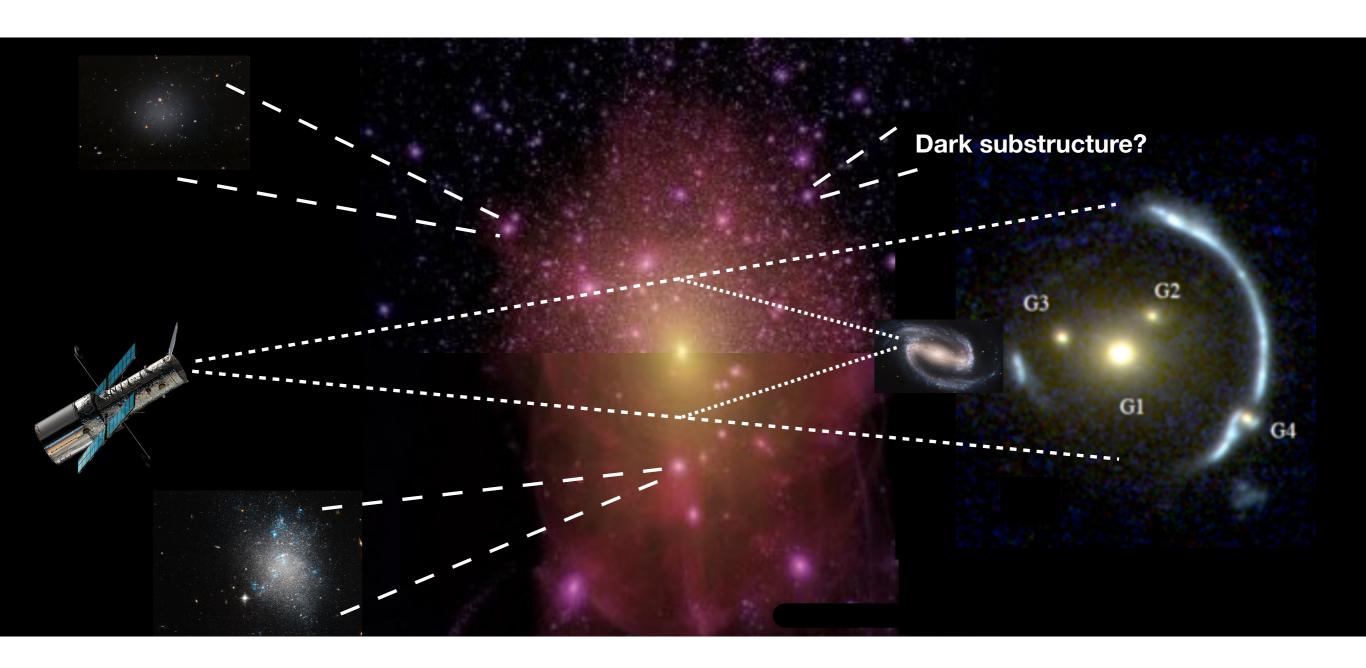
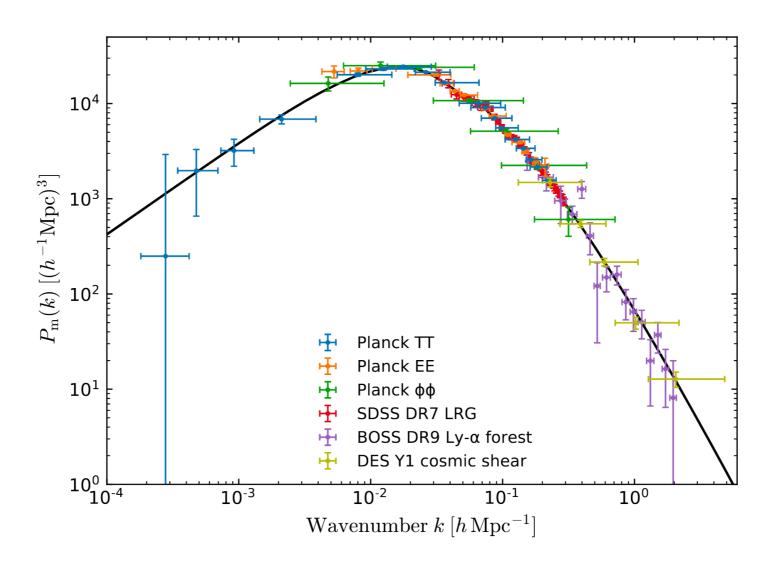
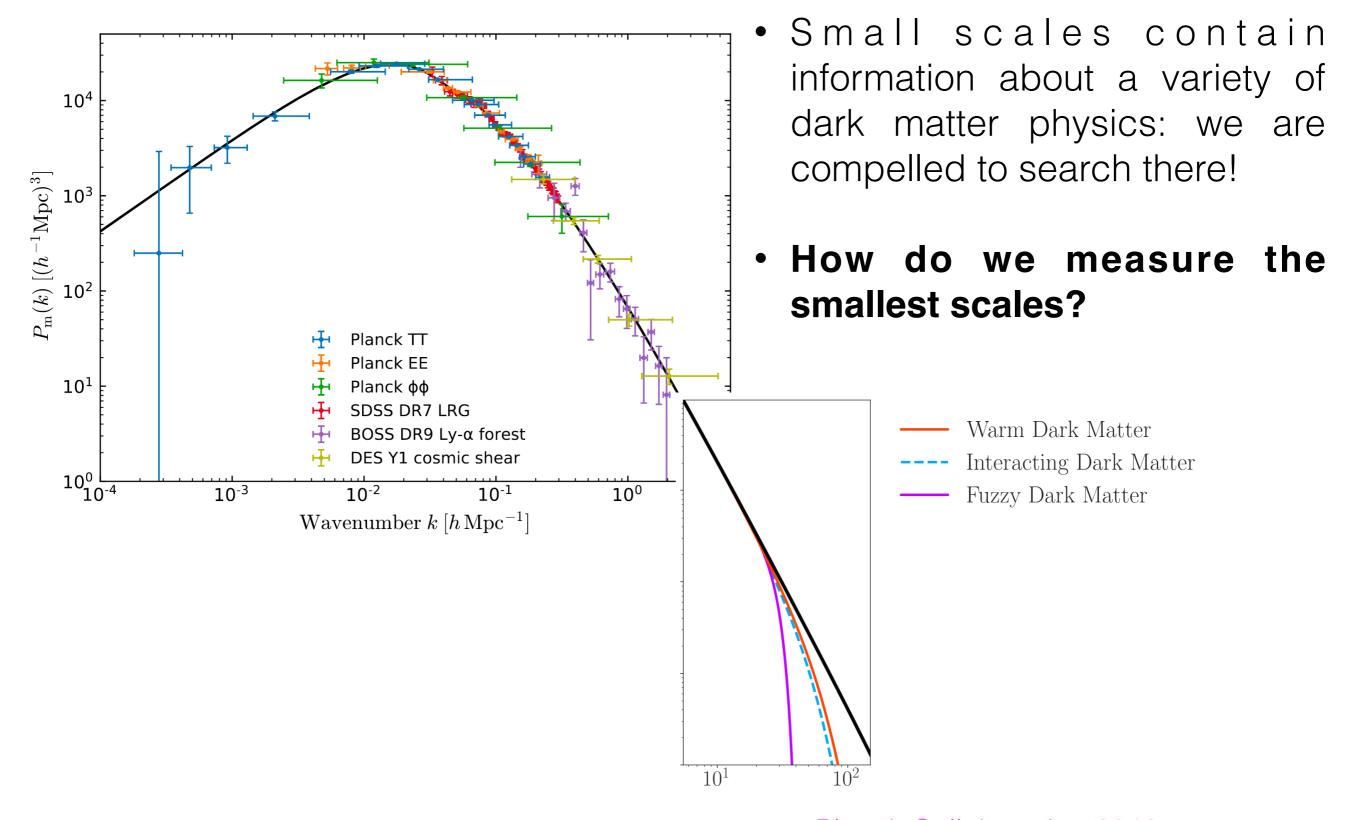


