

ICTP FÍSICA SIN FRONTERAS 2019  
JORNADA DE FÍSICA DE ALTAS ENERGÍAS EN ECUADOR

# KM3NeT

(A MULTIDISCIPLINARY OBSERVATORY IN THE ABYSS OF THE MEDITERRANEAN SEA)

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KM3NeT-Latinamerica Principal Investigator / LASF4RI National Representative / IAEACRP Chief Scientific Investigator



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AND NANOTECHNOLOGY

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29<sup>th</sup> August 2019

## PART 1

INTRODUCTION TO KM3NeT SCIENCE

## PART 2

THE KM3NeT LAYOUT

## PART 3

THE KM3NeT NEUTRINO DETECTOR

## PART 1

# INTRODUCTION TO KM3NeT SCIENCE

*Cities and Sites of KM3NeT*

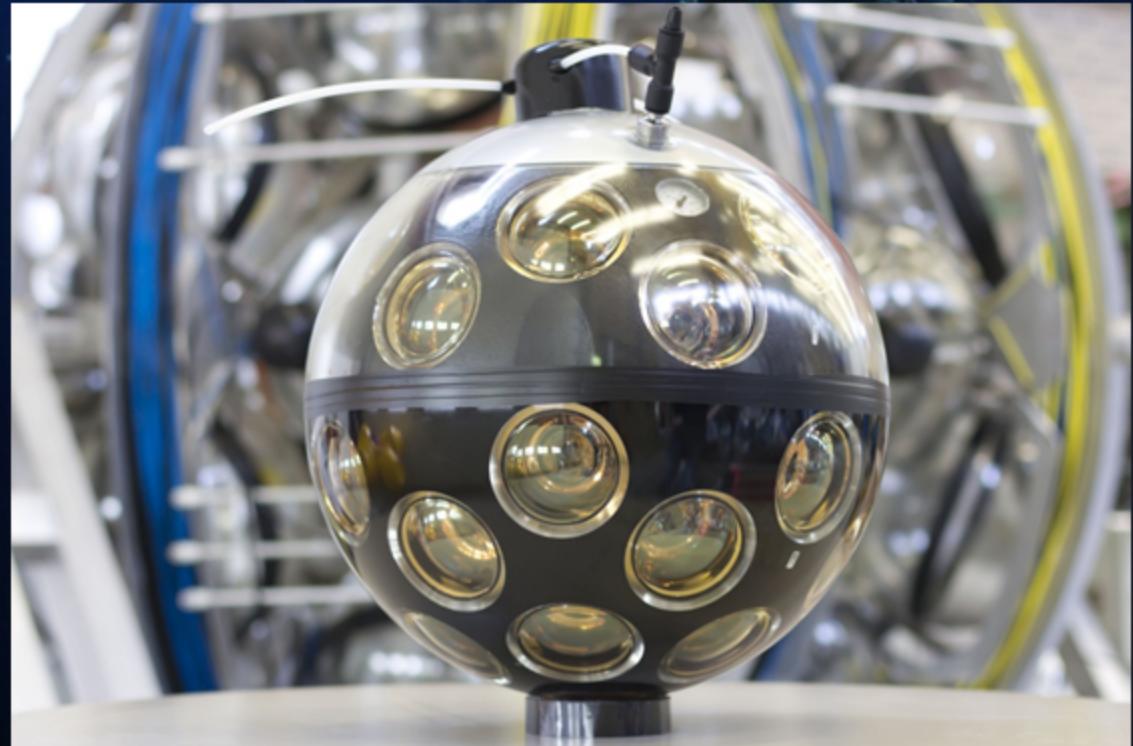
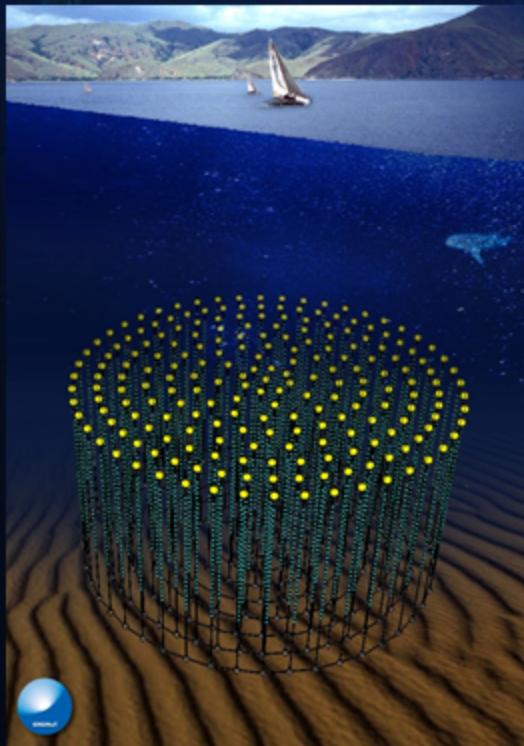




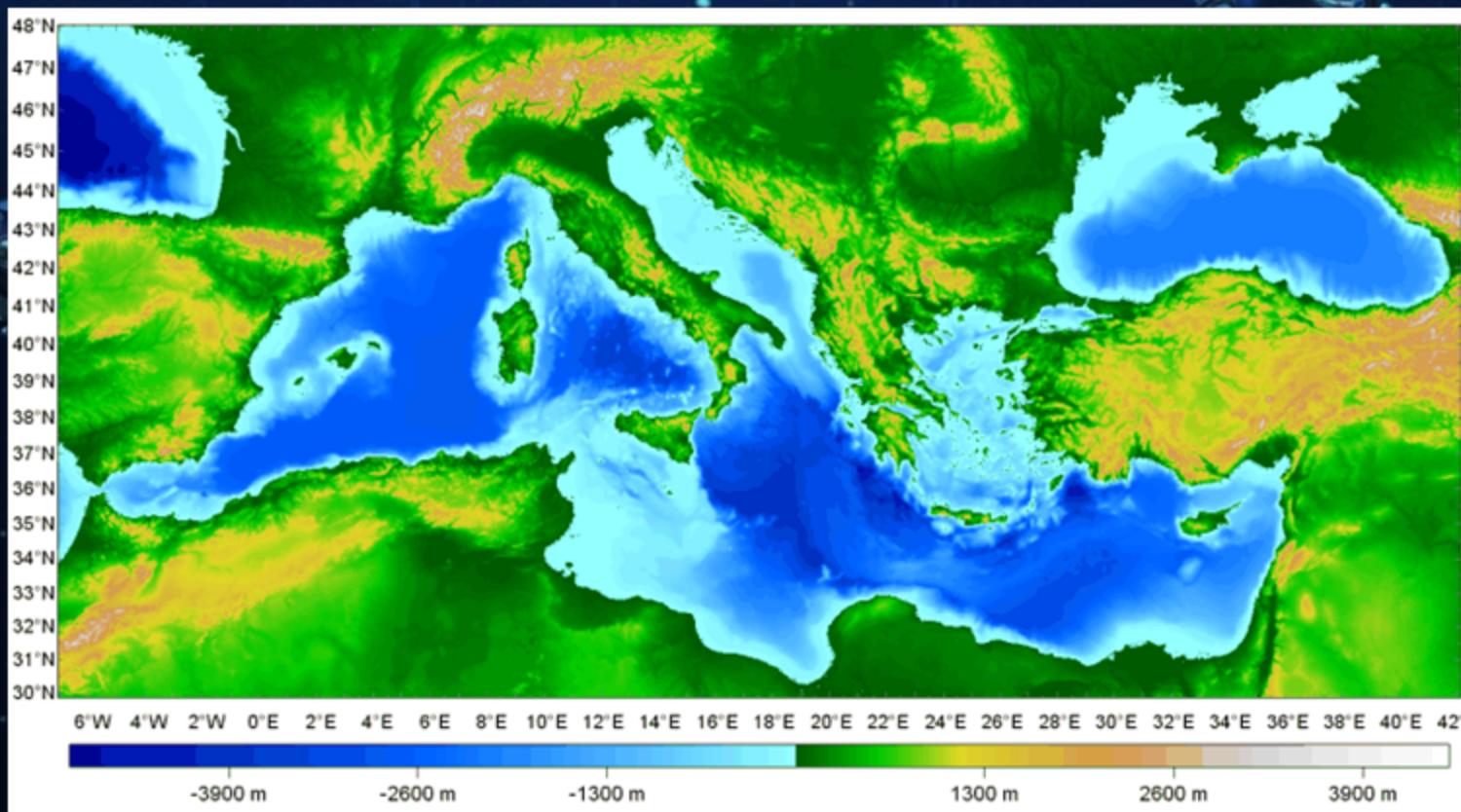
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To enhance our knowledge spanning Sea, Earth and Space sciences, looking for new discoveries and technology advance through the construction and operative experience of large-scale deep-sea multidisciplinary infrastructures and advanced detector systems.



