

NLO Predictions for VBS at the (HL-/HE-)LHC

$pp \rightarrow VV + jj + X$ with leptonic decays
for the LHC
at $\sqrt{s} = 13, 14,$ and 27 TeV

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Vector-boson scattering (VBS): VBF's big(ger) brother

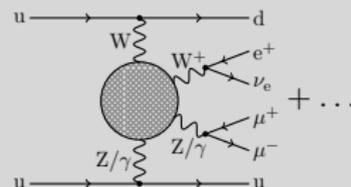
VBS process class: $pp \rightarrow VV + jj + X$ with lept. dec.:

- ① $W^+W^+ : pp \rightarrow e^+ \nu_e \mu^+ \mu^- + jj + X$
 - ② $W^+Z : pp \rightarrow e^+ \nu_e \mu^+ \nu_\mu + jj + X$
 - ③ $ZZ : pp \rightarrow e^+ e^- \mu^+ \mu^- + jj + X$
 - ④ $W^+W^- : pp \rightarrow e^+ \nu_e \mu^- \bar{\nu}_\mu + jj + X$
- + processes related by cc. and different lep. comb.

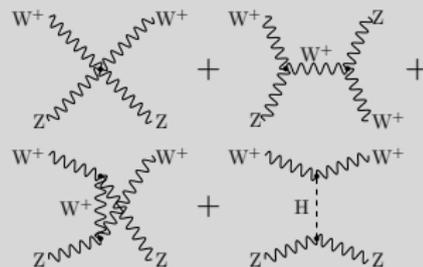
What can we learn?

- nature of QGCs (with triple-boson prod.)
 - form of Higgs-vector-vector couplings
- **EWSB**: interplay between QGC, TGC, and Higgs boson(s)

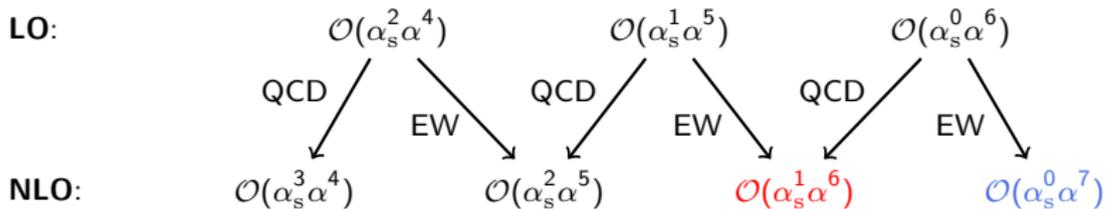
VBS@LHC: W^+Z scattering



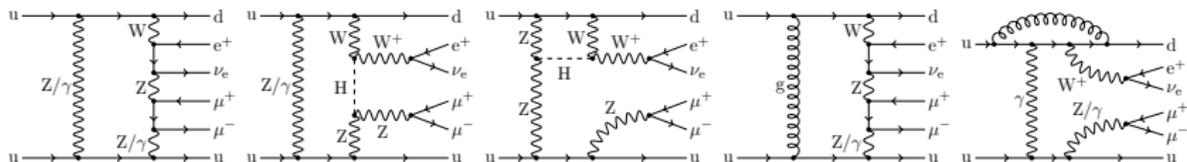
$W^+Z \rightarrow W^+Z$



Coupling structure of VBS: NLOs



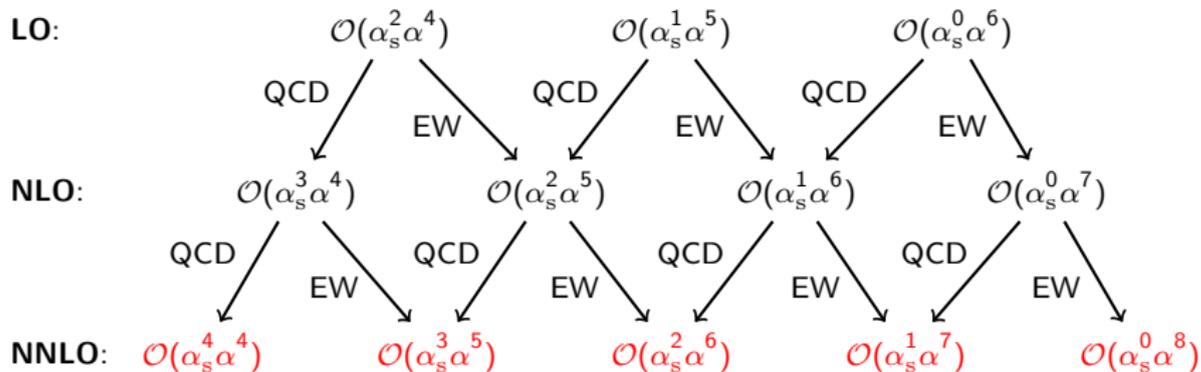
A few diagrams ($pp \rightarrow e^+ \nu_e \mu^+ \nu_\mu + jj + X$):



NLO doable, but only most important are done by now:

- [B. Biedermann, A. Denner, M. Pellen]: complete NLO tower for $pp \rightarrow e^+ \nu_e \mu^+ \nu_\mu + jj + X$
- [A. Denner, S. Dittmaier, P. Maierhöfer, M. Pellen, C.S.]: next-to-leading $\mathcal{O}(\alpha^7)$ and $\mathcal{O}(\alpha_s \alpha^6)$ for WZ
- ZZ and opposite-sign WW will be **very expensive** computationally
- QCD corrections are available, sometimes approximated

Coupling structure of VBS: NNLOs



→ NNLO will be hopeless for a long time (but probably not needed anytime soon?)

WW like-sign scattering: A few references

- [B. Jäger, C. Oleari, D. Zeppenfeld], [A. Denner, L. Hošeková, S. Kallweit]: **approx.** $\mathcal{O}(\alpha_s \alpha^6)$
- [T. Melia, K. Melnikov, R. Röntsch, G. Zanderighi]: $\mathcal{O}(\alpha_s^3 \alpha^4)$
- [M. Rauch]: VBF and VBS review
- [B. Biedermann, A. Denner, M. Pellen]: $\mathcal{O}(\alpha^7)$
- [B. Biedermann, A. Denner, M. Pellen]: complete NLO tower
- [A. Ballestrero et al.]: Tool/approximation comparison for $\mathcal{O}(\alpha_s \alpha^6)$ and PS, PDF uncertainties
- [HL-LHC Collaboration and HE-LHC Working Group; A. Denner, M. Pellen]: 14 and 27 TeV results

WW like-sign: 13 TeV [B. Biedermann, A. Denner, M. Pellen]

Leading orders:

| $\mathcal{O}(\alpha^6)$ [fb] | $\mathcal{O}(\alpha_s \alpha^5)$ [fb] | $\mathcal{O}(\alpha_s^2 \alpha^4)$ [fb] | Sum [fb] |
|------------------------------|---------------------------------------|---|-----------|
| 1.4178(2) | 0.04815(2) | 0.17229(5) | 1.6383(2) |

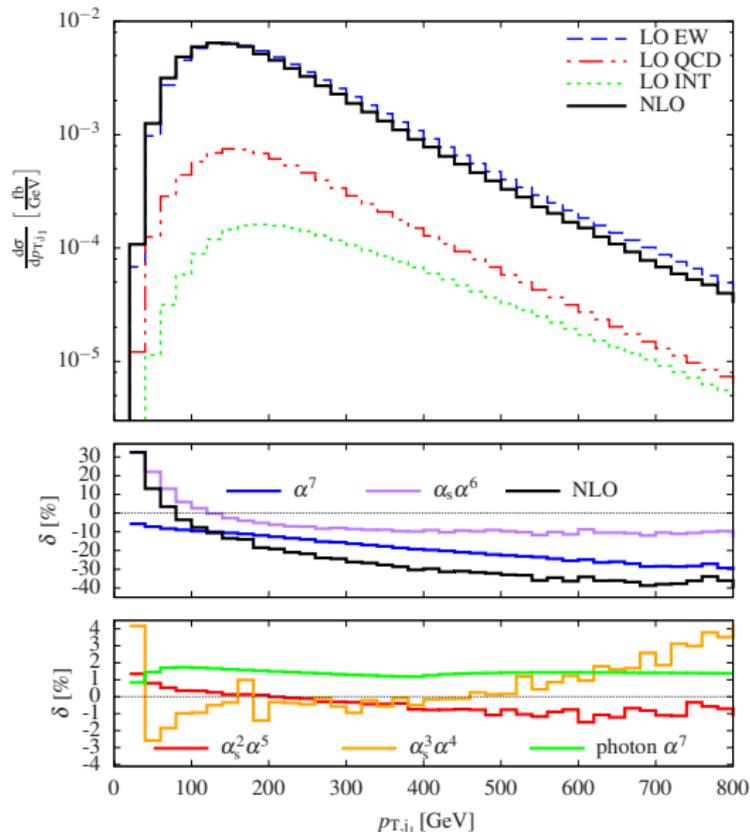
- $\mathcal{O}(\alpha_s^2 \alpha^4)$ kinematically suppressed
- $\mathcal{O}(\alpha_s \alpha^5)$ also colour suppressed

Next-to-leading orders:

| $\mathcal{O}(\alpha^7)$ [fb] | $\mathcal{O}(\alpha_s \alpha^6)$ [fb] | $\mathcal{O}(\alpha_s^2 \alpha^5)$ [fb] | $\mathcal{O}(\alpha_s^3 \alpha^4)$ [fb] | Sum [fb] |
|------------------------------|---------------------------------------|---|---|------------|
| -0.2169(3) | -0.0568(5) | -0.00032(13) | -0.0063(4) | -0.2804(7) |
| -13.2 % | -3.5 % | 0.0 % | -0.4 % | -17.1 % |

