

1 **Measurement of the evolution of a plasma column**
 2 **by modulation of a high-energy proton beam**

3 S. Gessner,¹ E. Adli,² A. Ahuja,¹ O. Apsimon,^{3,4} R. Apsimon,^{5,4} A.-M. Bachmann,^{1,6,7}
 4 F. Batsch,^{1,6,7} C. Bracco,¹ F. Braunmüller,⁶ S. Burger,¹ G. Burt,^{5,4} B. Buttenschön,⁸
 5 A. Caldwell,⁶ J. Chappell,⁹ E. Chevallay,¹ M. Chung,¹⁰ D. Cooke,⁹ H. Damerau,¹
 6 G. Demeter,¹¹ L.H. Deubner,¹² A. Dexter,^{5,4} S. Doebert,¹ J. Farmer,¹³
 7 V.N. Fedosseev,¹ R. Fiorito,^{14,4} R.A. Fonseca,¹⁵ F. Friebe,¹ L. Garolfi,¹ B. Goddard,¹
 8 I. Gorgisyan,¹ A.A. Gorn,^{16,17} E. Granados,¹ O. Grulke,^{8,18} E. Gschwendtner,¹
 9 A. Hartin,⁹ A. Helm,¹⁹ J.R. Henderson,^{5,4} M. Hüther,⁶ M. Ibison,^{14,4} S. Jolly,⁹
 10 F. Keeble,⁹ M.D. Kelisani,¹ S.-Y. Kim,¹⁰ F. Kraus,¹² M. Krupa,¹ T. Lefevre,¹
 11 Y. Li,^{3,4} S. Liu,²⁰ N. Lopes,¹⁹ K.V. Lotov,^{16,17} M. Martyanov,⁶ S. Mazzoni,¹
 12 V.A. Minakov,^{16,17} J.C. Molendijk,¹ J.T. Moody,⁶ M. Moreira,^{19,1} P. Muggli,^{6,1}
 13 H. Panuganti,¹ A. Pardons,¹ F. Peña Asmus,^{6,7} A. Perera,^{14,4} A. Petrenko,^{1,16}
 14 A. Pukhov,¹³ S. Rey,¹ P. Sherwood,⁹ L.O. Silva,¹⁹ A.P. Sosedkin,^{16,17}
 15 P.V. Tuev,^{16,17} M. Turner,¹ F. Velotti,¹ L. Verra,^{1,21} V.A. Verzilov,²⁰ J. Vieira,¹⁹
 16 C.P. Welsch,^{14,4} M. Wendt,¹ B. Williamson,^{3,4} M. Wing,⁹ B. Woolley,¹ and G. Xia^{3,4}

17 (The AWAKE Collaboration)

18 ¹*CERN, Geneva, Switzerland*

19 ²*University of Oslo, Oslo, Norway*

20 ³*University of Manchester, Manchester, UK*

21 ⁴*Cockcroft Institute, Daresbury, UK*

22 ⁵*Lancaster University, Lancaster, UK*

23 ⁶*Max Planck Institute for Physics, Munich, Germany*

24 ⁷*Technical University Munich, Munich, Germany*

25 ⁸*Max Planck Institute for Plasma Physics, Greifswald, Germany*

26 ⁹*UCL, London, UK*

27 ¹⁰*UNIST, Ulsan, Republic of Korea*

28 ¹¹*Wigner Research Center for Physics, Budapest, Hungary*

29 ¹²*Philipps-Universität Marburg, Marburg, Germany*

30 ¹³*Heinrich-Heine-University of Düsseldorf, Düsseldorf, Germany*

31 ¹⁴*University of Liverpool, Liverpool, UK*

32 ¹⁵*ISCTE - Instituto Universitário de Lisboa, Portugal*

33 ¹⁶*Budker Institute of Nuclear Physics SB RAS, Novosibirsk, Russia*

34 ¹⁷*Novosibirsk State University, Novosibirsk, Russia*

35 ¹⁸*Technical University of Denmark, Lyngby, Denmark*

36 ¹⁹*GoLP/Instituto de Plasmas e Fusão Nuclear,*
37 *Instituto Superior Técnico, Universidade de Lisboa, Lisbon, Portugal*

38 ²⁰*TRIUMF, Vancouver, Canada*

39 ²¹*University of Milan, Milan, Italy*

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41 **Abstract**

42 The AWAKE experiment at CERN is a proof-of-concept proton beam-driven plasma wakefield
43 accelerator. The plasma source is a 10 meter-long Rubidium vapor cell ionized by a high-intensity
44 laser pulse. The proton beam is modulated by the plasma, which results in micro-bunching at the
45 plasma frequency and resonant excitation of the plasma wave. The modulation process is sensitive
46 to the relative timing of the laser pulse with respect to the proton beam. We measure the frequency
47 of the micro-bunching when the laser and proton beams are coincident, and when the proton beam
48 propagates through the plasma long after the passage of the laser pulse. The plasma density is
49 inferred from the frequency of modulation of the proton beam. We observe a microsecond-long
50 plateau in the plasma density, followed by a power-law decay of the density versus time. Finally, we
51 discuss the implications of this measurement for high-repetition rate plasma wakefield accelerators
52 based on the seeded self-modulation concept.

53 I. INTRODUCTION

54 The Advanced Wakefield Experiment (AWAKE) at CERN is the first plasma wakefield
55 acceleration (PWFA) experiment to use a high-energy proton beam driver [1–4]. The 400
56 GeV proton beam from the Super Proton Synchrotron (SPS) accelerator contains approx-
57 imately 3×10^{11} particles per bunch, and the total stored energy in the beam is nearly 20
58 kJ, three orders of magnitude greater than the energy of electron beam drivers [5] and laser
59 pulse drivers [6] used in state-of-art plasma acceleration experiments. The large amount of
60 energy stored in the proton beam has the potential to simplify the design of a plasma wake-
61 field accelerator: a trailing beam of electrons can be accelerated to extremely high energies
62 in a single plasma stage, eliminating the need for complicated staging mechanisms [7–9].

63 While the SPS proton bunch delivers an unprecedented amount of energy for a PWFA
64 experiment, it also poses a novel challenge in that the bunch length σ_z is much longer than
65 the plasma wavelength λ_p . In typical PWFA experiments, the drive bunch is shorter than
66 the plasma wavelength and can drive a high-amplitude wakefield as soon as it enters the
67 plasma, but the proton beam at AWAKE is too long to effectively couple to the plasma
68 wave. However, the proton beam excites a self-modulation instability (SMI) as it passes
69 through the plasma, which causes the beam to form micro-bunches separated by the plasma
70 wavelength [10–12]. The micro-bunches act together to resonantly drive a high-amplitude
71 wakefield.

