# Searches for Very Long-Lived Particles at the LHC

W

UNIVERSITY of

WASHINGTON

Cristiano Alpigiani

Interplay between Particle and Astroparticle physics Vienna, 09 September 2022

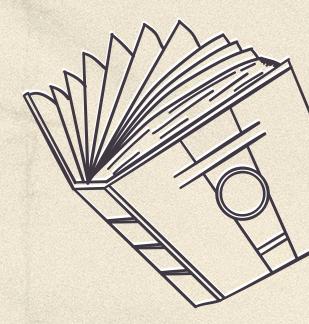
### **Table of Contents**

#### **01** VLLP: Motivations

Why the current LHC detectors may not be enough?

- **02** Transverse Neutral LLP Detectors
- **03** Transverse Charged LLP Detectors
- **04** Forward LLP Detectors
- **05** Some Other New Ideas

**Summary & Conclusions** 





# 01 Motivations

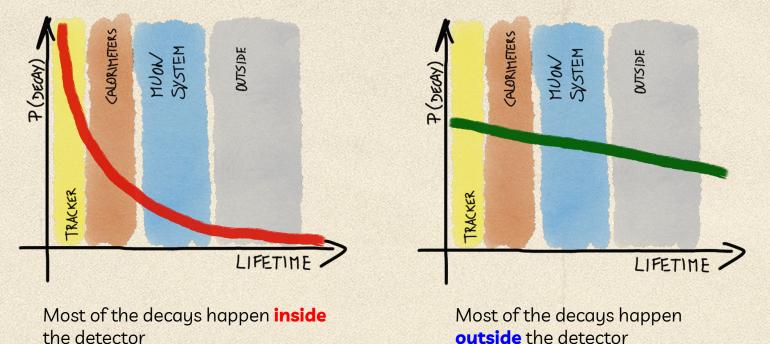
Theory intro in <u>Susanne's talk</u> today LLPs in current LHC experiments in <u>Daniele's talk</u> on Wednesday

### Long-Lived Particles Lifetime

The particle lifetime is a free parameter in the model

- It is sampled from an exponential
- Detector signature strongly depend on boost/mass of the LLP

Need to adapt the search strategy!

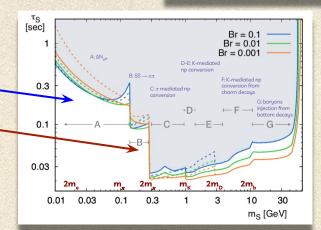


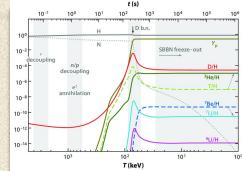
#### arXiv:1706.01920

### LLP Lifetime: Any Limit?

The lifetime of a metastable particle can be limited by cosmology, in particular by the Big Bang Nucleosynthesis (BBN)

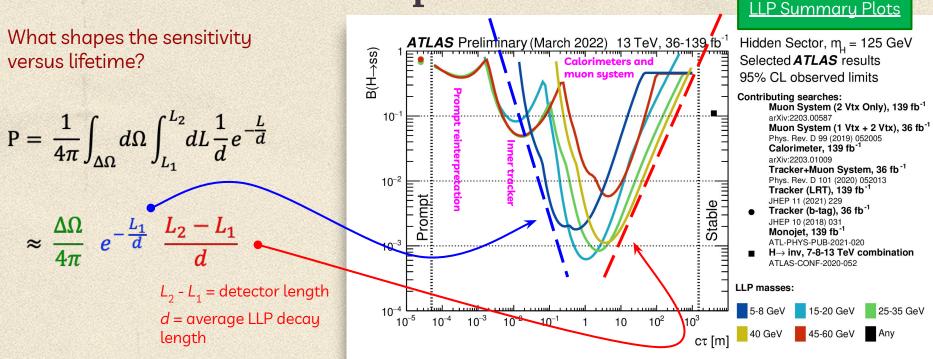
- BBN very well understood within the SM physics and well constrained
  - Happened between ~10 s 15 mins after the Big Bang
- LLP lifetime should be smaller or n/p ratio should have been increased by mesonic and nucleonic LLP decays spoiling the final light nuclei abundances
  - Possible constraint studied on a scalar model coupled through Higgs portal (h→ss)
    - For  $m_s < 2m_u \rightarrow lifetime$  can go up to 1 s
    - For  $2m_{\mu} < m_s < m_h/2 \rightarrow \text{lifetime} < 0.1 \text{ s}$
- Conclusions do not depend strongly on BR( $h \rightarrow ss$ )



