

PIERRE
AUGER
OBSERVATORY



INSTITUTO DE FÍSICA
Universidade Federal Fluminense

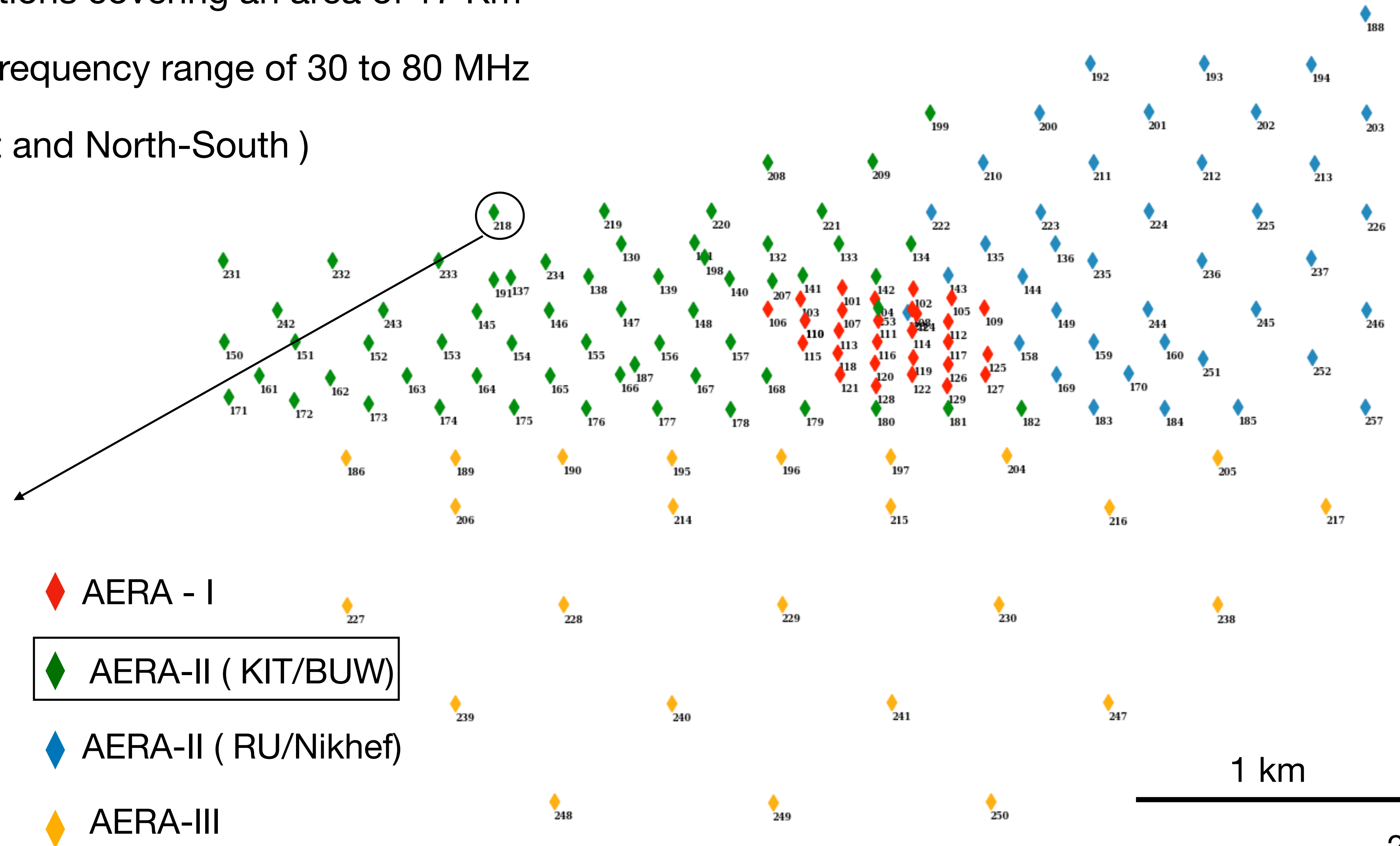
Absolute calibration and investigation of ageing of the AERA Radio Detector

R. M. de Almeida, for the Pierre Auger Collaboration

AERA - Auger Engineering Radio Array

- 153 Autonomous radio stations covering an area of 17 Km²
- Antennas sensitive in the frequency range of 30 to 80 MHz
- 2 polarizations (East-West and North-South)

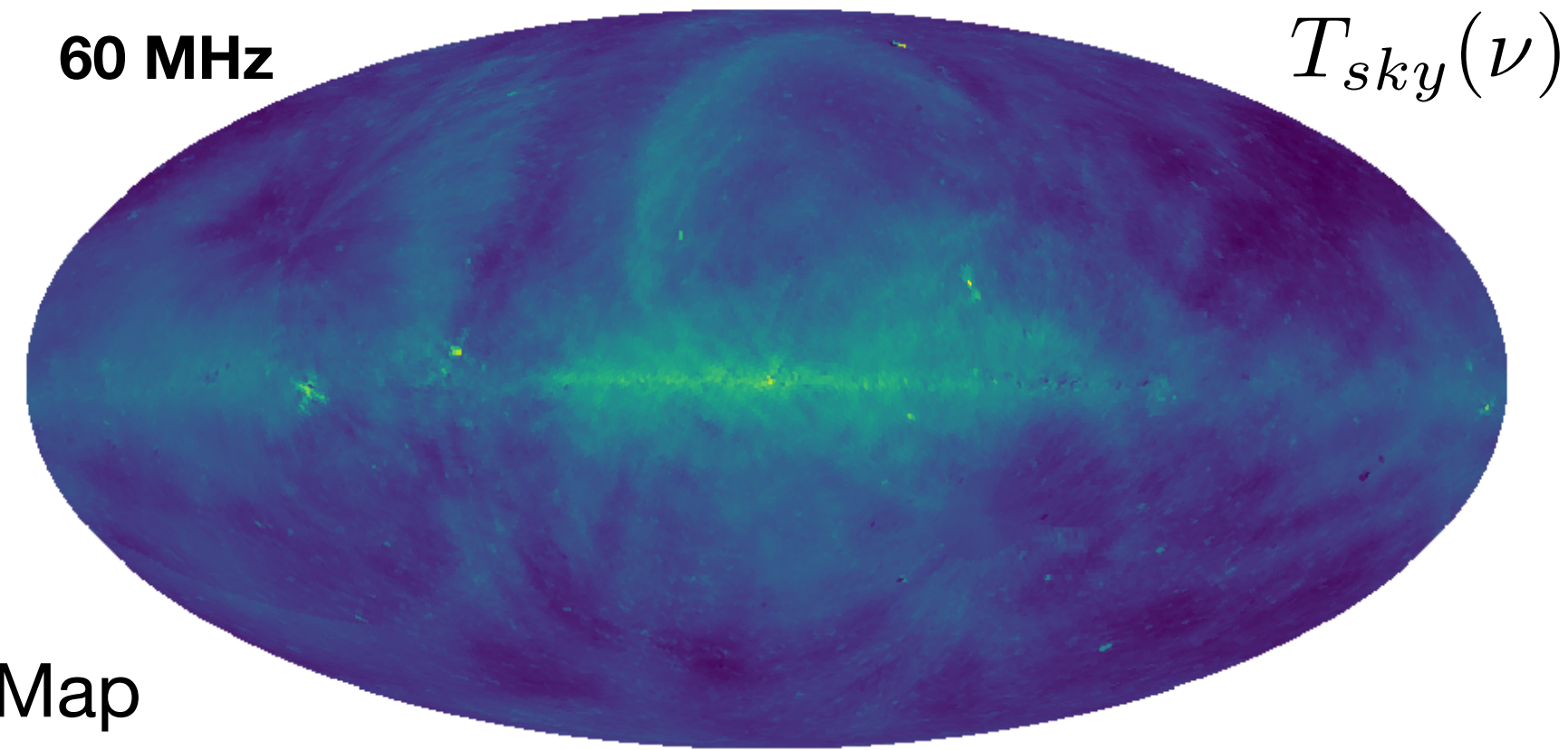
Butterfly antennas



Galactic calibration: Radio sky model

Antenna directional response

$$P_{sky}(t, \nu)$$



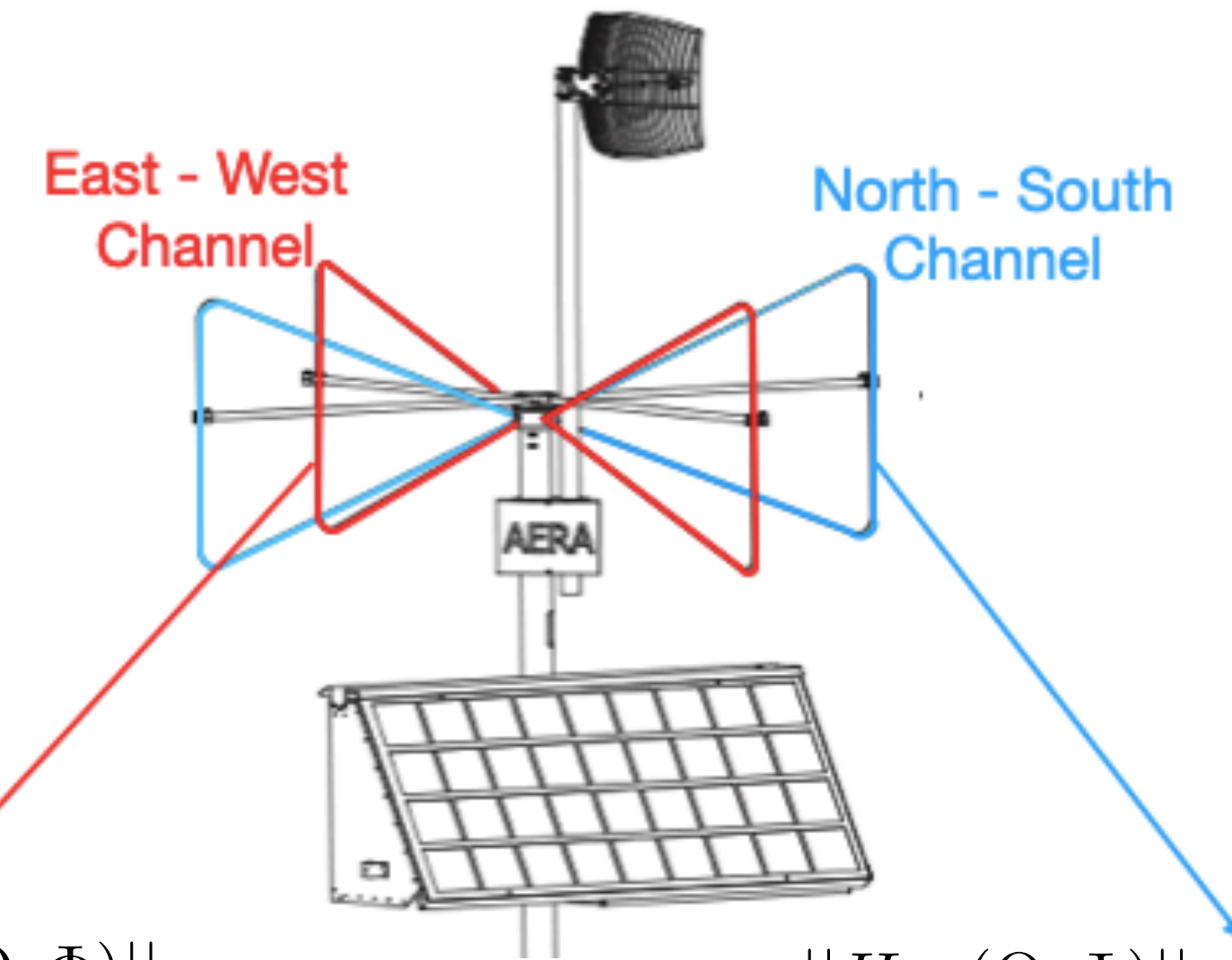
By LFMap

$$B(\nu, \alpha, \delta) = \frac{2K_b\nu^2}{c^2} T_{sky}(\nu, \alpha, \delta)$$

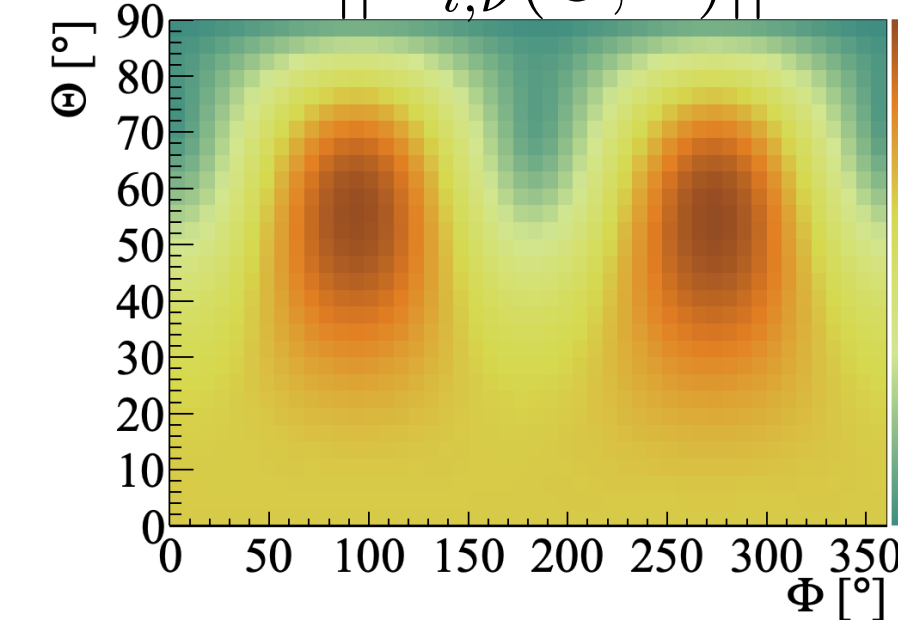
$$I_{\Omega, \nu}(\alpha, \delta) = B_\nu(\alpha, \delta) \|H_{i, \nu}(\alpha, \delta)\|^2$$

$$P_{sky}(t, \nu) = \frac{1}{2} \frac{Z_0}{Z_L} \int_{\Omega} I_{\Omega, \nu}(\alpha, \delta) d\Omega$$

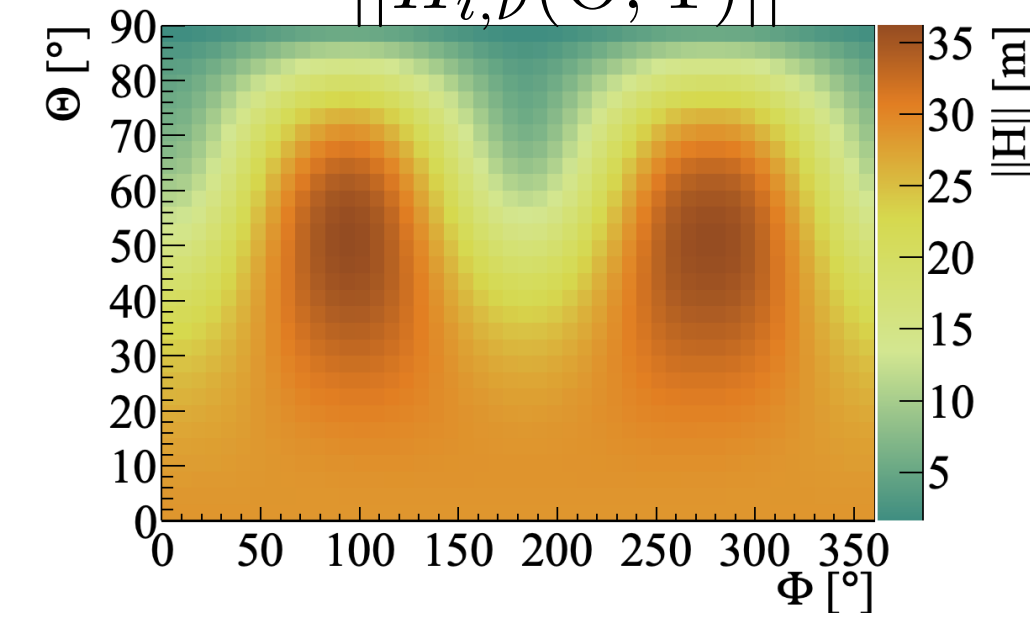
$$P_{model}(t, \nu) = P_{sky}(t, \nu) G_{ant}(\nu) G_{RCU}(\nu) S(\nu) + N_{tot}(\nu)$$



