

Antideuteron and antihelium cross sections for cosmic-ray studies

Diego Gomez

Department of Physics & Astronomy

University of Hawaii at Manoa

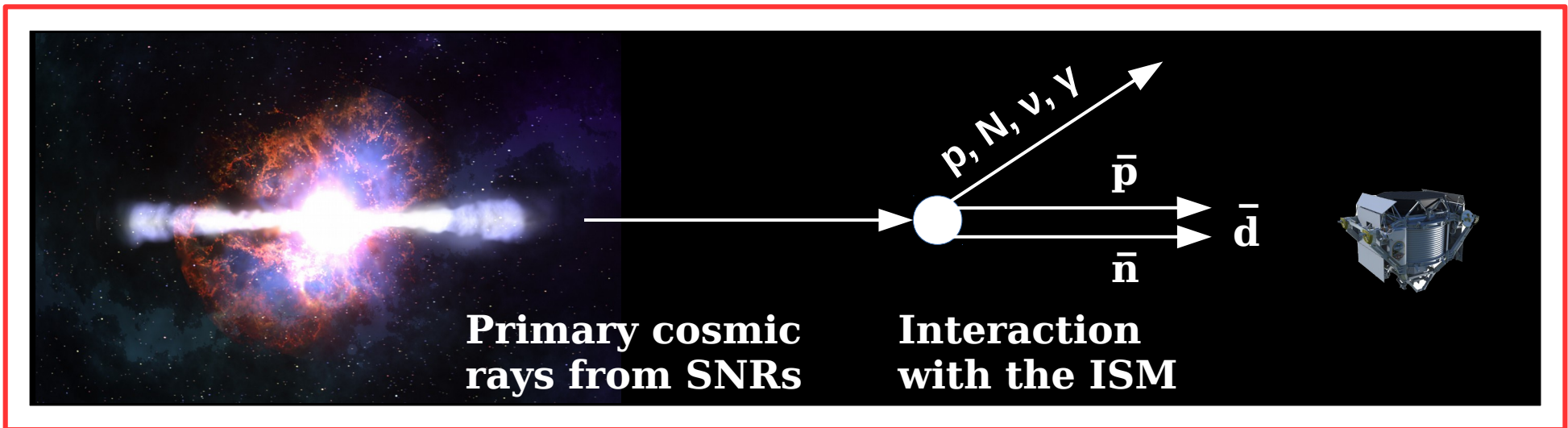
dgomezco@hawaii.edu



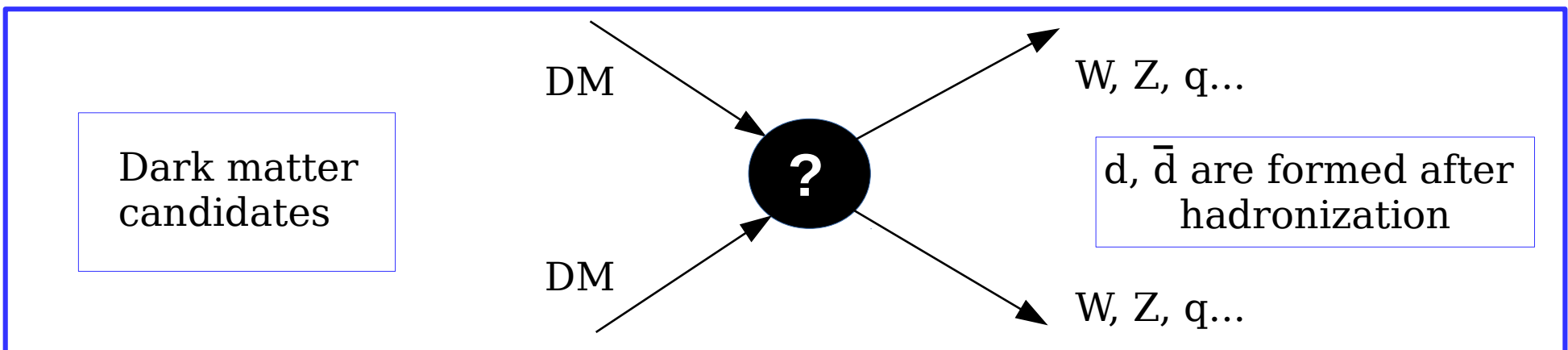
XSCRC workshop
CERN Nov 2019

Indirect search for dark matter

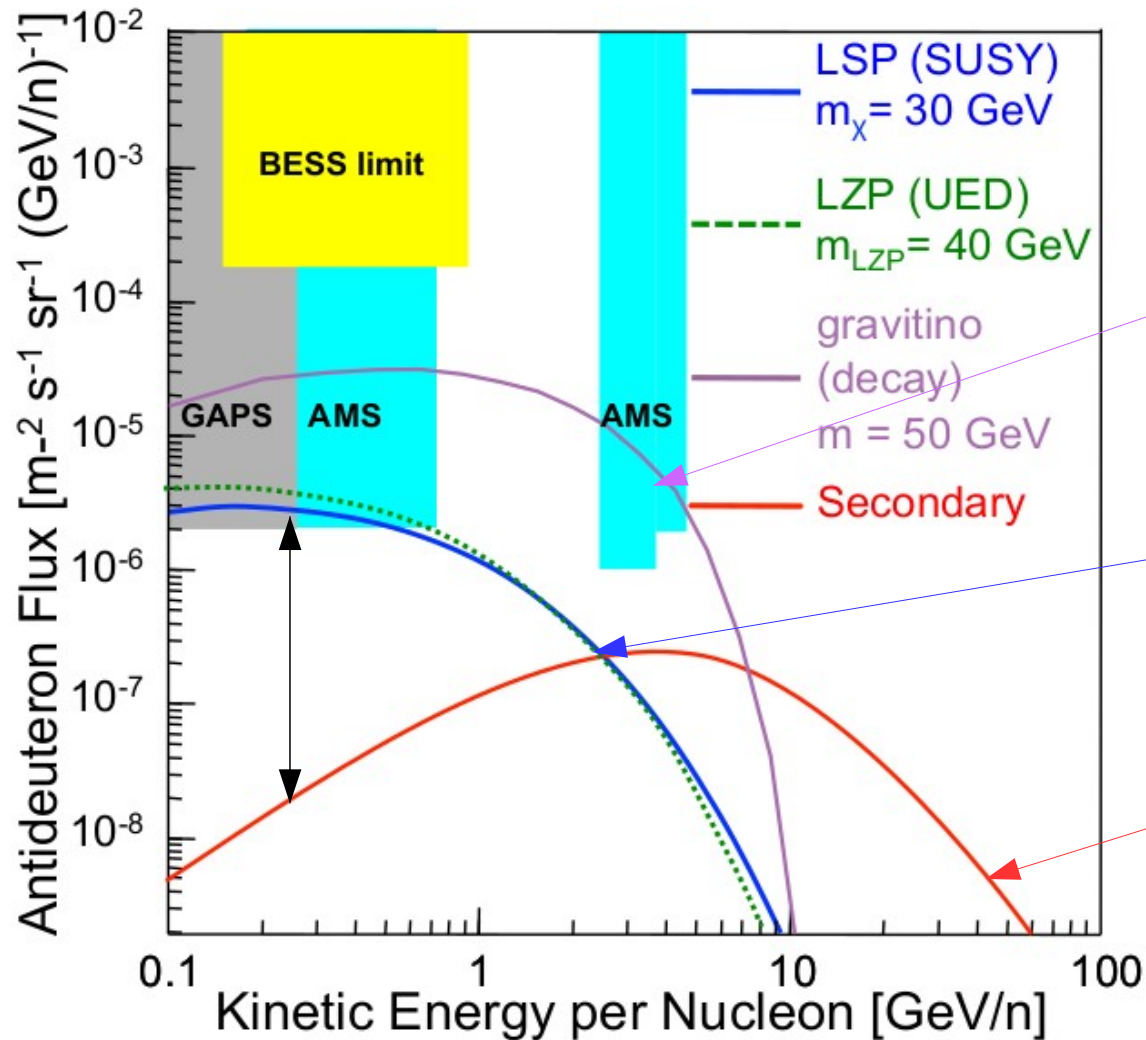
Expected antideuteron production in cosmic-ray interactions



Hypothetical antideuteron cosmic-rays from dark matter



Indirect search for dark matter



Examples of antideuteron signals from dark matter candidate interactions.

Late decays of unstable gravitinos

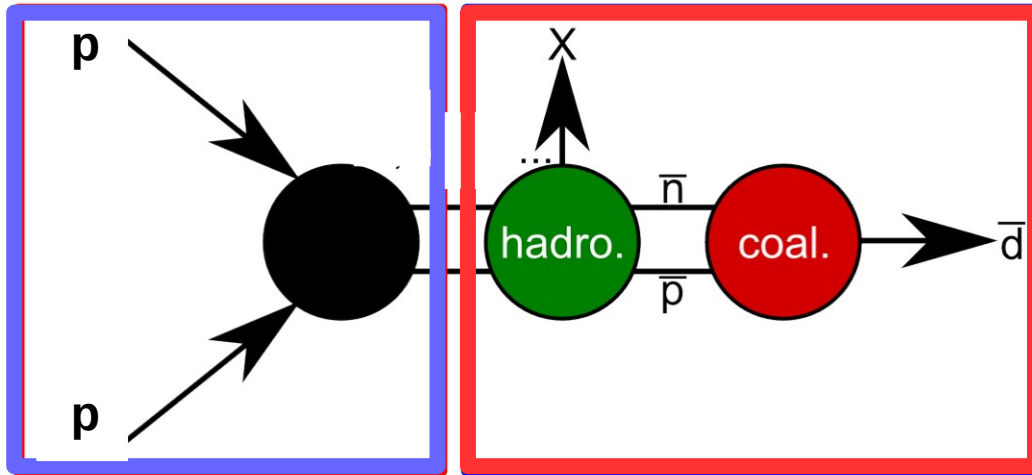
Neutralino: SUSY lightest supersymmetric particle, decay into $b\bar{b}$

Astrophysical background: Cosmic-ray collisions with the interstellar medium

Antideuterons are an important unexplored indirect detection technique.

