



Contribution ID: 36

Type: not specified

The early dark energy resolution to the Hubble tension in light of large-scale structure data and the CMB lensing anomalies

Thursday 10 December 2020 16:30 (20 minutes)

Early Dark Energy (EDE) contributing a fraction $f_{\text{EDE}}(z_c) \sim 10\%$ of the energy density of the universe around $z_c = 3500$ and diluting as or faster than radiation afterwards, can provide a simple resolution to the Hubble tension, the $\sim 5\sigma$ discrepancy –in the Λ CDM context –between the H_0 value derived from early- and late-universe observations. However, the inclusion of Large-Scale Structure (LSS) data, which are in $\sim 3\sigma$ tension with both Λ CDM and EDE cosmologies, might break some parameter degeneracy and alter these conclusions.

I will discuss the viability of the EDE scenario in view of a host of high- and low-redshift measurements, including LSS observations from recent weak lensing surveys, CMB, Baryon Acoustic Oscillation (BAO), growth function and Supernova Ia (SNIa) data, as well as the full-shape galaxy power spectrum from BOSS/SDSS, analyzed using the effective field theory (EFT) of LSS. I will show that the EDE cosmology still provides a potential resolution to the Hubble tension when confronted against current LSS data, though upcoming spectroscopic galaxy surveys, such as Euclid and DESI, will put it under stringent new tests. Finally, I will reassess the EDE scenario in light of the CMB lensing anomalies in Planck data, and I will outline further theoretical extensions that could allow to fully restore cosmological concordance.

Primary author: Dr MURGIA, Riccardo (LUPM, CNRS & Montpellier University)

Presenter: Dr MURGIA, Riccardo (LUPM, CNRS & Montpellier University)

Session Classification: Contributed talks