



## WG3 - Summary

Convenors: Guillaume Rosaz Teng Tan Tobi Junginger

### S1: Nb/Cu thin films

Development of coating system for the Wide Fabio Avino Open Waveguide (WOW) Crab Cavity (10'+8')

Progresses on ECR plasma Anne-Marie Valente-Feliciano deposited Nb thin films (10'+8')

14:36 - 14:54

15:12 - 15:30

Legnaro Thick films (10'+8') Cristian Pira

500/1-001 - Main Auditorium, CERN

Test results of re-built LHC spare M. Franck Peauger cavities (10'+8')

ARIES WP15 progresses (10'+8') Oliver Julius Kugeler

500/1-001 - Main Auditorium, CERN

How to coat complex shapes?

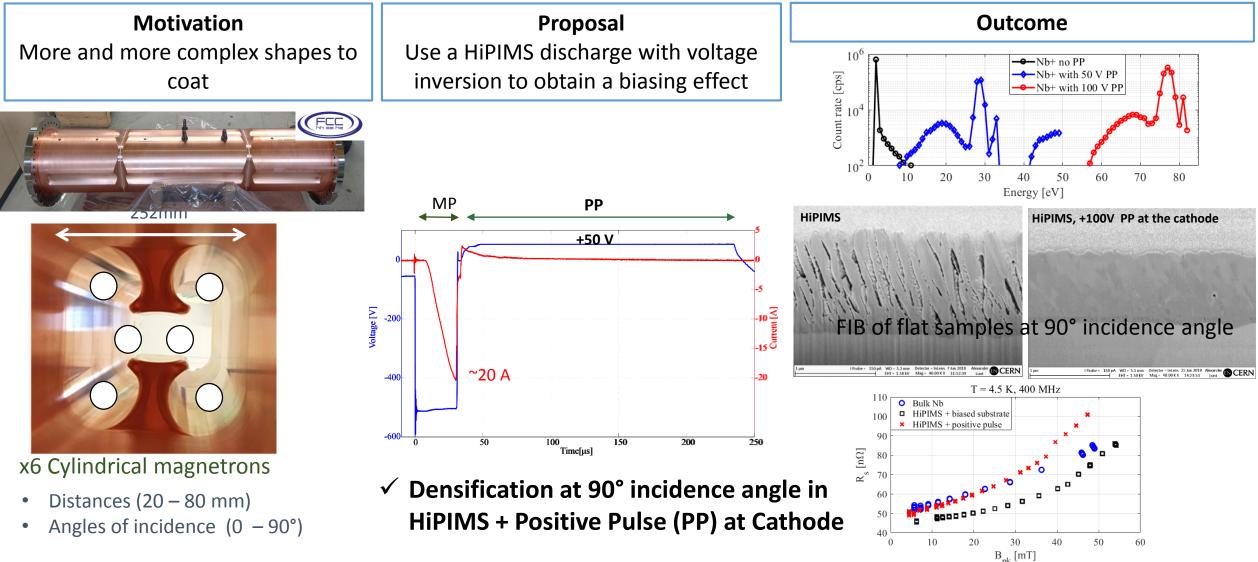
Can thin films compete with bulk Nb?

How to mitigate Q-slope?

How does thin film structure impacts SRF performance?

