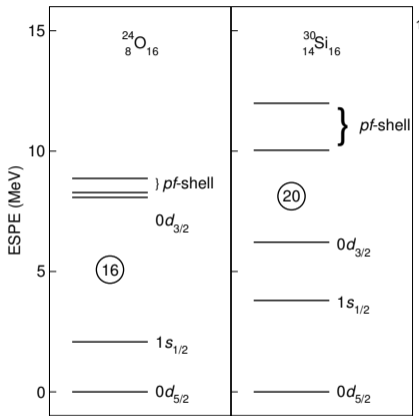
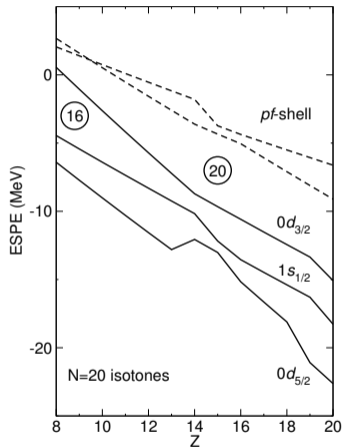


Measuring the $^{28}\text{Mg}(d,p)$ reaction with ISS

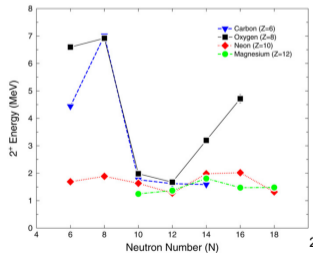
Patrick MacGregor
ISOLDE Solenoidal Spectrometer Workshop
27 August 2019



Motivation — $N = 20$ shell gap evolution

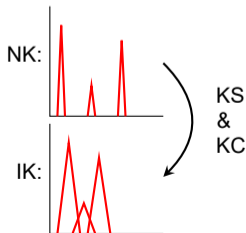
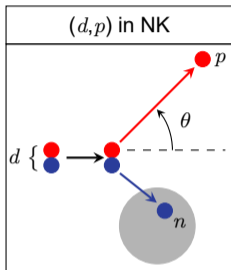
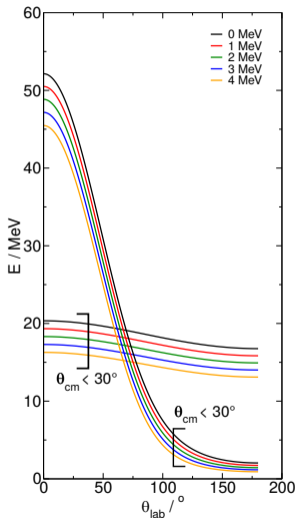


- Observed weakening in the $N = 20$ shell closure with decreasing Z .
- Weakening caused by relative strength of interaction between neutrons and protons in different orbits ($\nu d_{3/2}$ and $\pi d_{5/2}$).
- Mapping size of shell gaps crucial for understanding this region.
- Transfer reactions can be used here to map evolution of pf -states and how separation evolves.



¹ Adapted from T. Otsuka et al. *Eur. Phys. J. A.* 15 (2002), pp. 151–155.

² C.R. Hoffman et al. *Phys. Lett. B* 672.1 (2009), pp. 17–21.



- Ejectile, p , provides information for the populated state in ^{29}Mg :
 - Yields \rightarrow cross section.
 - θ \rightarrow angular momentum.
 - Ejectile energy \rightarrow excitation of residual nucleus.

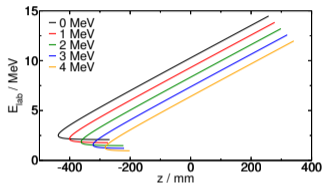
- Single-particle strength split by correlations between nucleons.
- Deduce spectroscopic factors, $S_{j\ell}$, which measure how close each state looks like a n in IPM state:

$$\left(\frac{d\sigma}{d\Omega}\right)_{\text{EXP}} = S_{j\ell} \left(\frac{d\sigma}{d\Omega}\right)_{\text{DWBA}}$$

$$\text{where } |^{29}\text{Mg}; j, \ell\rangle \sim \sum_{j, \ell} S_{j\ell} (|^{28}\text{Mg}; 0, 0\rangle \otimes |n; j, \ell\rangle).$$

- NK works if the target is stable/long-lived BUT $^{28}\text{Mg} \rightarrow ^{28}\text{Al}$ via β^- ; $\tau_{1/2} = 21$ h.
- IK allows transfer on radioactive nuclei, but introduces some non-trivial problems:

