

# The Kinetic Mass of Heavy Quarks at Three Loops

14th international workshop on Heavy Quarkonium - QWG 2021

Matteo Fael | March 18, 2021

INSTITUTE FOR THEORETICAL PARTICLE PHYSICS - KIT KARLSRUHE

based on [Fael, Schönwald, Steinhauser, PRL 125 \(2020\) 052003, JHEP 10 \(2020\) 087](#)

## Inclusive semileptonic $B$ -meson decays

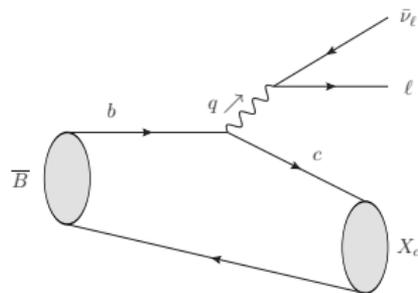
- $|V_{cb}|$

- $m_b$  and  $m_c$

- Use **kinetic scheme**

Gambino, Schwanda, Phys.Rev.D 89 (2014) 1, 014022

Alberti, Gambino, Healey, Nandi, Phys.Rev.Lett. 114 (2015) 6, 061802.



## Why three-loop corrections?

- Precision predictions need short mass schemes for fast convergence of  $\alpha_s$  series.
- Precise conversion between  $\overline{m}_b$  and  $m_b^{\text{kin}}$ .
- Improve the SM prediction by including  $\alpha_s^3$  corrections in pQCD.

# A short distance mass for $B \rightarrow X_c l \bar{\nu}_l$

$$\Gamma_{\text{sl}} = \frac{G_F^2 |V_{cb}|^2 (m_b^{\text{OS}})^5}{192\pi^3} f(0.25) \left[ 1 - 1.78 \left(\frac{\alpha_s}{\pi}\right) - 13.1 \left(\frac{\alpha_s}{\pi}\right)^2 - 163.3 \left(\frac{\alpha_s}{\pi}\right)^3 \right] + O\left(\frac{1}{m_b^2}\right)$$

See: Bigi, Shifman, Uraltsev, Vainshtein PRD 50 (1994) 2234; Beneke, Braun, NPB 426 (1994) 301;  
Ball, Beneke, Braun, PRD 52 (1995) 3929; Melnikov, van Ritbergen, PLB 482 (2000) 99.

- Mass scheme change:  $m_b^{\text{OS}} \rightarrow \tilde{m}_b \left(1 + c \frac{\alpha_s}{\pi}\right)$

$$\Gamma_{\text{sl}} \propto (\tilde{m}_b)^n \left[ 1 + (nc + a_1) \left(\frac{\alpha_s}{\pi}\right) + \left(\frac{n(n+1)}{2} c^2 + nc a_1 + a_2\right) \left(\frac{\alpha_s}{\pi}\right)^2 + \dots \right]$$

- Can we resum the power enhanced  $(n\alpha_s)^k$  terms (with  $n = 5$ )?

**Solution:** use the relation between the  $B$  and bottom mass:

$$\overline{M}_B = m_b + \overline{\Lambda} + \frac{\mu_\pi^2}{2m_b} + \dots$$

- $\overline{\Lambda}$ : the  $B$ -meson binding energy.
- $\mu_\pi$ : the kinetic energy induced by the residual motion of the heavy quark.

The relevant parameter in  $\Gamma_{\text{sl}}$  is  $m_b^5$ , not  $M_B^5$ :

$$\Gamma_{\text{sl}} \simeq \frac{G_F^2 |V_{cb}|^5}{192\pi^3} (M_B - \overline{\Lambda})^5$$

$$m_b^{\text{OS}} = m_b^{\text{kin}}(\mu) + [\bar{\Lambda}(\mu)]_{\text{pert}} + \frac{[\mu_\pi^2(\mu)]_{\text{pert}}}{2m_b^{\text{kin}}(\mu)} - \dots$$

Bigi, Shifman, Uraltsev, Vainshtein, PRD 56 (1997) 4017.  
see also: Czarnecki, Melnikov, Uraltsev, PRL 80 (1998) 3189;  
Gambino, JHEP 09 (2011) 055;  
Fael, Schönwald, Steinhäuser, PRL 125 (2020) 052003.