#### Development of coating system for the Wide Open Waveguide (WOW) crab cavity

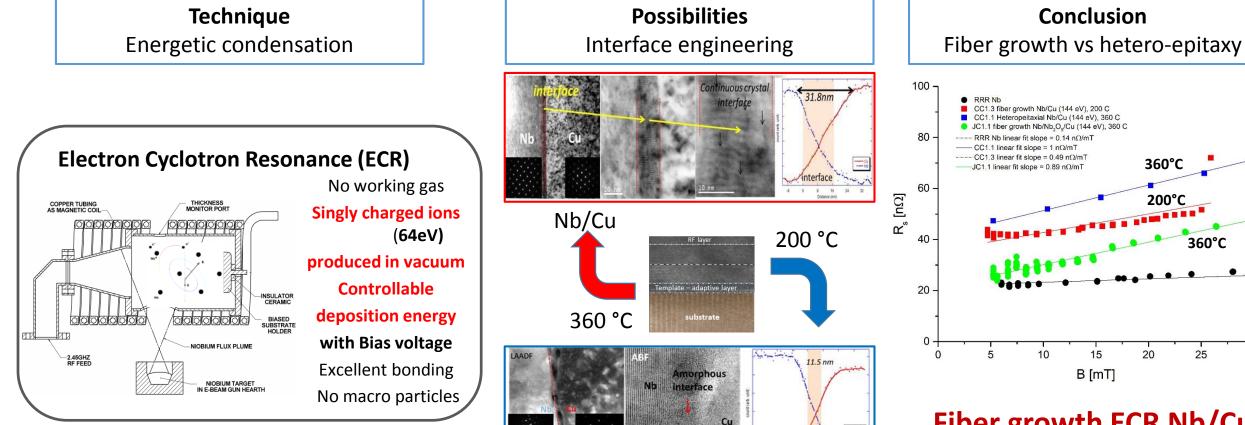
F. Avino *et al* 



CERN

#### Progress with Nb/Cu film engineering with energetic Jefferson Lab condensation

A.-M. Valente-Feliciano et al



Fiber growth ECR Nb/Cu show better mitigation trend of the Q-slope

360°C

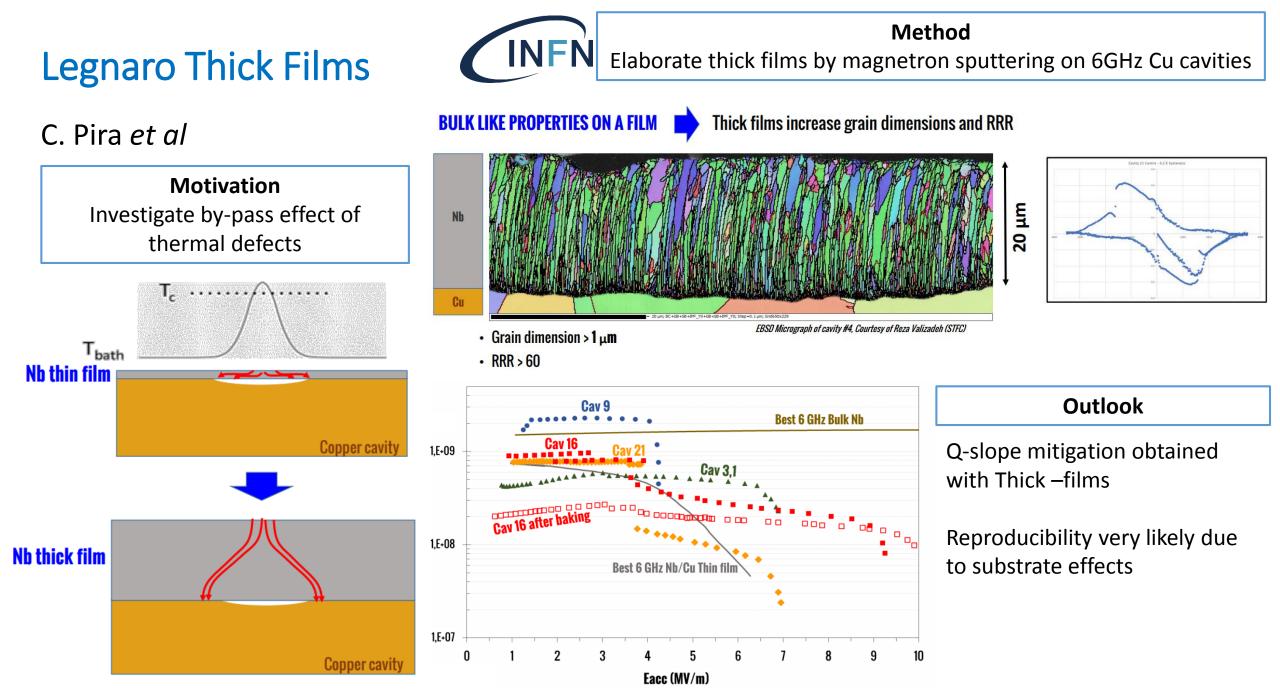
200°C

20

360°C

25

30

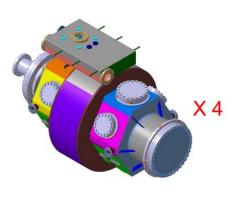


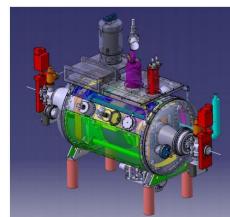
#### Test results of re-built LHC spare cavities

#### F. Peauger et al

#### Spare cavities program

- Only one spare dressed cavity and one spare cryomodule available
- New project started in 2015 to re-build and qualify <u>four cavities</u> and <u>one quarter cryomodule\*</u>





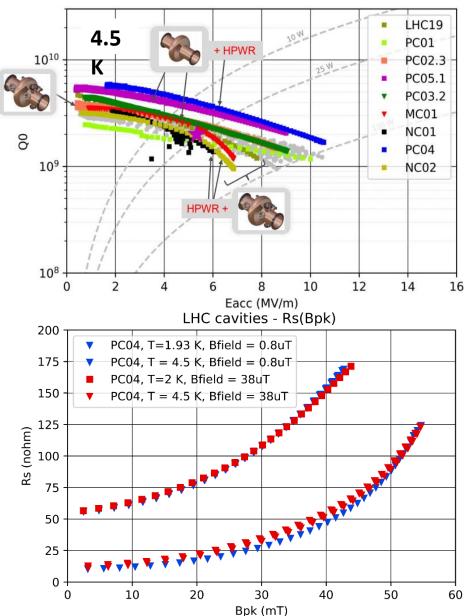
Simplified and "real" cavities

Simplified systematically perform better

Investigations on going to indentify the root-cause

Best simplified cavity tested in actively shielded cryostat

Very good performance Rres = 12nOmhs Very low sensitivity to trapped flux





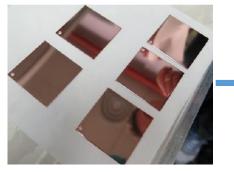


#### **ARIES WP15 Thin Films progresses**

O. Kugeler et al

From samples to QPR

Surface preparation optimization



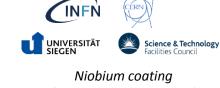


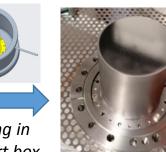


Surf. preparation, polishing (INFN LNL)









Shipping in transport box

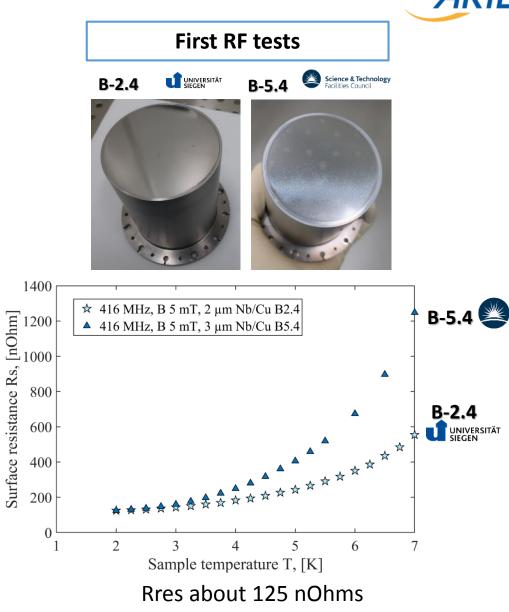


Shipping in transport box



RF testing (HZB)