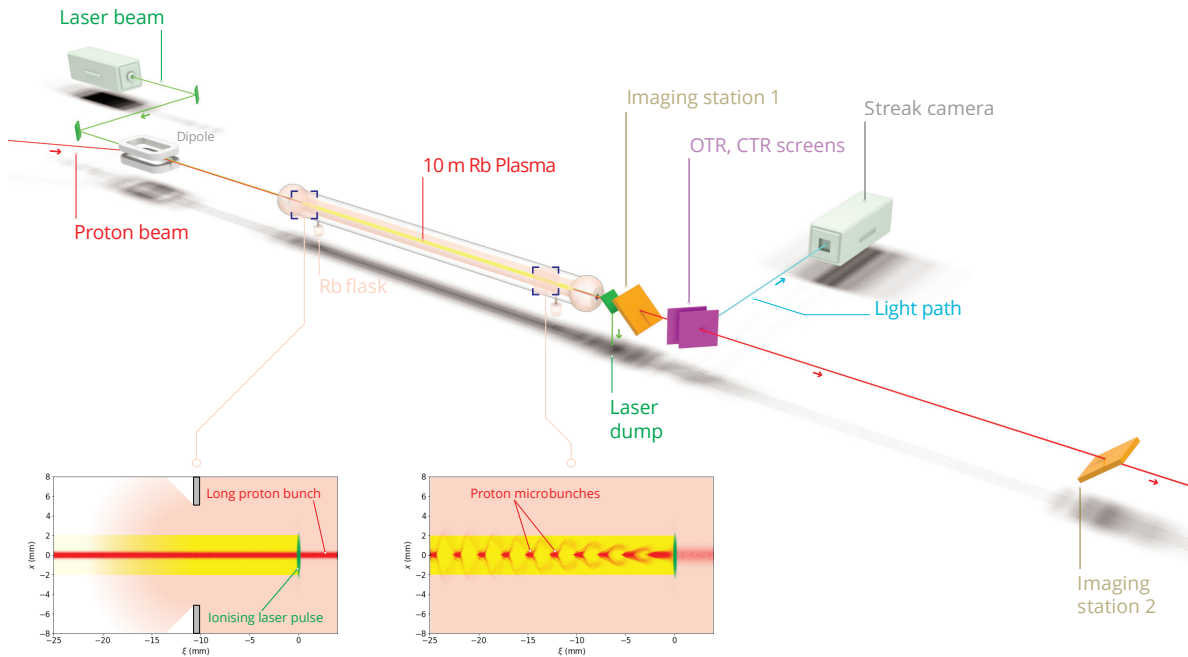
72 In order to control the onset of SMI, we use a high-power laser pulse to seed the instability.
73 This process is referred to as seeded self-modulation (SSM) [13]. The seed pulse passes
74 through and ionizes a rubidium (Rb) vapor source coincident with the passage of the proton
75 beam. The seed pulse is temporally centered within the proton beam, such that the head of
76 the beam transits through vapor and the tail transits through plasma. As the SSM develops,
77 the micro-bunches and the plasma wave that they excite have a fixed phase with respect to
78 the seed pulse. The SSM phenomenon was recently demonstrated in a series of experiments
79 at AWAKE [14, 15].

80 The onset of micro-bunching is sensitive to the timing of the laser pulse with respect to
81 the proton beam. If there is ionization of the Rb vapor ahead of the seed pulse, for instance
82 caused by a laser pre-pulse or residual plasma from the previous shot, the head of the proton
83 beam will also travel through plasma and excite a wake with a random phase with respect

84 to the seed pulse. This wake competes with the wake created through the SSM process
 85 and should be suppressed. In addition, a deleterious transverse hose instability may arise
 86 in the case of un-seeded SMI [16]. It is therefore important to understand the degree to
 87 which a pre-ionized plasma affects the proton beam and how long the plasma persists after
 88 ionization, as this will influence the repetition rate of a plasma accelerator based on the
 89 SSM mechanism.

90 In this work, we report on the lifetime and evolution of the plasma channel by inferring
 91 the plasma density from the frequency of a self-modulated proton bunch.

92 II. EXPERIMENTAL OVERVIEW



93

94 FIG. 1. Experimental layout of the AWAKE experiment showing key components. Dipole magnets
 95 and optical mirrors are used to align the proton and laser beams upstream of the vapor cell.
 96 Downstream of the cell, the proton beam passes through an OTR foil and optical light is generated
 97 in the process. The light is transported to a streak camera for temporal imaging. The inset
 98 illustrates a uniform, gaussian proton bunch entering the plasma which is micro-bunched by the
 99 time it exits the plasma.

100 The 400 GeV proton beam from the SPS is delivered to the AWAKE experiment via an 800
101 m beamline. For the dataset considered here, the proton beam contained $3.03 \pm 0.16 \times 10^{11}$
102 particles per bunch and a bunch length $\sigma_z = 8.24 \pm 0.15$ cm ($\sigma_t = 275 \pm 5$ ps). The
103 transverse emittance was approximately 3.5 mm·mrad in both planes. The beam is focused
104 to a transverse spot size $\sigma_r \approx 200$ μ m at the entrance of the vapor cell. A sketch of the
105 experimental layout highlighting the key components is shown in Figure 1.

106 The core of the AWAKE experiment is the 10 meter long Rb vapor cell [17, 18]. Rb vapor
107 flows into the cell from reservoirs at either end of the chamber. The heating of Rb reservoirs
108 is adjusted until the vapor density is the same at both ends of the cell. The uniform heating
109 of the cell implies that if the density is the same at both ends, it is uniform throughout. There
110 are diagnostic ports at both ends of the cell where a white-light laser passes through the
111 vapor. The density of the vapor is inferred from spectroscopic measurements of white-light
112 absorption at the 780 nm and 795 nm absorption lines, with an uncertainty of 0.5% [19, 20].
113 The temperature of the cell was held at 468 K for this measurement.