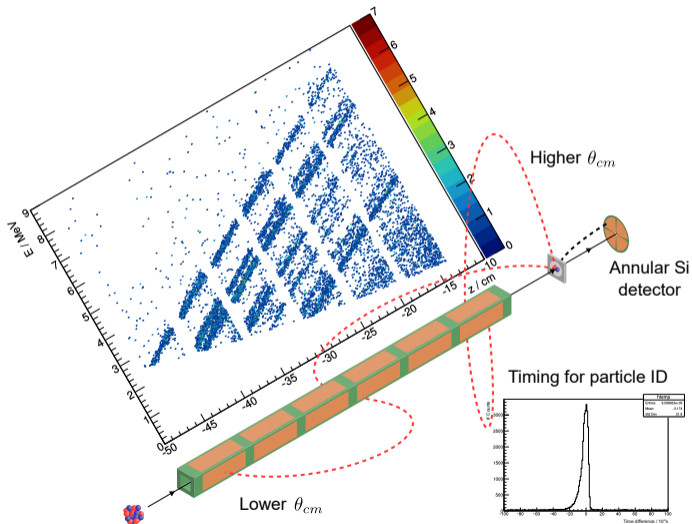
- Kinematic shift (KS)** - broadens peaks because of large $\frac{dE}{d\theta}$ for a finite angular acceptance, $\Delta\theta$.
- Kinematic compression (KC)** reduces energy difference between states.

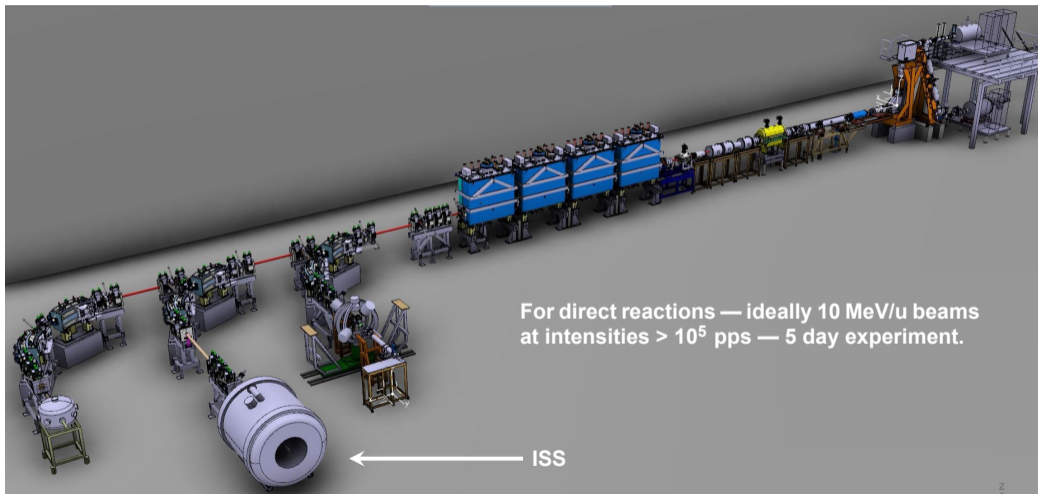


Measure z rather than θ .

For a given z , $E_{lab} \propto \Delta E_x$.
 \Rightarrow no compression in the solenoid.
 \Rightarrow better resolution.