- In pQCD, we can *peel off* the IR renormalon from the on-shell mass identifying:

$$m_b(\mu) \rightarrow m_b^{\text{kin}}(\mu),$$

$$\bar{M}_B \rightarrow m_b^{\text{OS}},$$

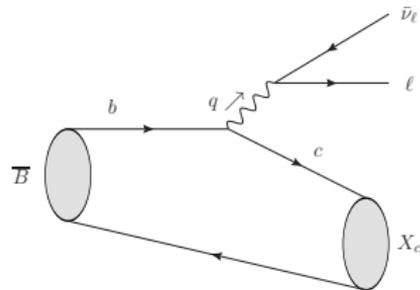
$$\bar{\Lambda}(\mu) \rightarrow [\bar{\Lambda}(\mu)]_{\text{pert}},$$

$$[\mu_\pi^2(\mu)] \rightarrow [\mu_\pi^2(\mu)]_{\text{pert}}.$$

# The Small Velocity sum rules

Moments of the excitation energy  $\omega = E_X - M_D$

$$I_n(\vec{q}^2) = \int d\omega \omega^n \frac{d\Gamma_{\text{tree}}}{d\omega d\vec{q}^2}$$



In the small velocity limit  $|\vec{v}| = |\vec{q}/m_c| \ll 1$ :

$$I_0(\vec{q}^2) = |\vec{q}| \frac{G_F^2 |V_{cb}|^2}{8\pi^3} (m_b - m_c)^2 + \mathcal{O}\left(|\vec{v}|^2, \frac{\Lambda_{\text{QCD}}}{m_b}\right)$$

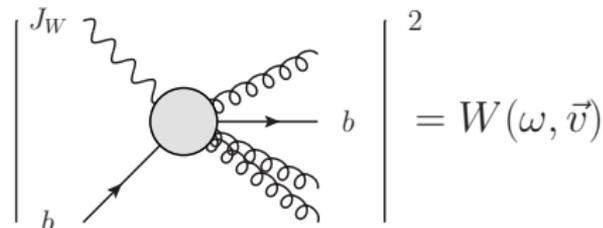
$$I_1(\vec{q}^2) = I_0 \frac{\vec{v}^2}{2} \bar{\Lambda} + \mathcal{O}\left(|\vec{v}|^3, \frac{\Lambda_{\text{QCD}}^2}{m_b^2}\right)$$

$$I_2(\vec{q}^2) = I_0 \frac{\vec{v}^2}{3} \mu_\pi^2 + \mathcal{O}\left(|\vec{v}|^3, \frac{\Lambda_{\text{QCD}}^3}{m_b^3}\right)$$

$$[\bar{\Lambda}(\mu)]_{\text{pert}} = \lim_{\vec{v} \rightarrow 0} \lim_{m_b \rightarrow \infty} \frac{2}{\vec{v}^2} \frac{\int_0^\mu d\omega \omega W(\omega, \vec{v})}{\int_0^\mu d\omega W(\omega, \vec{v})}$$

$$[\mu_\pi^2(\mu)]_{\text{pert}} = \lim_{\vec{v} \rightarrow 0} \lim_{m_b \rightarrow \infty} \frac{3}{\vec{v}^2} \frac{\int_0^\mu d\omega \omega^2 W(\omega, \vec{v})}{\int_0^\mu d\omega W(\omega, \vec{v})}$$

- Consider only soft radiation  $\Lambda_{\text{QCD}} \ll \mu \ll m_b$



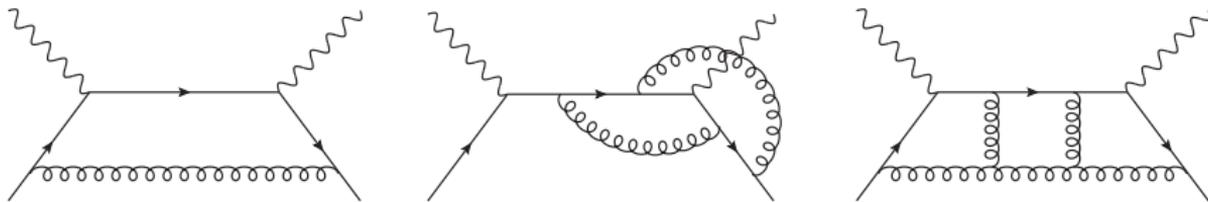
# The Small Velocity sum rules

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- The structure function  $W(\omega, \vec{v})$  up to  $O(\alpha_s^3)$  (we chose scalar and vector current)
- $W(\omega, \vec{v})$  is given by the imaginary part of forward scattering amplitudes like these:



- We need the leading term of  $W(\omega, \vec{v})$  in an expansion in  $\omega$  and  $\vec{v}^2$ .

$$W(\omega, \vec{v}) = W_{\text{el}}(\vec{v}) \delta(\omega) + \frac{\vec{v}^2}{\omega} W_{\text{real}}(\omega) \theta(\omega) + \mathcal{O}\left(v^4, \frac{\omega}{m_b}\right)$$

- We compute it via the *method of regions* in a threshold expansion:  $y = s - m_b^2$ .

