

# Practical and Accurate Calculations of Radio Emission from Extensive Air Showers

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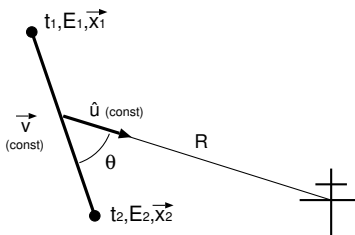
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# Today: e.g. ZHAireS in the time domain

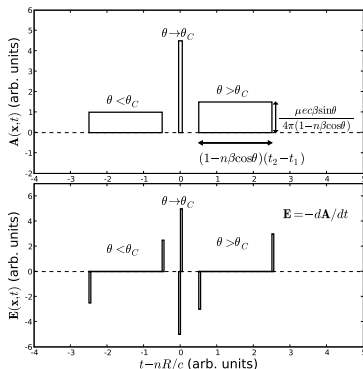
- Contribution to  $\vec{A}$  calculated separately for EVERY SINGLE TRACK!

$$\vec{A}(t, \hat{u}) = \frac{\mu e}{4\pi R c} \vec{\beta}_\perp \frac{\Theta(t-t_1^{det}) - \Theta(t-t_2^{det})}{1 - n\vec{\beta} \cdot \hat{u}}$$



$t_{1,2}^{det}$  depend on  $n_{eff} = \frac{1}{R} \int_0^R n(h) dl$  which is **VERY EXPENSIVE!**

(Phys.Rev.D81:123009,2010)

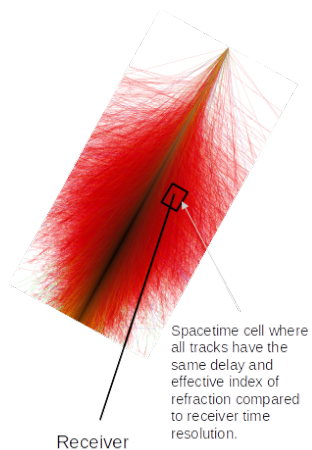


## Basic idea: Divide the shower in 4-D volumes

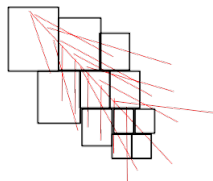
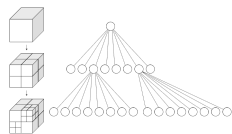
- Shower is divided into 4-D volumes of spatial sides  $L$  and time length  $L/c$
- Volumes must obey Fraunhofer condition for all observers:

$$L < \frac{1}{\sin\theta} \sqrt{\frac{\lambda R}{2\pi}} \quad (\theta = 60^\circ)$$

- Cell is sufficiently small so that many terms of the  $\vec{A}$  contribution are constant for the whole cell (depends on receiver time resolution)
- Contribution of all the tracks inside the cell is calculated only once, based on an average track for the cell
- Almost amounts to a macroscopic treatment of the shower, but retaining the microscopic precision



# Cells have different sizes: Octree binning



- Cell size optimization: Octree binning
- Start with spatial cells set to a large side length  $L$
- Checks the Fraunhofer condition for that side  $L$
- If side length does not meet the condition,  $L$  is divided by 2
  - We now have 8 cubes instead of the original
  - Far away from the observers,  $L$  is large
  - Close to the observers,  $L$  is small
- Calculate  $n_{eff}$  once per cell and store it
- Each cube is uniquely defined by its center point



# Effective average track and $\vec{A}$ calculation

All tracks inside a volume are represented by a single “effective” track



- For a cell  $i$  and track  $j$ :  
(different notation but equivalent to the ZHS single track formula)

$$\vec{A}_i(\vec{x}, t) = \frac{\mu_0(Q\vec{v})_{i\text{eff}\perp}}{4\pi R_i} \left| \frac{dt'_{i\text{eff}}}{dt} \right|_{t'=t'_{\text{ret}}} \Pi(t', t'_{1\text{eff}}, t'_{2\text{eff}}),$$

where  $\Pi$  is a boxcar function,

$t'$  is source time,  $\vec{x}$  and  $t$  are observer position and time,

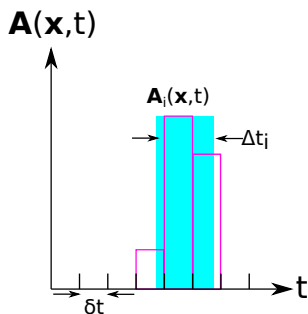
$$(dt')_{i\text{eff}} = |t'_{2\text{eff}} - t'_{1\text{eff}}| = \left| \frac{\sum_j w_{ij} q_{ij} v_{ij} dt'_{ij}}{\sum_j w_{ij} q_{ij} v_{ij}} \right| \text{ and}$$

$$(Q\vec{v})_{i\text{eff}} = \frac{\sum_j w_{ij} q_{ij} \vec{v}_{ij} dt'_{ij}}{\sum_j dt'_{ij}}$$

