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# INTEGRAL X-RAY CONSTRAINTS ON SUB-GEV DARK MATTER

Speaker: Elena Pinetti

PONT, 10th December 2020

# INTEGRAL X-ray constraints on sub-GeV Dark Matter

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# INTEGRAL X-ray constraints on sub-GeV Dark Matter

$1 \text{ MeV} \leq m_\chi \leq 5 \text{ GeV}$

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# Photon searches

INTEGRAL

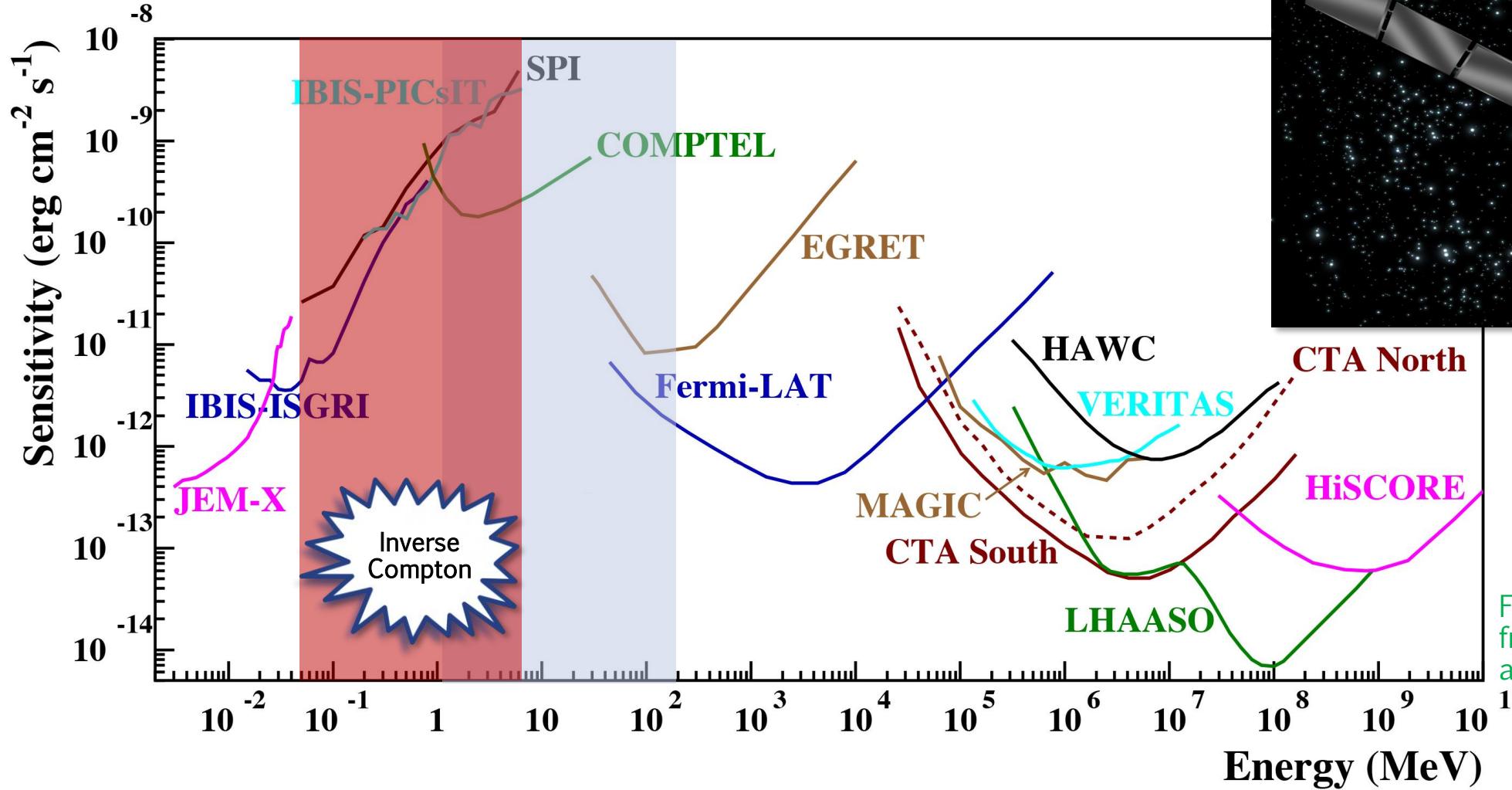


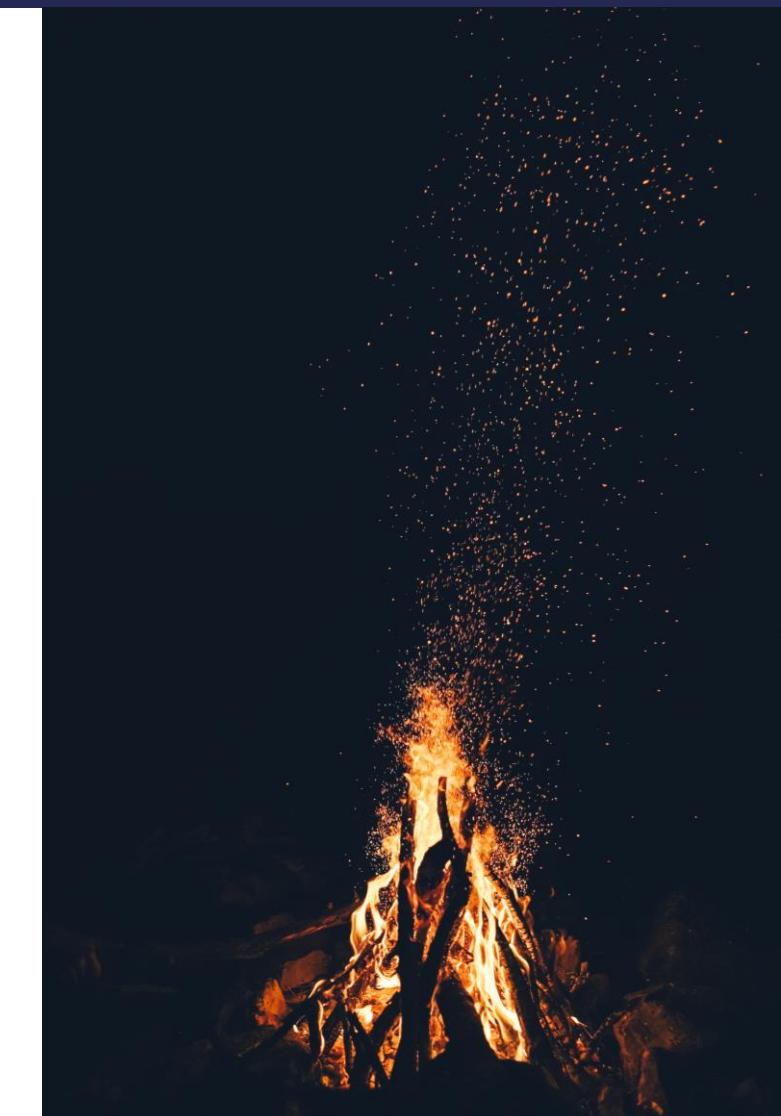
Figure adapted  
from Tatischeff+  
arxiv:1805.06435