$$E_{cm} = E_{lab} + \frac{1}{2} m V_{cm}^2 - \frac{m V_{cm} z}{T_{cyc}}$$



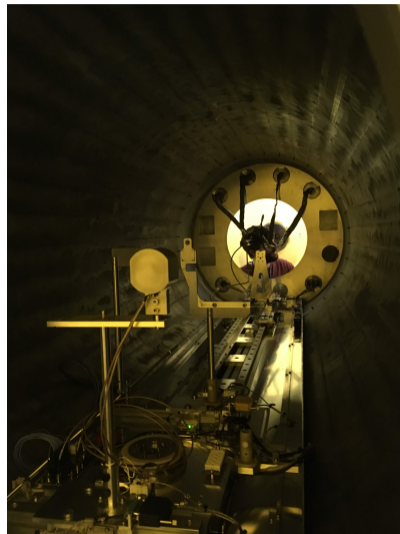


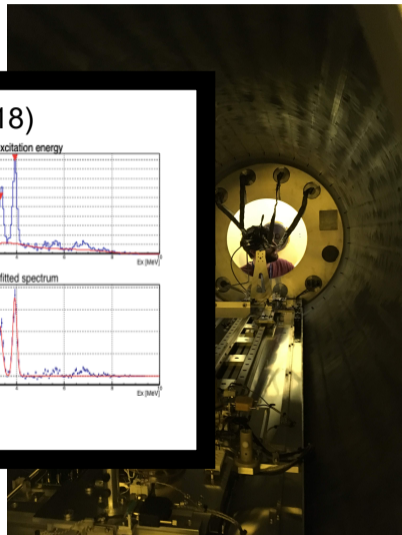
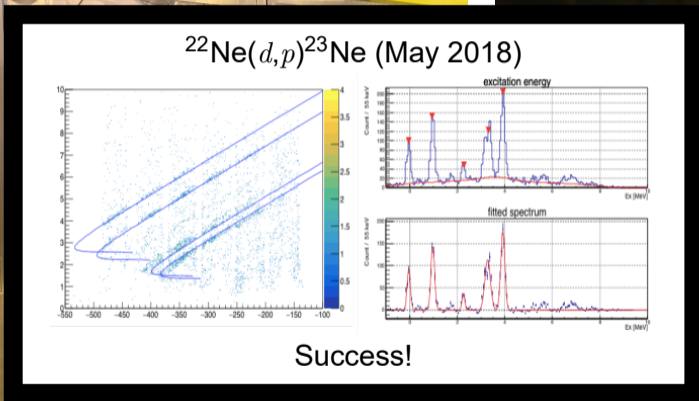
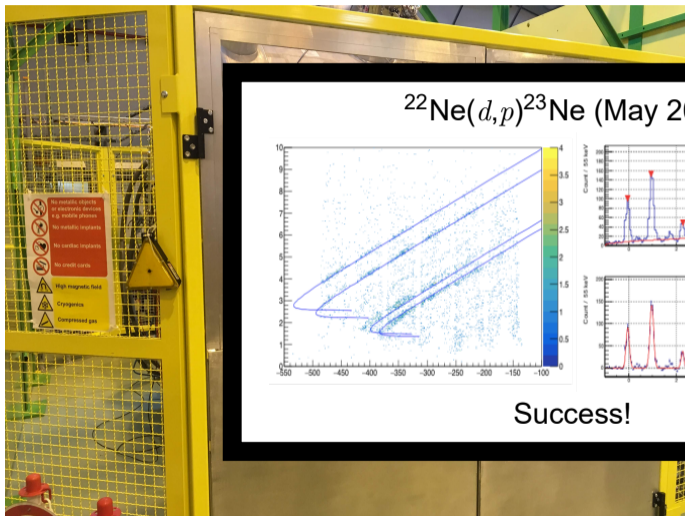
3 URL: <http://hie-isolde-project.web.cern.ch/about-hie-isolde>.

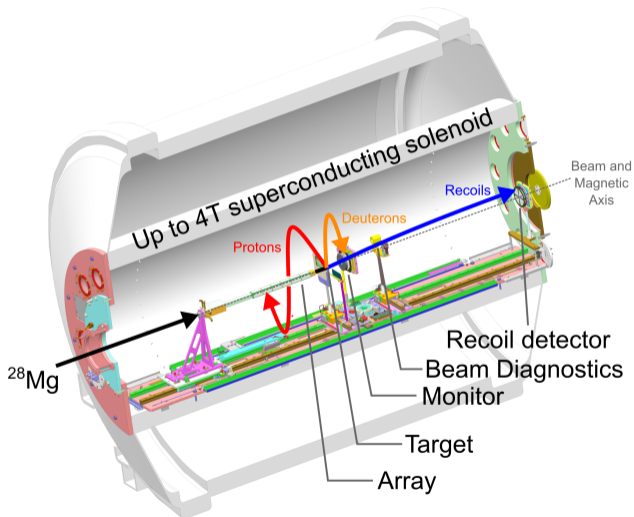
ISOLDE Solenoidal Spectrometer (ISS)

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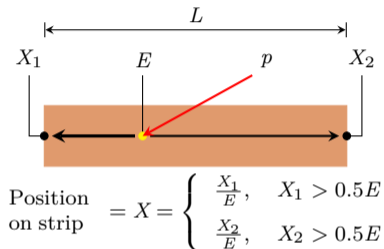






- Based on HELIOS spectrometer at Argonne National Laboratory.
- Uses HELIOS array on loan from Argonne to allow for experiments before LS2.
- Detectors:
 - ▶ HELIOS array — backwards scattered p .
 - ▶ S1 detector — elastically scattered d .
 - ▶ $E\Delta E$ recoil detector — scattered beam.
- Measure energies, position, and times on each detector.
- Beam energy: 9.473 MeV/u ($dE/E = 0.3\%$).
- Maximum 10^6 pps.

Positions:



- Calculate position on strip using gain-matched X_1 , X_2 , and E .
- Use laser alignment from CERN team to calculate distance from target:

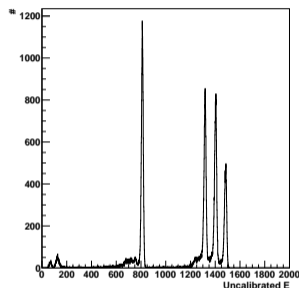
$$z = L \left(X - \frac{1}{2} \right) - z_{\text{off}} - d_i,$$

z is the distance along the beam axis from the target.

d_i is the distance along the array to the centre of strip i .

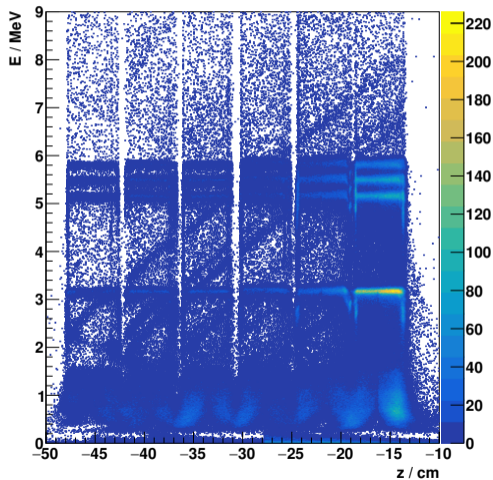
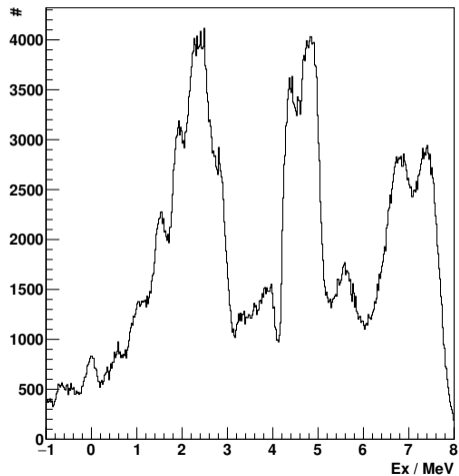
z_{off} is the distance from the target to the array.