- Average track inside cell  $i$  is completely defined by  $(dt')_{i\text{eff}}$  and  $(Q\vec{v})_{i\text{eff}}$

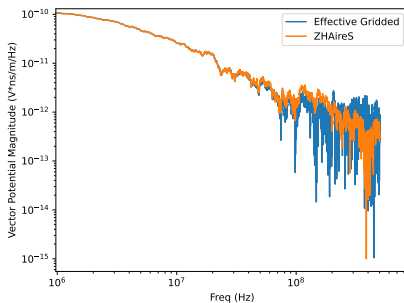
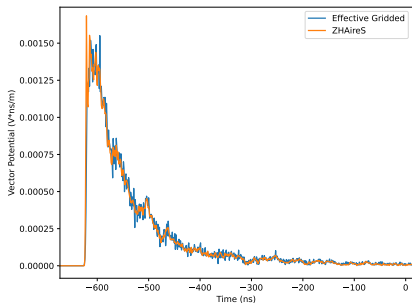
## Time binning the Radiation at the Observer

- Add the contribution of all cells to a resulting  $\vec{A}$  at the observer
- Similar to ZHAIREs
- Boxcar function  $\Pi(t', t'_{1\text{eff}}, t'_{2\text{eff}})$  gives rise to time window  $\Delta t_i$
- $\delta t$  is a fixed time bin width set by the receiver
- Bins that fully overlap with  $\Delta t_i$ ; filled with the full value of  $\vec{A}_i(\vec{x}, t)$
- Contribution of partially overlapping bins scaled by the fraction of bin overlap



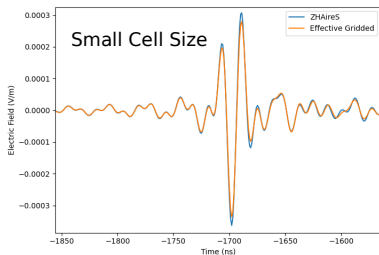
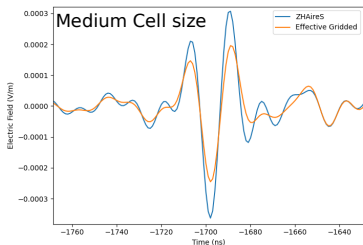
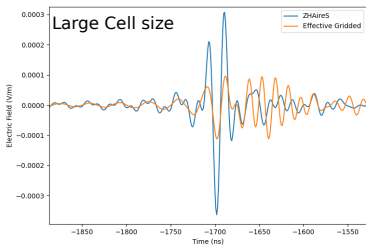
# Comparison with full ZHA<sub>IRE</sub>S simulations

- Input: particle tracks obtained from ZHA<sub>IRE</sub>S
- Calculations using the ZHS algorithm compared to this work
- Thinning of  $10^{-3}$  and  $\theta = 70^\circ$



# Importance of correct cell sizing

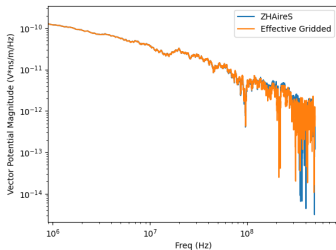
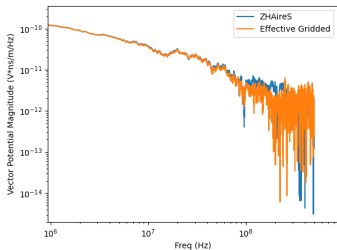
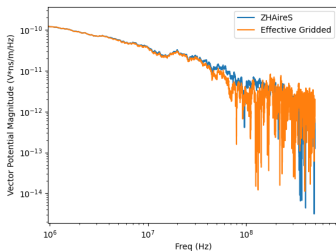
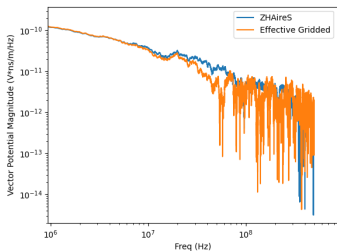
- Bandpass filtered pulse 30-80 MHz, thinning  $10^{-3}$  and  $\theta = 70^\circ$