Beneke, Smirnov, NPB 522 (1998) 321; Smirnov Springer Tracts Mod. Phys. 250 (2010)

- 4, 66 and 1586 diagrams at one, two and three loops
- Partial fractioning and mapping between families implemented in FORM thanks to code LIMIT.

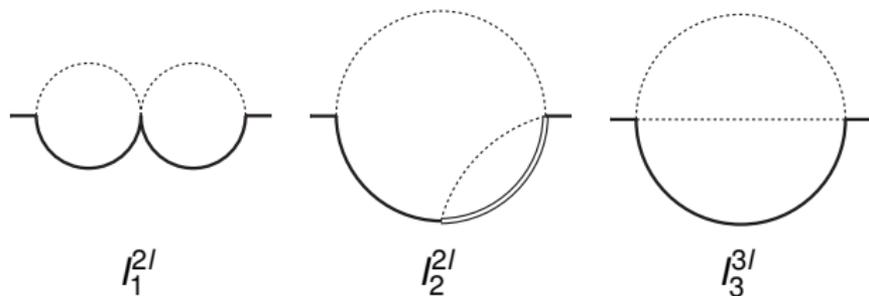
Herren, PhD thesis, KIT, 2020

- FIRE and LiteRed reduction of integral families:

Smirnov, Chuharev, hep-ph/1901.07808; Lee, hep-ph/1212.2685.

- $\{2, 2\}$  in the  $\{(uu), (uh)\}$  regions at two loops
- $\{14, 4, 3\}$  in the  $\{(uuu), (uuh), (uhh)\}$  regions at three loops
- **New master integrals:**
  - 3 in the  $(uu)$  region
  - 20 in the  $(uuu)$  region
- Heavy quark form factors up to  $O(\alpha_s^2)$  (static limit)

Lee, Smirnov, Smirnov, Steinhauser, JHEP 1805 (2018) 187; Blümlein, Marquard, Rana, PRD 99 (2019) 016013



- The new master integrals contains linear-massive propagators:

$$I_2^{2l} = \int d^d k_1 d^d k_2 \frac{1}{k_1^2 (k_1 - k_2)^2 (2k_1 \cdot p - y) (2k_2 \cdot p)}$$

## ■ Mellin-Barnes method

MB package

Czakon, *Comput. Phys. Commun.* 175 (2006) 559;

Smirnov<sup>2</sup>, *EPJC* 62 (2009) 445.

## ■ PSLQ

## ■ Analytic summation of residues

HarmincSums

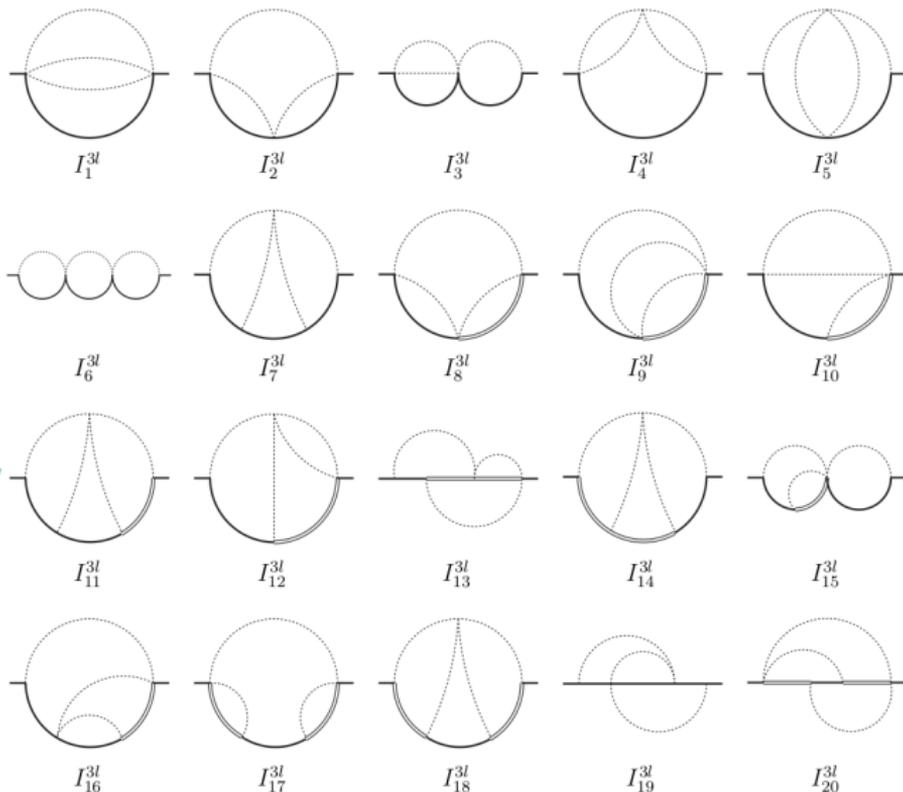
[www3.risc.jku.at/research/combinat/software/HarmonicSums/](http://www3.risc.jku.at/research/combinat/software/HarmonicSums/)