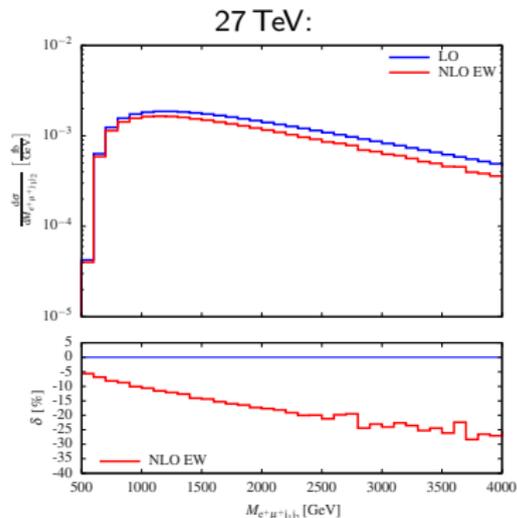
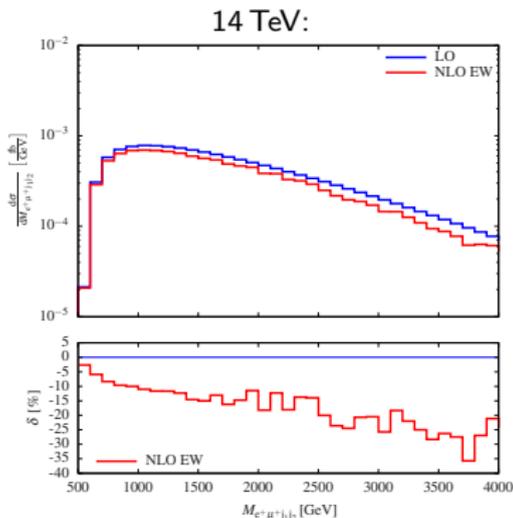
- large EW corrections, only weakly dep. on M_{jj} and Δy_{jj} cuts
- small QCD corrections
- both mixed correction tiny/negligible

WW like-sign: 13 TeV [B. Biedermann, A. Denner, M. Pellen]



- large EW corrections
- even larger EW corrections towards higher transverse momenta (Sudakov logs)
- $\mathcal{O}(\alpha_s \alpha^5)$: large at small p_{j_1} (only soft jets); shape-changing
- $\mathcal{O}(\alpha_s^3 \alpha^4)$ corrections small
- $\mathcal{O}(\alpha_s^2 \alpha^5)$ corrections tiny

WW like-sign: 14 and 27 TeV [A. Denner, M. Pellen]



- slightly different setup w.r.t. 13 TeV
- Large EW corrections
- Slightly increasing NLO EW for 27 TeV
- 14 TeV to 27 TeV: Increase by a factor ~ 3

| | LO ¹ [fb] | NLO EW [fb] |
|--------|----------------------|-------------|
| 14 TeV | 1.4282 | 1.213 |
| | 100 % | -15.1 % |
| 27 TeV | 4.7848 | 3.881 |
| | 100 % | -18.9 % |

¹only $\mathcal{O}(\alpha^6)$

WZ scattering: references

- [G. Bozzi, B. Jäger, C. Oleari, D. Zeppenfeld]: **Approx.** $\mathcal{O}(\alpha_s \alpha^6)$
- [F. Campanario, M. Kerner, L.D. Ninh, D. Zeppenfeld]: $\mathcal{O}(\alpha_s^3 \alpha^4)$
- [J. Bendavid et al.] (SM Les Houches 2017 report, Sec. V.3): LOs, Tool comparison
- [B. Jäger, A. Karlberg, J. Scheller]: Parton-shower effects to $\mathcal{O}(\alpha^6)$
- [A. Denner, S. Dittmaier, P. Maierhöfer, M. Pellen, C.S.]: NLO $\mathcal{O}(\alpha^7)$ and $\mathcal{O}(\alpha_s \alpha^6)$
- [A. Ballestrero, E. Maina, G. Pelliccioli]: Polarisation studies

LO integrated cross sections

Integrated xs for $pp \rightarrow e^+ \nu_e \mu^+ \mu^- jj$ @ $\sqrt{s} = 13$ TeV:

| Sum [fb] | EW [fb] | QCD [fb] | Int. [fb] |
|----------|-----------------------------|----------------------------|-------------------------------|
| 1.55 | $0.255^{+9.03\%}_{-7.75\%}$ | $1.10^{+37.0\%}_{-24.9\%}$ | $0.00682^{+18.4\%}_{-14.4\%}$ |
| 100 % | 16.4 % | 70.6 % | 0.439 % |

| Photons [fb] | Bottom-quarks [fb] |
|--------------------------------|-----------------------------|
| $0.000988^{+11.5\%}_{-9.47\%}$ | $0.195^{+3.59\%}_{-7.22\%}$ |
| 0.0636 % | 12.5 % |

- very large **QCD** contributions mainly due to gluon-PDF
- small **interference** (colour and kinematical suppression)
- smaller **EW** contribution compared to like-sign VBS (\rightarrow Z-boson)
- **photon** contributions completely irrelevant \rightarrow leave out photon-initiated at NLO
- important: **bottom-quark** contributions

NLO integrated cross section

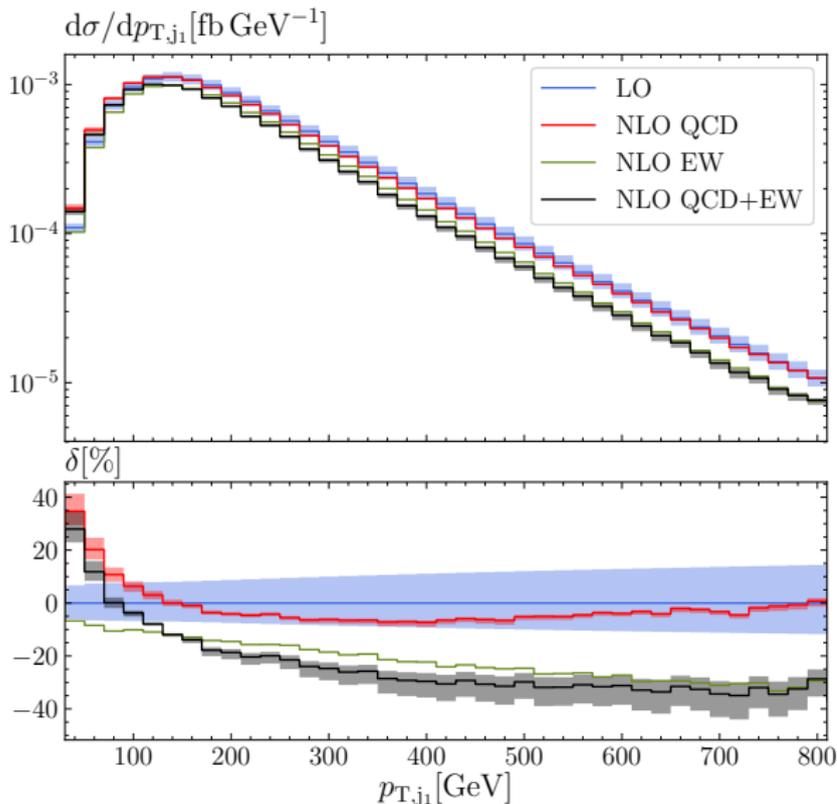
Integrated xs for $pp \rightarrow e^+ \nu_e \mu^+ \mu^- jj + X$:

| \sqrt{s} | LO ² [fb] | NLO EW [fb] | NLO QCD [fb] | NLO EW+QCD [fb] |
|------------|--|------------------|---|--|
| 13 TeV | 0.2551 ^{+9.0%} -7.8% 100.0% | 0.2142 -16.0% | 0.2506 ^{+1.0%} -1.0% -1.8% | 0.2097 ^{+1.3%} -2.2% -17.8% |
| 14 TeV | 0.299 ^{+8.5%} -7.4% 100.0% | 0.251 -16.1% | 0.294 ^{+0.7%} -1.0% -1.8% | 0.245 ^{+0.8%} -0.8% -17.9% |
| 27 TeV | 1.031 ^{+4.6%} -4.3% 100.0% | ? ? | ? ? | ? ? |

- Uncertainties are 7-point QCD-scale variations → NLO EW not accounted for!
- **Large corrections** on the integrated cross section, very similar to like-sign scattering
- QCD corrections small, large EW corrections
- 14 TeV to 27 TeV: Increase by a factor ~ 3

²only $\mathcal{O}(\alpha^6)$

Jet observables 13 TeV



- Leading jet p_T peaks around 140 GeV

- Note that

$$p_{T,j1} > p_{T,j2} > p_{T,j3} > 30 \text{ GeV}$$

- Large positive QCD corrections for small $p_{T,j1}$ (all jets have small transv. momentum)

- EW corr. become increasingly negative; Sudakov logs
- QCD uncertainty band small for large $p_{T,j1}$ due to

$$\mu = \sqrt{p_{T,j1} \cdot p_{T,j2}}$$

Technical challenges of the calculation of $pp \rightarrow e^+ \nu_e \mu^+ \mu^- jj + X$

NLO calculation straightforward thanks to OPENLOOPS and RECOLA, but ...

Virtuals:

- Up to 83 000 Feynman diagrams for each partonic channel
- **ME evaluation expensive**: evaluation of each partonic process takes about ten seconds → parallelisation with MPI [MPI Forum]
- Computational costs: $\sim 8 \times 10^5$ CPU hours for each one of $\{\mathcal{O}(\alpha^7), \mathcal{O}(\alpha_s \alpha^6)\}$; comparable to NNLO QCD calculations, see T. Gehrmann's talk last Thursday

Reals:

- Most complicated example: $\mathcal{O}(\alpha_s \alpha^6)$
 - 40 qq partonic processes for the QCD reals, each has 12–14 dipoles
 - 14 qg partonic processes for the QCD reals, each has 10–11 dipoles
 - 16 qq partonic processes for the EW reals, each has 42 dipoles
- around **1500 dipoles**: automation is key → most complicated NLO computation to date
- **phase space integration** complicated: 5×10^{10} points before cuts (eff.: 14%) → MPI

In general:

- Time consuming **checks** of two calculations against each other, but necessary!