114 The Rb vapor is ionized by a terawatt-class Ti:sapphire laser. The laser pulse energy can
115 be varied from 40 to 450 mJ and the pulse length is approximately 120 fs [21]. The laser
116 is focused into the vapor cell with a Rayleigh length of 15 m, resulting in a laser spot size
117 of 2 mm over the entire length of the cell, much larger than the 200 μ m proton beam spot
118 size. The main laser pulse is derived from an 88 MHz oscillator. The oscillator phase is
119 locked to the RF phase of the cavities in the SPS ring. This allows for synchronization of
120 the amplified laser pulse with the extracted proton beam with picosecond accuracy.

121 The principle diagnostic for the measurements described in this paper is a streak camera.
122 The proton bunch passes through a metallic foil two meters downstream of the exit of the
123 Rb cell. The beam produces optical transition radiation (OTR) as it passes through the
124 foil, and the OTR light has the same spatio-temporal pattern as the micro-bunched beam.
125 The light is sent via an optical transport system to a dark room where it is imaged onto the
126 aperture of a streak camera. Inside the streak camera, the OTR photons are converted into
127 electrons by a photocathode and accelerated through a “streak tube” towards a phosphor
128 screen [22]. A time-varying, transverse voltage is applied to the streak tube such that the
129 electrons receive a kick that depends on when they were produced at the photocathode. The
130 electrons arrive at different transverse positions on the phosphor screen and the emitted light
131 is imaged onto a CMOS camera. The streak camera manufacturer provides a calibration

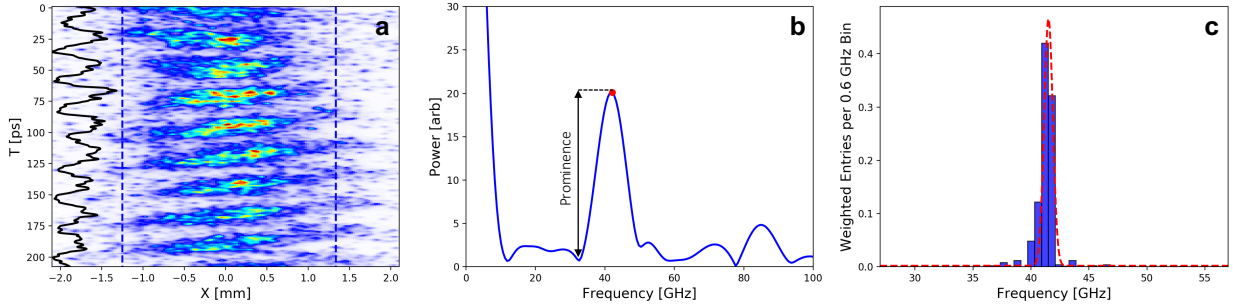


FIG. 2. a) Streak camera image of a modulated proton beam in a 200 ps-long time window. The vertical axis is the time axis, with $t = 0$ denoting the start of the window and not necessarily the start of the proton bunch. The horizontal axis is the transverse spatial coordinate with $x = 0$ located at the beam centroid. The ROI is denoted by dashed blue lines. The solid black line shows temporal projection of the image data in the region of interest. b) The absolute value of the FFT of the projection shown in a) after application of the analysis procedure described in Section III A. A peak in the spectrum is identified at 42.09 GHz using the peak-prominence algorithm and is denoted by the red circle. The prominence is denoted by the black arrow. c) Histogram of the peak frequencies identified with the column-by-column analysis technique. The entries to the histogram are weighted by their prominence and the histogram is fit with a gaussian centered at 41.57 ± 0.07 GHz, where the error on the centroid is extracted from the fit.

132 of the streak axis so that the vertical dimension of the image is mapped to the co-moving
 133 temporal dimension of the beam. The AWAKE streak camera is a Hamamatsu C10910-05
 134 model with a 16-bit, 2048×2048 pixel ORCA-Flash4.0 CMOS sensor, binned to 512×512
 135 pixels for streak operation. The streak camera provides time windows ranging from several
 136 ns down to 73 ps. For the smallest time window, the resolution is limited by the streak
 137 dynamics rather than the pixel size and is 1 ps [23]. An optical target (Ronchi ruling with
 138 5 lines per mm) was placed at the location of the foil and imaged by the streak camera
 139 to determine the point spread function (PSF) in the non-streaked (transverse) plane. The
 140 full-width half-maximum (FWHM) resolution in the transverse plane was determined to be
 141 0.187 mm, or roughly 8 pixels.

142 **III. STREAK IMAGE ANALYSIS AND FREQUENCY IDENTIFICATION**

143 The streak camera was operated with a time window of 208 ps near the center of the
144 proton bunch. The streak camera records images of the modulated bunch which are then
145 analyzed to extract the frequency of modulation.

146 **A. Fast Fourier Transform of the Projection**

147 The first step of the analysis to identify the signal region in the streak camera image and
148 take a projection of the data in the region of interest (ROI) along the time axis. A Hann
149 Filter is applied to the projection which is then embedded in a zero-padded array. A Fast
150 Fourier Transform (FFT) is taken of the zero-padded array. The purpose of zero-padding
151 the array is to smooth the result of the FFT, while the Hann Filter prevents aliasing due
152 to embedding. Figures 2a) and 2b) illustrate the signal identification and resulting FFT for
153 a sample image. The modulation frequency is identified by a peak-prominence algorithm.
154 The peak-prominence is defined as the amplitude of the local maximum with respect to the
155 nearest local minimum on the low frequency side of the maximum. This definition reduces
156 sensitivity to the high-amplitude DC component in the signal.

157 **B. Statistical Approach**

158 The FFT of the projection identifies the frequency of modulation averaged transversely
159 across the bunch, since the projection itself is a transverse average across the bunch. It is
160 possible that multiple frequencies appear in a single image, or that the same frequency ap-
161 pears with a phase offset across the image. This is what is observed in the SSM phenomenon,
162 where focused protons appear at the center of the image and defocused protons appear at
163 the edge of the image, phase shifted by 180 degrees [14]. In this case, the projection averages
164 both modulations together out of phase, and the result is a reduced amplitude of the peak
165 frequency.

166 In order to account for the variation in frequency across the image, the image is subdivided
167 into 40 columns, each 8 pixels wide, corresponding to the FWHM PSF resolution. The FFT
168 is taken for each column and the peak frequency is identified using the peak-prominence
169 algorithm described above. The identified frequencies are added to a histogram, weighted

170 by their prominence. Finally, a gaussian is fitted to the resulting histogram and the mean
171 peak frequency is determined from the centroid of the fit. The error on the mean is extracted
172 from the covariance matrix of the fit.