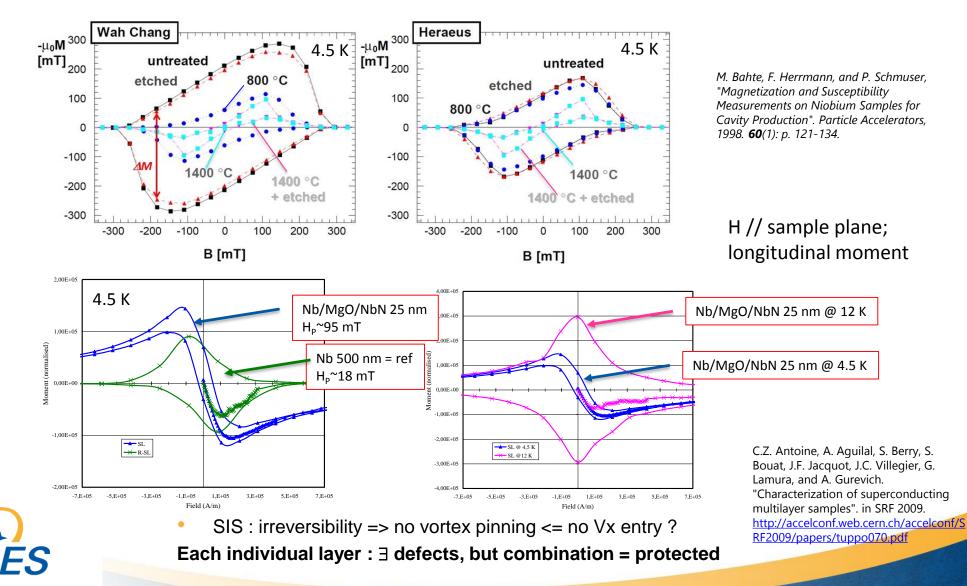
**Optimization on-going** 

### S2: Alternative structure/materials

Recent SIS layer results (10'+8')	Claire Antoine 🥝
500/1-001 - Main Auditorium, CERN	16:00 - 16:18
Overview Multilayers RF results (10'+8')	Sebastian Keckert 🥝
Nb3Sn/V3Si results (10'+8')	Stephanie Fernandez 🖉
500/1-001 - Main Auditorium, CERN	16:36 - 16:54
Two-stage coating of MgB2 system	(10'+8') Hiroshi Sakai 🥝
500/1-001 - Main Auditorium, CERN	16:54 - 17:12
CVD Thick Nb film and cavity coatin (10'+8')	ng Zeming Sun 🖉
500/1-001 - Main Auditorium, CERN	17:12 - 17:30

- SIS structures
- Sputtering A15 thin film
- MgB2 come back to life in LANL and KEK.
- CVD thick Nb films: new surface morphology.

#### Bulk Niobium vs thin films

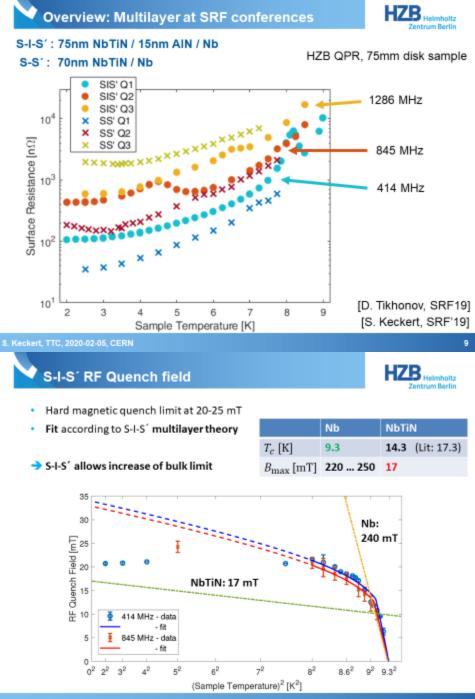


• Bulk : irreversibility (no vortex pinning) only if well recrystallized

### S. Keckert RF measurement of SIS

- SRF characterization of multilayer structures with sample test cavities
- ➔ So far: Mostly surface resistance data at high frequency and low field
- ➔ Consistently: Lower R<sub>s</sub> than for Nb at higher temperature So far: Severe limitations by residual resistance
- Penetration depth measurement agrees with S-I-S´ multilayer theory
- ➔ First RF critical field measurements of S-I-S´ and S-S´ structures show low-field quenches

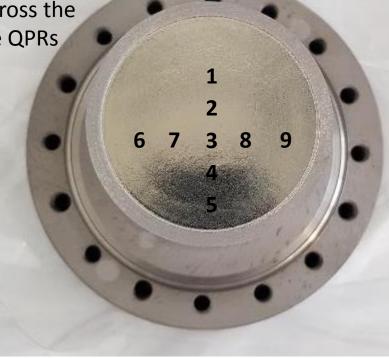
 $\rightarrow$  We need more data on the RF quench field !



