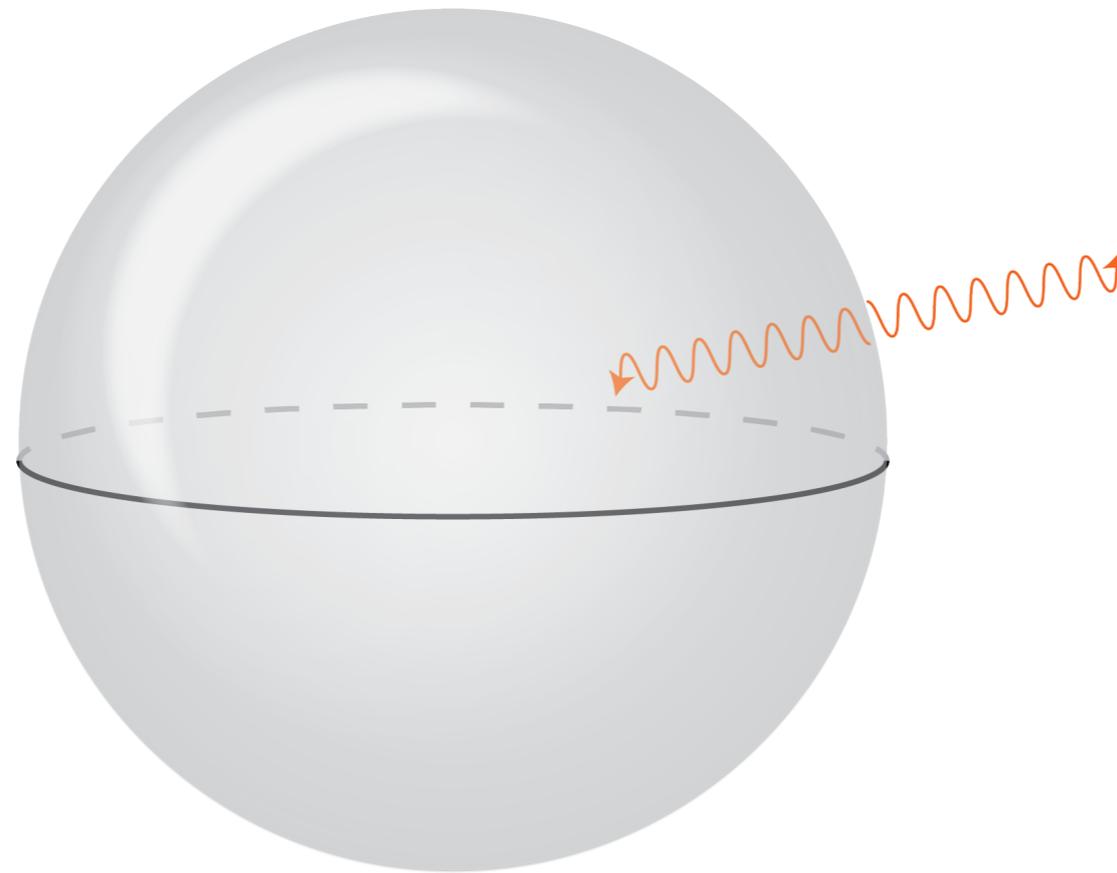


The Price of Curiosity:

Information Recovery in de Sitter Space



Based on arXiv:2104.00006 (*accepted in JHEP*) with W. Sybesma



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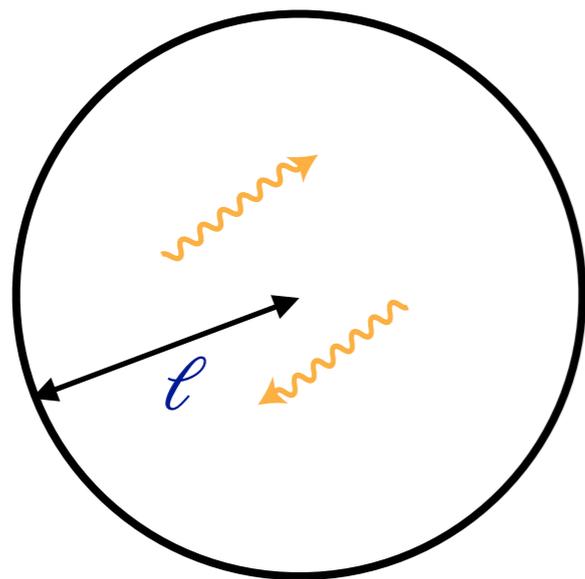
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De Sitter Space