- ✓ To guarantee the Mediterranean deep sea as of prime interest for investigations in the fields of marine and earth sciences featured by the permanent connection to the shore.



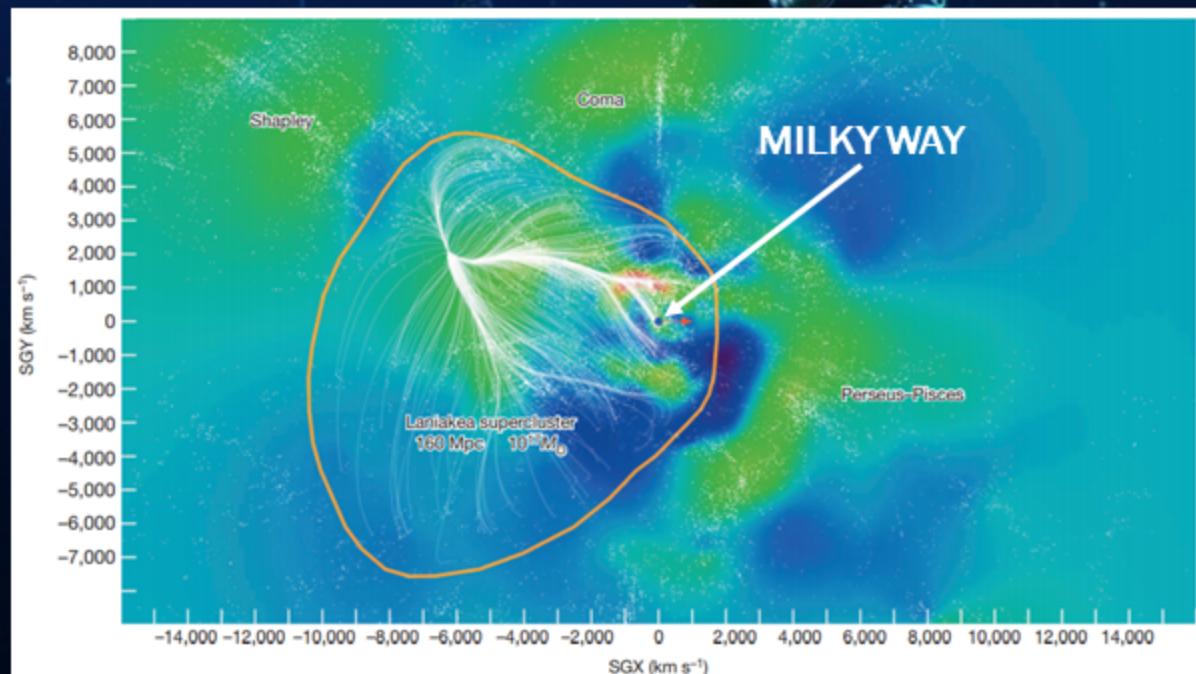
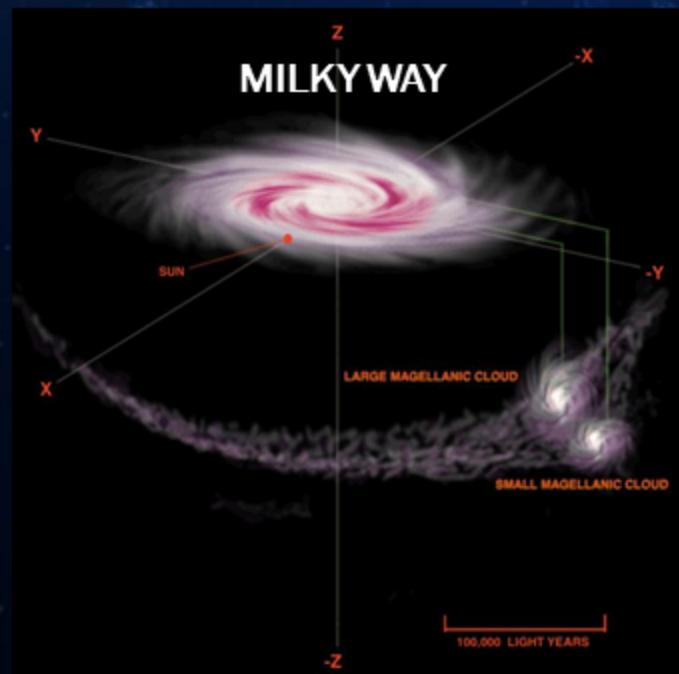
Such objective is strongly motivated by:

- ✓ Long-term monitoring, since permanent connections (powering and reading out instruments) from the deep sea to the shore are only available via cabled sites, at present in Europe, those of the neutrino telescope pilot projects (e.g., ANTARES).
- ✓ Complementary goals and collaborative activities with an infrastructure hosting multi-cubic-kilometer neutrino telescopes in several regions of the Mediterranean.
- ✓ Associated-sciences, infrastructure as a world leading observatory with enormous discovery potential, important node in a global network of deep-sea observatories.



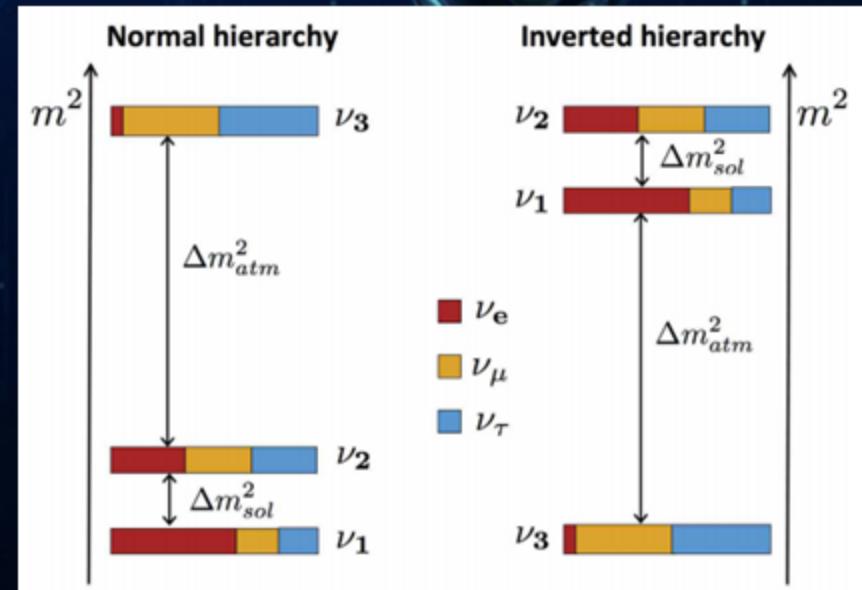
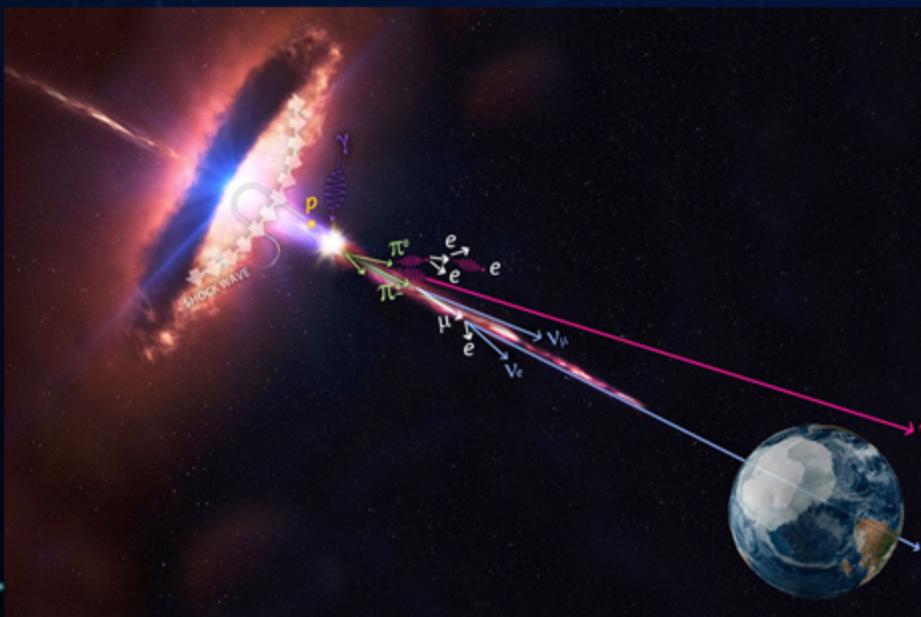
+ DONET Japan

- ✓ Discovery and subsequent observation of **high-energy neutrino sources** in the Universe.
- ✓ Determination of the **mass hierarchy of neutrinos**.



