# Magnetic monopoles in heavy ion collisions

## Arttu Rajantie

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Collaborators: Oliver Gould, David Ho, Cheng Xie MoEDAL Collaboration

## **Magnetic Fields in Heavy Ion Collisions**



(Video: Brookhaven National Laboratory)

# **Magnetic Fields in Heavy Ion Collisions**

- Magnetic field strengths:
  - LHC magnets
  - Magnetars
  - Fixed-target Pb collisions at SPS
  - 5.02 TeV Pb-Pb collisions at LHC
- New physics phenomena?
  - Magnetic monopoles
  - Baryon number violation

 $\begin{aligned} |\vec{B}| &\sim 8.3 \text{ T} \sim 1.6 \times 10^{-15} \text{ GeV}^2 \\ |\vec{B}| &\sim 2 \times 10^{11} \text{ T} \sim 4 \times 10^{-5} \text{ GeV}^2 \\ |\vec{B}| &\sim 5 \times 10^{13} \text{ T} \sim 10^{-2} \text{ GeV}^2 \\ |\vec{B}| &\sim 4 \times 10^{16} \text{ T} \sim 7 \text{ GeV}^2 \end{aligned}$ 

## Magnetic Charges

$$\vec{\nabla} \cdot \vec{E} = \rho_{\rm E}$$

$$\vec{\nabla} \cdot \vec{B} = \rho_{\rm M}$$

$$\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t} - \vec{J}_{\rm M}$$

$$\vec{\nabla} \times \vec{B} = \frac{\partial \vec{E}}{\partial t} + \vec{J}_{\rm E}$$

• Duality  $\vec{E} + i\vec{B} \rightarrow e^{i\theta} (\vec{E} + i\vec{B})$ 

• Dirac quantisation condition  $g \in g_D \mathbb{Z}$ ,  $g_D = \frac{2\pi}{e_0}$ 

## **Monopole Searches in Colliders**



## **Monopole Searches in Colliders**



- Upper bounds on production cross section ATLAS Collaboration, 2019
- To obtain a bound on the monopole mass, one needs to calculate the cross section from theory

## **Monopole Searches in Colliders**



# **Production Amplitude**

- Semiclassical argument for solitonic monopoles: pair production from two-particle collisions suppressed by ~  $e^{-4/\alpha}$  ~  $10^{-238}$ (Witten, Drukier&Nussinov)
- Confirmed numerically for kinks in 1+1D (Demidov&Levkov 2011)
- Production of solitonic monopoles may be practically impossible in two-particle collisions



## **Schwinger Pair Production**

• Energy of a charged particle-antiparticle pair in uniform electric field  $\vec{E}$ :

$$E(\vec{r}) = 2m - \frac{e^2}{4\pi r} - e\vec{E}\cdot\vec{r}$$

 Sauter 1931, Heisenberg&Euler 1936, Schwinger 1951: Tunneling through potential barrier
 pair production from vacuum



# **Schwinger Rate**

Pair production rate per spacetime volume

 $\Gamma \sim \exp(-S_{\text{inst}})$ 

- $S_{\text{inst}}$  is the instanton action
- Prefactor from functional determinant
- Affleck, Alvarez & Manton 1981:
  - Circular worldline instanton with radius  $r = m/e \left| ec{E} 
    ight|$

• Action 
$$S_{\text{inst}} = \frac{\pi m^2}{e|\vec{E}|} - \frac{e^2}{4}$$
  

$$\Gamma = \frac{e^2 |\vec{E}|^2}{8\pi^3} e^{-\frac{\pi m^2}{e|\vec{E}|} + \frac{e^2}{4}}$$
Arbitrary coupling but weak field  $e|\vec{E}| \ll m^2$ 

# **Schwinger Production of Monopoles**

- Calculation does not require weak coupling:  $g \gg 1$  is not a problem
- Pair production rate (constant field, pointline monopoles):

$$\Gamma = \frac{g^2 |\vec{B}|^2}{8\pi^3} e^{-\frac{\pi M^2}{g|\vec{B}|} + \frac{g^2}{4}}$$