Observations suggest that the accelerated expansion of the universe is driven by a **positive cosmological constant**.

The corresponding solution to the vacuum Einstein's equations is **de Sitter space**.

In de Sitter space, there is a **cosmological horizon**.



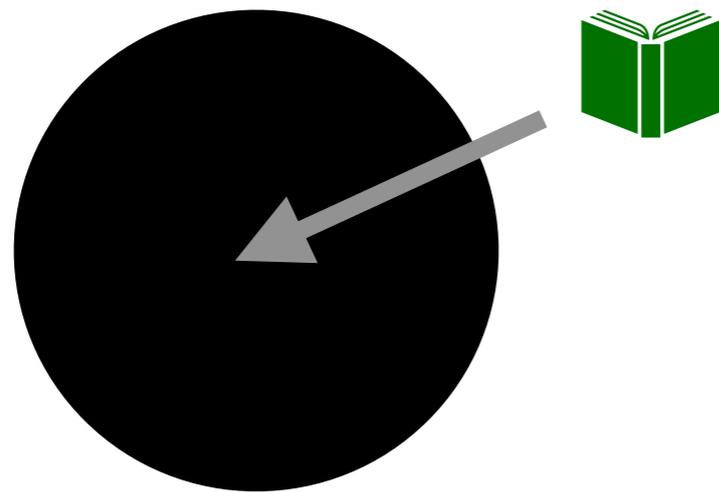
Taking into account (small) quantum effects, the horizon has thermodynamic properties.

Shares similarities with a black hole horizon, such as **temperature** and **entropy**.

$$\beta = 2\pi\ell \quad S = \frac{\text{Area}}{4G_N}$$

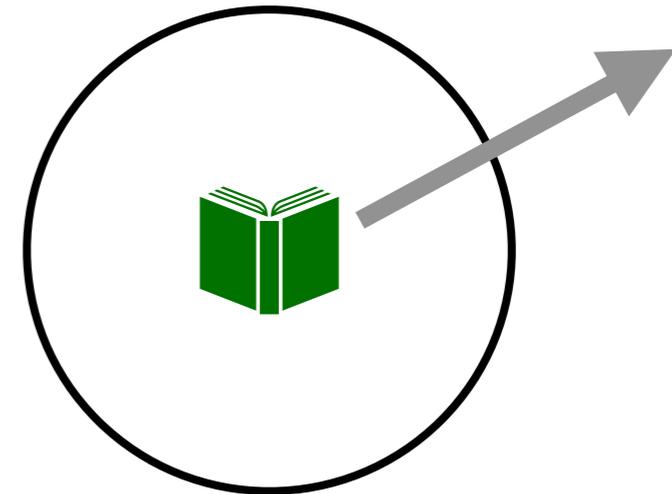
Quantum Information

Information that falls through a horizon is lost (classically).