## ■ Differential equations in auxiliary variable

Kotikov, *PLB* 254 (1991), 158

Gehrmann, Remiddi, *NPB* 580 (2000) 485

Henn, *PRL* 110 (2013), 251601.



$$\begin{aligned}
 \frac{m^{\text{kin}}}{m^{\text{OS}}} = & 1 - \frac{\alpha_s^{(n_l)}}{\pi} C_F \left( \frac{4}{3} \frac{\mu}{m^{\text{OS}}} + \frac{1}{2} \frac{\mu^2}{(m^{\text{OS}})^2} \right) + \left( \frac{\alpha_s^{(n_l)}}{\pi} \right)^2 C_F \left\{ \frac{\mu}{m^{\text{OS}}} \left[ C_A \left( -\frac{215}{27} + \frac{2\pi^2}{9} + \frac{22}{9} l_\mu \right) + n_l T_F \left( \frac{64}{27} - \frac{8}{9} l_\mu \right) \right] \right. \\
 & + \left. \frac{\mu^2}{(m^{\text{OS}})^2} \left[ C_A \left( -\frac{91}{36} + \frac{\pi^2}{12} + \frac{11}{12} l_\mu \right) + n_l T_F \left( \frac{13}{18} - \frac{1}{3} l_\mu \right) \right] \right\} + \left( \frac{\alpha_s^{(n_l)}}{\pi} \right)^3 C_F \left\{ \frac{\mu}{m^{\text{OS}}} \left[ C_A^2 \left( -\frac{130867}{1944} \right. \right. \right. \\
 & + \left. \left. \frac{511\pi^2}{162} + \frac{19\zeta_3}{2} - \frac{\pi^4}{18} + \left( \frac{2518}{81} - \frac{22\pi^2}{27} \right) l_\mu - \frac{121}{27} l_\mu^2 \right) + C_A n_l T_F \left( \frac{19453}{486} - \frac{104\pi^2}{81} - 2\zeta_3 \right. \right. \\
 & + \left. \left. \left( -\frac{1654}{81} + \frac{8\pi^2}{27} \right) l_\mu + \frac{88}{27} l_\mu^2 \right) + C_F n_l T_F \left( \frac{11}{4} - \frac{4\zeta_3}{3} - \frac{2}{3} l_\mu \right) + n_l^2 T_F^2 \left( -\frac{1292}{243} + \frac{8\pi^2}{81} + \frac{256}{81} l_\mu - \frac{16}{27} l_\mu^2 \right) \right] \\
 & + \left. \frac{\mu^2}{(m^{\text{OS}})^2} \left[ C_A^2 \left( -\frac{96295}{5184} + \frac{445\pi^2}{432} + \frac{57\zeta_3}{16} - \frac{\pi^4}{48} + \left( \frac{2155}{216} - \frac{11\pi^2}{36} \right) l_\mu - \frac{121}{72} l_\mu^2 \right) + C_A n_l T_F \left( \frac{13699}{1296} - \frac{23\pi^2}{54} \right. \right. \right. \\
 & \left. \left. \left. - \frac{3\zeta_3}{4} + \left( -\frac{695}{108} + \frac{\pi^2}{9} \right) l_\mu + \frac{11}{9} l_\mu^2 \right) + C_F n_l T_F \left( \frac{29}{32} - \frac{\zeta_3}{2} - \frac{1}{4} l_\mu \right) + n_l^2 T_F^2 \left( -\frac{209}{162} + \frac{\pi^2}{27} + \frac{26}{27} l_\mu - \frac{2}{9} l_\mu^2 \right) \right] \right\}, (4)
 \end{aligned}$$

Fael, Schönwald, Steinhauser, PRL 125 (2020) 052003

- The mass relation is written in terms of  $\alpha_s^{(n_l)}$ .
- $n_l$  = number of **massless quarks**,  $l_\mu = \log(2\mu/\mu_s)$ .