T. Aramaki et al., Phys. Rept. 618, 1 (2016), arXiv:1505.07785 [hep-ph].

Antideuteron formation model



Coalescence Model

Deuterons and antideuterons can be formed by a pair p - n or \bar{p} - \bar{n} close in phase space.

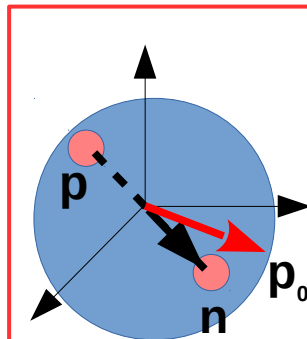
$$\gamma_d \frac{d^3 N_d}{dp_d^3} = \frac{4\pi}{3} p_0^3 \left(\gamma_p \frac{d^3 N_p}{dp_p^3} \right) \left(\gamma_n \frac{d^3 N_n}{dp_n^3} \right)$$

$p+p$, $p+\text{He}$,
 $\text{He}+p$, $p\bar{b}ar+p$
collisions

Coalescence

p_0 is extracted from
this comparison

Collisions are
simulated with
Monte Carlo
generators



Afterburner
 d and \bar{d} are
created
from the
pairs **event**
by event

The results from
simulations are
compared to data

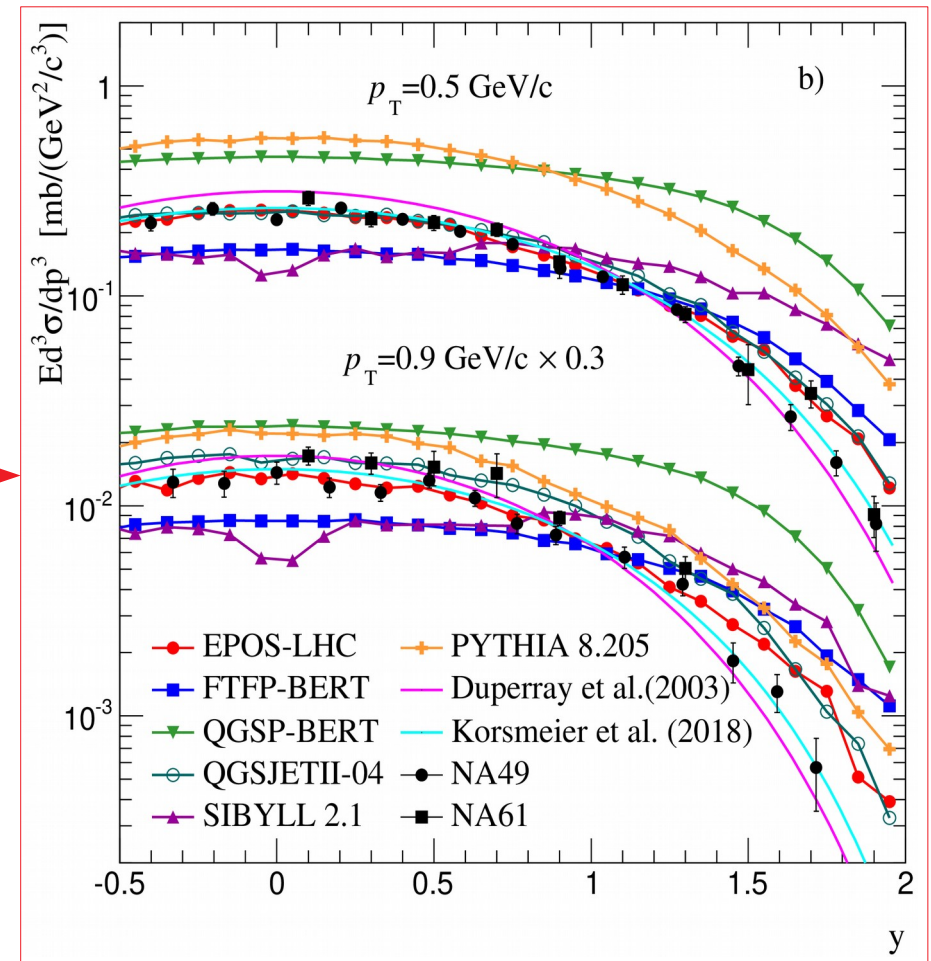
Antiproton production simulation

- To generate a correct prediction of antideuterons using MC, it is necessary to have a proper description of antiprotons.

Experiment or Laboratory	Reference	Collision	Final states	p_{lab} (GeV/c)	\sqrt{s} (GeV)
ITEP ¹	[192]	p+Be	p	10.1	4.5
CERN ¹	[193, 194]	p+p	p, \bar{p}	19.2	6.1
		p+Be	p, \bar{p}		
CERN ¹	[194]	p+p	p	24	6.8
NA61/SHINE	[195]	p+C	p	31	7.7
	[85]	p+p	p, \bar{p}		
NA61/SHINE	[85]	p+p	p, \bar{p}	40	8.8
Serpukhov ¹	[196, 197]	p+p	p, \bar{p}	70	11.5
	[198]	p+Be	p, \bar{p}		
	[199]	p+Al	p, \bar{p}		
NA61/SHINE	[85]	p+p	p, \bar{p}	80	12.3
CERN-NA49	[82]	p+p	p, \bar{p}	158	17.5
	[83]	p+C	p, \bar{p}		
CERN-NA61	[85]	p+p	p, \bar{p}		
CERN-SPS ¹	[200, 201]	p+Be	p, \bar{p}	200	19.4
		p+Al	p, \bar{p}		
Fermilab ¹	[202, 203]	p+p	p, \bar{p}	300	23.8
		p+Be	p, \bar{p}		
Fermilab ¹	[202, 203]	p+p	p, \bar{p}	400	27.4
		p+Be	p, \bar{p}		
CERN-ISR	[204]	p+p	p, \bar{p}	1078	45.0
CERN-ISR	[204]	p+p	p, \bar{p}	1498	53.0
CERN-LHCb	[86]	p+He	\bar{p}	6.5×10^3	110
CERN-ALICE	[84]	p+p	p, \bar{p}	4.3×10^5	900
CERN-ALICE	[84]	p+p	p, \bar{p}	2.6×10^7	7000

Proton and antiproton data list on p+p and p+A collisions to be compared to simulations. **D. Gomez-Coral et al. Phys. Rev. D 98, 023012 (2018) arXiv:1806.09303 [astro-ph.HE]**

