



Quantum Amplitudes and Classical Gravity

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Also Based on "Classical Solutions and their Double Copy in Split Signature" With Donal O'Connell, David Peinador Veiga, Ricardo Monteiro.

arXiv: 2012.11190.

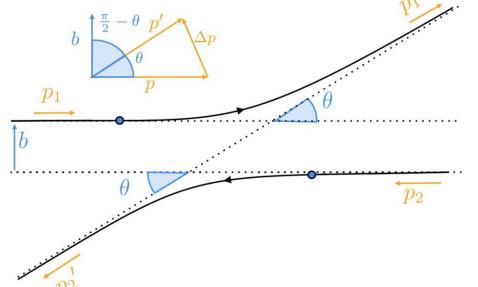
Classical observables from amplitudes



• Classical gravity/EM from amplitudes: $\Delta p^{\mu}, \ \Delta s^{\mu}, \ F^{\mu\nu}_{\rm rad}, \ \Psi^{\alpha\beta\gamma\delta}_{\rm rad}$

Kosower, Maybee & O'Connell, 2018 Maybee, O'Connell & Vines, 2019

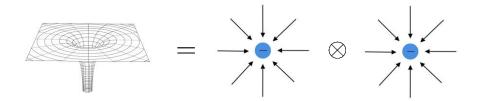
$$\Delta p_1^\mu = \int d^4q \, \delta(p_1 \cdot q) \delta(p_2 \cdot q) e^{-ib \cdot q} \bigg[q^\mu i \mathcal{A}(p_1 \to p_1 + q) \\ + \int d^4l \, \delta(p_1 \cdot l) \delta(p_2 \cdot l) l^\mu \mathcal{A}^*(p_1 + q_1 \to p_1 + l) \mathcal{A}(p_1 \to p_1 + l) \bigg]$$
 in momentum



Taken in the classical limit! " $\hbar \rightarrow 0$ "

Double copy duality





Bern, Carrasco & Johansson, 2008 Monteiro, O'Connell & White, 2014

$$b^{\mu}_{\perp} = \Pi^{\mu}_{\nu}(b + ia_1 + ia_2)^{\nu}$$

$$\Delta p_{\text{Kerr}}^{\mu} = -\frac{2m_1 m_2 G_N}{\sinh w} \text{Re} \left[\frac{\cosh 2w \, b_{\perp}^{\mu} + 2i \cosh w \varepsilon^{\mu\nu\rho\sigma} u_{1\rho} u_{2\sigma} b_{\perp\nu}}{b_{\perp}^2} \right] + \mathcal{O}(G_N^2 a_1^2 a_2^{\infty})$$

Vines, 2017 Arkani-Hamed, Huang & O'Connell, 2019

Guevara, Ochirov & Vines, 2018

 $\Rightarrow \Delta p_{\rm Schw}^{\mu} \propto G_N^4$ Determined up to 3 loops!

Bern, Parra-Martinez, Roiban, Ruf & Shen, 2021

We want to compute curvature invariants



$$S|\psi\rangle = \frac{1}{p} + \frac{1}{p} + \cdots + \cdots + \cdots$$

Lorentz invariant phase space

$$S|\psi\rangle = \int\!\! d\Phi(p)\varphi(p) \exp\left(\sum_h\!\!\int\!\! d\Phi(k)\,\delta(2p\cdot k)\,i\mathcal{M}_{-h}^{(3)}(k)\,a_h^\dagger(k)\right)|p\rangle$$
 Initial state wavepacket

$$\langle \psi | S^{\dagger} \Psi(x) S | \psi \rangle = 2\kappa \operatorname{Re} \int d\Phi(k) \delta(2p \cdot k) i \mathcal{M}_{+}^{(3)}(k) |k\rangle^{4} e^{-ik \cdot x}$$
 "Weyl spinor"
$$\kappa = \sqrt{32\pi G_{N}} \qquad \qquad \mathcal{M}_{h}^{(3)} = -\kappa \left(p \cdot \varepsilon_{h}(k)\right)^{2}$$

Monteiro, O'Connell, Peinador-Veiga & MS, 2020

Conclusion



QFT is useful for classical Physics too!

- Better understanding of loop computations
- Emitted radiation for bound states?
- Inclusion of spin effects

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Thank you!