# Probing Dark Matter and Dark Energy with strong gravitational lensing



## Dark Matter and Structure Formation

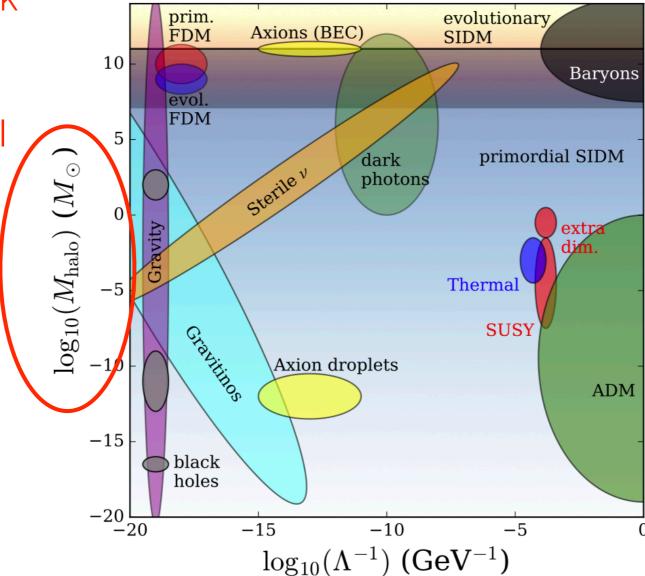




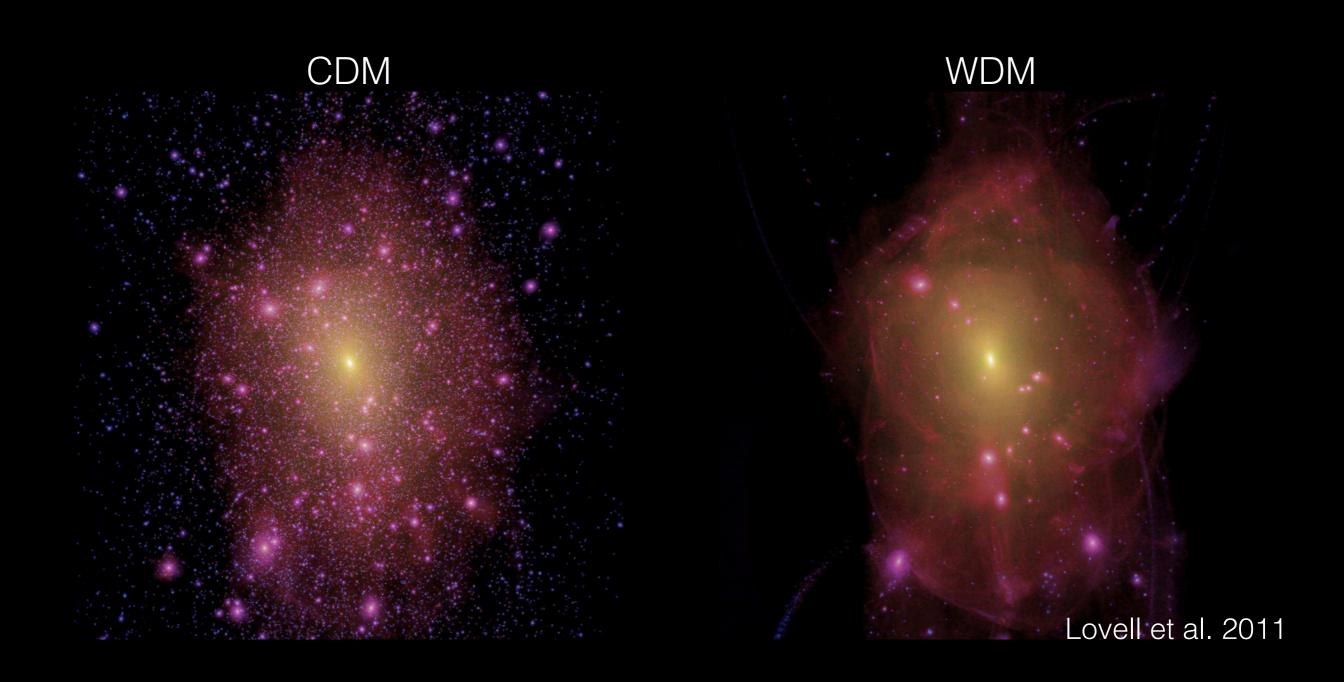
Planck Collaboration 2019 adopted by E. Nadler A. Drlica-Wagner

### Dark Matter and Structure Formation

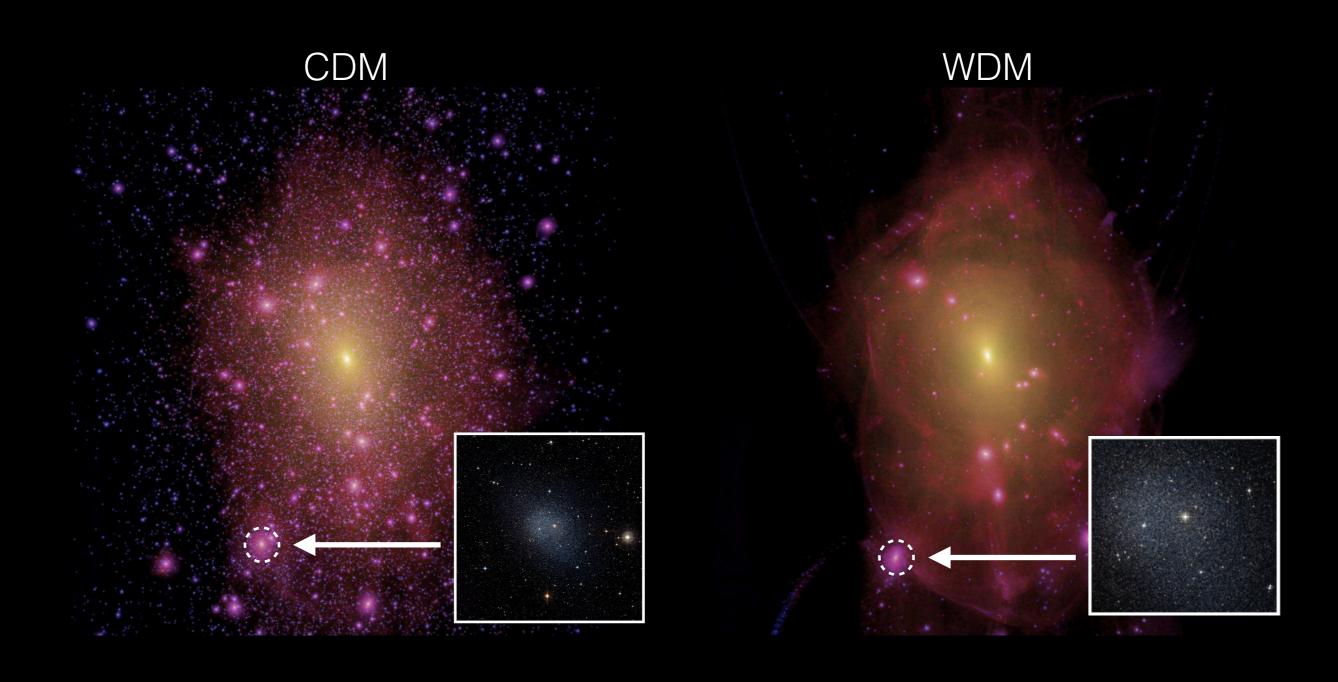
Microphysical dark matter properties affect structure formation on small scales