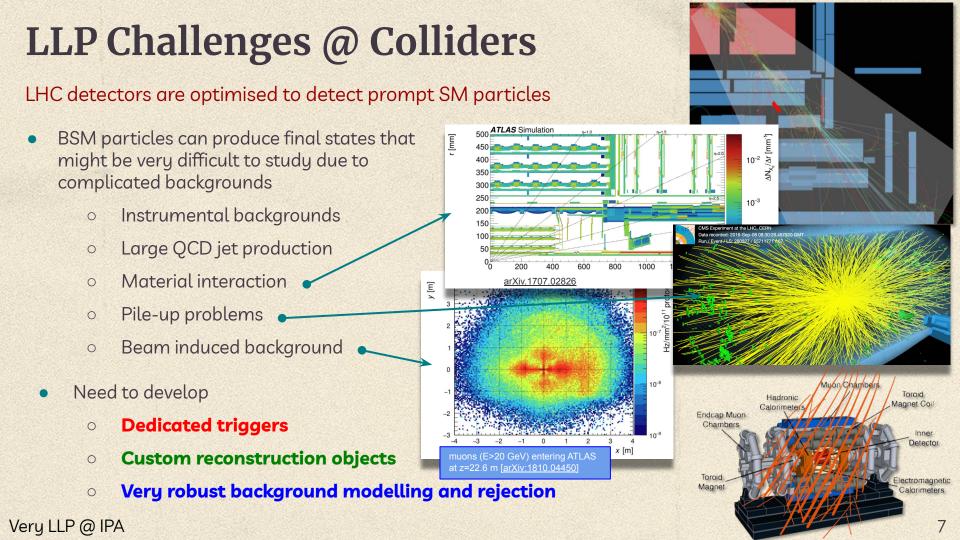


Very LLP @ IPA

### LLP Geometrical Acceptance

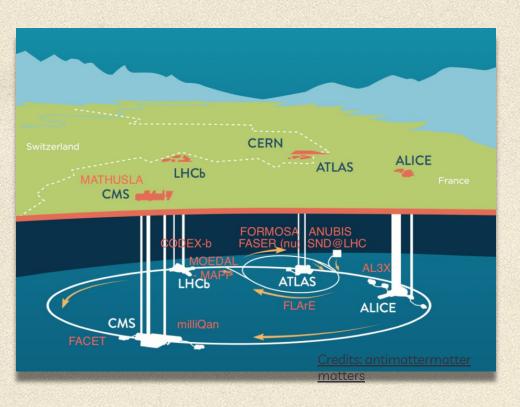


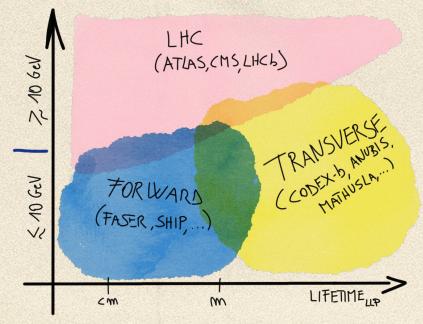
- Good solid angle coverage  $\rightarrow$  lifetime independent
- Sensitivity to smaller lifetimes  $\rightarrow$  need high efficiency close to the IP
- Sensitivity to larger lifetimes  $\rightarrow$  need longer detector



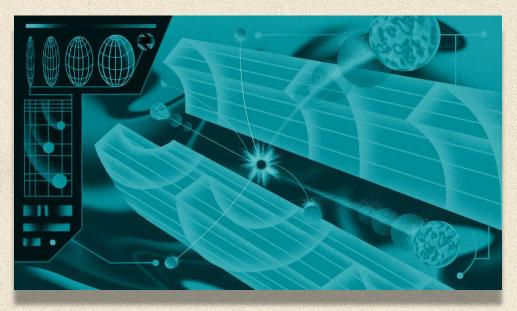
### The Proposed LLP Detectors @ LHC

#### Big variety of complementary detectors





 Allow to cover a big range of lifetimes up to BBN limit, couplings, masses, decay modes





## Transverse <u>Neutral</u> LLP detectors

### MATHUSLA



#### <u>arXiv:2005.02018</u> (Test stand) <u>arXiv:2009.01693</u> (update Lol) <u>arXiv:2203.08126</u> (Snowmass22)

Jura Side

LLP decay volum

Multi-layer tracker

Double layer tracker

Floor detector

Beam line

v(cm

LHC beam pipe

Surface

Dedicated detector sensitive to neutral LLP with lifetime up to BBN

- Proposed a large area surface detector located above CMS
  - Robust tracking + excellent background rejection
  - Floor detectors to reject interactions occurring near the surface
  - **Extruded scintillators + SiPMs** (good time/space resolution)
- **Cosmic**  $\mu$  rate of about **1.7 MHz** and **10 Hz LHC**  $\mu$  rejected with timing
- **LHC neutrinos:** expected 0.1 events from high-E neutrinos (W, Z, top, b), ~1 events from low-E neutrinos ( $\pi/K$ ) over the entire HL-LHC run
- Upward atmospheric neutrinos (70 evts/y above 300 MeV) "decaying" to low momentum proton
  - Advantages: sensitive to very long lifetimes, almost fully shielded from IP background, can do interesting cosmic ray studies

Very LLP @ IPA • Drawback: big detector, need to excavate 20 m (which increases the cost)



Core position

op plane

**RPC** lave

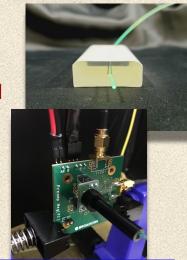
x(cm)

