## New Physics Searches in Cosmic Ray Showers

Oliver Fischer

with Marius Bertrand, Maximilian Reininghaus, Ralf Ulrich



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# The stage: Cosmic Rays



Albrecht et al. [2105.06148]

- Interactions above the LHC centre-of-mass energy.
- Cosmic rays provide an untapped source for such interactions.

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# The tool: CORSIKA8 a few words only



See talk by R. Ulrich here.

- New framework to investigate particle cascades in astroparticle physics.
- Open source, joint community project. gitlab.ikp.kit.edu
- ▶ Highly modular, flexible geometry and physics (PYTHIA8).

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#### On the side: Interactions of mesons from Cosmic Rays



Figure 1: Number of hadronic interactions by species in a single EAS (left) as function of energy, and folded with the CR spectrum (right)

M. Reininghaus, PoS ICHEP2020 (2021), 602.

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# Motivation: The muon deficit problem



R. Ulrich et al. [2105.06148]

- Simulated/observed muon densities in extensive air showers.
- Visible for showers above  $\sim 10 \text{ PeV} (10^{16} \text{ eV})$ .
- For all hadronic interaction models.
- Physics (beyond) the Standard Model?

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#### Hypothesis: new process

- 1. Becomes operative at threshold energy, e.g.  $\sim$  100 PeV.
- 2. Produce larger number of muons.

(Muons stem from decaying hadrons.)

3. Cross section  $\sim \sigma_{had}$ 

#### $\Rightarrow$ Large-multiplicity scattering

# Theory: Large-multiplicity scattering



Spannowsky et al., Phys. Rev. D 94 (2016), 085031

- Creation of  $n \gg 1$  of bosons in a single scattering process.
- Large number of diagrams: factorial growth of the amplitude. J. M. Cornvall, Phys. Lett. B 243 (1990), 271-278
- Proposed as solution to the hierarchy problem (Higgsplosion).

V. V. Khoze and M. Spannowsky, Nucl. Phys. B 926 (2018), 95-111

In principle testable at LHC (and other colliders).

J. S. Gainer, [arXiv:1705.00737 [hep-ph]]

•  $n_{
m max}$  limited by  $\sqrt{s}$  ( $\sigma$  estimated as  $\sim$  ab).

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#### Theoretical key questions

- Large multiplicity cross section  $\sigma_{LM}$  grows exponentially.
- Countered by the growth of the off-shell mediator width.
- Maximum at some threshold energy.
- What is its value?
  - 1.  $\sigma_{LM} \ll \sigma_{had}$ : small number of events.
  - 2.  $\sigma_{LM} \gtrsim \sigma_{had}$ : dominant process for energy above threshold.
- Agnostic approach: find observables and study data to constrain the cross section

# The Model

- Model parameters:  $f, \epsilon, b$ .
- A fixed fraction f of  $\sqrt{s}$  is used to to create bosons b in pairs.
- The kinetic energy  $(n_{\text{pairs}}\epsilon)$  is distributed among the *b* pairs.

$$n_{\mathrm{pairs}} = \left\lfloor rac{f\sqrt{s}}{2m_b(1+\epsilon)} 
ight
floor$$

- Isotropic momentum vectors in centre-of-mass system.
- Bosons decay via Pythia8 into long-lived particles.
- Electromagnetic particles are fed into CONEX to generate EM longitudinal profiles by solving the cascade equations.
- Secondary hadronic interactions: QGSJetII-04 above and with UrQMD below 60 GeV.

#### Implemented into CORSIKA8

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First we consider only the case: b = h.

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## Results

Standard Model shower:



Large multiplicity scattering for two values of f:



Kinematics: snapshot a few cm after the scattering.

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Observable I -  $\langle X_{\rm max} \rangle$ 



- Depth of the electromagnetic shower maximum.
- Measured in slant depth along the shower axis.

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# Observable I - $\langle X_{\rm max} \rangle$ continued



- Remember: f controls n<sub>pairs</sub>.
- $\blacktriangleright$  e controls  $E_{kin}$  of the bosons, increases the shower depth.
- For  $f \rightarrow 0$  we get the SM shower back.

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## Observable II - the muon spectrum



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## Outlook

- This was work in progress.
- Next we will also study other bosons b = Z, W.
- Other interesting observables: high-energy lepton bundles.
- Other detectors: IceCube, ATLAS/CMS, DECOR.
- Moreover ...