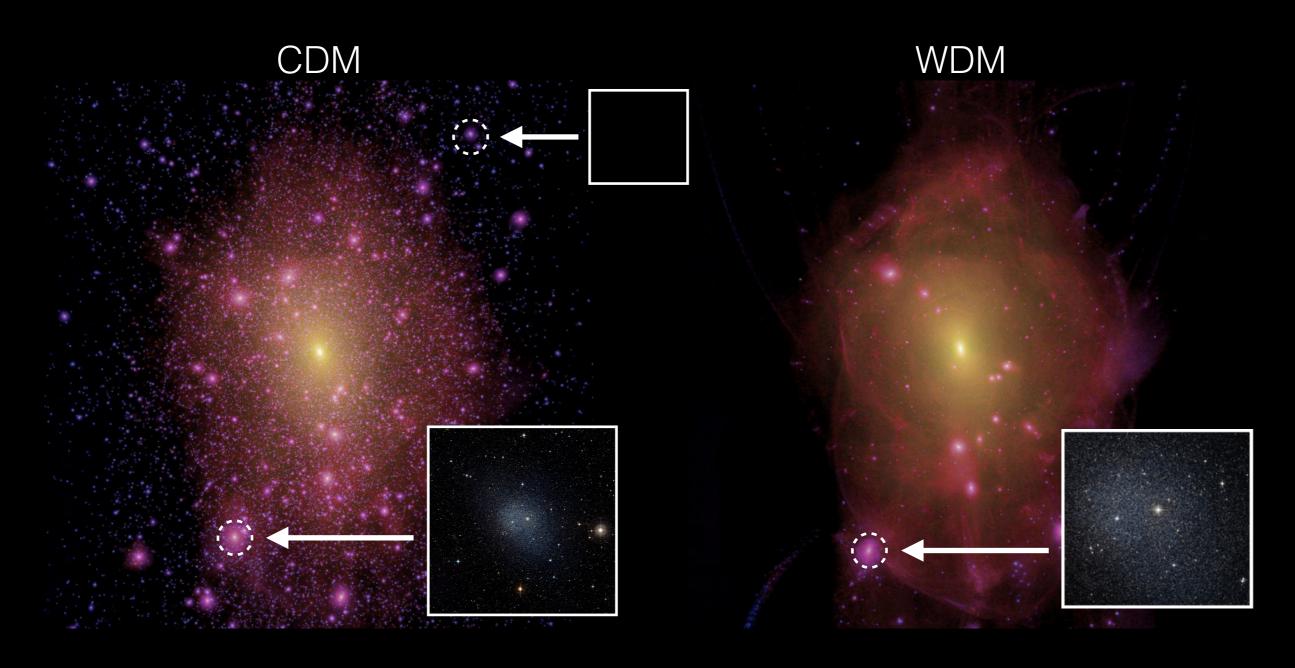
## Probing Dark Matter



## Probing Dark Matter



## What if structure is completely dark?



CDM predicts completely dark halos!