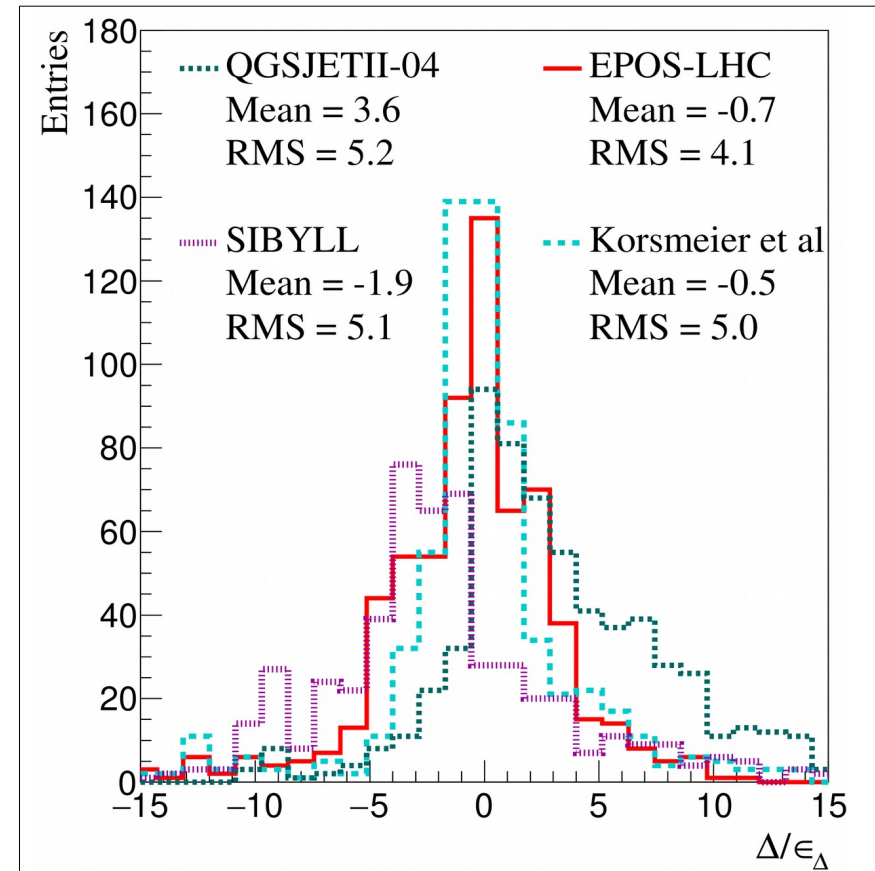
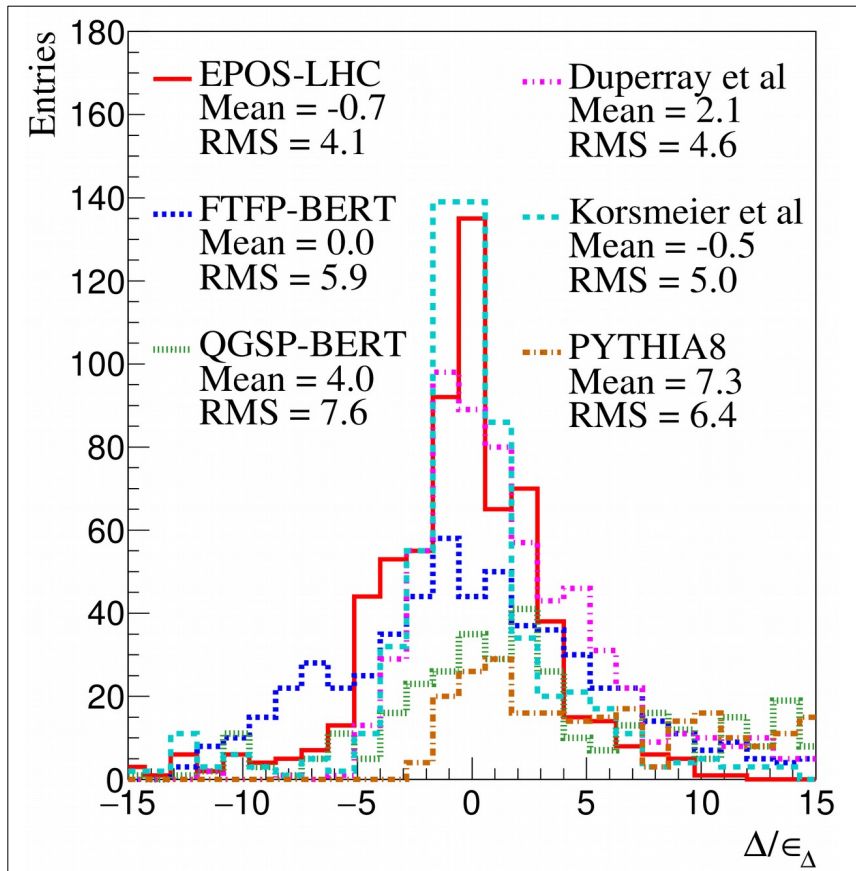
Invariant differential cross section for **antiprotons** in p+p collisions at 158 GeV/c, as function of rapidity



Antiproton production simulation

- Simulation is compared to data point by point.
- The most reliable MC model is selected from the comparison to data.

$$\frac{\Delta}{\epsilon_{\Delta}} = \frac{\left(E \frac{d^3\sigma}{dp^3}^{sim} - E \frac{d^3\sigma}{dp^3}^{data} \right)}{\sqrt{(\epsilon_{sim})^2 + (\epsilon_{data})^2}}$$



Antideuteron production simulation

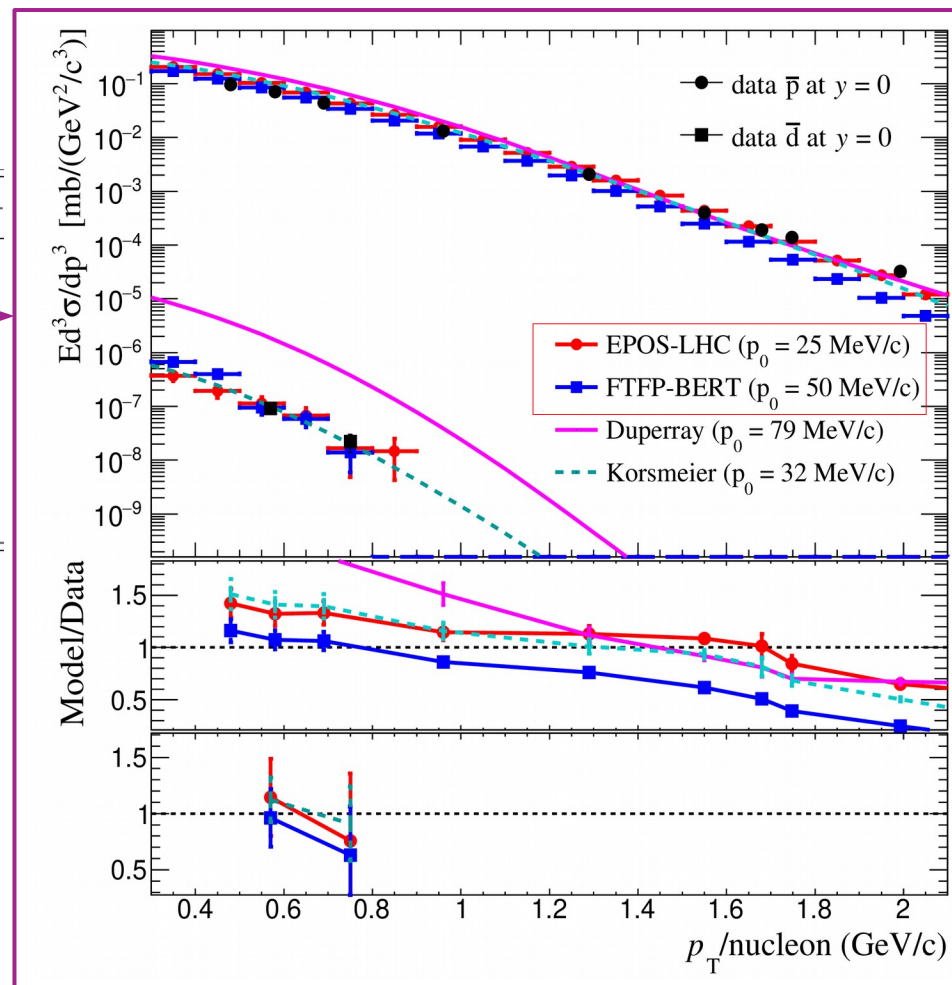
- The coalescence momentum (p_0) is determined from the fit of simulations to data.

Experiment or Laboratory	Reference	Collision	p_{lab} (GeV/c)	\sqrt{s} (GeV)	No. of points	
					d	dbar
CERN	[194]	p+p	19	6.15	6	0
CERN	[194]	p+p	24	6.8	4	0
Serpukhov	[198]	p+p	70	11.5	7	2
CERN-SPS	[200, 205]	p+Be	200	19.4	6	3
		p+Be			3	5
		p+Al			3	3
Fermilab	[203]	p+Be	300	23.8	4	1
CERN-ISR	[206, 207, 208]	p+p	1497.8	53	3	8
CERN-ALICE	[155, 209]	p+p	4.3×10^5	900	3	3
CERN-ALICE	[155, 209, 210]	p+p	2.6×10^7	7000	21	20

Deuteron and antideuteron data list on p+p and p+A collisions to be compared to simulations. **D. Gomez-Coral et al. Phys. Rev. D 98, 023012 (2018) arXiv:1806.09303 [astro-ph.HE]**

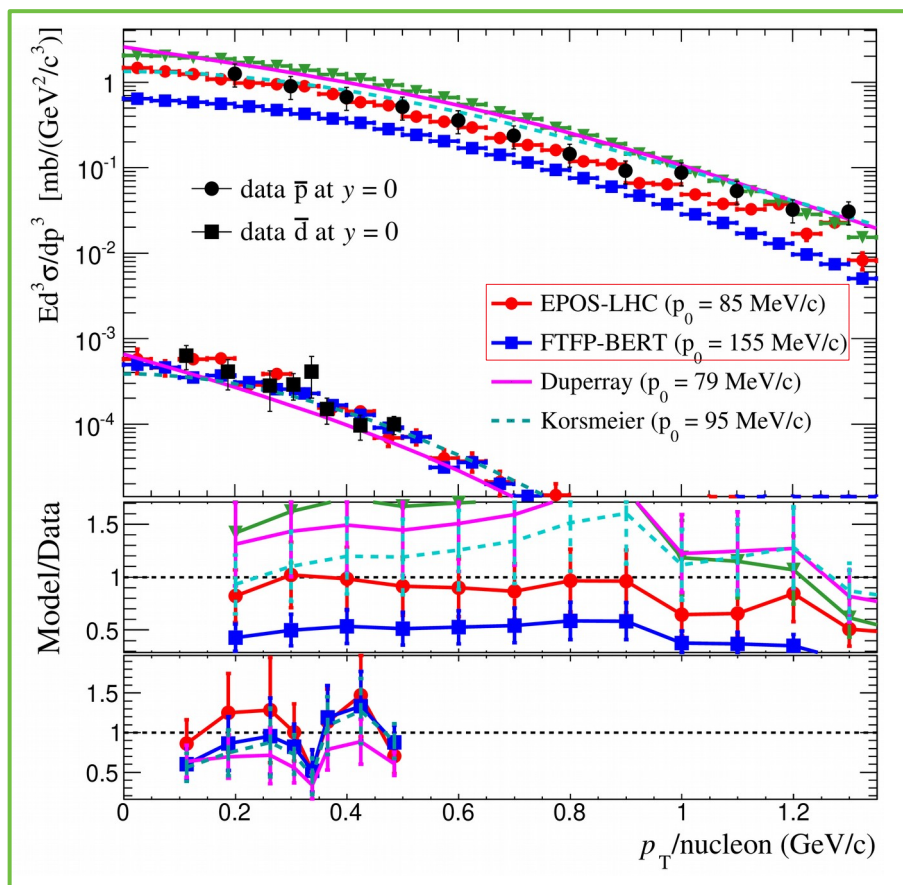
Antideuteron invariant differential cross section in p+p collisions at 70 GeV/c, as function of p_T