Energies:

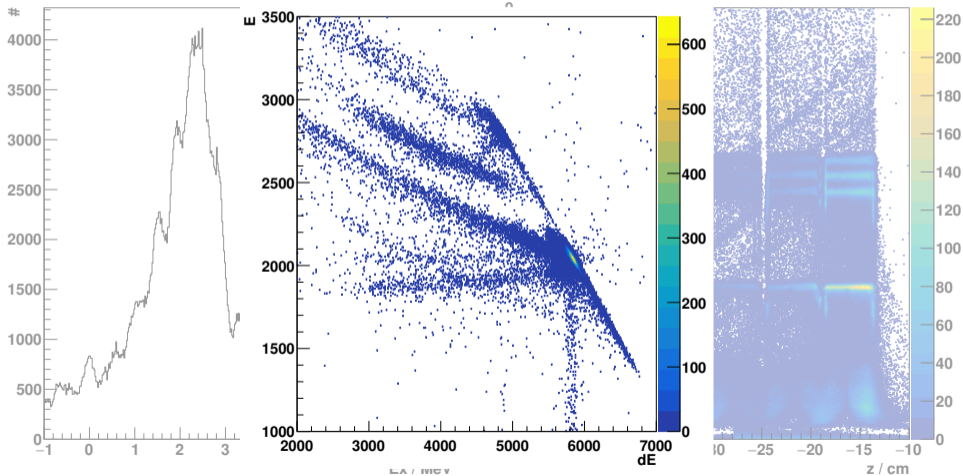


- Rough energy calibration with quadruple α -source.
- Gain match X_1 and X_2 to each other.
- Match X_1 and X_2 to E .
- α -source not sufficient, as α 's lose energy in the target.
- α -source calibration improved by calibrating to known states in ^{29}Mg .

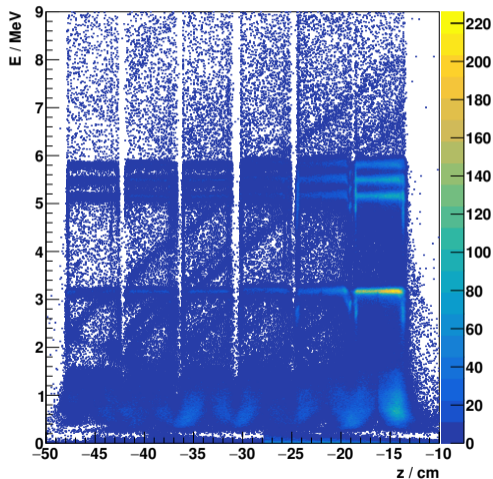
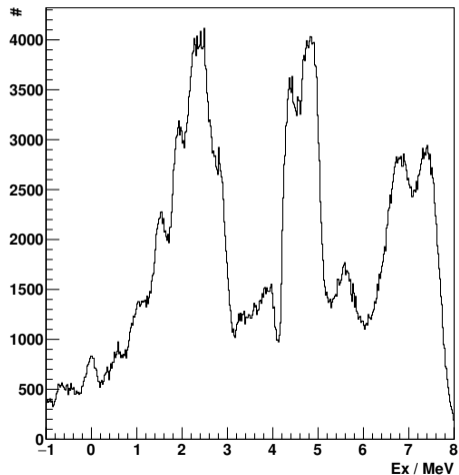
Raw Data