# Strong gravitational lensing

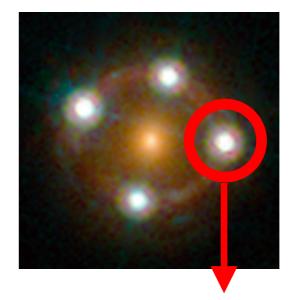
## Strong gravitational lensing

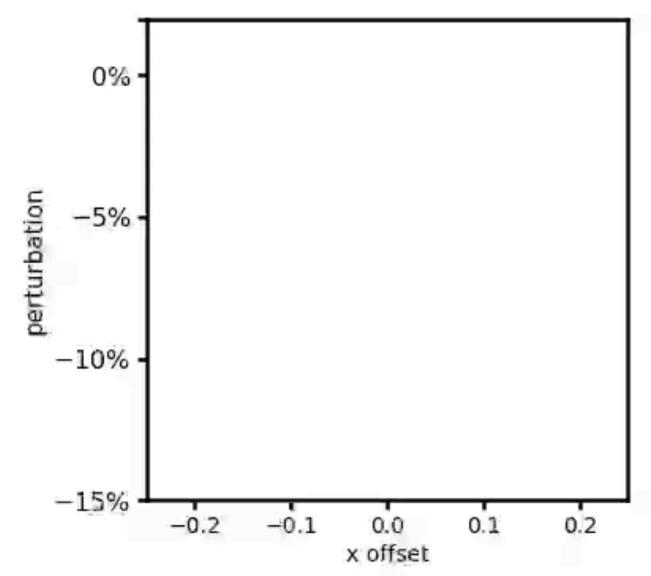
Y. Hezaveh

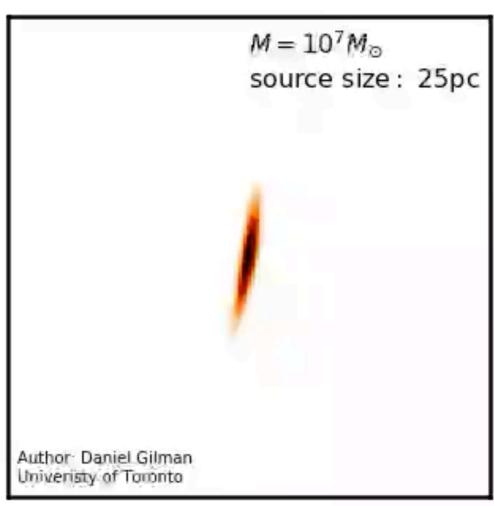
### Dark Matter and Structure Formation

- Dark matter as a collision-less cold fluid is extremely successful in describing cosmological observations
- The sub-galactic scales offer a laboratory for micro-physical properties of dark matter
- Below the galaxy-formation limit, only gravitational probes are sensitive to dark matter physics, in particular strong gravitational lensing

# Flux-ratios are sensitive to completely dark structure



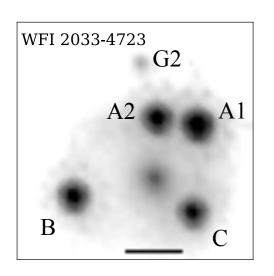