ZZ and WW opposite-sign scattering

- [B. Jäger, C. Oleari, D. Zeppenfeld]: WW approx. $\mathcal{O}(\alpha_s \alpha^6)$
 - [B. Jäger, C. Oleari, D. Zeppenfeld]: ZZ approx. $\mathcal{O}(\alpha_s \alpha^6)$
 - [T. Melia, K. Melnikov, R. Röntsch, G. Zanderighi]: WW $\mathcal{O}(\alpha_s^3 \alpha^4)$
 - [B. Jäger, G. Zanderighi]: WW approx. $\mathcal{O}(\alpha_s \alpha^6)$ matched with PS
 - ZZ $\mathcal{O}(\alpha_s^3 \alpha^4)$?
-
- Any EW corrections are missing

Conclusions

Like-sign WW:

- NLO tower available for like-sign WW scattering
- small QCD, large EW corrections: $\sim 16\%$ (13 and 14 TeV)
- for WW like-sign 27 TeV slightly larger: $\sim 19\%$

WZ:

- for WZ full $\mathcal{O}(\alpha_s \alpha^6)$ and $\mathcal{O}(\alpha^7)$ now known, results similar to like-sign WW,
- $\mathcal{O}(\alpha_s^2 \alpha^5)$ still is missing

ZZ/opposite WW:

- for ZZ and opposite-sign WW only QCD corrections are known
- full NLO will be expensive

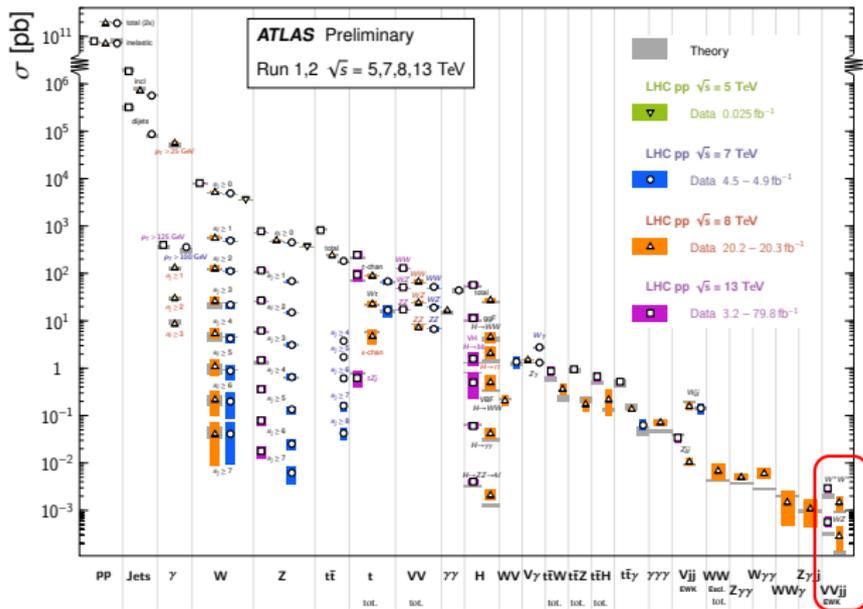
→ QCD \ll EW: FO uncertainty estimation?

VBS at the LHC, $pp \rightarrow e^+ \nu_e \mu^+ \mu^- jj + X$

VBS: $\sigma \approx \mathcal{O}(1 \text{ fb}) \rightarrow$ need large \sqrt{s} and \mathcal{L} : new class of rare processes **accessible in run II**

Standard Model Production Cross Section Measurements

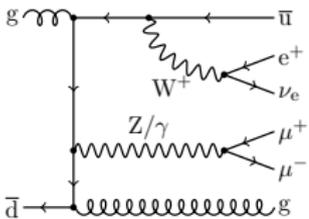
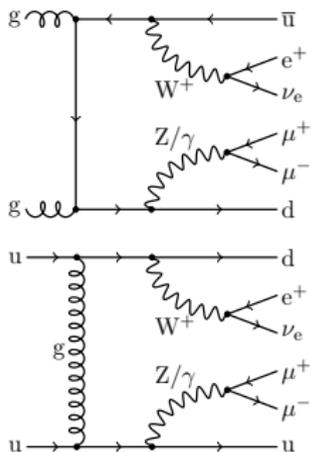
Status: July 2019



(staircase plot from the [ATLAS Collaboration])

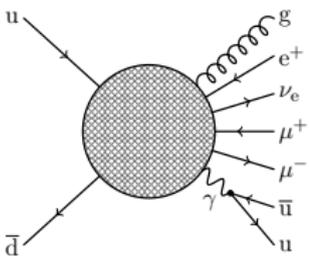
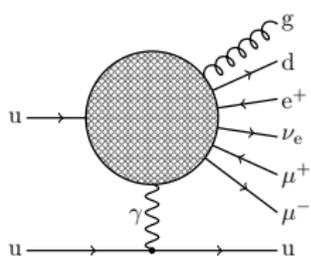
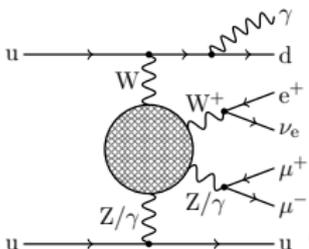
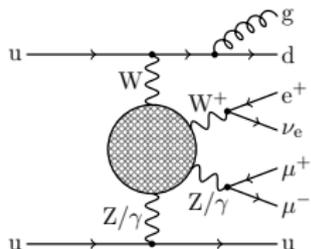
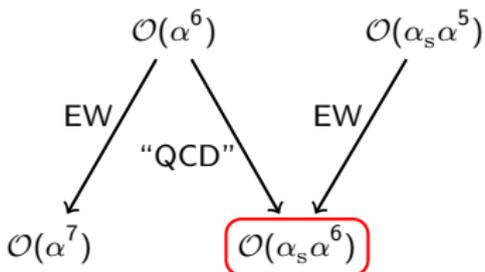
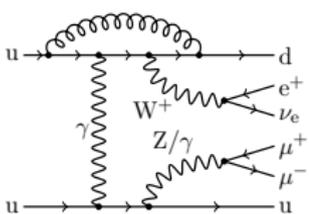
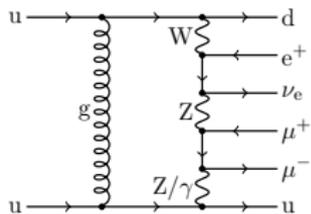
- Largest VBS channel is $W^+W^+ \rightarrow W^+W^+$, full NLO corrections available [Biedermann, Denner, Pellen]
- Second largest channel: $W^+Z \rightarrow W^+Z$

Strong production: $\mathcal{O}(\alpha_s^2 \alpha^4)$



- In comparison to like-sign W-scattering gluons are possible at LO (charge)
- 8 additional MEs with **two gluons**, making up 66 % of the cross section
- in total the $\mathcal{O}(\alpha_s^2 \alpha^4)$ is **4.3 times larger** than the electroweak LOs

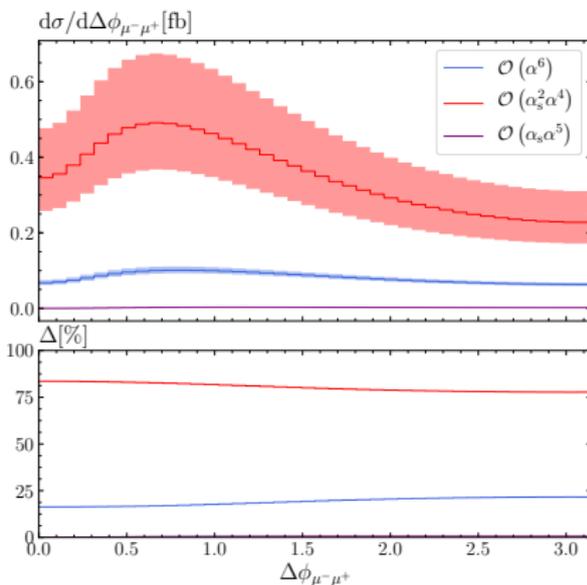
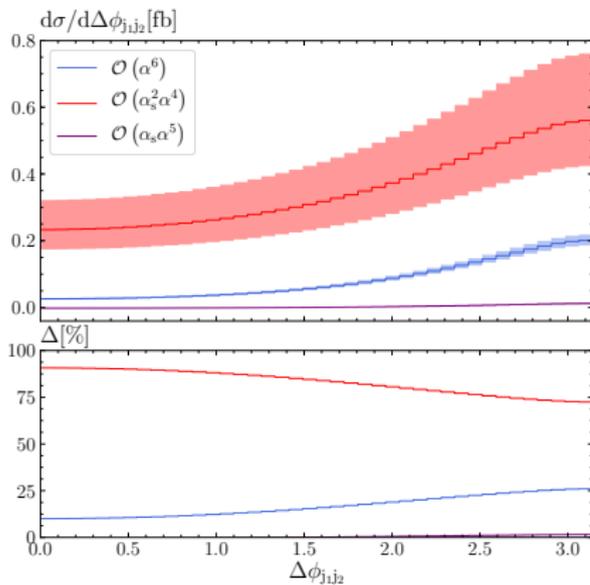
$\mathcal{O}(\alpha_s \alpha^6)$ mixed corrections: "QCD"



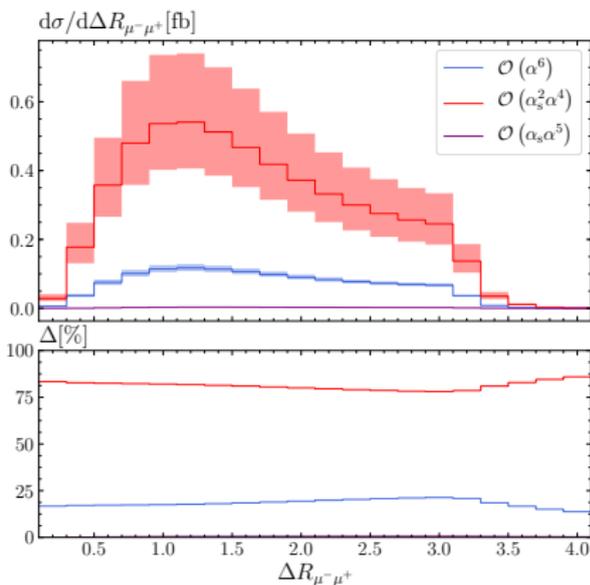
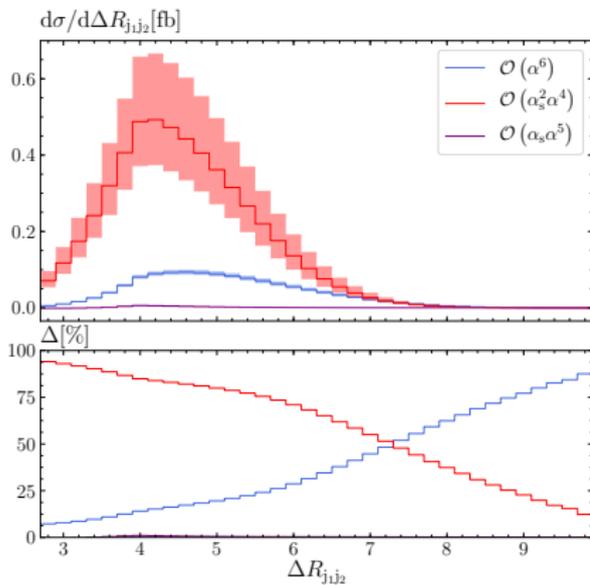
QED singularities in "QCD":