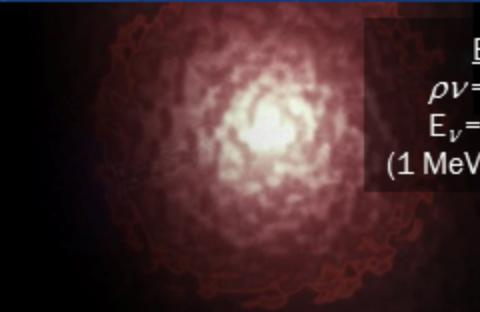
These objectives are strongly motivated by two “recent” important discoveries:

- ✓ The IceCube discovery of high-energy astrophysical neutrino source (**TXS0506+056**).
- ✓ The sizeable contribution of electron neutrinos to the **third neutrino mass eigenstate** as reported by Daya Bay, Reno and others.



AGNs/GRBs/others?

$$E^{-2} \Phi_{\nu}^{Earth} \sim 10^{-8} \nu \text{ GeV/ cm}^2 \text{s sr}$$
$$E_{\nu} \sim 100 \text{ TeV - PeV}$$



Big Bang

$$\rho_{\nu} = 330 / \text{cm}^3$$

$$E_{\nu} = 0.0004 \text{ eV}$$

$$(1 \text{ MeV} = 1.6 \times 10^{-13} \text{ J})$$



Blazar TXS 0506+056

$$E_{\nu} \sim 100 \text{ s TeV}$$

Sun

$$\Phi_{\nu}^{Earth} = 6 \times 10^{10} \nu / \text{cm}^2 \text{s}$$
$$E_{\nu} \sim 0.1 - 20 \text{ MeV}$$



Atmospheric neutrinos

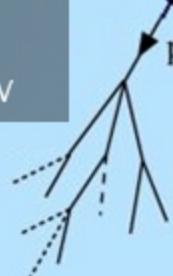
$$\nu_e, \nu_{\mu}, \bar{\nu}_e, \bar{\nu}_{\mu}$$

$$\Phi_{\nu} \sim 1 \nu / \text{cm}^2 \text{s}$$

$$E_{\nu} \sim 0.1 - 100 \text{ GeV}$$

SN1987

$$E_{\nu} \sim \text{MeV}$$



Human body

$$\Phi_{\nu} = 340 \times 10^6 \nu / \text{day}$$



Nuclear reactors

$$E_{\nu} \sim \text{few MeV}$$

Terrestrial radioactivity

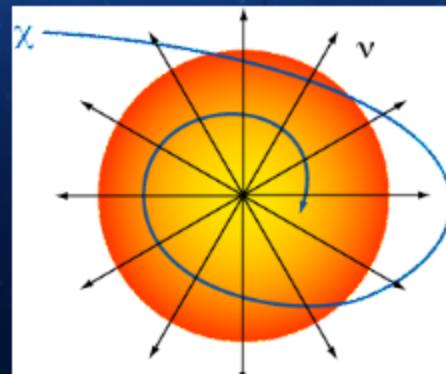
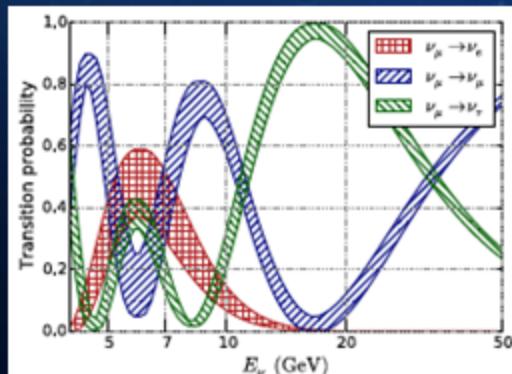
$$\Phi_{\nu} \sim 6 \times 10^6 \nu / \text{cm}^2 \text{s}$$



Accelerators

$$E_{\nu} \sim 0.3 - 30 \text{ GeV}$$





**Low Energy**  
( $\text{MeV} < E_\nu < 100 \text{ GeV}$ )

$\nu$  Oscillations Supernovae

**Medium Energy**  
( $10 \text{ GeV} < E_\nu < 1 \text{ TeV}$ )

Dark Matter, Monopoles,  
Nuclearites

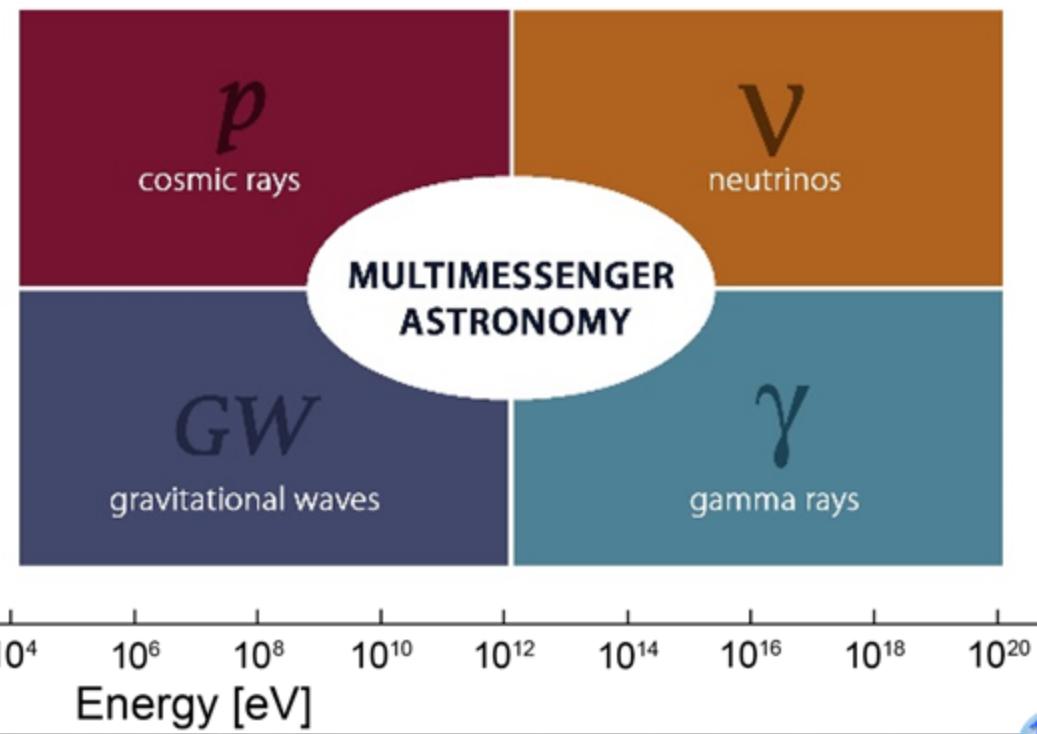
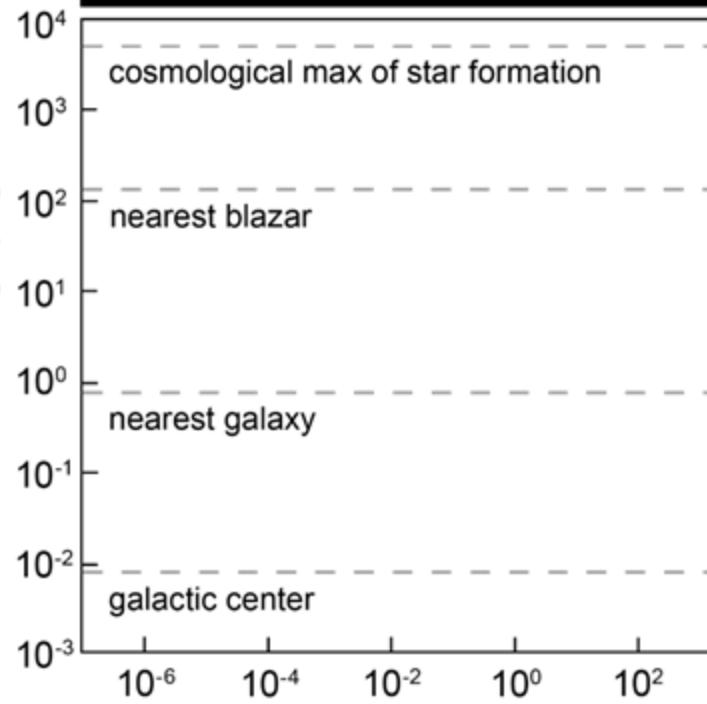
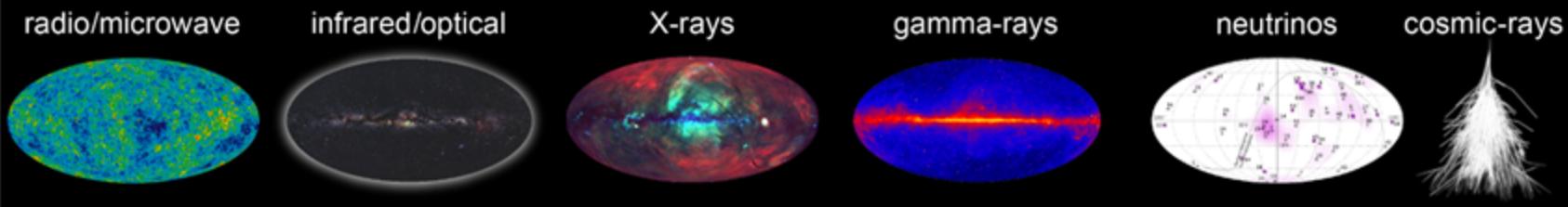
**High Energy**  
( $E_\nu > 1 \text{ TeV}$ )