# The kinetic- $\overline{\text{MS}}$ mass relation for bottom

- Input values:

$$\alpha_s^{(5)}(M_Z) = 0.1179 \quad \overline{m}_c(3 \text{ GeV}) = 0.993 \text{ GeV} \quad \overline{m}_b(\overline{m}_b) = 4.163 \text{ GeV}$$

- We study different scenarios of  $m_c$  effects (A-D).
- The charm quark wants to be treated as *heavy*.

scheme	$\alpha_s^{(n_f)}$	$m_c$ in $\overline{\text{MS}}\text{-OS}$	$m_c$ in kin-OS	
(B)	4	✓	✓	$m_b^{\text{kin}}(1 \text{ GeV}) = 4163 + 259 + 78 + 26 = 4526 \text{ MeV}$

Fael, Schönwald, Steinhauser, Phys.Rev.D 103 (2021) 1, 014005

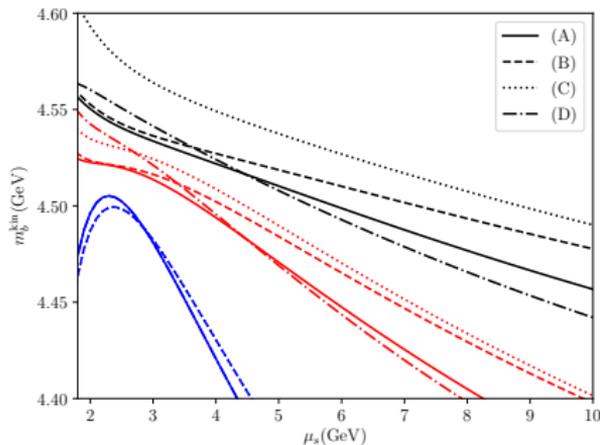
# Scheme conversion uncertainty

- Half of the 3 loop correction

$$\delta m_b^{\text{kin}} \simeq 15 \text{ MeV}$$

- Four-loop large- $\beta_0$  approximation

$$\delta m_b^{\text{kin}} \simeq 8 \text{ MeV}$$



Fael, Schönwald, Steinhauser, Phys.Rev.D 103 (2021) 1, 014005

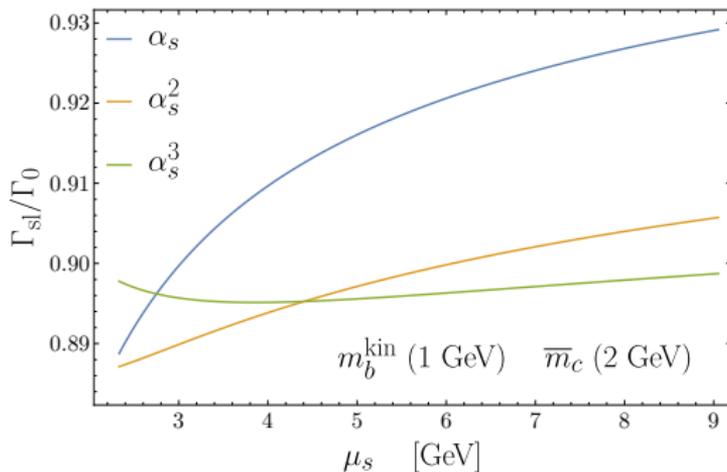
## Compared with:

- scheme conversion uncertainty at two-loops:  $\delta m_b^{\text{kin}} = 30 \text{ MeV}$  [Gambino, JHEP 09 \(2011\) 055](#)
- $m_b^{\text{kin}}(1 \text{ GeV}) = 4554 \pm 18 \text{ MeV}$  from  $B \rightarrow X_c \ell \bar{\nu}_\ell$  global fit [HFLAV 2019](#)

# Improved predictions for $B \rightarrow X_c l \nu$

$$\Gamma_{sl} = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} f(\rho) \left[ 1 + \sum_n Y_n \left( \frac{\alpha_s}{\pi} \right)^n \right]$$