- initial state: cancelled with collinear counterterm (PDFs)
 - final state: **photon-to-jet conversion function** [Denner, Dittmaier, Pellen, C.S.]
- Correction is neither purely QCD/EW, it is mixed

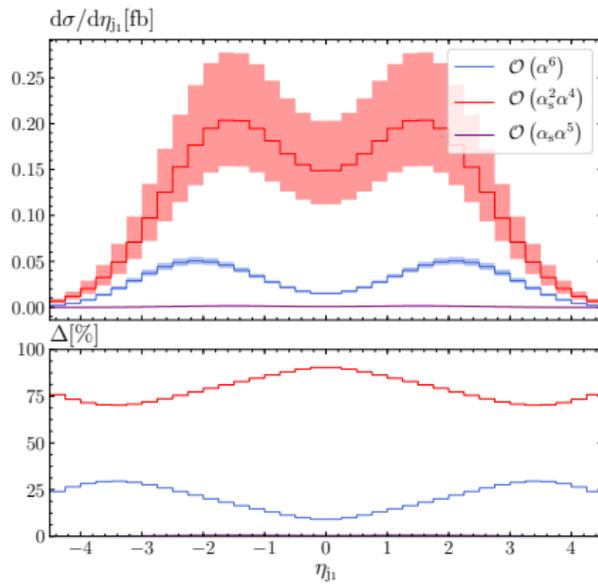
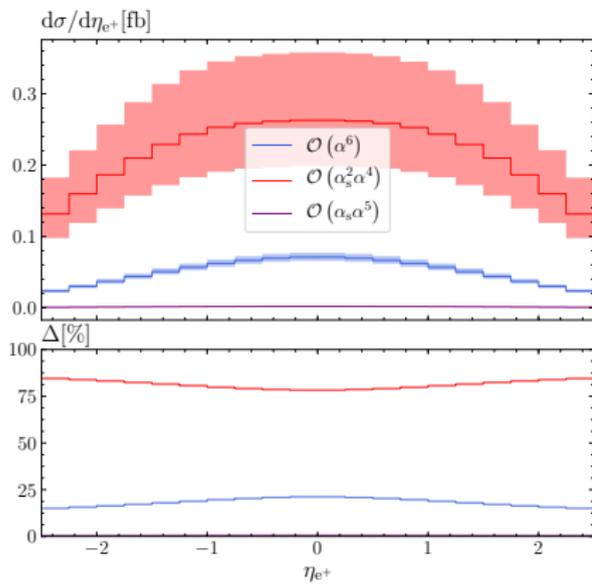
LO distributions (II)



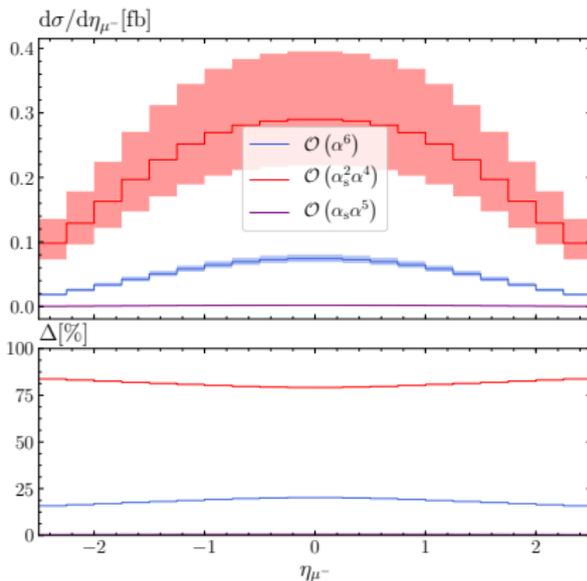
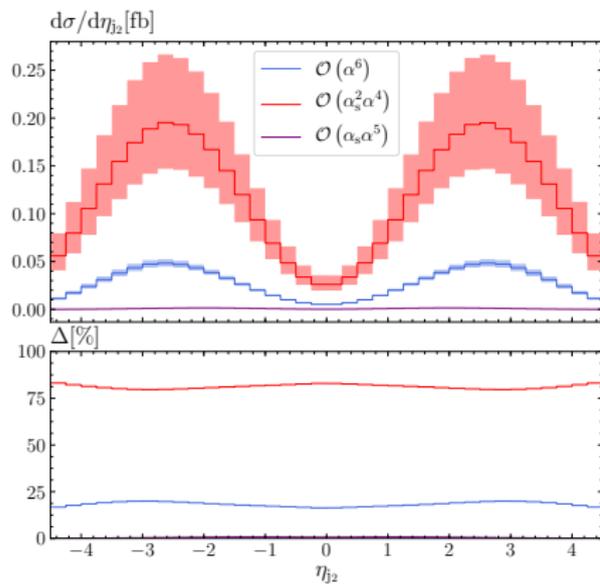
LO distributions (III)



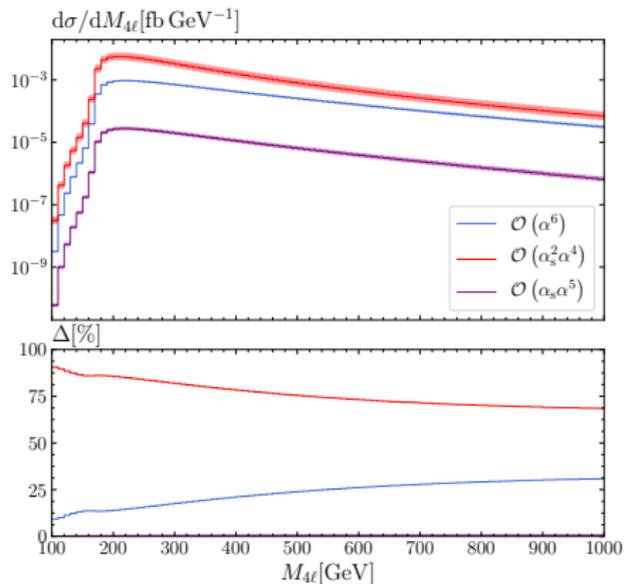
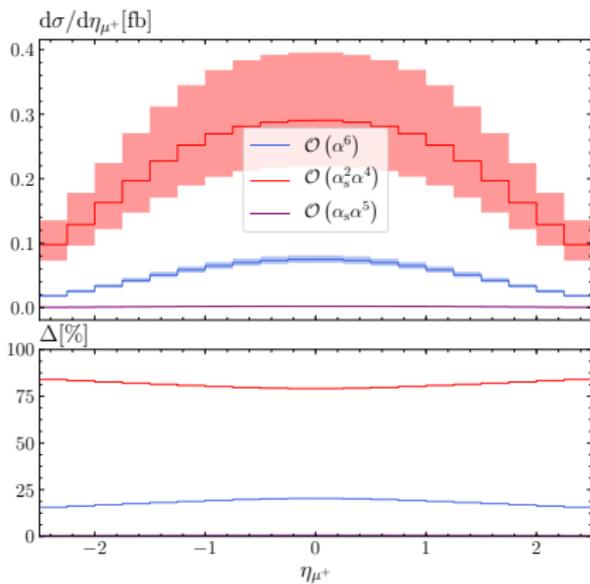
LO distributions (IV)



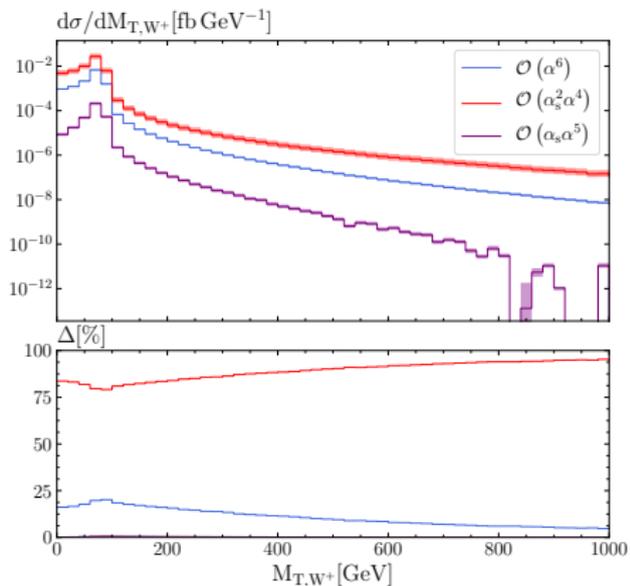
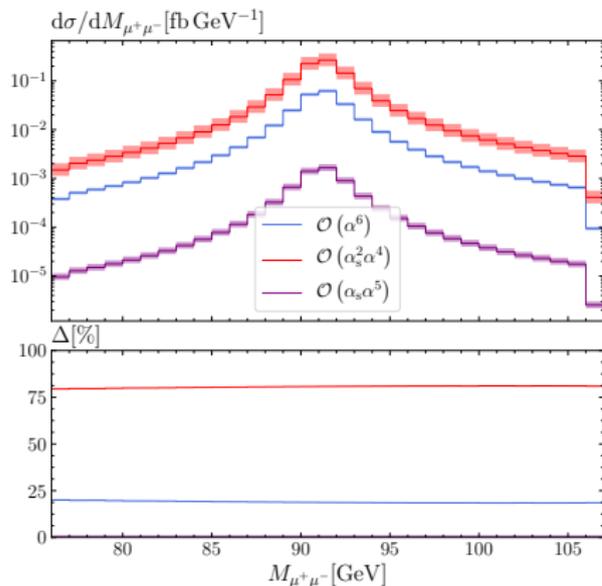
LO distributions (V)



