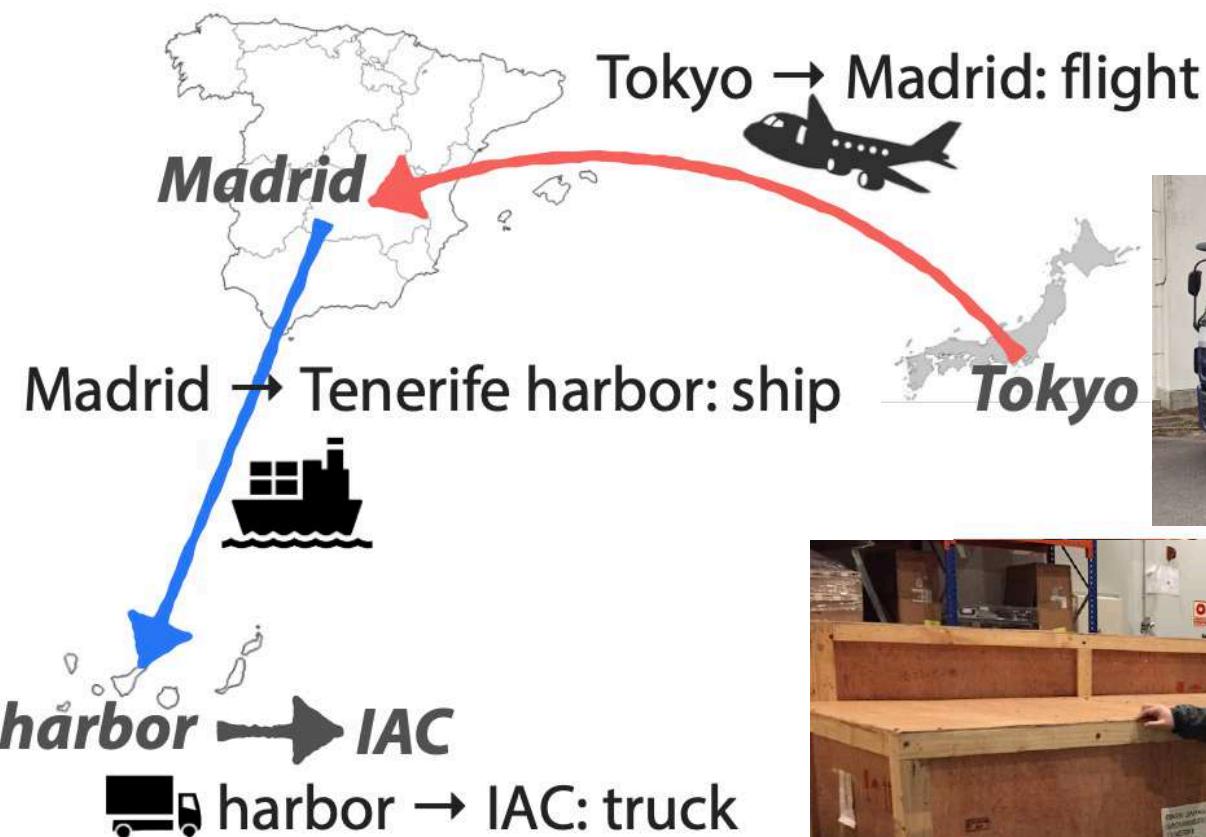


Az motion of bore-sight



Acknowledgement: MKID borrowed from SRON was used.

# Japan → the Canaries



Departure  
Feb. 13, 2019



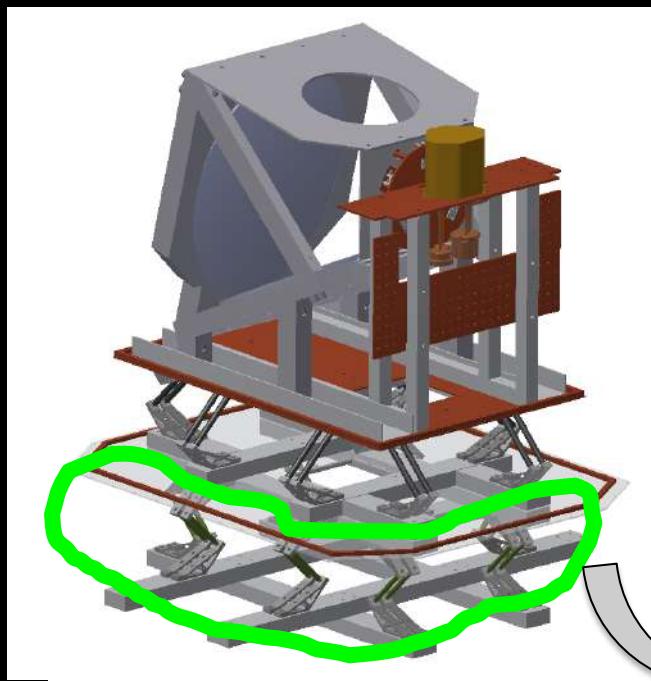
# Let's start "final" cool down test in IAC (lab at 500m alt.)