### The ANITA excess

#### Unusual Near-horizon Cosmic-ray-like Events Observed by ANITA-IV

P. W. Gorham,<sup>1</sup> A. Ludwig,<sup>2</sup> C. Deaconu,<sup>2</sup> P. Cao,<sup>3</sup> P. Allison,<sup>4</sup> O. Baneriee,<sup>4</sup> L. Batten,<sup>5</sup> D. Bhattacharva,<sup>6</sup> J. J. Beatty,<sup>4</sup> K. Belov,<sup>7</sup> W. R. Binns,<sup>8</sup> V. Bugaev,<sup>8</sup> C. H. Chen,<sup>9</sup> P. Chen,<sup>9</sup> Y. Chen,<sup>9</sup> J. M. Clem,<sup>3</sup> L. Cremonesi,<sup>5</sup> B. Dailev,<sup>4</sup> P. F. Dowkontt,<sup>8</sup> B. D. Fox,<sup>1</sup> J. W. H. Gordon,<sup>4</sup> C. Hast,<sup>10</sup> B. Hill,<sup>1</sup> S. Y. Hsu,<sup>9</sup> J. J. Huang,<sup>9</sup> K. Hughes,<sup>4</sup> R. Hupe,<sup>4</sup> M. H. Israel,<sup>8</sup> T.C. Liu,<sup>11</sup> L. Macchiarulo,<sup>1</sup> S. Matsuno,<sup>1</sup> K. McBride,<sup>4</sup> C. Miki,<sup>1</sup> J. Nam,<sup>9</sup> C. J. Naudet,<sup>7</sup> R, J, Nichol,<sup>5</sup> A, Novikov,<sup>12,13</sup> E, Oberla,<sup>2</sup> M, Olmedo,<sup>1</sup> R, Prechelt,<sup>1</sup> B, F, Rauch,<sup>8</sup> J, M, Roberts,<sup>1</sup> A, Romero-Wolf,<sup>7</sup> B. Rotter, <sup>1</sup> J. W. Russell, <sup>1</sup> D. Saltzberg, <sup>14</sup> D. Seckel, <sup>3</sup> H. Schoorlemmer, <sup>15</sup> J. Shiao, <sup>9</sup> S. Stafford, <sup>4</sup> J. Stockham, <sup>12</sup> M. Stockham,<sup>12</sup> B. Strutt,<sup>14</sup> M. S. Sutherland,<sup>2</sup> G. S. Varner,<sup>1</sup> A. G. Vieregg,<sup>2</sup> S. H. Wang,<sup>9</sup> and S. A. Wissel<sup>16</sup> <sup>1</sup>Dept. of Physics and Astronomy, Univ. of Hawaii, Manoa, HI 96822. <sup>2</sup>Dept. of Physics, Enrico Fermi Institute, Kavli Institute for Cosmological Physics, Univ. of Chicago, Chicago IL 60637. <sup>3</sup>Dept. of Physics, Univ. of Delaware, Newark, DE 19716. <sup>4</sup>Dept, of Physics, Center for Cosmology and AstroParticle Physics, Ohio State Univ., Columbus, OH 43210, <sup>5</sup>Dept. of Physics and Astronomy, University College London, London, United Kingdom, <sup>6</sup>Dept. of Mathematics, George Washington University, Washington D.C. <sup>7</sup>Jet Propulsion Laboratory, California Institute of Technology, Pasadena, CA 91109. <sup>8</sup>Dept of Physics & McDonnell Center for the Space Sciences, Washington Univ in St Louis, MO, 63130 <sup>9</sup>Dept. of Physics, Grad. Inst. of Astrophys., & Leung Center for Cosmology and Particle Astrophysics, National Taiwan University, Taipei, Taiwan, <sup>10</sup>SLAC National Accelerator Laboratory, Menlo Park, CA, 94025. <sup>11</sup>Dept. of Electrophysics, National Yang-Ming Chiao Tung University, Hsinchu 30010, Taiwan. <sup>12</sup>Dant of Physics and Astronomy Univ. of Kansas Lawrence, KS 66045

Balloon-borne experiment, [2008.05690]

#### IceCube: astrophysical explanation assuming SM is disfavoured.

"A search for IceCube events in the direction of ANITA neutrino candidates," [2001.01737]

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# BSM in super-LHC energy collisions



- Explanations with TeV-scale resonances and Long-Lived Particles (LLP)
  cf. [1812.00919], [2002.12910], [2004.09464]
- ► IceCube makes compatible observations. D. B. Fox et al. [1809.09615]
- New Physics from resonances with √s > 14 TeV, or from rare decays of SM particles.
- Visible signatures in 'large-scale' experiments.

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# LLP in CR showers



- LLP with masses ~ GeV can be produced in hadron decays.
- Possible in a CR shower even without TeV-scale mediator.
- LLP flux that has to be quantified for every model.
- CR at all energies and angles contribute.
- No new detectors necessary: Dedicated searches should be sensitive already.

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## This is an upcoming topic!

"Unified explanation of flavor anomalies, radiative neutrino masses, and ANITA anomalous events in a vector leptoquark model,"

P. S. Bhupal Dev et al., [2004.09464]

 "New Constraints on Millicharged Particles from Cosmic-ray Production"

R. Plestid, V. Takhistov, Y. D. Tsai, T. Bringmann, A. Kusenko and M. Pospelov [2002.11732]

 "Constraining strongly-coupled new physics from cosmic rays with machine learning techniques"

Spannowsky et al. [1906.09064]

"Searches for Atmospheric Long-Lived Particles"

Coloma, Argüelles et al. [1910.12839]

"Constraining New Physics with High Multiplicity : I. Ultra-High Energy Cosmic Rays on air-shower detector arrays"

Jho & Park [1806.03063]

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# Conclusions

Cosmic rays constitute a ubiquitous source for BSM:

- Resonances at super-LHC energies;
- Long lived particles with masses  $\sim$  GeV;
- $\Rightarrow$  Complementarity.
- A link between a CR simulation framework and BSM models is currently missing.
- Signatures in Cosmic Ray showers.
- Useful to study anomalies in IceCube, ANITA, and others.
- New Physics discovery potential!

Thank you.

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