p+p at $\sqrt{s} = 11.5$ GeV



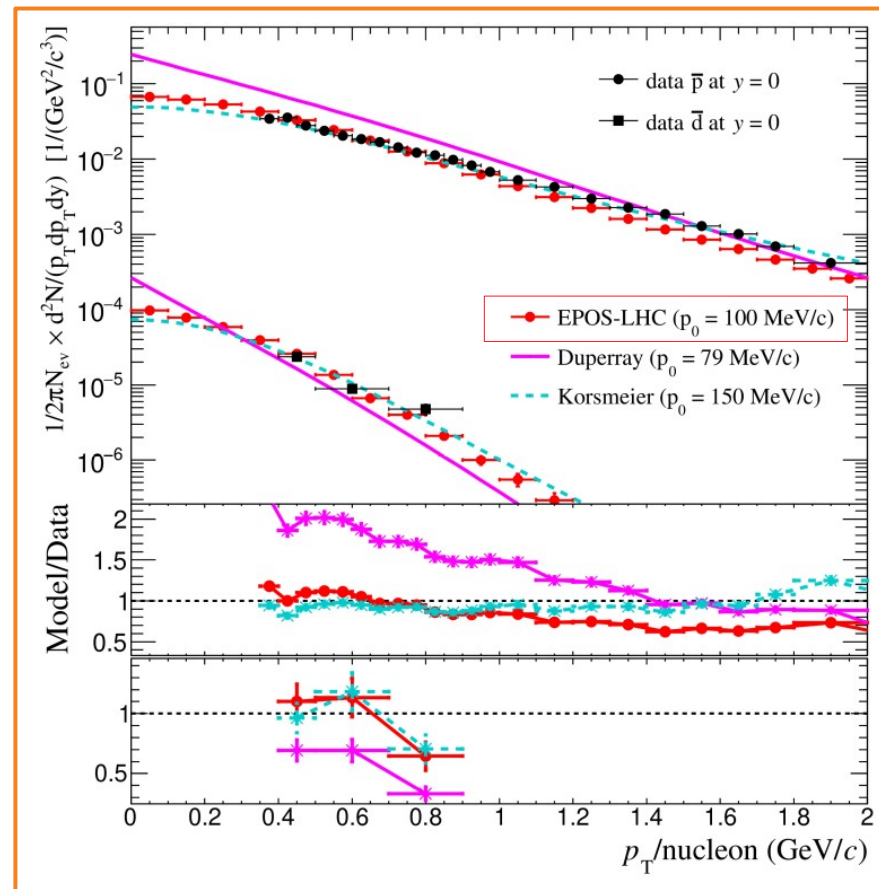
Antideuteron production simulation

p+p at $\sqrt{s} = 53$ GeV



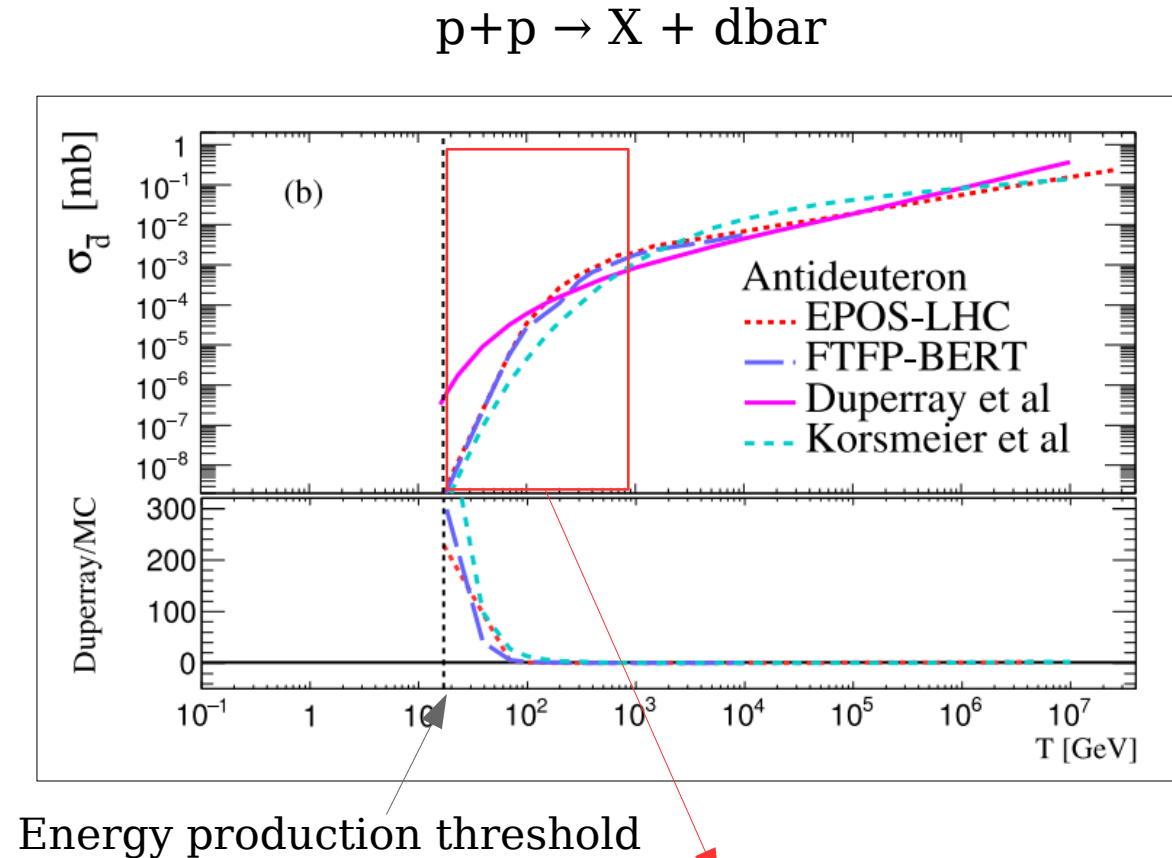
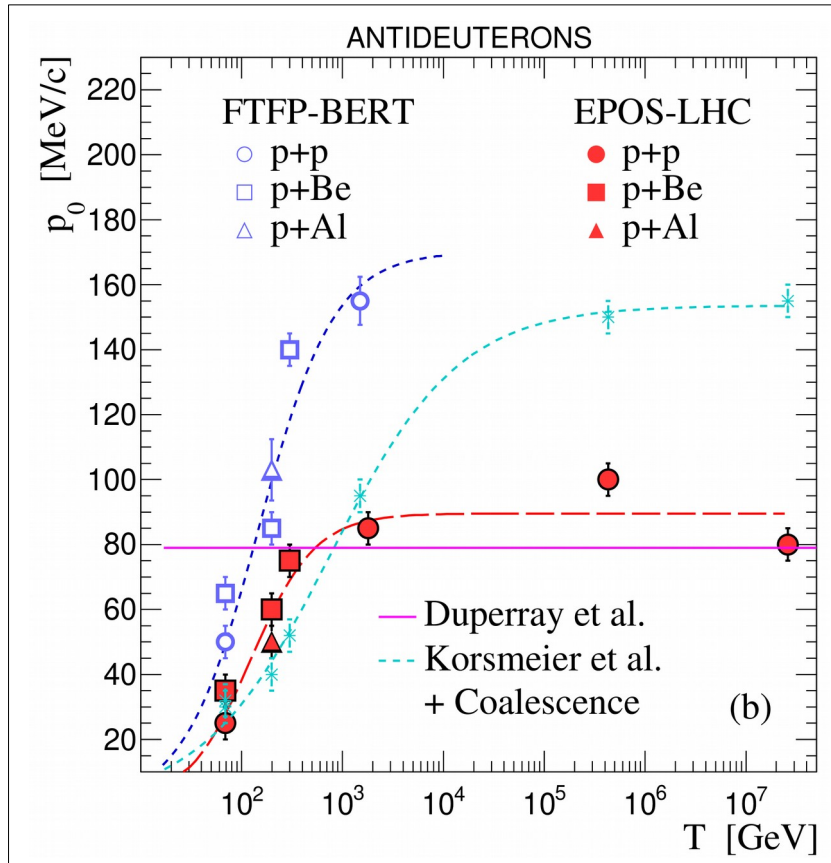
Antideuteron invariant differential cross section as function of p_T compared to ISR data.

p+p at $\sqrt{s} = 900$ GeV



Antideuteron invariant differential cross section as function of p_T compared to ALICE-LHC data.