### S. Fernandez Nb<sub>3</sub>Sn and V<sub>3</sub>Si from sputtering



Homogeneous composition across the surface of the QPRs



75 mm diameter Atomic composition (%) measured in nine points along the disk

	Nb	Sn
1	77.9 %	22.1 %
2	78.3 %	21.7 %
3	78.5 %	21.5 %
4	77.9 %	22.1 %
5	78.2 %	21.8 %
6	78.8 %	21.2 %
7	78.1 %	21.9 %
8	78.3 %	21.7 %
9	78.7 %	21.3 %

 Film thickness may be the key

 Nb<sub>3</sub>Sn

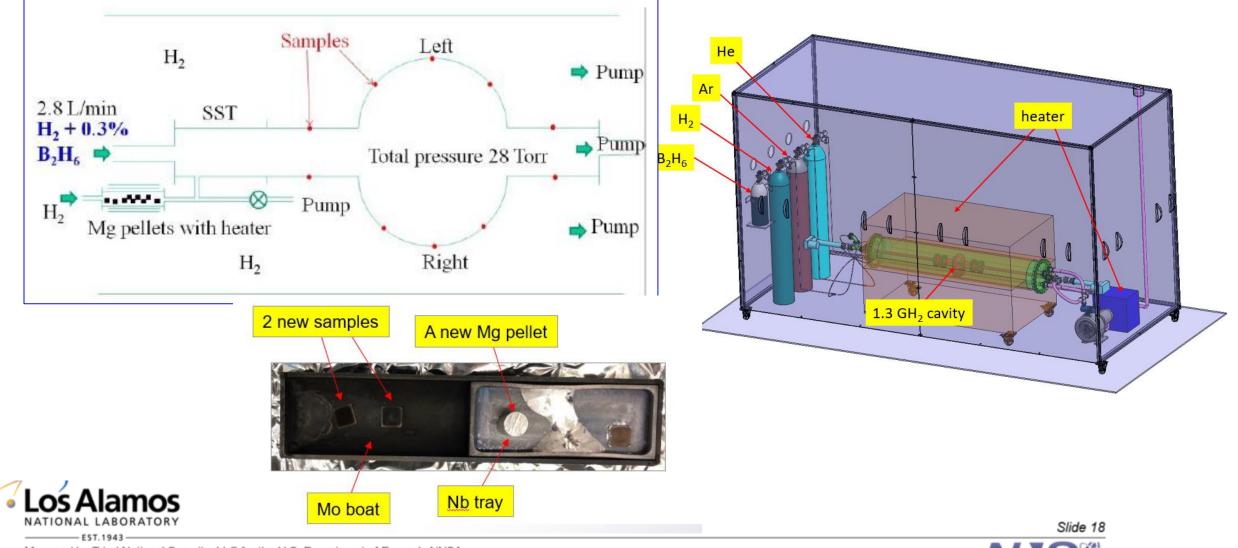
 Ta

 Cu

t vs London penetration depth

Sample	Recipe	Critical temperature (T <sub>c</sub> )	Surface resistance (R <sub>s</sub> )
QPR1	Nb <sub>3</sub> Sn/Nb	14.5 K	1000 nΩ
QPR2	Nb <sub>3</sub> Sn/Ta	11.5 K	25 nΩ
QPR3	Nb <sub>3</sub> Sn/Ta	15 K	2000 nΩ
QPR4	Nb <sub>3</sub> Sn (thick)/Ta	16 K	1000 nΩ
			11

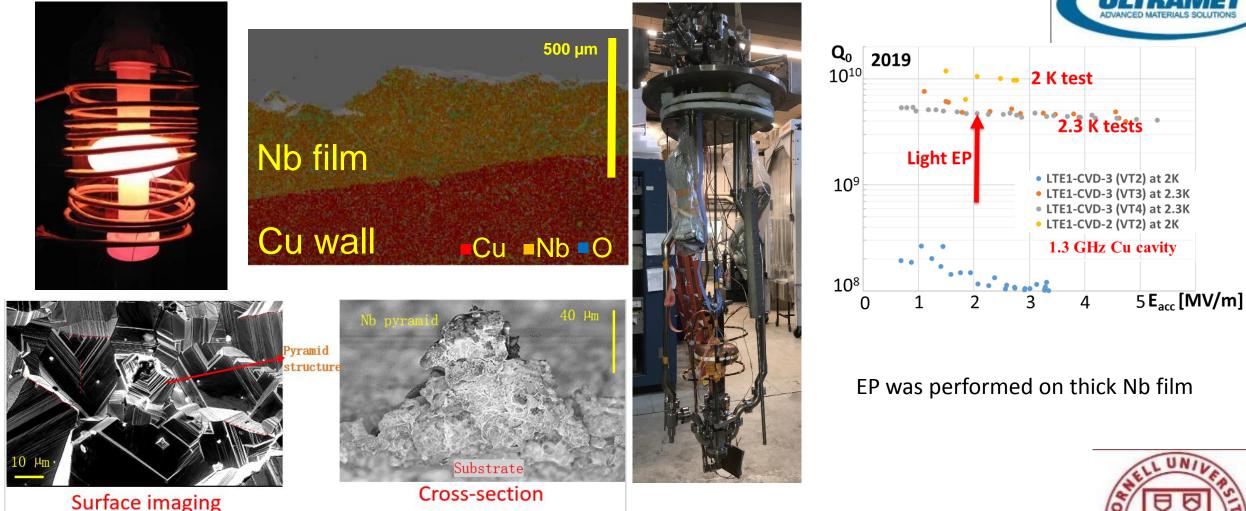
### H. Sakai: MgB<sub>2</sub> revived at KEK and LANL