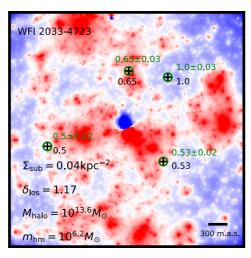


Mao & Schneider 1998, Dalal & Kochanek 2002, Moustakas & Metcalf 2003, Nierenberg+2014, 2017 Hsueh+2016, 2017, 2019, Gilman, **SB**+2018, 2019a,b,c

#### End-to-end inference of dark matter microphysics

#### Flux ratio statistics



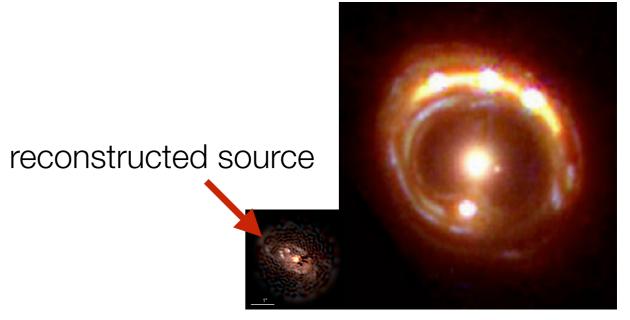


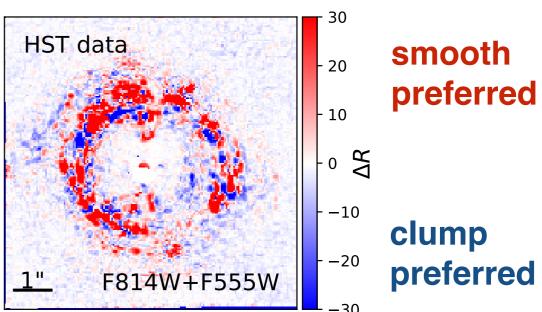
Forward modeling and simulation based inferences with Approximate Bayesian Computing

thermal relic mass > 5.2 keV from a sample of 8 quasar lenses, > 2 keV from1 lensing arc

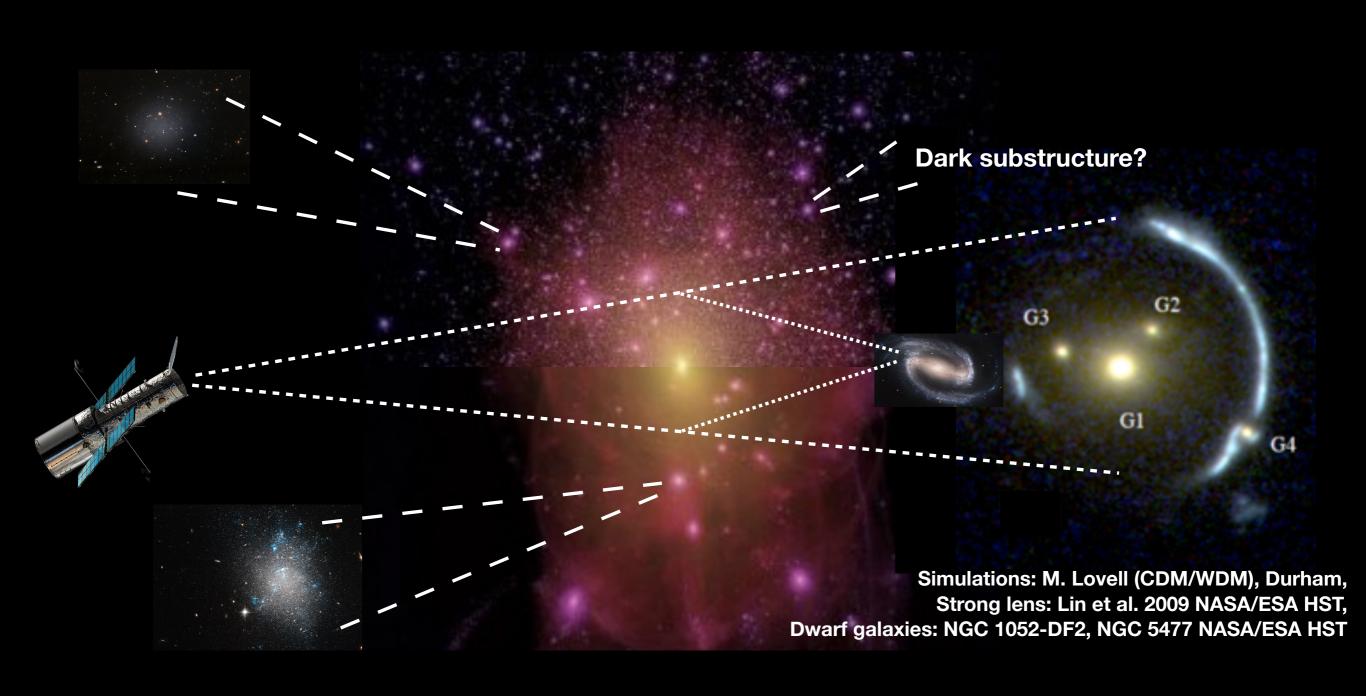
**SB+**2015, **SB+**2017a, b, Gilman, **SB+**2019, 2020a, b

