

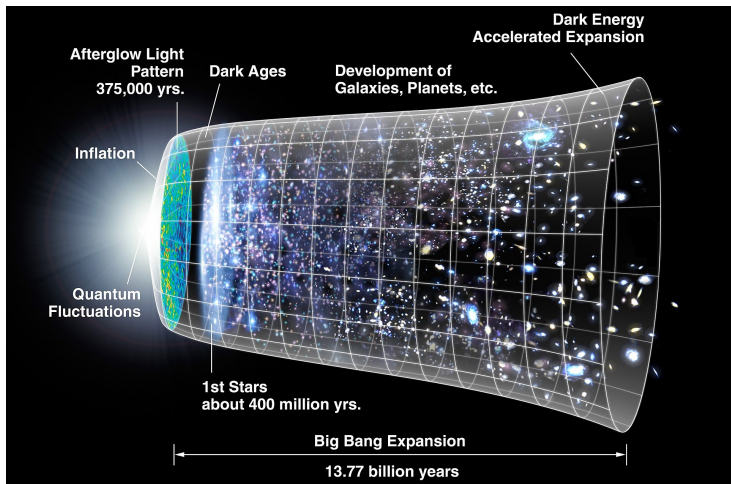
# Mimicking Alternatives of Inflation with Interacting Spectator Fields

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# Motivation



[Wikipedia]

- Prediction of minimal inflationary set up (Single Field Slow Roll):

$$\mathcal{P}_{\mathcal{R}} \sim k^{n_s-1}$$

- Assuming a power-law scale factor:

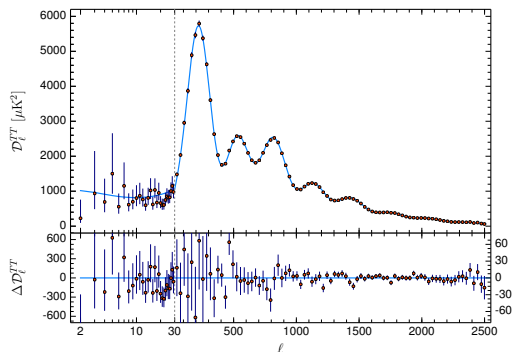
$$a(t) \sim t^p$$

- Resulting spectral tilt:

$$n_s - 1 = 3 - \left| 3 + \frac{2}{p-1} \right| \approx 0$$

- $n_s$  vanishes for  $p \gg 1$  (Inflation)  
and  $p = 2/3$  (Matter Contraction)

# Features in The CMB



- Cosmic Variance limits accuracy at small  $\ell$ .
- Dip at  $\ell \approx 30$
- Hint for features in the CMB

[Planck 2018: Constrains on inflation (1807.06211)]

# Heavy Fields / Primordial Standard Clock

- Heavy Fields ( $m \gg H$ ) during the primordial epoch oscillate around minimum.
- Oscillation will affect the evolution of the "inflaton".
- Studies on oscillatory features in CMB with  $m^2\chi^2$ -potential [Primordial Standard Clock, X. Chen et al (1411.2349)].
- Assumes only gravitational interaction between field.
- Can distinguish Inflation and alternatives.

- Introduce non-trivial interaction between fields:

$$\mathcal{L} = \frac{M_{\text{P}}^2}{2} R - \frac{1}{2} \omega^2(\chi) g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi) \\ - \frac{1}{2} f^2(\phi) g^{\mu\nu} \partial_\mu \chi \partial_\nu \chi - \frac{1}{2} m^2(\phi) \chi^2$$

- non-canonical kinetic term  $\omega(\chi)$  for the "inflaton"  $\phi$
- non-canonical kinetic term  $f(\chi)$  for heavy spectator field  $\chi$   
+ non-constant mass term  $m(\phi)$
- **red** terms appear generically in models with non-minimal coupling to gravity

[Guillem Domenech, Javier Rubio, JW (1811.08224)]

# Mechanism to Create Features

$$\chi \propto \sqrt{\frac{a^{-3}}{m_{\text{eff}} f^2}} \sin \left( \int m_{\text{eff}} dt \right) \quad \text{with} \quad m_{\text{eff}} = \frac{m}{f}$$

$\chi$	Spectator field oscillates under the assumption $m_{\text{eff}}/H \gg 1$ .
$\Delta H_{\text{osci}}$	Spectator field adds an oscillatory corrections to the Hubble parameter.
$\Delta \epsilon_{\text{osci}}$	Also the slow-roll parameter will get a correction.
$\Delta \mathcal{P}_{\mathcal{R}}$	Finally this will lead to oscillatory features in the power spectrum

$$\Delta \mathcal{P}_{\mathcal{R}} \propto \left( \frac{2k}{k_r} \right)^\nu \sin \left[ C \left( \frac{2k}{k_r} \right)^{\frac{\gamma}{p}} + \theta_r \right]$$

# Corrections to the power spectrum

$$\Delta \mathcal{P}_{\mathcal{R}} \propto \left( \frac{2k}{k_r} \right)^\nu \sin \left[ C \left( \frac{2k}{k_r} \right)^{\frac{\gamma}{p}} + \theta_r \right]$$

- $k_r$  : scale of first oscillation
- Scaling of amplitude:  $\nu = -\frac{3}{2} + \frac{1}{2} \frac{\gamma}{p} + \frac{\delta_m - 2\delta_f}{1 + \delta_m - \delta_f}$
- Scaling of frequency:  $\gamma = \frac{1 + p\delta_m - p\delta_f}{1 + \delta_m - \delta_f}$
- Assume constant change in:  $\delta_m = \frac{\dot{m}}{Hm}$  and  $\delta_f = \frac{\dot{f}}{Hf}$

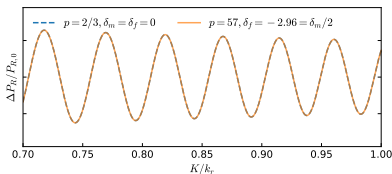
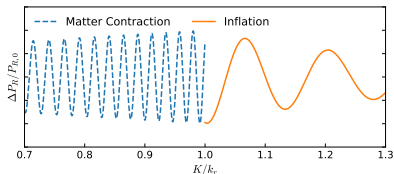


# Mimicking a Signal

Frequency:  $\bar{p} = p/\gamma$ ,      Amplitude:  $\nu = -\frac{3}{2} + \frac{1}{2}\frac{1}{\bar{p}} + \frac{\delta_m - 2\delta_f}{1 + \delta_m - \delta_f}$

- Use two parameters  $\delta_m$  and  $\delta_f$  to fix two observables  $\nu$  and  $\bar{p}$
- 1) Choose  $\delta_m = 2\delta_f$  to cancel additional contribution to  $\nu$
  - 2) Choose  $\delta_f$  such that  $\bar{p} = p_{\text{inf}}/\gamma = p_{\text{alt}}$

Resonance condition:  $k = \frac{m}{f} a$



- Constant  $m$  and  $f = 1$ , Resonance directly probes  $a(t)$
- Very distinct feature for inflation and alternatives
- $m$  and  $f$  depend on time, Resonance probes combination of  $m, f$  and  $a$
- Inflation can mimic signal from alternative scenarios

# Conclusions

- Time dependent mass makes different signals indistinguishable at the level of the power spectrum.
- Power spectrum is not enough to tell inflation and alternatives apart.
- Need full information (frequency scaling + amplitude scaling) on bispectrum to break degeneracy.