# Annihilation channels

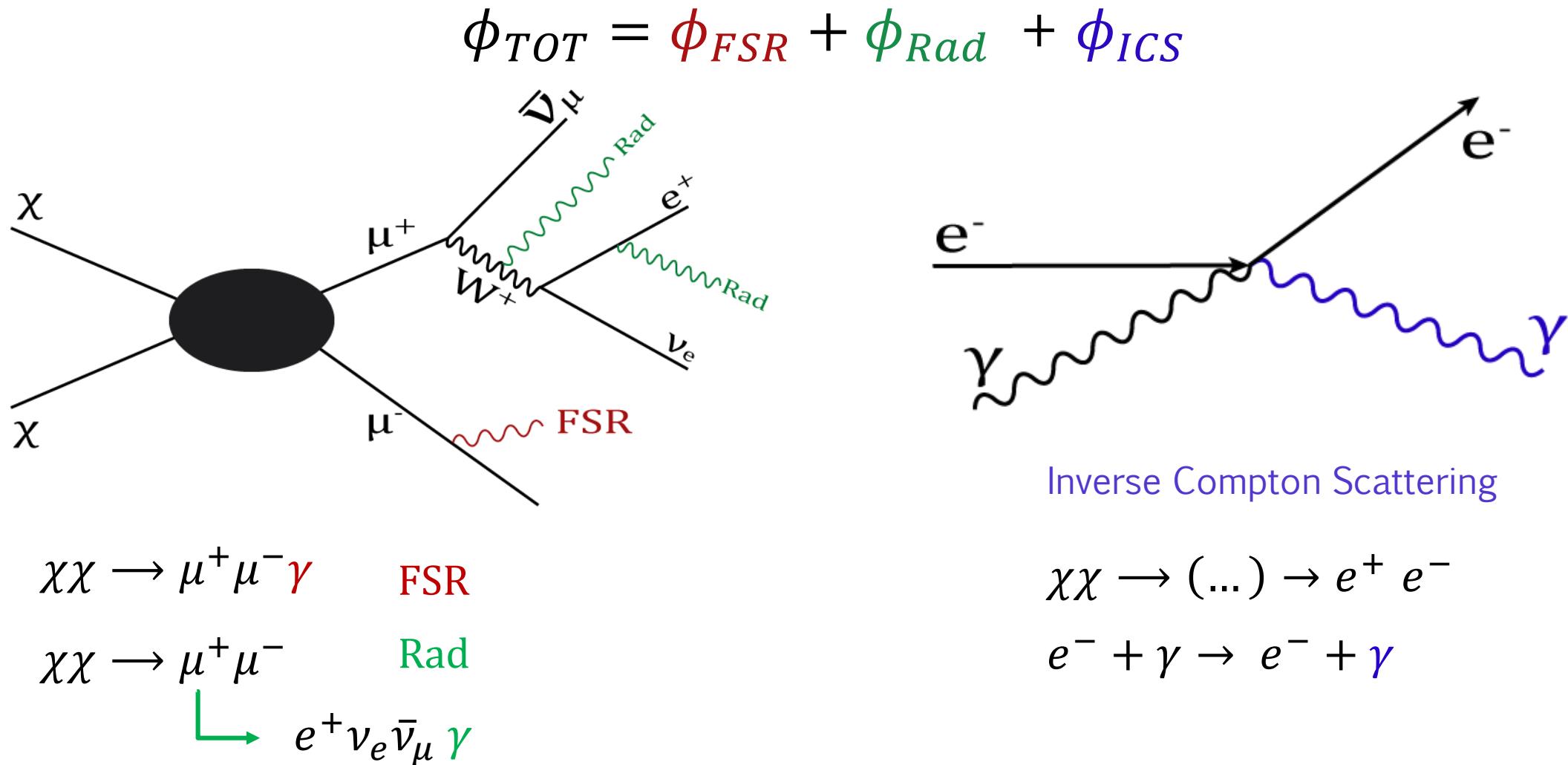
3 annihilation channels:  $\chi\chi \rightarrow e^+e^-$

$$\chi\chi \rightarrow \mu^+\mu^-$$
$$\chi\chi \rightarrow \pi^+\pi^-$$

Kinematically open:  $m_\chi > m_i \quad i = e, \mu, \pi$



# Total Flux



# Final State Radiation

$$\frac{d\phi_{FSR}}{dE_\gamma d\Omega}(E_\gamma, \theta) = \frac{1}{4\pi} \frac{\langle \sigma_{ann} v \rangle}{2m_{DM}^2} \frac{dN_{FSR}}{dE_\gamma} J(\theta)$$

Particle  
Properties

Energy spectrum per  
annihilation event

Angle in the sky DM density

$$J(\theta) = \int_{l.o.s} \rho^2(r(s, \theta)) ds$$

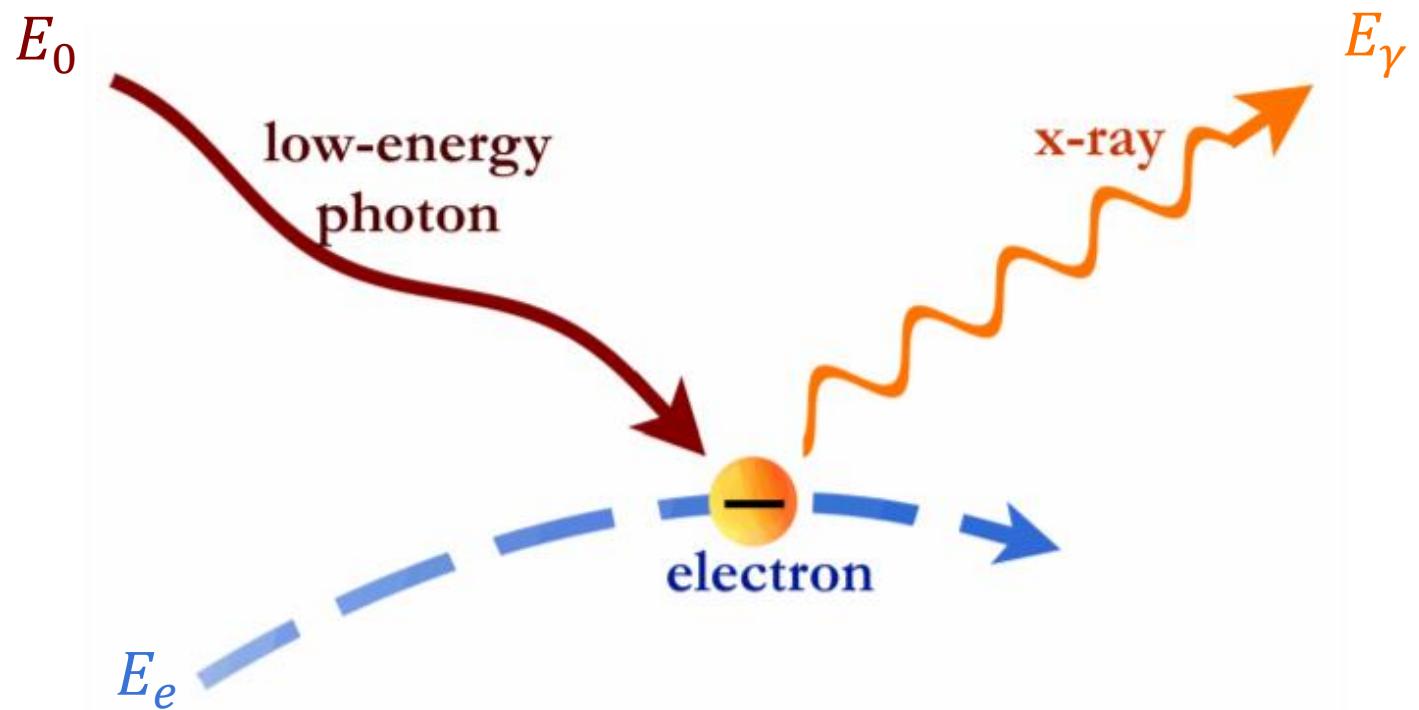
Line of sight

Radiative Decay:

$$\begin{aligned} \mu^- &\rightarrow e^- \bar{\nu}_e \nu_\mu \gamma \\ \pi^- &\rightarrow l \bar{\nu}_l \gamma \quad l = e, \mu \end{aligned} \quad \left. \right\} \quad \frac{dN_{Rad}}{dE_\gamma}$$

# Inverse Compton Scattering

$$\chi\chi \rightarrow (\dots) \rightarrow e^+ e^-$$



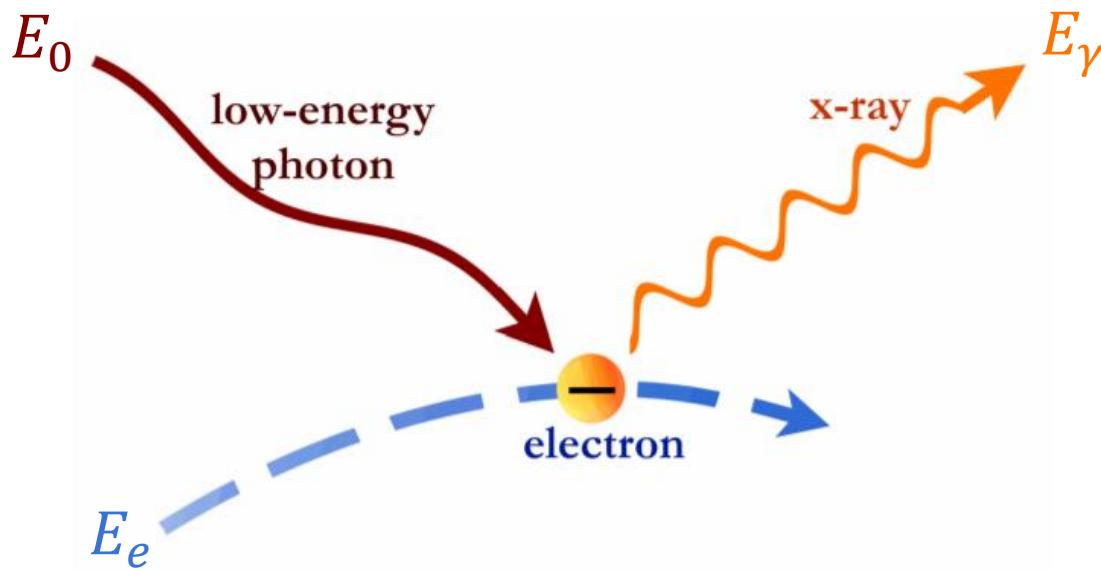
3 kind of photons:

- CMB
- IR (dust)
- Optical (starlight)

# X rays

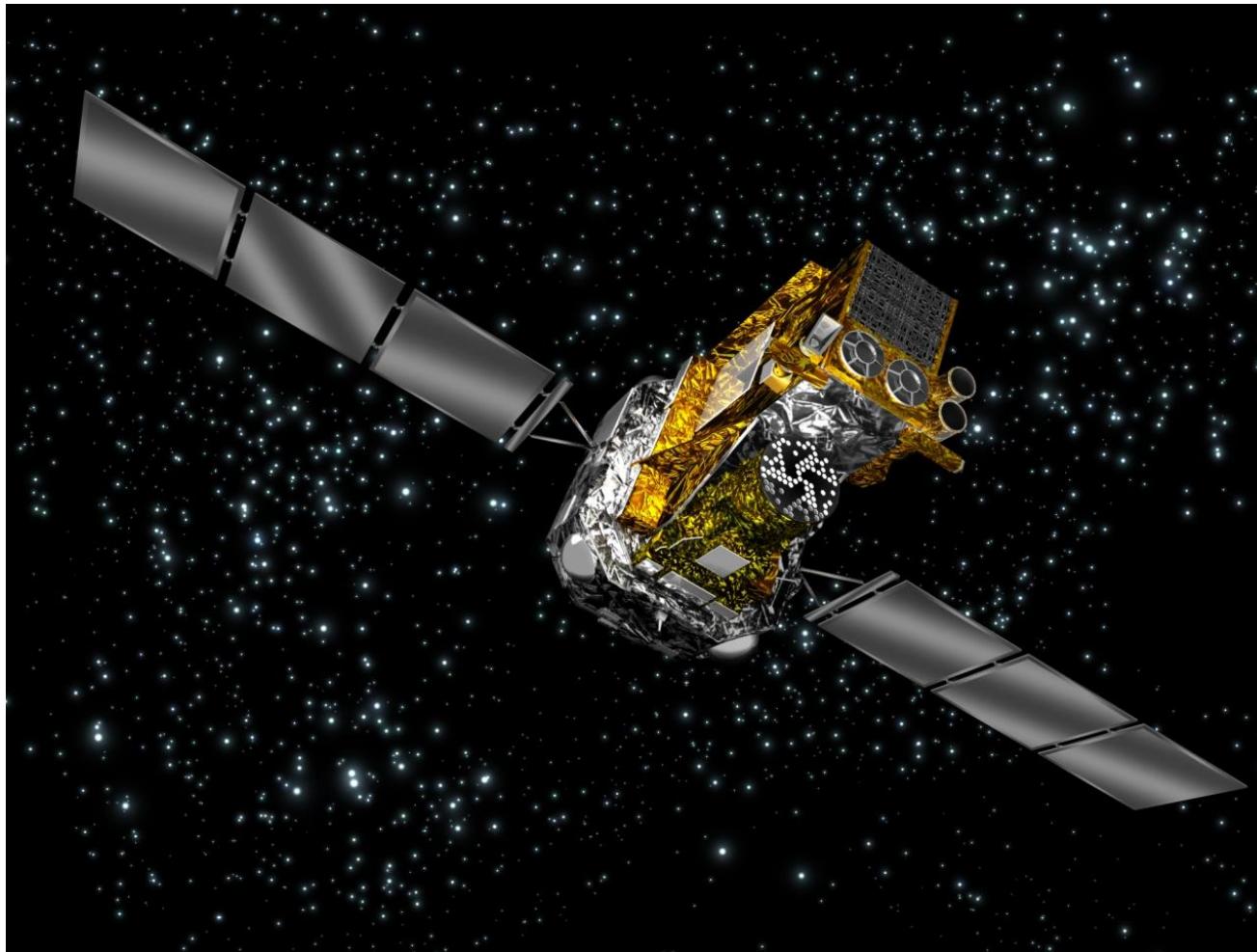
$$\gamma = \frac{E_e}{m_e}$$

$$E_\gamma \approx 4\gamma^2 E_0$$



Type	$E_0$ [eV]	$E_e$ [GeV]	$E_\gamma$ [keV]	
CMB	$10^{-4}$	5	40	
IR	$10^{-2}$	0.5	40	
Opt	10	0.05	400	X rays

# INTEGRAL Space Telescope



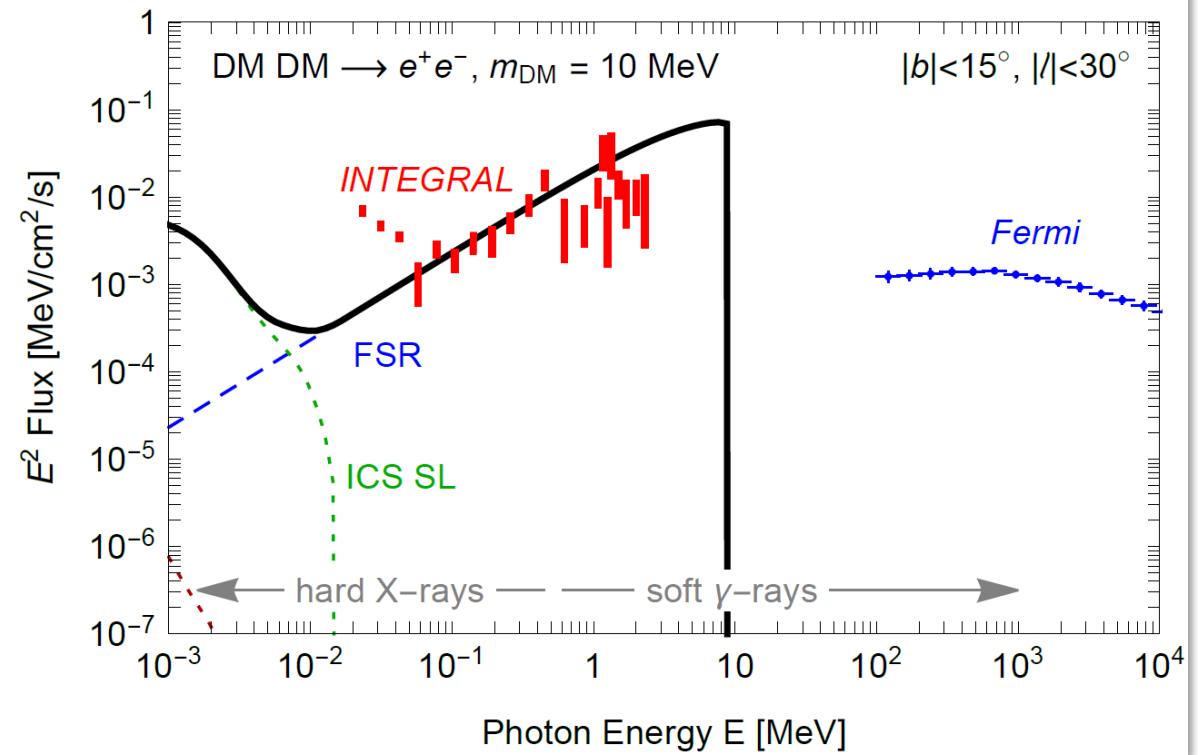
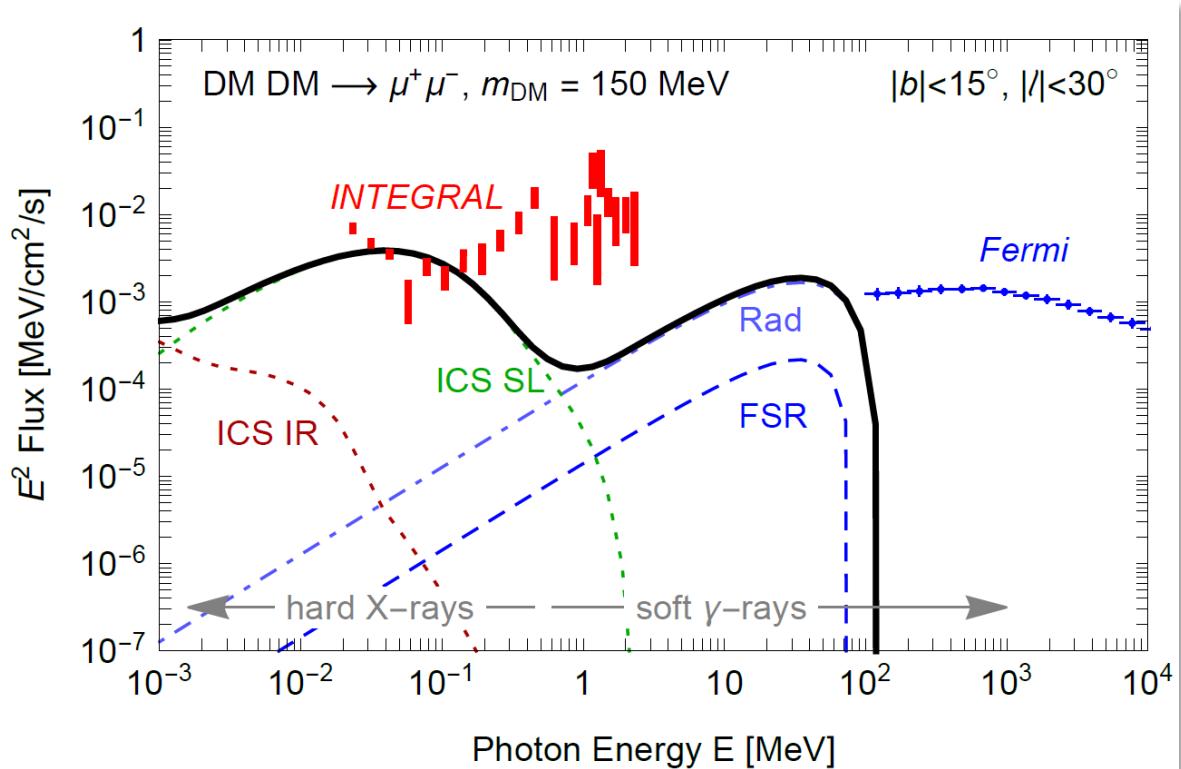
Data taking: 2003-2009