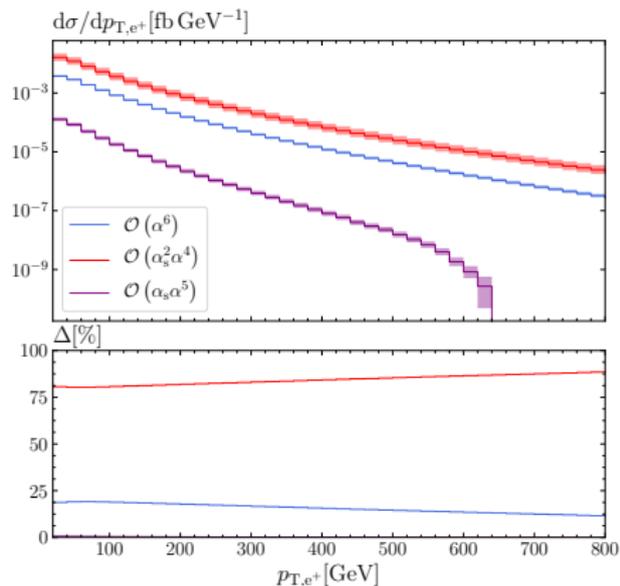
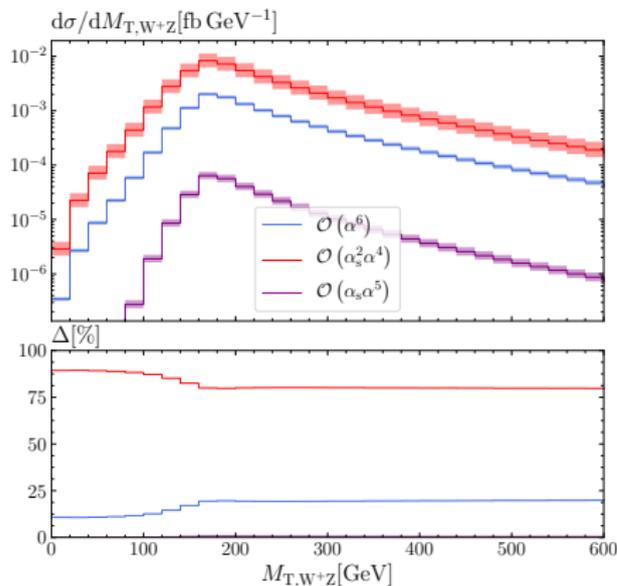
LO distributions (VI)



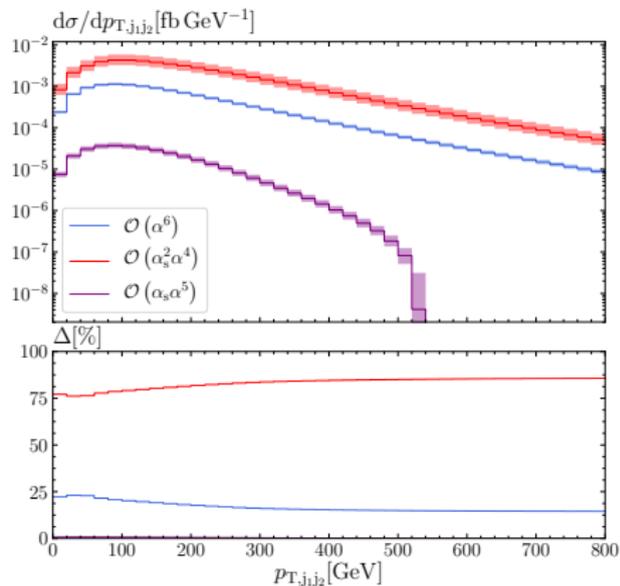
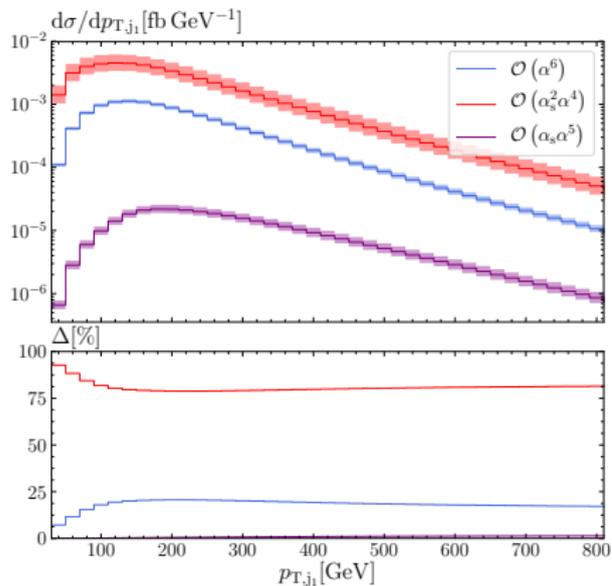
LO distributions (VIII)



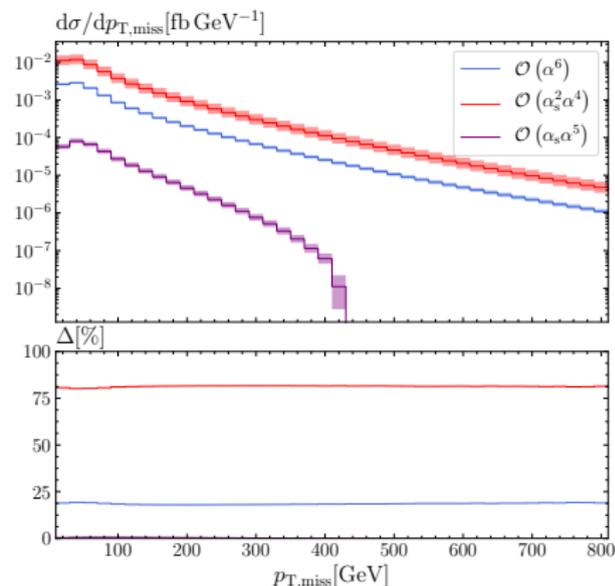
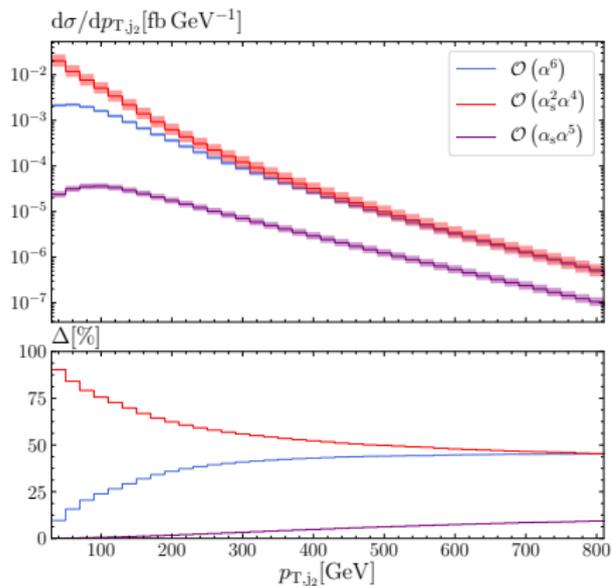
LO distributions (IX)



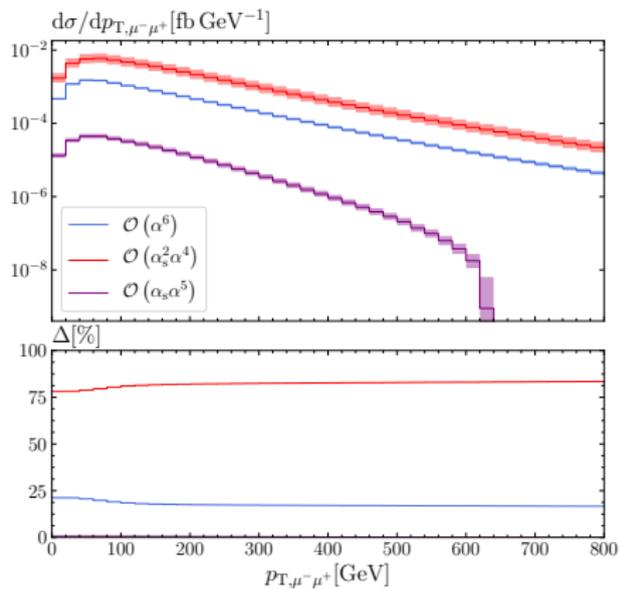
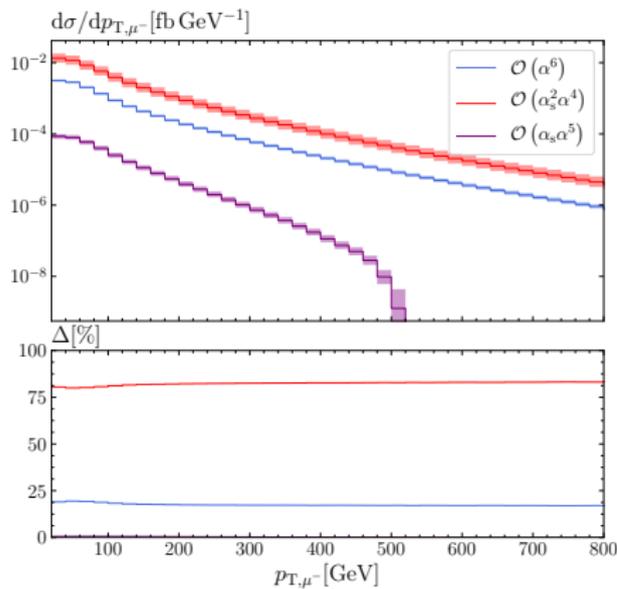
LO distributions (X)