Cosmic  $\nu$ , Origin and production  
mechanism of HE CR

KM3NeT-ORCA

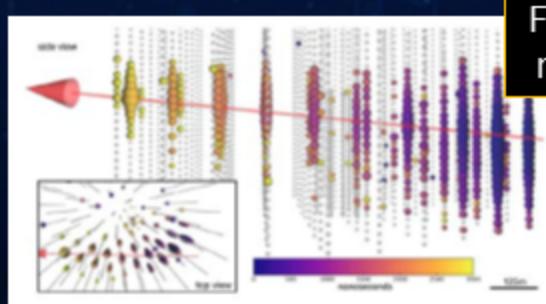
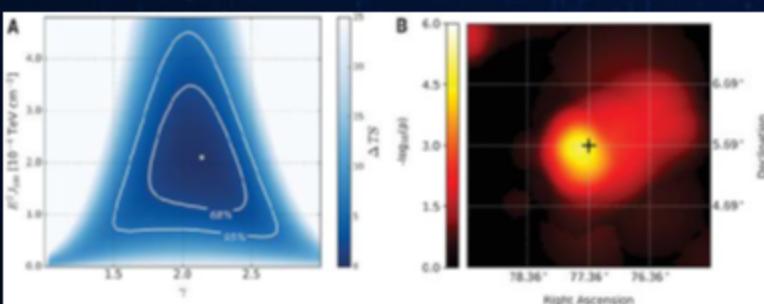
KM3NeT-ARCA

+ **Earth and Sea Science:** environmental sciences, geology and geophysics, marine biology and oceanography ....

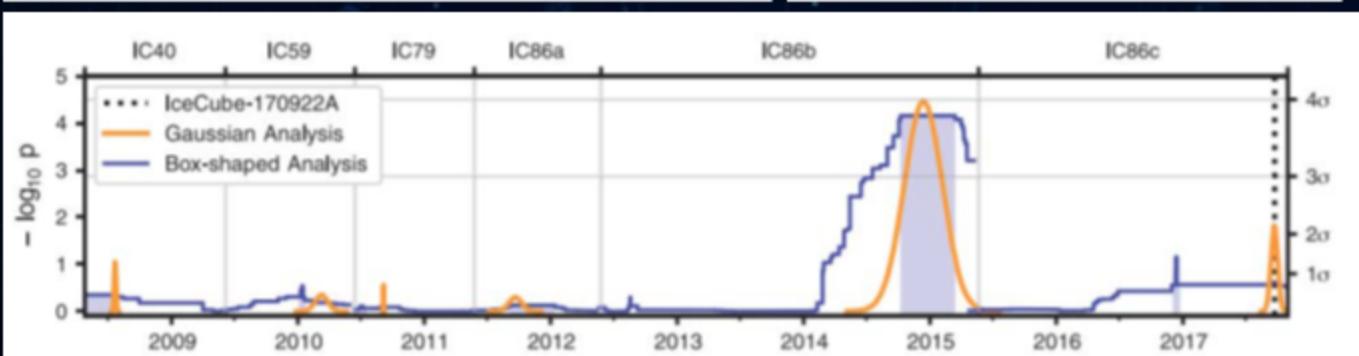


September 2017: A EHE alert is sent by IceCube (Ev~290 TeV). Two analyses:

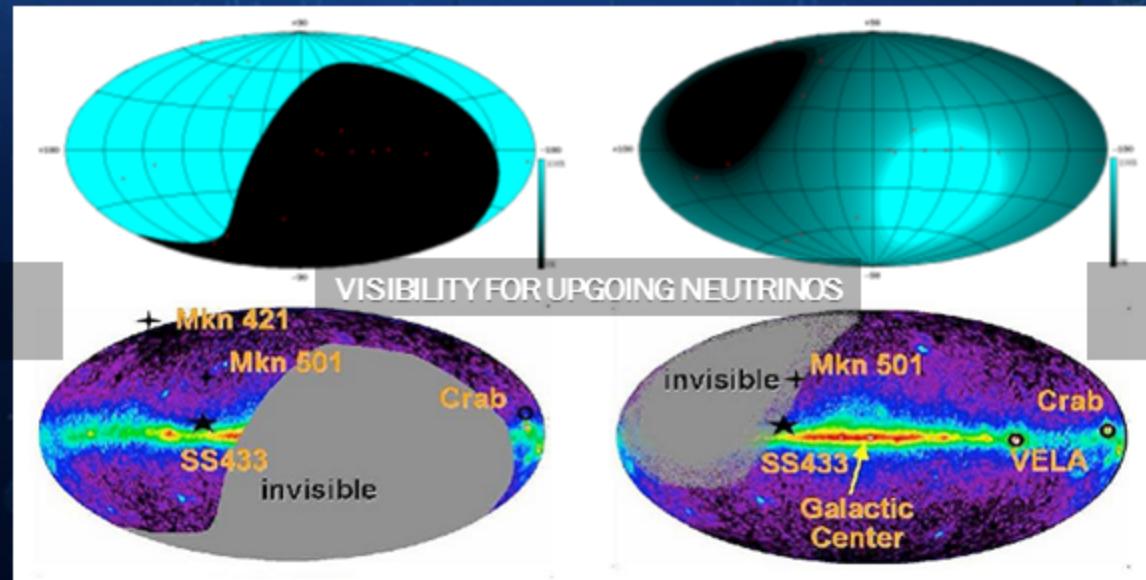
- ✓ IC170922A alert, which triggered positive observation by Fermi, MAGIC and others.
- ✓ Search for additional “neutrino flares” in previous IceCube data: signal of  $17 \pm 5$  over Background.