173 The column-by-column analysis has two advantages over the projection approach de-
174 scribed in Section III A. First, the algorithm is robust to noise when identifying the peak
175 frequencies in the signal region. Second, it produces a statistical error for each image, which
176 is a measure of our confidence in the signal identification. For the sample data shown in
177 Figure 2, the two approaches produce similar values, where the projection approach gives a
178 value of 42.09 GHz and the column-by-column approach yields 41.57 ± 0.07 GHz.

179 IV. RESULTS

180 We study the evolution of the plasma density as a function of time after ionization at
181 three laser pulse energies. The Rb vapor density is held at a constant value of $1.81 \pm 0.09 \times$
182 10^{14} cm⁻³, throughout the measurement. The relative laser-proton timing is adjusted via
183 the phase-locking system described in Section II. The time delay Δt is varied from 0 μ s
184 (corresponding to the laser seed pulse coincident with the center of the proton beam) to 80
185 μ s. The laser energy is controlled by changing the angle of an attenuating polarizer. For
186 every time delay setting, we record ten shots each for a laser pulse energy of 135 mJ, 95 mJ,
187 and 40 mJ, which we refer to as the high, medium, and low energy settings, respectively. In
188 addition, we record a single laser-off background shot for each setting, for an average of 33
189 events per time delay. Figure 3 shows example images of the self-modulated proton bunches
190 for different time delays after ionization with the high-power setting.

191 The images are grouped by time delay and laser energy setting. We apply the projection
192 analysis to each image (Section III A) and extract a first estimate of the peak frequency
193 for each time delay and laser energy. The results of the projection analysis for the high
194 energy laser setting are shown as a waterfall plot in Figure 4. The projection analysis shows
195 a nearly constant modulation frequency over the first 1 ms, followed by rapid decay. Note
196 that the larger time delays are logarithmically spaced. For time delays $\Delta t \geq 5$ μ s, a second
197 peak is observed in the FFT spectrum at twice the frequency of the main peak. The second
198 peak may be the result of events that have a transverse hose instability, which appears as a
199 zigzag pattern on the streak camera image with the same frequency as the SMI, or because

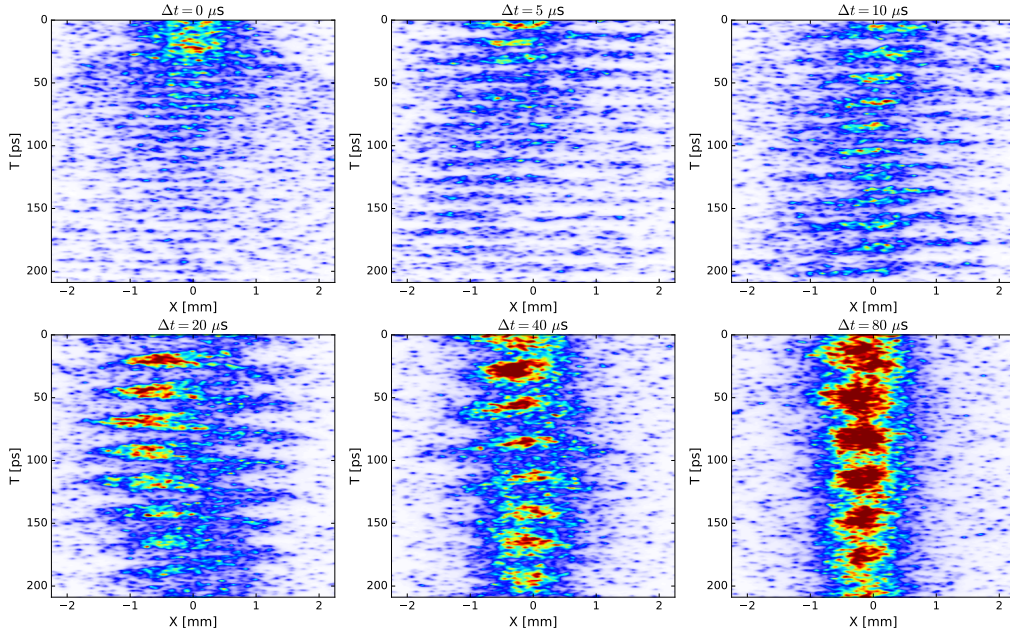


FIG. 3. Streak camera images of the modulated proton beam for example time delays Δt after ionization by the laser pulse with the high-power setting. The images share the same spatial, temporal, and color scales. The frequency of the modulation is clearly observed to be decreasing with time after ionization, indicating a decay in the plasma density. The image at a time delay of 20 μs has the signature of the hose instability which may arise in the case of un-seeded self-modulation.

200 of harmonics in the SSM signal.

201 The peak frequencies and error on the frequencies are determined using the statistical
 202 analysis described in Section III B. The data points are grouped into sets by time delay and
 203 laser pulse energy and the histograms for each image in the set are added together. The fit
 204 is performed on the ensemble histogram, with centroids and errors extracted from the fit.
 205 Finally, the frequencies ν are converted into plasma densities n with the equation

$$206 \quad n = \epsilon_0 m_e \left(\frac{2\pi\nu}{e} \right)^2 \quad (1)$$

207 where ϵ_0 is the permittivity of free space, m_e is the mass of the electron, and e is the charge
 208 of the electron. The error on the mean density is given by