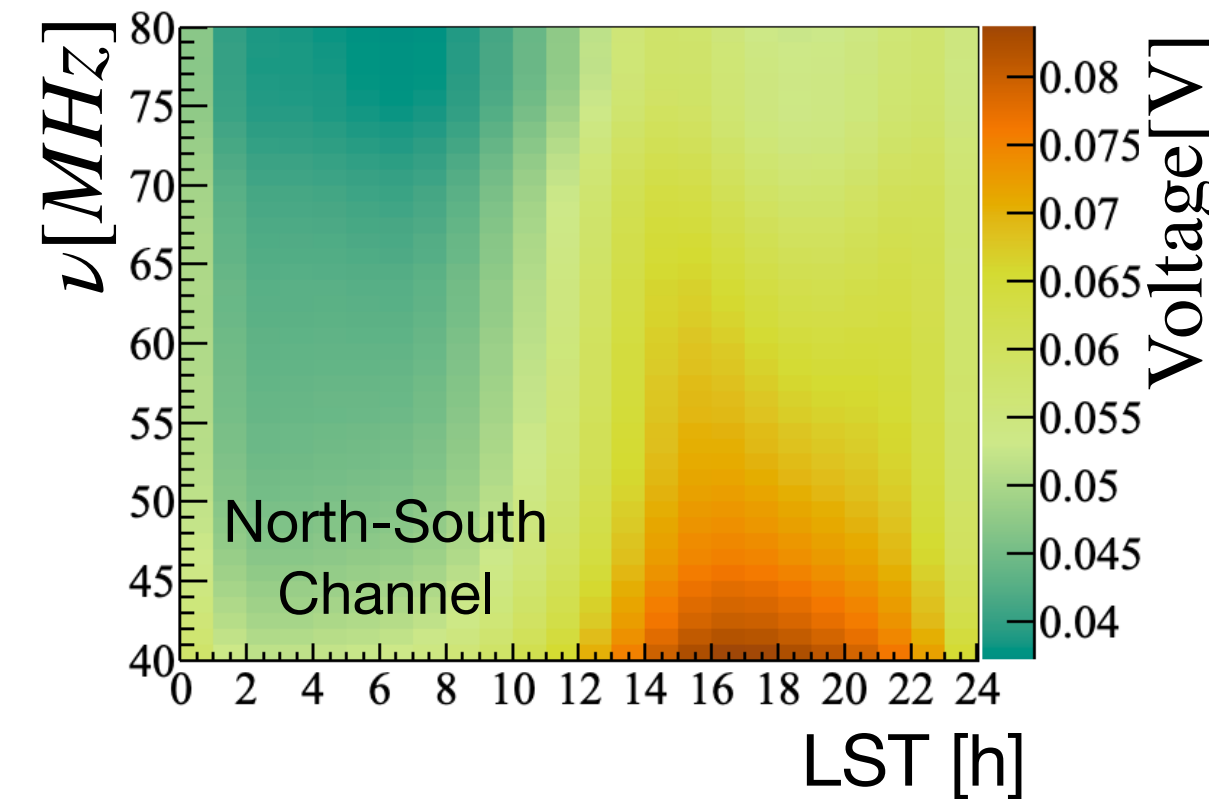
$$\|H_{i, \nu}(\Theta, \Phi)\|$$



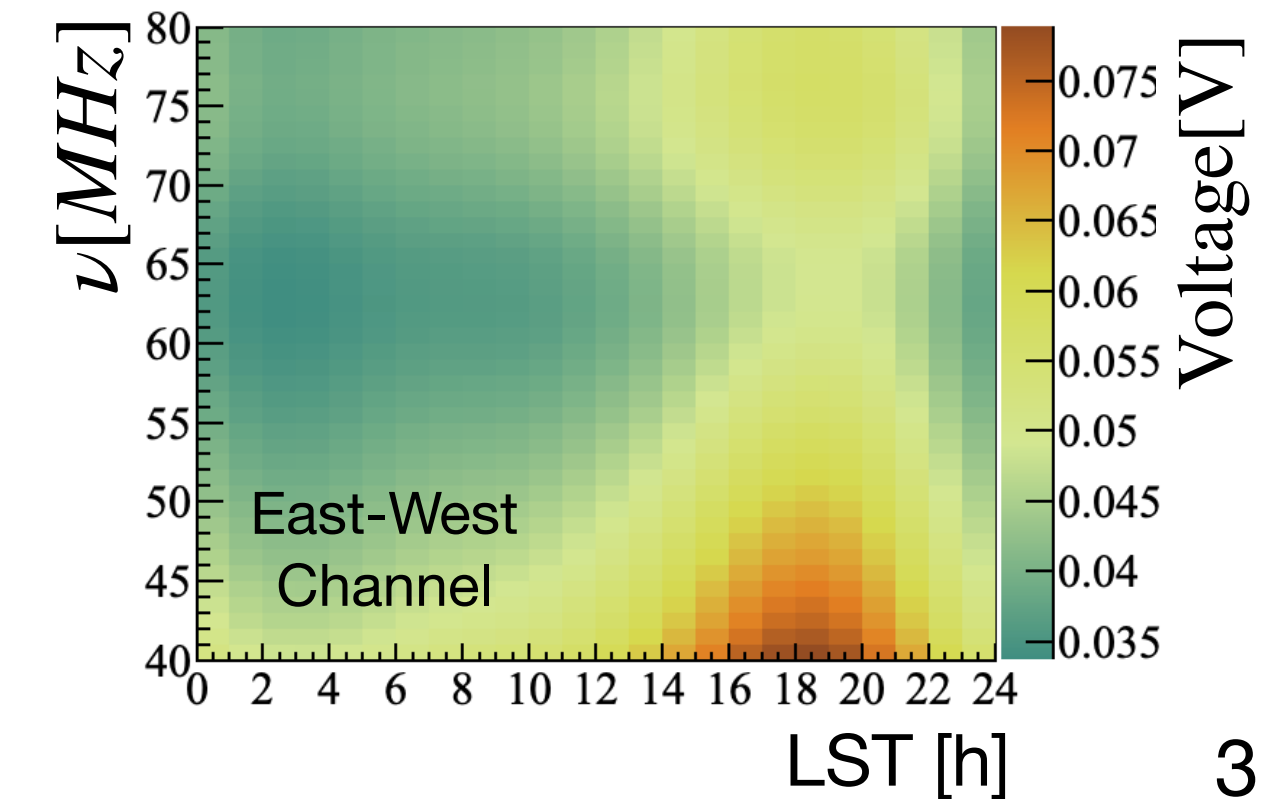
$$\|H_{i, \nu}(\Theta, \Phi)\|$$



$$U_{sky}(t, \nu)$$



$$U_{sky}(t, \nu)$$



Narrowband RFI and Broadband RFI

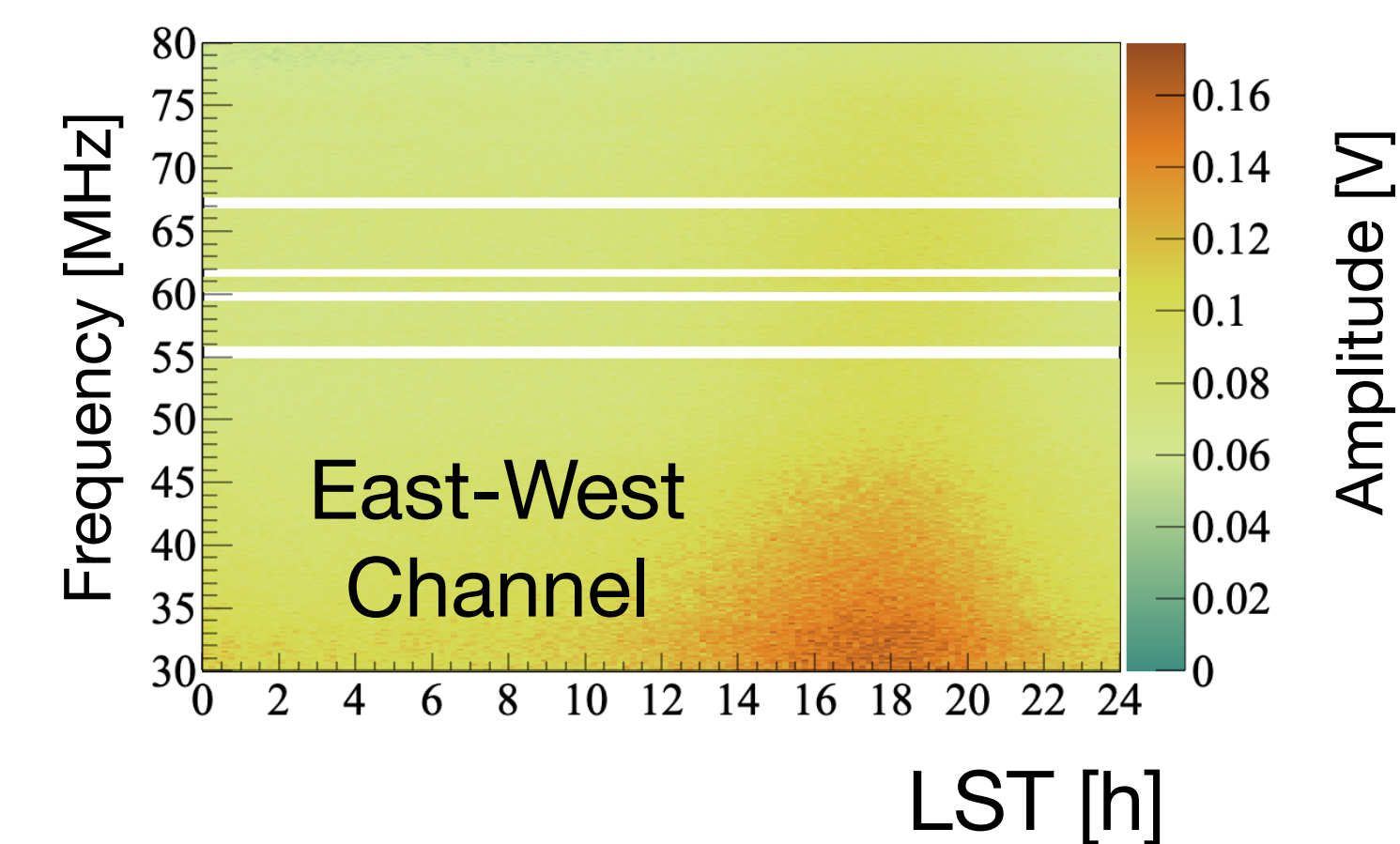
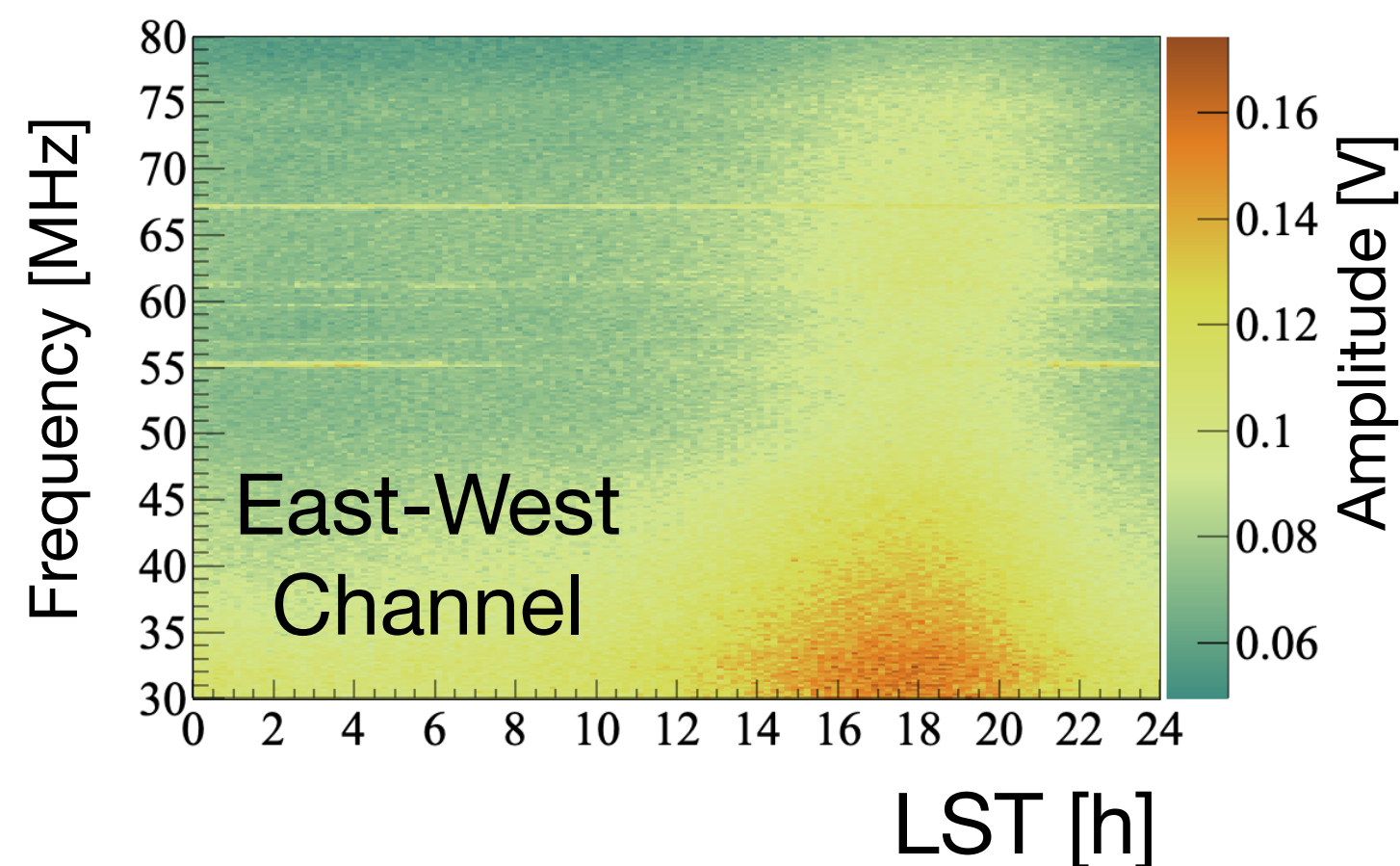
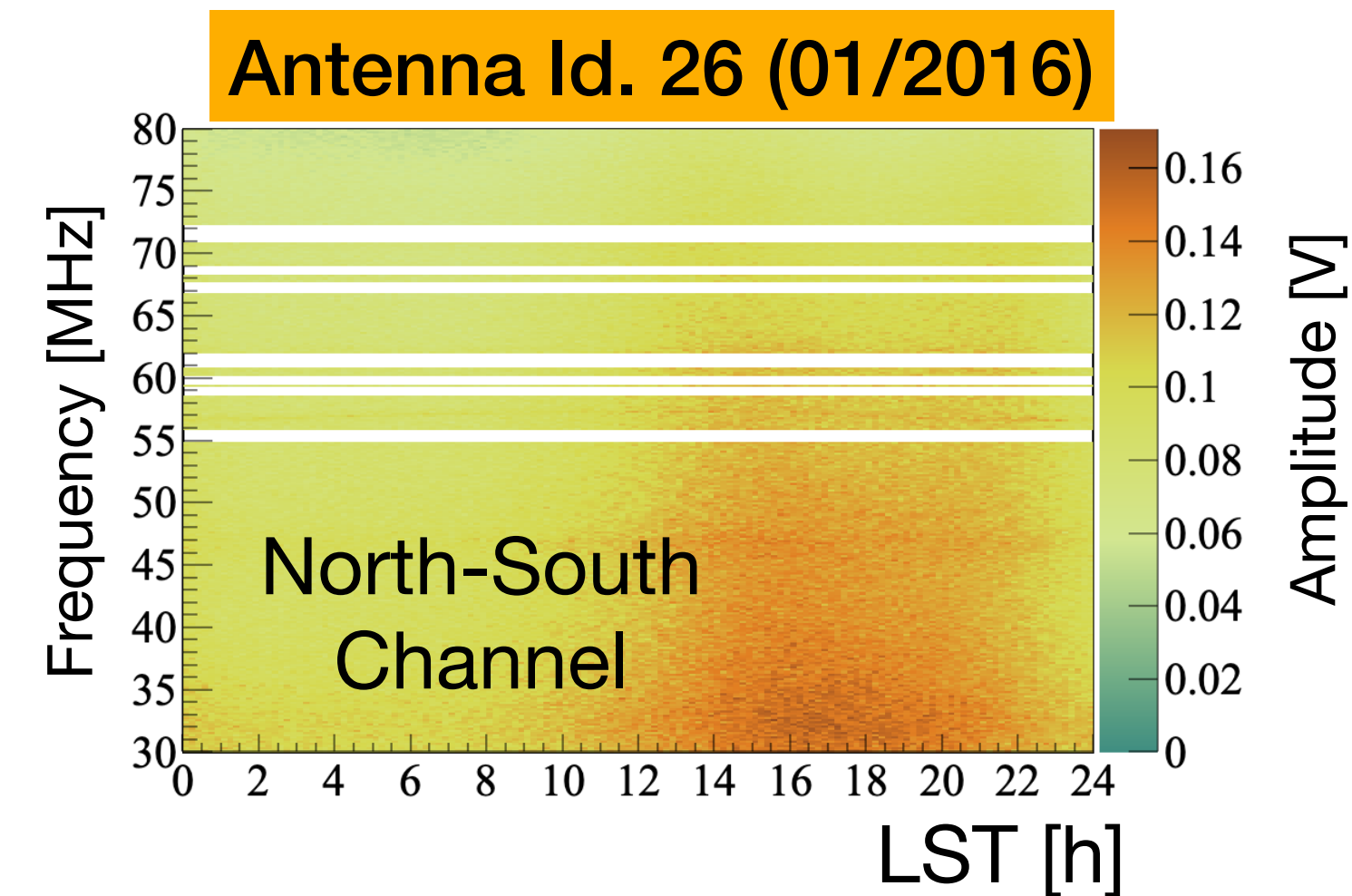
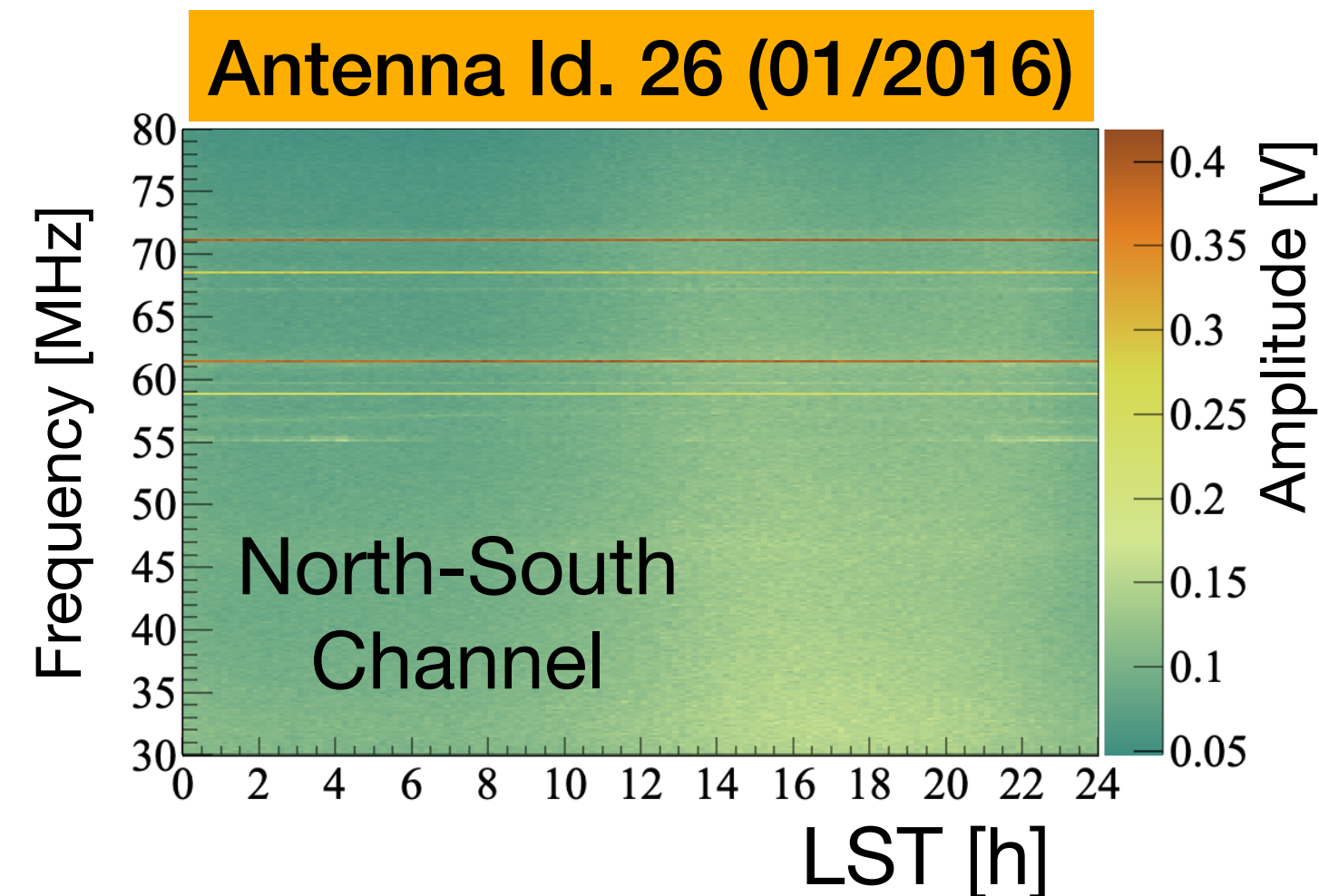
- We use 52 butterfly antennas (externally triggered setup)

Narrowband RFI removal

- We identify and remove narrowband RFI

Broadband RFI removal

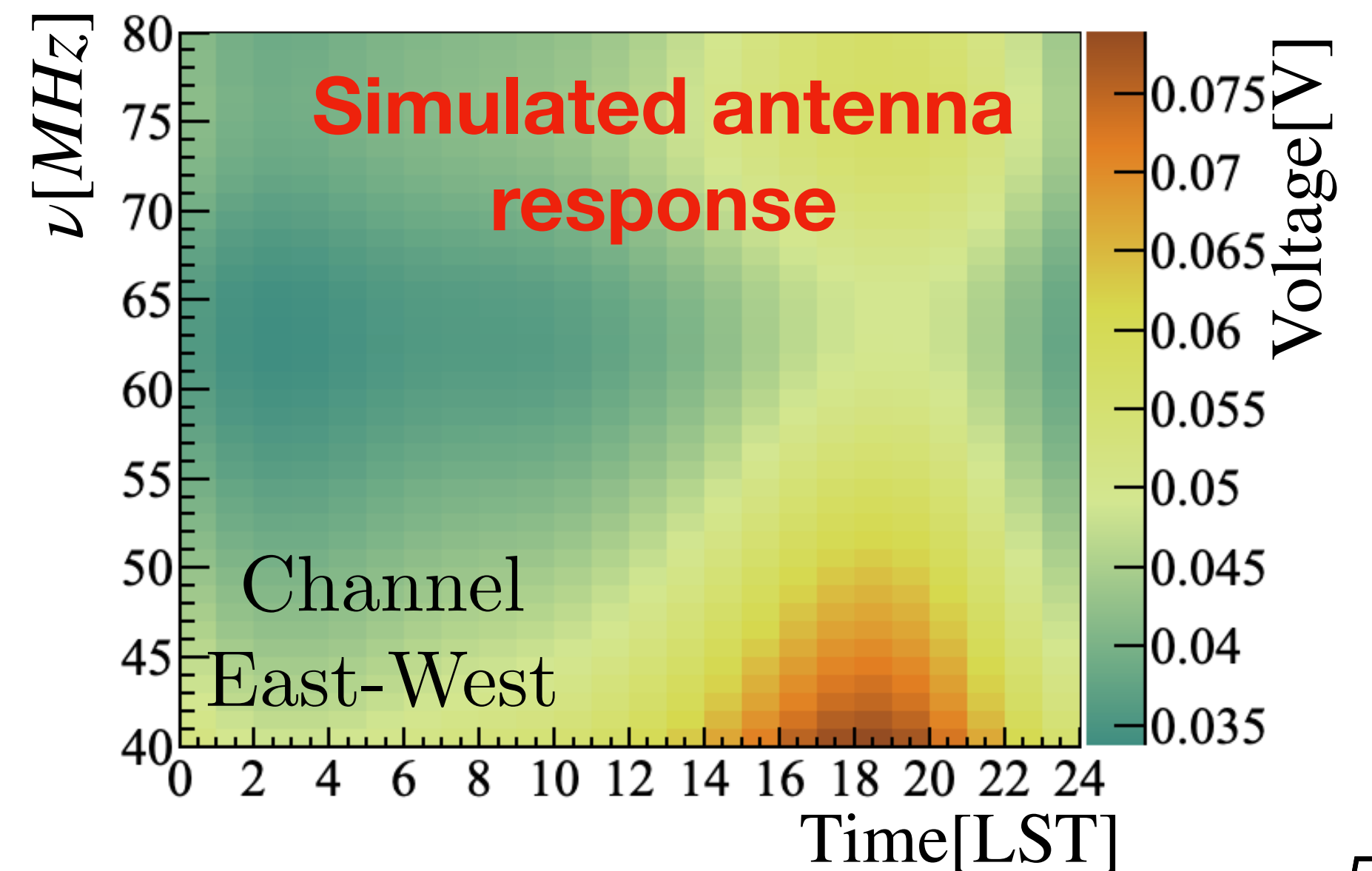
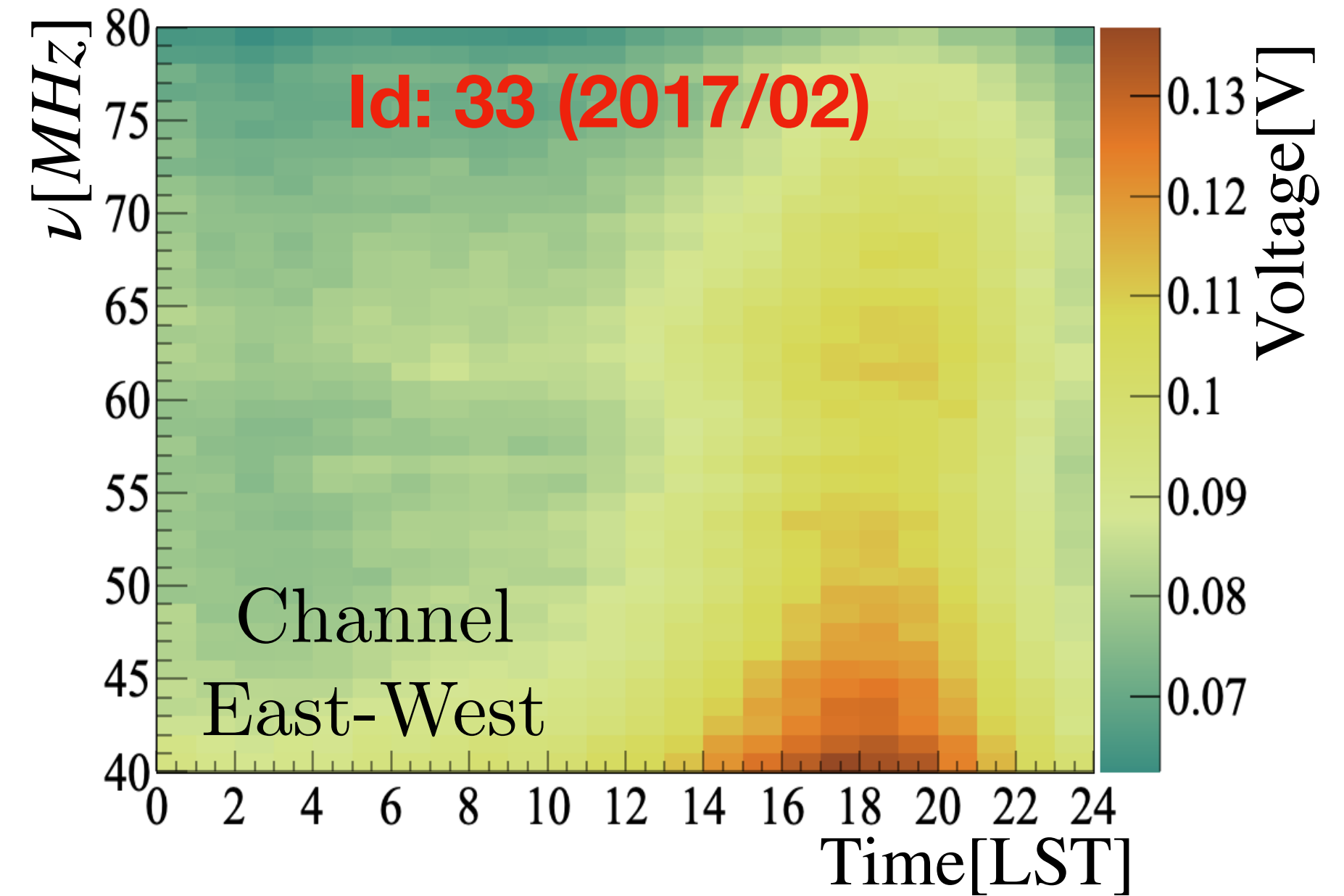
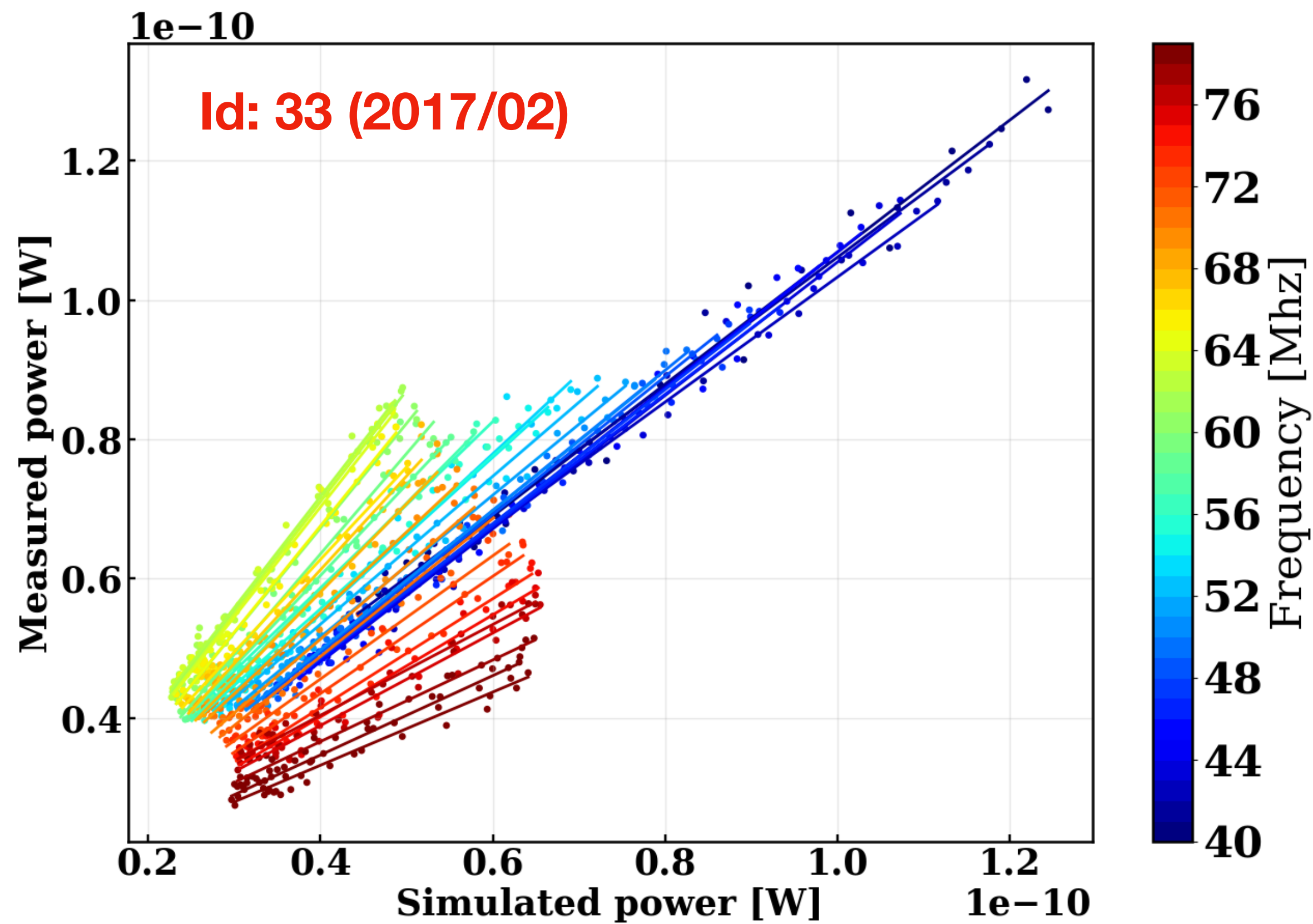
- We use a LST-dependent threshold approach
- Threshold as a function of the LST is determined for each antenna, month and channel