First identification of a (HE) cosmic neutrino source! (TXS 0506+056)



South Pole (SH)  
(AMANDA 1 TeV)



Mediterranean Sea (NH)  
(ANTARES 1 TeV)

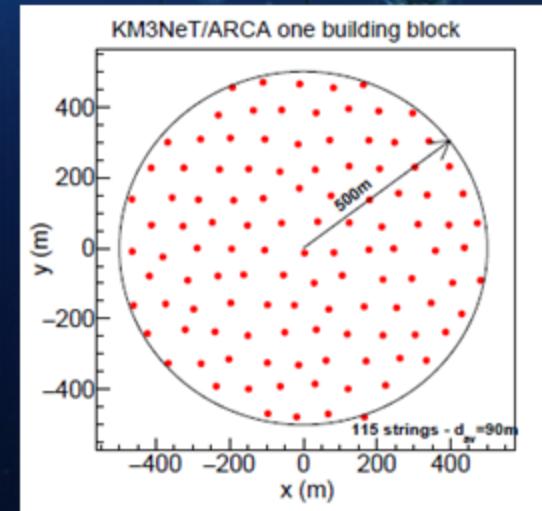
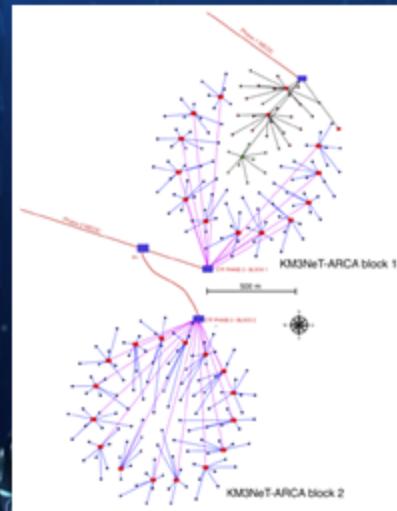
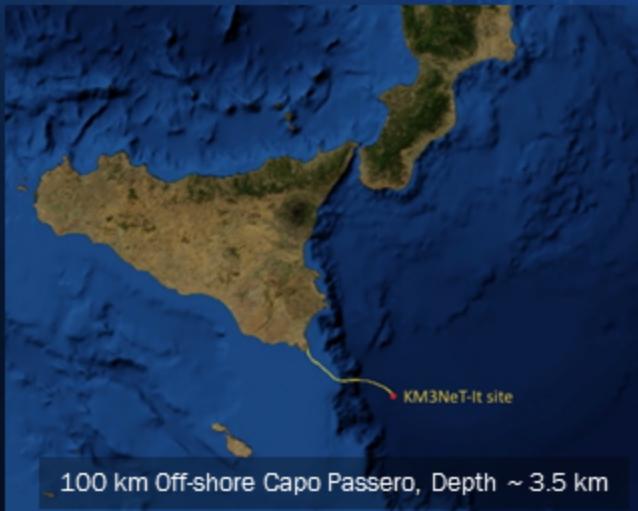
High-energy neutrinos (with IceCube data):

- ✓ Quasi-isotropic dominant component at high energy ( $E > 100\text{TeV}$ ) suggests extragalactic origin.
- ✓ Galactic origin component at lower energies.

→ A better angular resolution is reached in deep-sea water than in ice ←

# PART 2

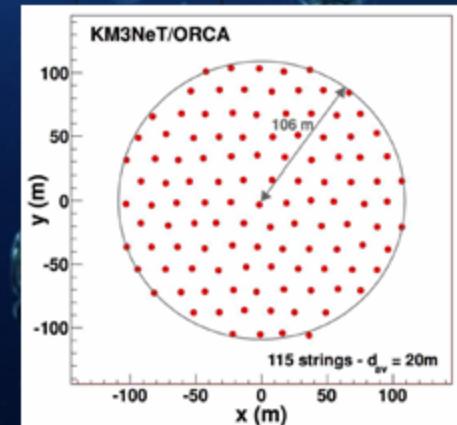
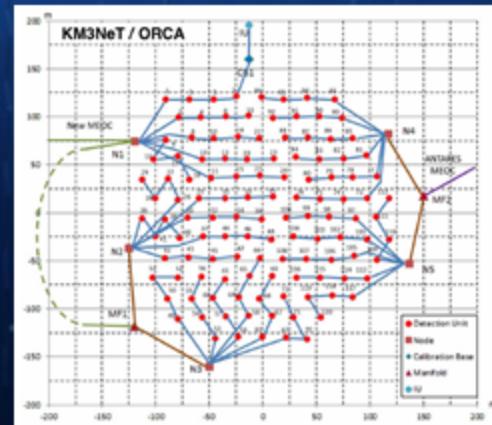
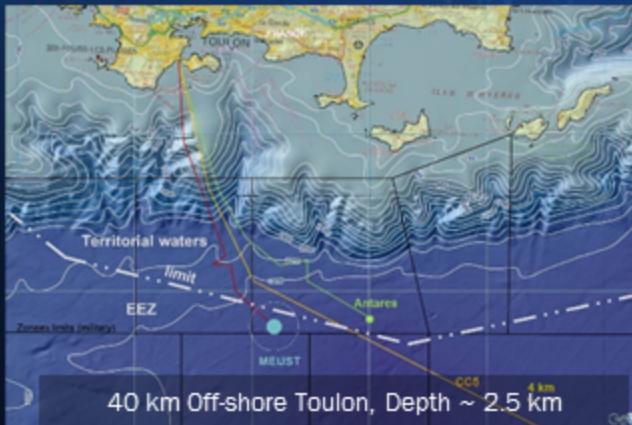
# THE KM3NeT LAYOUT



## ARCA (Astroparticle Research with Cosmics in the Abyss)

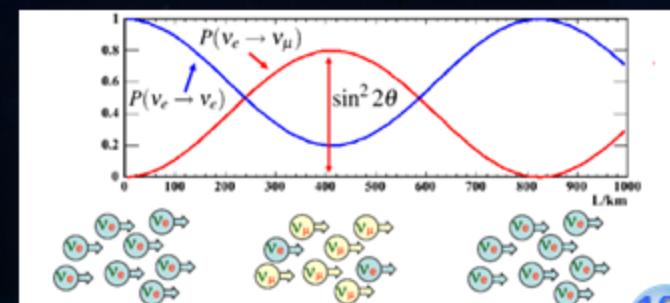
- ✓ High-energy neutrino astronomy and astrophysics (GRB's, AGN's, Blazars, etc.)
- ✓ Multimessenger studies (p, v, GW, γ) and Cosmic ray physics
- ✓ Particle physics with atmospheric muons and neutrinos
- ✓ Tau neutrinos
- ✓ Dark matter
- ✓ Exotics
- ✓ Violation of Lorentz invariance





## ORCA (Oscillation Research with Cosmics in the Abyss)

- ✓ Neutrino Mass Hierarchy and Oscillation parameters
- ✓ Tau Appearance
- ✓ Sterile neutrinos and non-standard interactions
- ✓ Dark matter
- ✓ Supernova neutrinos
- ✓ CP violation
- ✓ Tomography with Cosmics



### Phase-I:

- ✓ Already funded
- ✓ 30 lines or DU's (24 ARCA, 6 ORCA)
- ✓ Proof of feasibility and first science results

### Phase-II (2.0):

#### ARCA:

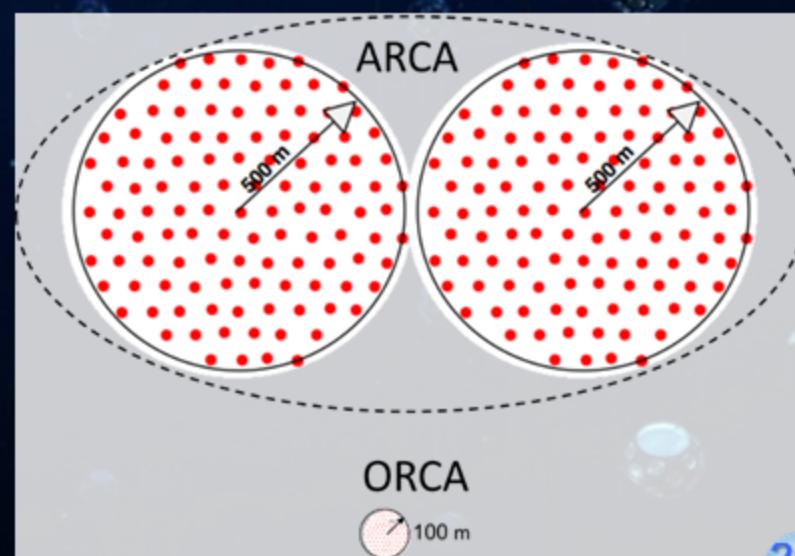
Study IC3 Signal  
2x115 DU's (=2BB)  
Sparse layout  
DU distance = 90 m  
DOM spacing = 36 m

#### ORCA:

Mass ordering (and DM)  
115 DU's (=1BB)  
Dense layout  
DU distance = 20 m  
DOM spacing = 9 m

### Phase-III:

- ✓ 6x115 DU's (ARCA+ORCA) in TOTAL
- ✓ 18 DOM / line; 31 X 3'' PMTs / DOM
- ✓ Neutrino Astronomy including Galactic Sources.