Black Hole

↔
“Inverse”



De Sitter Space

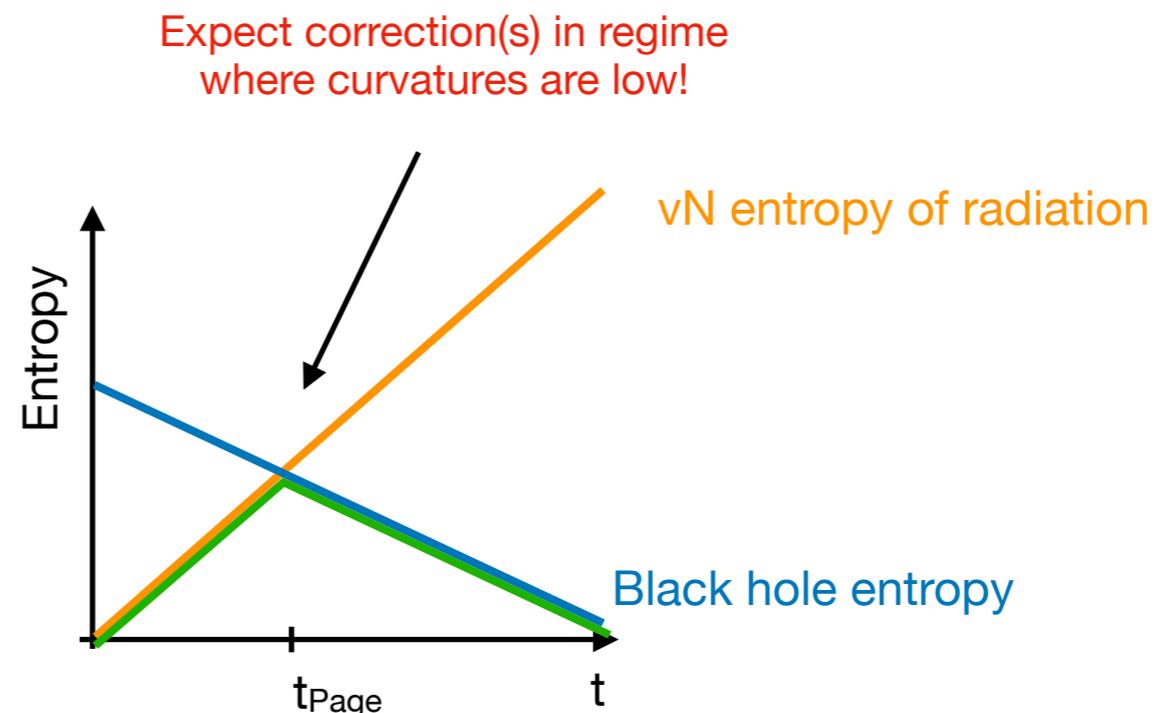
Quantum mechanically, **unitary evolution** implies that information is preserved. It should be encoded in the Hawking radiation. [Page '93]

Information Paradox

Precisely how information can be recovered was unknown until recently, this is the **information paradox**.

Quantum information is captured by the **von Neumann/entanglement entropy**.