- n=1 [Jezabek, Kühn, Jezabek, Kuhn, NPB 314 \(1989\) 1](#)  
 n=2 [Melnikov, PLB 666 \(2008\) 336; Pak, Czarnecki, PRD 78 \(2008\) 114015.](#)  
 n=3 [Fael, Schönwald, Steinhauser, hep-ph/2011.13654](#)



$$m_b^{\text{OS}} : m_c^{\text{OS}} \quad 1 - 1.78 \left( \frac{\alpha_s}{\pi} \right) - 13.1 \left( \frac{\alpha_s}{\pi} \right)^2 - 163.3 \left( \frac{\alpha_s}{\pi} \right)^3$$

$$\bar{m}_b(\bar{m}_b) : \bar{m}_c(3 \text{ GeV}) \quad 1 + 3.07 \left( \frac{\alpha_s}{\pi} \right) + 13.3 \left( \frac{\alpha_s}{\pi} \right)^2 + 62.7 \left( \frac{\alpha_s}{\pi} \right)^3$$

$$m_b^{\text{kin}}(1 \text{ GeV}) : \bar{m}_c(2 \text{ GeV}) \quad 1 - 1.24 \left( \frac{\alpha_s}{\pi} \right) - 3.65 \left( \frac{\alpha_s}{\pi} \right)^2 - 1.0 \left( \frac{\alpha_s}{\pi} \right)^3$$

- We computed the  $O(\alpha_s^3)$  relation between the kinetic and the on-shell mass.
- We studied finite charm mass effects in  $m_b^{\text{kin}}$ : the charm *wants* to be heavy!
- Scheme conversion uncertainty is reduced by a factor of two compared to two-loop.
- Mass formula is crucial to improve  $B \rightarrow X_c \ell \nu$  prediction in the SM.
- Our results are relevant for future extractions of  $|V_{cb}|$  and  $m_b$  from Belle-II data.

Spare

# Let's include radiative corrections ...

$$I_n(\vec{q}^2) = \int_{|\vec{q}|}^{q_0^{\max}} dq_0 \omega^n \frac{d\Gamma_{\text{tree}}}{dq_0 d\vec{q}^2} + \int_{q_0^{\max} - \mu}^{q_0^{\max}} dq_0 \omega^n \frac{d\Gamma_{\alpha_s}}{dq_0 d\vec{q}^2} + \int_{|\vec{q}|}^{q_0^{\max} - \mu} dq_0 \omega^n \frac{d\Gamma_{\alpha_s}}{dq_0 d\vec{q}^2}$$

- We introduce a **Wilsonian cutoff**  $\mu$ , with  $\Lambda_{\text{QCD}} \ll \mu \ll M_B$ , to separate gluons with
  - $\omega < \mu$  that belong to the non-perturbative regime,
  - $\omega > \mu$  that can be described in pQCD.

# Let's include radiative corrections ...

$$I_1(\vec{q}^2) = \underbrace{\int_{|\vec{q}|}^{q_0^{\max}} dq_0 \omega \frac{d\Gamma_{\text{tree}}}{dq_0 d\vec{q}^2} + \int_{q_0^{\max} - \mu}^{q_0^{\max}} dq_0 \omega \frac{d\Gamma_{\alpha_s}}{dq_0 d\vec{q}^2}}_{\text{use this to define in the SV limit } \bar{\Lambda}(\mu)} + \int_{|\vec{q}|}^{q_0^{\max} - \mu} dq_0 \omega \frac{d\Gamma_{\alpha_s}}{dq_0 d\vec{q}^2}$$

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# $m_b^{\text{kin}}(1 \text{ GeV})$ from $\overline{m}_b(\overline{m}_b)$

scheme	$\alpha_s^{(n_r)}$	$m_c$ in $\overline{\text{MS}}\text{-OS}$	$m_c$ in kin-OS		in MeV
(A)	3	✓	–	$4163 + 248 + (81 + 7_{\Delta_{m_c}} + 12_{dec} - 20_{n_c})$ $+ (30 + 14_{\Delta_{m_c}} + 16_{dec} - 30_{n_c} - 1_{n_c \times dec} + 0.4_{\Delta_{m_c} \times dec})$ $= 4163 + 248 + 80 + 30 = 4520$	
(B)	4	✓	✓	$4163 + 259 + (88 + 7_{\Delta_{m_c}} + 5_{\Delta_{m_c}^{\text{kin}}} - 22_{n_c})$ $+ (34 + 16_{\Delta_{m_c}} + 10_{\Delta_{m_c}^{\text{kin}}} - 34_{n_c})$ $= 4163 + 259 + 78 + 26 = 4526$	