Photons in the energy range:  
20 keV ~ 8 MeV

- Hard X rays
- Soft gamma rays

INTEGRAL:  
INTERnational Gamma-Ray Astrophysics Laboratory

# Photon Spectra



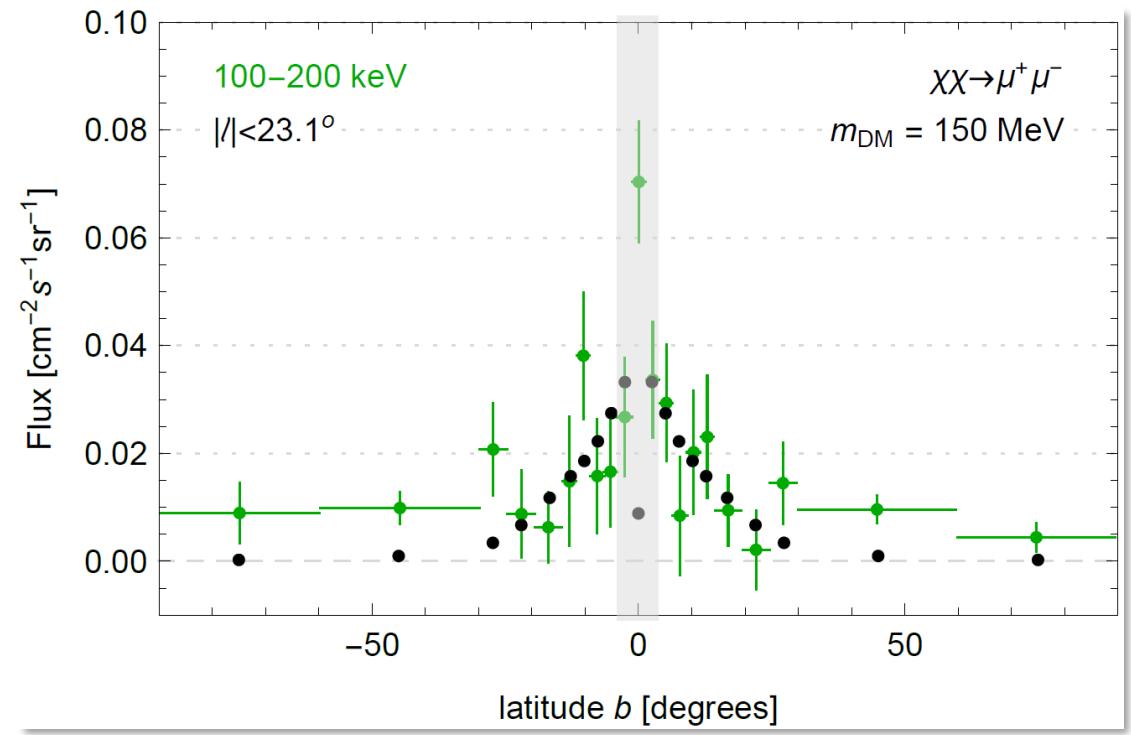
# Data sets

5 energy bands:

- 27-49 keV
- 49-90 keV
- 100-200 keV
- 200-600 keV
- 600-1800 keV

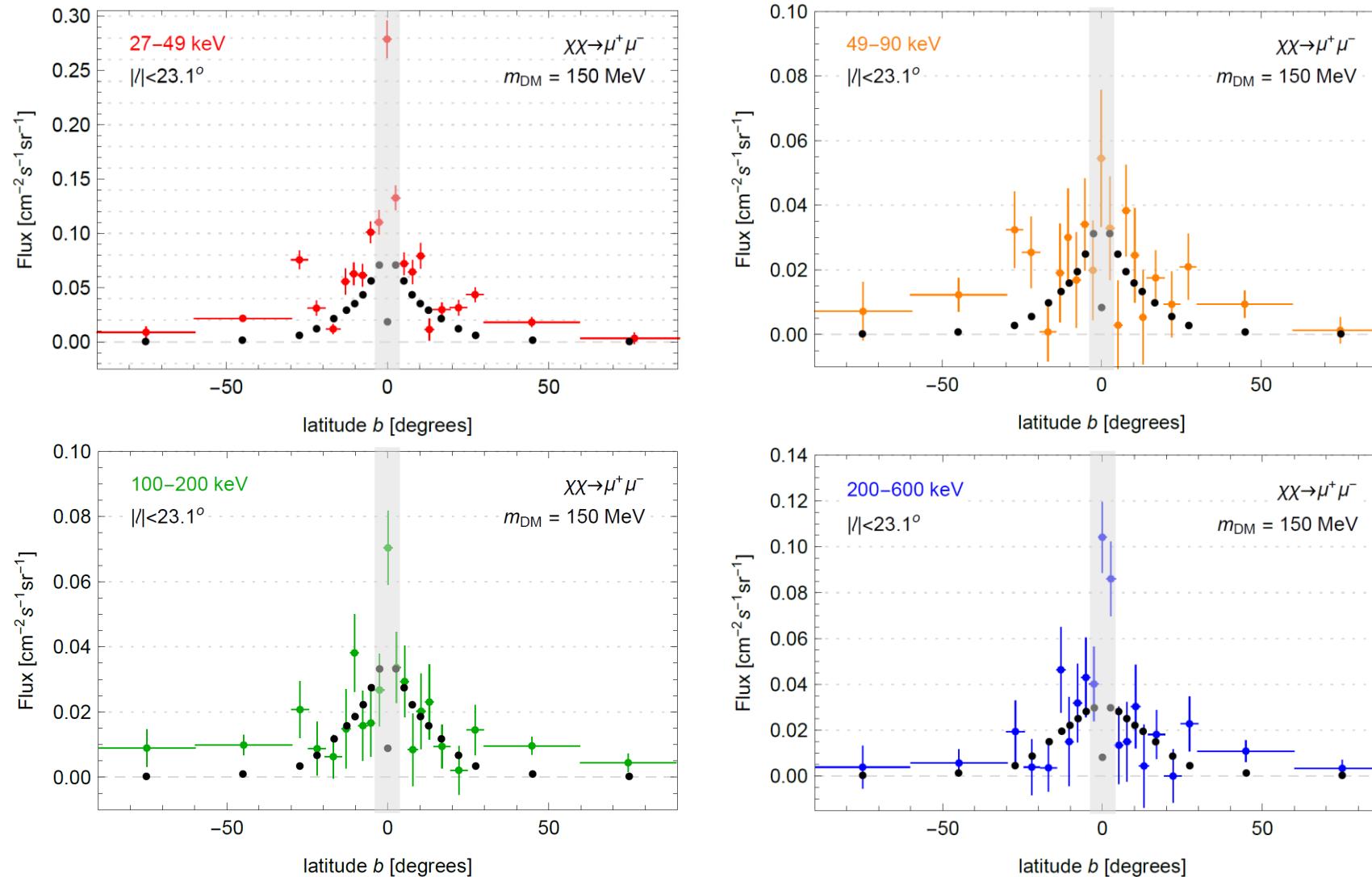
21 bins in latitude (15 for the fifth band)