Unitarity:
$$S_{\text{Rad}}(t) \leq S_{\text{BH}}(t)$$



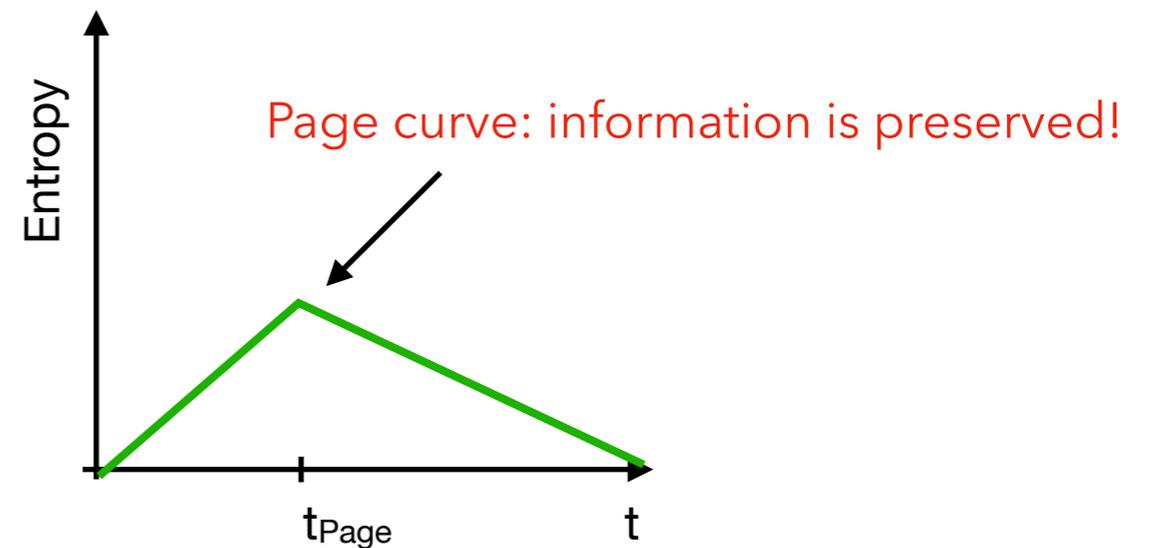
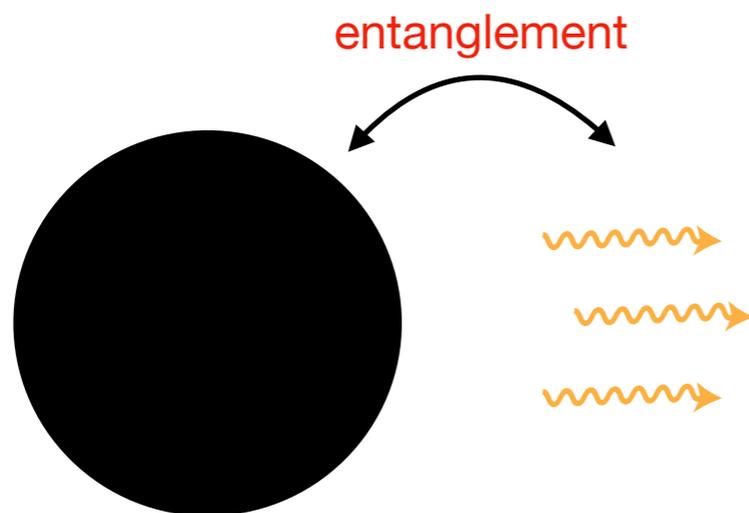
Recently, the required corrections have been found resulting in a **unitary Page curve**. [Almheiri, Hartman, Engelhardt, Maldacena, Marolf, Maxfield, Penington, Shaghoulian, Tajdini + many more works '19-'21]

Entanglement Islands

The entropy of a **gravitational systems** is given by:

$$S_{\text{gen}} = \frac{A}{4G_N} + S_{\text{vN}}$$

When there is a large amount of entanglement, the generalized entropy is given by the **"island formula"**.