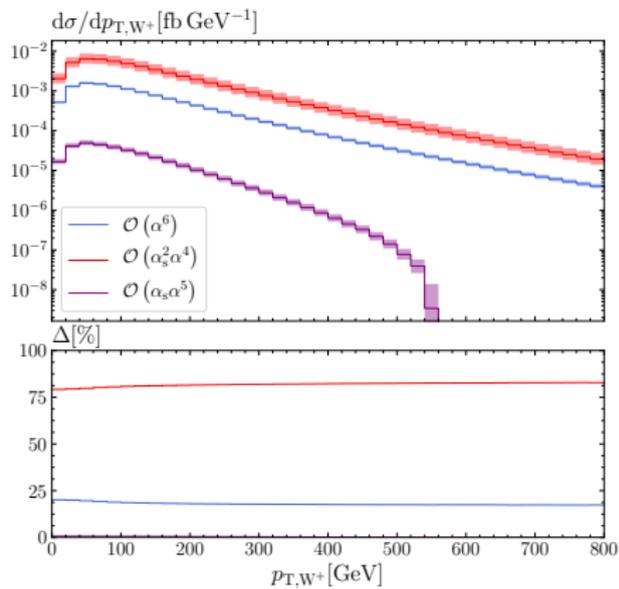
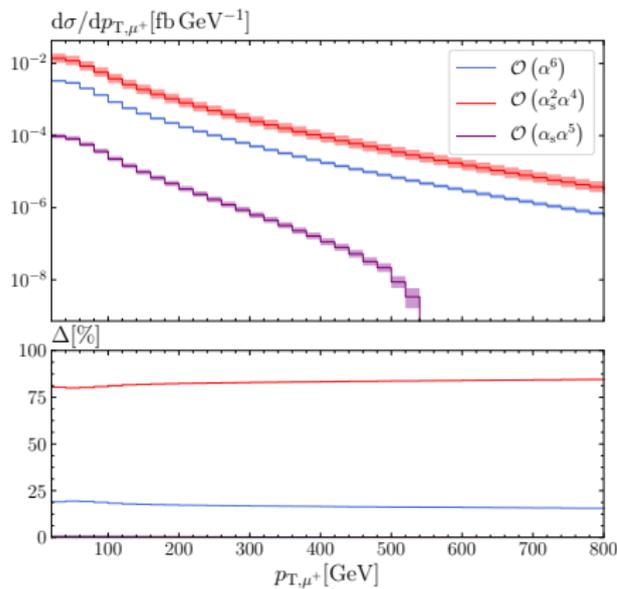
LO distributions (XI)



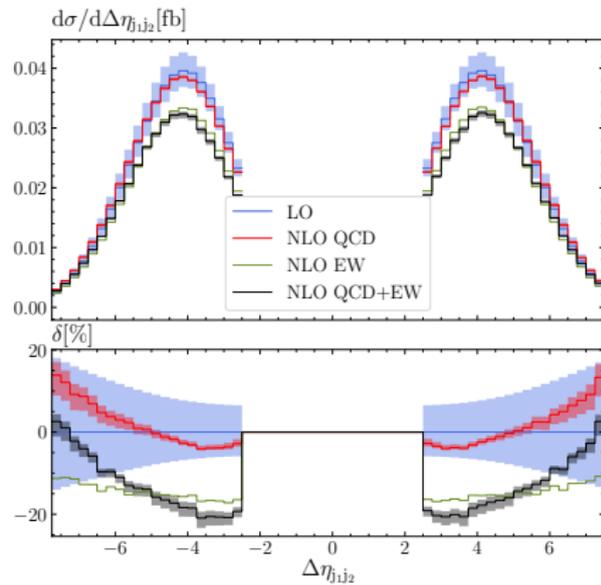
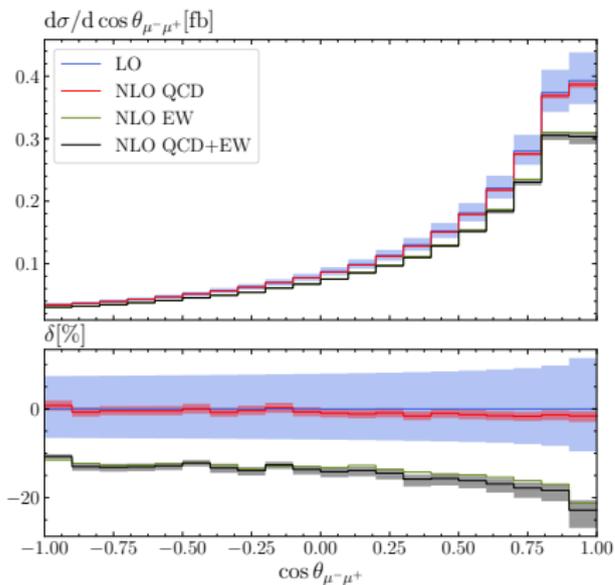
LO distributions (XII)



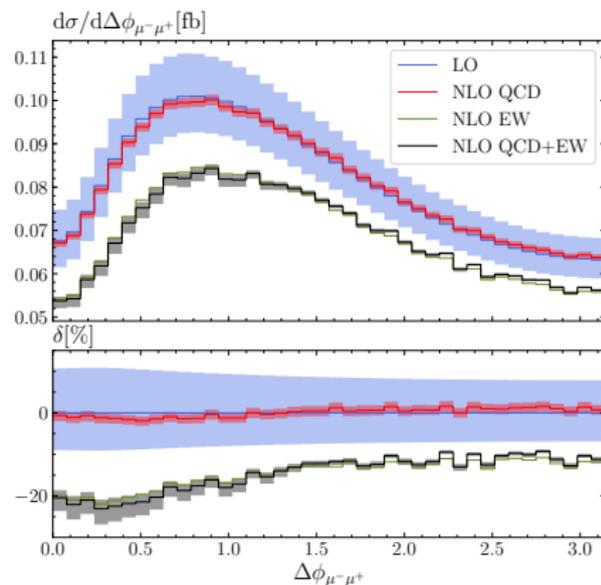
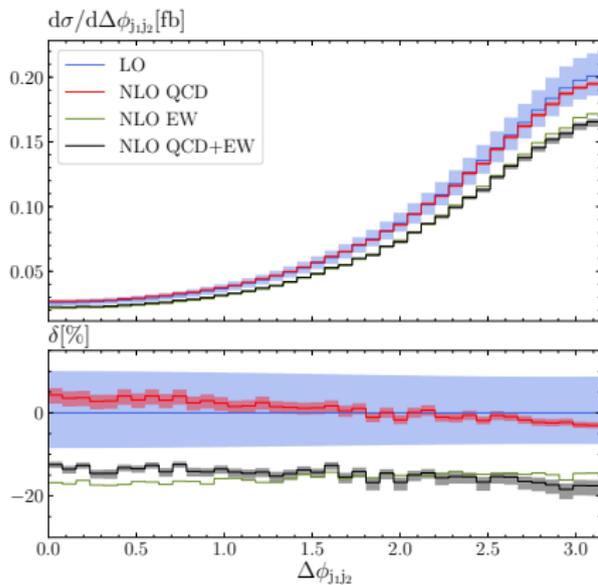
LO distributions (XIII)



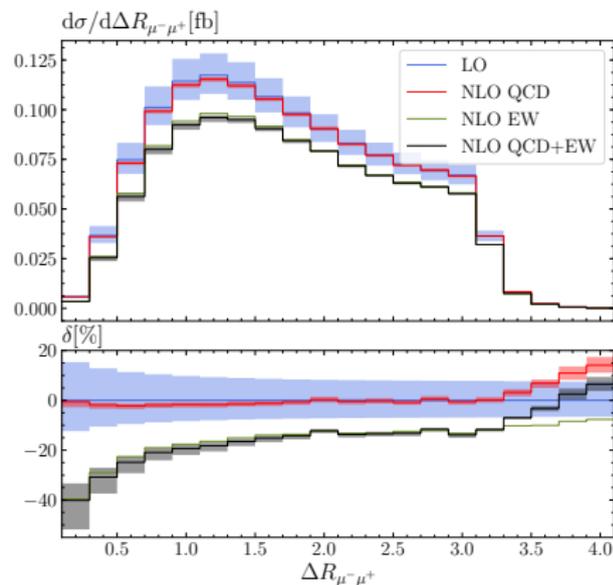
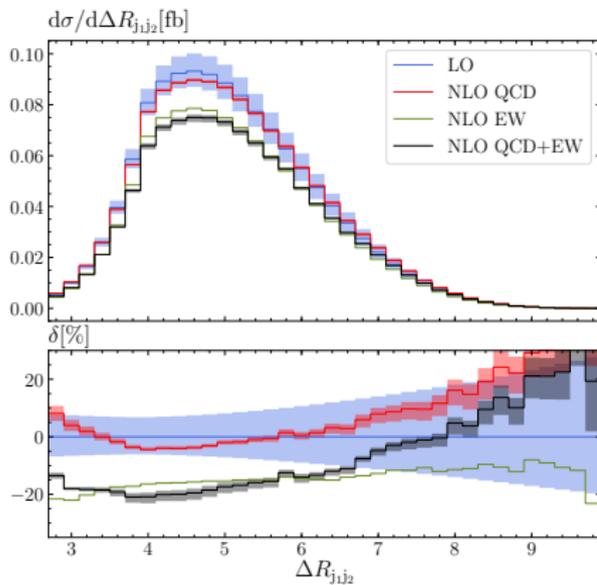
NLO distributions (I)