Study of C_0 as a function of the time

$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu) S(\nu) + N_{tot}(\nu)$$

- Independent linear fit for each frequency band



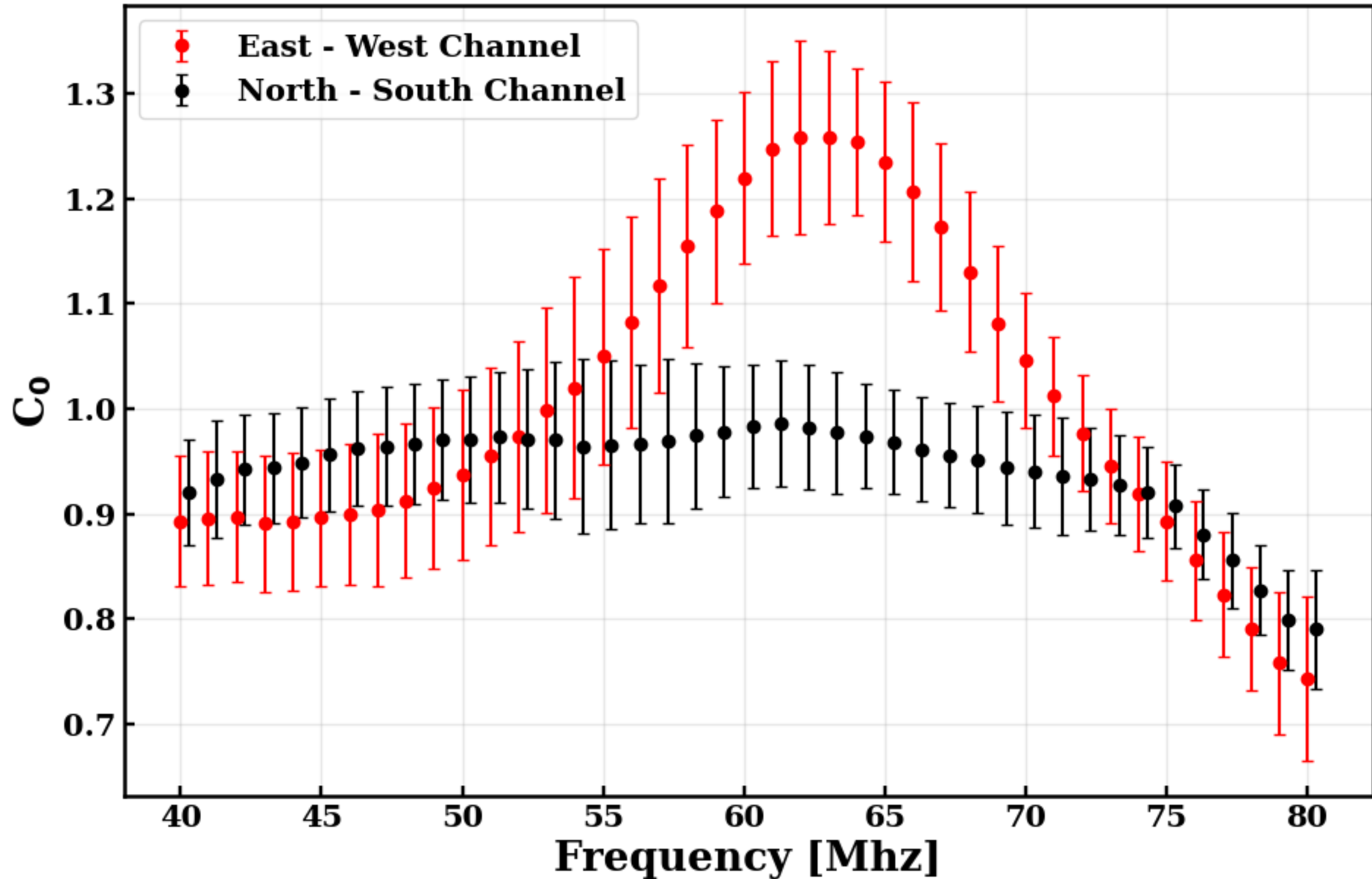
Study of C_0 as a function of the time

→ Perform a fit for each antenna (and channel) in a monthly basis from 2014 to 2020

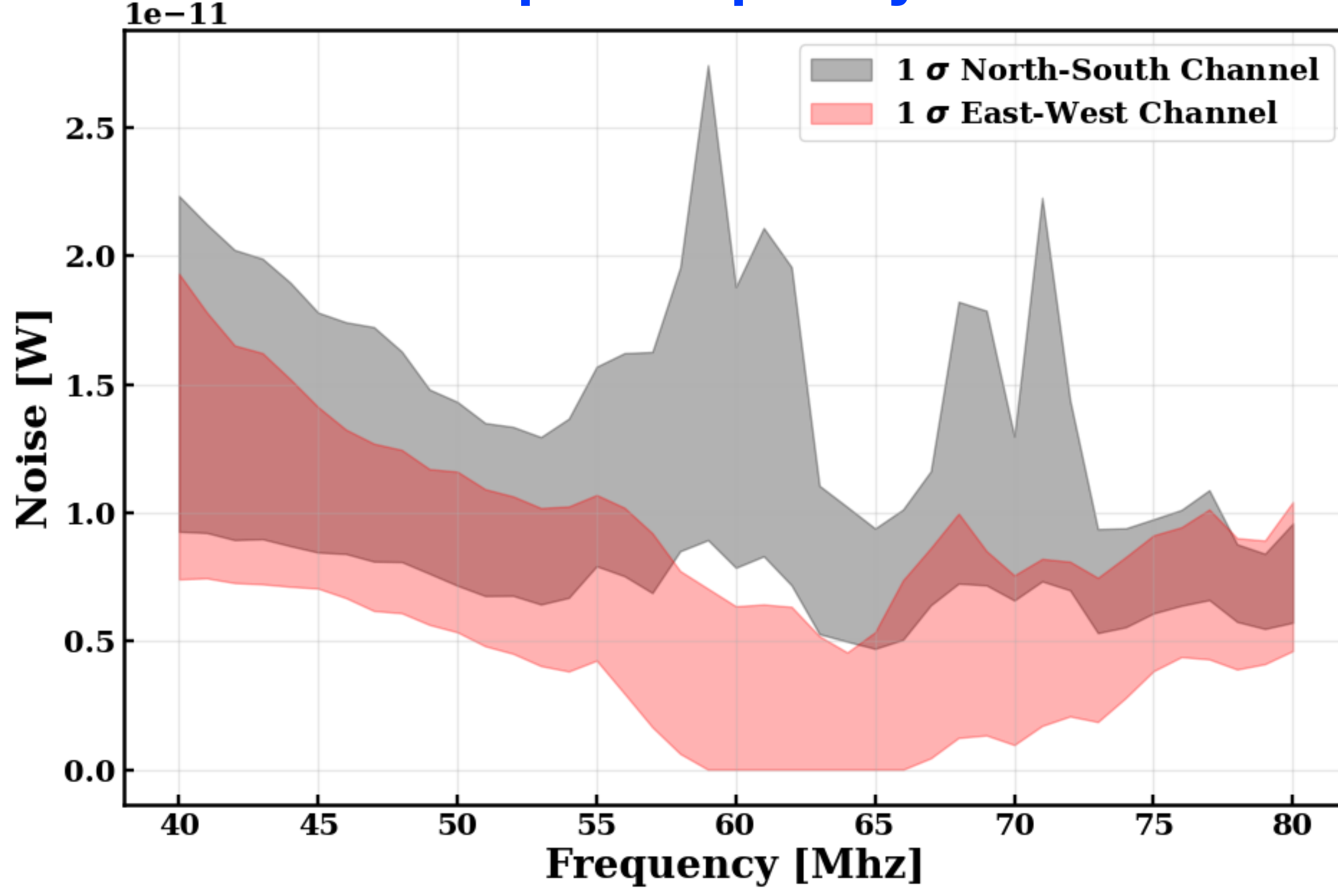
$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S(\nu) + N_{tot}(\nu)$$

$$C_0 = \sqrt{S}$$

C_0 per frequency factor results



Noise per frequency results



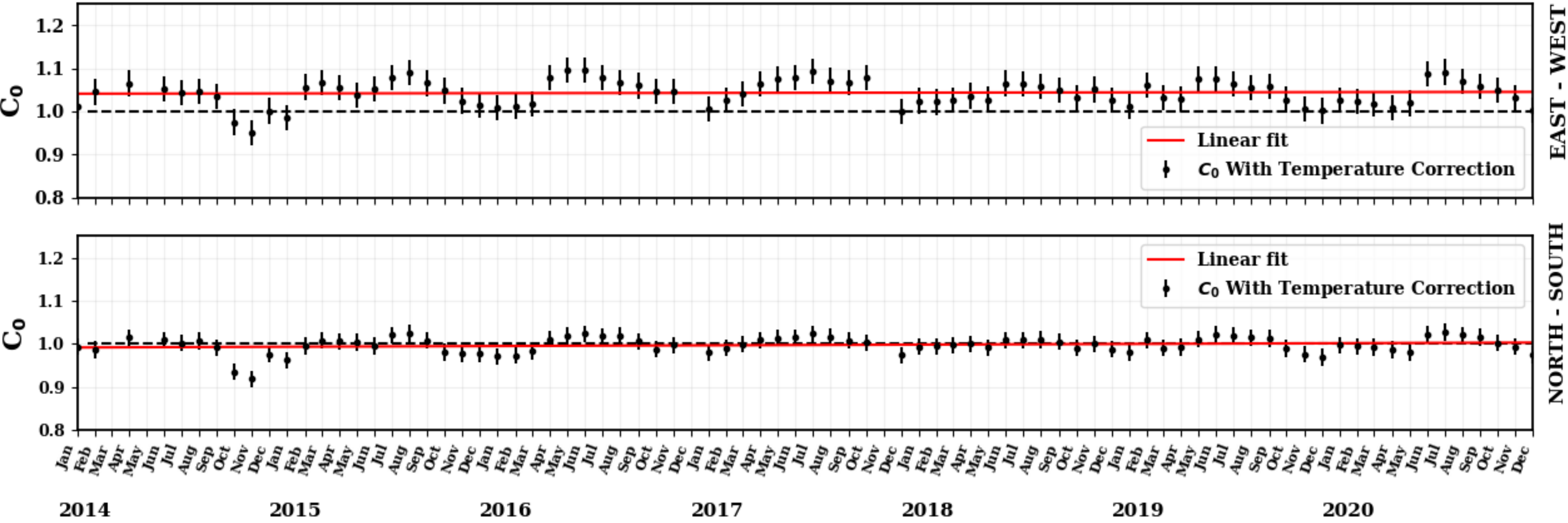
Distributions considering all antennas and months

Study of C_0 as a function of the time

- For each antenna and month, C_0 is given by $\langle C_0(\nu) \rangle$

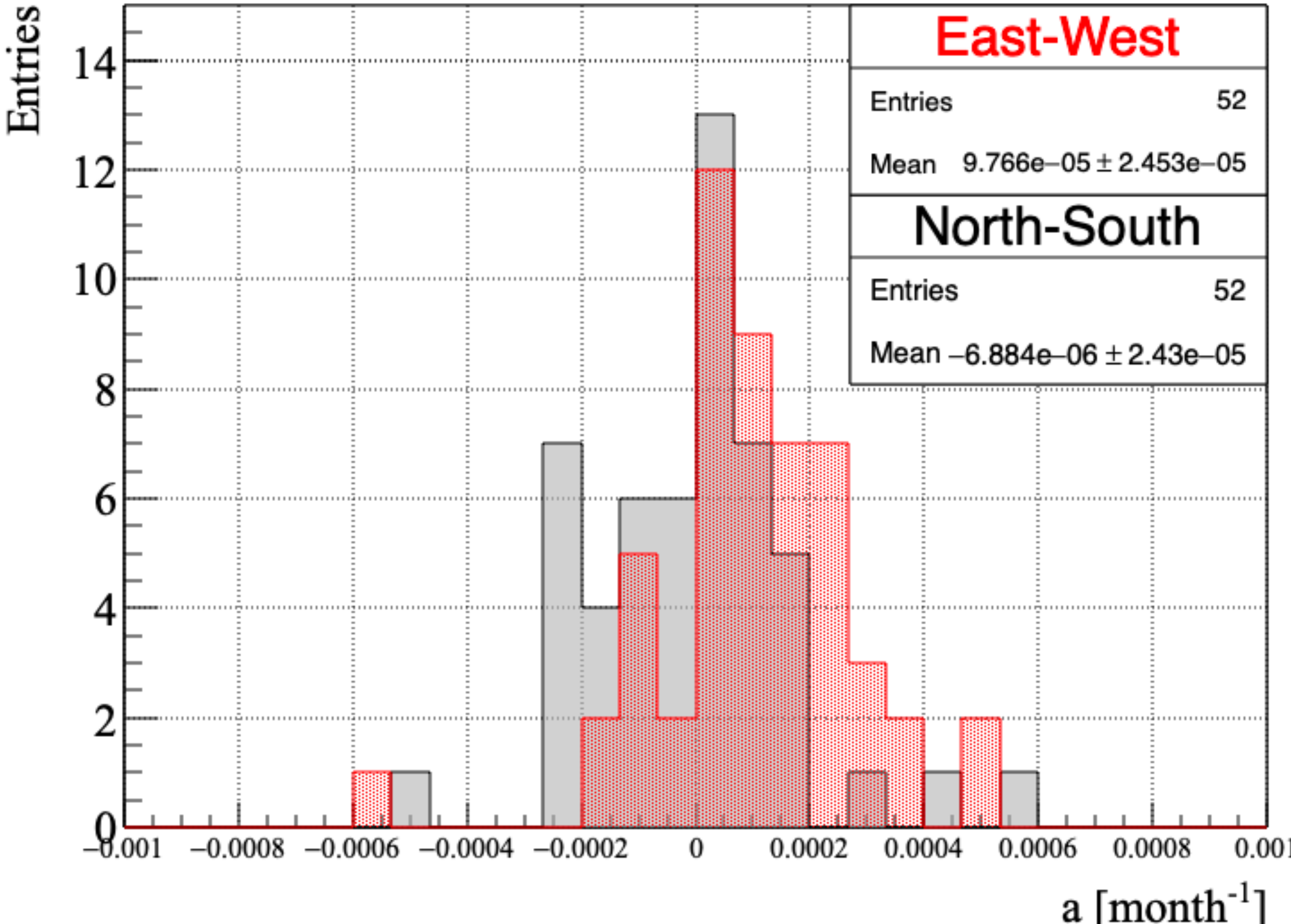
→ Perform a linear fit to check for possible ageing factor $C_0(t) = at + b$

Calibration constants as a function of the time (Id : 33)



Study of C_0 as a function of the time

→ Perform a linear fit to check for possible ageing factor $C_0(t) = at + b$

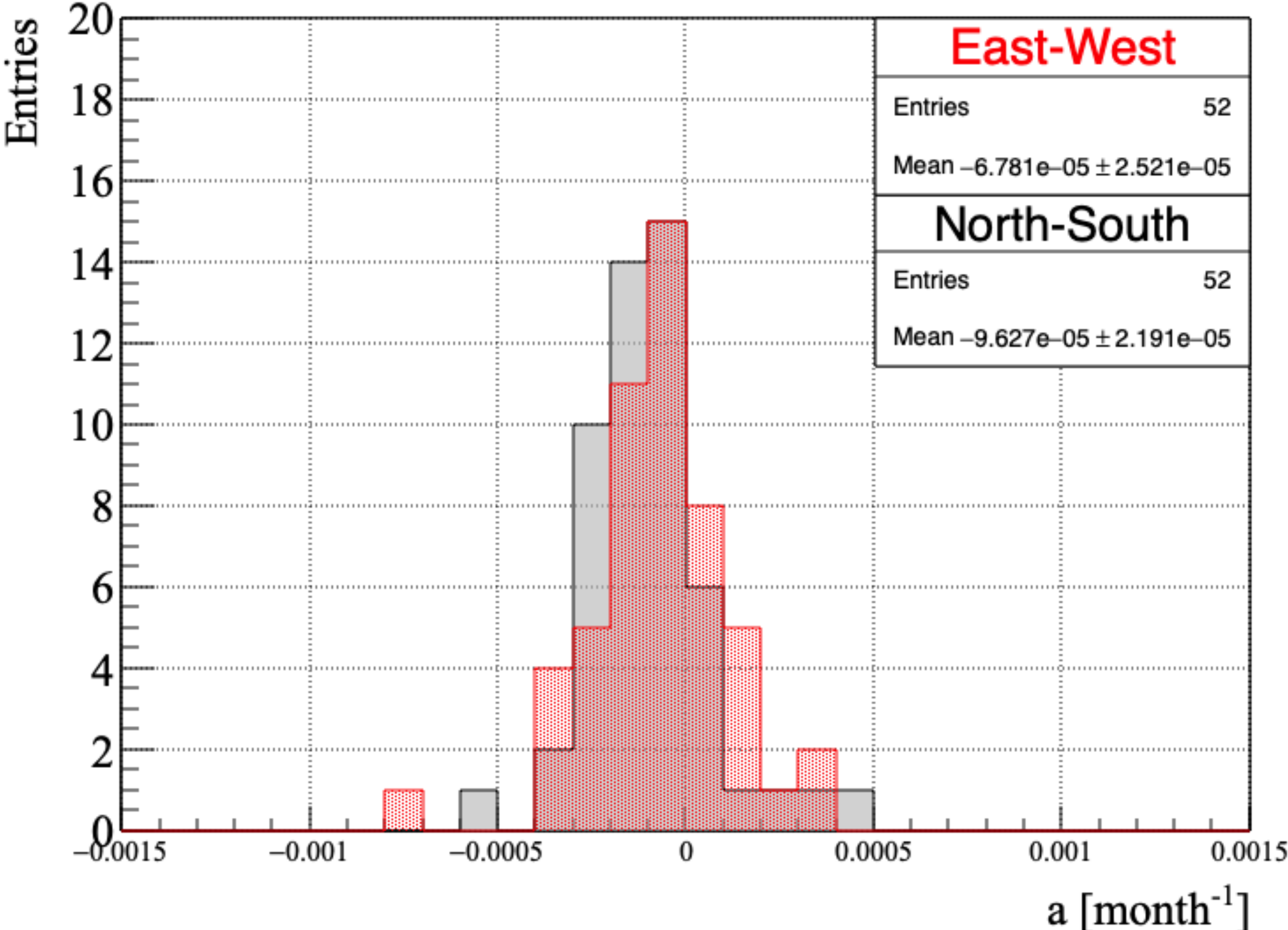


% per decade East-West Channel
$1.17 \pm 0.29 \%$
% per decade North-south Channel
$-0.08 \pm 0.29 \%$

Study of C_0 as a function of the time

Perform a linear fit to check for possible ageing factor $C_0(t) = at + b$

Removing points strongly fluctuated to smaller C_0 from 2014



**% per decade
East-West Channel**

$-0.80 \pm 0.30 \%$

**% per decade
North-south Channel**

$-1.10 \pm 0.26 \%$

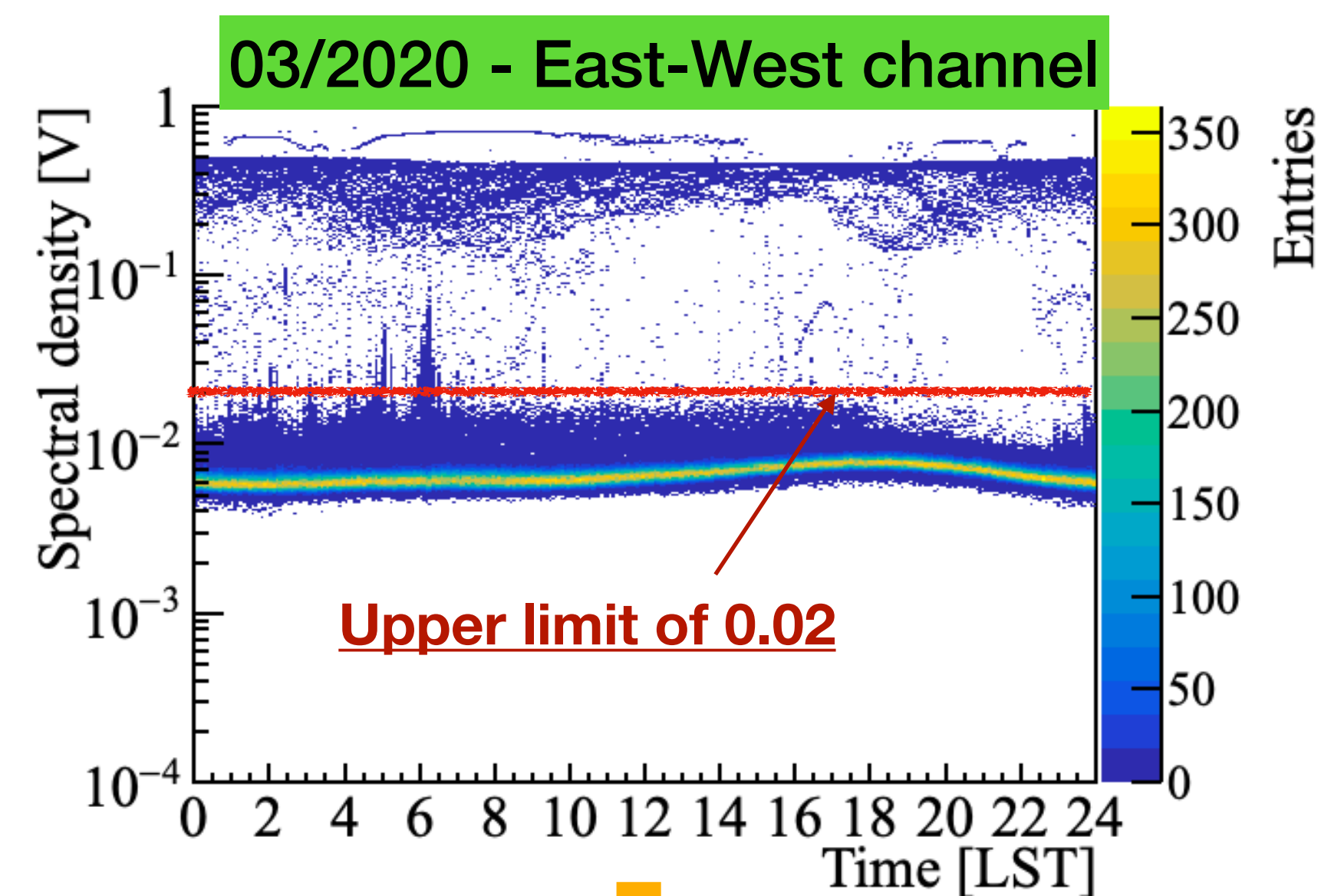
Conclusions

- We studied the behavior of the AERA stations calibration constants as a function of time from 2014 to 2020.
- Very small ageing effect (if exist) over nearly a decade.
 - Radio detectors could help to monitor possible ageing effects of other detector systems during long-term operations
 - Very important for determining an absolute energy scale