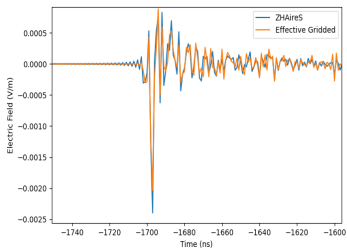
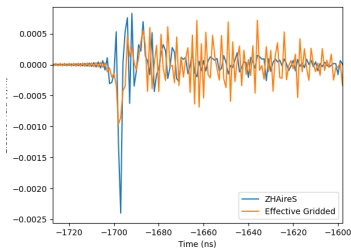
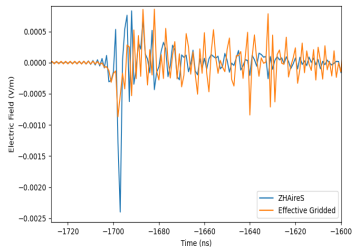
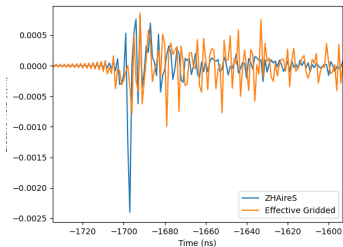
# Importance of correct cell sizing

- Spectra, thinning  $10^{-3}$  and  $\theta = 70^\circ$



# Importance of correct cell sizing

- Full bandwidth pulse, thinning  $10^{-3}$  and  $\theta = 70^\circ$  (cell size depends on  $\lambda$ )



# Advantages

- Only a single calculation per 4-D volume
- Theoretically much faster than track-by-track calculations
- Almost a macroscopic approach, but retaining the precision of the successful microscopic formalism.
  - Maybe parametrizations of average tracks could be devised
- More detailed calculations that are too expensive on a track-by-track basis can be performed
  - These can be calculated in advance for each cell
  - Atmospheric effects: e.g. curved propagation
    - Important effect specially for upgoing horizontal showers
    - Very relevant for satellite and balloon experiments
  - Antenna patterns could also be taken into account in the simulation

# Questions?

## Other applications of Radio...



# BACKUP

# Math

$$\vec{A}_i(\vec{x}, t) = \frac{\mu_0}{4\pi} \iiint_V d^3\vec{x}' \sum_i \frac{\vec{J}_{i\perp}(\vec{x}'_i, t'_i)}{|\vec{x} - \vec{x}'_i|} \bigg|_{t'=t'_i} \frac{dt'}{dt}$$
 In the case of a single track, we can use:

$$\vec{J}(\vec{x}', t') = \rho(\vec{x}', t') \vec{v}, \text{ where } \rho(\vec{x}', t') = q\delta(\vec{x}' - \vec{v}t') \Pi(t', t'_1, t'_2)$$

The  $\delta$  function models the track as an infinitely thin linear charge density with particle velocity  $\vec{v}$ . We then obtain:

$$\vec{A}(\vec{x}, t) = \frac{\mu_0 q}{4\pi R} \vec{v}_\perp \bigg|_{t'=t'_{ret}} \frac{dt'}{dt} \Pi(t', t'_1, t'_2)$$

If, instead of a single track, we have a charged current density vector  $\vec{J}_i$  at a location  $\vec{x}_i$  that is approximately constant over a volume  $\Delta V_i$ :

$$\vec{A}_i(\vec{x}, t) = \frac{\mu_0}{4\pi} \Delta V_i \frac{\vec{J}_{i\perp}(\vec{x}'_i, t'_{ret})}{|\vec{x} - \vec{x}'_i|} \bigg|_{t'=t'_{ret}} \frac{dt'}{dt}$$

In this case, our  $\vec{J}_\perp$  is no longer a single track, but a collection of tracks within a cell.  $\Delta V_i \vec{J}_{i\perp} = Q_i \vec{v}_{i\perp}$ , where  $Q_i$  is charge within the cell:

$$\vec{A}_i(\vec{x}, t) = \frac{\mu_0}{4\pi} \frac{Q_i \vec{v}_{i\perp}(\vec{x}'_i, t'_{ret})}{|\vec{x} - \vec{x}'_i|} \bigg|_{t'=t'_{ret}} \frac{dt'}{dt}$$