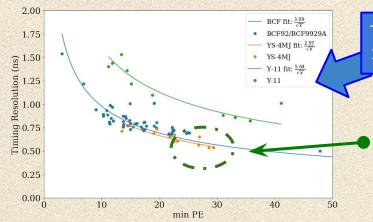
### **MATHUSLA - Status**

Extruded scintillator bars with wavelength shifting fibers coupled to SiPM (tested extrusion facilities - Fermilab and Uniplast)

- Critical features of the detector design
  - Separates downwards from upwards going tracks
  - Reject low beta particles from neutrino QIS

 $\rightarrow$  Need ~1 ns with >15 photoelectrons



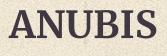


Timing measurement for a 5 m long fiber through a  $1 \times 4 \text{ cm}^2$  extrusion located at the center of the fiber

**Timing resolution of ~0.54 ns** (i.e. 9 cm RMS position resolution) well within MATHUSLA requirement. **Worst case light-yield was 29 PE** 

- Finalising first design of detector layout, DAQ and trigger
- Conceptual Design Report to be published soon

12





**AN U**nderground **B**elayed In-Shaft: i.e. instrumenting the ATLAS shaft with tracking detector for HL-LHC

- 4 stations of RPCs for tracking (2 triplet each)
- Timing to reject cosmic rays
- Can be combined with ATLAS information as veto and background estimation
- Assuming background similar to the LLP searches in ATLAS muon system (need 4-50 events for evidence)

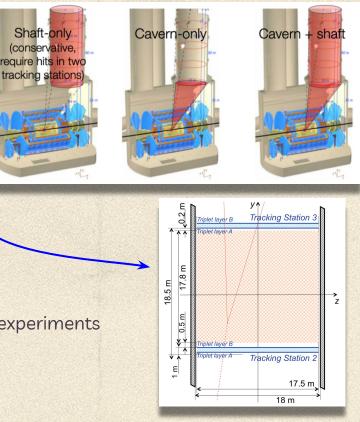
#### Advantages

- Solid angle comparable to MATHUSLA
- $\circ~$  Up to  $10^3$  better sensitivity wrt current/approved future experiments for neutral LLPs with ct  $\gtrsim 10^2$  m
- Moderate costs: O(10) MCHF

#### • Drawbacks

 Basically no shielding: need a very good understanding and modelling of background from the IP Very LLP @ IPA

#### **Conceptually similar to MATHUSLA**



### **ANUBIS - Status**

More details in <u>here</u> and <u>here</u> (XI LLP workshop)

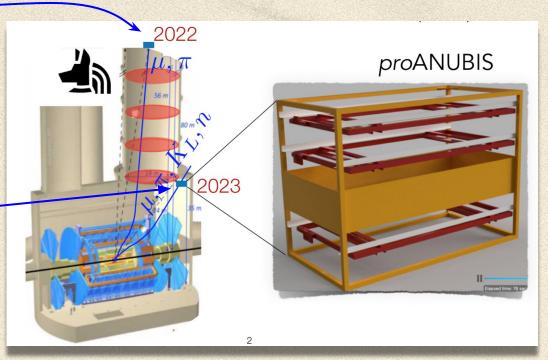
Currently building a 1.8 x 1 x 1 m<sup>3</sup> prototype to measure the flux and correlated to ATLAS during Run 3

#### • 2022 target

- Commissioning, study tracking performance, synchronisation with ATLAS
- Rate of secondaries from hadrons interacting with concrete, probability to see hadrons from punch-through jets

#### 2023 target

- Continue 2022 studies + measure rate of K<sub>L</sub>, n in events with jets pointing towards proANUBIS
- Validation of Geant4 simulation



• Some sensitivity to NP? I.e. charged massive particles with  $\beta \approx 1$  (not small enough to be seen by ATLAS) Very LLP @ IPA 13

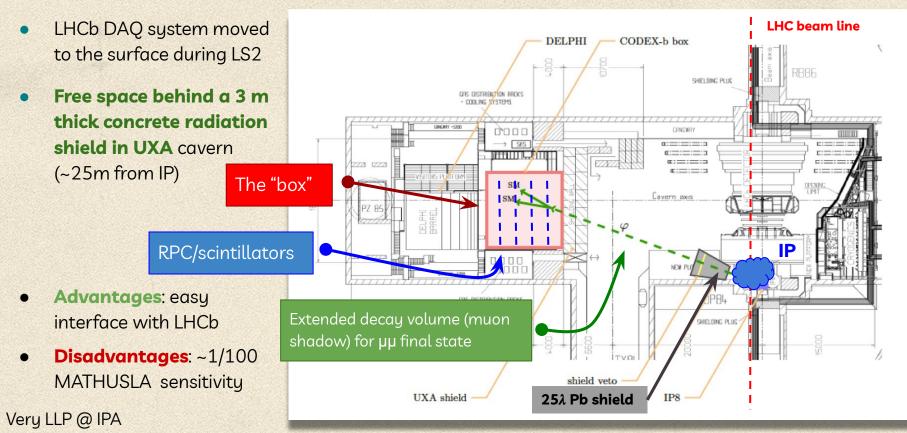
### **CODEX-b**



<u>arXiv:1911.00481</u> (proposal) <u>arXiv:2203.07316</u> (Snowmass22)