#### Gravitational imaging



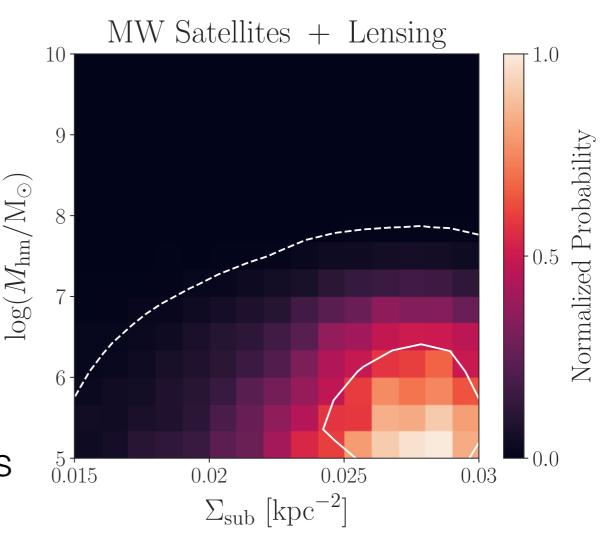


# Combining visible and invisible universe a self-consistent combined small-scale probe analysis



# Combining visible and invisible universe self-consistent combined small-scale probe analysis

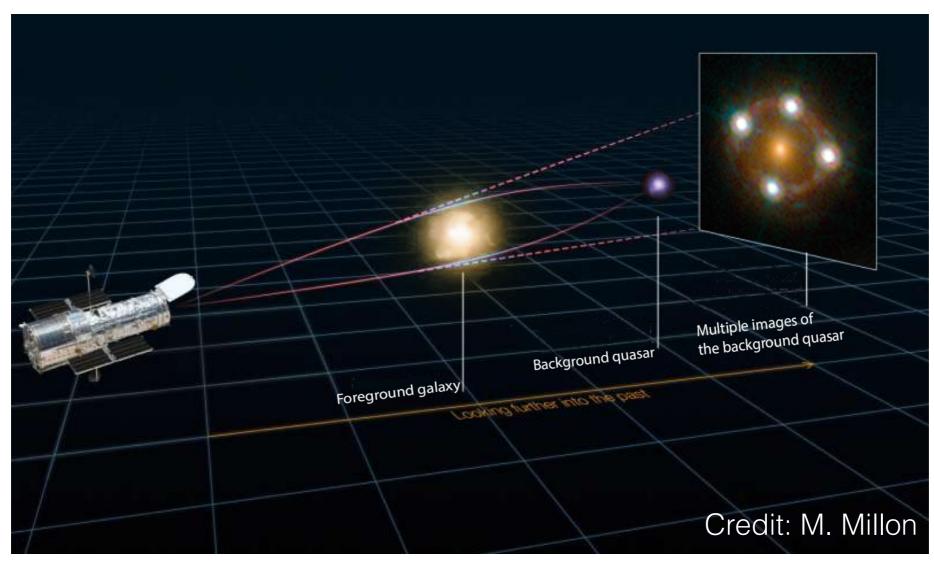
- Recent analyses of the Lyman-α forest, strong gravitational lenses, and Milky Way satellites achieve similar dark matter sensitivity
- Each individual measurement probes a **distinct aspect** of dark matter clustering
- Joint analyses of small-scale probes are key to break degeneracies and robustly detect non-CDM physics