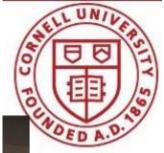


Managed by Triad National Security, LLC for the U.S. Department of Energy's NNSA



### Z. Sun CVD thick Nb/Cu: from deposition to RF test



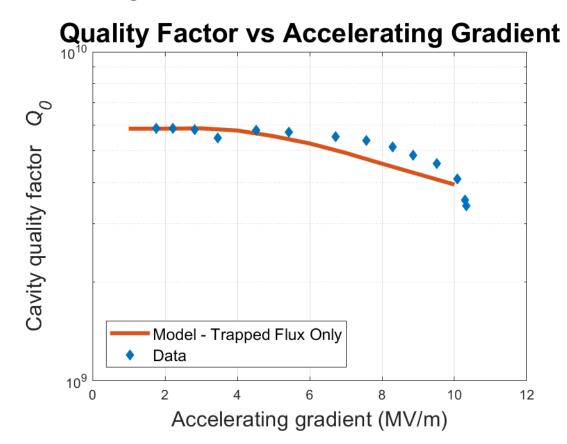


### S3: Progress on Nb<sub>3</sub>Sn towards accelerator



- Cryo-cooler driven, other deposition methods
- Multi-cell at Fermilab: two mode method for quench mechanism.
- New MP quench mechanism for Nb3Sn cavity. Hope for >24 MV/m
- Bronze route: another possibility
- Voices from outside TFSRF: what does the machine need

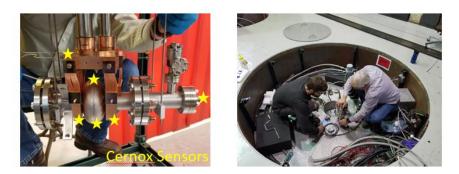
### R. Porter: Nb<sub>3</sub>Sn from Cornell



10 MV/m stable operation with cryo-cooler

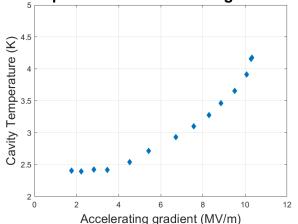
Some tries with other deposition techniques