$-23.1^\circ < |l| < 23.1^\circ$  ( $|l| < 60^\circ$  for the fifth band)

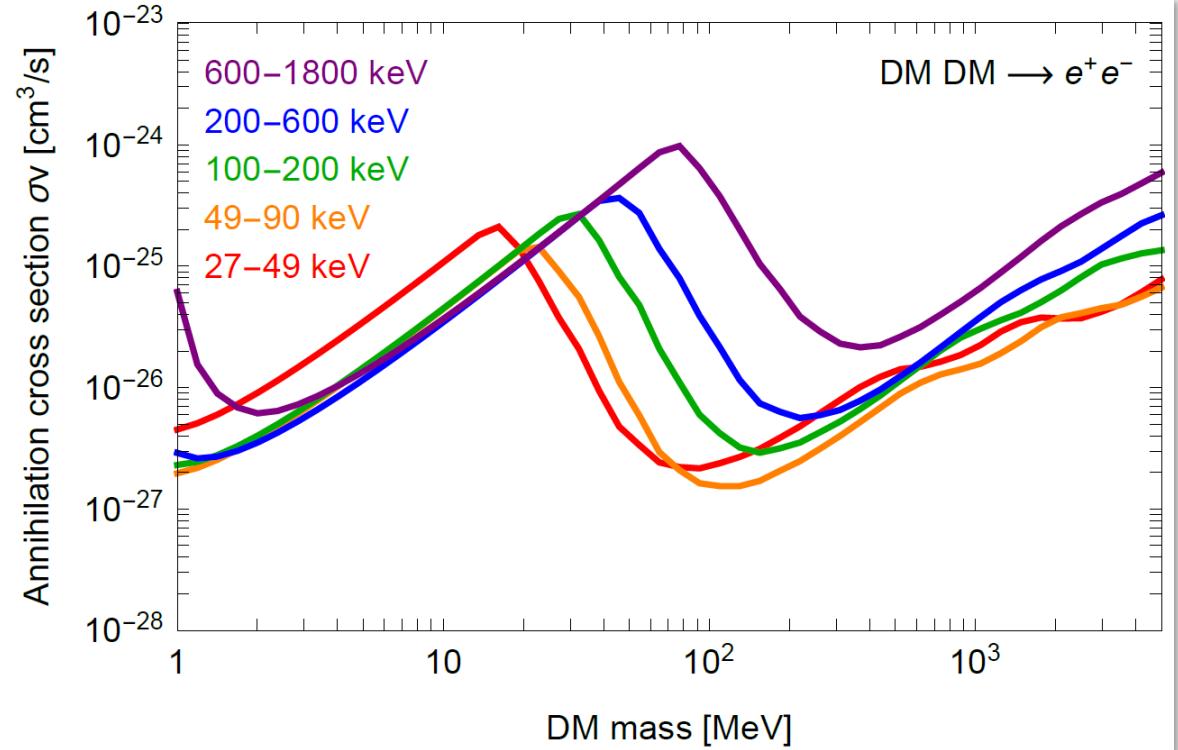
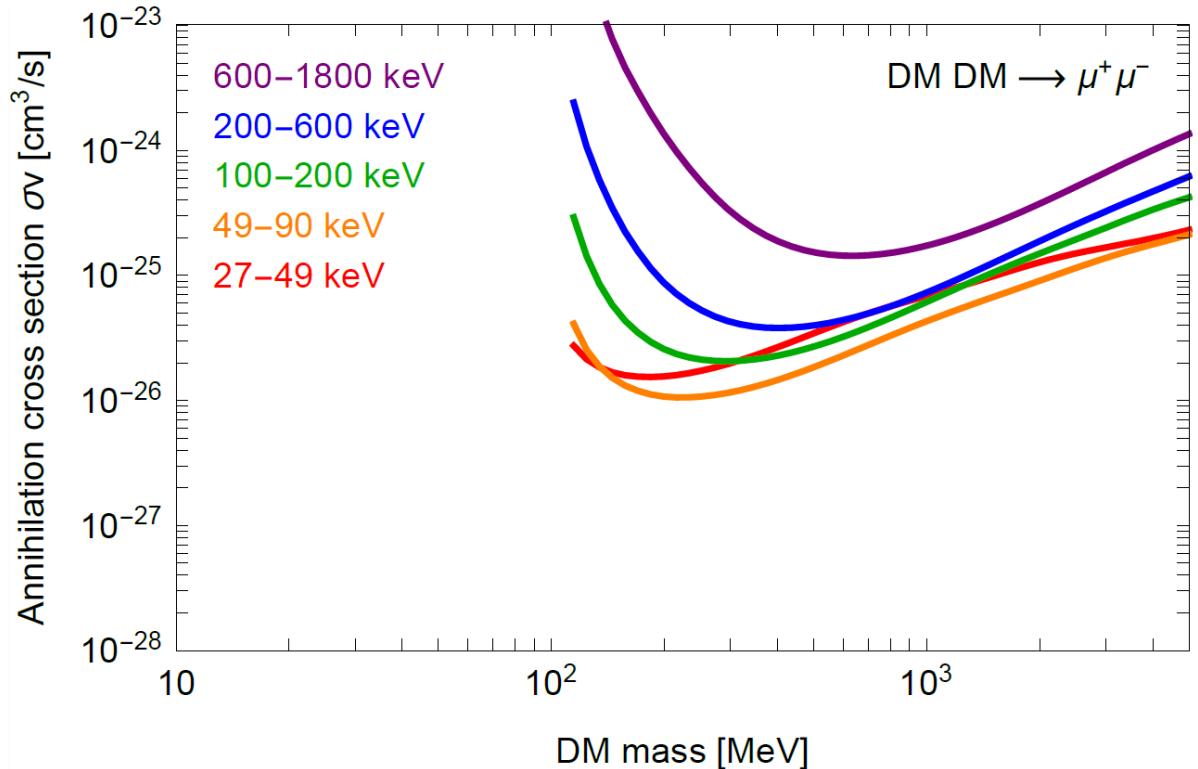


INTEGRAL data (green points)  
Bouchet et al. (2011)