Fael, Schönwald, Steinhauser, hep-ph/2011.XXXX

# $m_b^{\text{kin}}(1 \text{ GeV})$ from $\overline{m}_b(\overline{m}_b)$

$\alpha_s^{(n_f)}$        $m_c$  in  
MS-OS       $m_c$  in  
kin-OS

in MeV

(C)

4

✓

✗

$$\begin{aligned} & 4163 + 259 + (99 + 7\Delta_{m_c} - 22n_c) \\ & \quad + (59 + 16\Delta_{m_c} - 34n_c) \\ & = 4163 + 259 + 84 + 41 = 4547 \end{aligned}$$

(D)

3

✗

✗

$$4163 + 248 + 81 + 30 = 4521$$

Fael, Schönwald, Steinhauser, hep-ph/2011.XXXX

- $m_c^{\text{kin}}(0.5 \text{ GeV})$ :

$$m_c^{\text{kin}}(0.5 \text{ GeV}) = 993 + 191 + 100 + 52 \text{ MeV} = 1336 \text{ MeV} ,$$

$$m_c^{\text{kin}}(0.5 \text{ GeV}) = 1099 + 163 + 76 + 34 \text{ MeV} = 1372 \text{ MeV} ,$$

$$m_c^{\text{kin}}(0.5 \text{ GeV}) = 1279 + 84 + 30 + 11 \text{ MeV} = 1404 \text{ MeV} .$$

- $m_c^{\text{kin}}(1 \text{ GeV})$ :

$$m_c^{\text{kin}}(1 \text{ GeV}) = 993 + 83 + 35 + 14 \text{ MeV} = 1125 \text{ MeV} ,$$

$$m_c^{\text{kin}}(1 \text{ GeV}) = 1099 + 37 + 2 - 3 \text{ MeV} = 1135 \text{ MeV} ,$$

$$m_c^{\text{kin}}(1 \text{ GeV}) = 1279 - 73 - 61 - 17 \text{ MeV} = 1128 \text{ MeV} ,$$

where from top to bottom  $\mu_s = 3 \text{ GeV}, 2 \text{ GeV}$  and  $\bar{m}_c$  for  $\bar{m}_c(\mu_s)$  and  $\alpha_s^{(3)}(\mu_s)$ .

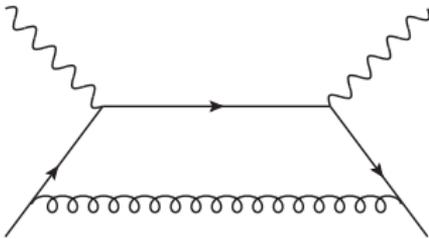
# Expansion by regions

- For one heavy particle threshold, there are two regions:

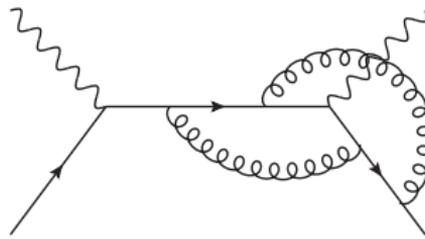
see also: [Smirnov Springer Tracts Mod. Phys. 250 \(2010\)](#)

hard (h):  $k_i \sim m_b$

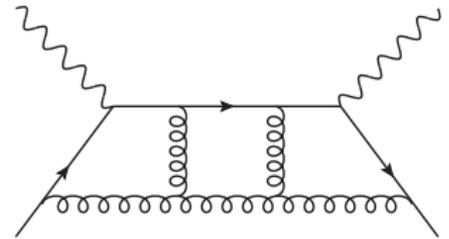
ultra-soft (us):  $k_i \sim y/m_b$



(u) (h)



(uu) (uh) (hh)



(uuu) (uuh) (uhh) (hhh)

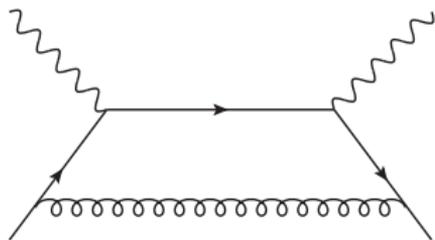
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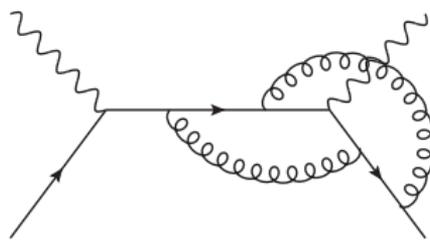
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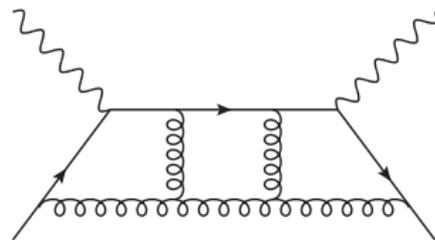
ultra-soft (us):  $k_i \sim y/m_b$