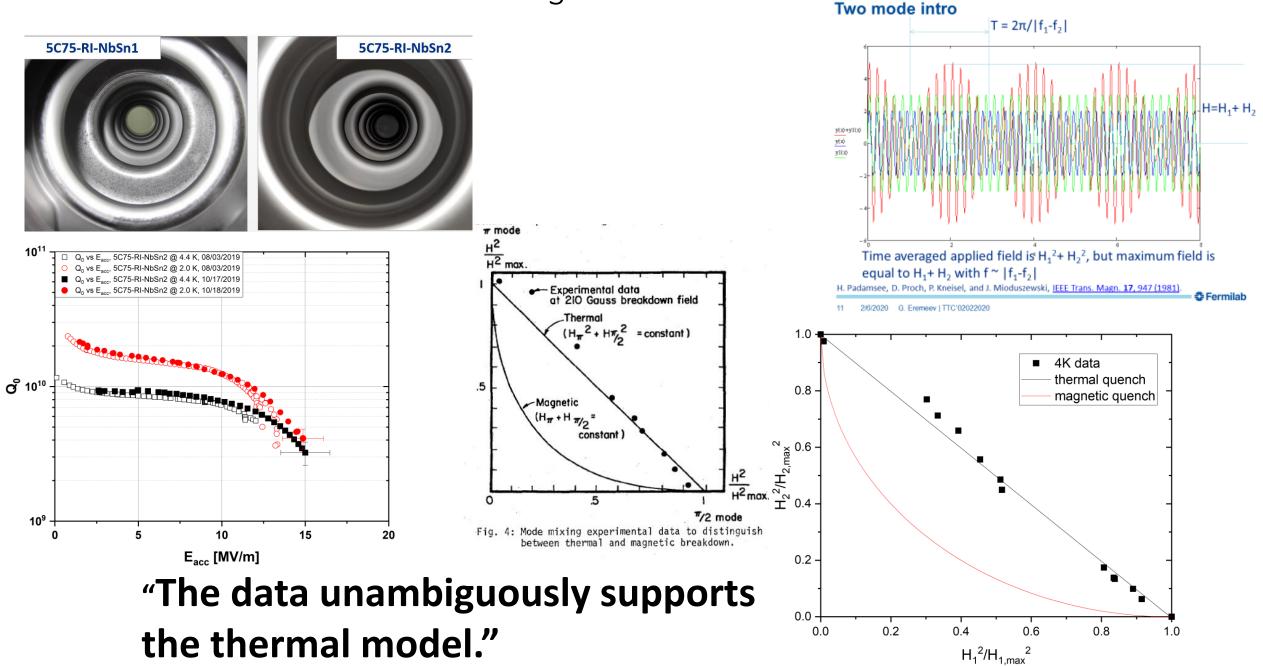


Temperature vs Accelerating Gradient

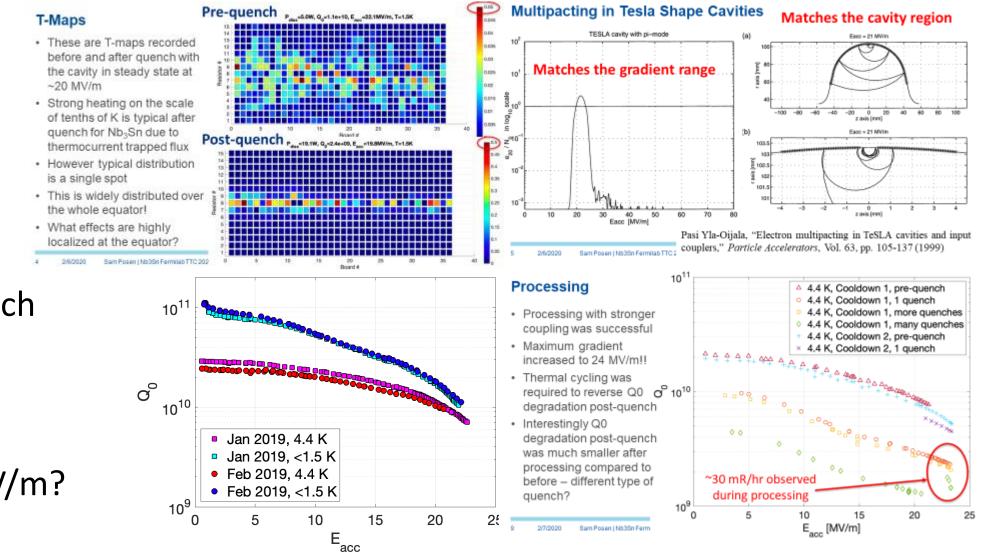


### G. Eremeev: Multi-cell Nb<sub>3</sub>Sn from FNAL

🛟 Fermilab



### S. Posen: New Quench Mechanism for Nb<sub>3</sub>Sn



MP induced quench Improved Q after processing

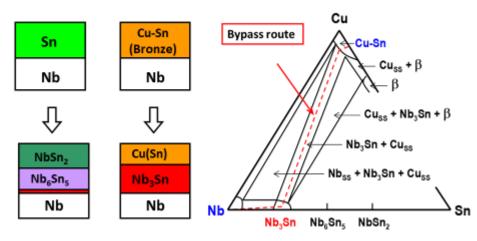
Going over 24 MV/m?

🛛 🛟 Fermilab

### A. Kikuchi: Nb<sub>3</sub>Sn thin film/layer via bronze route

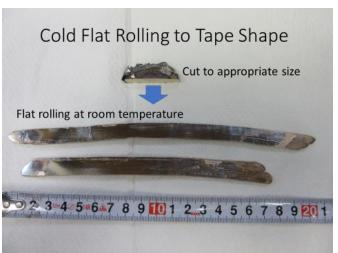
#### Discovered Effect of Cu in 1967

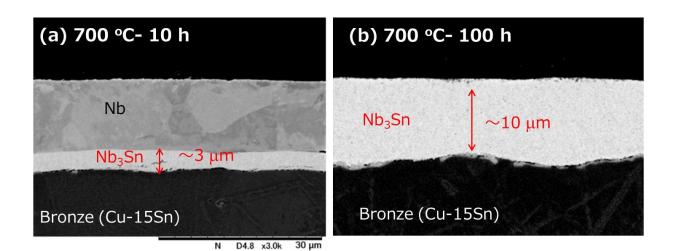
Bypass Route for Direct Forming of  $Nb_3Sn$  Phase at low temperature diffusion reaction. (*Tachikawa Method*)



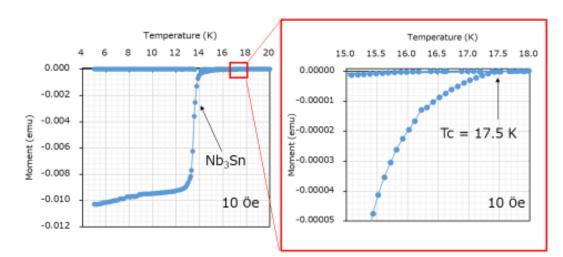
#### Method for Nb<sub>3</sub>Sn wire

**Composite material** 





#### Full Reacted at 700 °C for 100 h





### R. Calaga: Nb<sub>3</sub>Sn: towards machine

- Accelerator compatibility requires dealing with real world mess. So Nb3Sn needs an "industrialization" break-through
- Its applications could/should be decoupled from bulk-Nb (starting with the substrate)
- Think special applications
  - Higher frequencies (s/c-band), higher quench fields, HOM couplers
- Think cheap fabrication techniques
  - Machining/3D printing of cheaper-Nb, Cu, Al....
- One day in the far future, we might even dream of in-situ coating of Nb3Sn inside existing SC-accelerators

Voices from accelerator people, instead of thin film people themselves

#### SRF Accelerator "Landscape"

	Some Examples	Frequencies [MHz]	Voltage [MV/m]	?
Circular colliders & storage rings	LHC, FCC, KEKB, CESR, Light sources	350-500, 800	5 - 15	
Linear Colliders & FELs	XFEL, LCLS2, ILC	1300, 3900	20-35	
High Intensity	SNS, ESS, PIPII	650-800	7 – 20	
Nuclear Physics	FRIB, HIE-ISOLDE, ATLAS	~100-400	6 - 10, 20	
ERLs & RLs	CEBAF, Test facilities, e-cooling	700, 1300	15-20	
Special Applications	Crab/Deflecting cavities, Medical	400, 3900	~25	
Compact/cheap e-accelerators	Studies	600-800, 1300?	~10	

\*\*Note: By no mean comprehensive but some qualitative examples



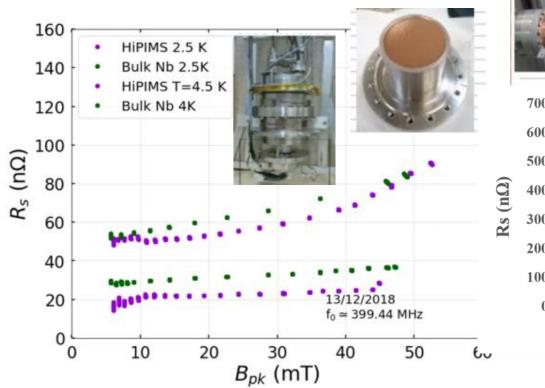
# Session 4: How to predict SRF performance without building an entire cavity