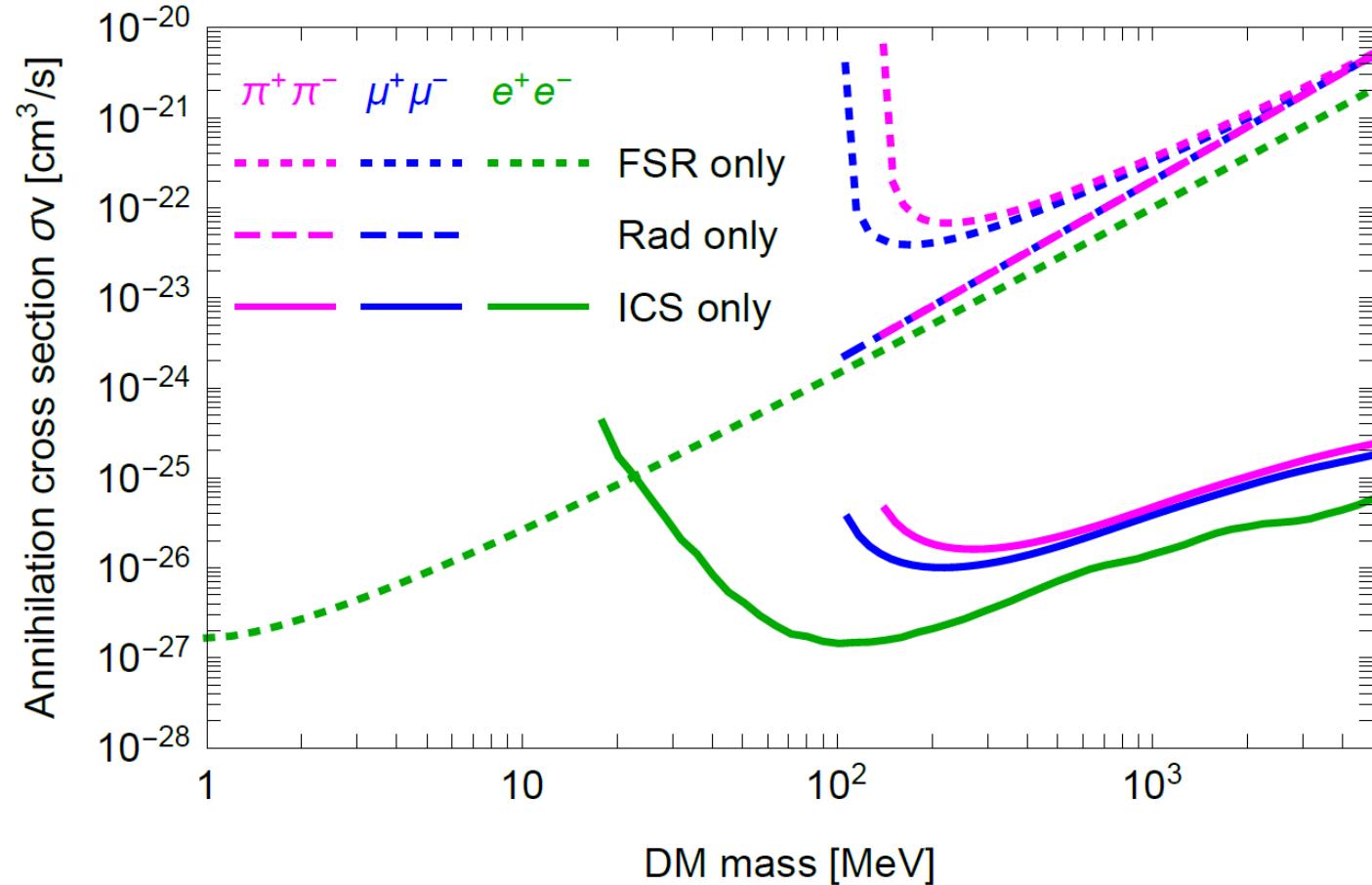
# Angular profile



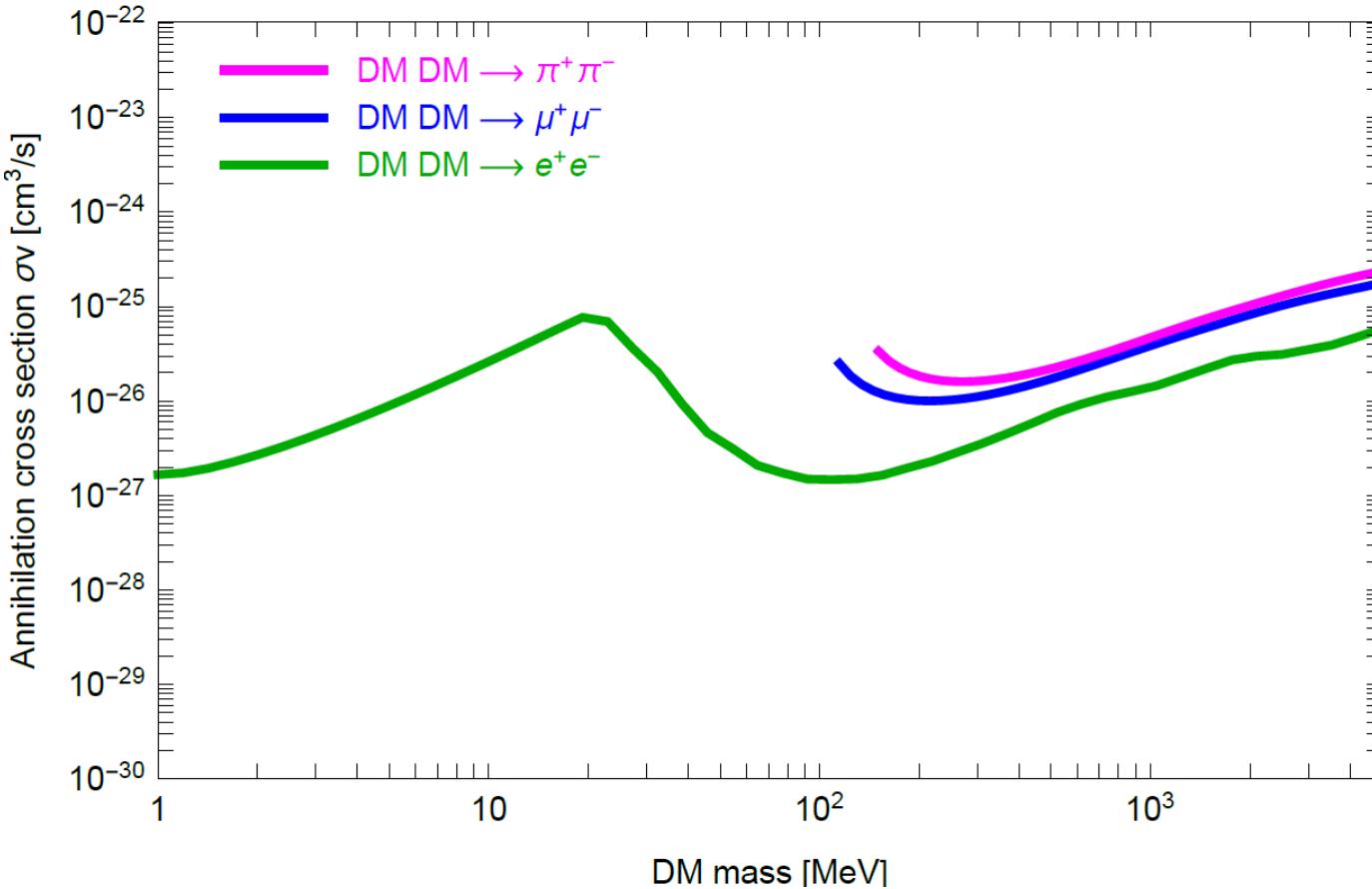
# Bounds in the energy bands



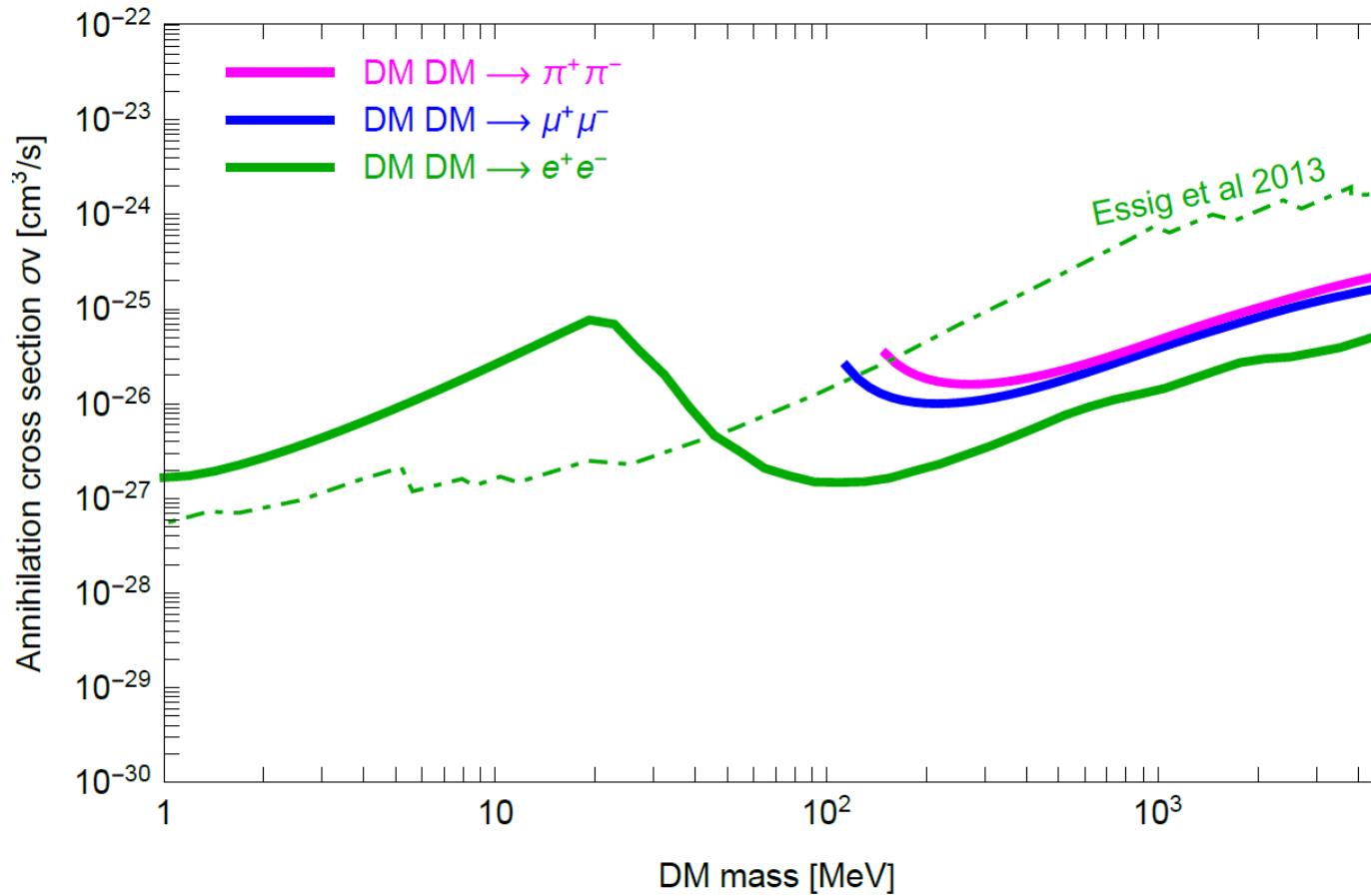