INSTRUMENTED VOLUME →

50 kTon

8 MTon

20 MTon

1 GTon

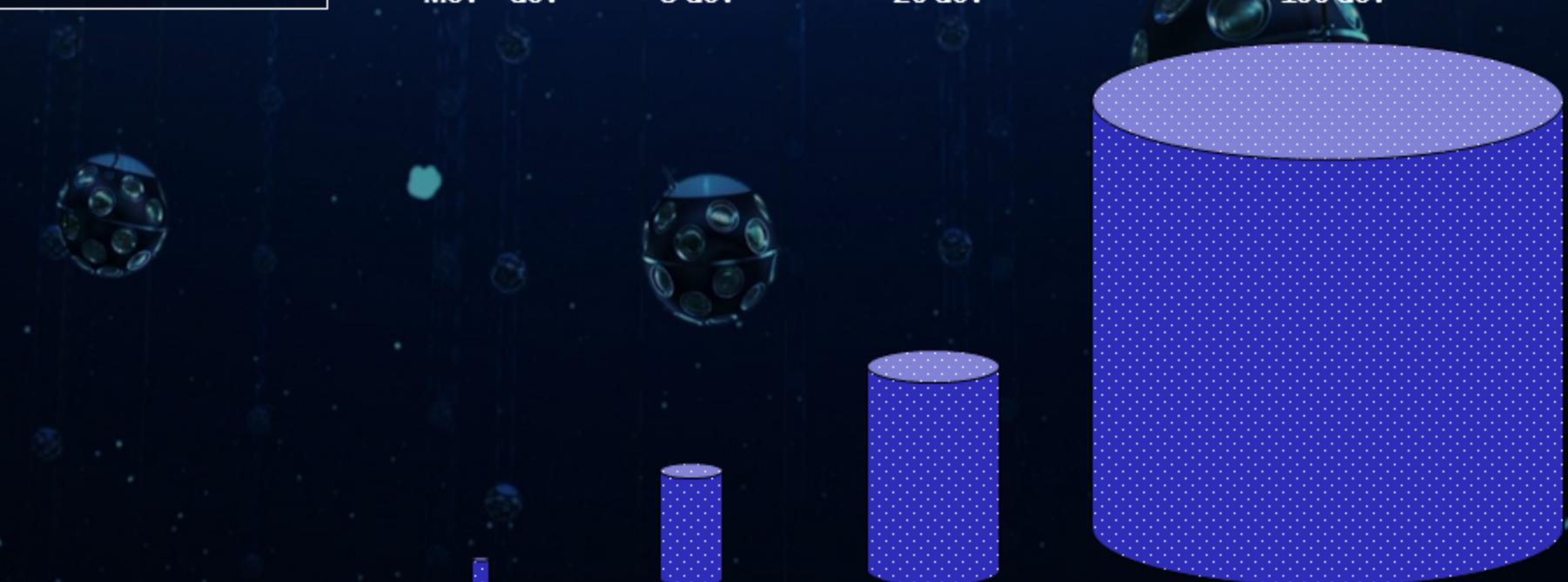
ENERGY THRESHOLD →

MeV – GeV

&gt; 3 GeV

&gt; 20 GeV

&gt; 100 GeV



WATER-BASED CHERENKOV →

Super Kamiokande

ORCA

ANTARES

IceCube, ARCA

THE KM3NeT ROADMAP

Start KM3NeT 2.0

User ports

ERIC

Open access

Confirmation  
of IC signal  
Determination of  
neutrino mass  
ordering

CONSTRUCTION

OPERATION

2016

2017

2018

2019

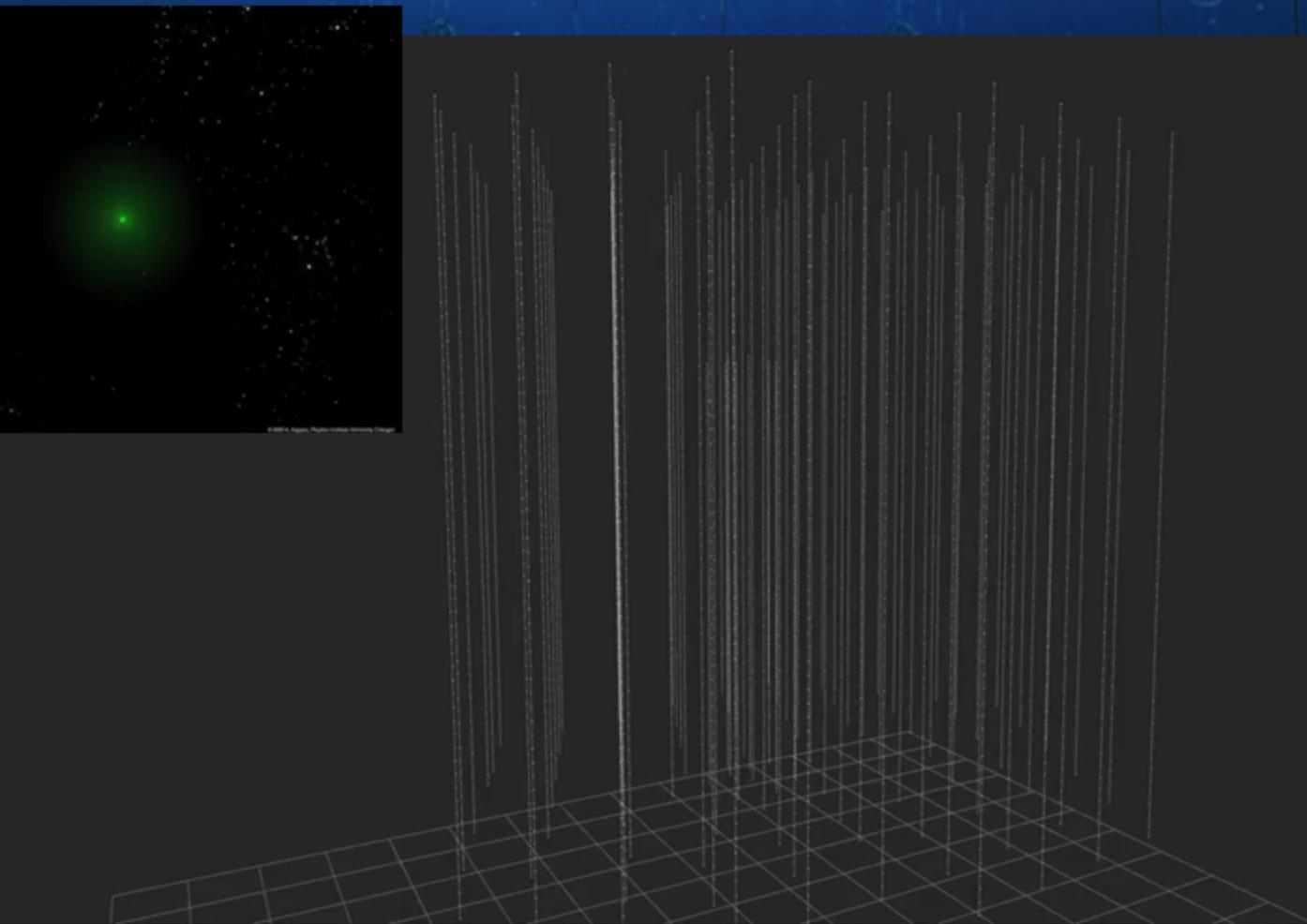
2020

2021

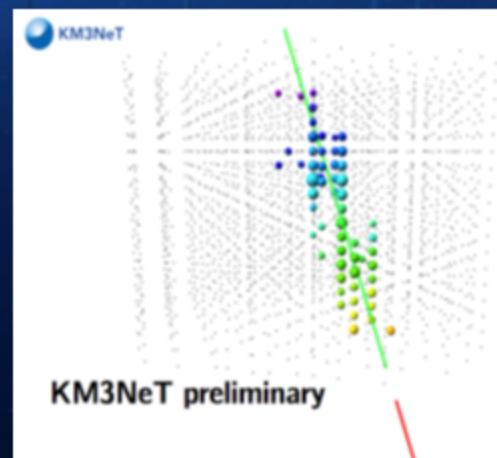