HIPIMS from QPR to 1.3GHz cavities (10'+8')	Lorena Vega Cid 🥝
500/1-001 - Main Auditorium, CERN	11:00 - 11:18
Predicting SRF Performance using muSR and betaNMR (10'+8')	Edward Thoeng 🧭
500/1-001 - Main Auditorium, CERN	11:18 - 11:36
RF local magnetometry (10'+8')	Steven Anlage 🧭
500/1-001 - Main Auditorium, CERN	11:36 - 11:54
Predicting SRF performance PCT (10'+8')	Thomas Proslier 🥝
500/1-001 - Main Auditorium, CERN	11:54 - 12:12
Local magnetometry (10'+8')	Daniel Turner 🥝
500/1-001 - Main Auditorium, CERN	12:12 - 12:30

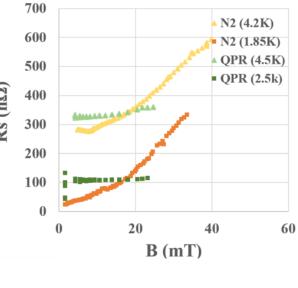
Quadrupole Resonator (**QPR**), Muon Spin Roatation and Relaxation (**muSR**), beta detected nuclear magnetic resonance (**beta-NMR**), RF local magnetometry, point contact tunneling (**PCT**), vibrating sample magnetometry (**VSM**), local DC magnetometry with 3<sup>rd</sup> harmonics, field penetration measurements

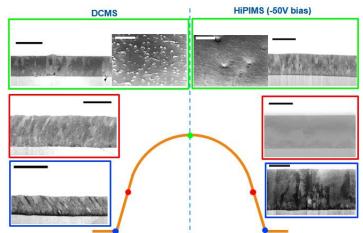
### Lorena Vega – HIPIMS from QPR to 1.3 GHz cavities











HiPIMS technique allows to achieve denser layer in all the orientations.

This technique may still need specific optimization for the cavity geometry

HiPIMS sample showing comparable performance to bulk niobium <u>in</u> terms of Q-slope

✓ When applying the same coating technique to 1.3 GHz cavities the RF performance is not as good as expected.

✓ Efforts should be put on:

Specific optimization of the coating technique for the cavity geometry. Have good substrates (seamless cavities with thermal mapping system).

#### Edward Thoeng – Predicting SRF Performance using muSR and betaNMR

#### LE-musR: 120°C bake study with LE-muSR @ PSI[3] limited to 30 mT (low-field)

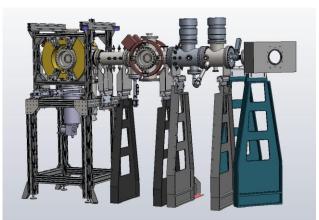
#### beta-NMR Probe

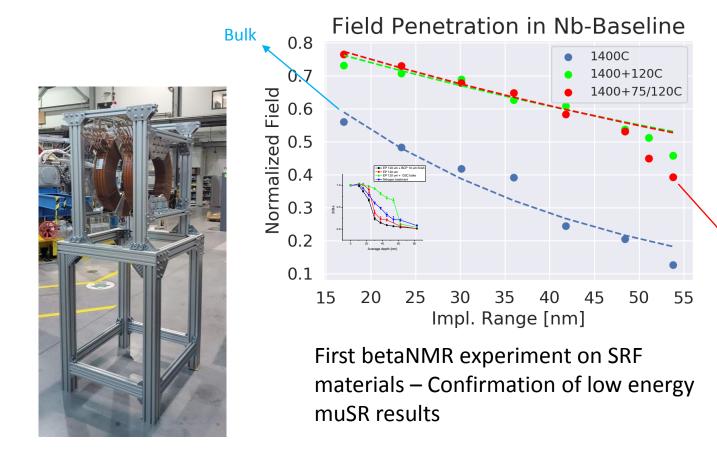
Implanted low energy radioactive Li-8 beam Field || sample: currently 24 mT – 200 mT after upgrade

#### Comparison with LE-muSR

Low-energy  $\rightarrow$  Surface & interface sensitive nm-scale Heavier ions (vs muons)  $\rightarrow$  Larger magnetic rigidity at high-field

HV deceleration [existing] → **Depth resolved** studies









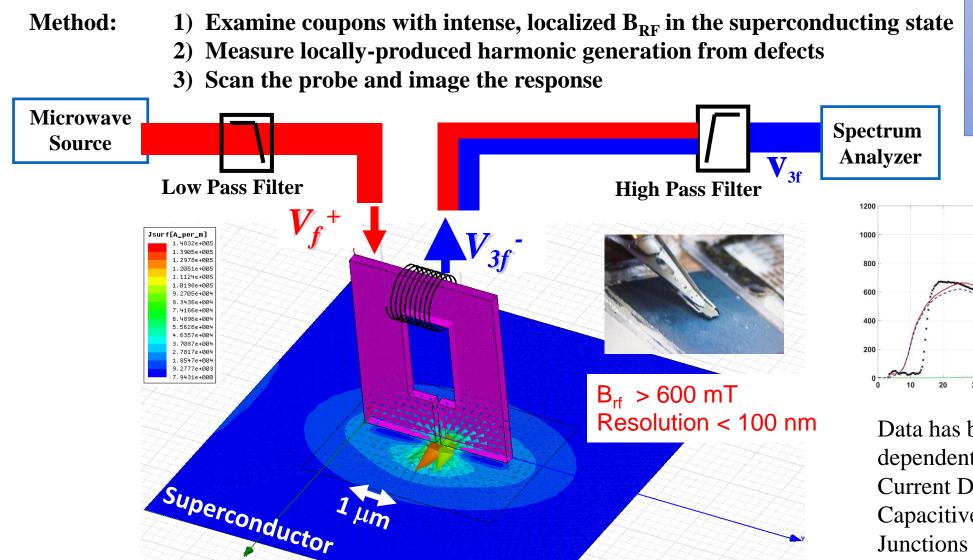
**Bi-layer** 

5

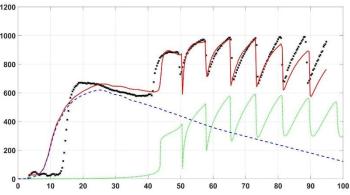
#### Steven Anlage – RF Local Magnetometry