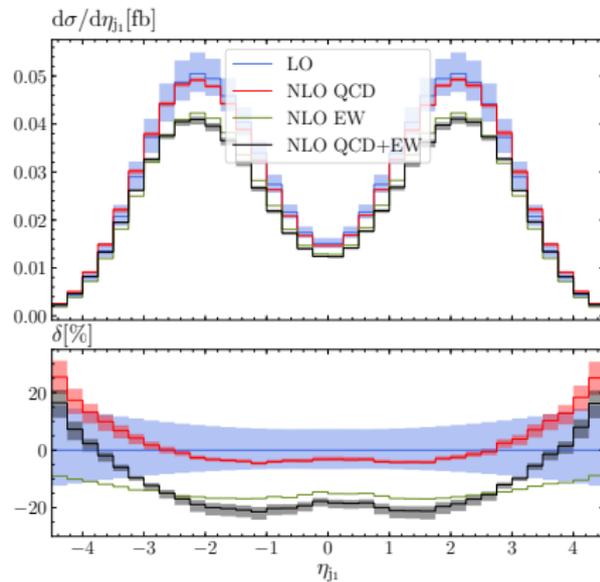
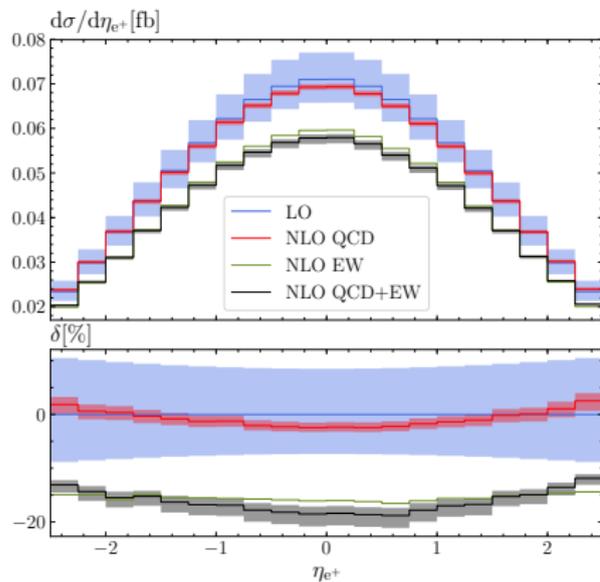
NLO distributions (II)



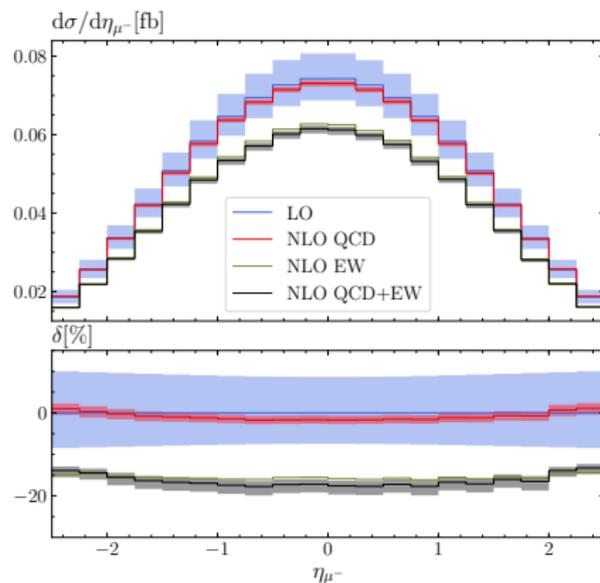
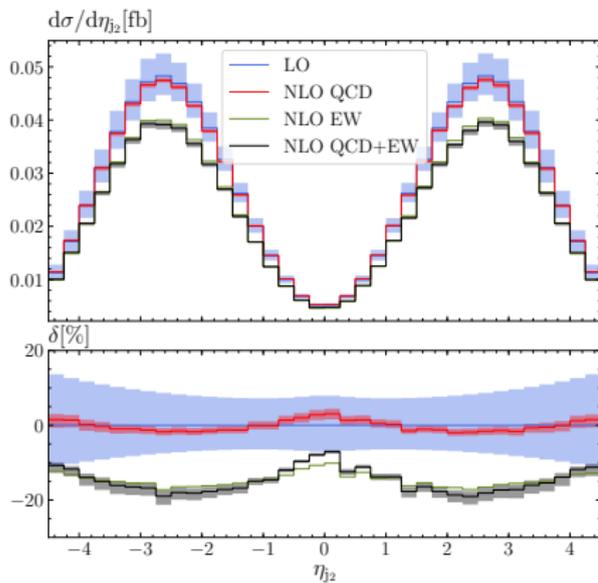
NLO distributions (III)



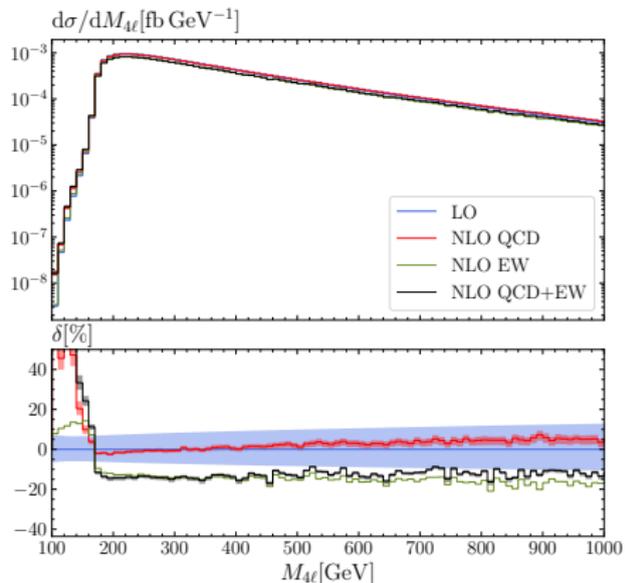
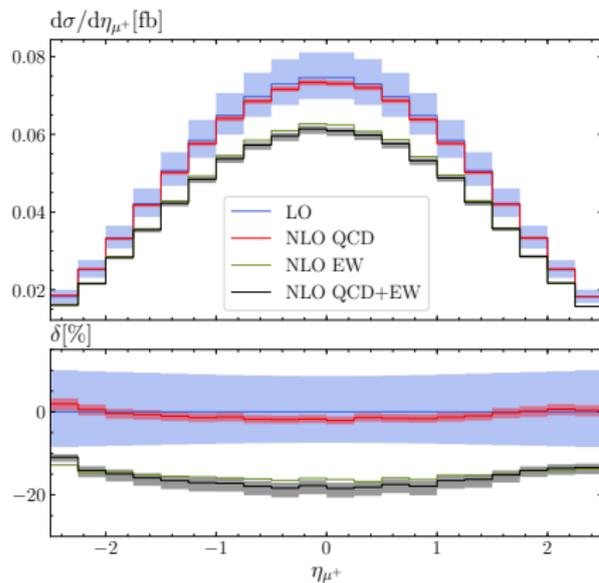
NLO distributions (IV)



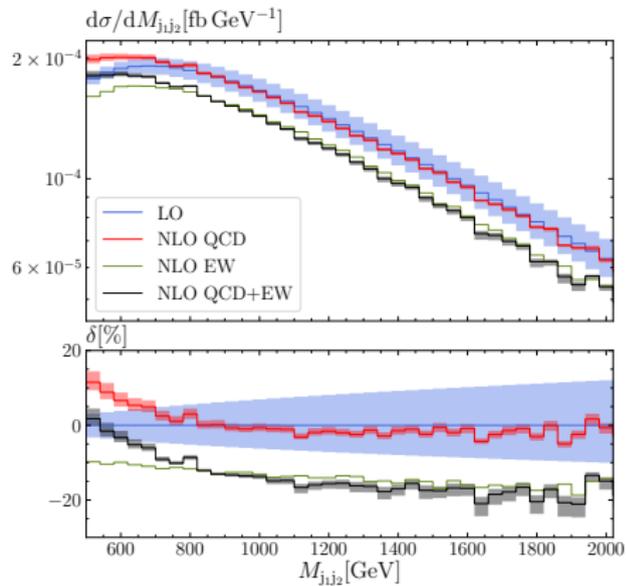
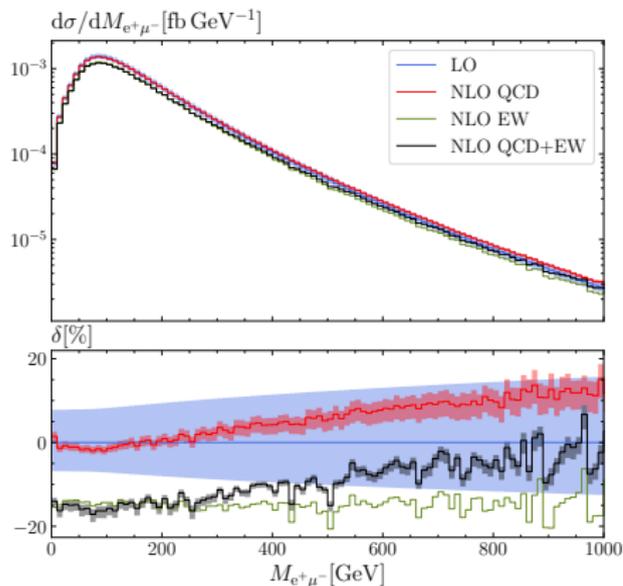
NLO distributions (V)



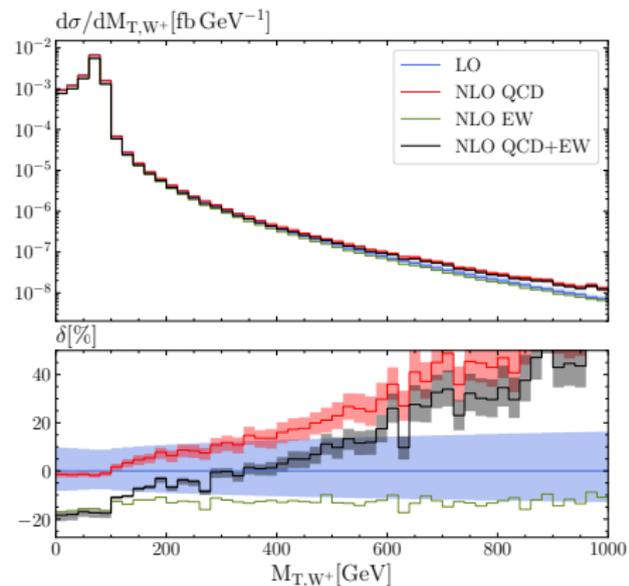
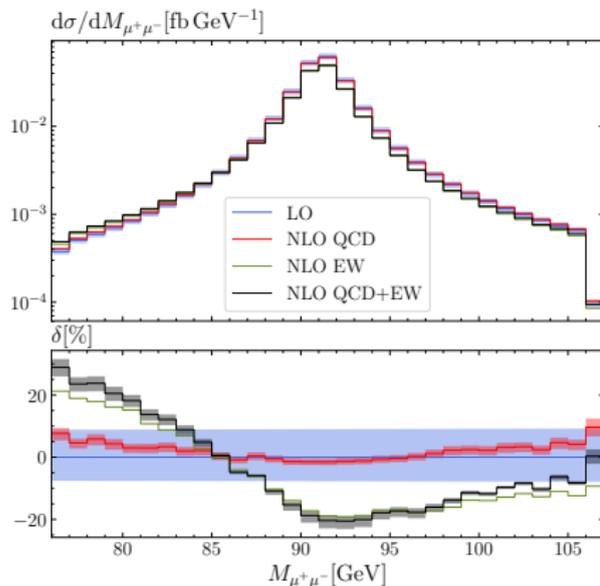
NLO distributions (VI)