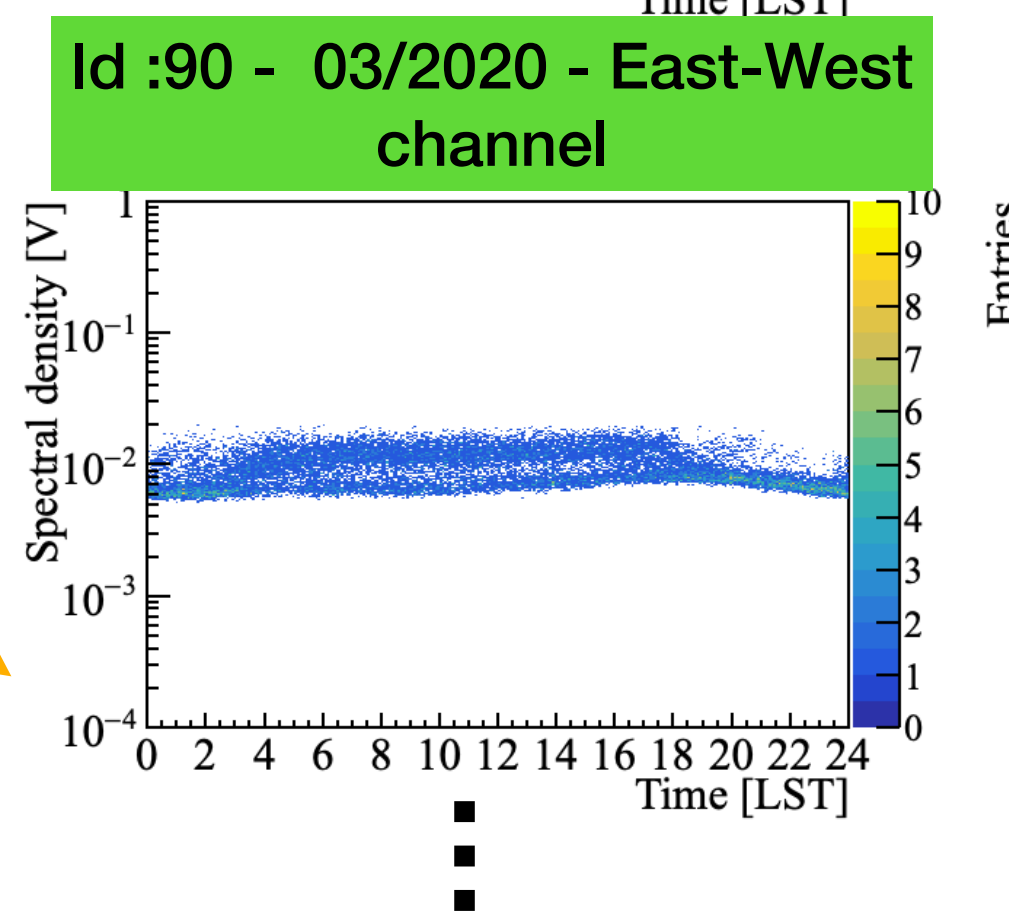
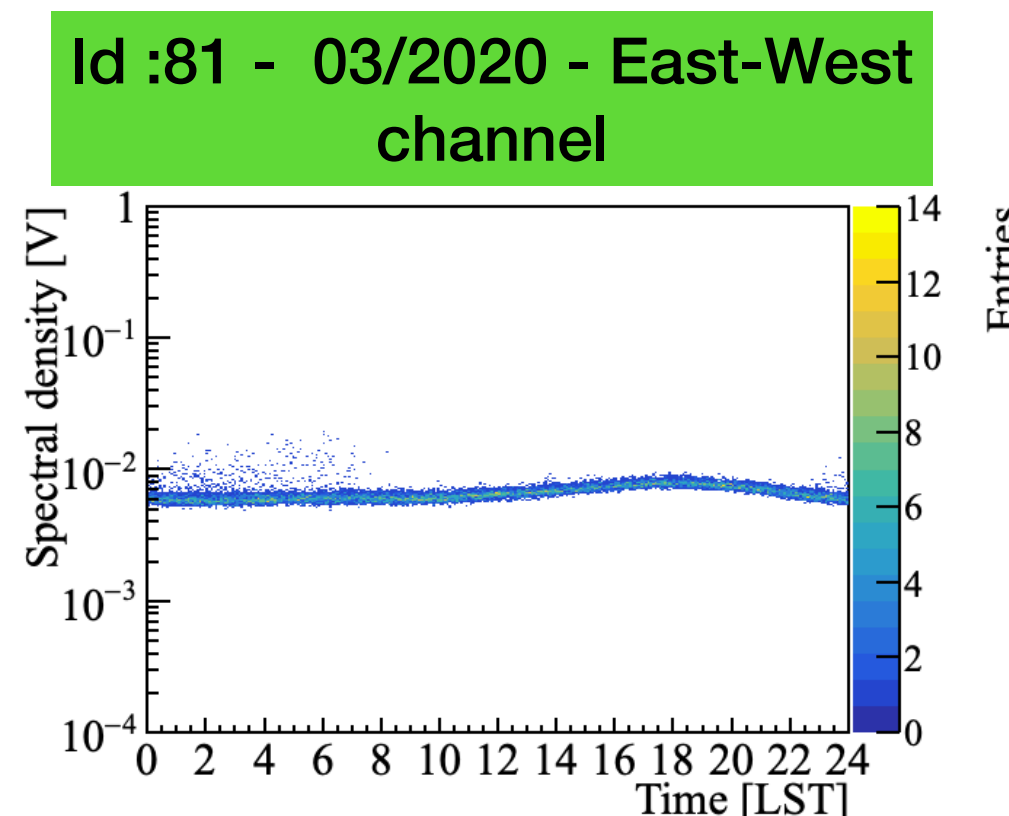
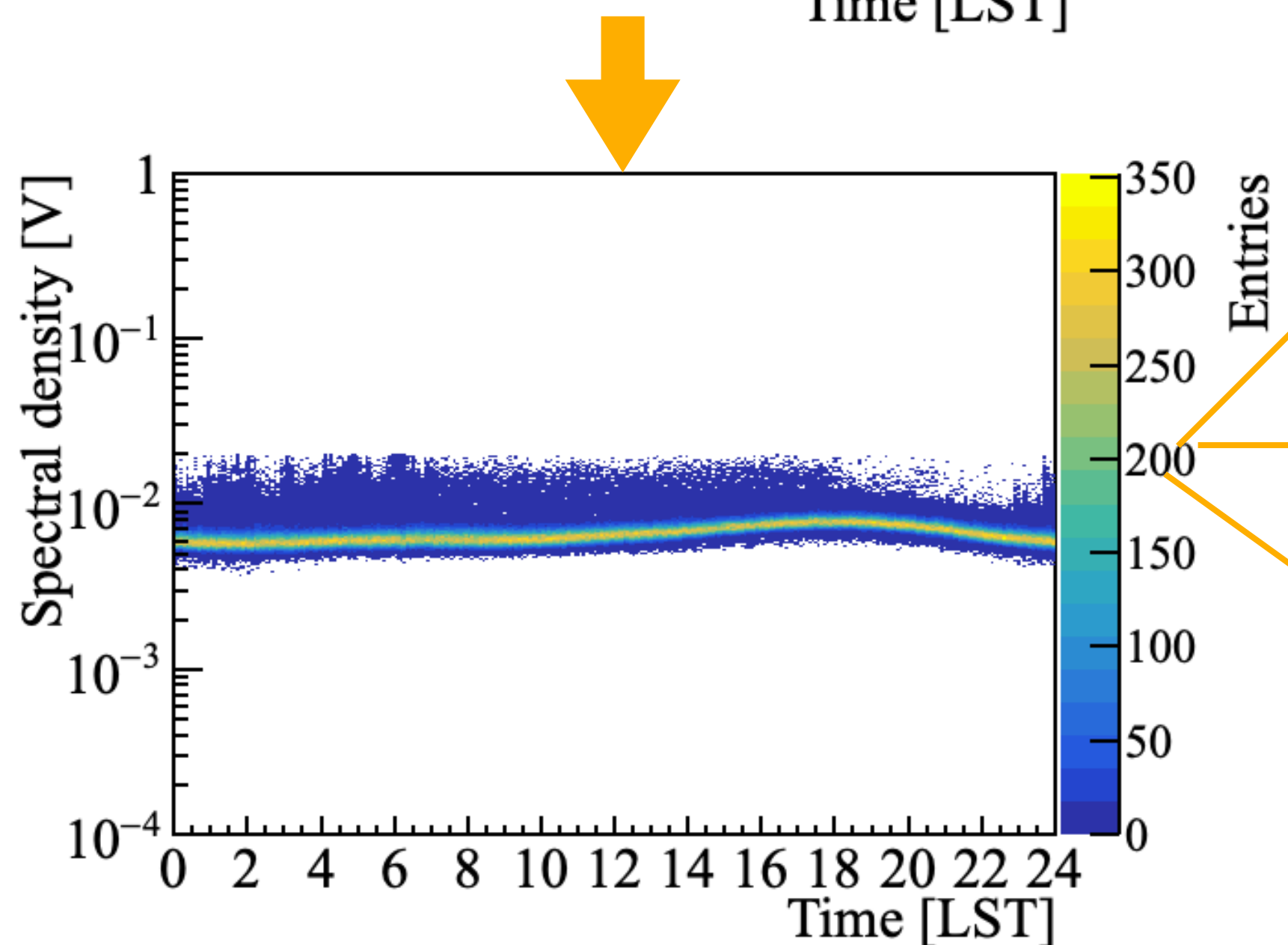
Backup

Broadband RFI

Procedure pipeline

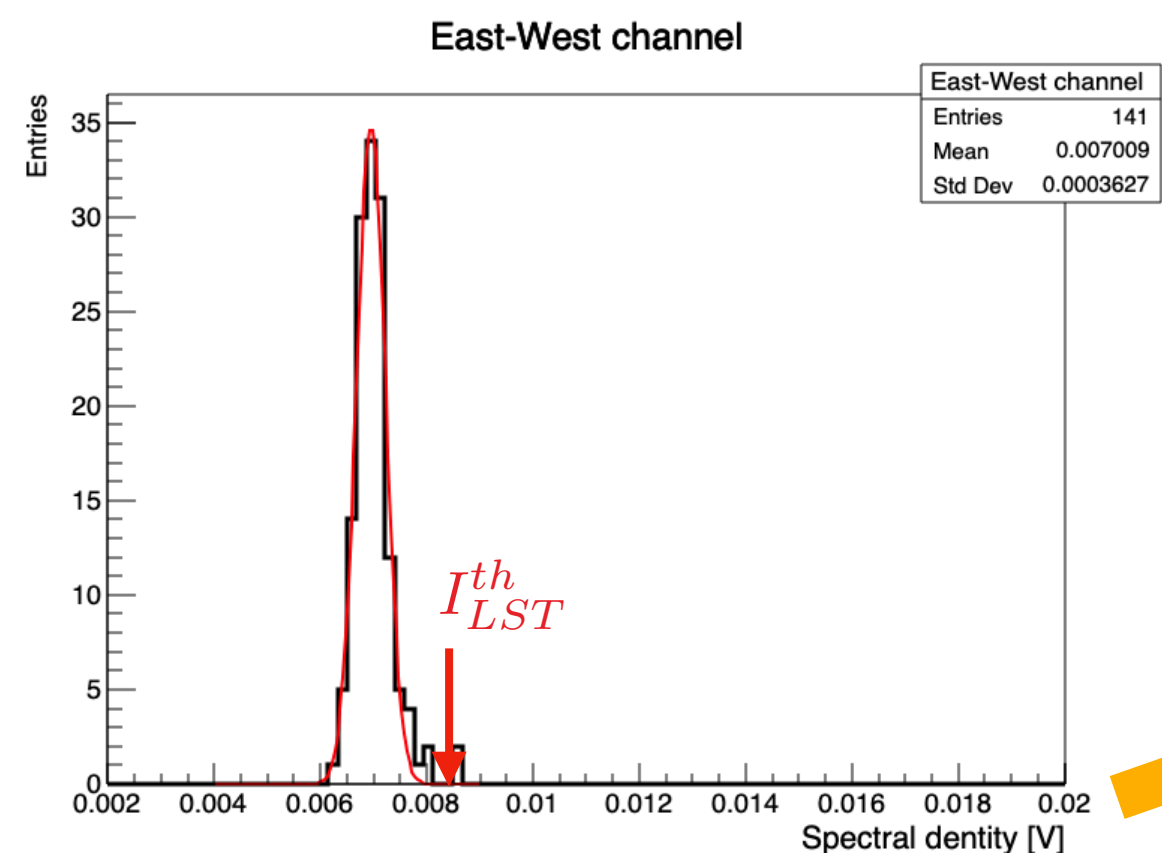
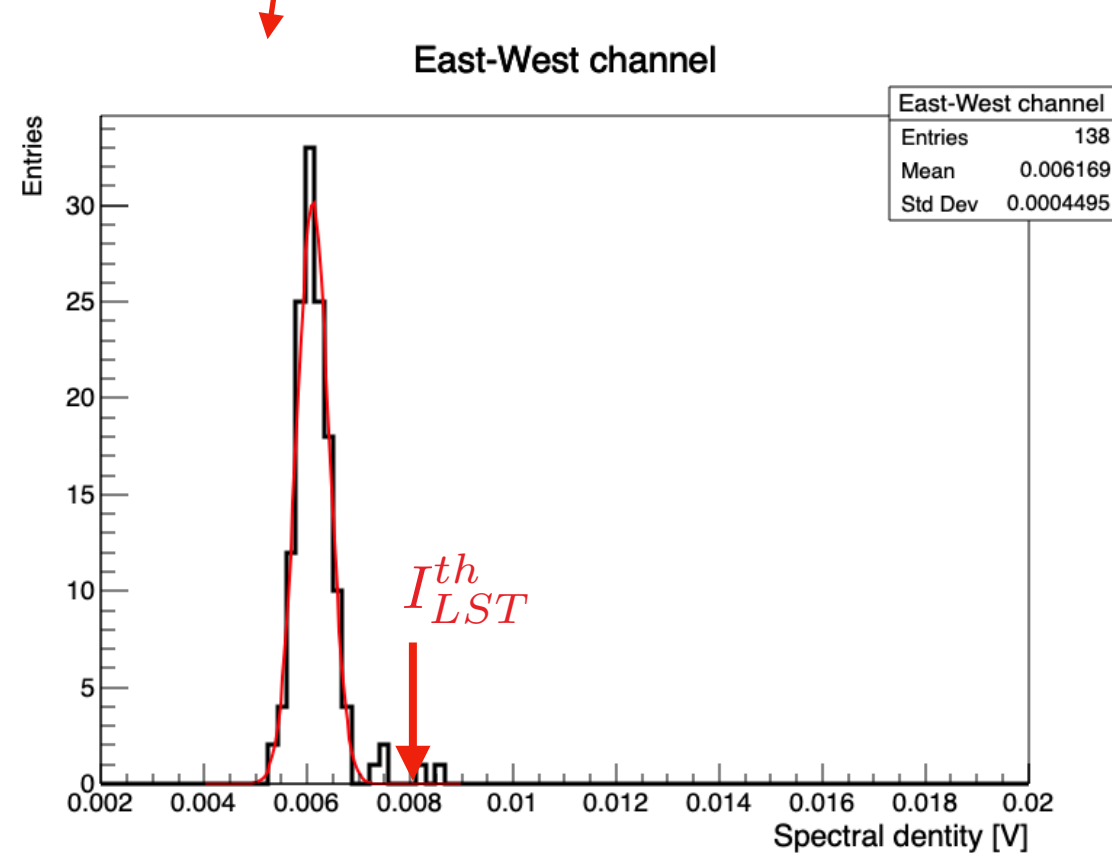
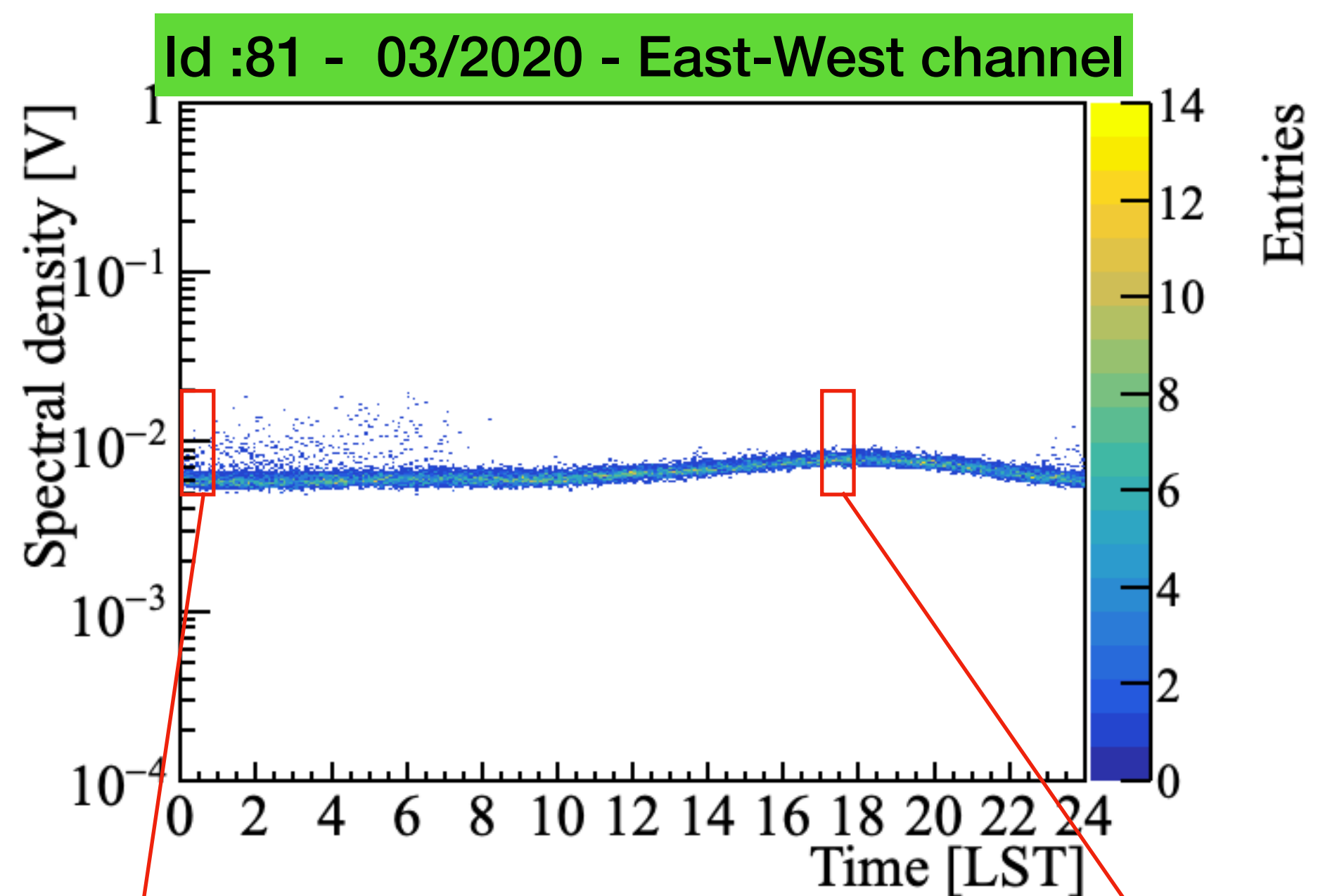


1. First, from the spectral density of all the antennas together, we reject signals larger than 0.02 V to avoid too high noise in our next steps/analyses.
2. We get the spectral density of each antenna and then we discretize the LST in 180 bins.



Broadband RFI

Procedure pipeline



3. We do a gaussian fit for the signal within each LST bin

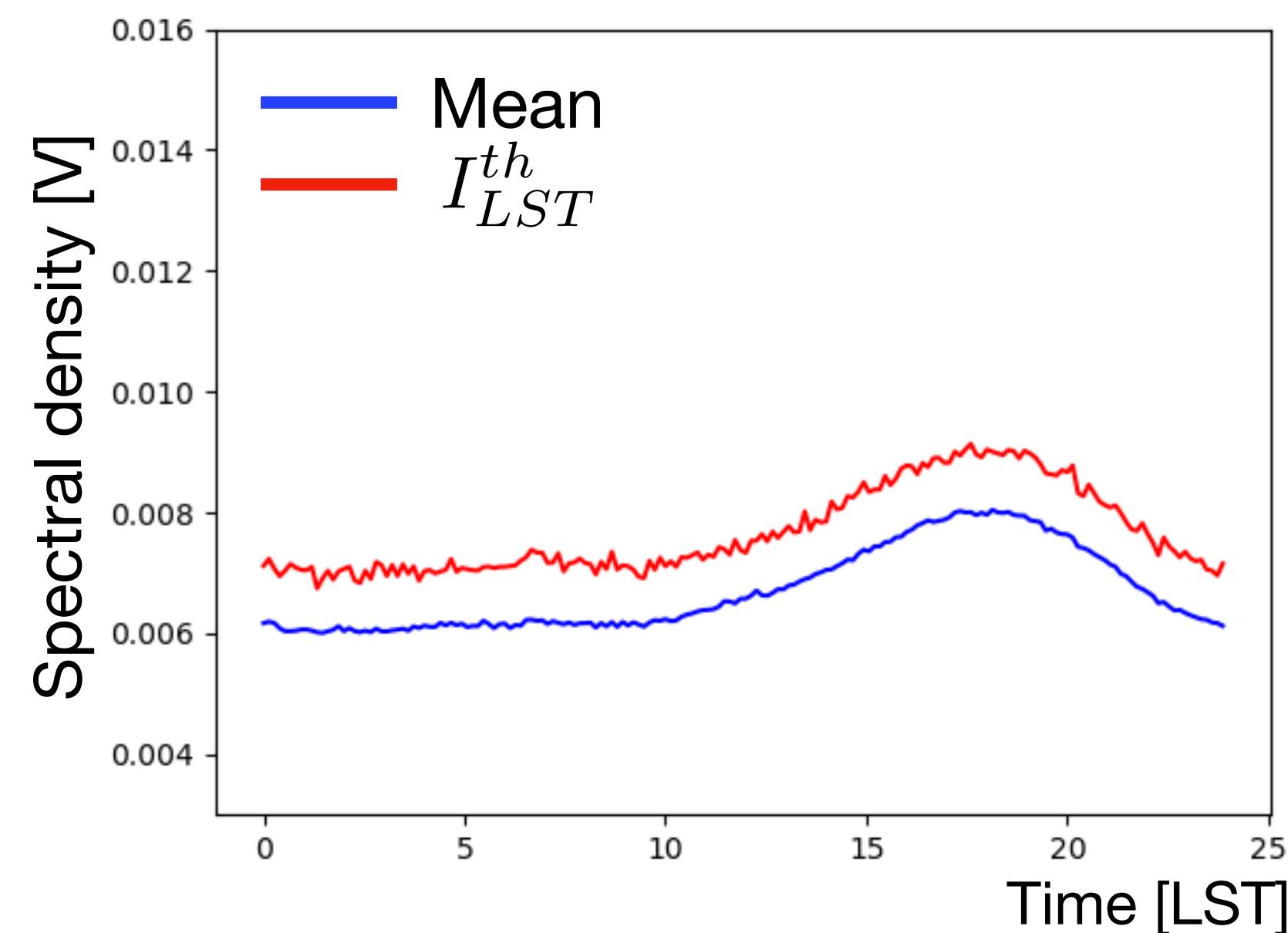
4. The threshold is defined according to the equation

$$\int_{-\infty}^{I_{LST}^{th}} G(I; \bar{I}, \sigma_I) dI = 0.9973$$

← Gaussian function

5. This procedure is repeated for all antennas and for both channels

Id:81 - 03/2020 - East-West channel

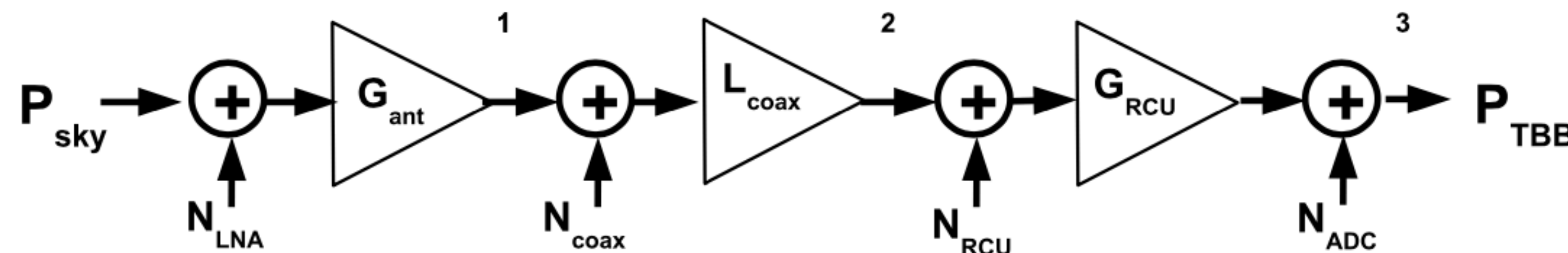


Galactic calibration: Nf model approach

Introduced by T.Fodran PoS (ICRC2021) 270

- Inspired in LOFAR calibration approach + frequency-dependent noise

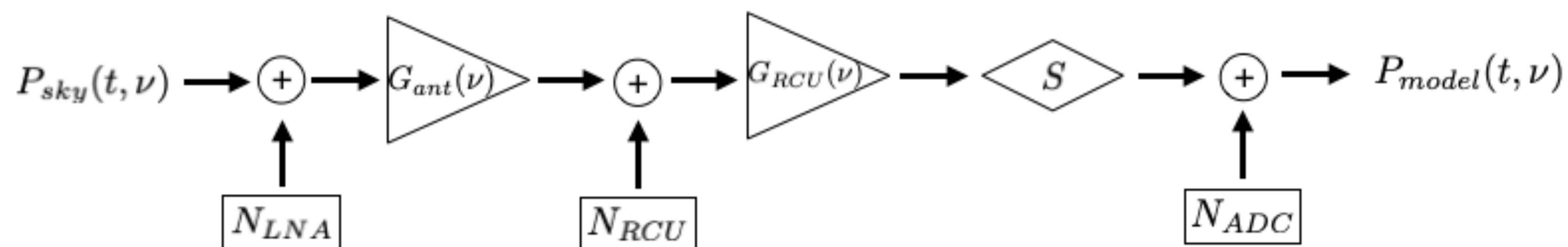
- LOFAR approach:** Electronic noise enters in the system at various stages of the signal chain



$$P_{model}(t, \nu) = [(P_{sky}(t, \nu) + N_{LNA})G_{ant}(\nu)A(\nu)L_{coax}(\nu) + N_{RCU}]G_{RCU}(\nu)S + N_{ADC}$$