**Objective: Identify microscopic defects that cause breakdown of SRF cavities** 



Ultimate goal is to map the surface resistance compare to theoretical models and relate to surface structure

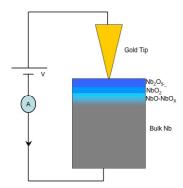


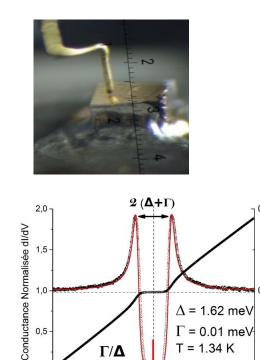
Data has been fitted with a time dependent GL model and Current Driven Resistively and Capacitively Shunted Josephson Junctions (RCSJ) model

#### **Thomas Proslier – Predicting Srf Performance with PCT**

ant I [µA]

Cou





2 4

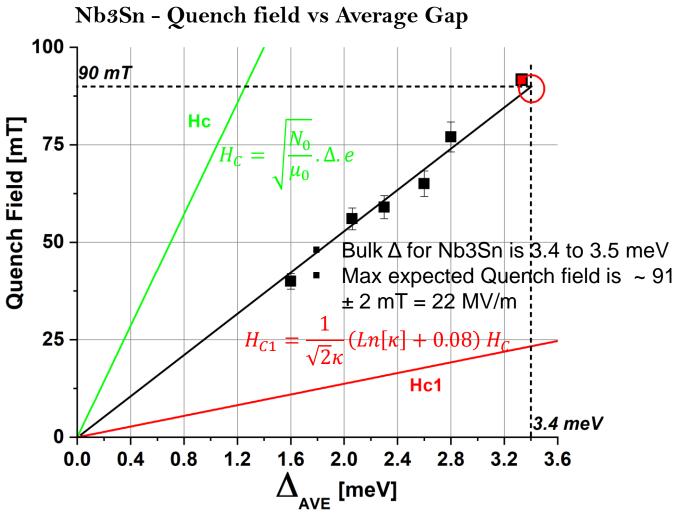
Tension [mV

6

 Measure the fundamental superconducting parameters: Δ, T<sub>C</sub>, H<sub>C2</sub>

0,0 -10 -8 -6 -4 -2

- Measure non-ideal signature: Γ.
- Shape of the DOS give clues to microscopic origins: *Proximity effect, magnetic impurties, deleterious phases.*
- Direct correlation to SRF cavity performances.
- Cartography.

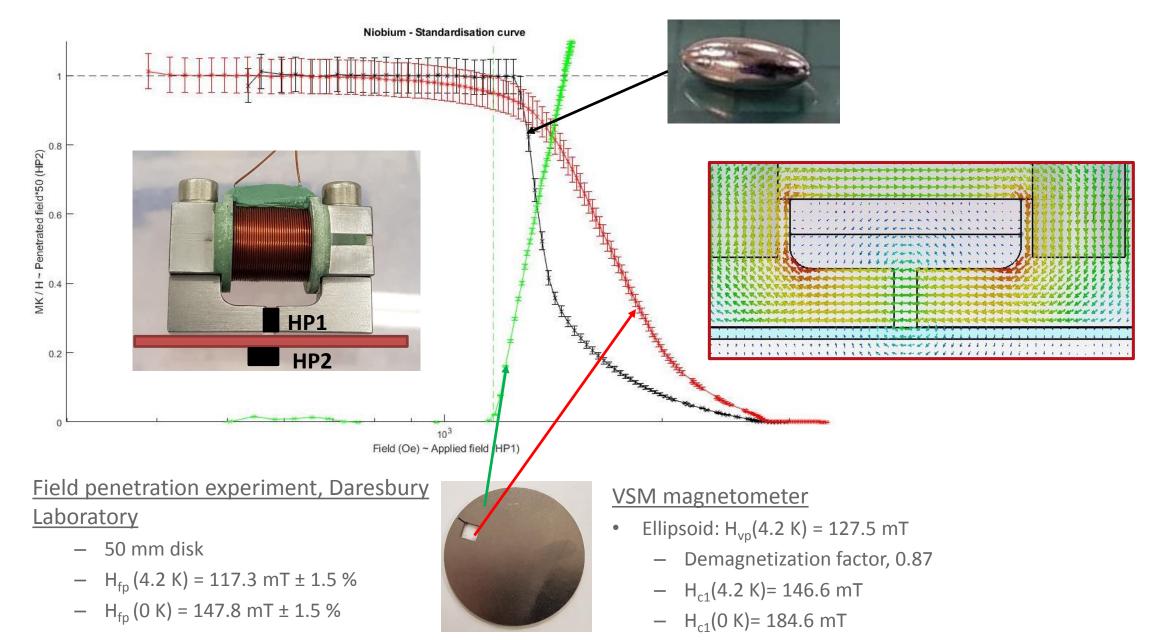


Linear dependence of Emax on the average surface gap H<sub>SH</sub>~3.5 times the H<sub>C1</sub> but why this gap dependence? Roughness, effective penetration depth?



### D. Turner – Local Magnetometry





# THANKS TO ALL CONTRIBUTORS FOR THE HIGH QUALITY TALKS