# Bounds from FSR, Rad, ICS



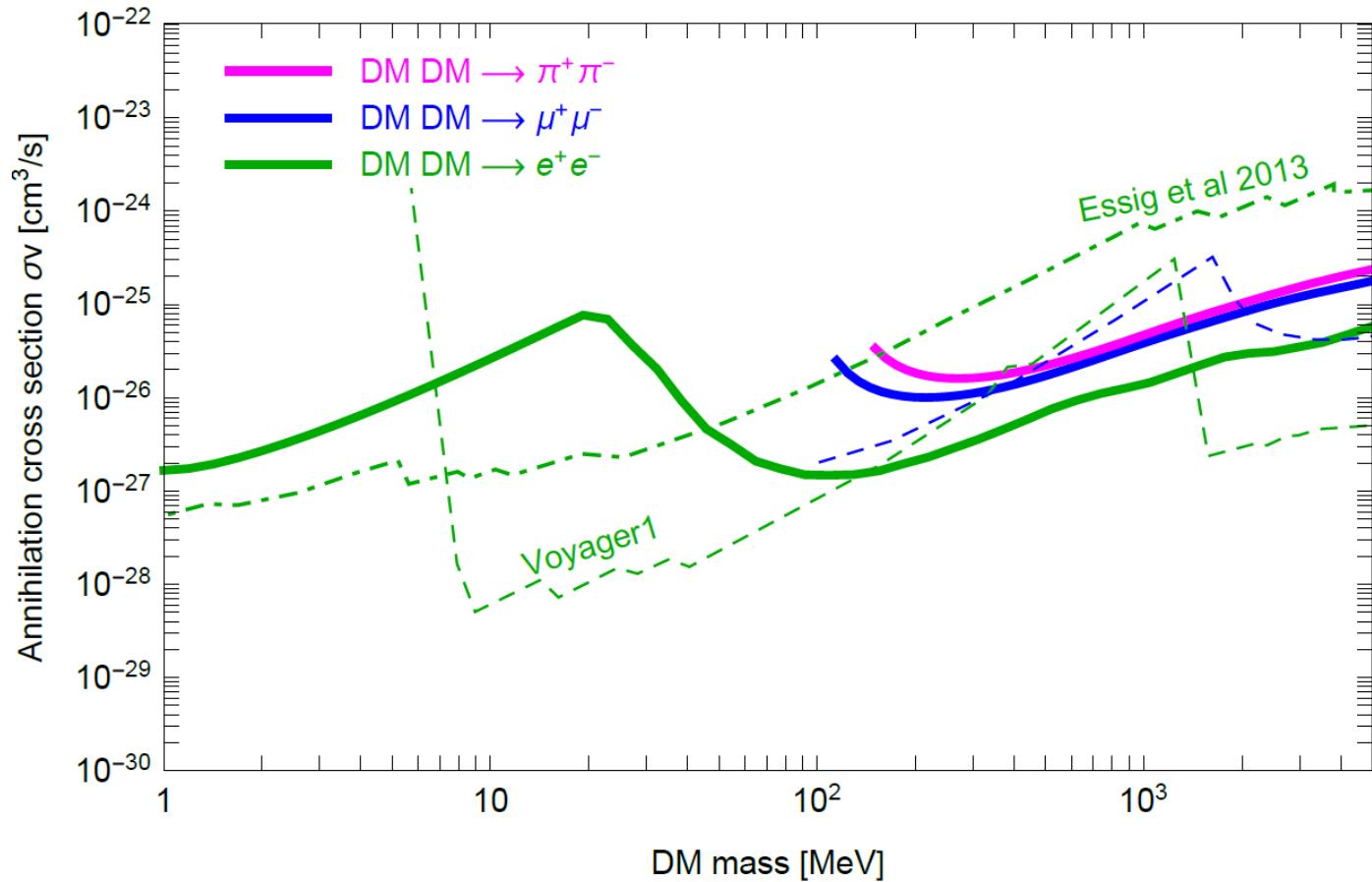
# Constraints on sub-GeV Dark Matter



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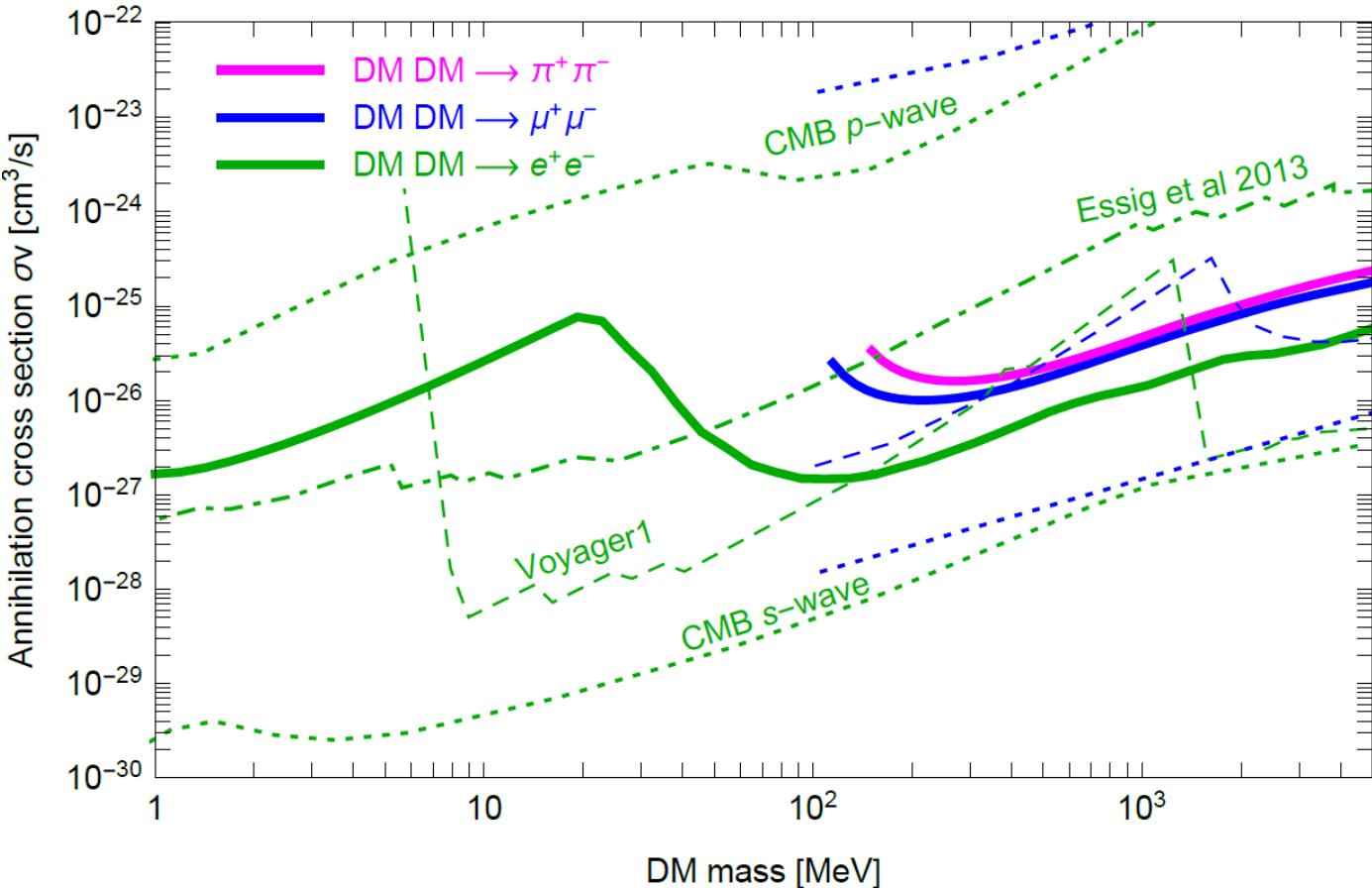