Astroparticle Physics Vol. 111, September 2019, pages 1-11

- Accounting for the AERA antenna model



$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S + \underline{N_{LNA}G_{ant}(\nu)G_{RCU}(\nu)S + N_{RCU}G_{RCU}S} + N_{ADC}$$

↘ $N_{sys}(\nu)$

- Considering a possible external noise $N_{ext}(\nu)$

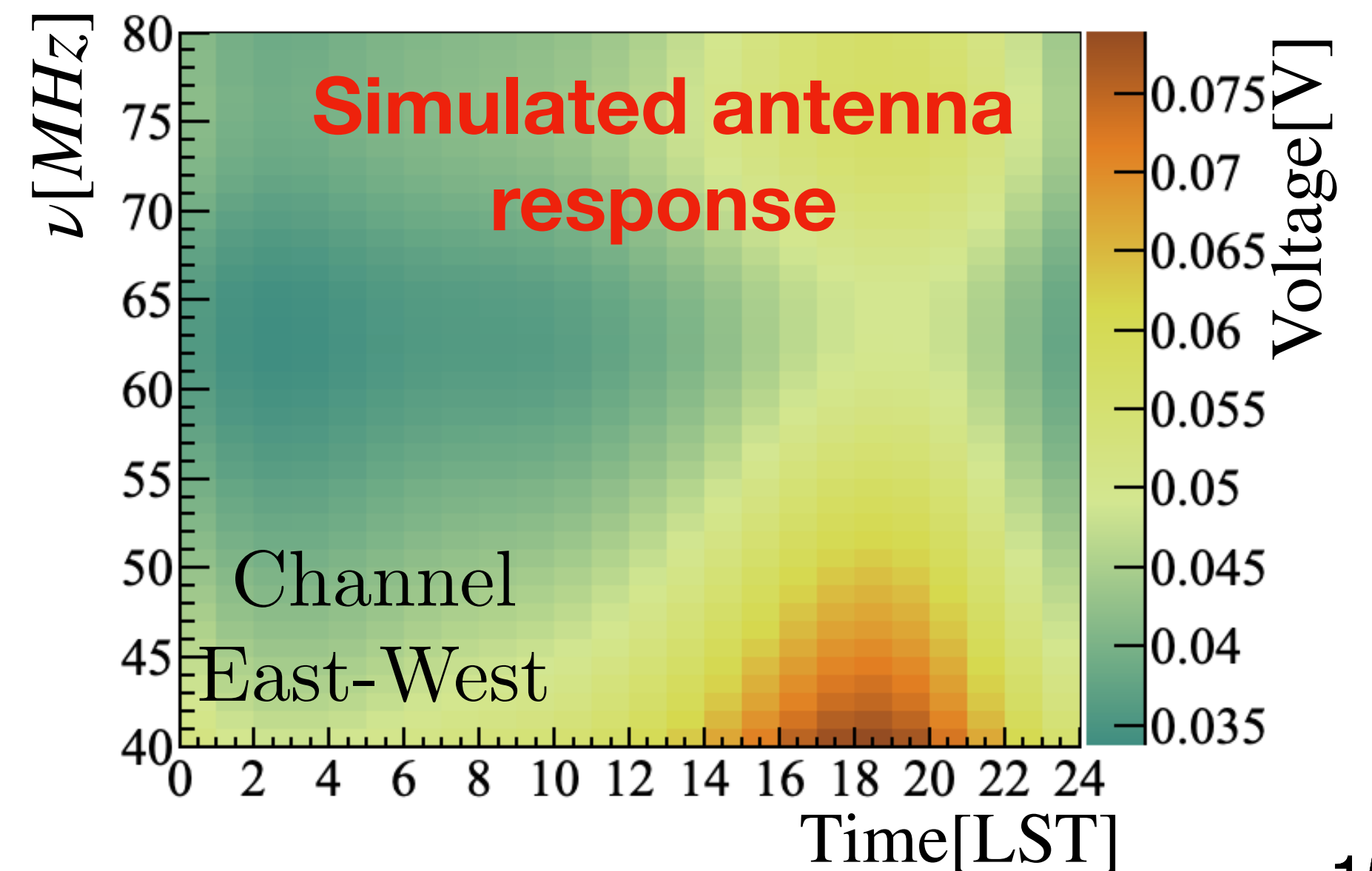
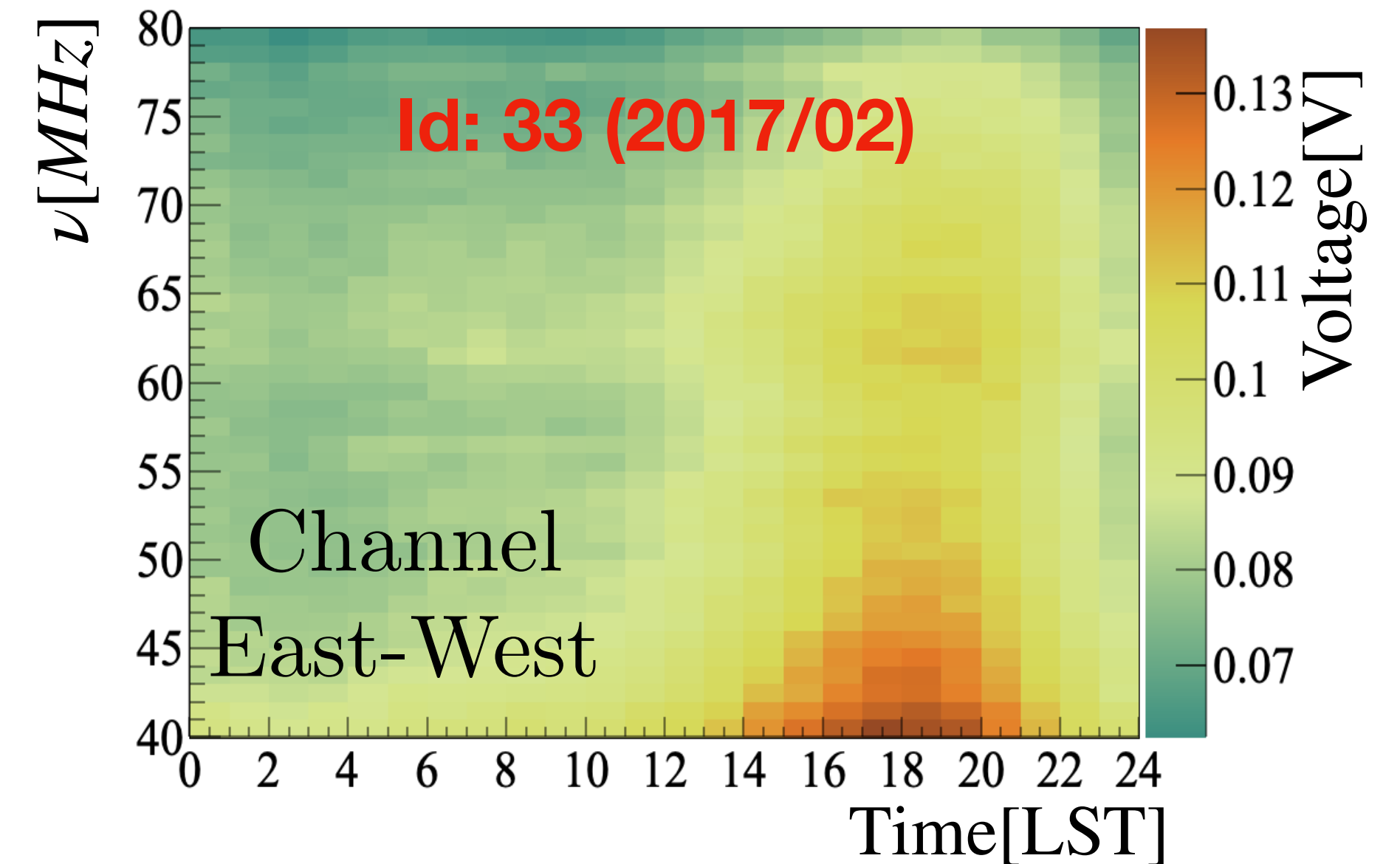
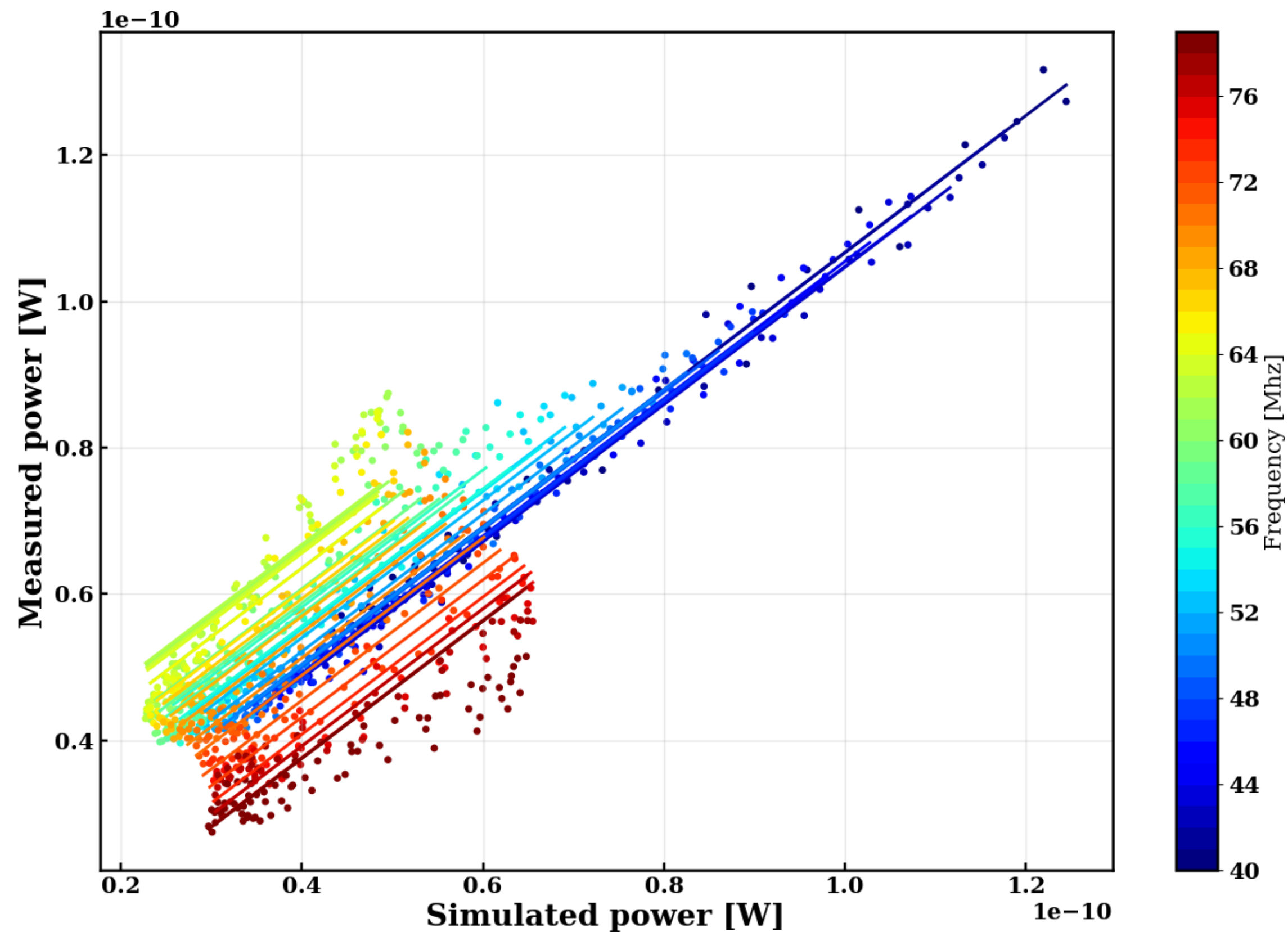
We can combine these terms into one single frequency-dependent term

So, $N_{tot}(\nu) = N_{ext}(\nu) + N_{sys}(\nu)$

$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S + N_{tot}(\nu)$$

Study of C_0 as a function of the time

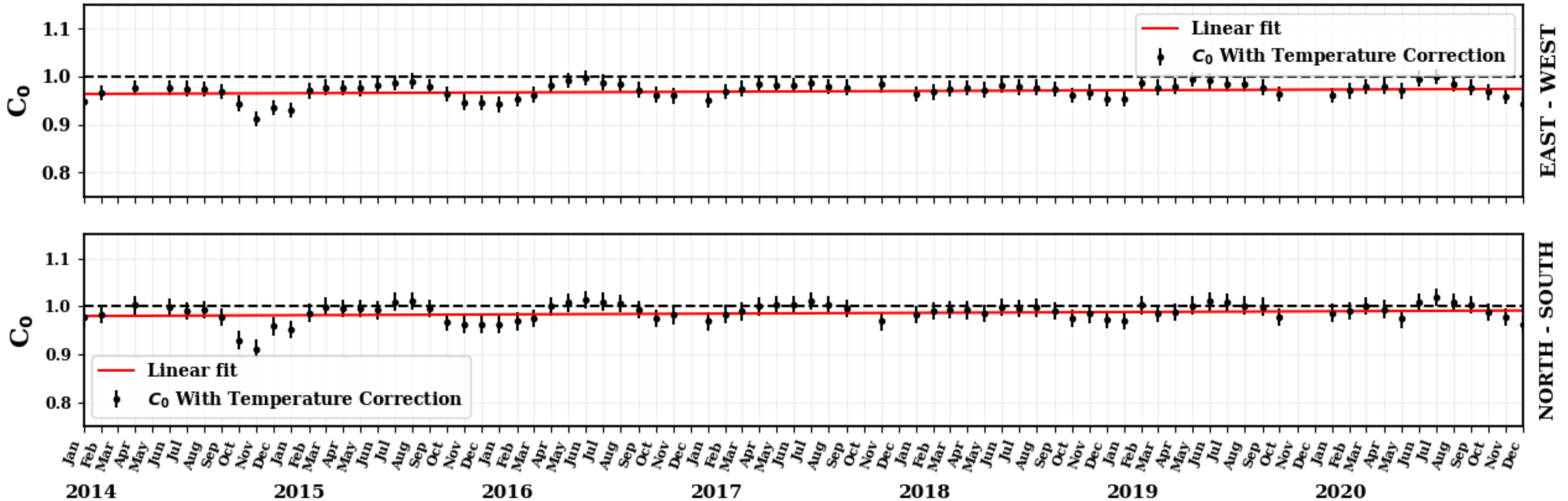
$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S + N_{tot}(\nu)$$



Study of C_0 as a function of the time

- Linear fit of all points to check ageing effect $C_0(t) = at + b$

Calibration constants as a function of the time (Id : 33)



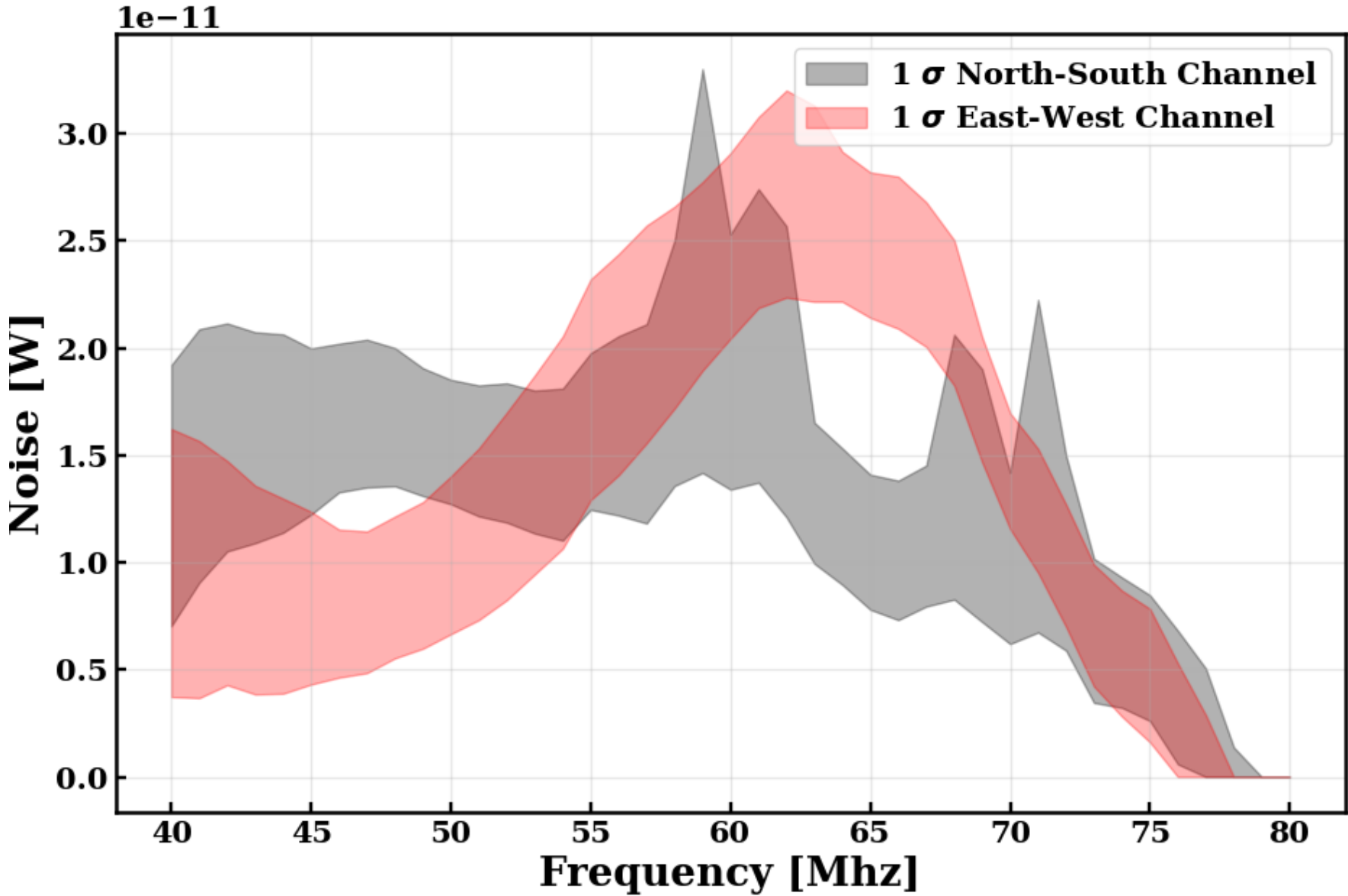
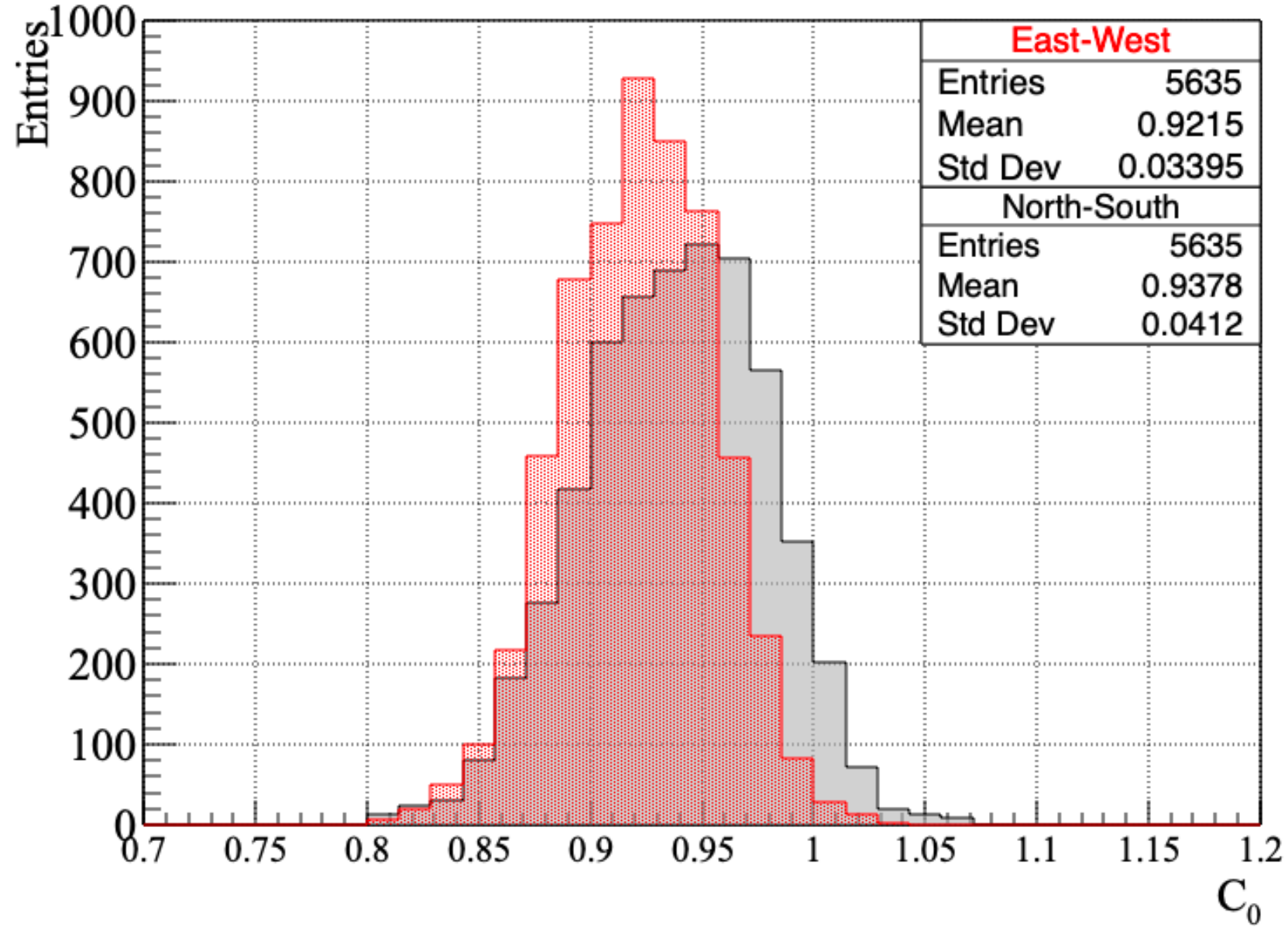
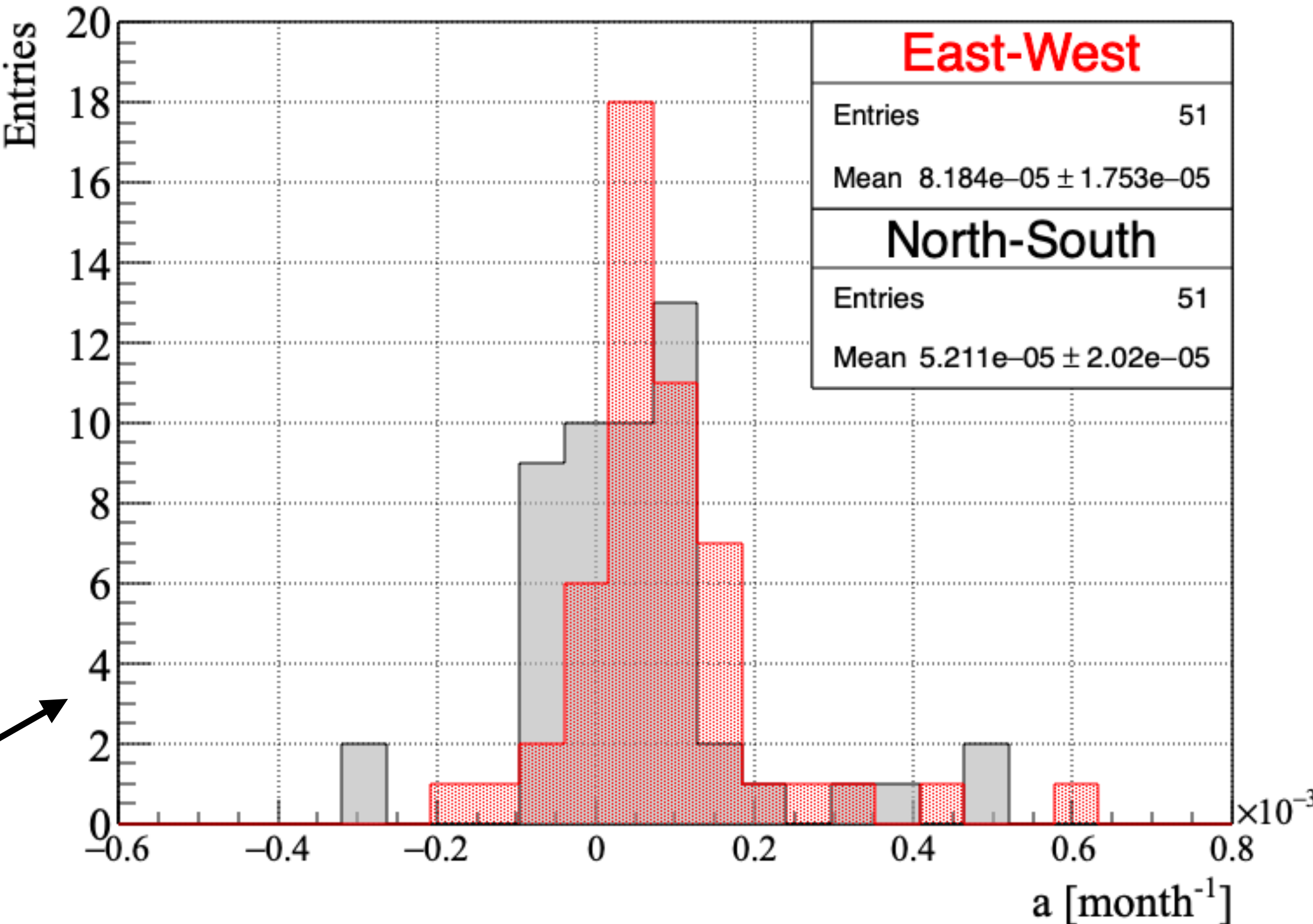
Study of C_0 as a function of the time