$$209 \quad \sigma_n = 2n \frac{\sigma_\nu}{\nu}, \quad (2)$$

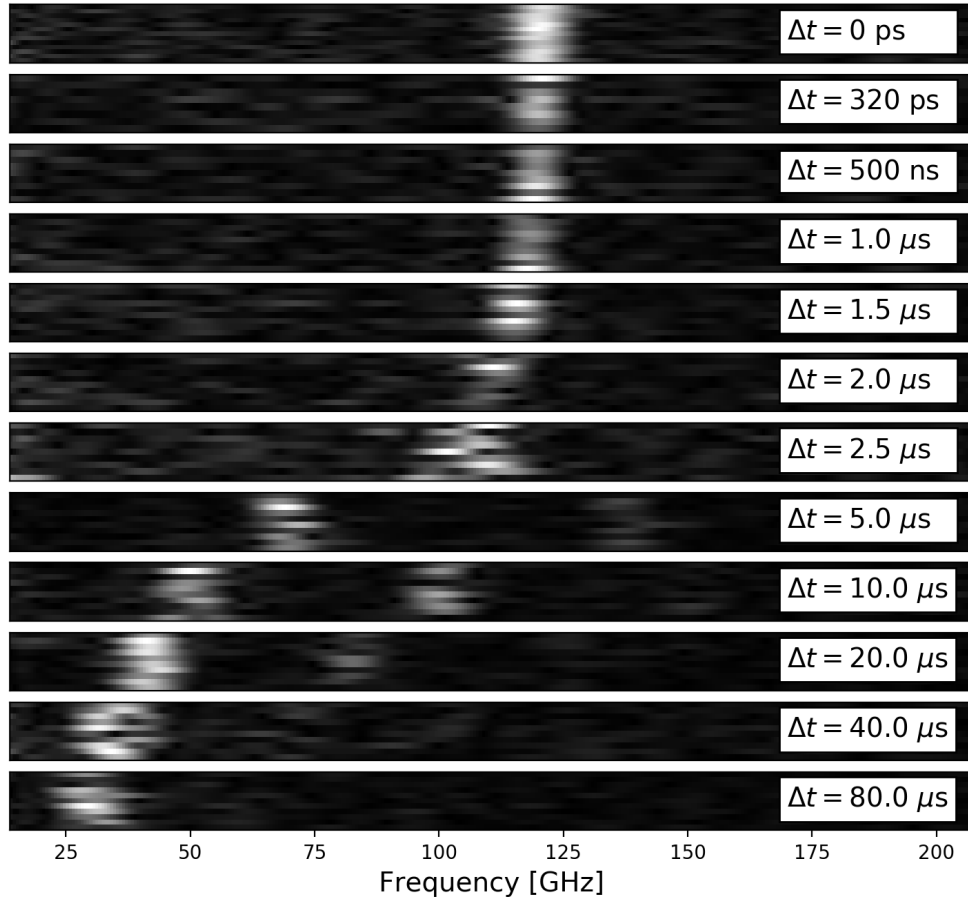


FIG. 4. Waterfall plot of the projected image data for the 135 mJ laser energy setting, with normalized color scale. Each subplot contains approximately 10 stacked Fourier spectra. The subplots are labeled by time after ionization, with 0 s at the top and 80 μ s at the bottom. The larger time delays are logarithmically spaced. For time delays $\Delta t \geq 5 \mu$ s, a second peak appears in the spectra at twice the frequency of the main peak. This may be due to the transverse hose instability, or harmonics in the main signal.

210 where σ_ν is the error on the mean frequency extracted from the gaussian fit of the histogram.
 211 Previous work has established that at zero time delay, the correlation between the modula-
 212 tion frequency and the plasma density is exact, and we can conclude that the plasma is fully
 213 ionized by the laser [15]. At longer time delays, the connection between the modulation
 214 frequency and plasma density is more nuanced, because the plasma may not be uniform
 215 across the proton bunch. We interpret the measured modulation frequency as the result of
 216 the average plasma density seen by the proton beam.

217 The results of the time delay scan are shown in Figures 5a) and 5b). Figure 5a) shows
 218 the evolution of the plasma density for delays less than 1 μs . The point at $\Delta T = 0$ s has been
 219 assigned an offset of 1 ps so that is visible on log x-axis plot. Further, the high, medium,
 220 and low energy data points are assigned small temporal offsets so that the error bars are
 221 visible without overlap, but all groups of points are taken at the same time delay, namely 0,
 222 320 ps, 500 ns, and 1 μs . The errors are multiplied by a factor of 5 so that they are visible.
 223 The measured plasma densities at 0 s, 320 ps, and 500 ns, are close to, but slightly less than
 224 the measured vapor density, shown with a 0.5% uncertainty in gray. The discrepancy could
 225 result from less than 100% ionization, or systematic error in the frequency measurement,
 226 but of particular significance is the fact that the measured density is unchanging over the
 227 first 500 ns. A roll-off in density begins to appear around 1 μs after ionization.

228 Figure 5b) shows the evolution of the plasma density for time delays of 500 ns and longer,
 229 plotted on a log-log scale. The low energy data point at $\Delta t = 2.0 \mu\text{s}$ has been removed due
 230 to noise in the signal. Error bars are plotted but are not visible on the the log-log scale. For
 231 time delays 10 μs and greater, the data are fit with a power law of form at^b . The significance
 232 of the power law evolution is discussed in Section V.

233 V. MODEL AND ANALYSIS

234 The data show two important features which we attempt to model. First, we observe
 235 a plateau in the plasma density for delays up to 1 μs . Second, we observe a power law-
 236 like decay of the plasma density at long time delays. Our model starts by considering the
 237 formation of the plasma column.

238 A. Plasma Formation

239 The ionizing laser has a nearly constant spot size of 2 mm throughout the 10 m long
 240 vapor cell. The peak intensity of the laser is 9.0×10^{12} W/cm², 6.3×10^{12} W/cm², 2.6×10^{12}
 241 W/cm², for the 135 mJ, 95 mJ, and 40 mJ laser energy settings, respectively. The first
 242 energy level of Rb is $U = 4.18$ eV, while the Ti:sapphire laser photon energy is $h\nu = 1.59$
 243 eV at a central wavelength of 780 nm. In all cases, the Keldysh parameter