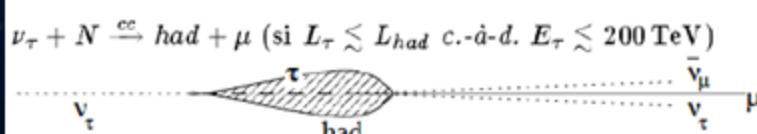
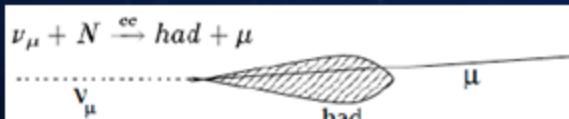
2022

# PART 3

# THE KM3NeT NEUTRINO DETECTOR



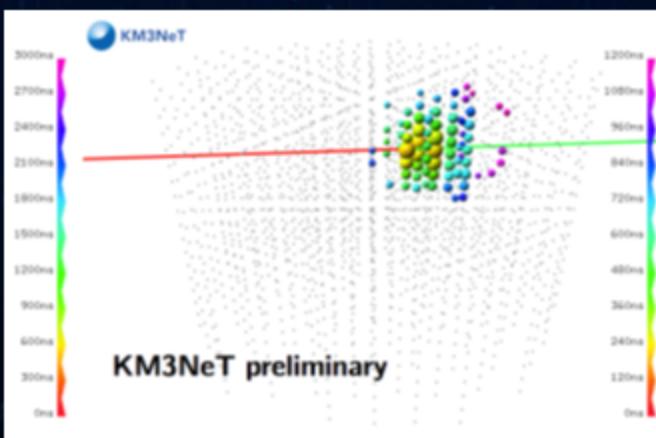
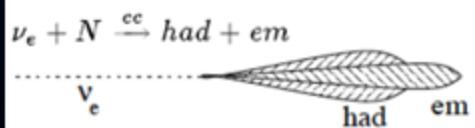
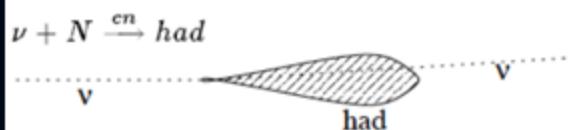
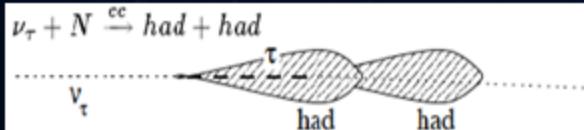
→ TRACK-LIKE: contains both a cascade and one track



Muon track from CC muon neutrinos (golden channel for nu astronomy)

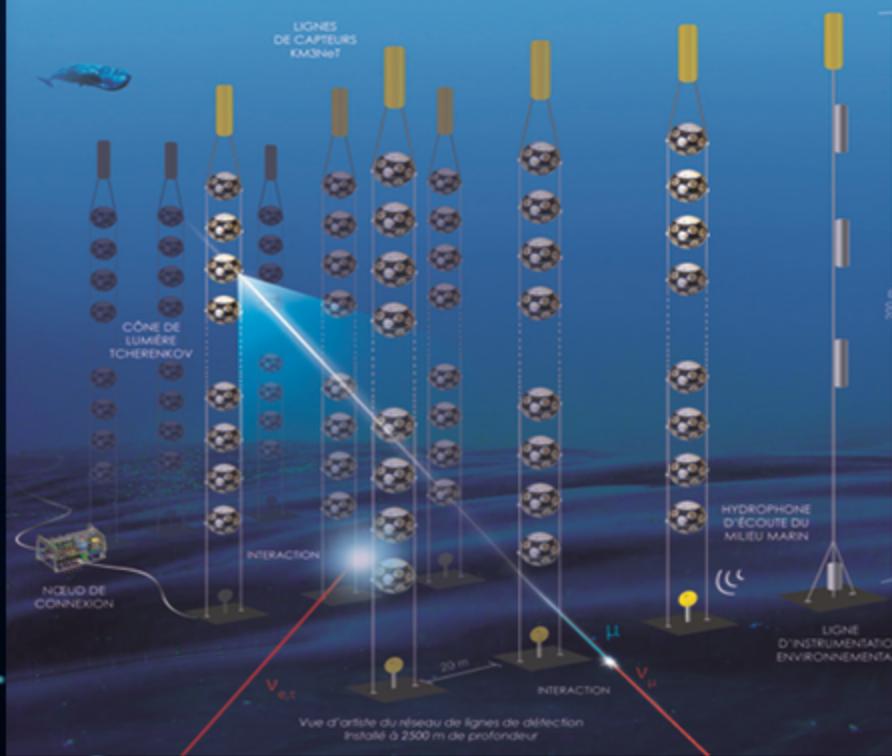
- ✓ Angular resolution  $0.5^\circ/0.1^\circ$  for ice/water.
- ✓ dE/dx resolution factor 2-3.