$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S + N_{tot}(\nu)$$

- Linear fit of all points to check ageing effect

$$C_0(t) = at + b$$

- 52 antennas with calibration constants from 2014-2020



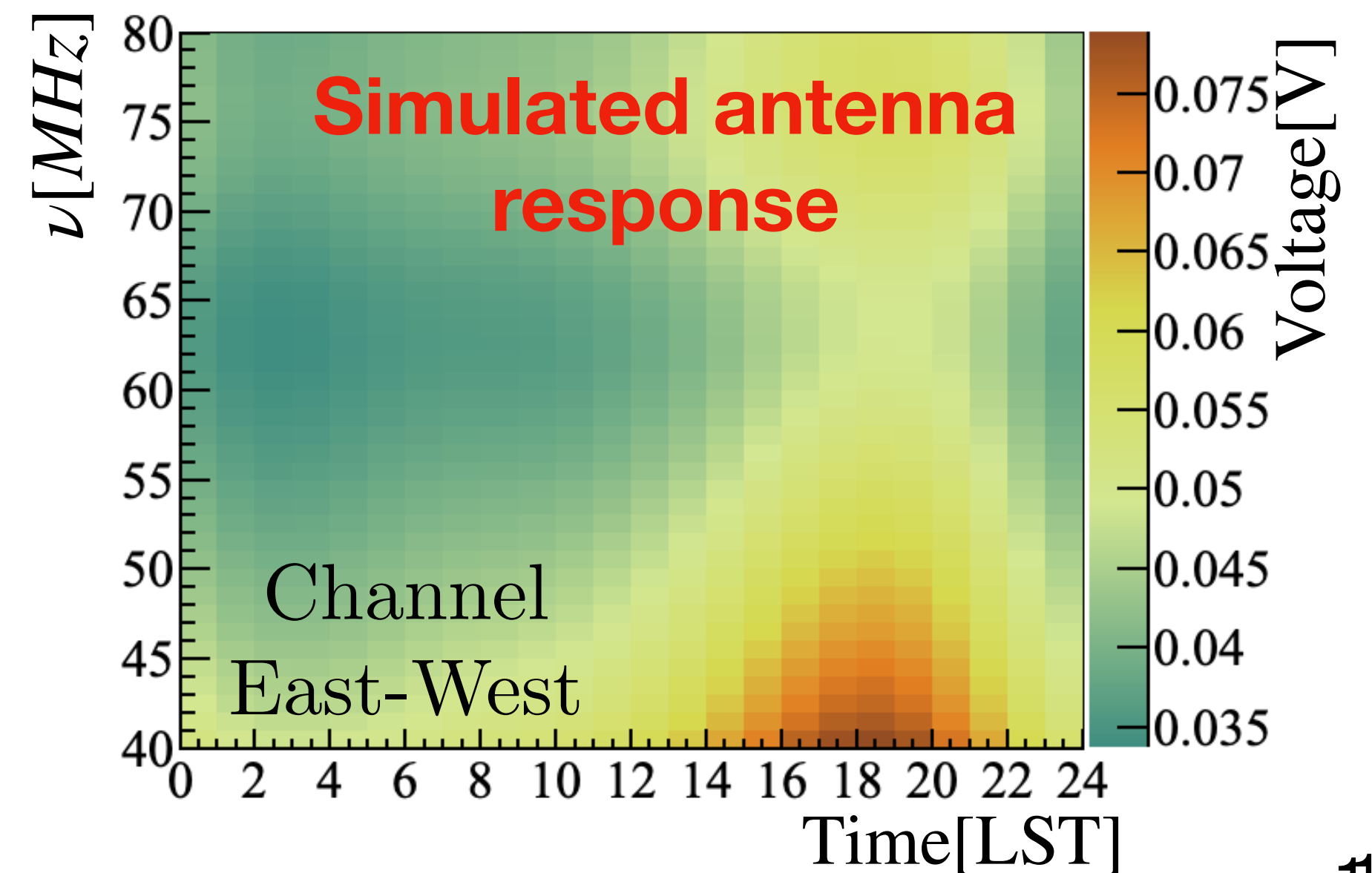
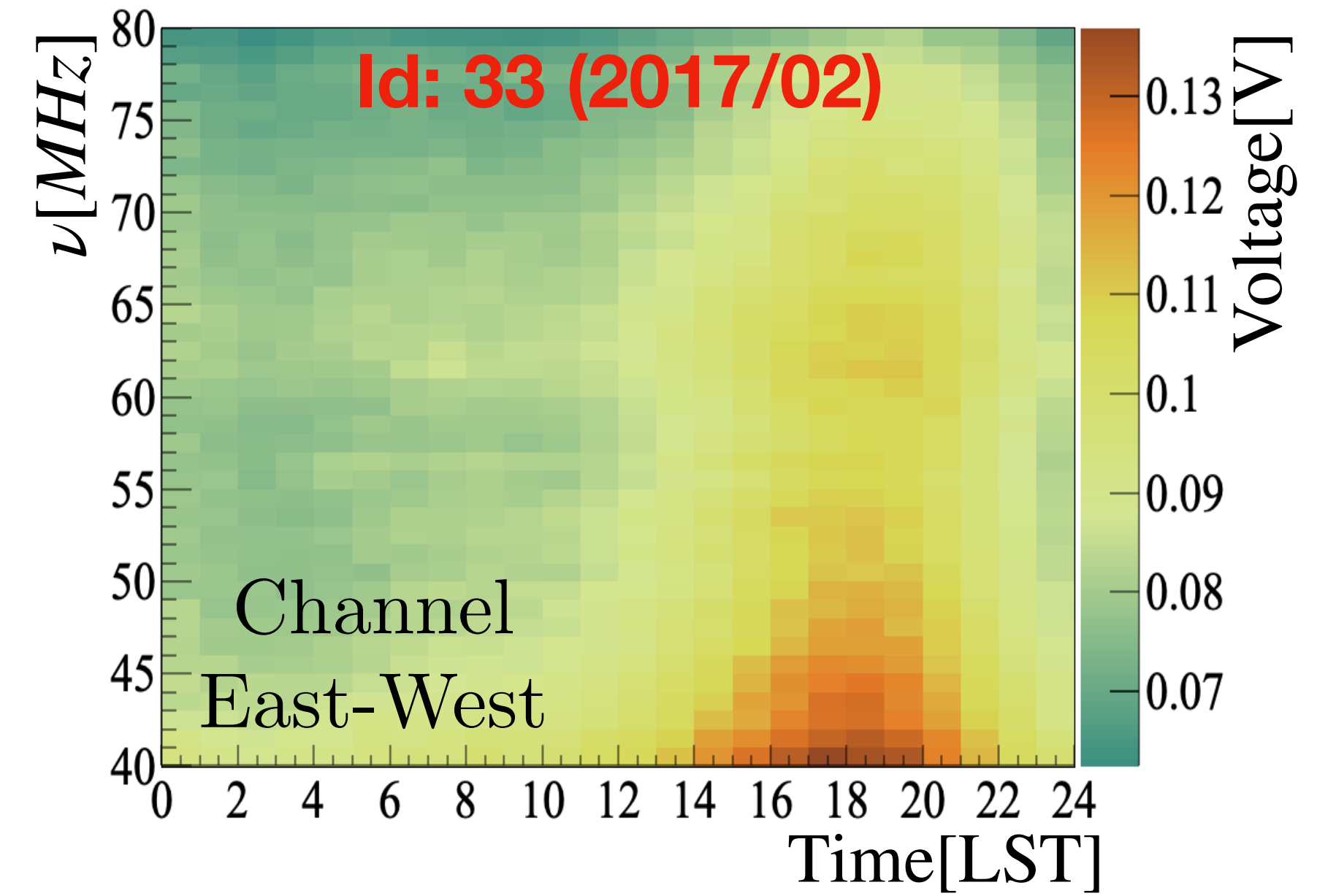
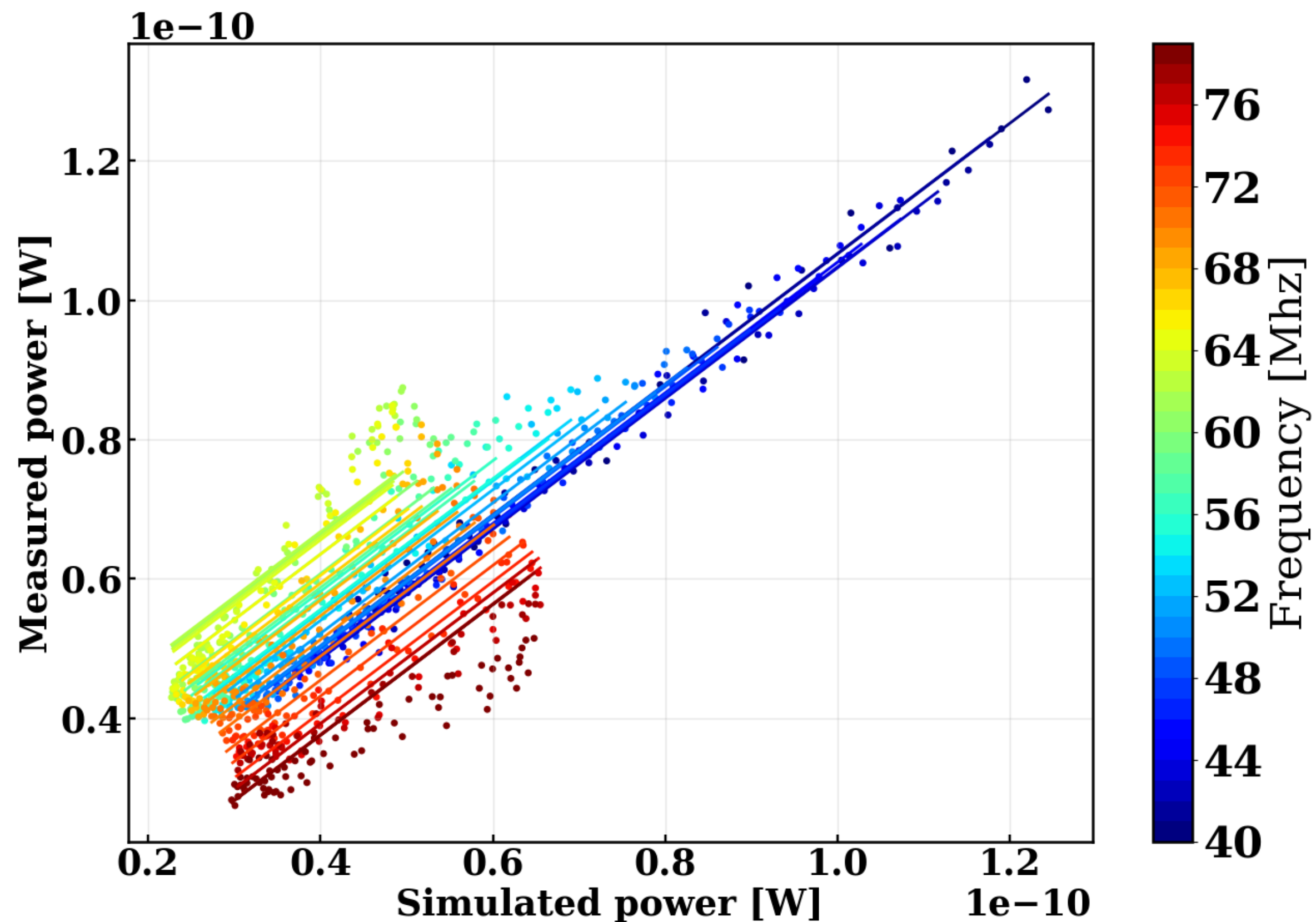
Ageing

% per decade East-West Channel
$0.98 \pm 0.20 \%$
% per decade North-south Channel
$0.62 \pm 0.24 \%$

Study of C_0 as a function of the time

$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S + N_{tot}(\nu)$$

Removing points strongly
fluctuated to smaller C_0
from 2014

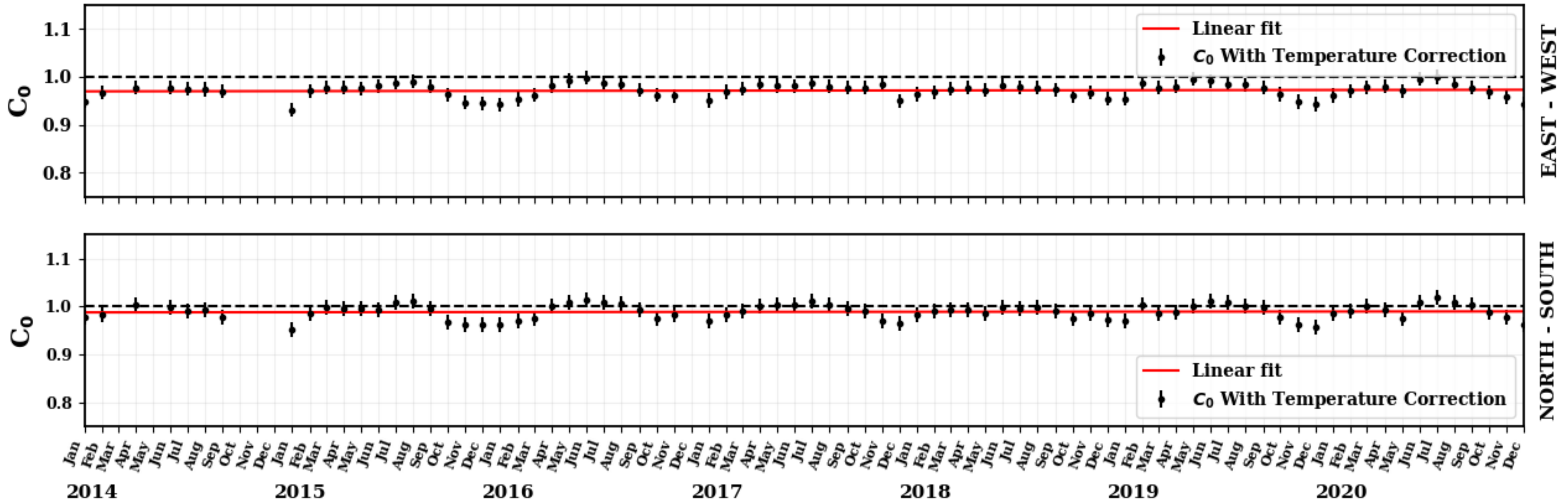


Study of C_0 as a function of the time

Removing points strongly fluctuated to smaller C_0 from 2014

- Linear fit of all points to check ageing effect $C_0(t) = at + b$

Calibration constants as a function of the time (Id : 33)



Study of C_0 as a function of the time

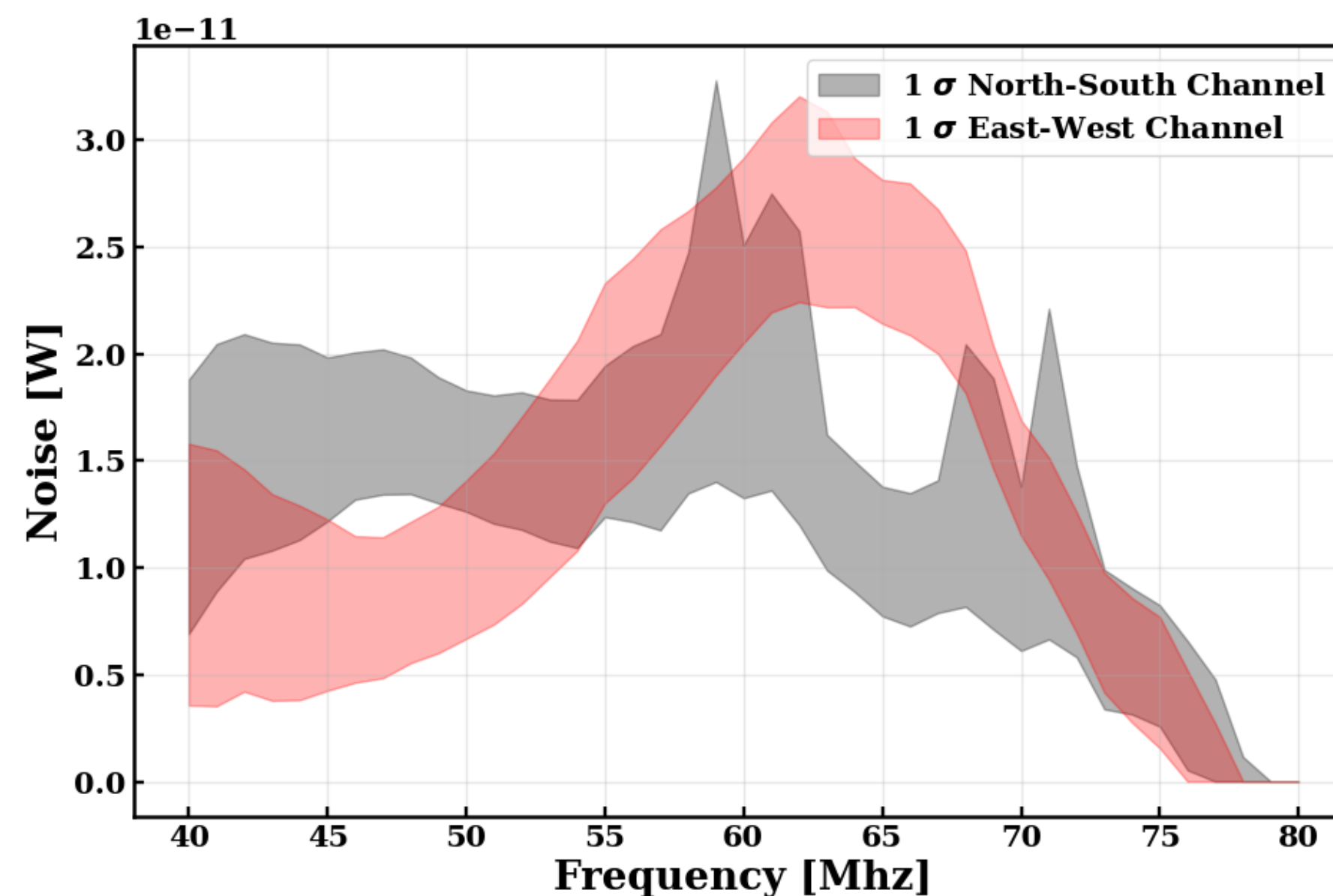
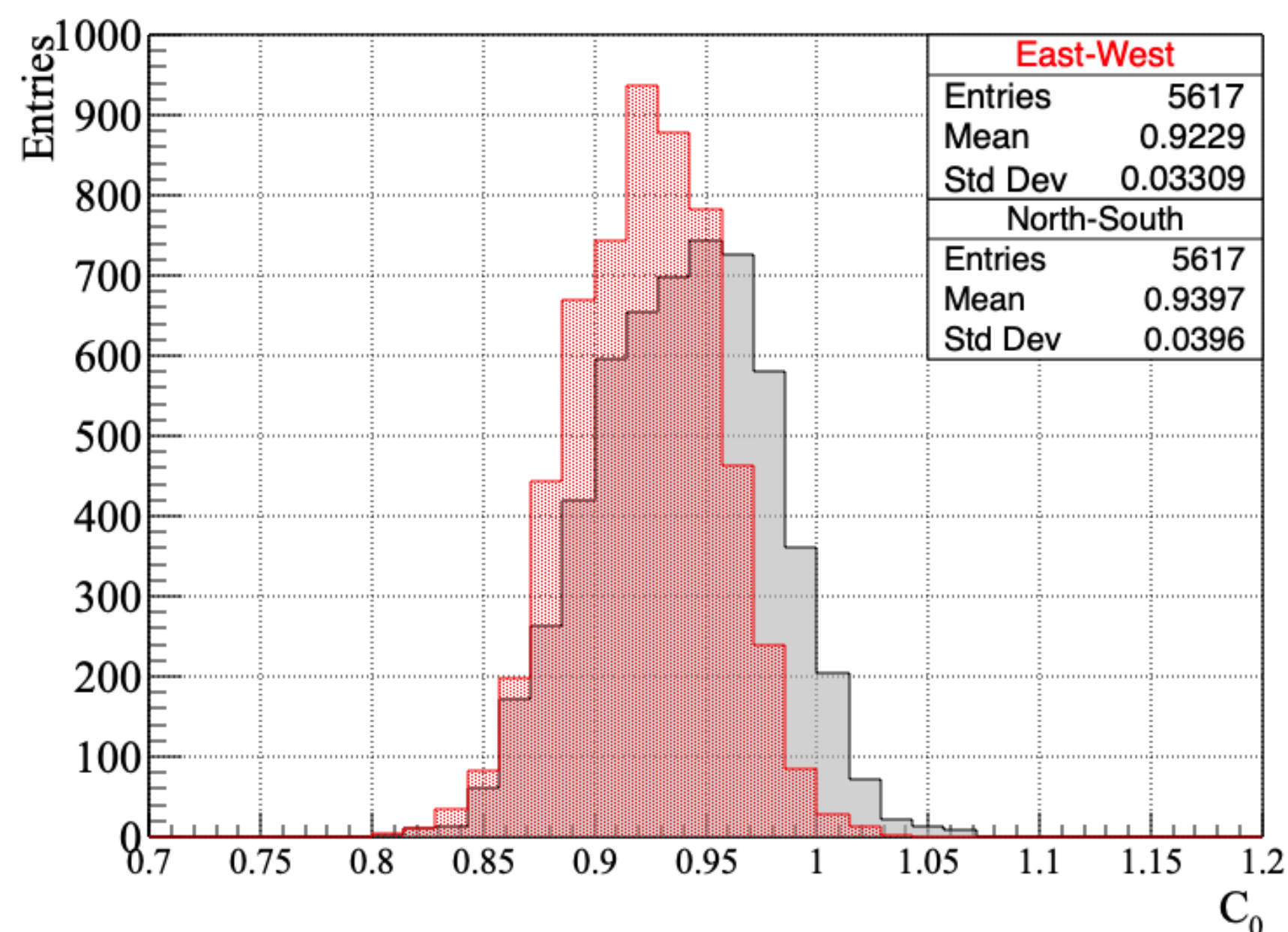
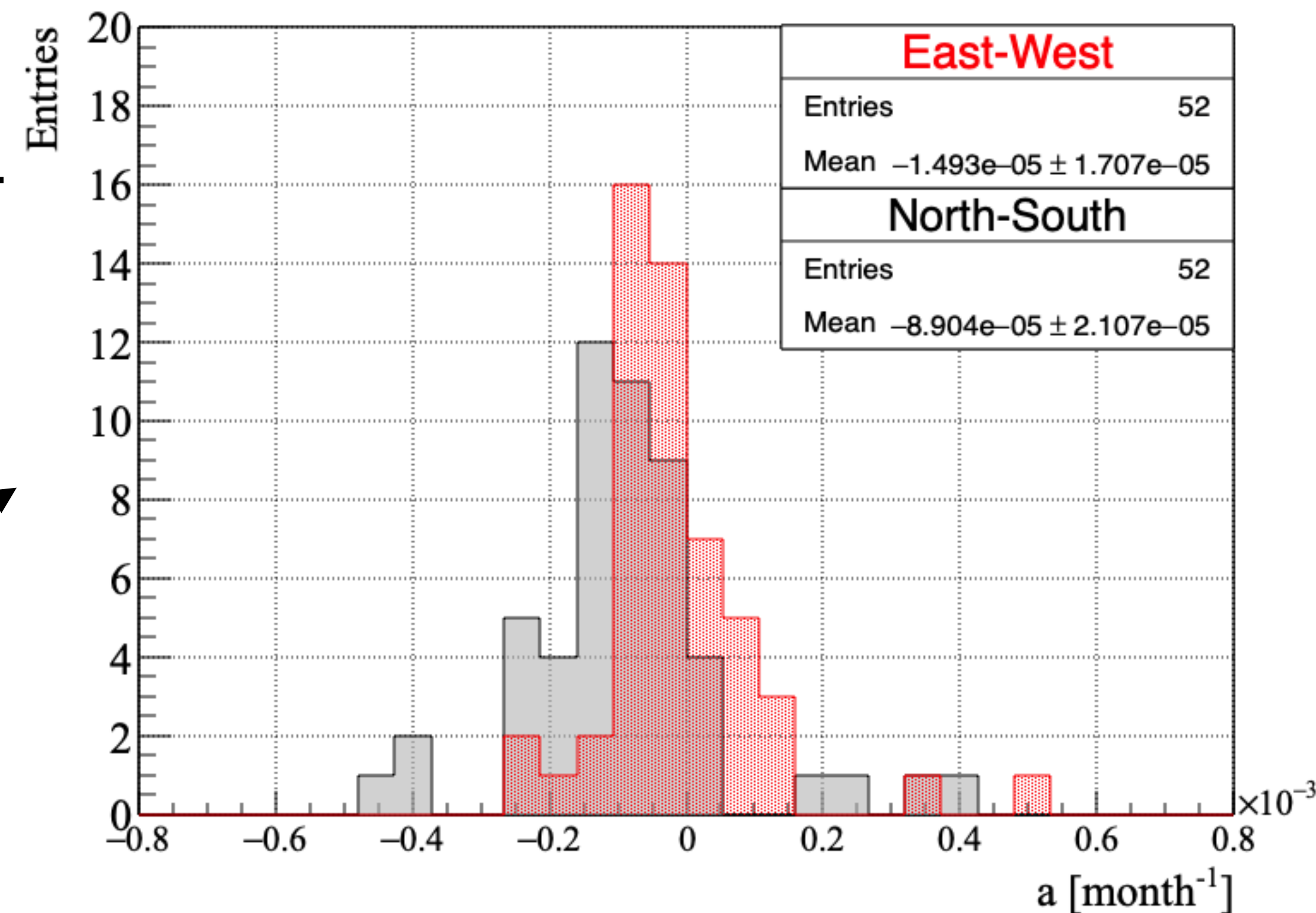
$$P_{model}(t, \nu) = P_{sky}(t, \nu)G_{ant}(\nu)G_{RCU}(\nu)S + N_{tot}(\nu)$$