**Conceptually similar to MATHUSLA** 

COmpact Detector for EXotics at LHCb (theoretically well-motivated ≤O(10 GeV) LLP)



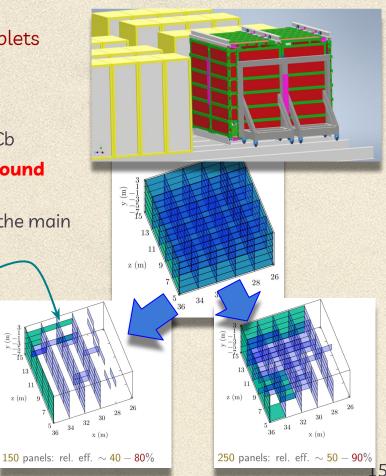
### **CODEX-b - Status**

A small demonstrator detector 2 x 2 x 2  $m^3$  made or RPC triplets will be installed in the old LHCb HLT server room

#### • Main goals

- Test tracking technology and integration with LHCb
- Demonstrate the **ability to reconstruct SM background** inside the detector
- Prove the mechanical structure and its scalability to the main detector
- Optimise the detector shape and tracking layout
  - Considering also possible different geometry/orientation/position
- The tracking layer surface area is the main driver of costs and installation time
- Module production should start by the end of this year

#### More details in this talk (XI LLP Workshop)

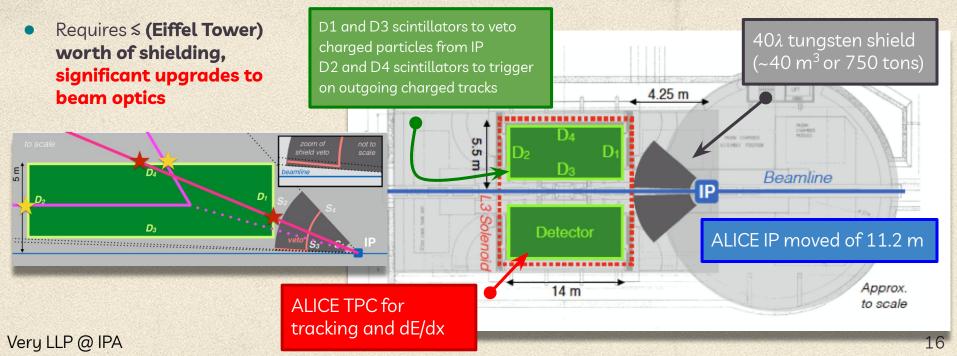


### AL<sub>3</sub>X

#### **Conceptually similar to MATHUSLA**

#### A Laboratory for Long-Lived eXotics

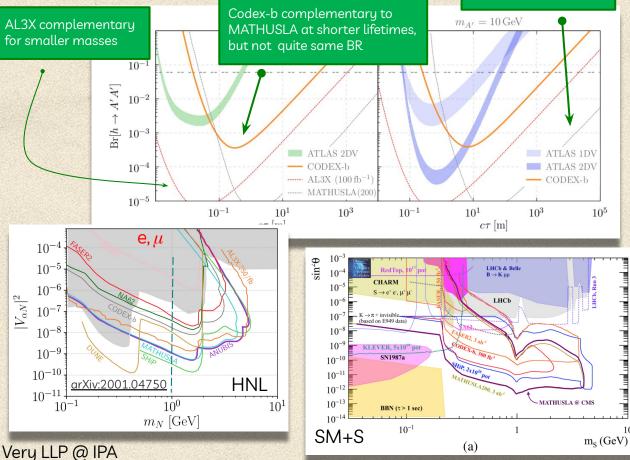
- Reconfigure ALICE detector and its collision point at HL-LHC for dedicated LLP search
- 1/10 MATHUSLA sensitivity at long lifetimes, MUCH BETTER at short lifetimes