$$244 \quad \gamma_k = \frac{\omega\sqrt{2m_eU}}{eE} > 1, \quad (3)$$

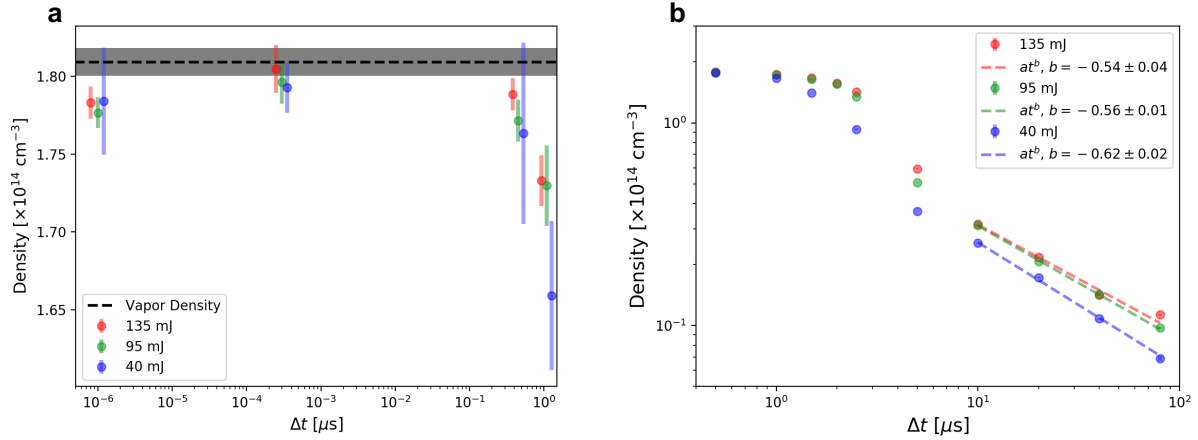
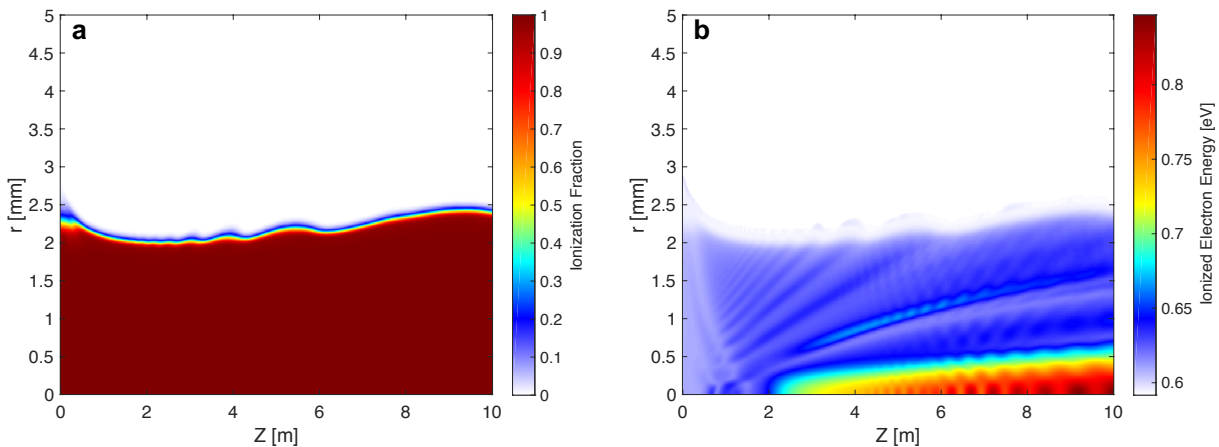


FIG. 5. a) Plasma density versus time after ionization for delays up to $1 \mu\text{s}$, for the high, medium, and low energy laser settings, plotted on a log x-axis. The errors are shown multiplied by a factor of five so that they are visible in the plot. The measured vapor density with a 0.5% uncertainty is also shown. The data points corresponding to $\Delta t = 0$ s are assigned an offset of 1 ps so that they are visible in the plot. In addition, the high, medium, and low energy data points are assigned small temporal offsets so that the error bars are visible without overlap, but all groups of points are taken at the same time delay, namely 0, 320 ps, 500 ns, and $1 \mu\text{s}$. b) Plasma density versus time after ionization for delays of 500 ns and greater. The data point corresponding to the low energy setting at $\Delta t = 2.0 \mu\text{s}$ has been excluded due to noise in the signal. A power law fit of the form at^b is applied to the data for $\Delta t \geq 10 \mu\text{s}$.

245 where ω is the laser frequency and E is the peak electric field [24]. For $\gamma_k > 1$, the dominant
 246 ionization mechanism is multi-photon absorption. We use simulations based on the PPT
 247 model to describe the ionization of the Rb vapor by the laser [25]. In particular, we are
 248 interested in the spatial distribution and temperature of the ionized electrons, as this will
 249 serve as the initial conditions for our plasma evolution model.

250 Figure 6a) shows the radial and longitudinal ionization fraction of the Rb vapor along
 251 the 10 m length of the cell for the 135 mJ laser energy setting. The average radius of
 252 ionization is 2.2 mm. The transverse laser profile used in the model was gaussian, but
 253 the laser mode profile in the experiment has a top-hat shape. Further simulations were
 254 performed with super-gaussian modes. The results of these simulations show less uniformity
 255 at the boundary of the ionized region, but produce a similar value for the average ionized

256 radius. Figure 6b) shows the average energy of the plasma electrons in the ionized region.
 257 The minimum energy of ionization is 0.59 eV, corresponding to three-photon absorption
 258 $E_{min} = U - 3h\nu$. The median electron energy after ionization is 0.62 eV. The model does
 259 not include plasma electron energy gain in the ponderomotive potential of the laser pulse,
 260 which has a maximum amplitude of $U_p = 0.51$ eV. We expect the energy contribution from
 261 U_p to be negligible given that the laser is linearly polarized and plasma electrons will be
 262 preferentially ionized at the maximum of the electric field which corresponds to the zero-
 263 crossing of the ponderomotive potential. Simulations were also performed for the 95 mJ and
 264 40 mJ laser energy settings. At 95 mJ, the average ionized radius is 2.0 mm and at 40 mJ
 265 it is 1.7 mm. In all cases, the median energy of the plasma electrons after ionization is 0.62
 266 eV.



268 FIG. 6. a) Ionization fraction for the 135 mJ laser pulse setting with a gaussian transverse pulse
 269 shape based on the PPT model. The average radius of the ionized plasma column is 2.2 mm. b)
 270 Energy of ionized electrons. The minimum energy after ionization is 0.59 eV, while the median
 271 energy is 0.62 eV.

272 B. Thermalization, Diffusion, and Recombination

273 Several processes contribute to the evolution of the plasma column after ionization. We
 274 consider the following three effects in detail: diffusion, thermalization, and recombination.
 275 For both diffusion and thermalization, the key parameter of interest is the collision fre-

276 quency [26]

$$277 \quad \nu_{\alpha,\beta} = 1.8 \times 10^{-19} \frac{(m_\alpha m_\beta)^{1/2} n \lambda_{\alpha,\beta}}{(m_\alpha T_\beta + m_\beta T_\alpha)^{3/2}}, \quad (4)$$

278 where m is the particle mass, T is the temperature, n is the density, and λ is the Coulomb
 279 logarithm. Here, α and β represent two species in the plasma, so the same equation can be
 280 used to calculate the electron-electron collision frequency ν_{ee} , the ion-ion collision frequency
 281 ν_{ii} , and the electron-ion collision frequency ν_{ei} . Immediately after ionization, the electron
 282 distribution is non-thermal, with a median electron energy of $E_{e0} = 0.62$ eV. The electron
 283 thermalization time is $t_{ee} = \nu_{ee}^{-1} \approx 100$ ps. The temperature of the electrons after thermal-
 284 ization is $T_{e0} = 2E_{e0}/3 = 0.41$. The ions are already thermal at the time of ionization with
 285 a temperature of $T_{i0} = 0.040$ eV and will come to equilibrium with the electrons in time
 286 $t_{ei} = \nu_{ei}^{-1} \approx 7$ μ s, which is much longer than the electron thermalization time.