- Linear fit of all points to check ageing effect

$$C_0(t) = at + b$$

- 52 antennas with calibration constants from 2014-2020

Removing points strongly fluctuated to smaller C_0 from 2014

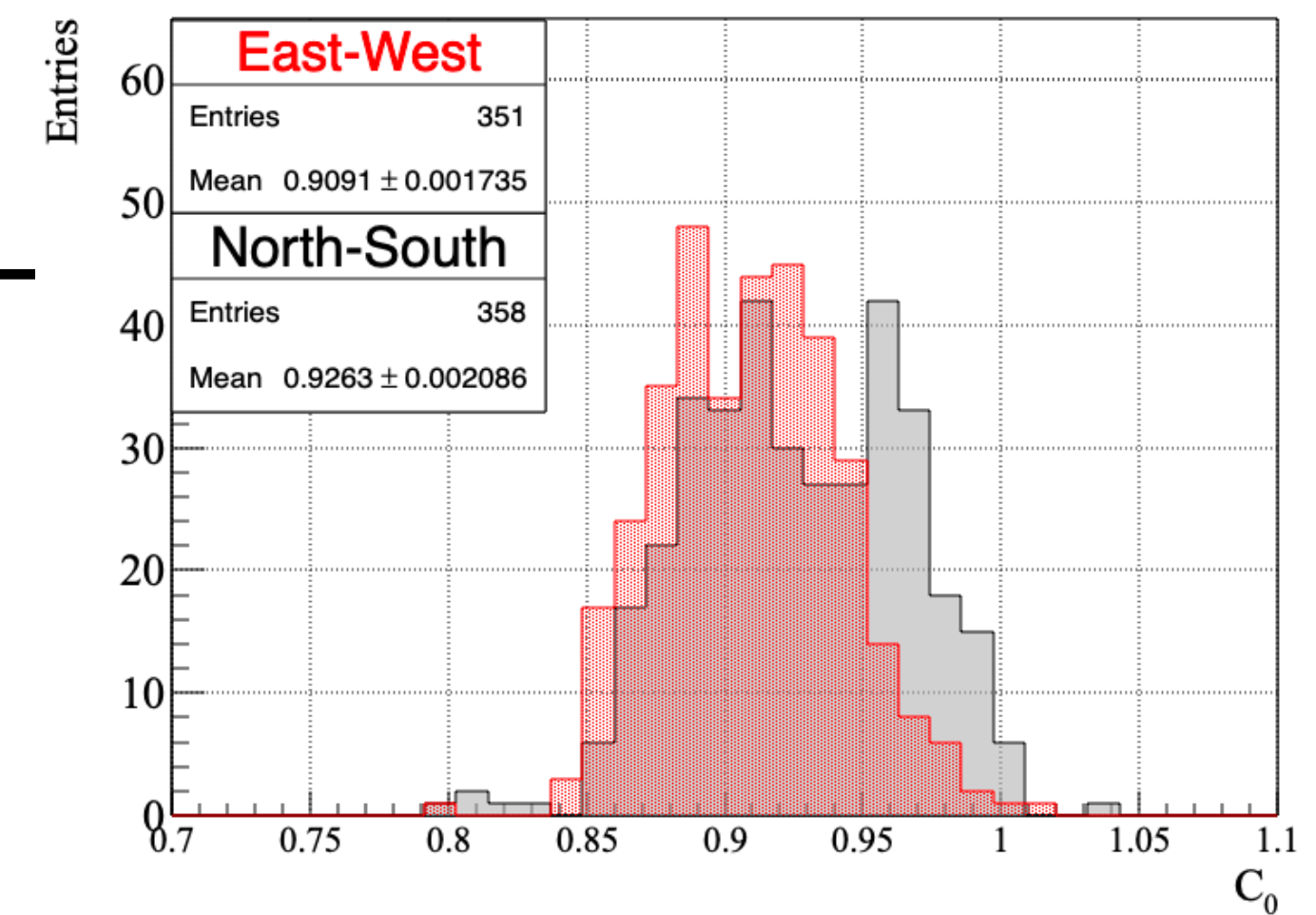


Ageing

% per decade East-West Channel
$-0.17 \pm 0.20 \%$
% per decade North-south Channel
$-1.07 \pm 0.25 \%$

Study of temperature x C_0

- Spectrum of each antenna obtained considering the whole year of 2019
- Ambient temperature used (CRS)

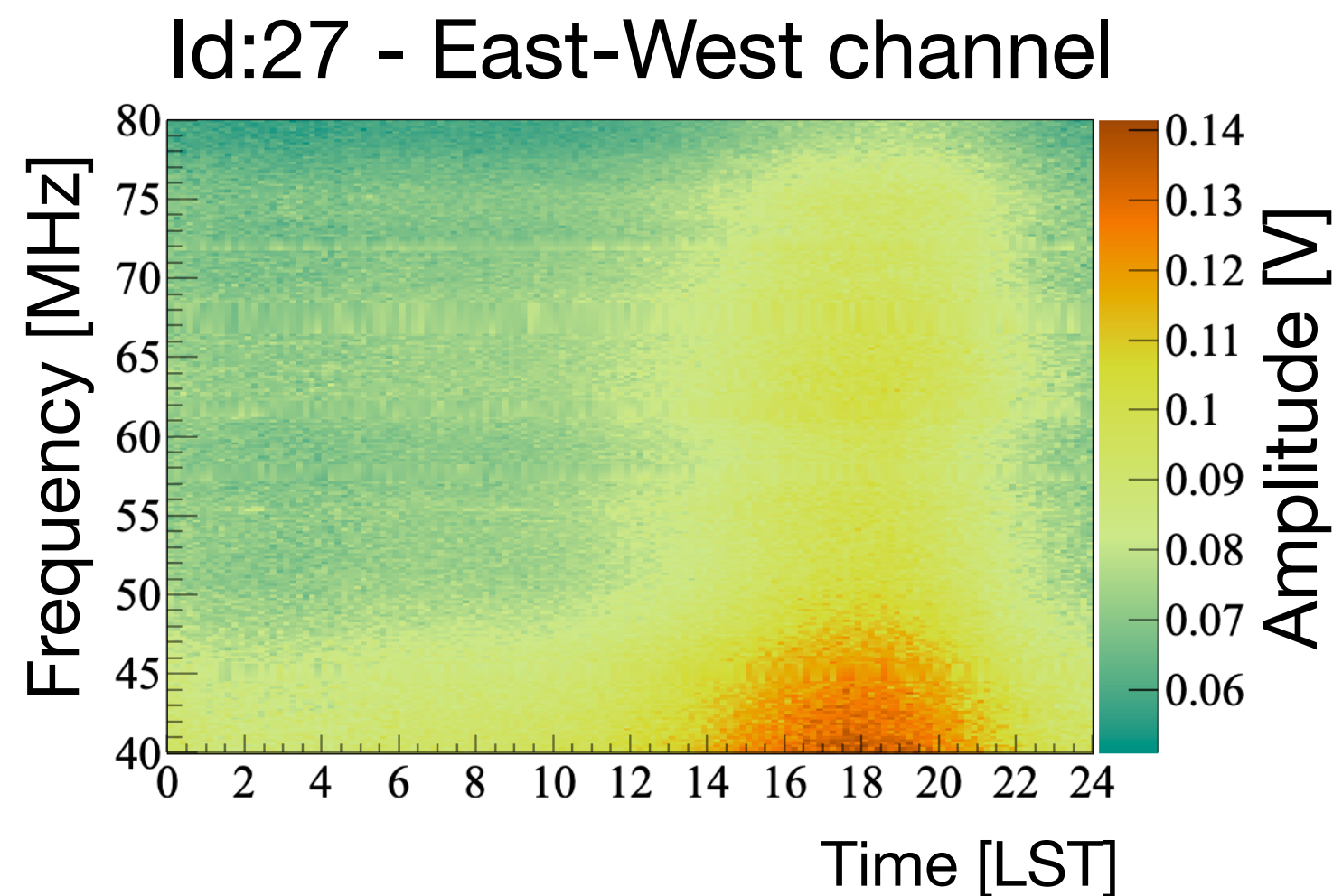


Procedure

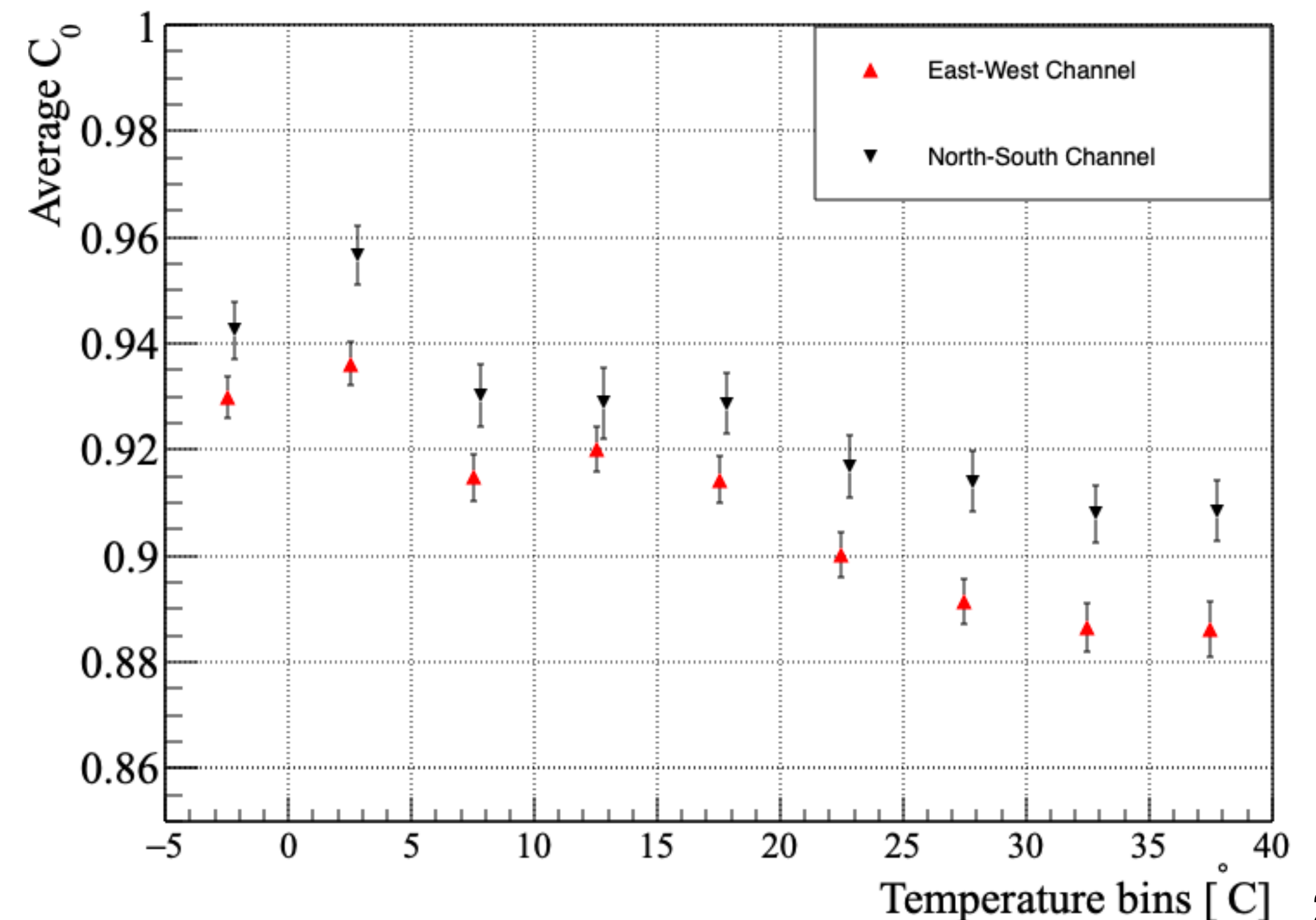
Fix a temperature bin (ex -5 to 0 degree)

In a whole year (2019) it only keeps traces when the temperature is within the temperature bin

★ Temperature correction is used



Calibration



Study of Irradiance x C_0

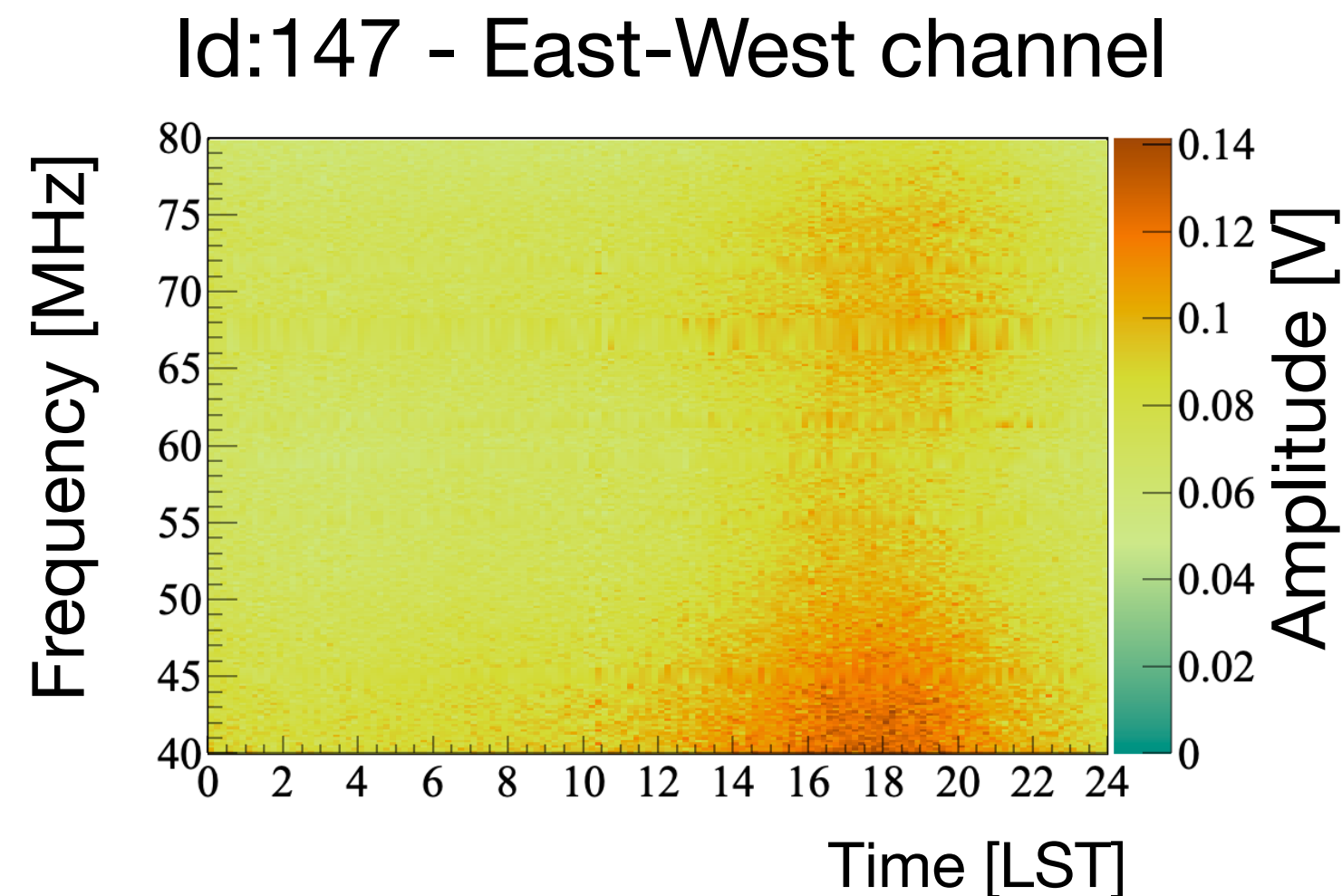
- Spectrum of each antenna obtained considering the whole year of 2020
- Irradiance data used

★ Temperature correction is used

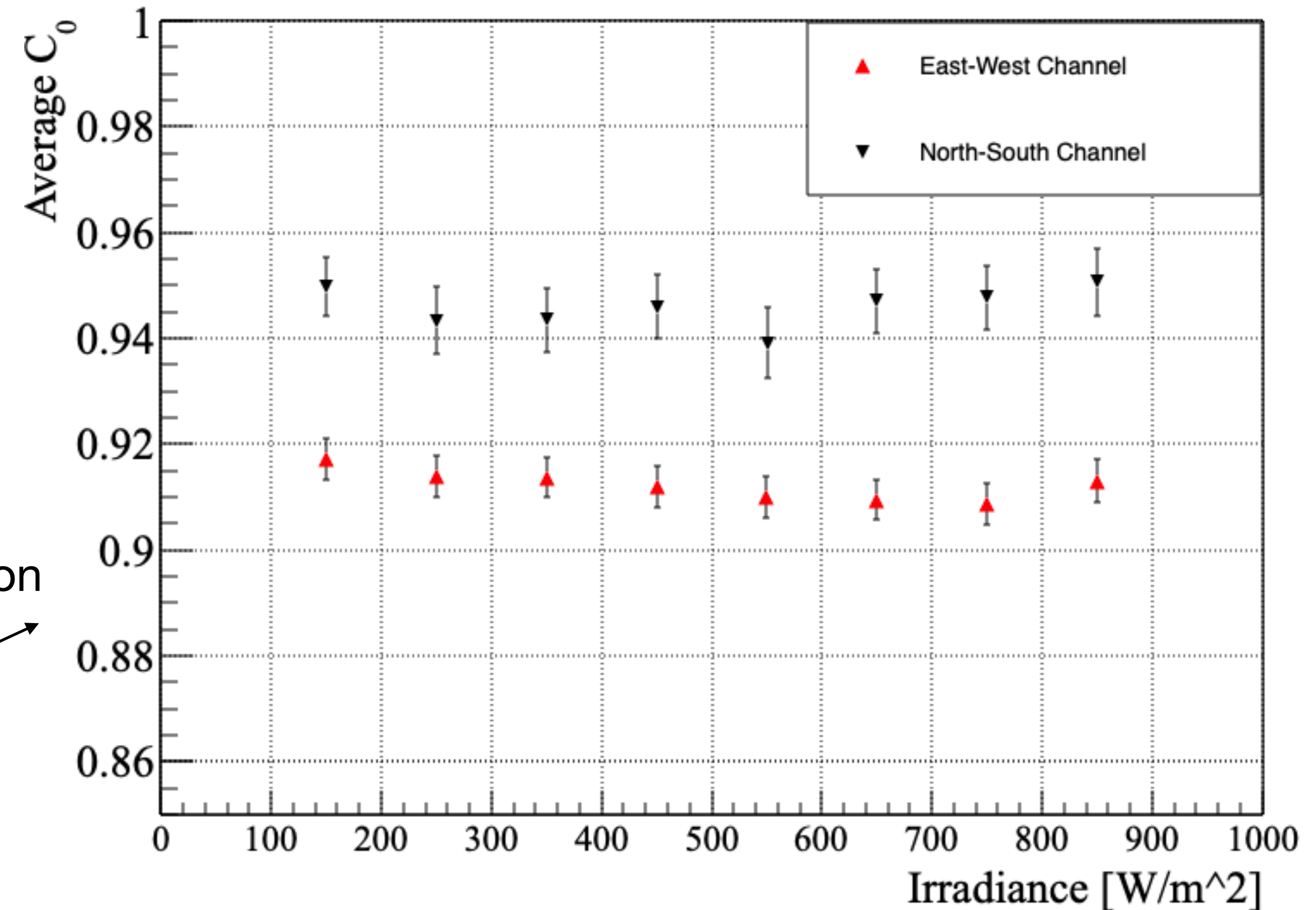
Procedure

Fix an Irradiance bin (ex 200 to 300 W/m^2)

In a whole year (2020) it only keeps traces when the irradiance is within the irradiance bin