thermal relic mass > 9.7 keV at 95% confidence

Nadler, SB, Gilman et al. arXiv:2101.07810

# Measuring the Hubble constant with time-delay cosmography



absolute scale

time delay

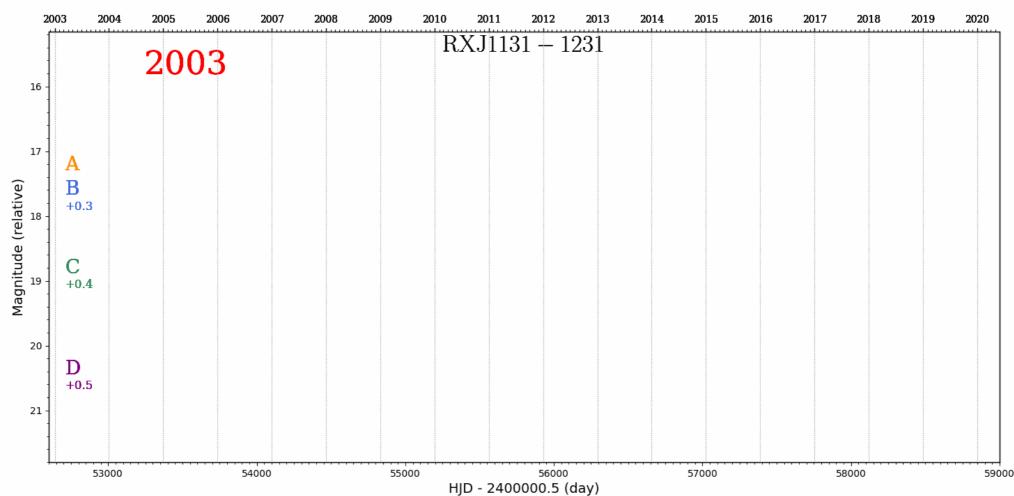
path difference

lensing potential

 $t(\theta, \beta) = \frac{(1+z_{\rm d})}{c} \frac{D_{\rm d}D_{\rm s}}{D_{\rm ds}} \left[ \frac{(\theta-\beta)^2}{2} - \psi(\theta) \right]$ 