287 Diffusion describes the expansion of the plasma column into the unionized vapor. The
 288 collision frequency determines the ion diffusion rate

$$289 \quad D_i = \frac{T_i}{m_i \nu_{tot}}, \quad (5)$$

290 where $\nu_{tot} = \nu_{ii} + \nu_{i0}$ is the sum of the ion-ion collision rate and the ion-neutral collision
 291 rate. In the case of a plasma with hot electrons and cold ions, the electrons increase the
 292 diffusion rate of the ions such that both species evolve at the ambipolar diffusion rate
 293 $D_a \approx D_i(1 + T_e/T_i)$ [27]. Inside the column, where there is complete ionization of the plasma,
 294 the diffusion rate is dominated by ion-ion collisions ($\nu_{tot} = \nu_{ii}$) and the mean free path is
 295 short $l = v_t/\nu_{ii} = 1.75$ μ m, where v_t is ion thermal velocity. Outside the plasma column, the
 296 mean free path of the ions is determined by ion-neutral collision frequency ($\nu_{tot} = \nu_{i0}$). The
 297 ion-neutral charge exchange cross section is given by $\sigma_{Rb} = (a - b \ln v_t)^2$, with $a = 42 \times 10^{-8}$
 298 cm and $b = 1.85 \times 10^{-8}$ cm [28]. For an ion thermal velocity of $v_t = 3.69 \times 10^4$ cm/s,
 299 we find $\sigma_{Rb} = 5 \times 10^{-14}$ cm², which at a neutral density $n_0 = 1.81 \times 10^{14}$ cm⁻³ yields
 300 a mean free path $l = v_t/\nu_{i0} = 1.1$ mm, much larger than the mean free path inside the
 301 column. The ambipolar diffusion rate varies greatly between the ionized and unionized
 302 regions, jumping from 10 cm²/s to 15,000 cm²/s across the boundary of the plasma column.
 303 In this scenario, the expansion of the plasma column into the unionized vapor is described
 304 by the time-dependent ambipolar diffusion wave equation [29]

$$305 \quad \frac{1}{v_a^2} \frac{\partial^2 n}{\partial t^2} + \frac{1}{D_a} \frac{\partial n}{\partial t} = \nabla n, \quad (6)$$

306 where $v_a = (\gamma_r T_e / m_i)^{1/2}$ is the ambipolar velocity, with γ_r the ratio of the specific heats
 307 of the plasma and vapor. For times $t < D_a / v_a^2 = 1.5 \mu\text{s}$, we neglect the diffusion term
 308 and the expansion of the plasma column is described by a wavefront propagating into the
 309 neutral vapor at speed $v_a \approx 9.6 \times 10^4 \text{ cm/s}$, with $\gamma_r = 2$ for hot plasma diffusing into neutral
 310 gas following the convention used in Reference [30]. While the edge of the column expands
 311 outwards, a density depression propagates inwards, etching away at the plasma column.
 312 The on-axis density remains constant until the inward propagating wave reaches the axis,
 313 at which point it starts to drop rapidly.

314 Equation 6 is evaluated numerically for different electron temperatures T_e and plasma
 315 column radii r_0 and compared with the data. Assuming an initial electron temperature of
 316 0.41 eV, and an initial plasma density $n_0 = 1.785 \times 10^{14} \text{ cm}^{-3}$, corresponding to the average
 317 measured plasma density at $\Delta t = 0$, we fit the model to the data and find the initial plasma
 318 radii to be 1.73 mm, 1.70 mm, and 1.63 mm for the high, medium, and low energy laser
 319 settings, respectively. On the other hand, higher electron temperatures are consistent with
 320 larger initial radii, and vice versa, because the two quantities can be combined to produce
 321 the same etching time of the plasma column $t_{etch} = r_0 / v_a$. Figure 7 shows the data in the
 322 range $0.5 \leq \Delta t \leq 1.5 \mu\text{s}$ compared to the on-axis density curves from the model. The
 323 dashed lines indicate the fit to the data at a fixed temperature of 0.41 eV, while varying
 324 the initial plasma radius. Note that the model cannot simultaneously provide a good fit to
 325 the data at $\Delta t = 1.0$ and $\Delta t = 1.5 \mu\text{s}$. We also show a gray region corresponding to other
 326 model parameters consistent with the data. These parameters include initial plasma electron
 327 temperatures in the range of $0.39 \leq T_e \leq 0.73 \text{ eV}$. The lower limit corresponds to the case
 328 where electrons are ionized with the minimal possible energy (3-photon absorption), while
 329 the upper limit corresponds to the case where the electrons gain an additional 0.51 eV of
 330 energy in the ponderomotive potential of the laser. As discussed in Section V A, we expect
 331 the energy gain from the ponderomotive potential to be small, and therefore the upper
 332 bound chosen for the electron temperature is a conservative estimate. The initial plasma
 333 column radii that are consistent with the data in the range of temperatures cited above is
 334 $r_0 = 1.58 \text{ mm}$ at $T_e = 0.39 \text{ eV}$ and $r_0 = 2.30 \text{ mm}$ at $T_e = 0.73 \text{ eV}$. In future experiments, we
 335 will make an independent measurement of the initial plasma radius via Schlieren imaging
 336 and this information can be used to break the degeneracy of the model.