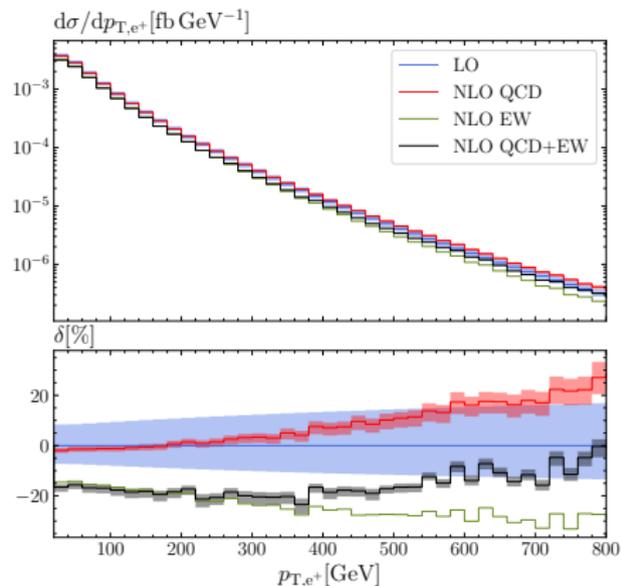
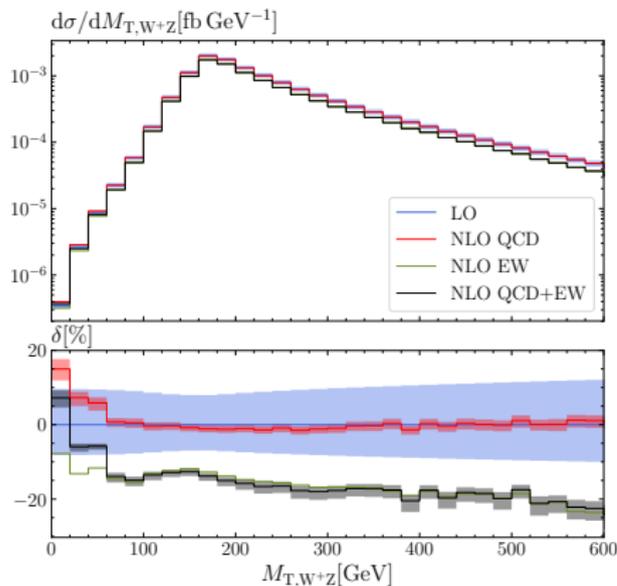
NLO distributions (VII)



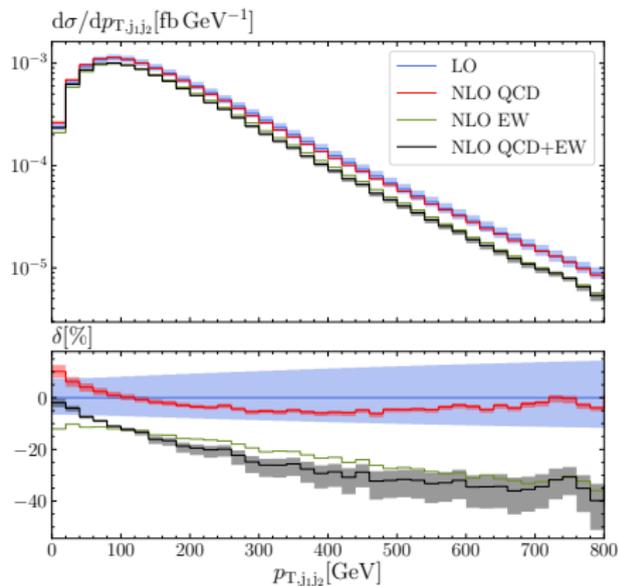
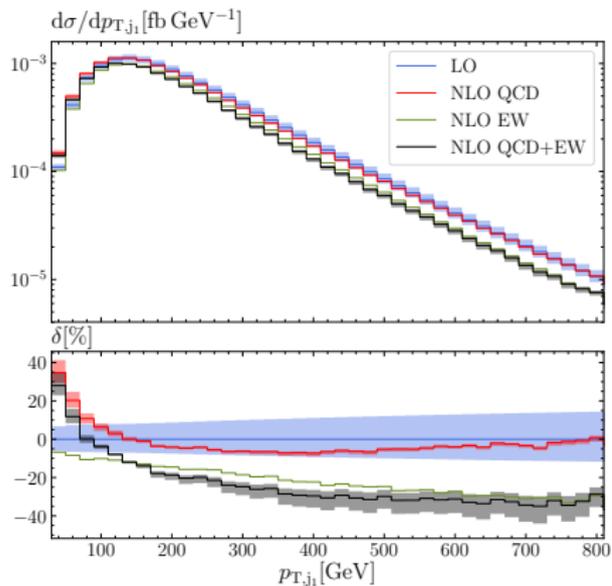
NLO distributions (VIII)



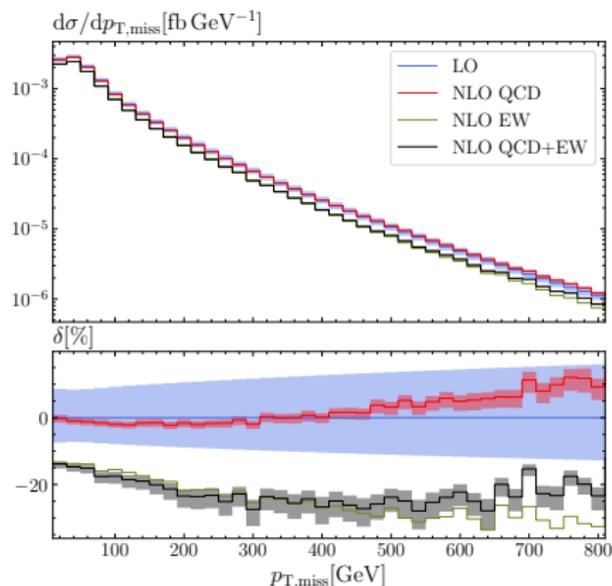
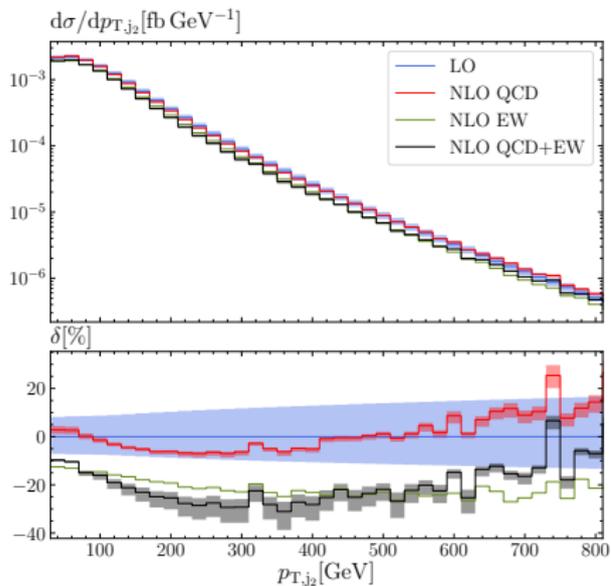
NLO distributions (IX)



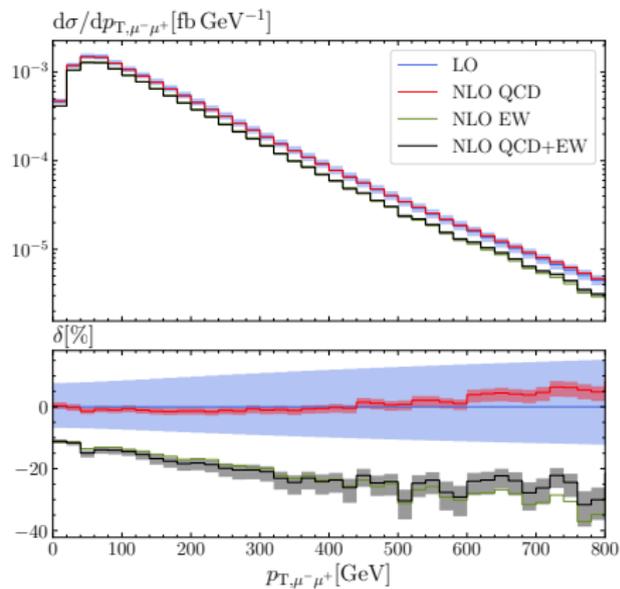
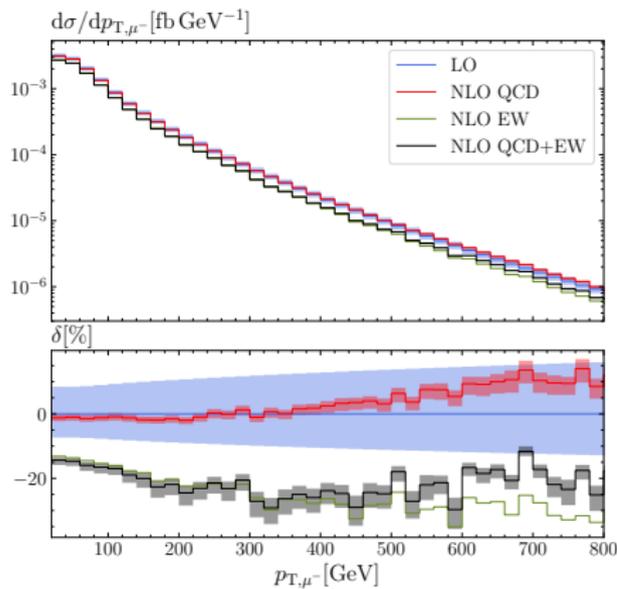
NLO distributions (X)



NLO distributions (XI)



NLO distributions (XII)



NLO distributions (XIII)