Coalescence momentum (p_0) and production cross section



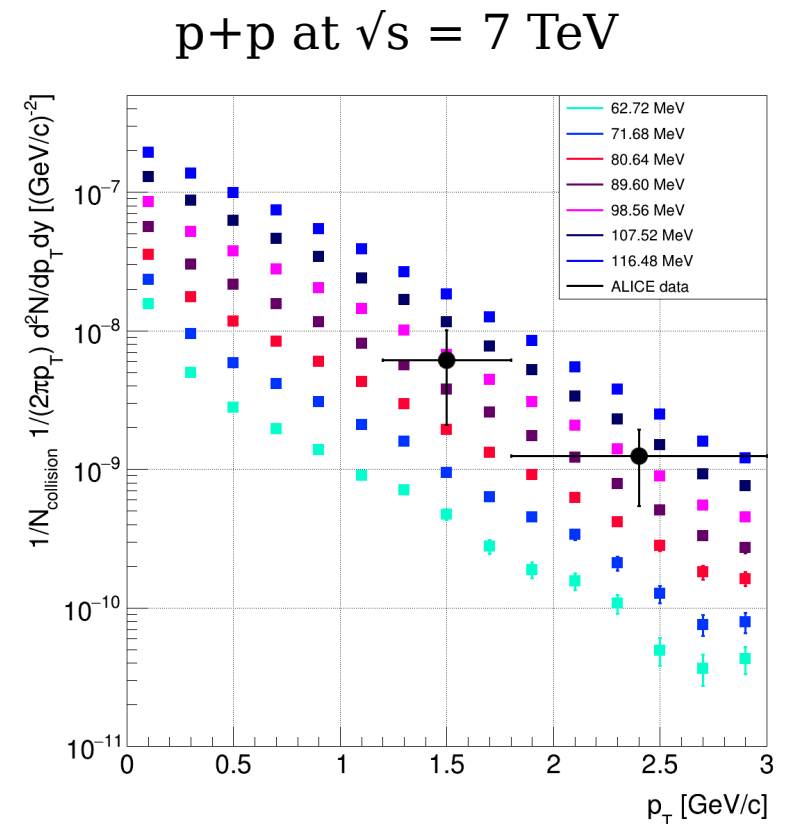
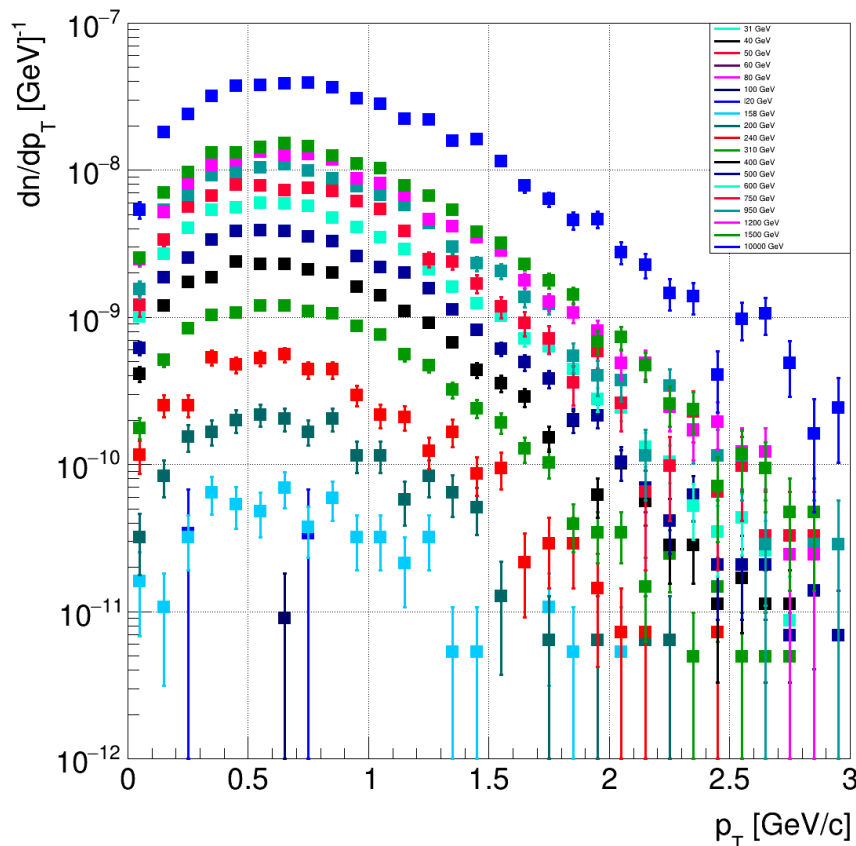
- p_0 is parameterized as function of the projectile kinetic energy.
- p_0 is similar for p+p and p+Be collisions.

- p_0 changes in the energy region of major importance for cosmic ray production.

Antihelium production simulation

- MC coalescence is expanded to merge 3 antinucleons from p-p interactions.
- High computing power is required ~ 2000 years so far.

- Using same p_0 as for dbar shows very good agreement with ALICE antihelium-3 data



Anirvan Shukla PhD student UH

Propagation with Galprop56

$$\frac{\partial f(p, \vec{r}, t)}{\partial t} = \vec{\nabla} \cdot (D_{xx}(p, \vec{r}) \vec{\nabla} f - \vec{V} f) + \frac{\partial}{\partial p} p^2 D_{pp} \frac{\partial}{\partial p} \frac{1}{p^2} f - \frac{\partial}{\partial p} \left[\dot{p} f - \frac{p}{3} (\vec{\nabla} \cdot \vec{V}) f \right] - \frac{1}{\tau_f} f - \frac{1}{\tau_r} f + q(p, \vec{r}, t),$$

Antideuteron source term
↓

Set 1 DR
Without convection

z [kpc]	$D_0/10^{28}$ [cm ² s ⁻¹]	δ	V_{alf} [km s ⁻¹]
6	4.37	0.494	7.64

Helium Proton

R1	R2	γ_1	γ_2	γ_3
5.78 GV	304 GV	1.74	2.35	2.178

R1	R2	γ_1	γ_2	γ_3
5.78 GV	304 GV	1.69	2.29	2.12

T. A. Porter et al., 2017

Set 2 DCR
With convection

z [kpc]	$D_0/10^{28}$ [cm ² s ⁻¹]	δ	V_{alf} [km s ⁻¹]	V_{conv} [km s ⁻¹]	dV/dz_{conv} [km s ⁻¹ kpc ⁻¹]
4	4.3	0.395	28.6	12.4	10.2