337 For times delays between $1.5 \leq \Delta t \leq 10 \mu\text{s}$, the dominant processes are diffusion and

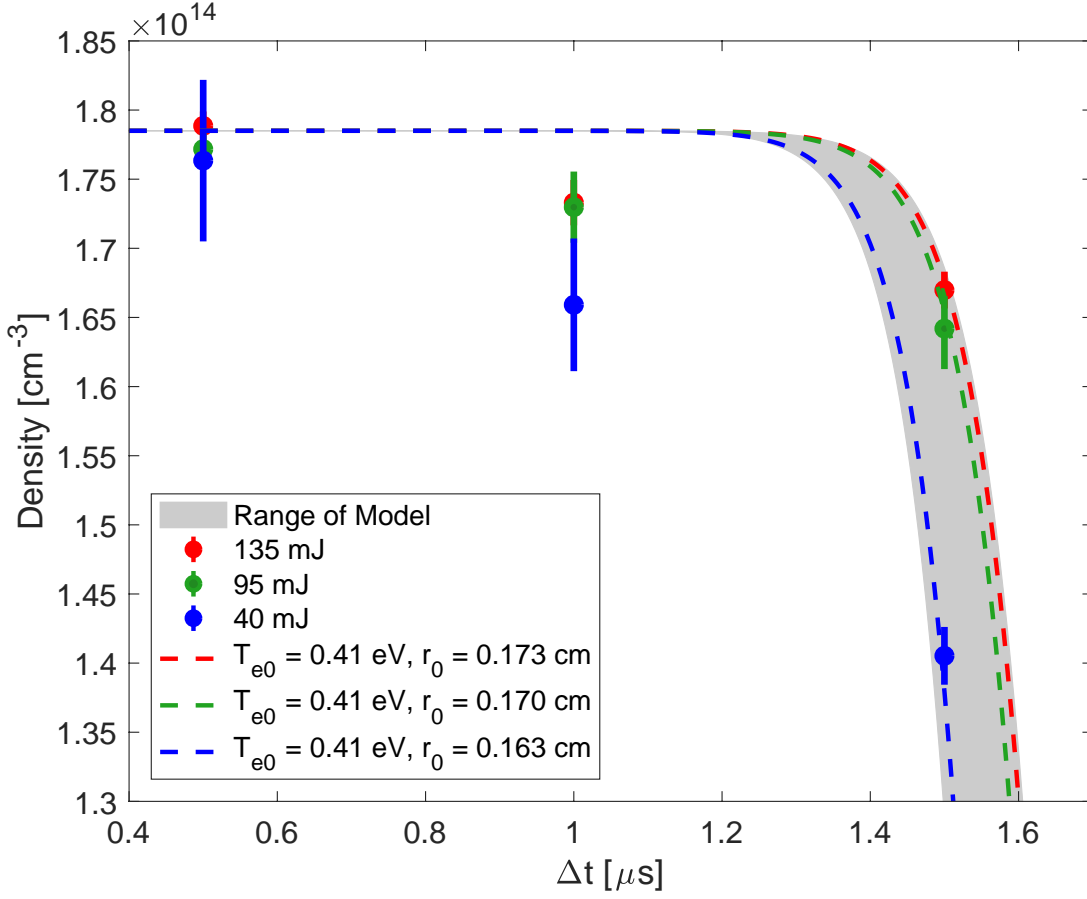


FIG. 7. Plasma density versus time compared to the on-axis density curves from the model. The red, blue and green data points correspond to the high, medium, and low energy laser settings, respectively, shown with error bars multiplied by five. The dashed lines are fits to the data at a fixed initial electron temperature of 0.41 eV. The shaded gray region shows the region of degeneracy of the model, where different values of T_e and r_0 produce similar on-axis density curves, so long as $t_{etch} = r_0/v_a \approx 1.5 \mu s$.

338 thermalization. During this time, the plasma spreads out and cools as the electrons come
 339 to equilibrium with the ions. The cool electrons may then start to recombine with the ions
 340 through the 3-body recombination process

$$341 \quad \frac{dn}{dt} = -\alpha_3 n^3, \quad (7)$$

342 with α_3 given by [26]

$$343 \quad \alpha_3 = 8.75 \times 10^{-27} T_e^{-4.5}. \quad (8)$$

344 The strong dependence on temperature in equation 8 explains why no recombination occurs
 345 until the electrons are cooled by coming to equilibrium with the ions. After thermalization,
 346 α_3 is constant and Equation 7 has solution

$$347 \quad n(t) = (2\alpha_3 t + n_0^{-2})^{-1/2}. \quad (9)$$

348 Figure 5b) shows the data at times $10 \leq \Delta t \mu\text{s}$ fit with a power law. The extracted powers
 349 are close to but slightly more negative than -0.5. It is likely that there are other plasma
 350 loss mechanisms at extremely long time delays, such as plasma reaching the wall of the cell,
 351 which contribute to a higher rate of decay of the plasma density.

352 VI. DISCUSSION AND SUMMARY

353 In this work, we measure the decay of a singly-ionized Rubidium plasma for times after
 354 ionization $0 < \Delta t < 80 \mu\text{s}$. The plasma density is inferred from the frequency of modulation
 355 of a high energy proton beam. We observe two features in the data: a plateau in the plasma
 356 density for times $\Delta t \lesssim 1 \mu\text{s}$, and a power law decay for $10 \mu\text{s} \lesssim \Delta t$. The short-time delay
 357 behavior is modeled using the time-dependent ambipolar diffusion wave equation, while the
 358 long-time delay behavior is described by 3-body recombination.

359 This measurement was motivated by interest in the self-modulation instability and opti-
 360 mal conditions for proton beam-driven plasma wakefield. We find that it is indeed important
 361 to suppress ionization prior to the transit of the proton beam driver in order to control the
 362 seeding of the wakefield. Possible sources of ionization include laser pre-pulses or impact
 363 ionization by the head of the proton driver. The latter source is intrinsic to the physics
 364 of SSM. Therefore, the plasma density arising from impact ionization can be considered
 365 the background plasma density that the system should return to prior to the passage of a
 366 subsequent bunch. For a vapor density of $1.81 \times 10^{14} \text{ cm}^{-3}$, we extrapolate the results of
 367 Reference [31] to estimate a background density of $4 - 8 \times 10^{11} \text{ cm}^{-3}$. Extrapolating from
 368 the fit presented in Figure 5b), we find that it takes approximately 25 ms after ionization to
 369 reach this density. This time scale is short compared to the 20 second repetition rate of the
 370 AWAKE experiment, but long compared to the repetition rate of other accelerators used in
 371 high-energy physics applications, which often operate with sub-microsecond bunch spacing.

372 This work considers the decay of a cold plasma after ionization by a laser pulse, but

373 there will be an additional contribution to the plasma electron energy from the passage of
374 previous proton beams driving large amplitude wakefields. In this case the plasma electron
375 energy will be in the keV range and the subsequent thermalization and recombination will
376 require even more time. In future experiments, we aim to probe this effect by sending two
377 proton beams through the vapor cell and measuring the effect of the leading on bunch on
378 the trailing bunch as a function of the separation between the beams.

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