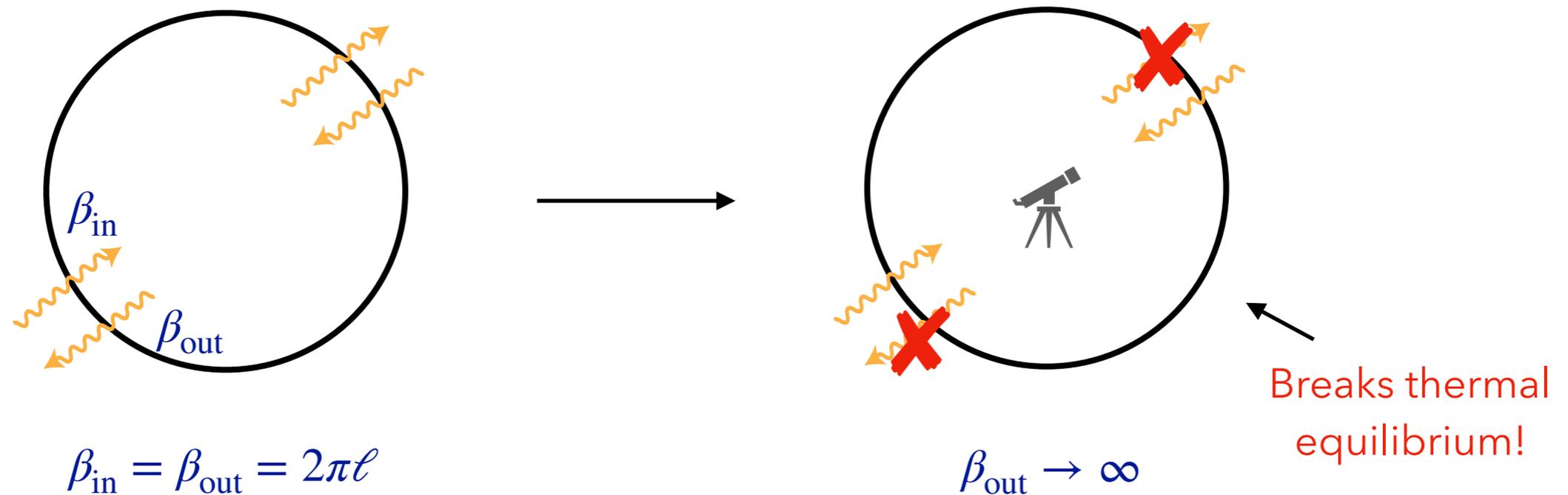


$$S_{\text{Rad}} = \frac{A_{\text{Island}}}{4G_N} + (S_{\text{Island}}^{\text{vN}} \cup S_{\text{Rad}}^{\text{vN}}) \leftarrow \text{island found by extremizing}$$

Info Recovery in dS

[LA, Parikh, van der Schaar '19]

De Sitter does not “evaporate”, but an observer can collect radiation to try to recover information.



Explicit expression for **entropy** and **backreaction** are difficult to compute in general.

We consider a simplified model: **two-dimensional JT gravity**. [Hartman,

Jiang, Shaghoulian '20] [Chen, Gorbenko, Maldacena '20] [Chen, Gorbenko, Maldacena '20]

[Balasubramanian, Kar, Ugajin '20][Sybesma '20][Geng, Nomura, Sun '21]

JT Gravity in dS_2

Jackiw-Teitelboim gravity is a two-dimensional dilaton model of gravity. Obtained as reduction of four-dimensional black hole in dS.

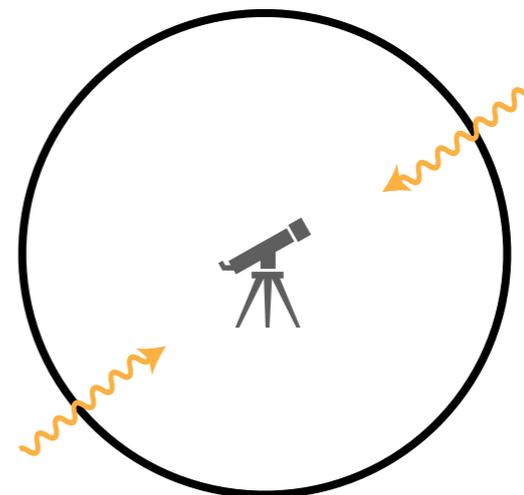
$$I = I_0(\Phi_0) + \int d^2x \Phi \left(R - \frac{2}{\ell^2} \right) + I_{\text{CFT}}$$

Constant entropy Dynamics Large c matter sector

$$\text{EOM: } \Phi g_{ab} - \ell^2 \nabla_a \nabla_b \Phi + \ell^2 g_{ab} \square \Phi = \pi \ell^2 \langle T_{ab} \rangle$$

$$\langle T_{\text{in}} \rangle = \frac{\pi c}{12\beta_{\text{in}}^2}$$

$$\langle T_{\text{out}} \rangle = 0$$



Islands in JT Gravity

[LA, Sybesma '21]

The solution to the EOMs in the state where an observer collects radiation is:

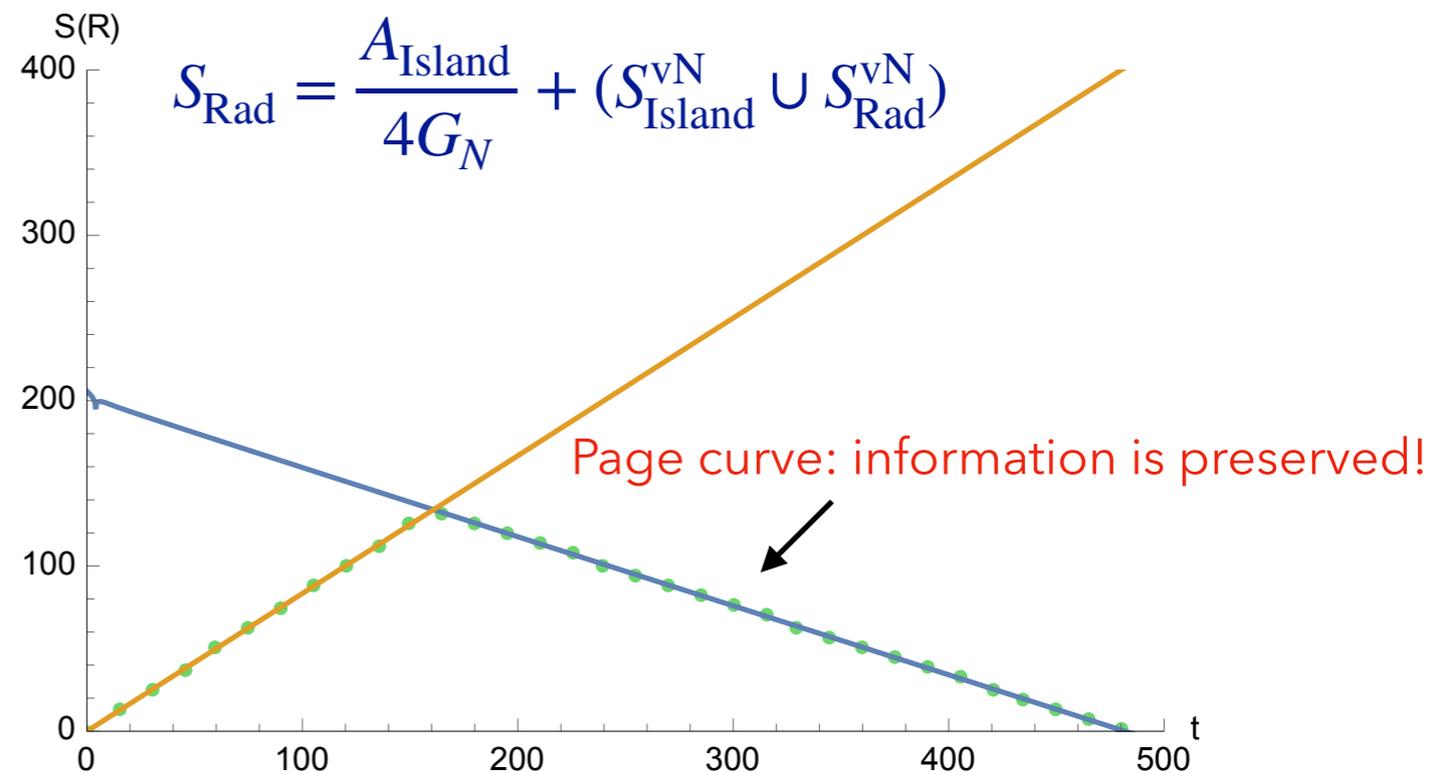
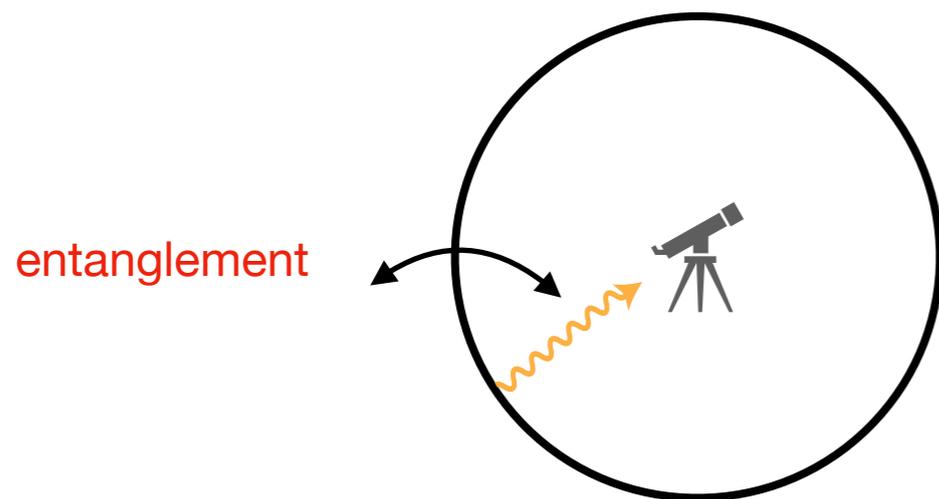
Metric:

$$ds^2 = - (1 - r^2/\ell^2)dt^2 + (1 - r^2/\ell^2)^{-1}dr^2$$

Dilaton:

$$\Phi = -\frac{cr}{\ell^2}t + \dots$$

Using the **island formula**, we find that islands contribute!



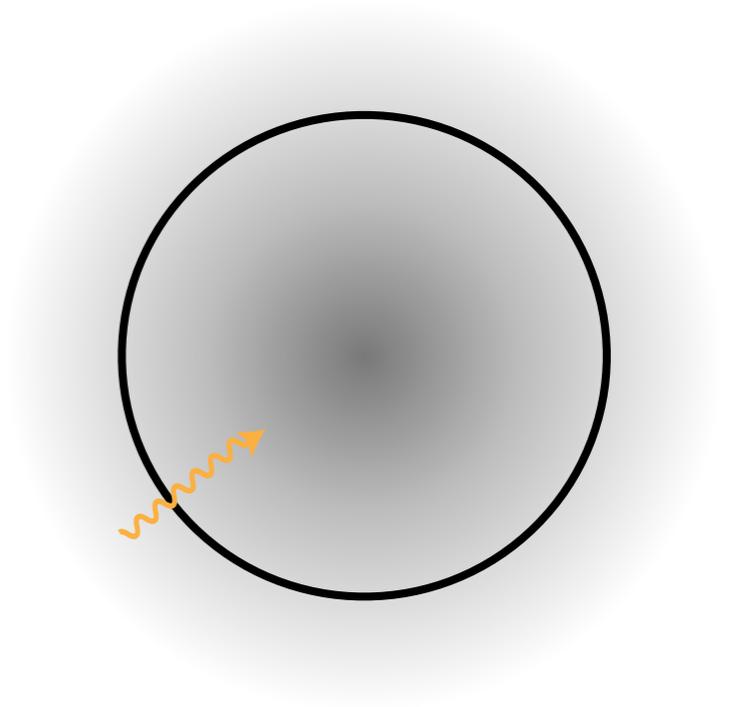
The Price of Curiosity

Contrary to a black hole, all radiation has to be stored in a **finite volume**. This will (eventually) lead to large backreaction.

Look for singularities using a **quantum singularity theorem**. [Wall '10] (see also [Freivogel, Kontou, Krommydas '20])

When information recovery is possible, formation of a singularity is **unavoidable**.

Information recovery is possible, but comes at a price!



Conclusions

- Using recent developments in black hole physics, we studied **information recovery** in de Sitter space.
- For analytical results, we focused on **JT gravity**.
- Computing the entropy of radiation, we found that **information is not lost**.
- Curiosity has a pricetag: a **singularity** forms.
- Focused on a static observer, consequences for **inflation**?

Thank you!

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