→ SHOWER-LIKE: no track is identified

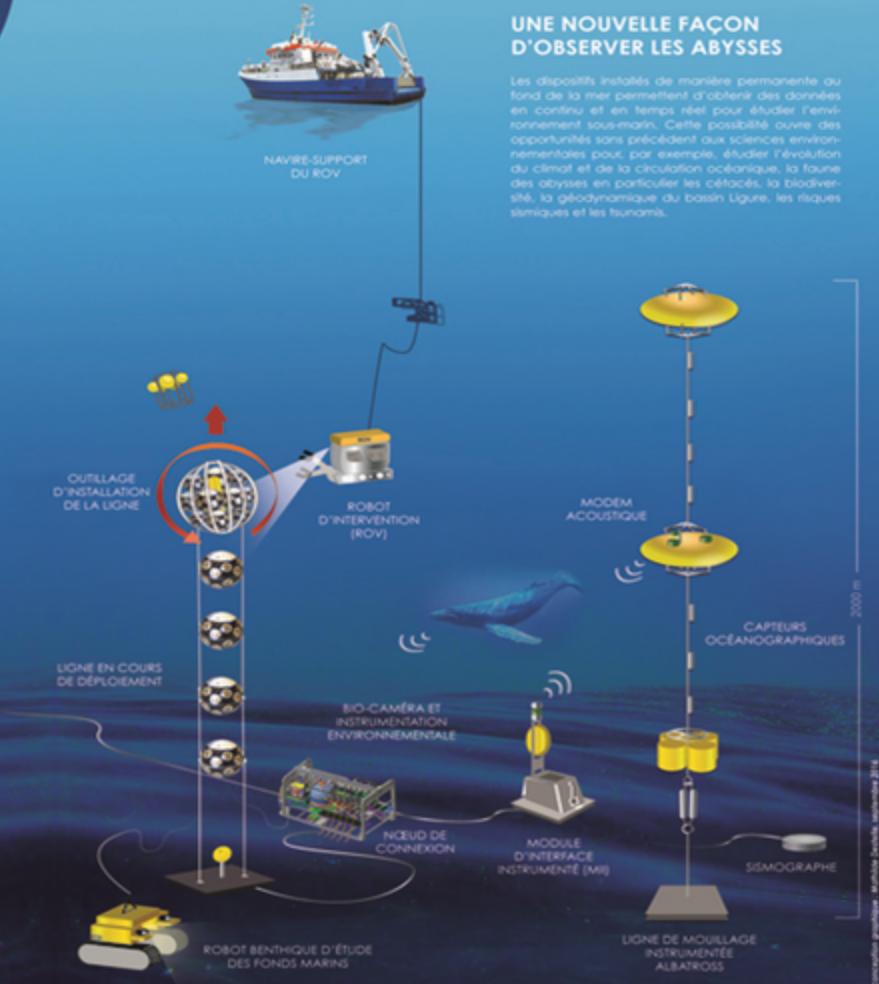


CC electron/tau and NC all flavour

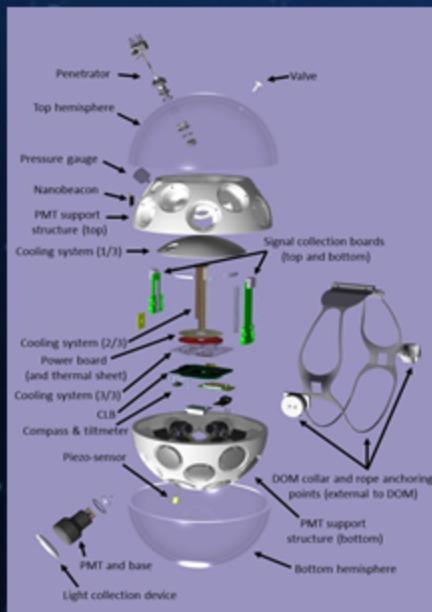
- ✓ 80% of all nu interactions.
- ✓ Angular resolution  $10^\circ/1^\circ$  at 100 TeV for ice/water.
- ✓ Energy resolution  $\sim 10\%$ .



Vue d'artiste du réseau de lignes de détection installé à 2500 m de profondeur



→ 18 DOM integrated on vertical slender strings supported by two parallel Dynema ropes.

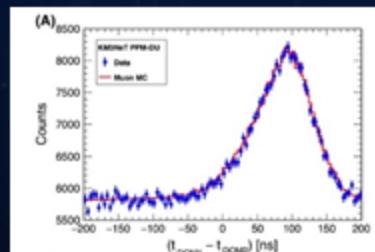


→ The DOM, a high pressure resistant glass sphere housing PMTs, acoustic, light devices and associated electronics.

- ✓ Multi PMT DOMs:  
31 X 3" PMTs + expansion cones
- ✓ Time synchronization:  
White rabbit
- ✓ Optical data transmission:  
Base module with DWDM at string anchor
- ✓ All data to shore concept (optical data transmission):  
Filtering/Trigger on shore in computer farm
- + nodes for long term high bandwidth connection for Earth and Sea sciences

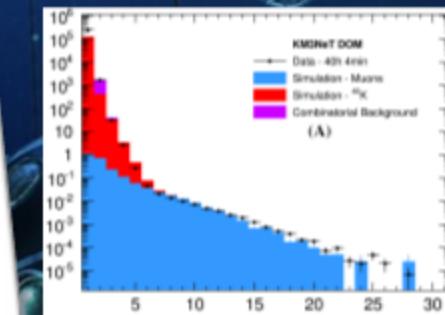
→ Strings arranged on the LOM, mounted on the anchor and ready for deployment

1. Prototype DOM deployed at Antares site April 2013.
2. Prototype DU (three DOMs) deployed in Capo Passero May 2014.

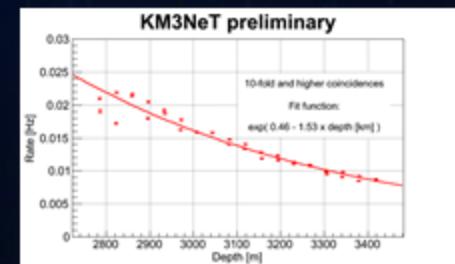


Test of DU structure functionality, Test of intra-DOM and inter-DOM calibration - *Eur. Phys. J. C* (2016) 76:54

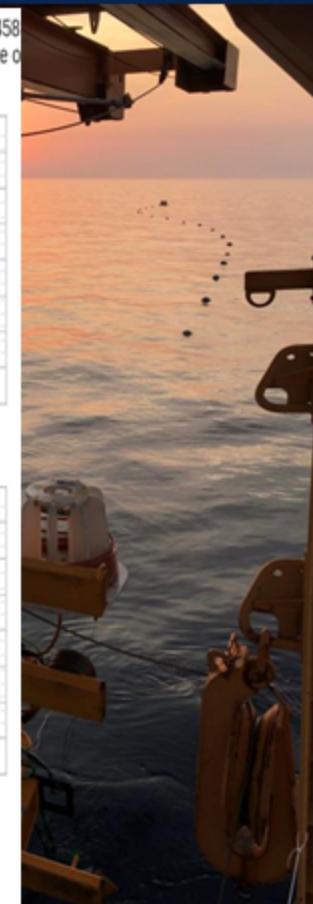
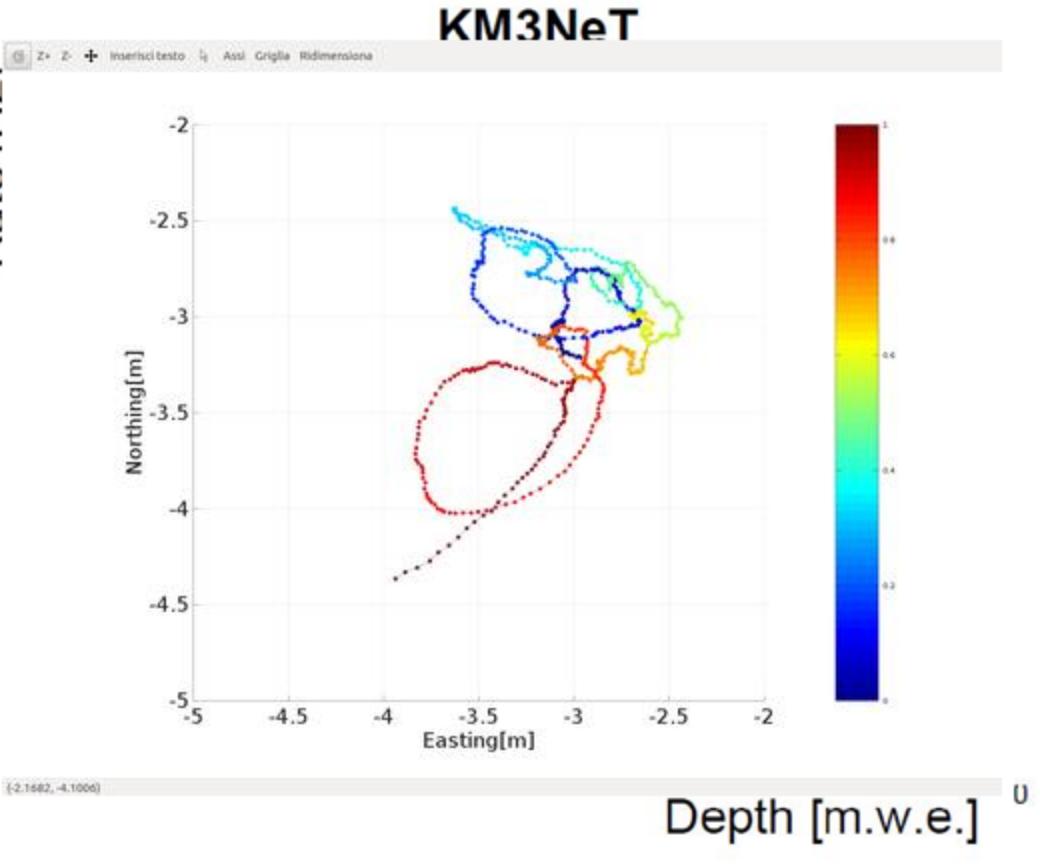
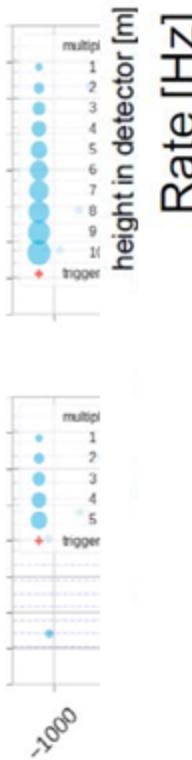
3. First ARCA DU deployed in Capo Passero December 2015.



Test of photon counting capabilities and directional sensitivity of DOM. *Eur. Phys. J. C* (2014) 74:3056



Muon flux dependence on depth, DU calibration, Trigger implementation, Track reconstruction + MC comparison







**WE WANT YOU!**