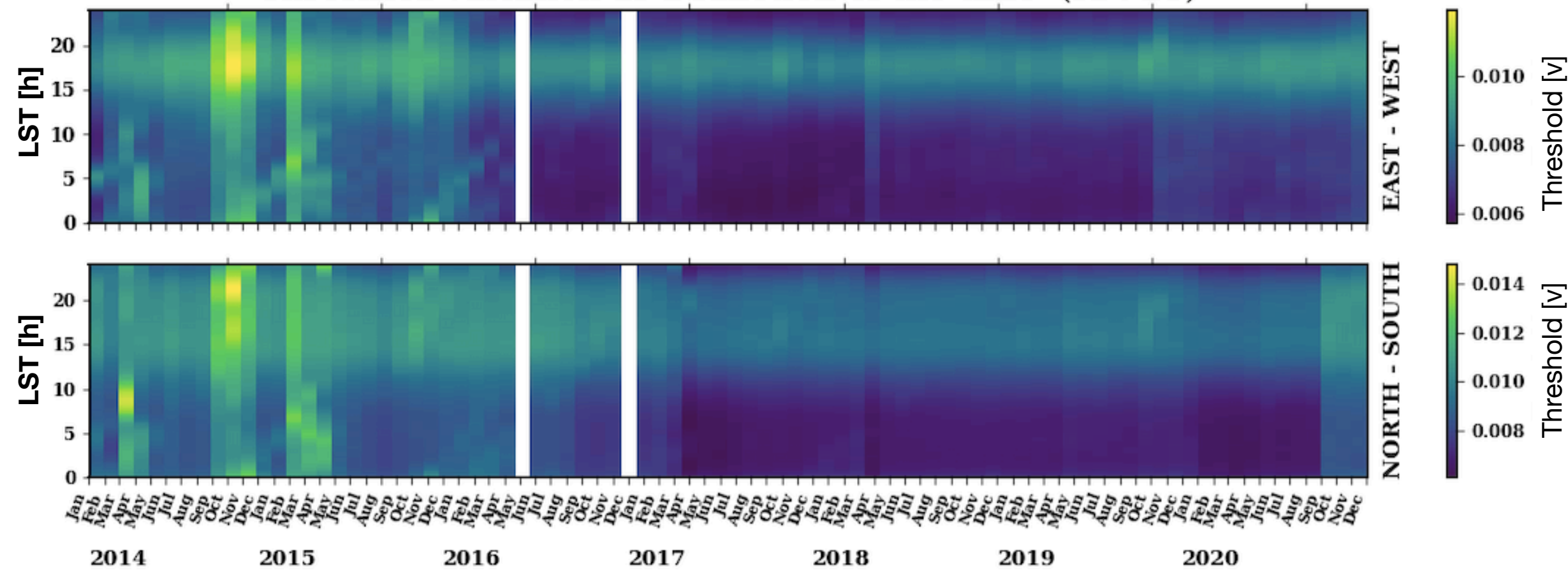
Calibration



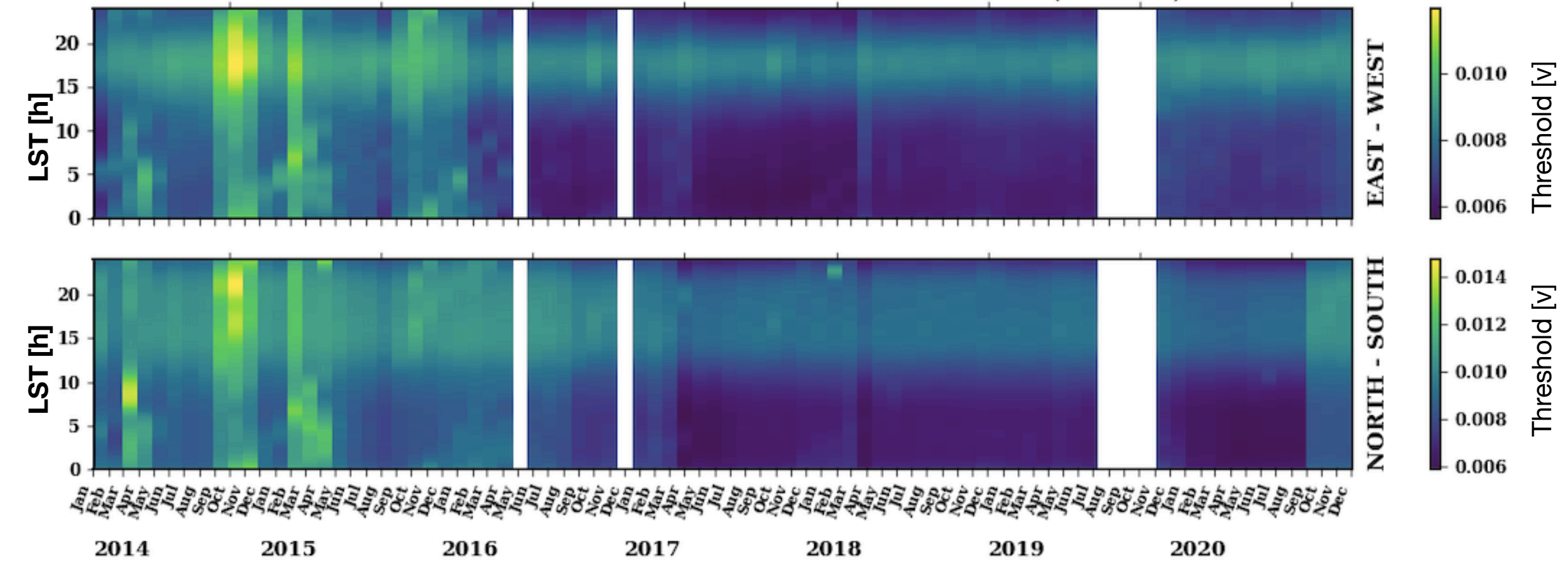
Time-dependent threshold approach

- Threshold as a function of time for each antenna

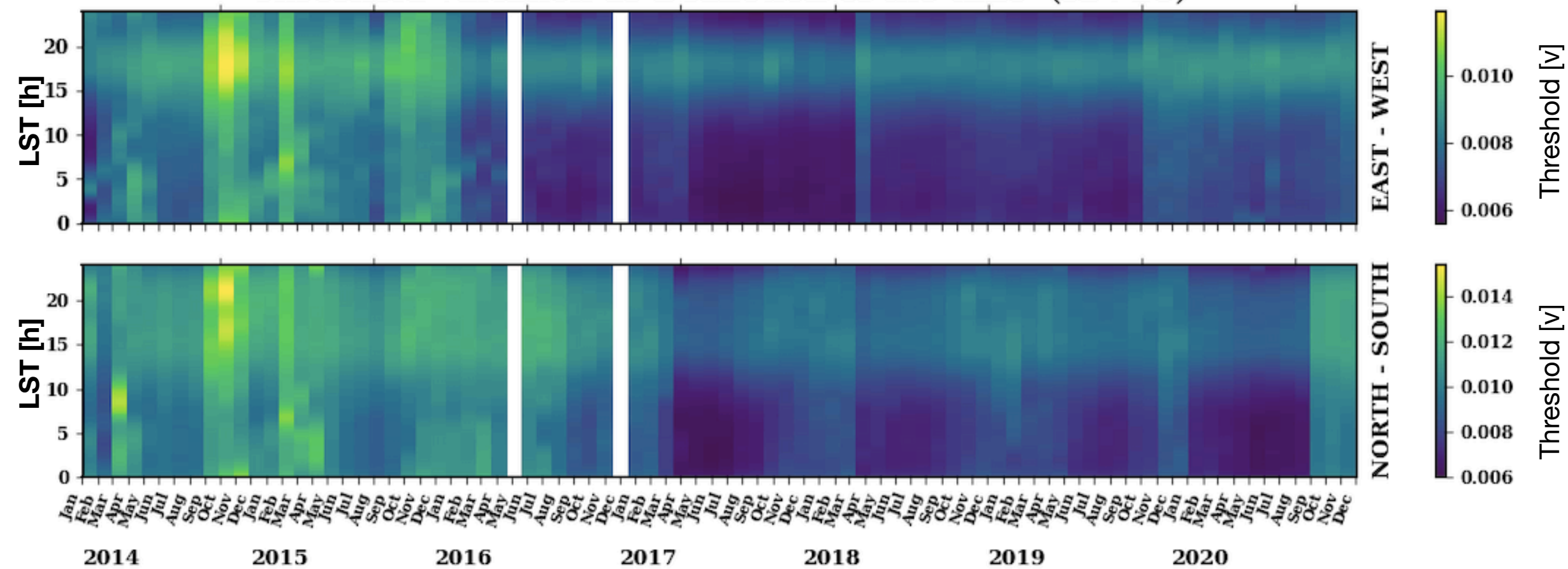
Threshold variation as a function of the time (Id : 25)



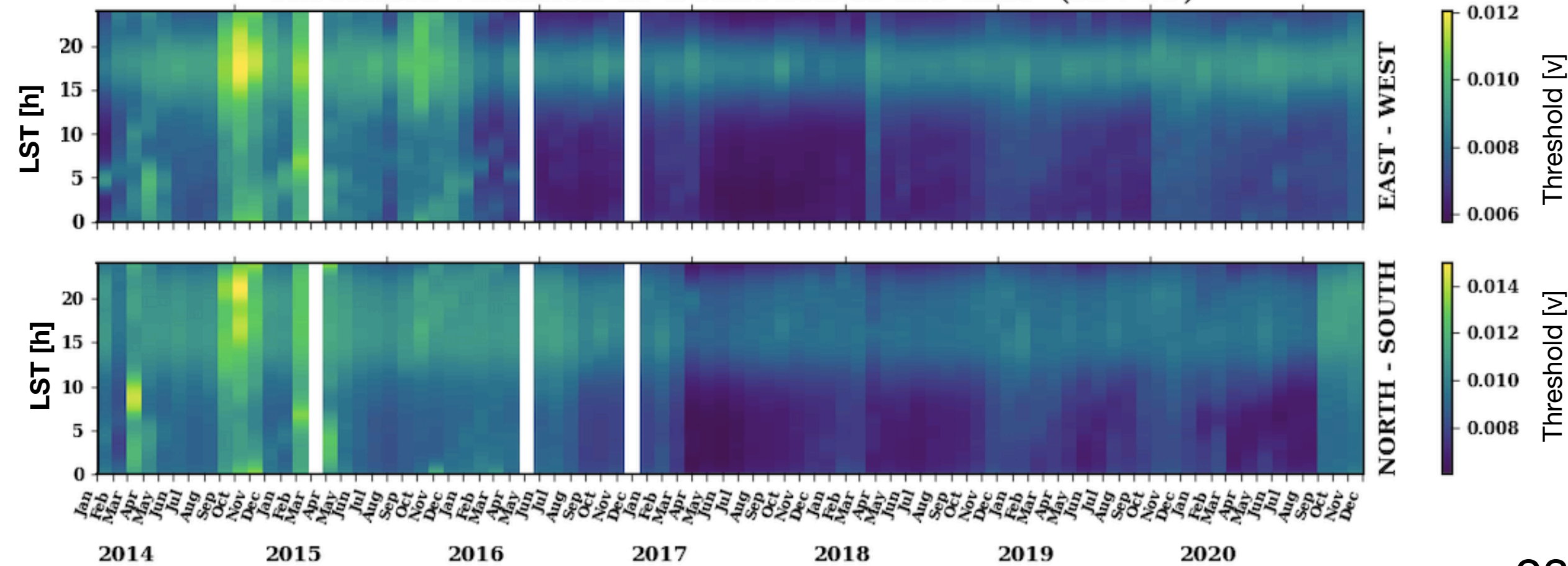
Threshold variation as a function of the time (Id : 27)



Threshold variation as a function of the time (Id : 28)



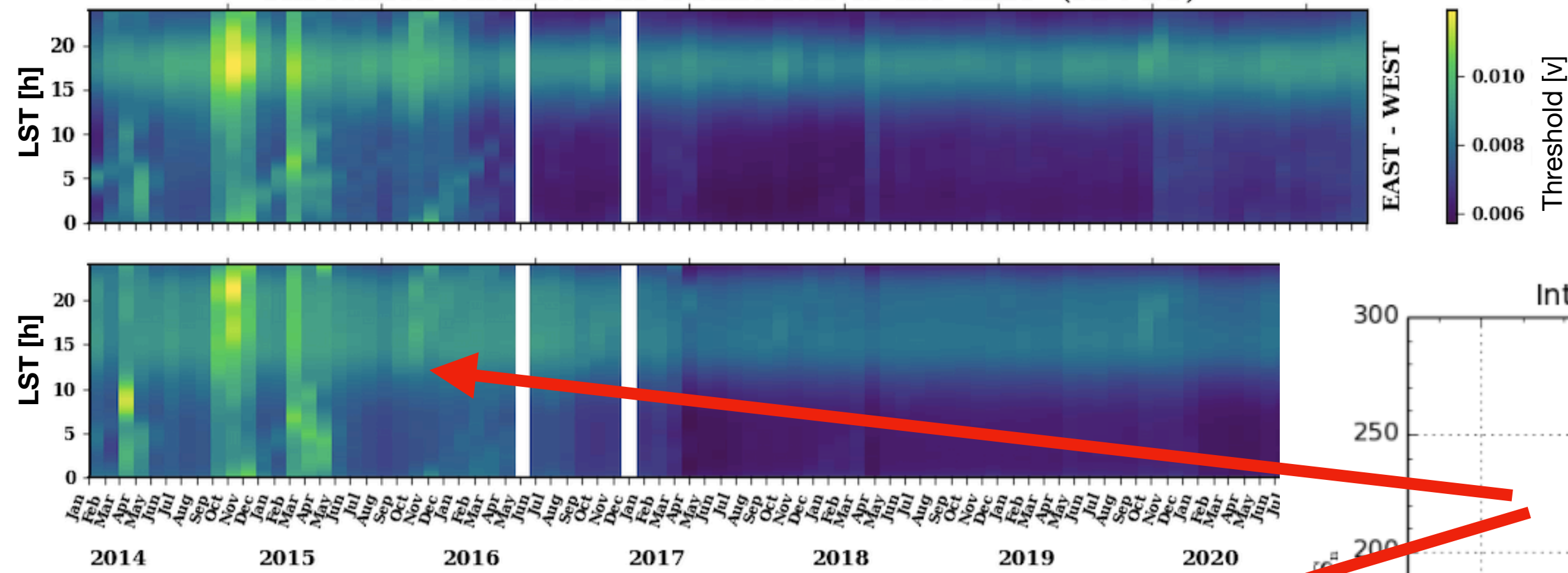
Threshold variation as a function of the time (Id : 29)



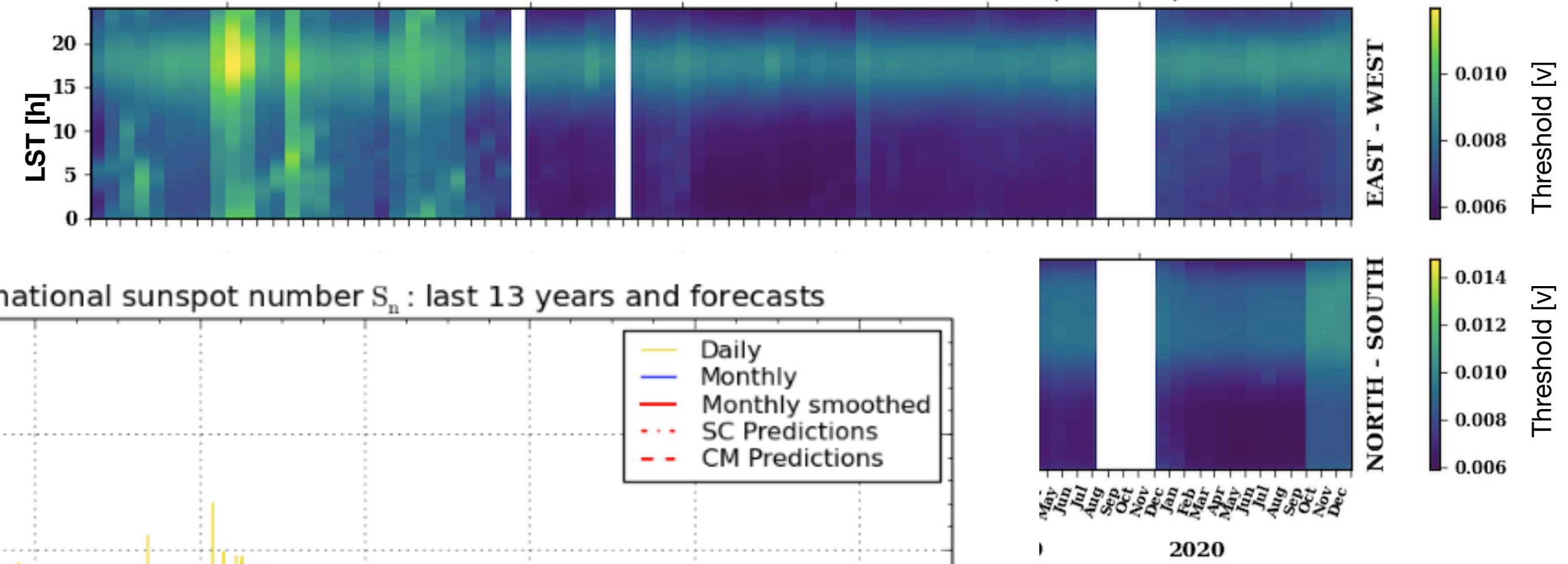
Time-dependent threshold approach

- Threshold as a function of time for each antenna

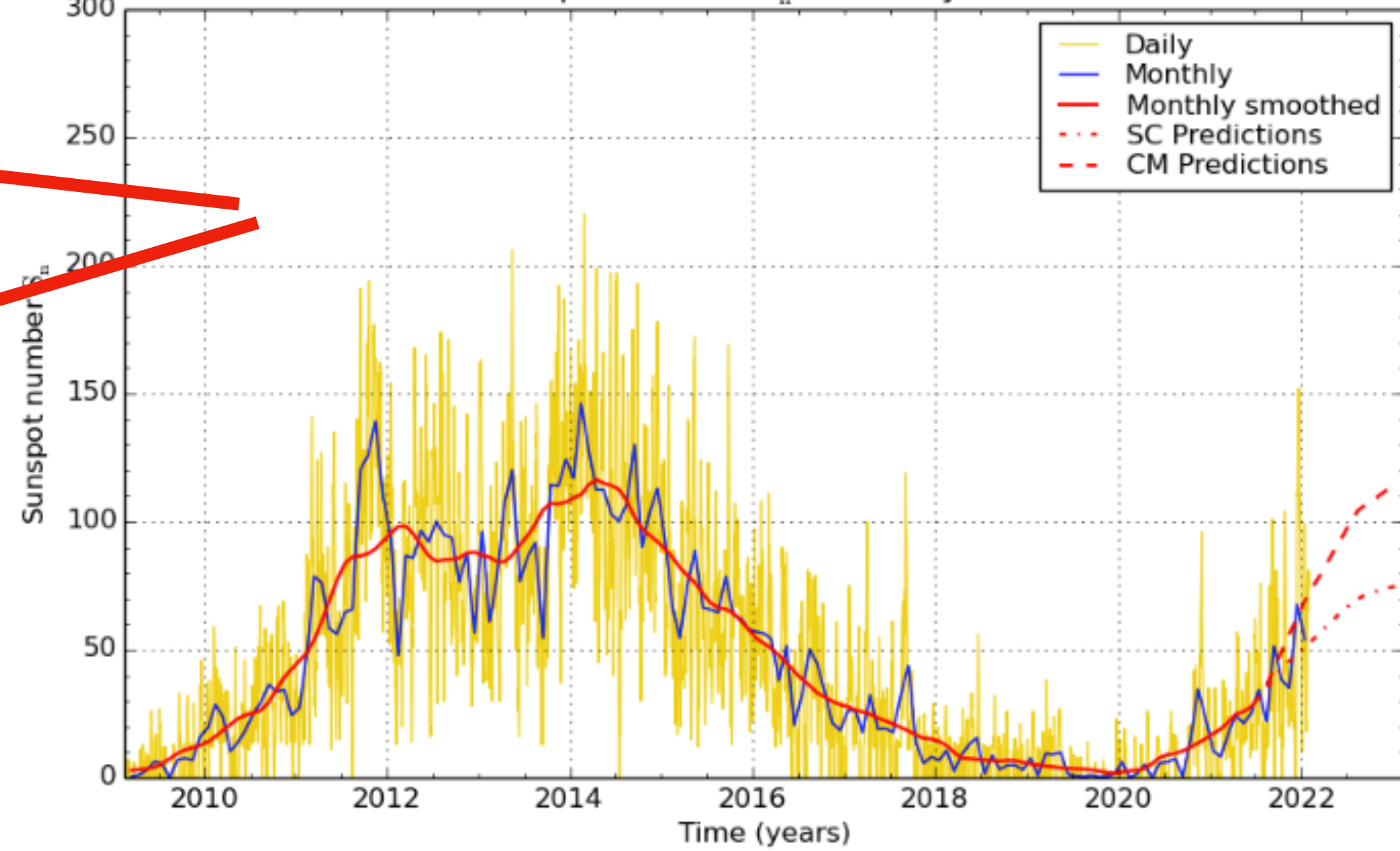
Threshold variation as a function of the time (Id : 25)



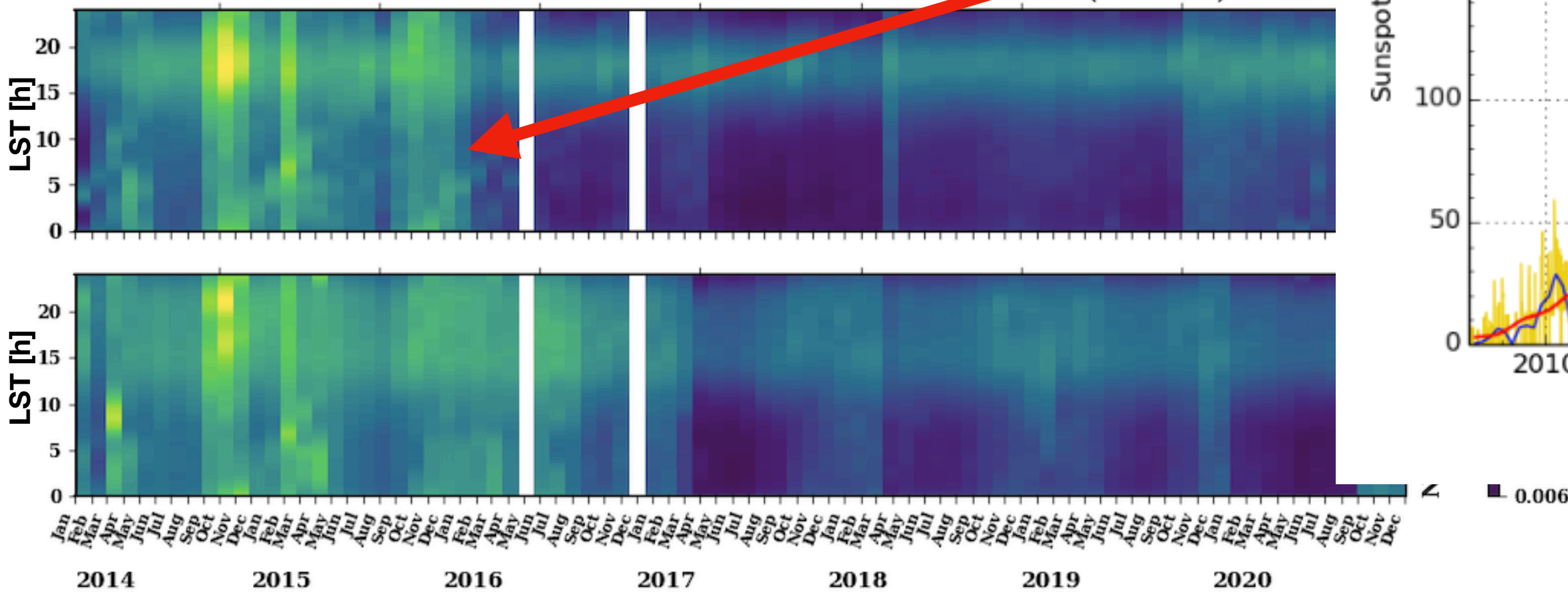
Threshold variation as a function of the time (Id : 27)



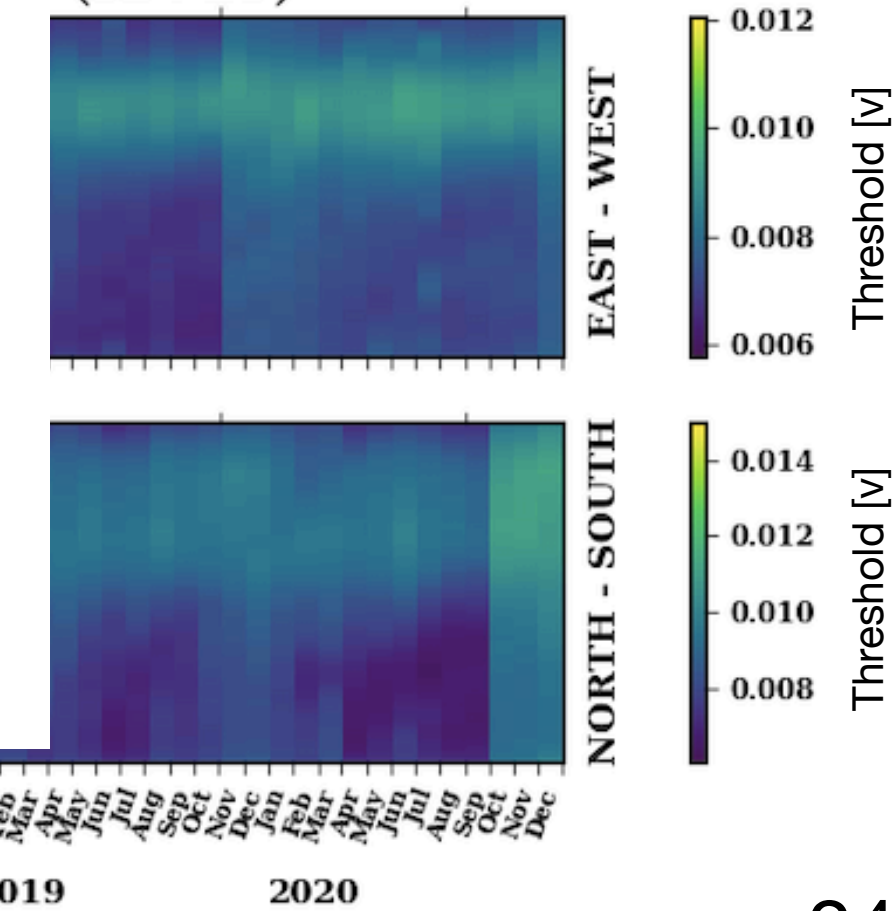
International sunspot number S_n : last 13 years and forecasts



Threshold variation as a function of the time (Id : 28)



(Id : 29)



SILSO graphics (<http://sidc.be/silso>) Royal Observatory of Belgium 2022 February 1