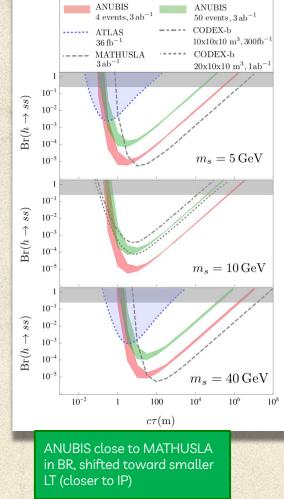


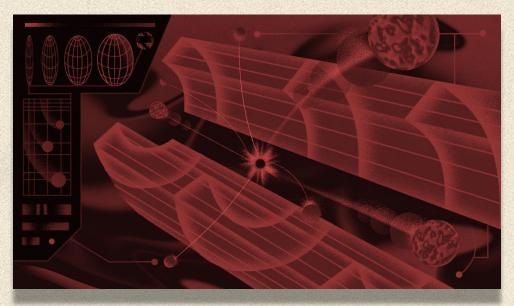
### (Some) Sensitivities

MATHUSLA has good sensitivity for masses >5 GeV and long LT

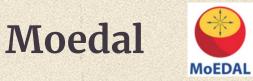
10







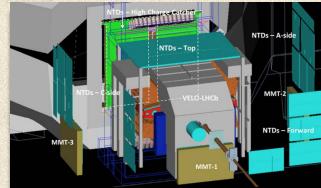
# Transverse Charged LLP detectors



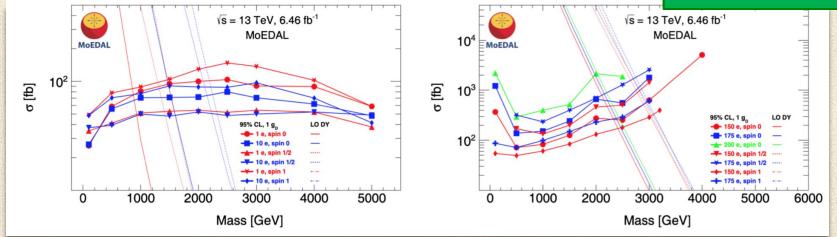
#### Moedal (PhysRevLett.126.071801, PhysRevLett.123.021802)

Experiment located in the LHCb cavern looking for highly ionizing particles, magnetic monopoles, pseudo-stable charged particles

- It uses magnetic monopole traps and nuclear track detectors
- Future plans
  - **MAPP** (in Run 3) to detect millicharged particle (0.001 e)
  - MALL to detect charged very LLP



#### First search for Dyons!



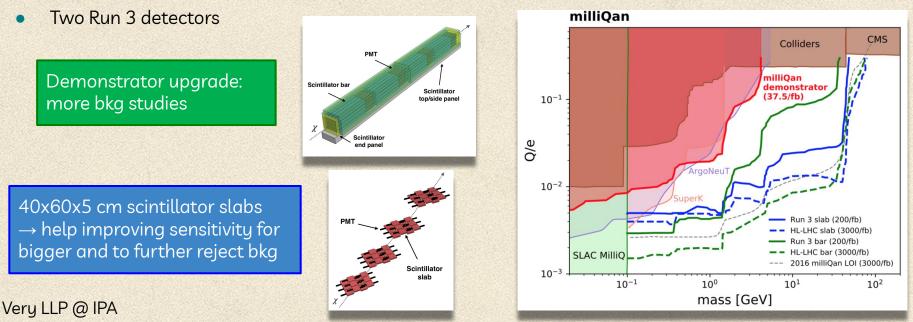
orward region

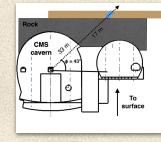
### Milliqan



**Milliqan** (Lol, <u>arXiv:1607.04669</u>, Run 3 updates <u>arXiv:2104.07151</u>) Experiment to detect "<u>milli-charged</u>" particles (0.3-0.001e) with O(GeV) masses produced by pp collisions at CMS

- 17 m rock shielding, 1x1x3 m<sup>3</sup> plastic scintillator array + high-gain PMTs with long axis pointing to CMS IP
- Small prototype (~1% size) took data in 2018 (and confirmed background expectation)





20





 $()/_{L}$ 

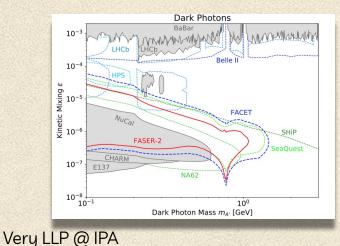
FASER: see Monica's talk

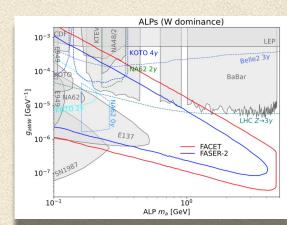
#### arXiv:2201.00019 (proposal)

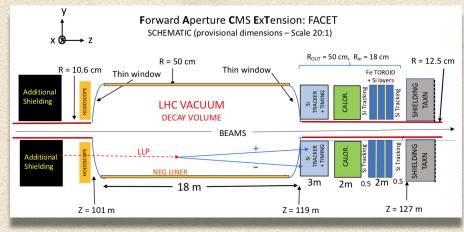
### FACET

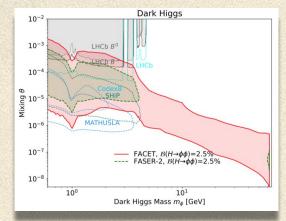
#### Forward- Aperture CMS ExTension

- Replace 18 m long section of the LHC beam pipe on one side of the interaction region with a circular pipe of 50 cm radius
  - **1.** Additional shielding placed upstream of the first detector hodoscope (made of radiation-hard quartz pads)
  - 2. Silicon trackers + timing and calorimeters (as CMS Phase-2 upgrade) to measure the LLP decay products
  - **3.** Iron toroid (1.75 T) with additional silicon trackings