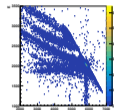
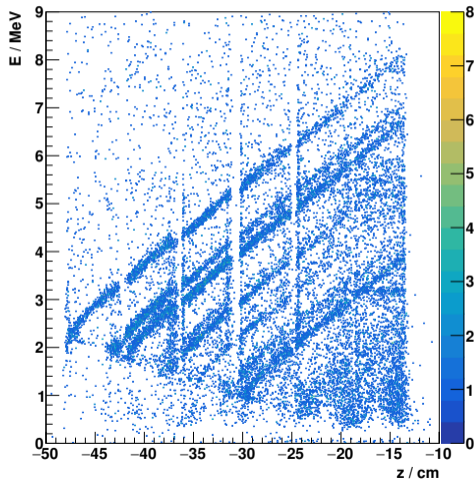
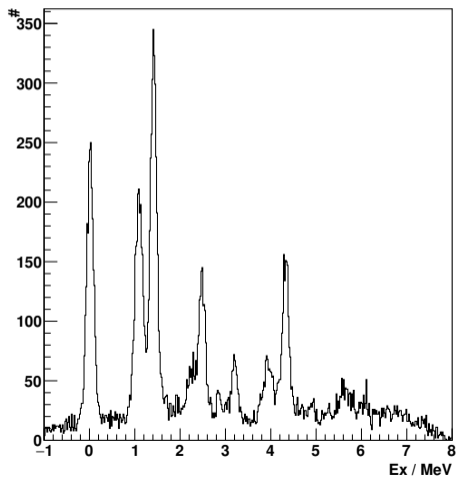
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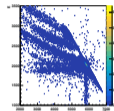
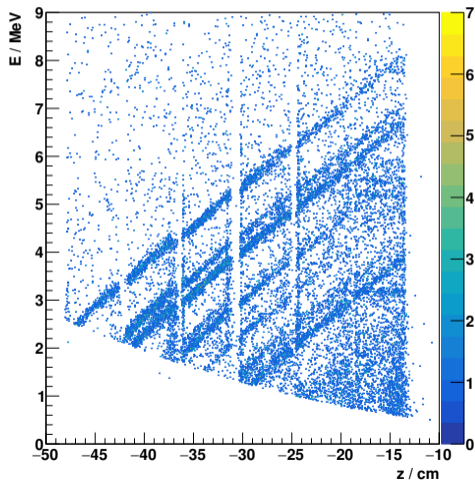
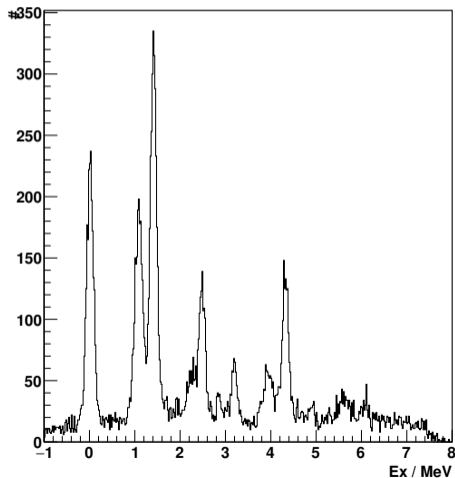
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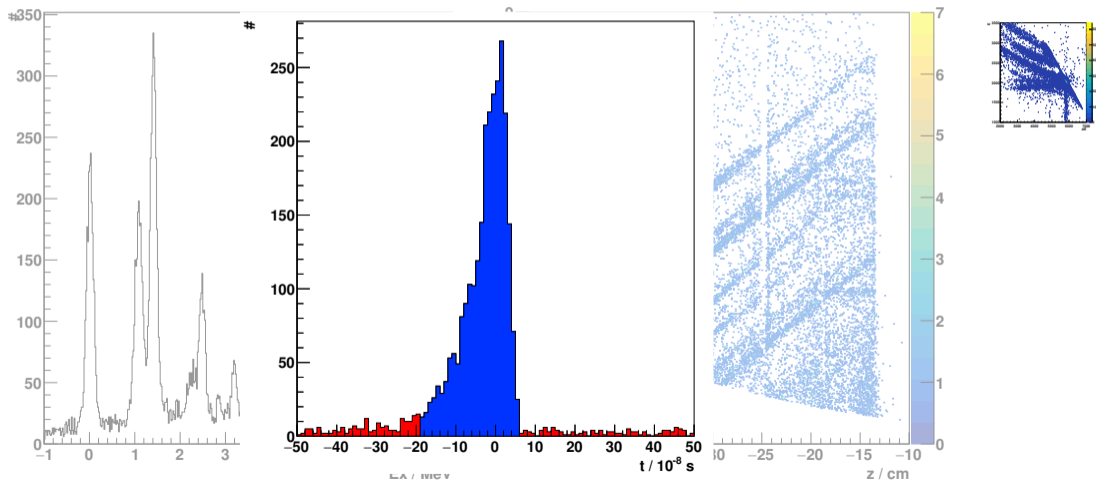
RDT Cuts



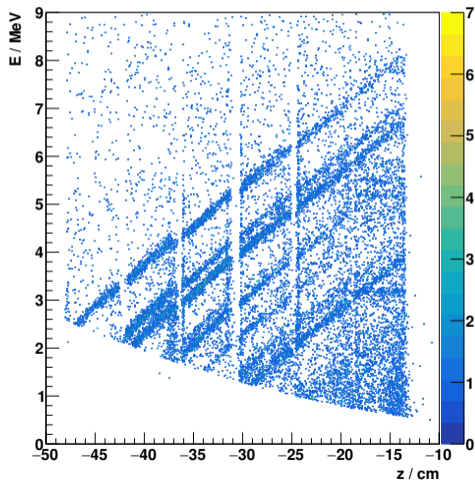
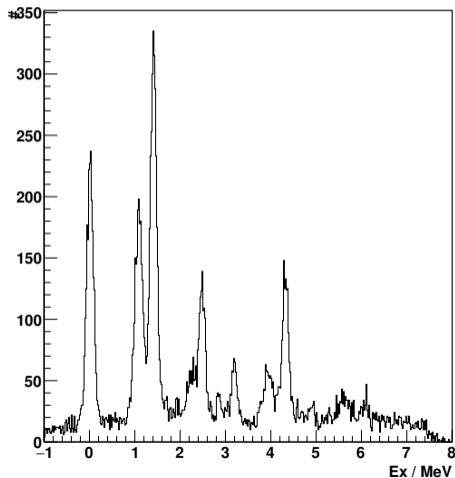
RDT + CM Angle Cuts