### Measuring time delays with long-term monitoring

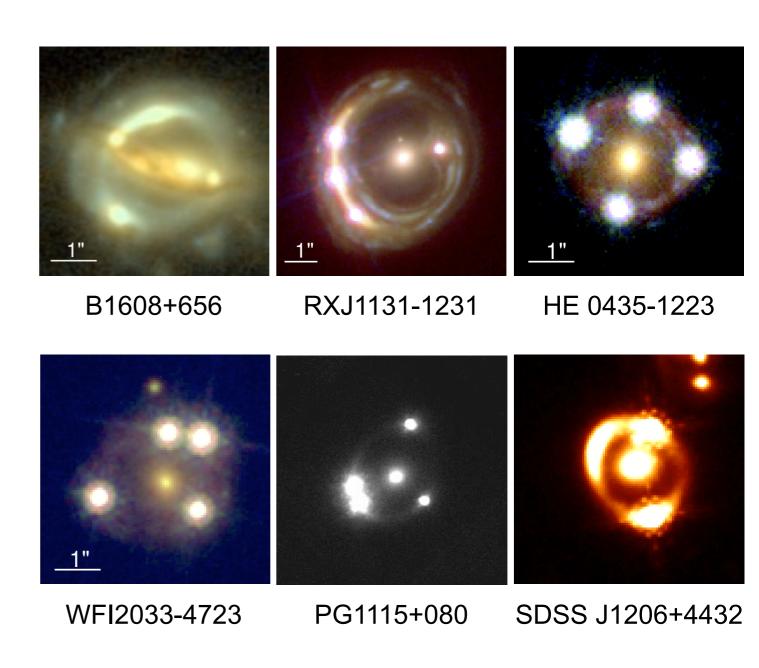




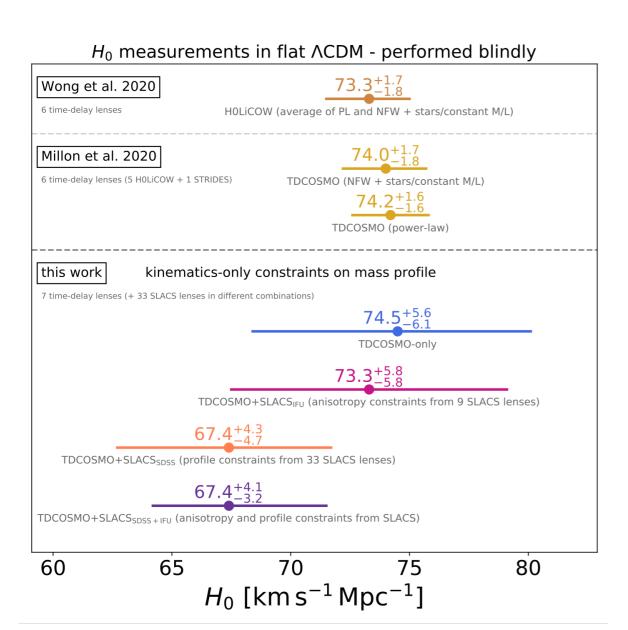
## TDCOSMO project

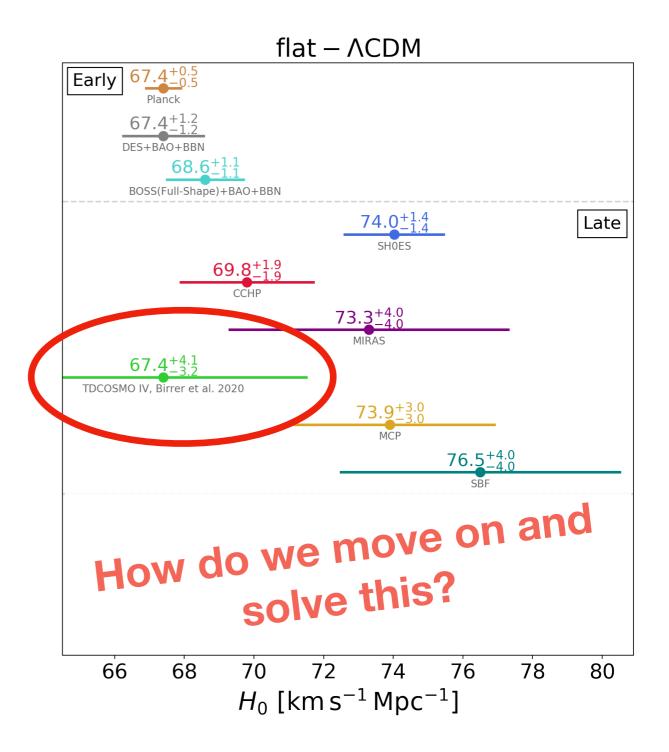
(H0LiCOW+STRIDES+SHARP+COSMOGRAIL)

- Detailed analysis of several time-delay lenses (Suyu+2017)
  - long term monitoring from COSMOGRAIL (Courbin+2011) for accurate time delays
  - high-resolution HST imaging for detailed lens modeling
  - wide-field imaging/spectroscopy to characterize mass along LOS
- Goal is to constrain H<sub>0</sub> to ~few % precision
- Seven lenses have been analyzed (Suyu+2010, 2013; Wong+2017, Birrer+2019, Rusu+2019, Chen+2019, Shajib+2019), more coming



## H0 - present





SB et al. 2020 TDCOSMO IV

### H0 - near future - road to 1%

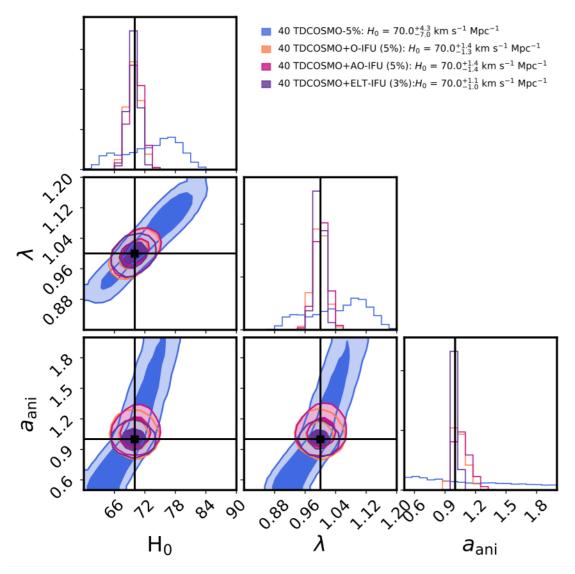
SN 'Refsdal'

7% Hn estimat

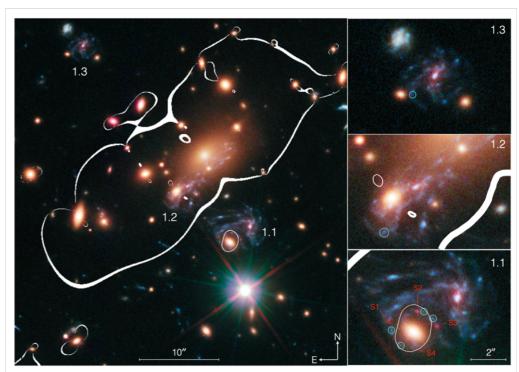
Kally et al (2015), Science 347(6226):1123=1126



(approved JWST program)



SB & Treu 2021 TDCOSMO V

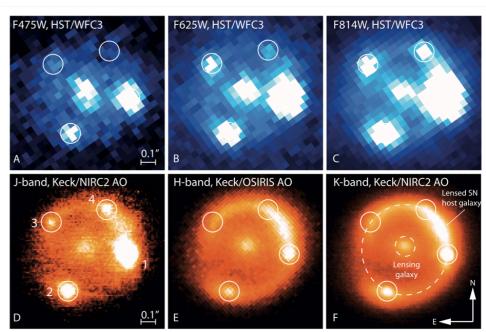


Kelly et al. 2015, Kelly (incl SB) et al. in prep

#### Lensed supernovae

iPTF16geu

idena: 356/63251:291–295



# Summary

#### **Gravitational lensing is...**



probes small (dark) matter structure



- competitive with other cosmological probes
- advancing with increased sample size and improved observational capabilities!