# Let's start "final" cool down test in IAC (lab at 500m alt.)



# Serious damage ☹ in support structure for 40K – 300K (during the transportation)



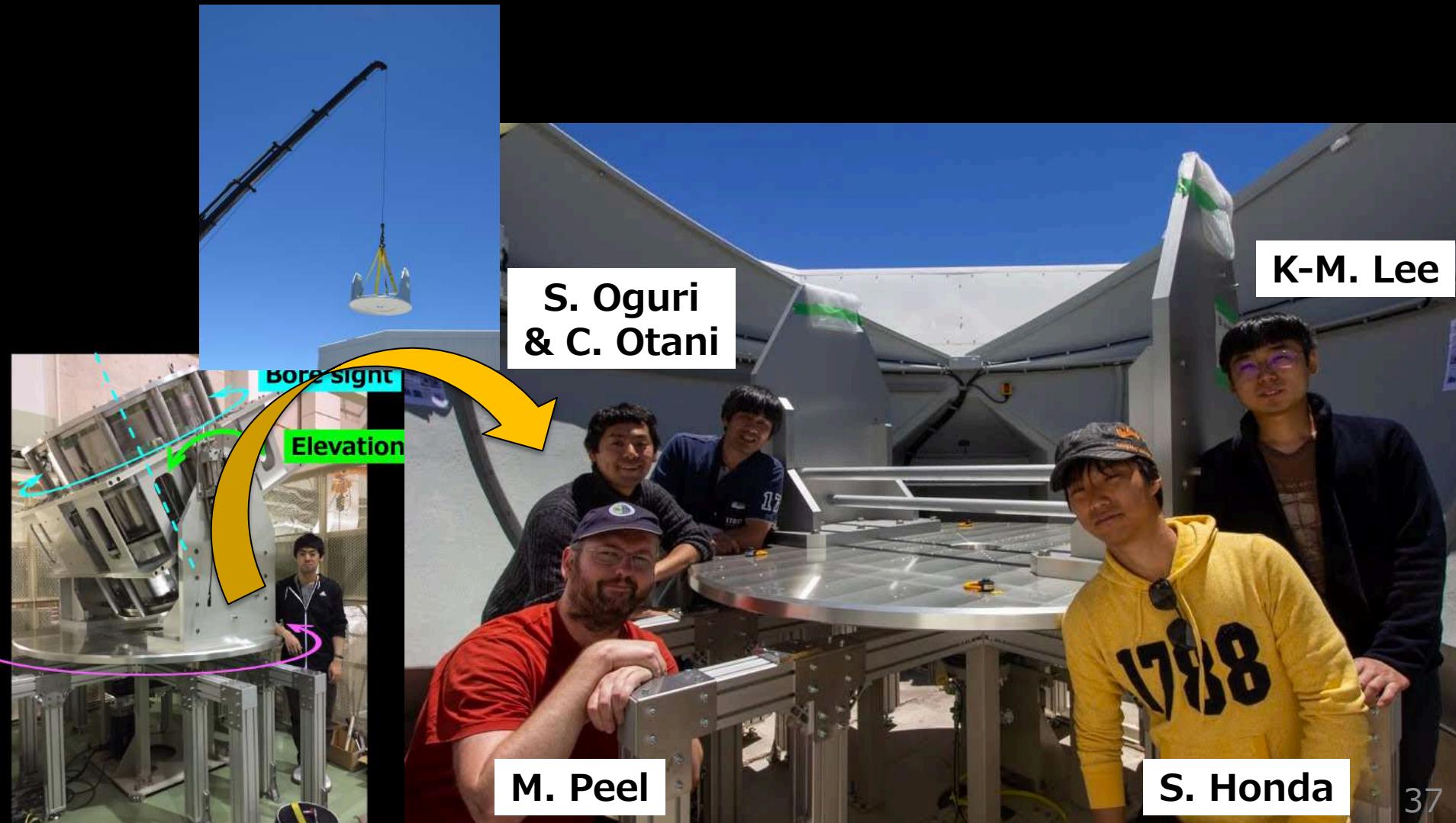
These parts will be replaced in next week

# Others are survived ☺



Cooldown tests will be done in July  
“First light” is planned in Aug.

# Telescope mount was set on the observatory in June 11 !



# Summary of GB

- Low-ell CMB for PGW &  $\tau$
- Unique concepts & techs.  
*High-speed Rotation-scan, MKID, ...*
- “Scan” demonstration in Japan
  - Transportation to the Canaries
  - Deployment is underway
- “First light” soon

# SNS like Summary

User name: GB

# The Canries NOW