R1	R2	γ_1	γ_2	γ_3
7 GV	360 GV	1.69	2.44	2.28

R1	R2	γ_1	γ_2	γ_3
7 GV	330 GV	1.71	2.38	2.21

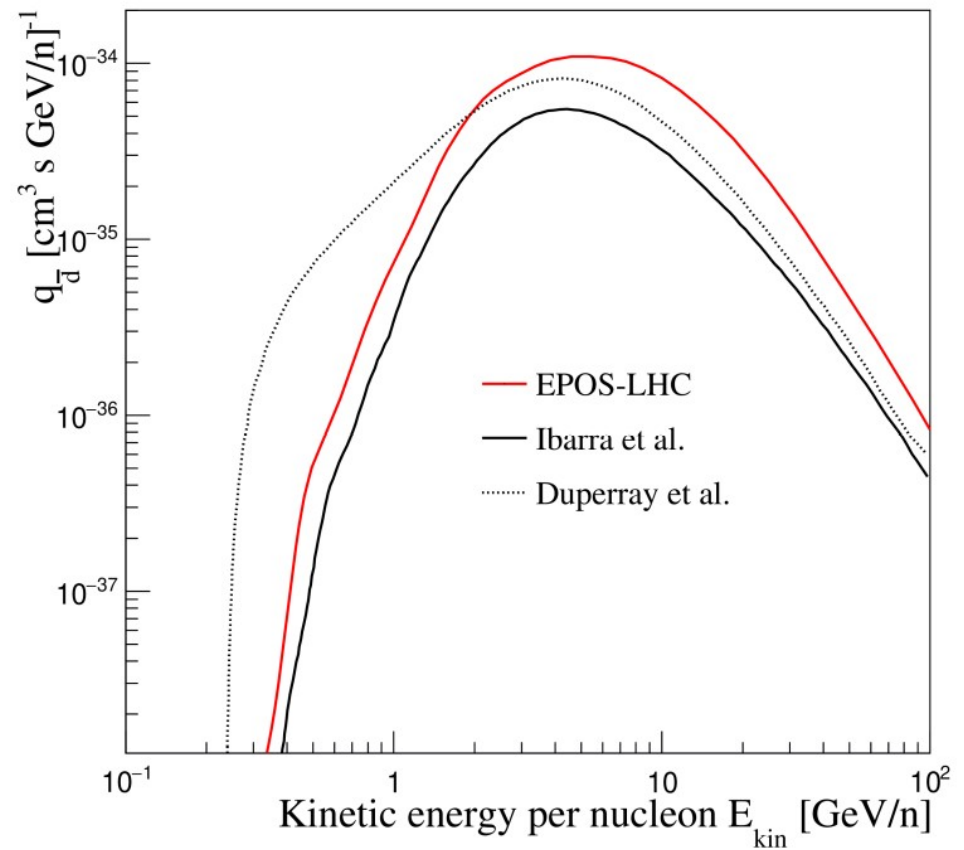
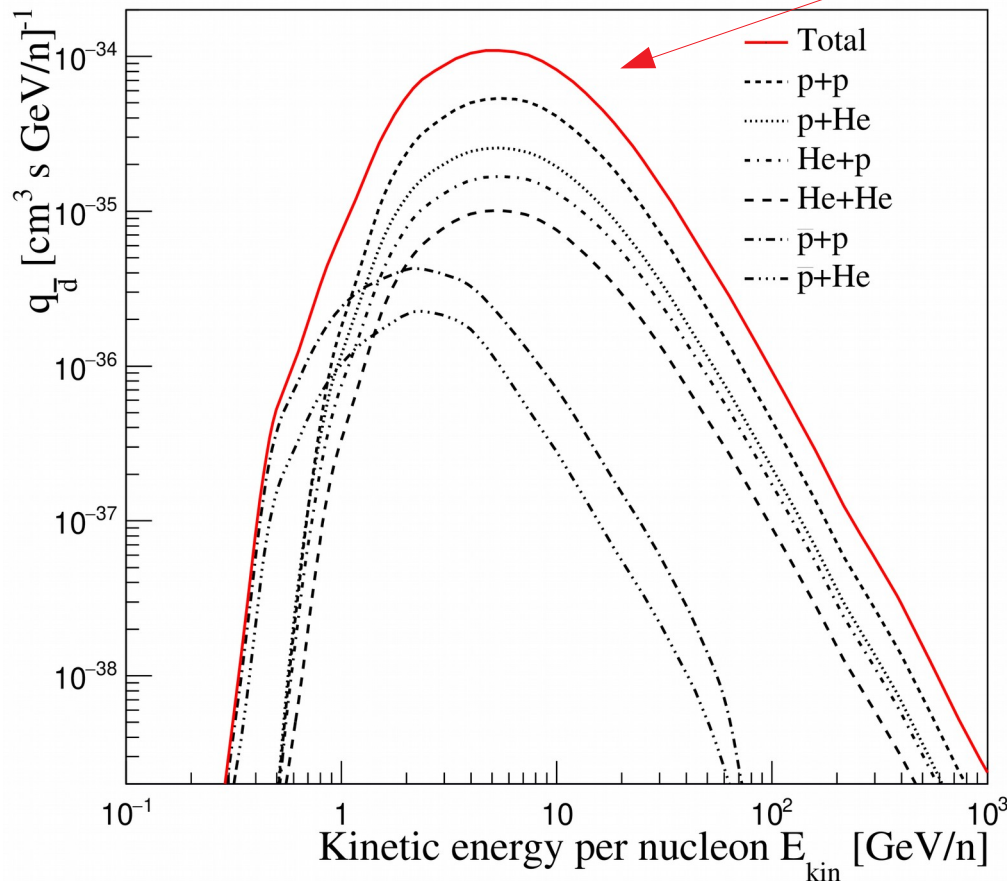
M. J. Boschini et al 2017 PoS(ICRC2017)278

Antideuteron source term

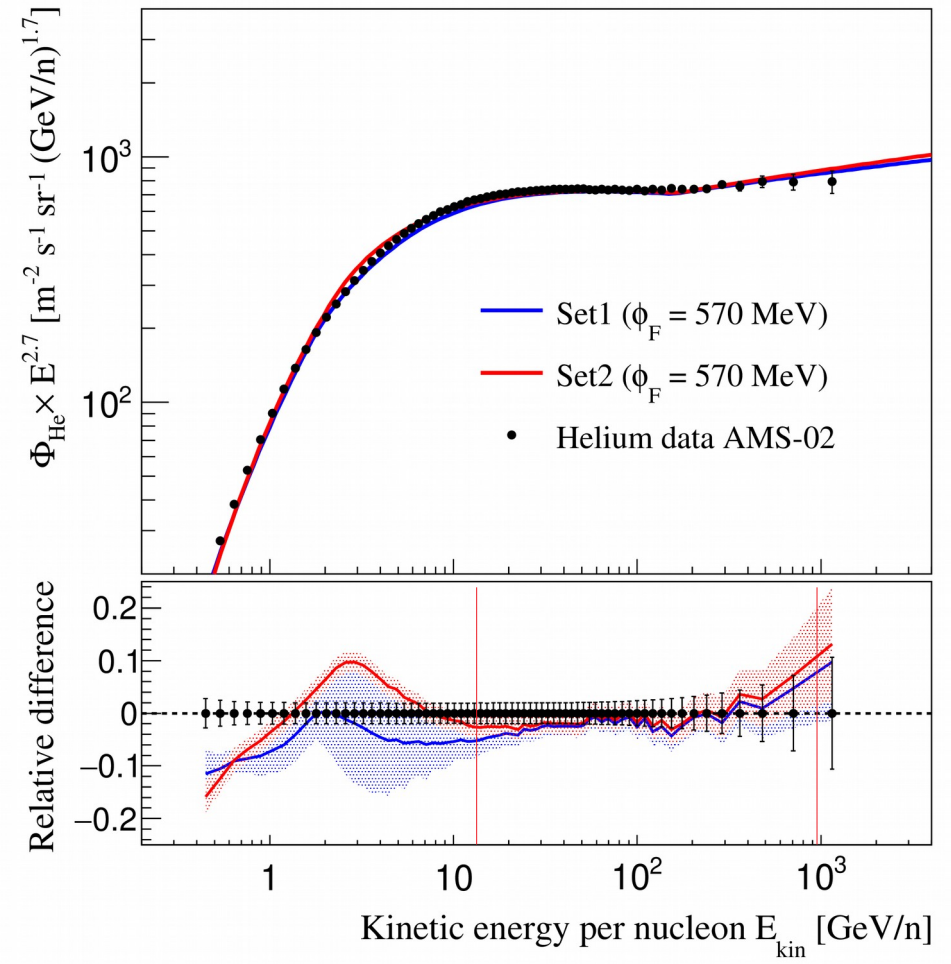
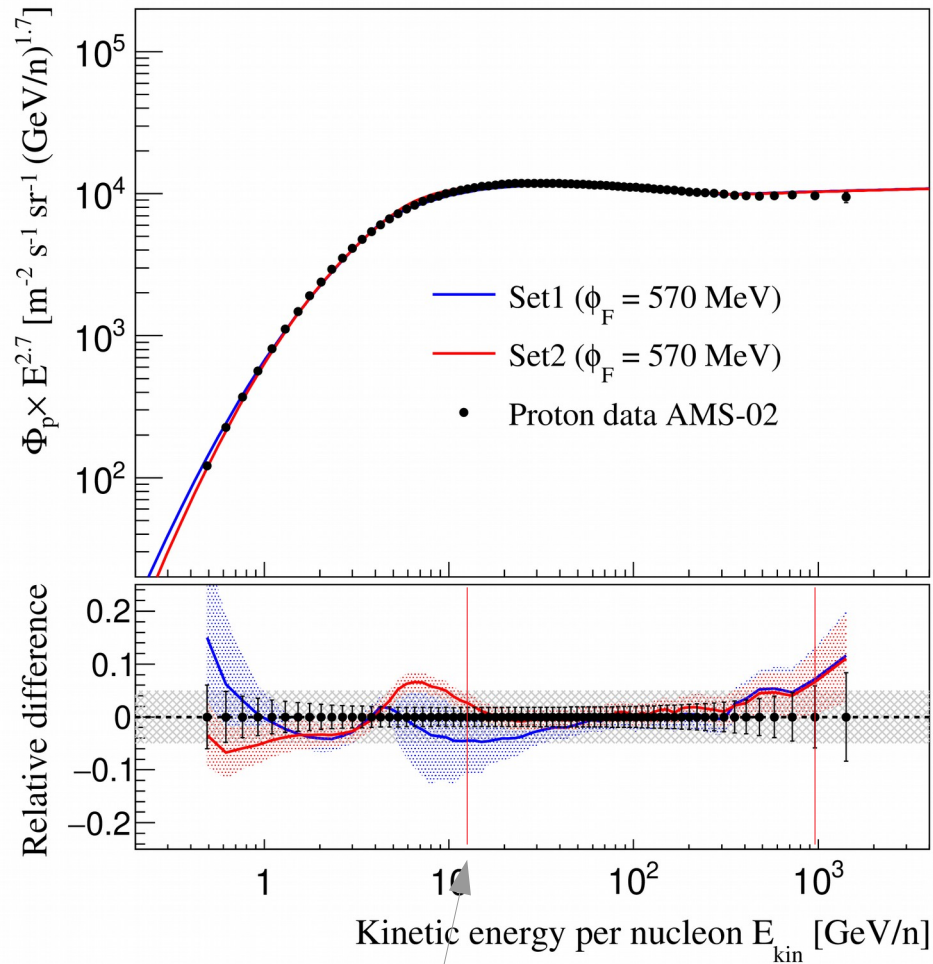
$$q_{\bar{d}}^{sec}(E_{kin}^{\bar{d}}) = \sum_{i=p,He,\bar{p}} \sum_{j=p,He} 4\pi n_j \int_{T_{min}}^{\infty} dT_i \left(\frac{d\sigma}{dE_{kin}^{\bar{d}}} \right)_{ij} \Phi_i(T_i)$$