- Pair production needs a strong magnetic field  $\left| \vec{B} \right| \gtrsim M^2/g^3$
- LHC magnets  $|\vec{B}| \sim 8.3 \text{ T} \sim 1.6 \times 10^{-15} \text{ GeV}^2$

$$M \gtrsim 1.5 \left(\frac{g}{g_D}\right)^{3/2} \text{keV}$$

► Magnetars  $|\vec{B}| \sim 2 \times 10^{11} \text{ T} \approx 4 \times 10^{-5} \text{ GeV}^2$  $M \gtrsim 0.3 \dots 0.7 \text{ GeV}$ 

## **LHC Heavy Ion Collisions**

Pb-Pb collisions in Nov 2018

B

- $\circ \sqrt{s_{NN}} = 5.02 \text{ TeV},$
- Time-dependent field

$$\vec{B} \approx \frac{B\hat{y}}{2} \left( \left( 1 + \omega^2 \left( t - \frac{z}{v} \right)^2 \right)^{-3/2} + \left( 1 + \omega^2 \left( t + \frac{z}{v} \right)^2 \right)^{-3/2} \right)^{-3/2}$$

where

$$pprox rac{Ze\gamma}{2\pi R^2} pprox 7.3 \ \mathrm{GeV}^2,$$

$$\omega \approx \frac{\gamma}{R} \approx 73 \text{ GeV}$$



## **Spacetime Dependence**

Parameterised by

$$\xi = \frac{M\omega}{gB} \approx M/(2n \text{ GeV})$$

- Constant field:  $\xi \to 0$
- Find worldline instanton in a time-dependent background (Gould, Ho & AR, PRD2019)
- > Zeroth order in self-interaction: Ellipse with axes  $\frac{M}{gB\sqrt{1+\xi^2}}, \frac{M}{gB(1+\xi^2)}$
- Production rate at large  $\xi$ :  $\Gamma \sim \exp(-4M/\omega)$



## **Spacetime Dependence**

- Space and time dependence enhances the production rate (Gould, Ho & AR, 2019)
   ⇒ Stronger mass bounds
- Cannot reach LHC parameters yet:
  - Self-interactions
  - Finite monopole size



Ignoring self-interactions gives the (conservative) estimate M ≥ 80 GeV if no monopoles found

# **Sphalerons for Solitonic Monopoles**

- Georgi-Glashow model: SU(2)+adjoint scalar,  $z = m_s/m_v$
- Unstable classical sphaleron solutions in external field B<sub>ext</sub>: Barrier height
- Energy determines the thermal rate  $\Gamma \propto \exp(-E_{\rm sph}/T)$ : Lower than for pointlike



(Ho and AR, 2020)

• Energy vanishes at Ambjørn-Olesen critical field  $B_{crit} = m_v^2/e$ : Classical monopole production even at zero temperature

# **Instantons for Solitonic Monopoles**

- Georgi-Glashow model: SU(2)+adjoint scalar ,  $\beta = m_{\rm s}/m_{\rm v}$
- 4D instanton solution
- Action determines the quantum Schwinger rate  $\Gamma \propto \exp(-S_{inst})$



<sup>(</sup>Ho and AR, 2021)

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<sup>(</sup>Ho and AR, 2021)

# **Monopoles from Magnetic Fields**

• Constant field 
$$\Gamma = \frac{g^2 |\vec{B}|^2}{8\pi^3} e^{-\frac{\pi M^2}{g|\vec{B}|} + \frac{g^2}{4}}$$

- Time dependence
  - Enhances production
  - Rapid pulse  $\Gamma \sim \exp(-4M/\omega)$
- Spatial inhomogeneity
  - Effect not known
- (Solitonic) monopole size
  - Enhances production
  - Classical instability at  $B_{\rm crit} = m_{\rm v}^2/e$

## **Baryon Number Violation**

- Chiral anomaly in the Standard Model
  - ⇒ Electroweak sphaleron transitions change baryon number
- Rate per spacetime volume

 $\Gamma \sim \exp(-E_{\rm sph}/T)$ 

- Zero field  $E_{\rm sph} \approx 1.9 \times \frac{4\pi\nu}{g} \approx 9 \, {\rm TeV}$
- Non-zero magnetic field: Numerical calculation (Ho&AR 2020)



## **Baryon Number Violation**

Unsuppressed B violation when

 $\left|\vec{B}\right| > \frac{m_{H}^{2}}{e} \approx 5.2 \times 10^{4} \text{GeV}^{2}$ 

Requires

 $\gamma \sim 10^7$ ,  $\sqrt{s_{NN}} \sim 2 \times 10^4 \text{ TeV}$ 

Effect of time-dependence?