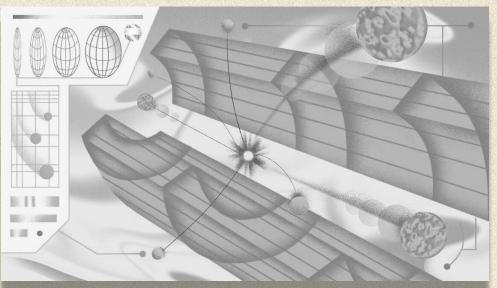








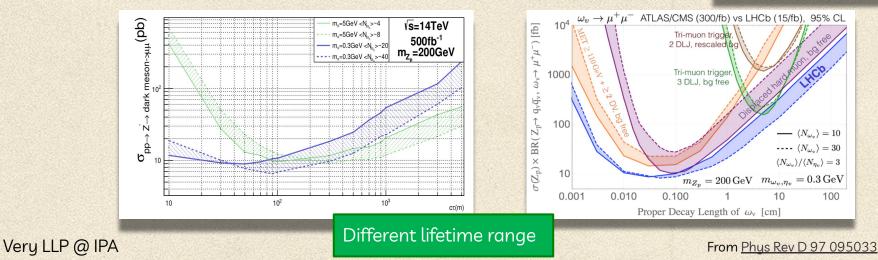
22



# Some Other New Ideas

### LHC Detector Array for LLP

- Use the ATLAS detector (biggest muon system) to detect di- $\mu$  LLP decays from other detectors
- Main backgrounds from
  - Cosmic muons / bkg from collisions (removed using directional information)
  - Radioactive environment (removed: muon pt > 1 GeV)
  - High energy muons from LHCb/ALICE, neutrino conversion (expected to be negligible)





JHEP02 (2022) 069

### LLP in Cosmic Rays

LLP with masses ~GeV can be produced in hadron decays

- Possibly produced in a cosmic ray showers even without TeV-scale mediator
- LLP flux need to be quantified for every model
- Cosmic ray from all directions can contribute
- In principle no detector necessary and dedicated searches can already be sensitive (?)



#### Many theoretical discussions arXiv:1906.09064 arXiv:1910.12839 arXiv:1806.03063

Observation of an Unusual Upward-going Cosmic-ray-like Event in the Third Flight of ANITA

P. W. Gorham,<sup>1</sup> B. Rotter,<sup>1</sup> P. Allison,<sup>2</sup> O. Banerjee,<sup>2</sup> L. Batten,<sup>3</sup> J. J. Beatty,<sup>2</sup> K. Bechtol,<sup>4</sup> K. Belov,<sup>5</sup> D. Z. Besson,<sup>6,7</sup> W. R. Binns,<sup>8</sup> V. Bugaev,<sup>8</sup> P. Cao,<sup>9</sup> C. C. Chen,<sup>10</sup> C. H. Chen,<sup>10</sup> P. Chen,<sup>10</sup> J. M. Clem,<sup>9</sup> A. Connolly,<sup>2</sup> L. Cremonesi,<sup>3</sup> B. Dailey,<sup>2</sup> C. Deaconu,<sup>4</sup> P. F. Dowkont,<sup>8</sup> B. D. Fox,<sup>1</sup> J. W. H. Gordon,<sup>2</sup> C. Hast,<sup>11</sup> B. Hill,<sup>1</sup> K. Hughes,<sup>2</sup> J. J. Huang,<sup>10</sup> R. Hupe,<sup>2</sup> M. H. Israel,<sup>8</sup> A. Javaid,<sup>9</sup> J. Lam,<sup>12</sup> K. M. Liewer,<sup>5</sup> S. Y. Lin,<sup>10</sup> T.C. Liu,<sup>10</sup> A. Ludwig,<sup>4</sup> L. Macchiarulo,<sup>1</sup> S. Matsuno,<sup>1</sup> C. Miki,<sup>1</sup> K. Mulrey,<sup>9</sup> J. Nam,<sup>10</sup> C. J. Naudet,<sup>5</sup> R. J. Nichol,<sup>3</sup> A. Novikov,<sup>6</sup> E. Oberla,<sup>4</sup> M. Olmedo,<sup>1</sup> R. Prechelt,<sup>1</sup> S. Prohira,<sup>6</sup> B. F. Rauch,<sup>8</sup> J. M. Roberts,<sup>1</sup> A. Romero-Wolf,<sup>5</sup> J. W. Russell,<sup>1</sup> D. Saltzberg,<sup>12</sup> D. Seckel,<sup>9</sup> H. Schoorlemmer,<sup>1</sup> J. Shiao,<sup>10</sup> S. Stafford,<sup>2</sup> J. Stockham,<sup>6</sup> M. Stockham,<sup>6</sup> B. Strutt,<sup>12</sup> G. S. Varner,<sup>1</sup> A. G. Vieregg,<sup>4</sup> S. H. Wang,<sup>10</sup> and S. A. Wissel<sup>13</sup>