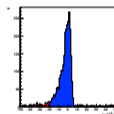
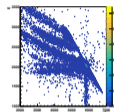
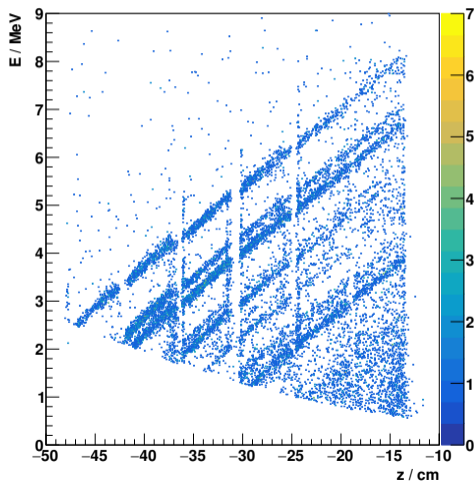
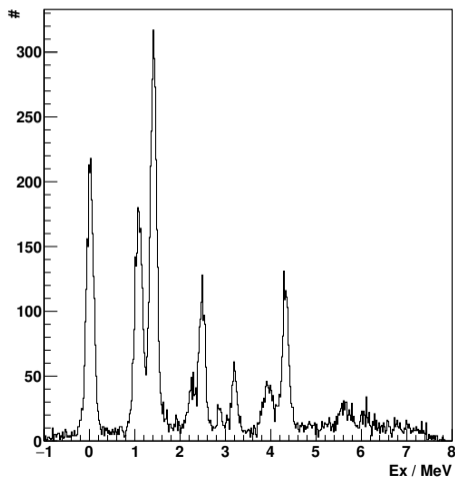
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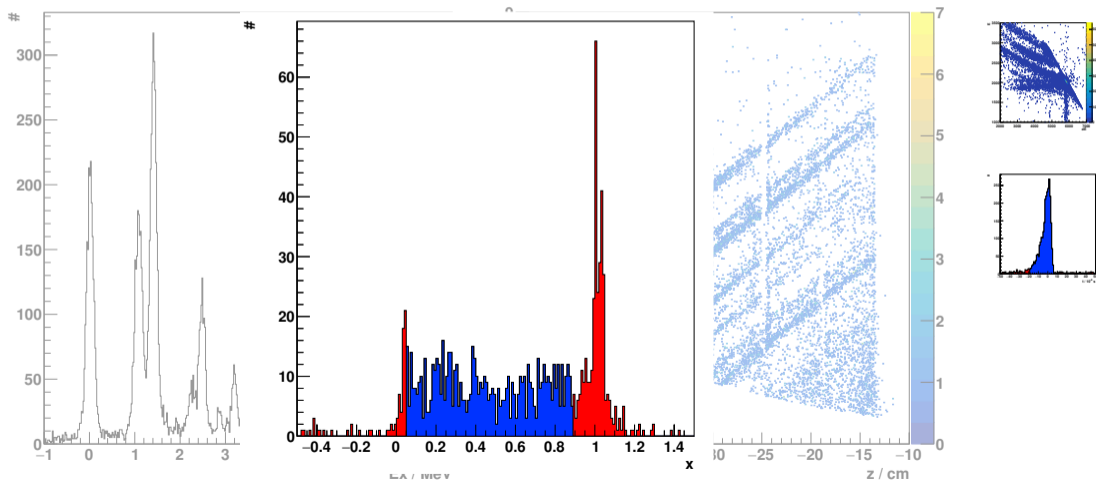
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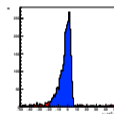
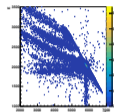
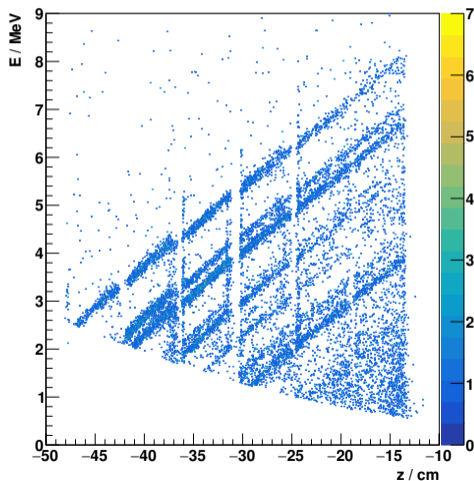
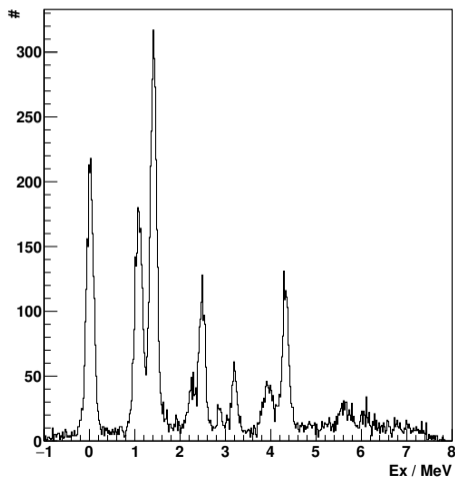
RDT + CM Angle + Timing Cuts



