## The synergy between CMB spectral distortions and anisotropies

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Based on:

Lucca et al. 2019 [1910.04619], Fu et al. 2020 [2006.12886], Lucca 2020 [2008.01115], and Schöneberg et al. 2020 [2010.07814]

In collaboration with:

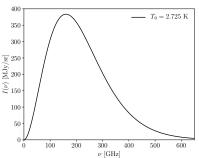
Elia S. Battistelli, Jens Chluba, Hao Fu, Silvia Galli, Deanna C. Hooper, Julien Lesgourgues, and Nils Schöneberg

Presentation for PONT 2020

### What are SDs?

#### Step 1:

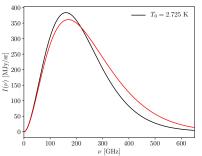
- Forget about the whole universe and everything that there is in it, including the CMB: all we need fits in a box with a hole
- Let's fill the box with only two types of particles, **electrons** and photons, which are in thermal equilibrium
- ► Then, if you look at the energy spectrum of the photons, you should see this:



→ Pure black body (BB) spectrum

#### Step 2:

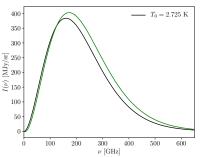
- Let's now imagine to **heat up** one side of **the box** for a tiny amount of time
- ▶ In this time interval we are then effectively heating up the photons that hit the wall
- ► Then, if you look at the energy spectrum of the photons, you should see this (red curve):



→ Shift of low energy photons to higher energies

#### Step 3:

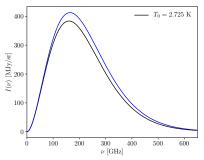
- Let's introduce in the box a scattering process such as Compton scattering
- Given enough time, the scatterings between photons and electrons will redistribute the "injected" energy
- ► Then, if you look at the energy spectrum of the photons, you should see this (green curve):



- → Overall shift of the spectrum
- → Not yet a BB: too many highenergy photons

#### Step 4:

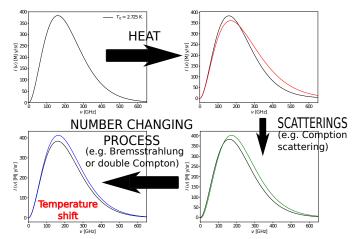
- Let's introduce in the box a **number changing process** such as Bremsstrahlung  $(\gamma + e \rightarrow \gamma + \gamma + e)$
- Given enough time, the few high energy photons will transform in many lower energy photons
- ► Then, if you look at the energy spectrum of the photons, you should see this (blue curve):



→ BB restored but with increased temperature!

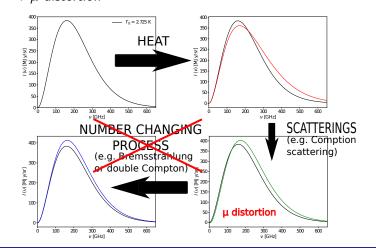
#### Step 5:

- Let's allow the box to **expand**, like the universe!
- At first everything is efficient → Temperature shift



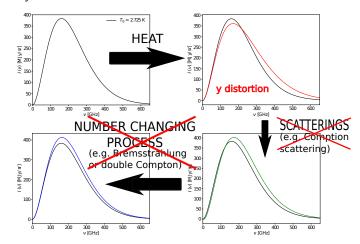
#### Step 5:

▶ Then Bremsstrahlung becomes inefficient (at  $z \simeq 10^6$ )  $\rightarrow \mu$  distortion



#### Step 5:

- ▶ Then **Compton scattering becomes inefficient** (at  $z \simeq 10^4$ )
  - $\rightarrow$  y distortion



#### In summary:

- SDs are deviations from a pure BB in the CMB energy spectrum (caused by the inefficiency of scattering and number-changing processes)
- SDs are very sensitive to energy injections
- Since the primordial SD signal is frozen at decoupling, SDs are a fantastic window into the early universe
- $\rightarrow$  All implemented in CLASS v3.0: First code able to evaluate CMB power spectra, SDs and much more with a single run!

See also Chluba '13 [1304.6120], Chluba&Jeong '14 [1306.5751]

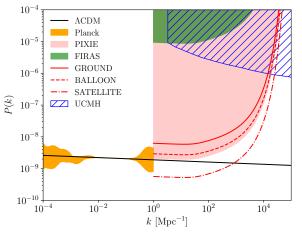
# What are possible sources of SDs?

- In ΛCDM:
  - 1. Adiabatic cooling of electrons and baryons
  - 2. Dissipation of acoustic waves  $(\rightarrow \{n_s, A_s\})$
  - 3. Sunyaev-Zeldovich effect
  - 4. ...
- In exotic scenarios:
  - 1. Non-standard inflationary scenarios
  - 2. **DM decay** or annihilation
  - 3. PBHs evaporation
  - 4. ...

See also Chluba&Sunyaev '12 [1109.6552], Chluba '13 [1304.6121] Chluba '16 [1603.02496], Poulin et al. '17 [1610.10051], Stöcker et al. '18 [1801.01871], Hart et al. '20 [2006.04826]

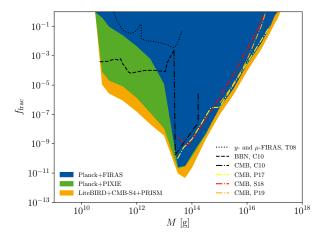
## Concrete examples

#### 1) ACDM (considering different experimental setups)



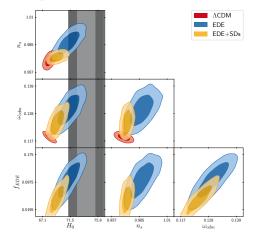
From Fu et al. '20 [2006.12886]

#### 2) Evaporating PBHs (similar results also for decaying DM)



From Lucca et al. '19 [1910.04619]

#### 3) Hubble tension (proof of principle applied to Early Dark Energy)



From Lucca '20 [2008.01115]

#### What can still be done?

- ► Improve CLASS implementation even further including e.g. COSMOSPEC, Chluba&Ali-Haïmoud '16 [1510.03877] SZPACK, Chluba et al. '12 [1205.5778] DARKAGES, Stöcker et al. '18 [1801.01871]
- Constrain a variety of new models such as
   Primordial Magnetic Fields (relevant also for the H<sub>0</sub> tension [2004.09487]) and
   Non-standard DM interactions (e.g. with photons or baryons)
- Apply "Hubble tension idea" to other models like Self-interacting neutrinos [1902.00534] and Interacting majoron-neutrinos [1909.04044]