<sup>1</sup>Dept. of Physics and Astronomy, Univ. of Hawaii, Manoa, HI 96822.
<sup>2</sup>Dept. of Physics, Center for Cosmology and AstroParticle Physics, Ohio State Univ., Columbus, OH 43210.
<sup>3</sup>Dept. of Physics and Astronomy, University College London, London, United Kingdom.
<sup>4</sup>Dept. of Physics and Astronomy, University College London, London, United Kingdom.
<sup>5</sup>Jet Propulsion Laboratory, Pasadena, CA 91109.
<sup>6</sup>Dept. of Physics and Astronomy, University, Rasadena, CA 91109.
<sup>6</sup>Dept. of Physics and Astronomy, University, Taiper, Kansas, Lawrence, KS 66045.
<sup>7</sup>National Research Nuclear Univ., Moscow Engineering Physics Inst., Moscow, Russia.
<sup>8</sup>Dept of Physics, Grad. Inst. of Astrophys., & Leung Center for Cosmology and Particle Astrophysics, National Taivan University, Taipei, Taiwan.
<sup>11</sup>SLAC National Accelerator Laboratory, Mento Park, CA, 94025.
<sup>12</sup>Dept. of Physics and Astronomy, Univ. of Kousis, Substitution, Candon State Univ.

We report on an upward traveling, radio-detected cosmic-ray-like impulsive event with characteristics closely matching an extensive air shower. This event, observed in the third flight of the Antarctic Impulsive Transient Antenna (ANITA), a NASA-sponsored long-duration balloon payload, is consistent with a similar event reported in a prev

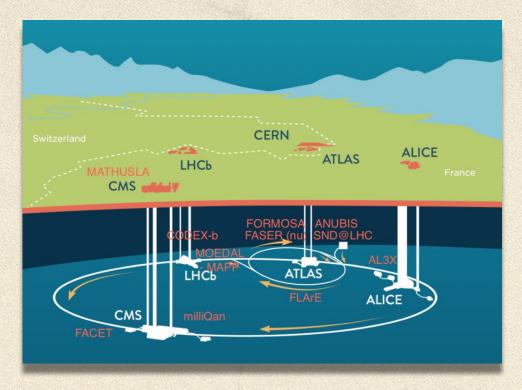
#### Maybe can explain ANITA excess?

Adding a layer of RPC/scintillator on top of ATLAS roof and use to create a decay volume between surface-ATLAS?



### Conclusions

- Impressive number of complementary detectors
- Can make the LHC LLP search program more comprehensive
- Can have the potential to significantly enhance and extend the new physics reach and capabilities of the current LHC detectors
- More ideas may come soon...



# Backup

### **Comparison of Detector Design - Neutral LLP**

	Collision point	Distance from IP	Fiducial volume	Use main experiment?	Shielding cosmics	Shielding collision	Technology
MATHUSLA	CMS	~90 m	25m x 100m x 100m	Under study	NO	YES	Scintillators (+ 1 RPC)
ANUBIS	ATLAS	~25 m	~56m x (9m)²	YES	Partial	NO	RPC (scintillators to be explored)
CODEX-b	LHCb	~35 m	10m x 10m x 10m	Under study	YES	YES	RPC
AL3X	ALICE	~4.25 m	~12m x (2.5m) <sup>2</sup>	NO	YES	YES	Gas TPC

• For a given decay volume,

- More solid angle if closer to the IP
- Number of decays higher if closer to the IP
- LHC collision backgrounds more important if closer to the IP (depending on shielding)

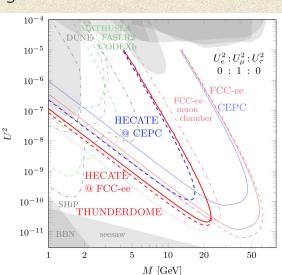
Credits: Emma Torró

### LLP @ FCC-hh and FCC-ee

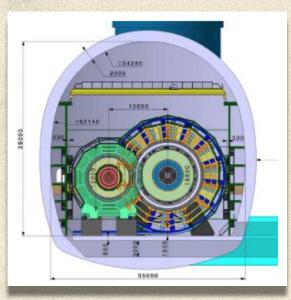
#### HErmetic CAvern TrackER

- FCC-hh / FCC-ee main detectors will be relatively smaller than the cavern
- Cover detector cavern walls with scintillator plates or RPCs
  - Use FCC detector as active veto
  - $\circ$  >= 2 layers of 1 m<sup>2</sup> separated by a sizeable distance
  - $\circ$  >= 4 layers for good tracking