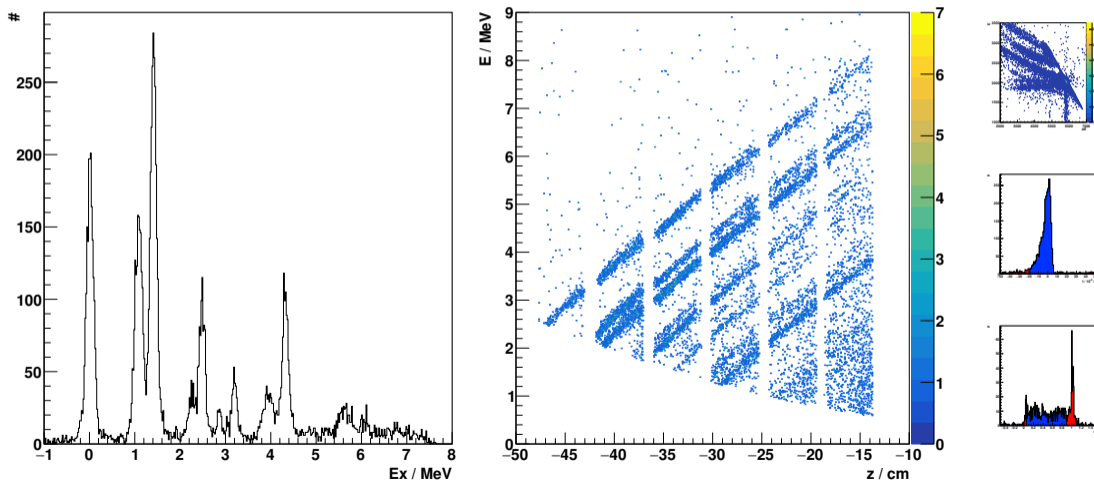
RDT + CM Angle + Timing Cuts

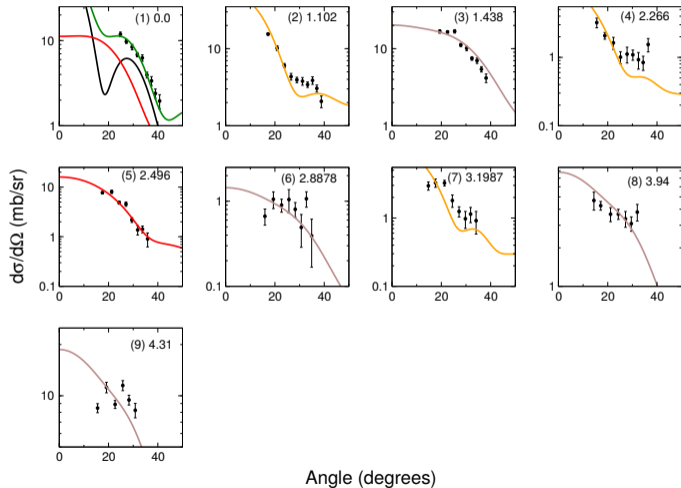


RDT + CM Angle + Timing Cuts

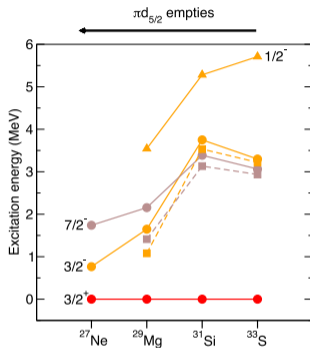
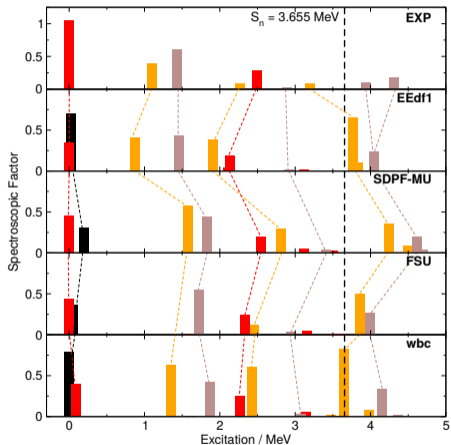


RDT + CM Angle + Timing + Si Cuts





- Angular distributions fitted for all states.
- Assigned angular momenta using DWBA calculations.
- Ground state is a $\ell = 0, 2$ doublet (total shown in green).
- Unbound states ($S_n = 3.655$ MeV) not fitting well with chosen optical potentials for DWBA.
- Spectroscopic factors extracted.