EPOS-LHC cross-section

p, He Fluxes

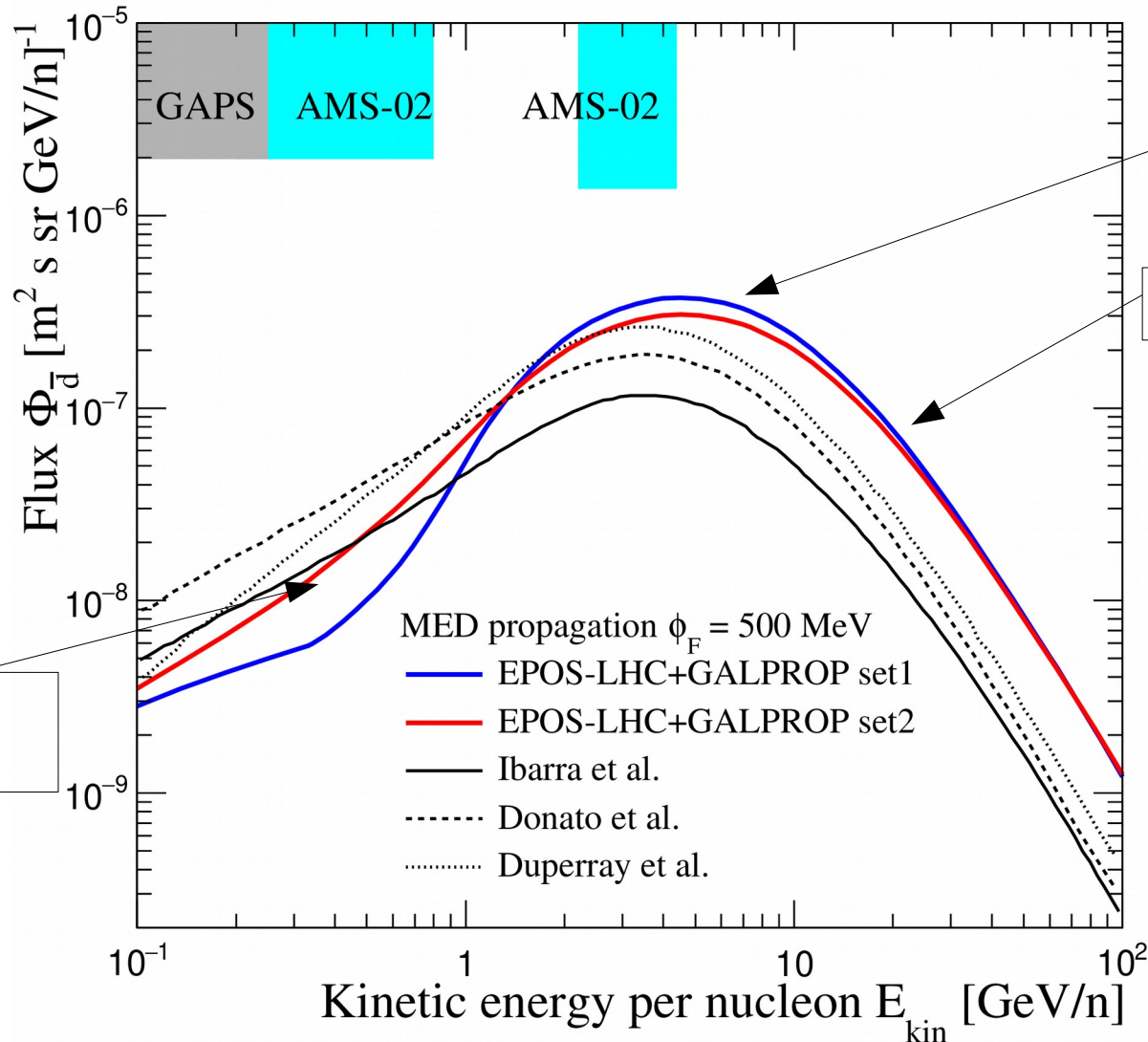


Proton and Helium fluxes



Energy production threshold

Secondary Antideuteron Flux

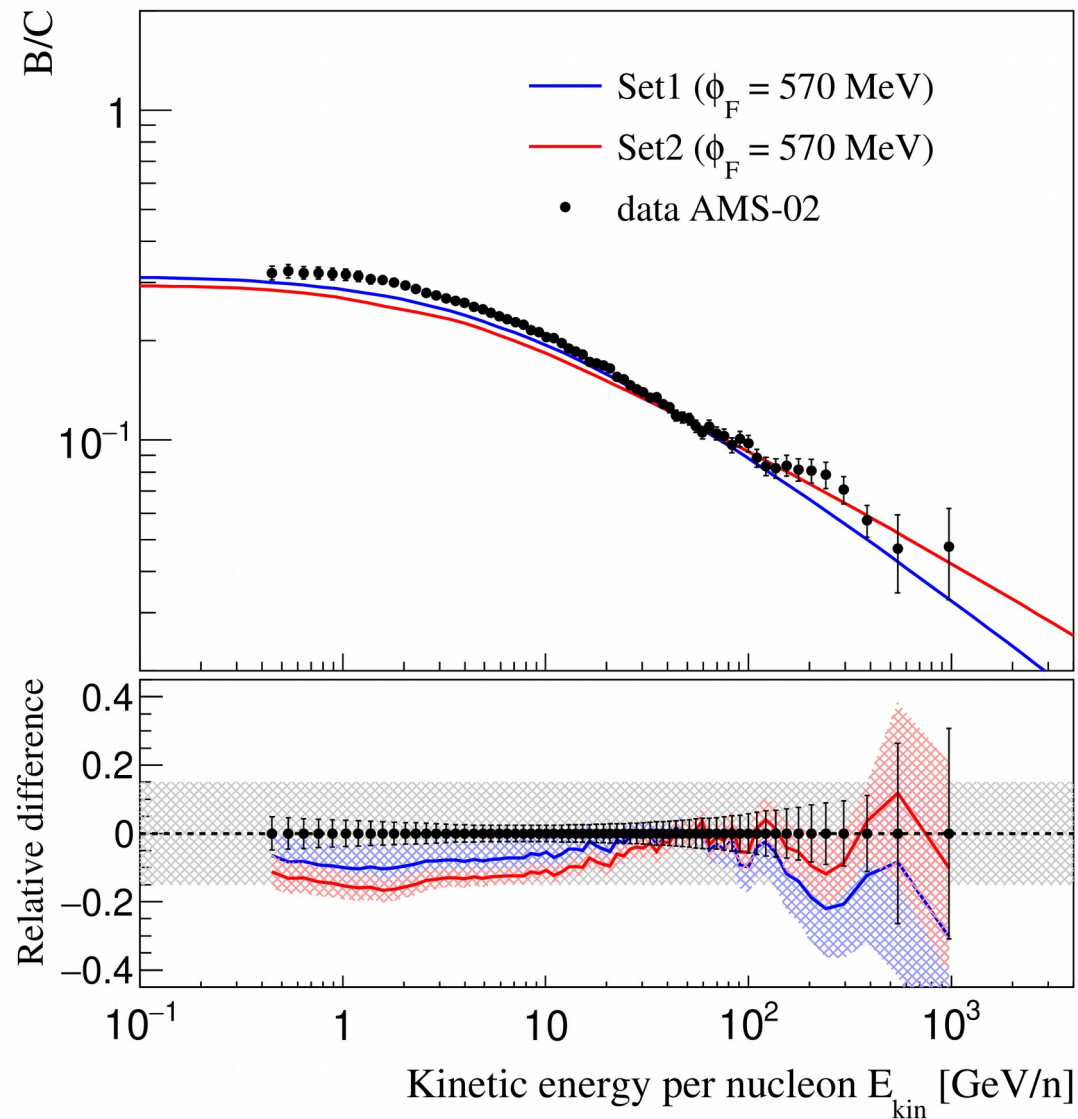


Summary

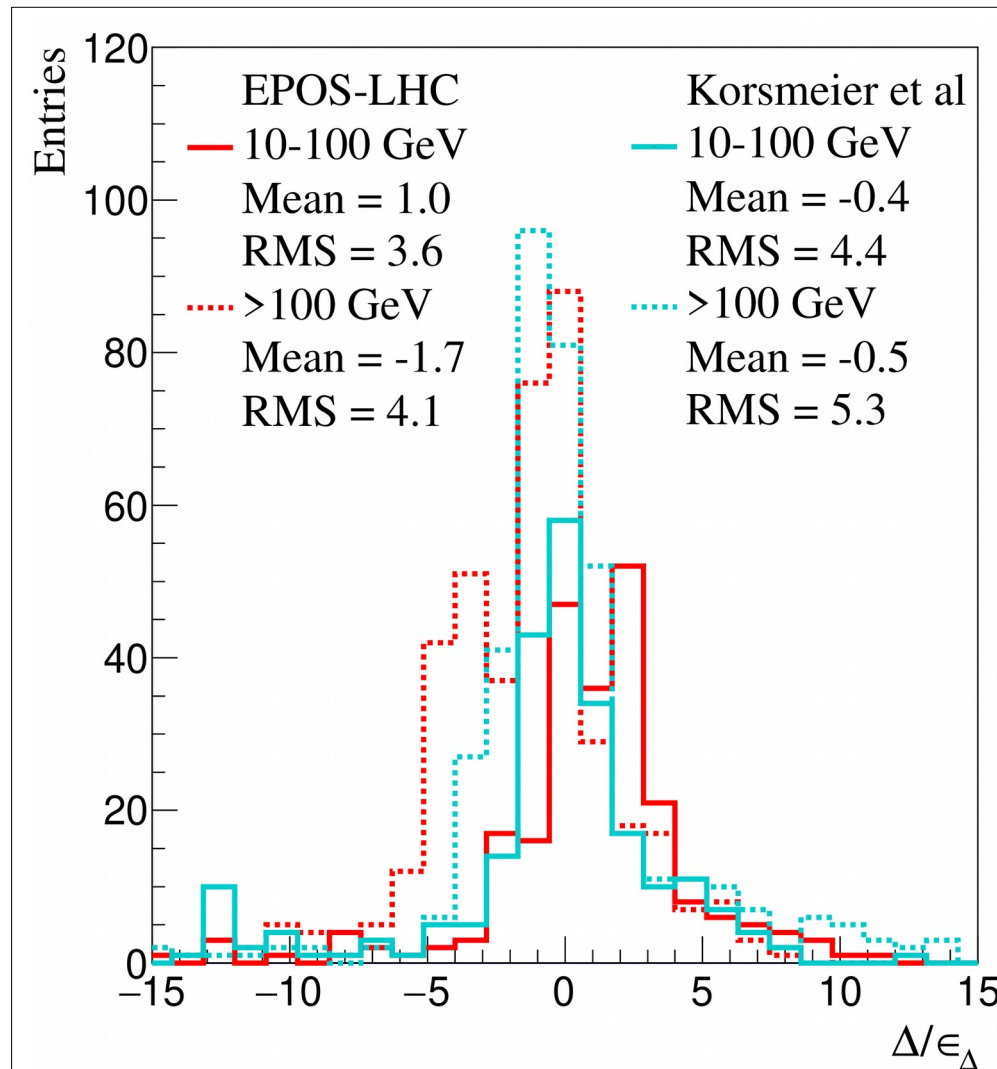
- A new study on the secondary antideuteron production was presented, using a high-energy MC generator (EPOS-LHC) and the coalescence model.
- Simulations were compared to an extensive data set, including new measurements from NA61 and ALICE-LHC to obtain the coalescence momentum (p_0).
- From the comparison to data, it seems the coalescence momentum (p_0) depends on the collision energy. As consequence:
 - ◆ Antideuteron production cross section shows important differences with respect to previous calculations in the region of interest for CR antideuteron production.
 - ◆ Antideuteron flux shape is slightly modified and its maximum is shifted above 4 GeV/n compared to other works.
 - ◆ Antideuteron flux is higher than previous estimations but it remains below experiment expected sensitivities.

Thank you!

Boron to Carbon ratio

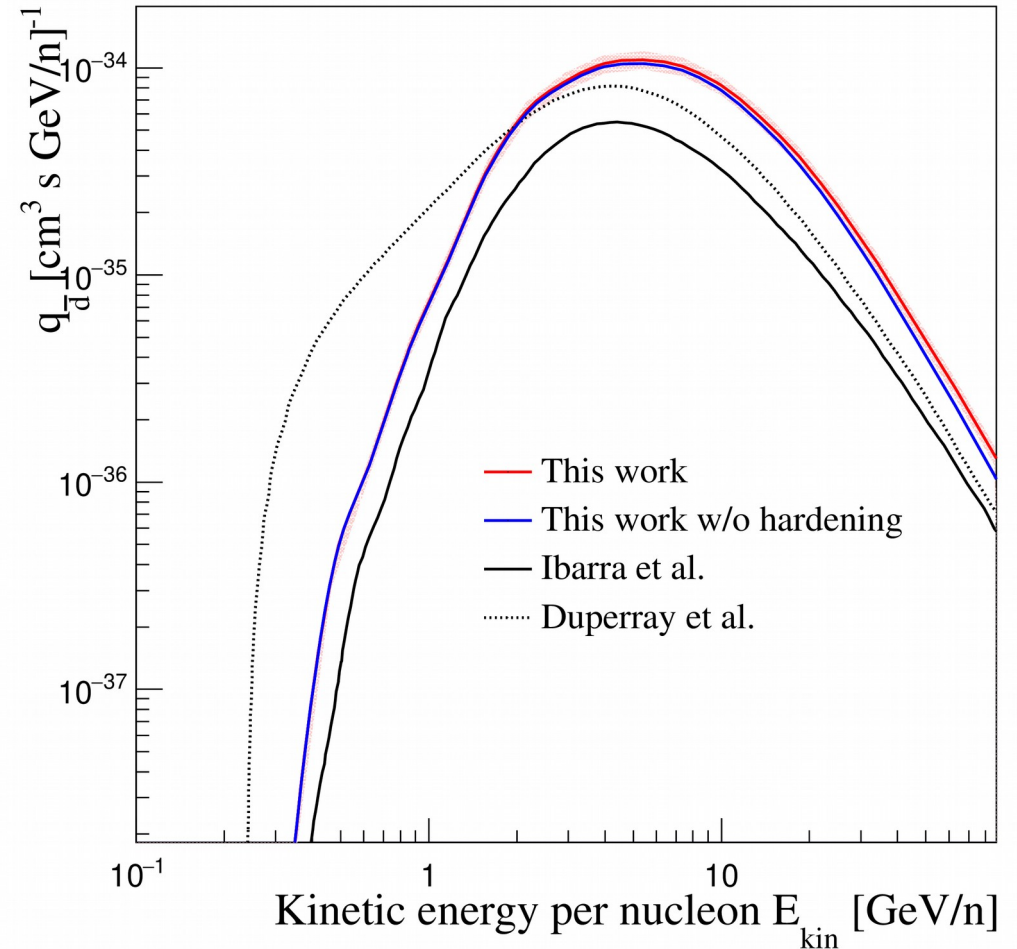
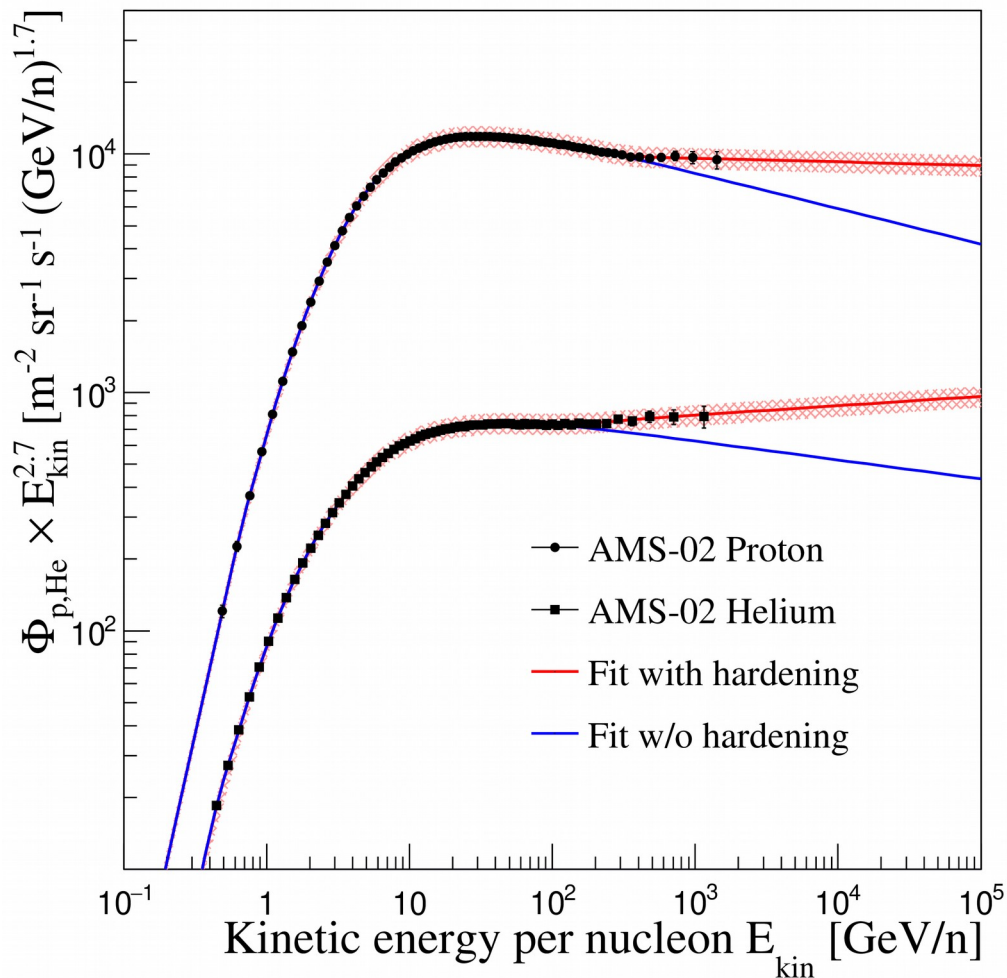


Antiproton production simulation



Antideuteron source term

- Proton and helium fluxes with and without hardening are inserted in the convolution.



- Hardening increases dbar flux by less than 10%
- dbar production is higher than in previous works.