THUNDERDOME = Totally Hyper-UNrealistic DEtectoR in a huge DOME



#### arXiv:2011.01005 (proposal)

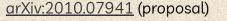


Cavern size: r~15 m and z~50 m Main detector size = 10m

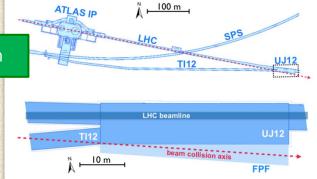
### Formosa

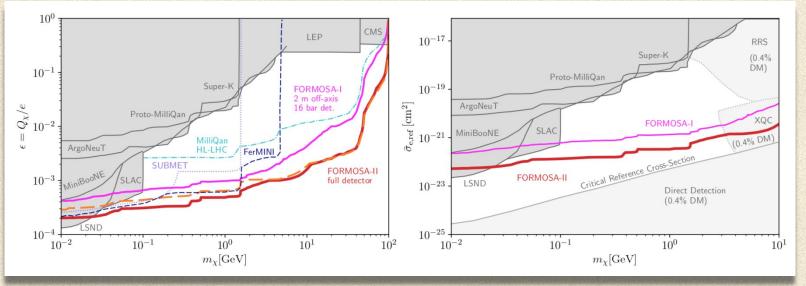
#### FOrward MicrOcharge SeArch

- Looking for millicharge particles in the 10 MeV -100 GeV region in a large and unexplored parameter space and study strongly interacting DM
- Scintillator based experiment (similar to Milliqan)



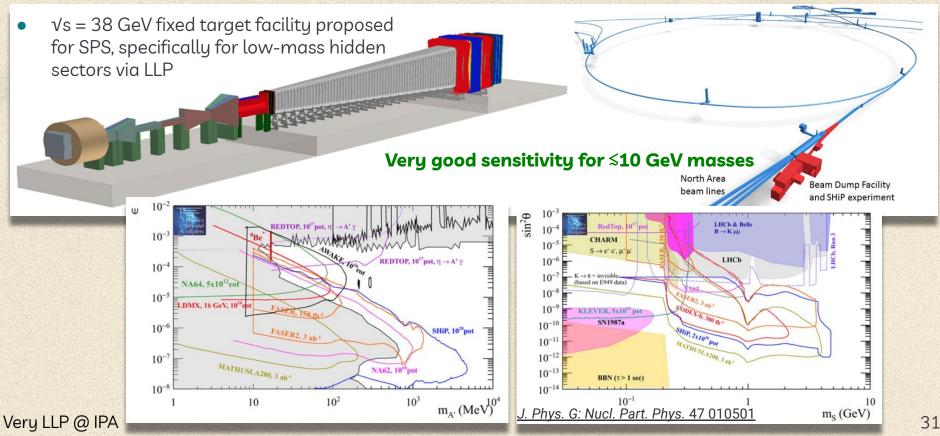
World's most sensitive location





### North Area Experiments - SHiP

#### Search for Hidden Particles

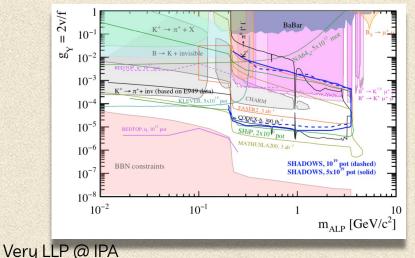


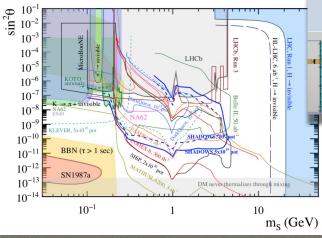
### **North Area Experiments - SHADOW**

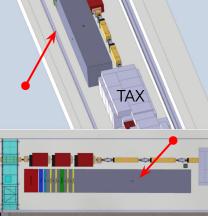
VV (expression of interest) https://cds.cern.ch/record/2799412

#### Search for Hidden Particles

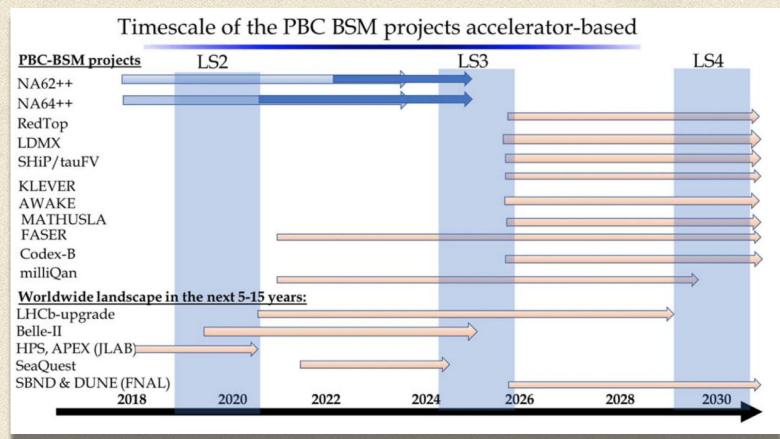
- Proposal for a beam dump experiment to complement NA62 beam dump facility
- "Low cost" detector installed slightly off axis of the TAX shield zone
  - Less affected by  $\mu/\nu$  bkg from beam interaction with dump
  - Series of decay volumes + muon spectrometers
- ~10<sup>9</sup> protons/year on target to study a large variety of FIPs in the mass range MeV-GeV
  - Strongest bounds exist up to K mass; above bounds weaken significantly







### Some Rough Timeline



### **HL-LHC Forward Facility**

#### Proposal to build dedicated forward physics facility for HL-LHC