# Constraints on sub-GeV Dark Matter



Boudaud et al., Phys. Rev. Lett. 119  
(2017) 021103

# Constraints on sub-GeV Dark Matter

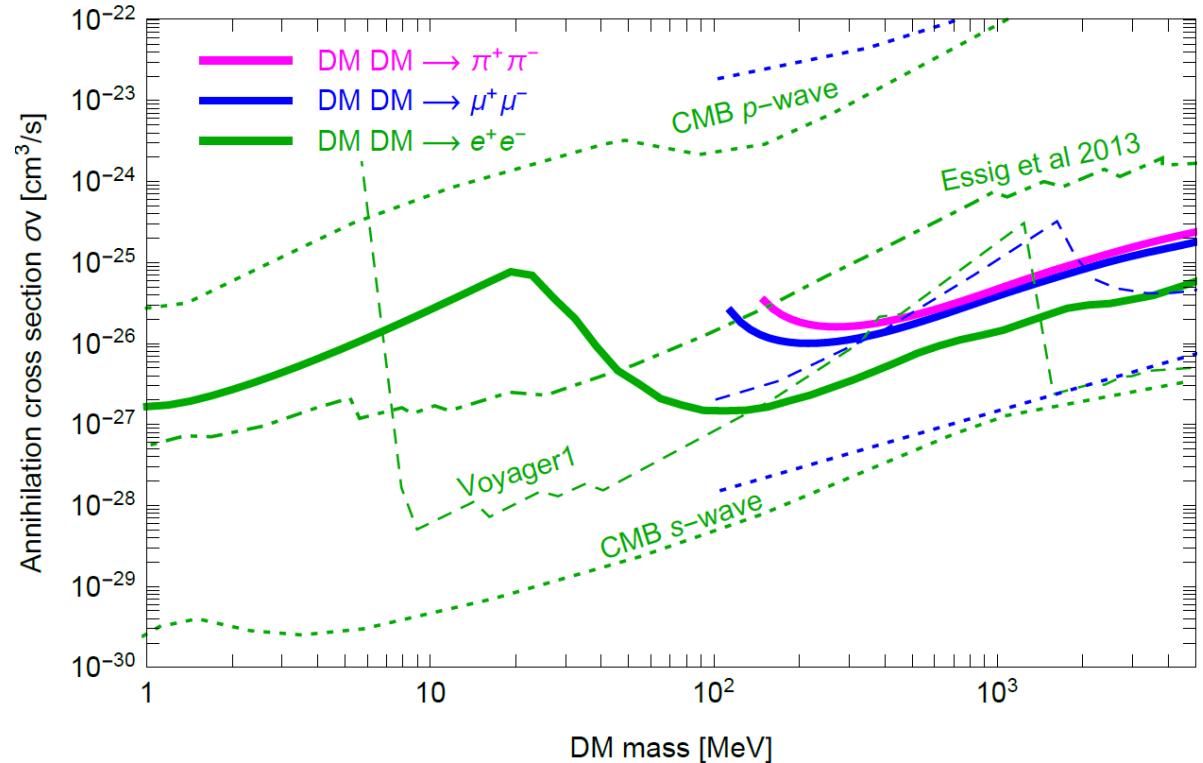


Slatyer, Phys. Rev. D 93 (2016)  
023527  
Lopez-Honorez et al., JCAP 07  
(2013) 046  
Diamanti et al., JCAP 02 (2014) 017  
Liu et al., arXiv:2008.01084

# Conclusions

- $1 \text{ MeV} \leq m_\chi \leq 5 \text{ GeV}$
- Include Inverse Compton Scattering
- X-ray data from INTEGRAL telescope

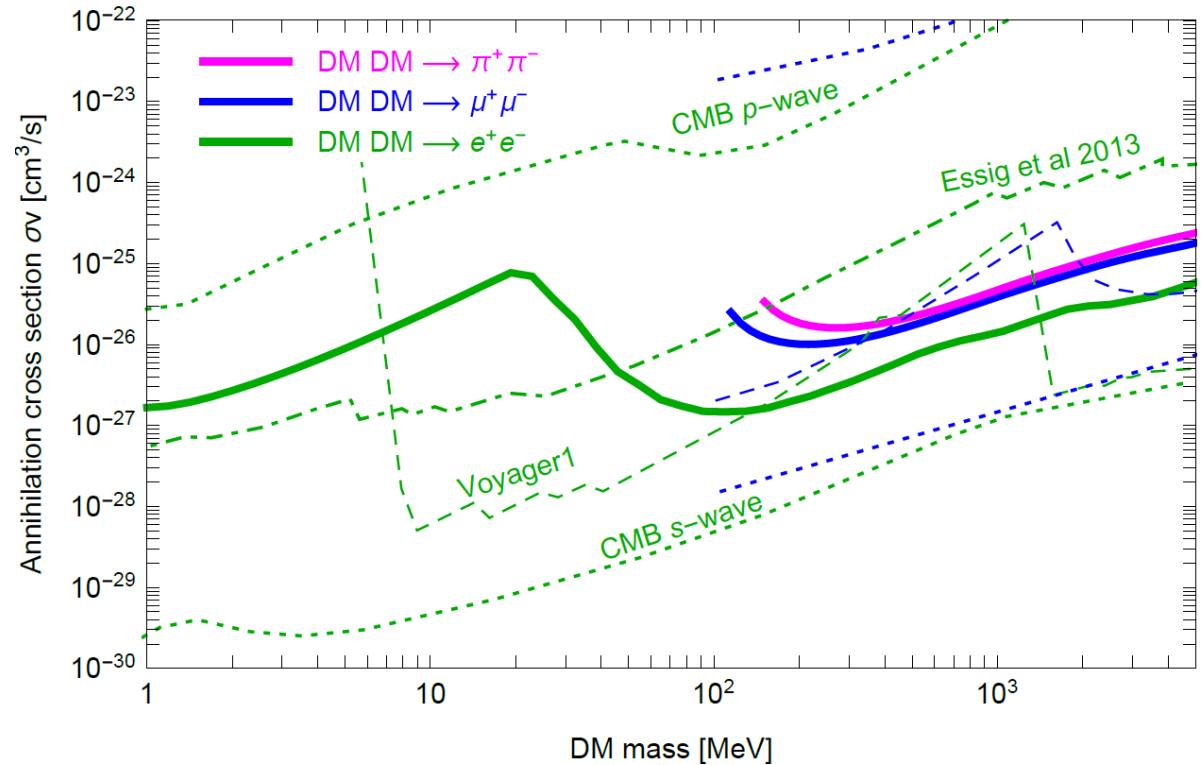
Strongest bound (if p-wave):  
 $150 \text{ MeV} \leq m_\chi \leq 1.5 \text{ GeV}$



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## Thank you for your attention!