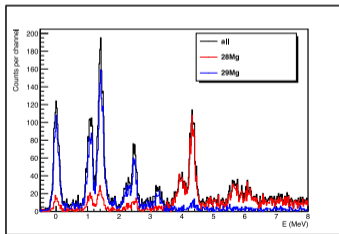
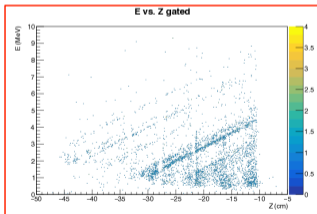
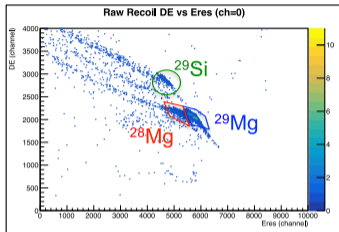
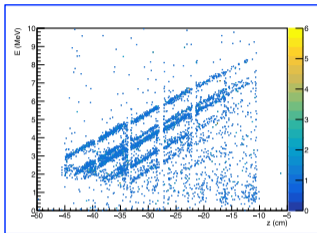
Drop in energies relative to $\nu d_{3/2}$ as $\pi d_{5/2}$ empties.

- Compared to different shell-model calculations^{4,5,6}.
- Experimental states have clear analogues in theoretical models.
- Modern calculations reproduce ordering of doublet ($\ell = 2$ g.s.)
- Generally good calculations overall.
- Seen some unbound states.

⁴ A. Matta et al. *Phys. Rev. C* 99 (4 2019), p. 044320.

⁵ R. Lubna. Priv. communication.

⁶ Y. Otsuna. Priv. communication.



7

- Recoil detector able to distinguish $^{28}\text{Mg}/^{29}\text{Mg}$.
- Look at the difference between bound and unbound states in ^{29}Mg - is there a γ -branch (unbound state decaying to bound state before neutron emitted)?
- Work ongoing...

⁷ D. K. Sharp. Private Communication.

- First experiment of ISS in early implementation stage.
- Comparable resolution to HELIOS.
- More beam than expected - requested 6×10^5 pps and got 1×10^6 pps. Allowed measurements at two array positions.
- Energy resolution of excitation not as good as hoped. Many possible reasons for this, e.g. beam properties, alignment etc. → this helps us learn about beam delivery into ISS for future runs.
- Highest energy per nucleon in a HIE-ISOLDE radioactive beam experiment.
- Data looks promising - in final stages of analysis before publication.
- A successful experiment!

Acknowledgements

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Peter Butler
Riccardo Raabe
Jiecheng Yang
Ismael Martel

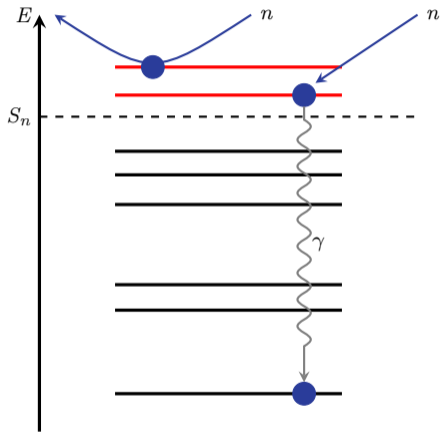
Joonas Konki
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Wilton Catford
Marc Labiche
Francesco Recchia
Oleksii Poleshchuk
Giacomo de Angelis



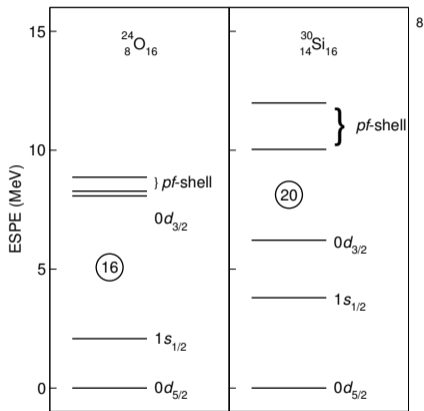
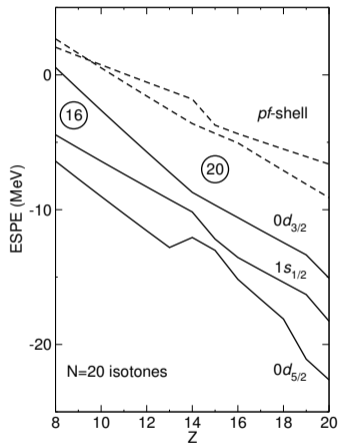
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- Neutrons above separation energy form unbound states.
- Neutrons forming unbound states close to S_n have *potential* to γ -decay to a bound state — is called a γ -branch.
- No direct measure of γ -rays in ISS.
- This information can be extracted using Si recoil and array: detect protons corresponding to states above S_n , and detect recoiling ^{29}Mg .



$$j_{>} = \ell + \frac{1}{2} = \pi d_{5/2}$$

$$j_{<} = \ell - \frac{1}{2} = \nu d_{3/2}$$

⁸ Adapted from Otsuka,