(u) (~~h~~)



(uu) (uh) (~~hh~~)



(uuu) (uuh) (uhh) (~~hhh~~)

- All hard** regions don't contribute: no imaginary part.
- After renormalization and decoupling  $\alpha_s^{(n_l+n_h)} \rightarrow \alpha_s^{(n_l)}$ ,  
**only all ultra-soft part remains**

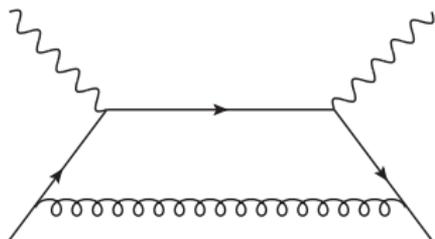
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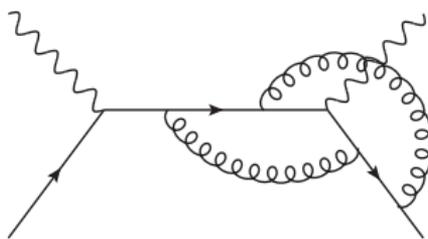
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hard (h):  $k_i \sim m_b$

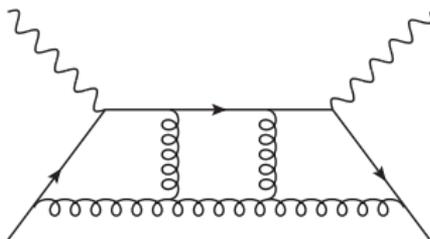
ultra-soft (us):  $k_i \sim y/m_b$



(u) (~~h~~)

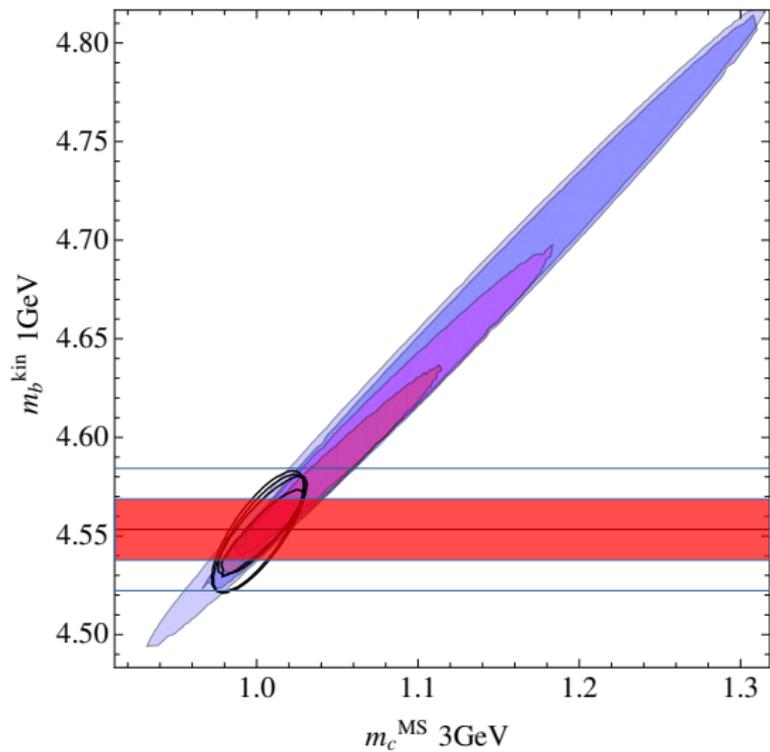


(uu) (~~uh~~) (~~hh~~)



(uuu) (~~uuh~~) (~~uhh~~) (~~hhh~~)

- All hard** regions don't contribute: no imaginary part.
- After renormalization and decoupling  $\alpha_s^{(n_l+n_h)} \rightarrow \alpha_s^{(n_l)}$ ,  
**only all ultra-soft part remains**



Original plot from  
[Gambino, Schwanda, PRD 89 \(2014\) 014022](#)  
Horizontal error bands superimposed by MF